



16th Patras Workshop on Axions, WIMPs and WISPs, 14-28 June 2021

# Quark Nugget models of dark matter revisited

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# Dark Matter models

## Beyond the Standard Model

- WIMs
- Sterile neutrino
- Axion
- Supersymmetric particles
- ...

## Standard Model based

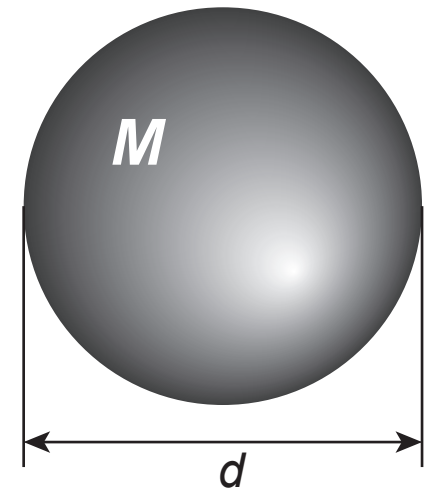
- Black holes
- Strangelets
- Compact Composite objects
- QCD balls
- Axion quark nuggets

# Compact composite objects as DM candidates

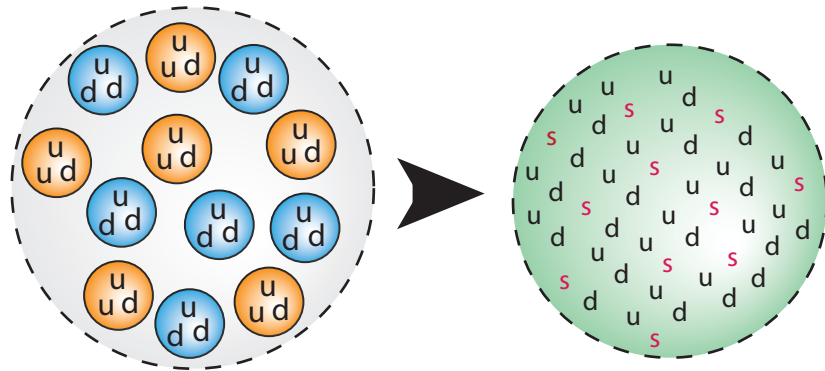
- Can be composed of SM particles (baryons, quarks), do not require new physics
- In spite of strong interaction with visible matter, such objects cannot be detected because of a small ratio of cross section to mass

$$\frac{\sigma}{M} \ll 1 \frac{\text{cm}^2}{g}$$

- Such objects are not excluded by any dark matter detecting experiments and cosmological observations



# Example: Strangelets



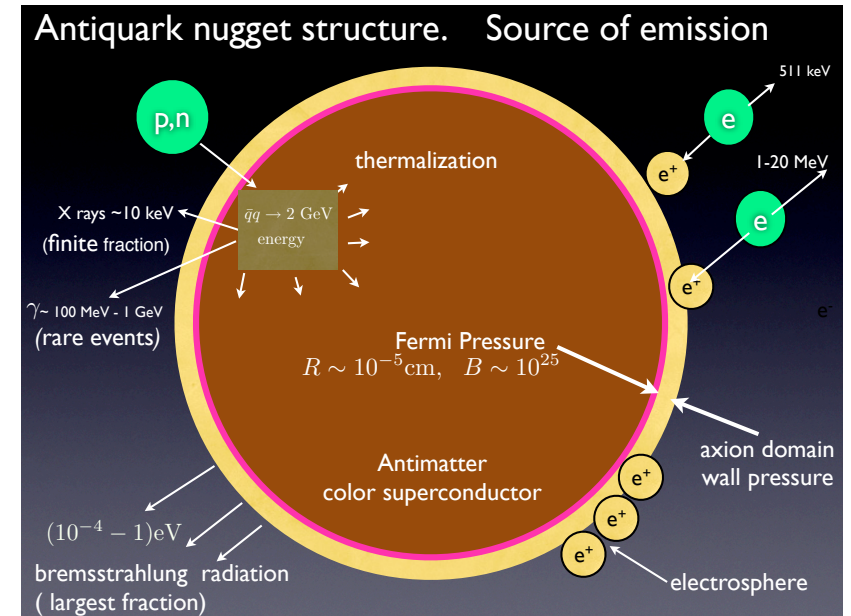
E. Farhi and R. L. Jaffe, Phys. Rev. D 30, 2379 (1984)  
E. Witten, Phys. Rev. D 30, 272 (1984)  
A. R. Bodmer, Phys. Rev. D 4, 1601 (1971)

- **Quark matter**: consists of quarks not bound to baryons
- Presence of **strange quark** can lower the energy of such objects and make them stable
- A stable composite particle which consists of large number of up down and strange quarks – dark matter particle candidate
- Neutron stars could have a quark core with strange quarks (debated)

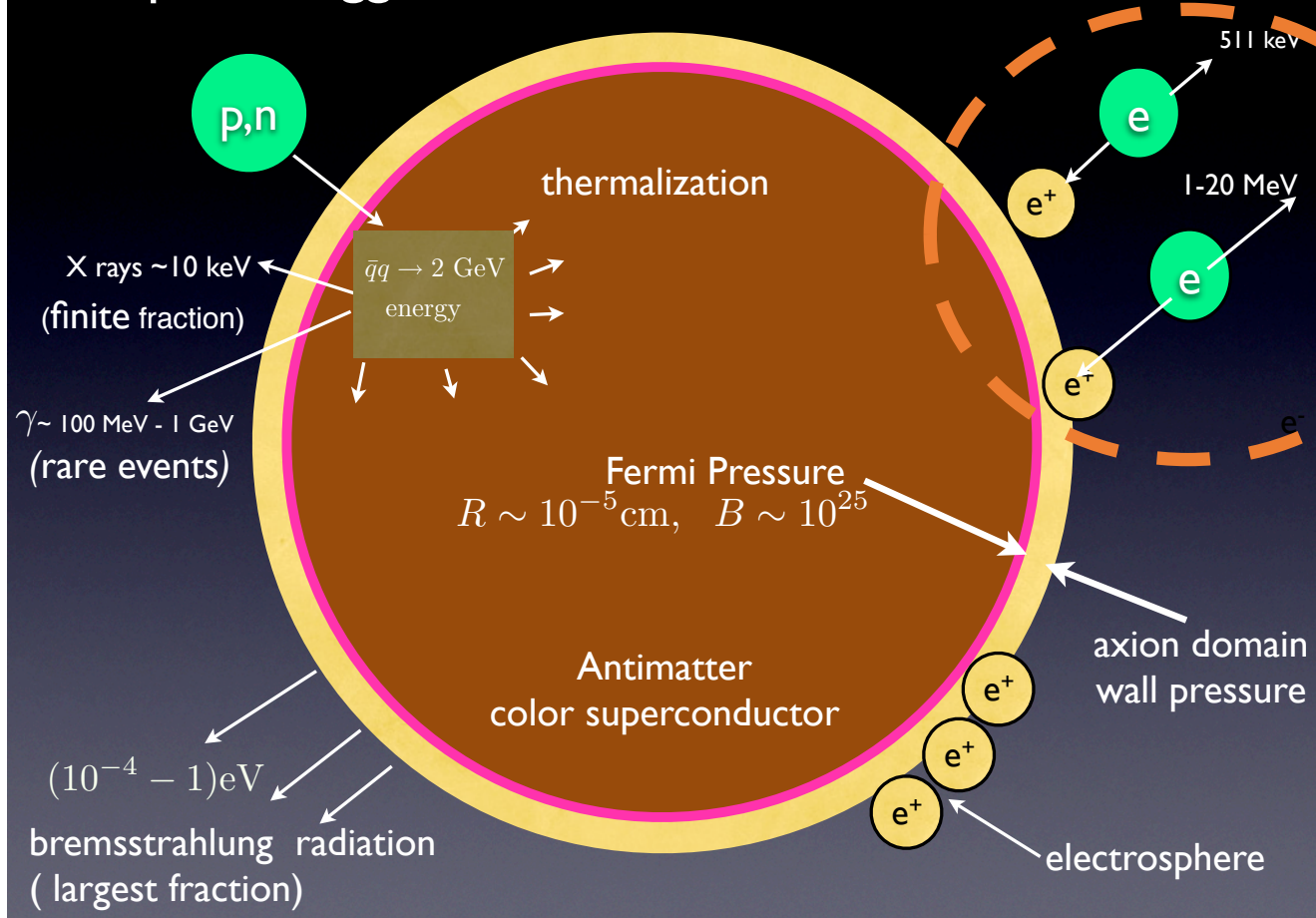
# Axion-quark nuggets, QCD balls, Compact composite objects, etc.

- Quark matter nuggets are composed of large number of quarks surrounded by electron cloud
- Anti-quark nugget consist of large number of anti-quarks, surrounded by the positron cloud
- Both quark and anti-quark nuggets amount to Dark Matter
- Assumption on ratio of abundances anti-quark nuggets : quark nuggets : baryonic matter = 3 : 2 : 1
- Explains matter-antimatter asymmetry in nature: anti-matter is hidden in anti-quark nuggets
- Has radiation which may (potentially) be detected

A. R. Zhitnitsky, JCAP10, 010 (2001)  
And many subsequent papers



# Antiquark nugget structure. Source of emission

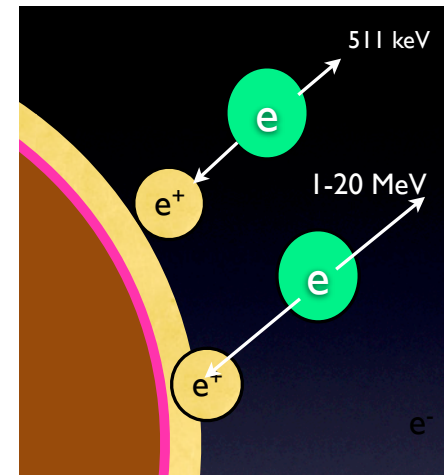


Adopted from the talk by A. Zhitnitsky

- This source of radiation produces very specific 511 keV gamma-photons [D. H. Oaknin and A. Zhitnitsky, Phys. Rev. D 71, 023519 (2005), M. M. Forbes, K. Lawson, A. R. Zhitnitsky, Phys. Rev. D 82, 083510 (2010)]
- Compare with observations of INTEGRAL satellite, B. Teegarden et al., Astrophys. J. 621, 296 (2005).

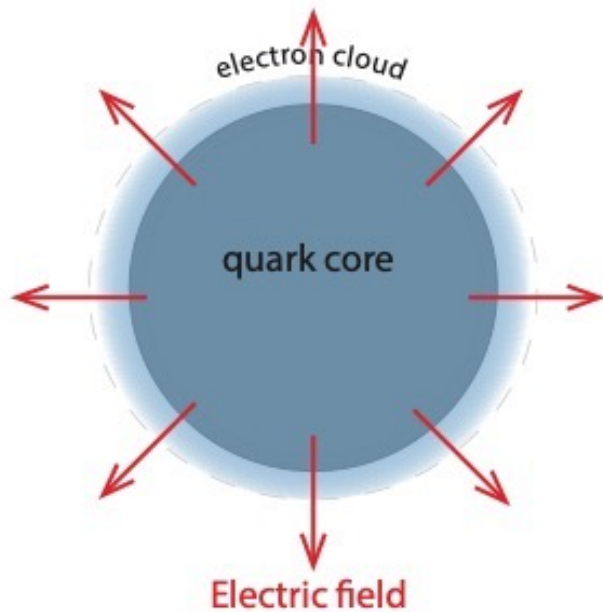
# Our goal

- To accurately consider the electron-positron annihilations in collisions of visible matter with anti-quark nuggets
- To study the radiation in such collisions and compare it with satellite observations of the center of our galaxy
- We consider Quark Nugget model which includes AQN as a particular case

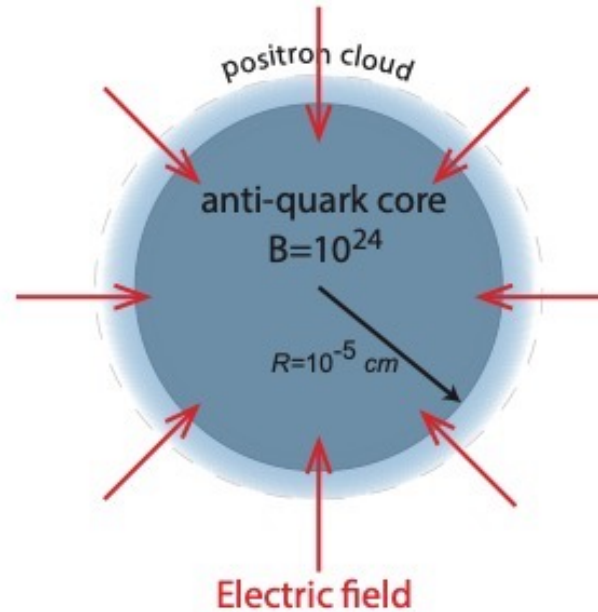


# Quark Nugget model

## Quark Nugget



## Anti-Quark Nugget

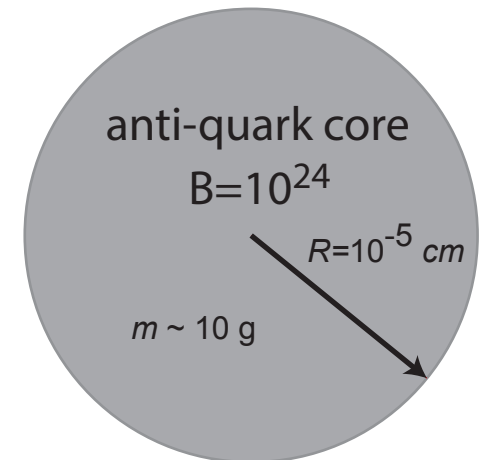


- Anti-QN are 1.5 more abundant
- Consist of anti-matter
- May be detected when collide with visible matter



# (anti)Quark core

- May be considered as a ball or characteristic **radius**  $R=10^{-5}$  cm
- Characteristic **mass** is of order 10 g
- **Density**  $n_0(r)$  is 1.5-3 times nuclear matter density
- **Baryonic charge** varies  $10^{23} - 10^{28}$ . We adapt  $B=10^{24}$ .
- **Electric charge** of quarks  $Z$  is unknown. It depends on the equation of state of the quark matter which is not known. If the quarks are in the color-flavor-locked phase, the electric charge is localized at the boundary,  $Z=0.3 B^{2/3}$ . If the quarks are in the 2CS phase,  $Z \sim 0.001 B$ .
- Huge negative charge of anti-quark matter attracts **positrons** => **positron cloud**



# Positron cloud

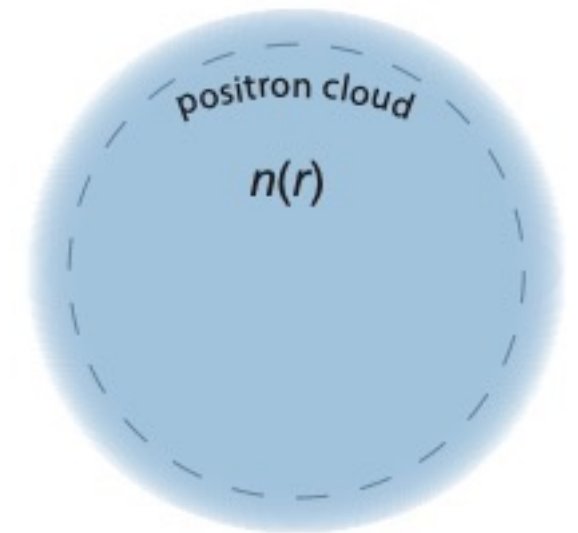
- In the Thomas-Fermi model **charge density**  $n(r)$  which may be found as a solution of Poisson equation with  $n_0(r)$  being the quark core charge density

$$\Delta\varphi = 4\pi e[n(r) - n_0(r)]$$

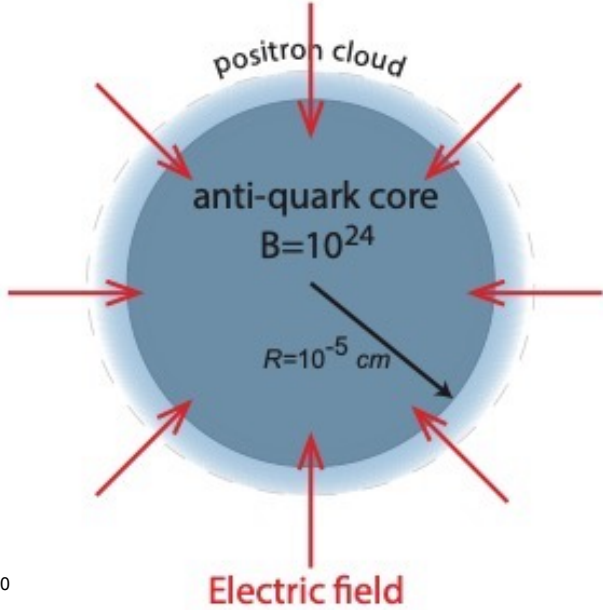
- The positron density, together with the quark core charge density, create the **electric field**

$$E = -\nabla\varphi$$

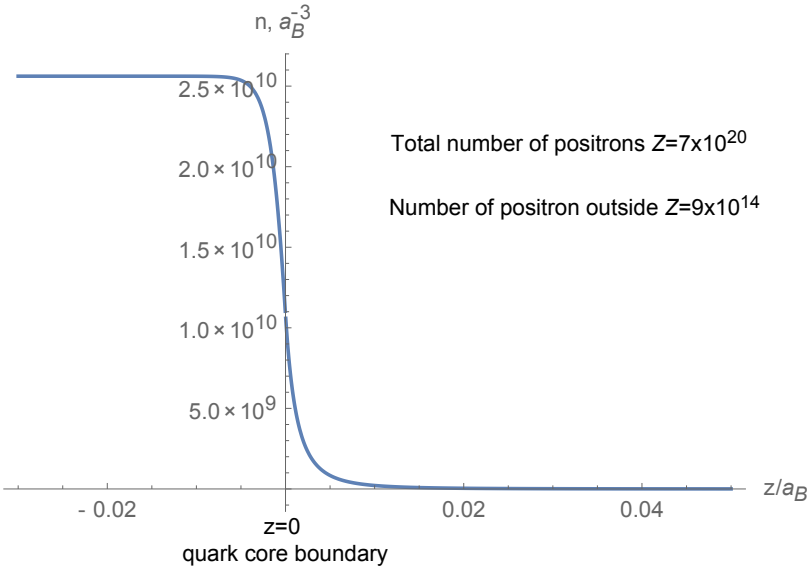
- Given the positron density and the electric field, we study the **annihilation probability** for incident electrons, atoms and molecules.



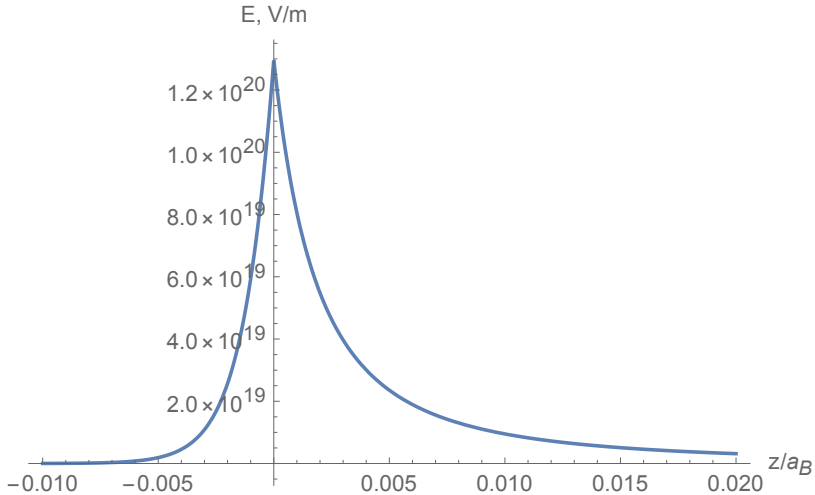
# Properties of positron cloud



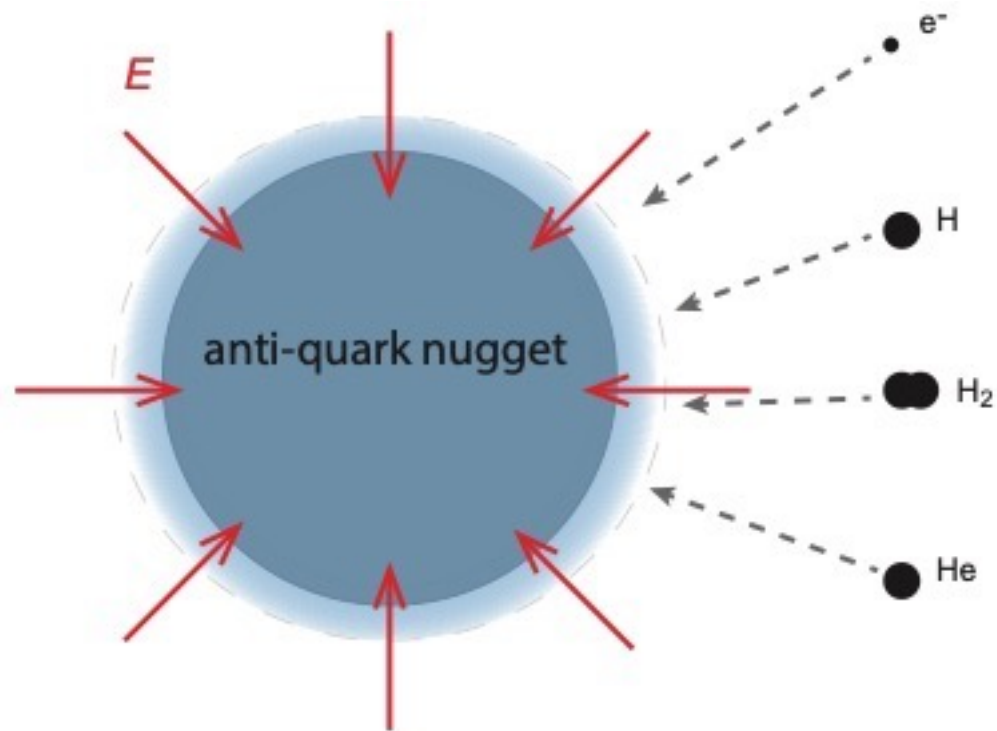
Charge density



Electric field near quark core boundary

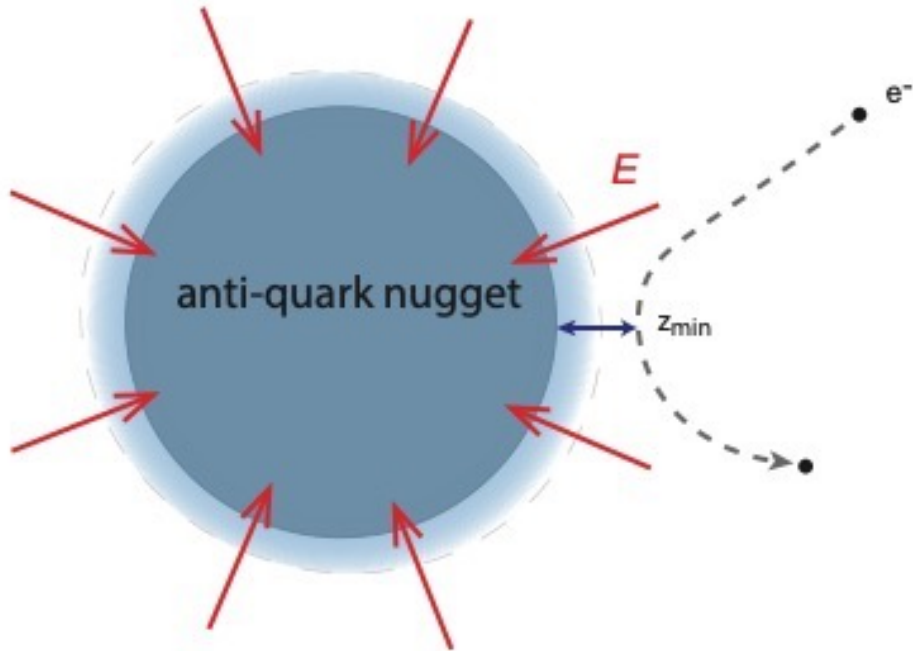


# Matter scattering off the Quark Nuggets



- Assume that these particles have characteristic velocities  $v=10^{-3} c$  (virial velocities in interstellar medium)
- **Problem:** probability of electron-positron annihilation with emission of 511 keV photons?

# Electron scattering



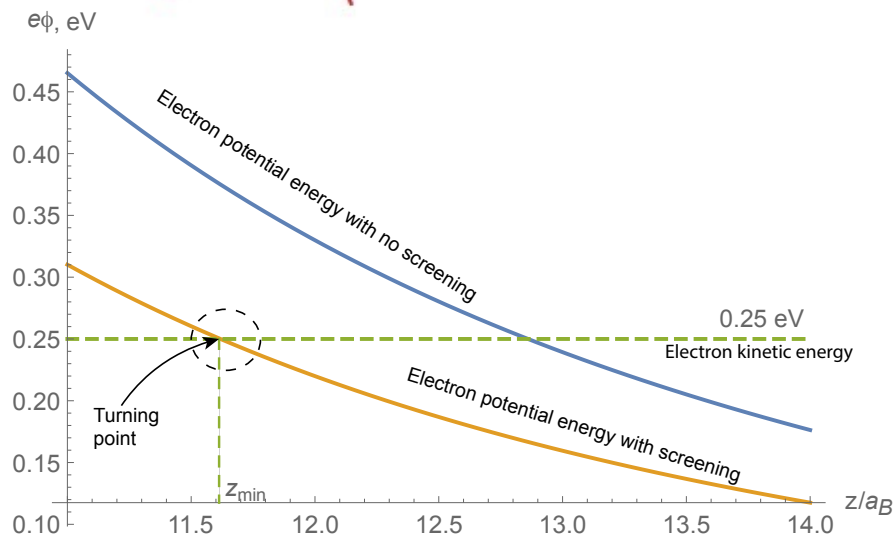
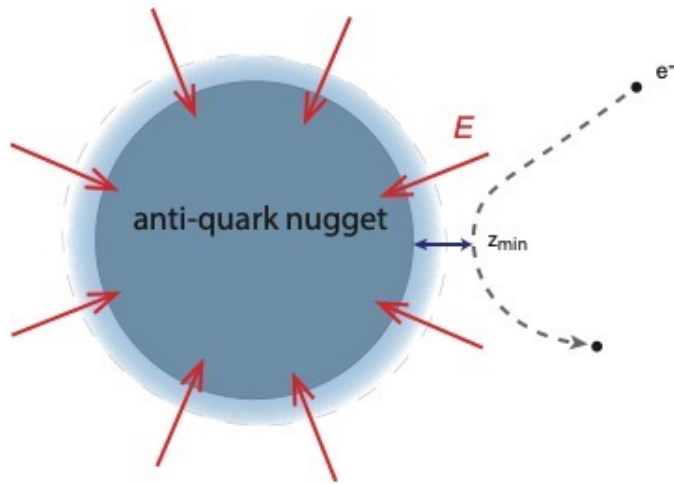
- Electron is **repelled** by the QN electric field
- **Debye screening** should be taken into account
- Electron has **effective charge** when moving through the positron cloud

$$q_{\text{eff}} = \frac{1}{3} \frac{e\varphi + 2m}{e\varphi + m}$$

- In the **non-relativistic case** (far from QN core)  $e\varphi \ll m$

$$q_{\text{eff}} = \frac{2}{3}$$

# Electron scattering



- Electron can approach QN core to characteristic distance  $z_{\min}=11-12 a_B$

- Direct annihilation **cross section**

$$\sigma \approx \pi r_e^2 c/v$$

- **Annihilation probability**

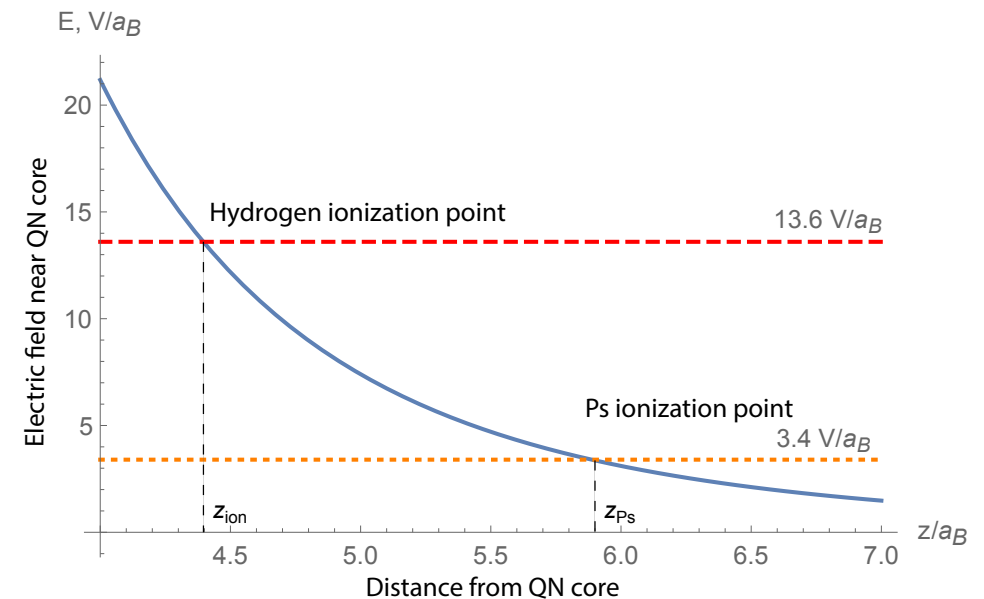
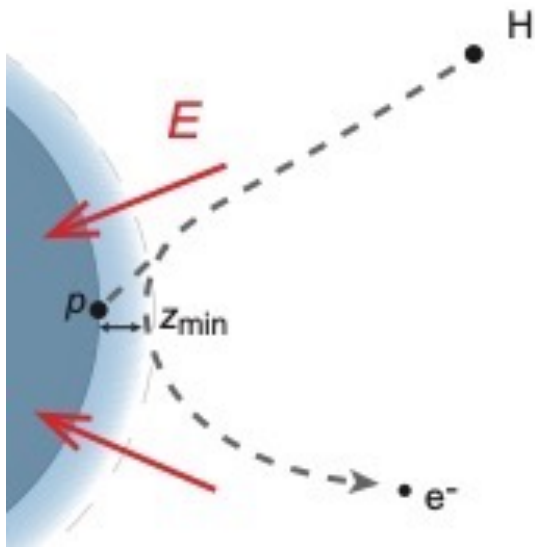
$$P = 1 - \exp \left[ -2\sigma \int_{z_{\min}}^{\infty} n(x) dx \right]$$

$$P \approx 10^{-9}$$

- Thus, this process **plays minor role**

# Hydrogen & helium scattering

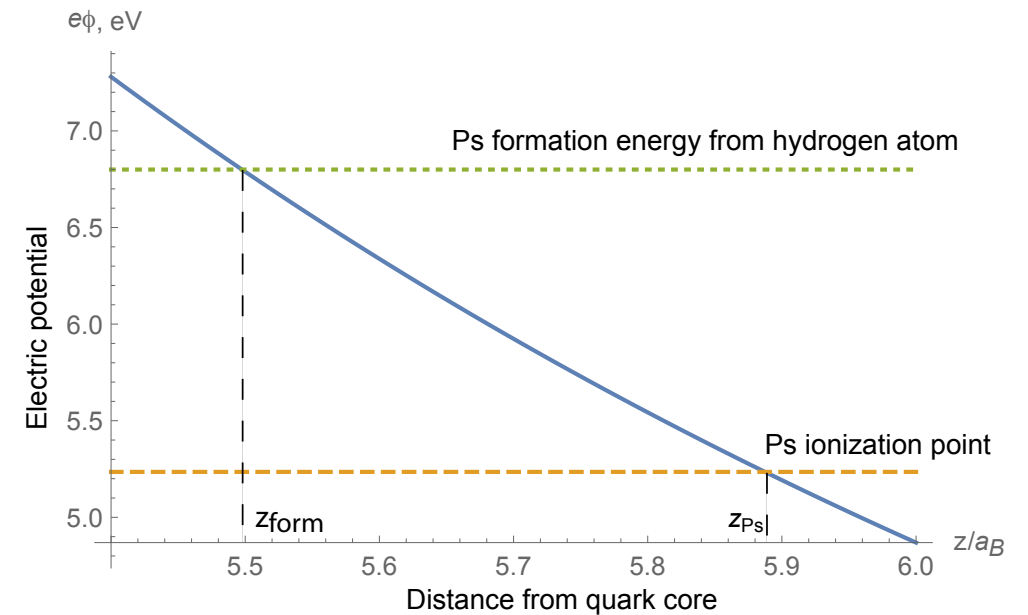
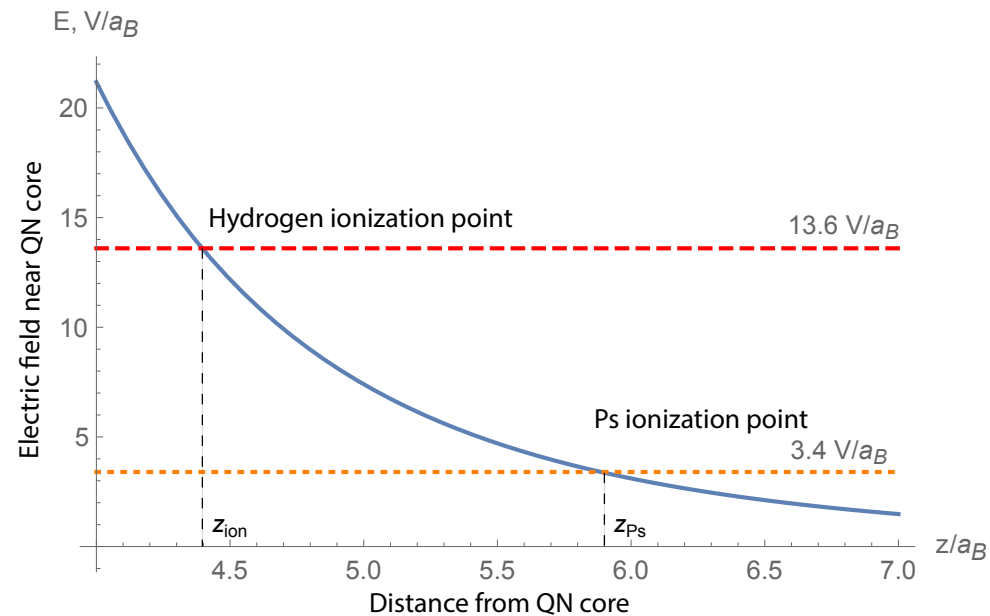
- H, H<sub>2</sub>, He or He<sup>+</sup> can approach quark core up to  $z_{\text{ion}}=3-7 a_B$ .
- Neutral atoms and molecules are attracted due to polarization potential
- Electric field ionizes these atoms and molecules,  $E=2-30 \text{ V}/a_B$
- Proton falls on the quark core while the electron escapes



# Ps formation in Hydrogen & helium scattering

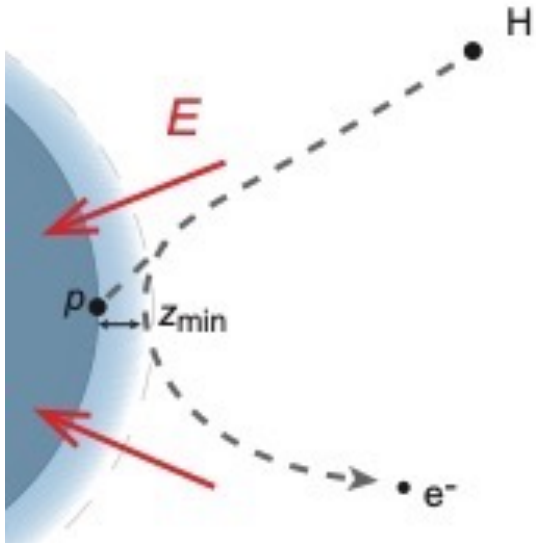
- Positronium is NOT stable at such a strong electric field. In addition, there is strong Pauli blocking at this high positron density. Positronium cannot be formed

$$z_{\text{formation}} < z_{\text{Ps decay}}$$





# Direct positron annihilation in Hydrogen & helium



- Cross section of direct positron annihilation on atoms  $\sigma \approx Z_{\text{eff}}\pi r_e^2, 4 < Z_{\text{eff}} < 15$

- Probability of direct electron-positron annihilation is

$$P = 1 - \exp \left[ -2\sigma \int_{z_{\text{min}}}^{\infty} n(x) dx \right]$$

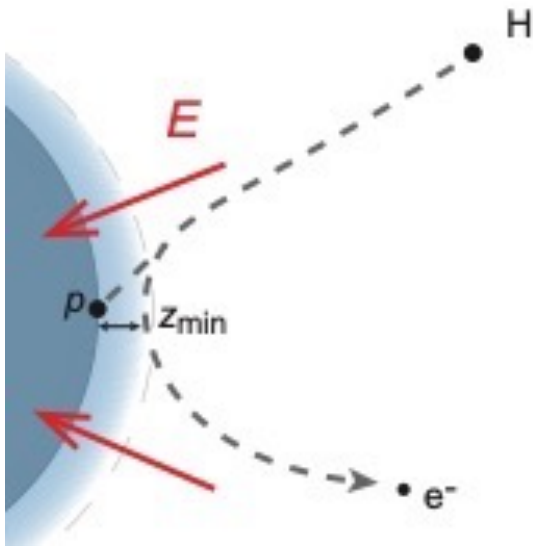
$$P \approx 10^{-5}$$

- Only photons with 511 keV energy are emitted.
- **NO photons with energies 1-20 MeV from electron-positron annihilation since there are no electrons at small z.**

## Our conclusions:

- Positronium atoms **cannot** be created in the positron cloud in the vicinity of quark core
- Electron-positron annihilation in the positron cloud **cannot** produce MeV-energy photons. **MeV-range photons may be emitted as the bremsstrahlung radiation of proton falling on the quark core and its annihilation**
- **A different mechanism of emission of 511 keV photons should be explored**

# Alternative mechanism



- **Proton** collides with the quark core and annihilates
- Charge of quark core is reduced
- Temperature of positron cloud rises
- Outer positrons can “evaporate”
- Thus, (anti)quark nuggets are **sources of positrons** in galaxy center
- These positrons annihilate with atoms and molecules and produce 511 keV photons. Flux of these photons may be compared with INTEGRAL satellite observations.

# Summary

- In collisions of Quark Nuggets with atoms and molecules the **positronium cannot be created** in the positron cloud close to quark core.
- Only **direct** electron-positron annihilation in atoms is possible, the probability is of order  $P=10^{-5}$ .
- Electron-positron and positron-atom collisions **cannot** produce photons with energies in the MeV range. However, such photons may be produced by bremsstrahlung and proton annihilation.
- However, **proton** can annihilate in the collision with the quark core. This reduces quark core charge and raises the temperature of the positron gas. This makes the positrons to evaporate.
- Thus, QNs may be a source of positrons which create the excess of 511 keV photons observed in the center of the galaxy.
- **More details in forthcoming preprint [arXiv:2106.xxxxx](#).**