

# Lab activities @LNF Status Report

A. Calcaterra, R. de Sangro, G. Felici,  
G. Finocchiaro, P. Patteri, M. Piccolo  
INFN LNF

XII SuperB General Meeting

DCH-II parallel session  
Annecy, 17 March 2010

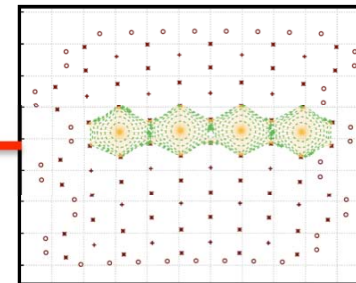
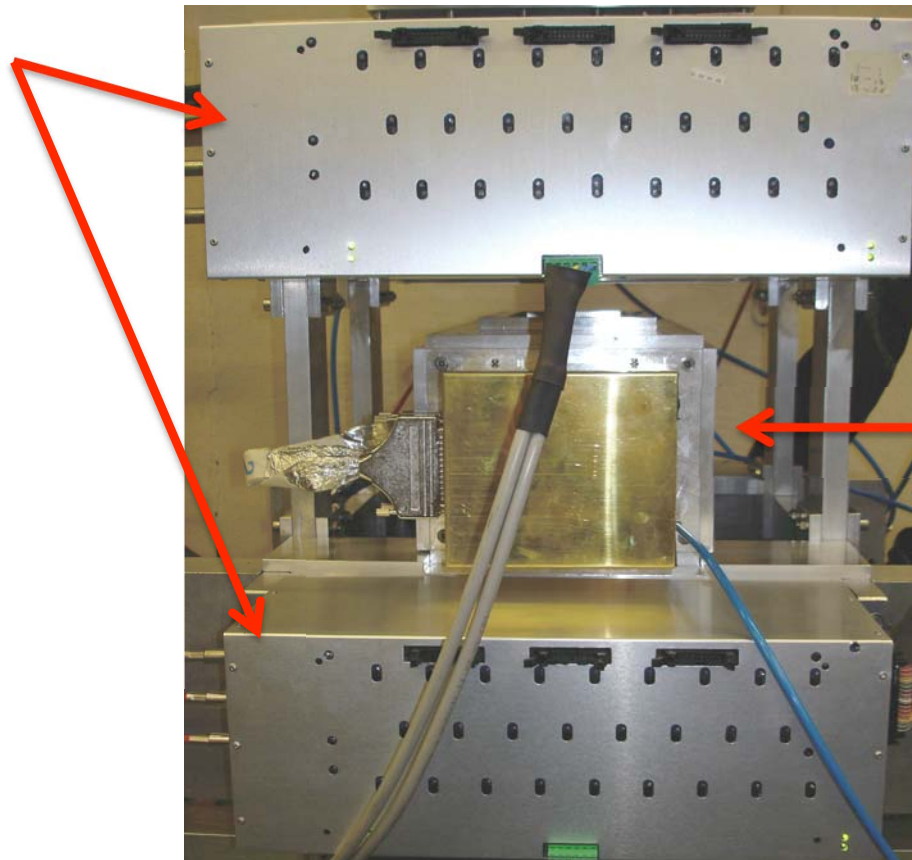
# Experimental Setup

## LST Telescope:

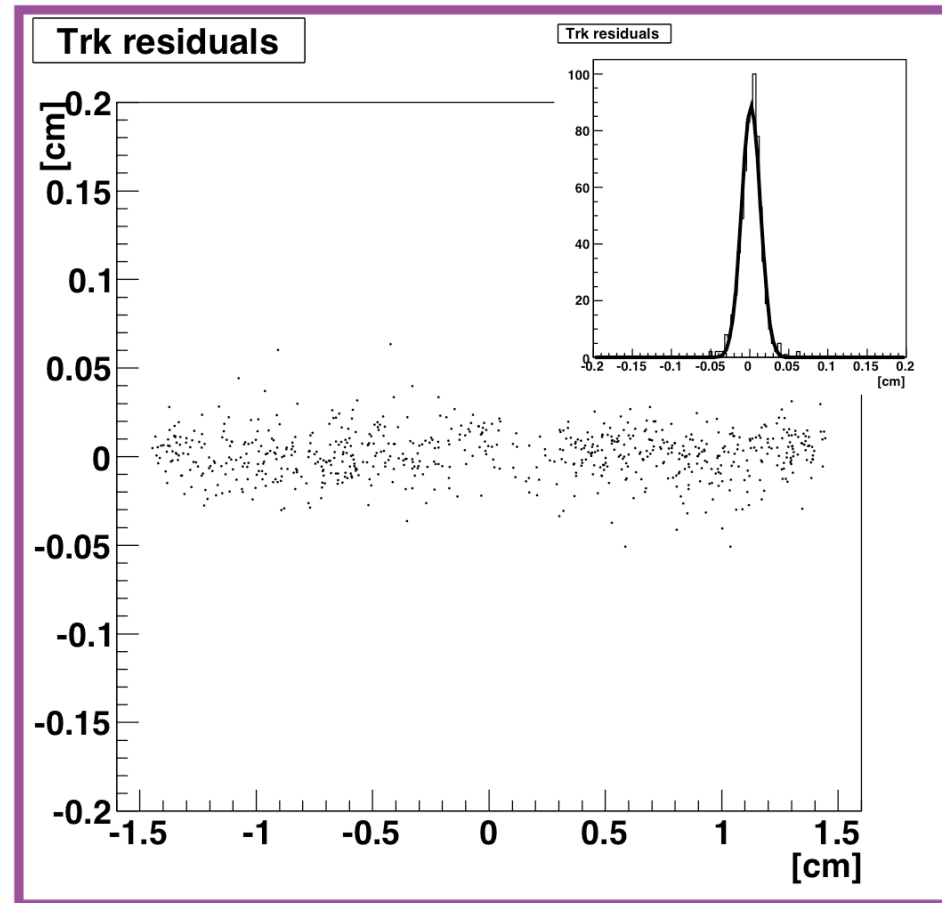
- Two identical assemblies of 26 tubes each
- 3 cm diameter, 100  $\mu\text{m}$  wires
- 40%-60% Ar- $\text{iC}_4\text{H}_{10}$  mixture

## Prototype 1

- 6x4 hexagonal cells à la *BABAR*
  - Guard wires to ensure uniformity of electric field among cells
  - Aluminized mylar windows on entrance-exit faces

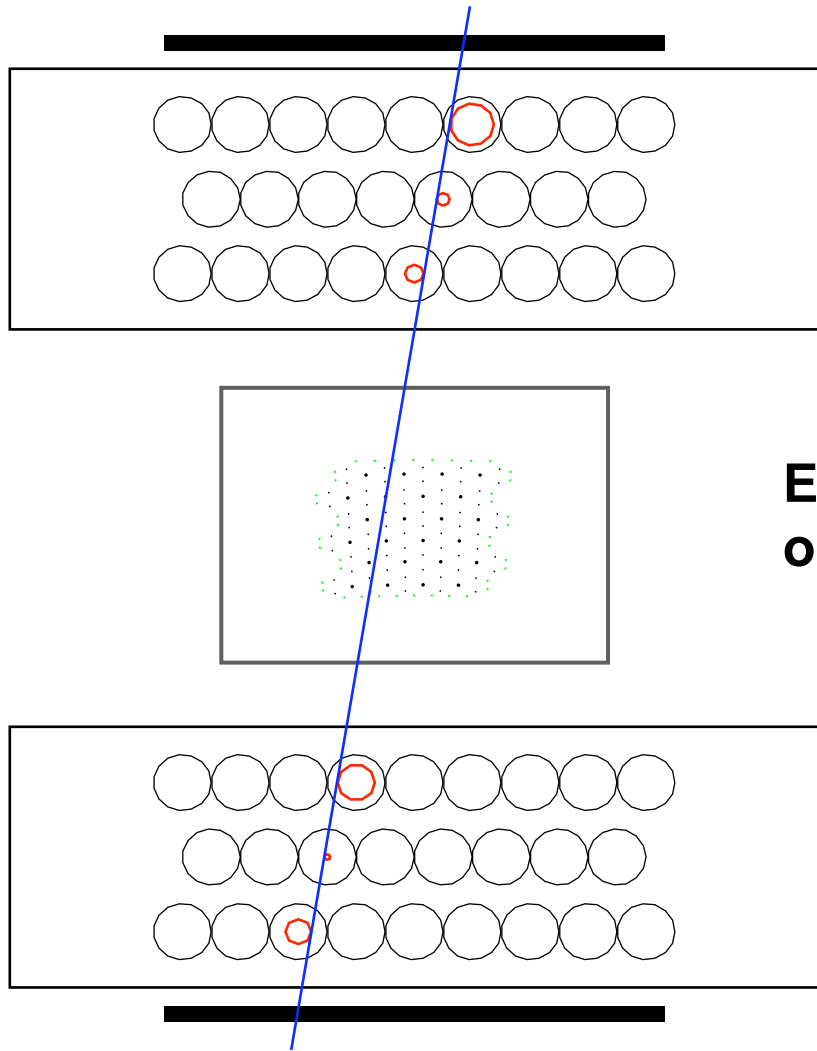


# Track Impact Parameter Resolution



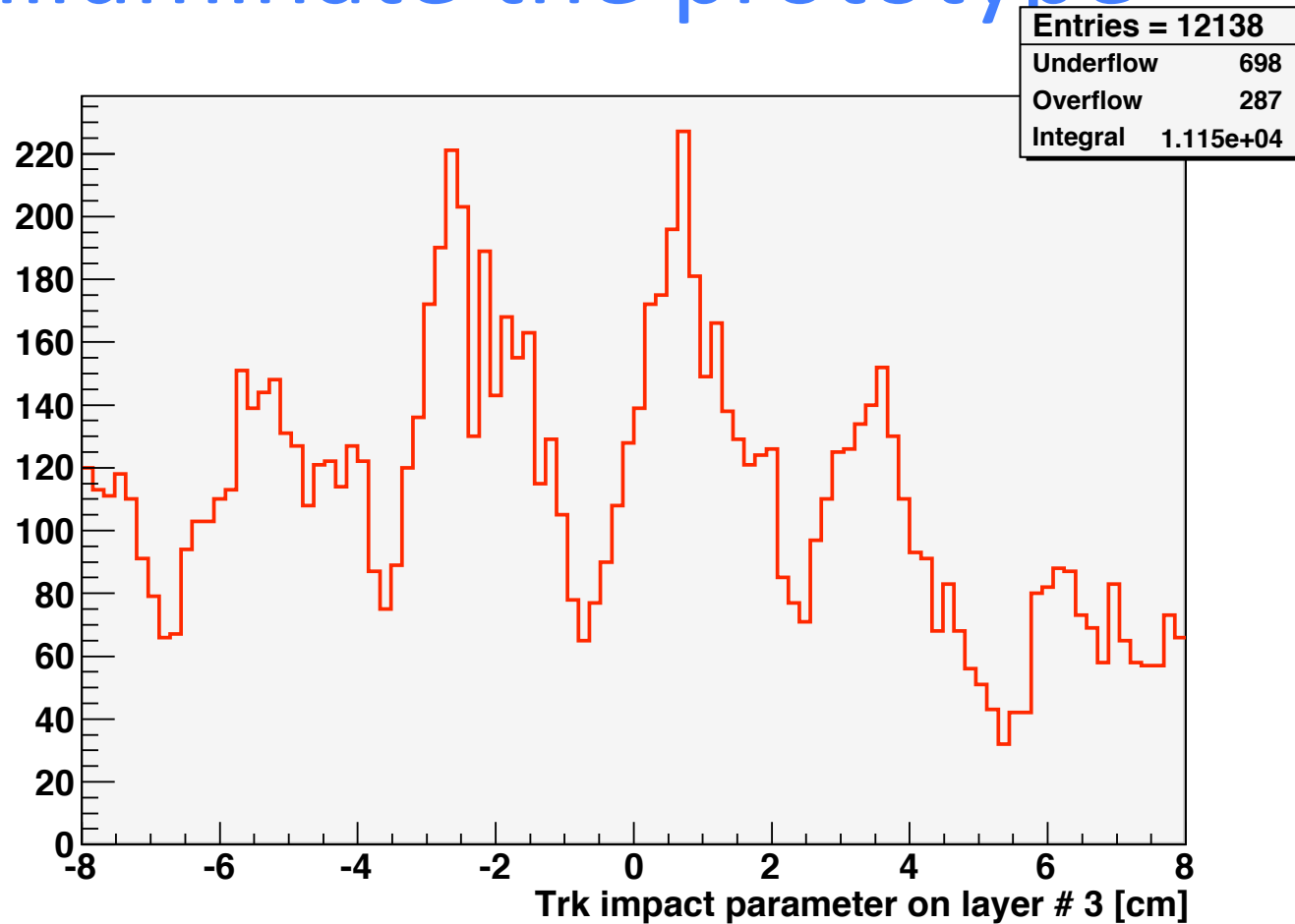
Telescope single tube spatial resolution  $\sim 100\mu\text{m}$

# Tracks extrapolated in Proto1

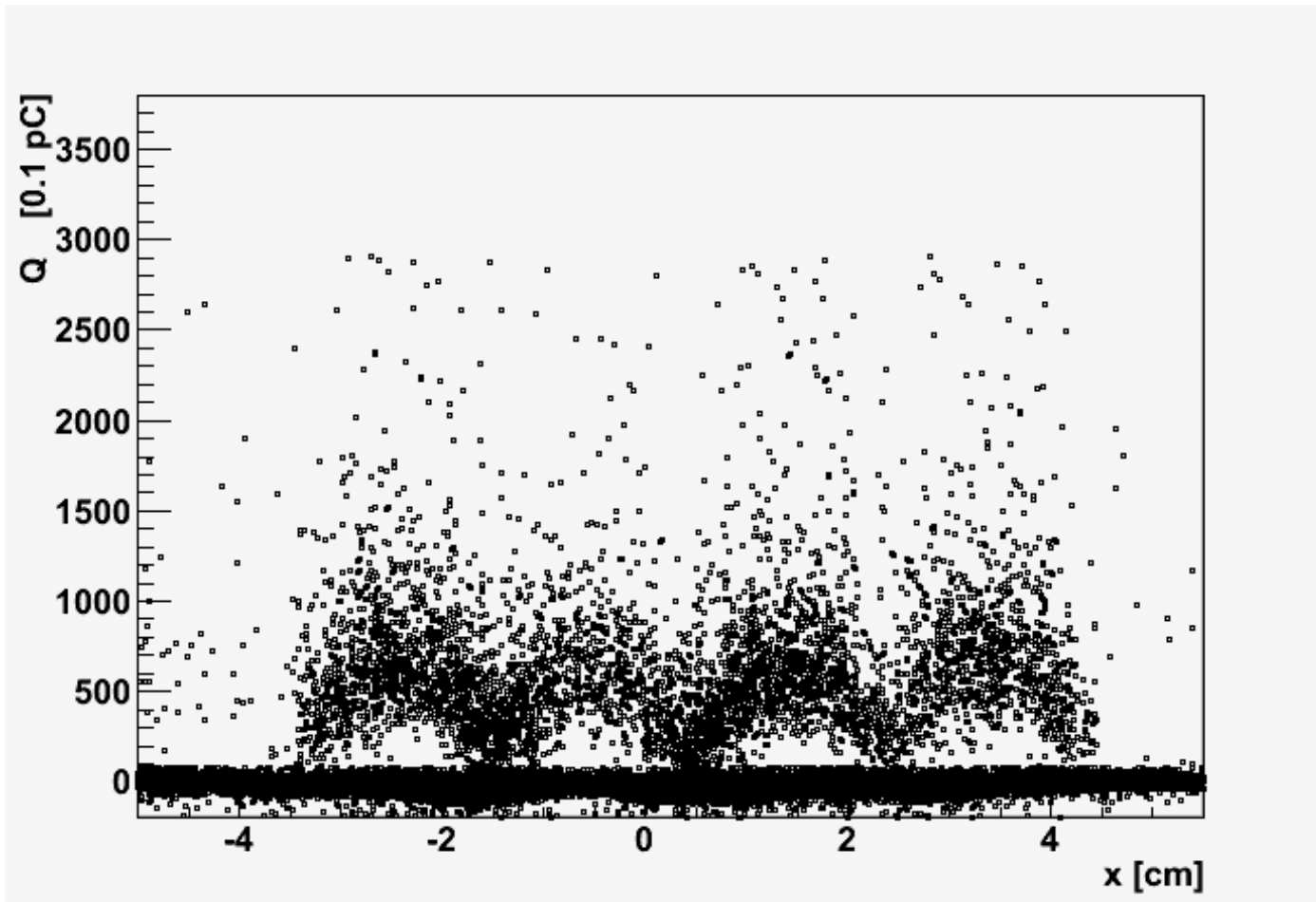


**Expected extrapolation accuracy  
on drift chamber prototype  $\leq 70\mu\text{m}$**

# How tracked cosmic rays illuminate the prototype

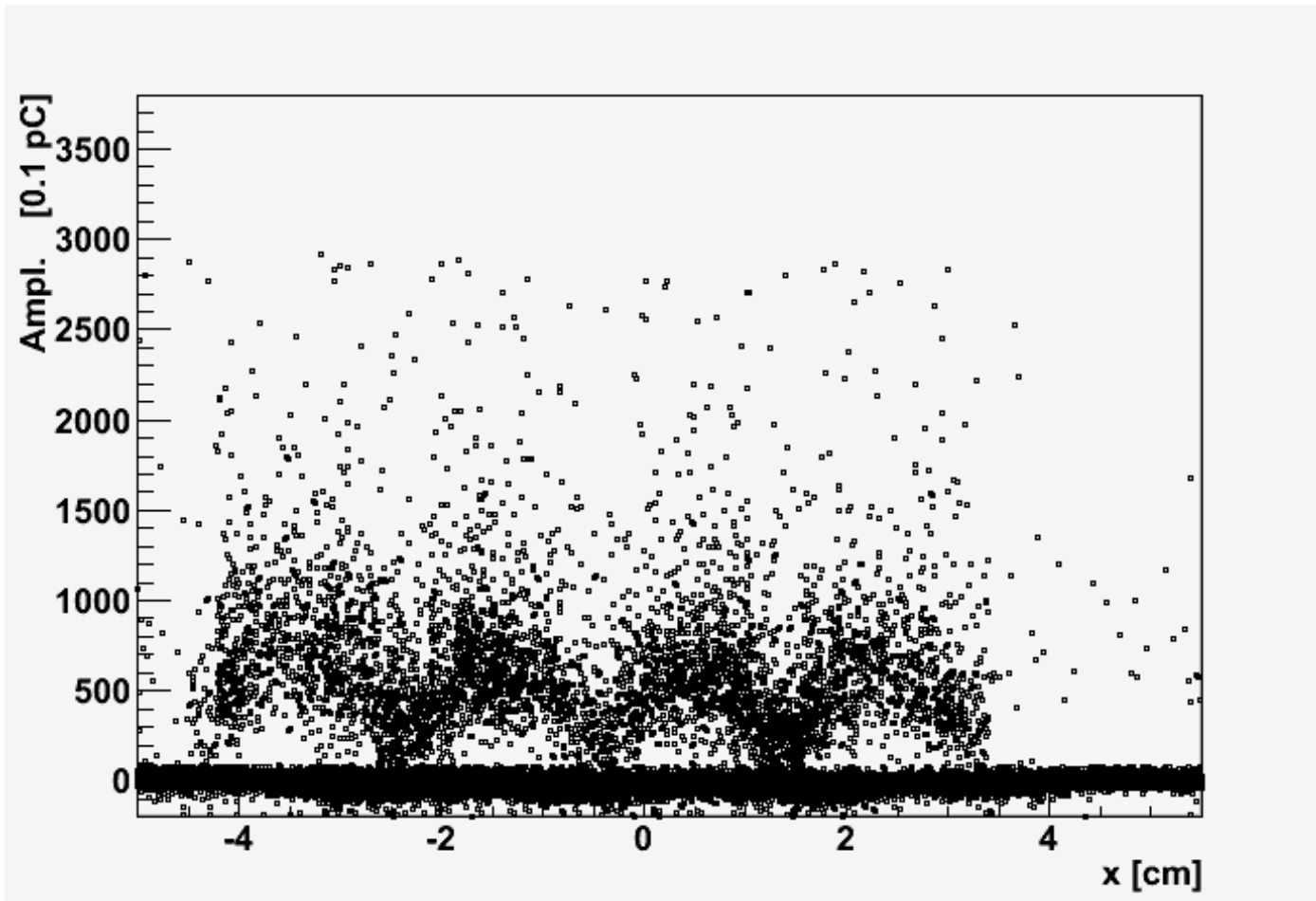


# Charge profile in a layer



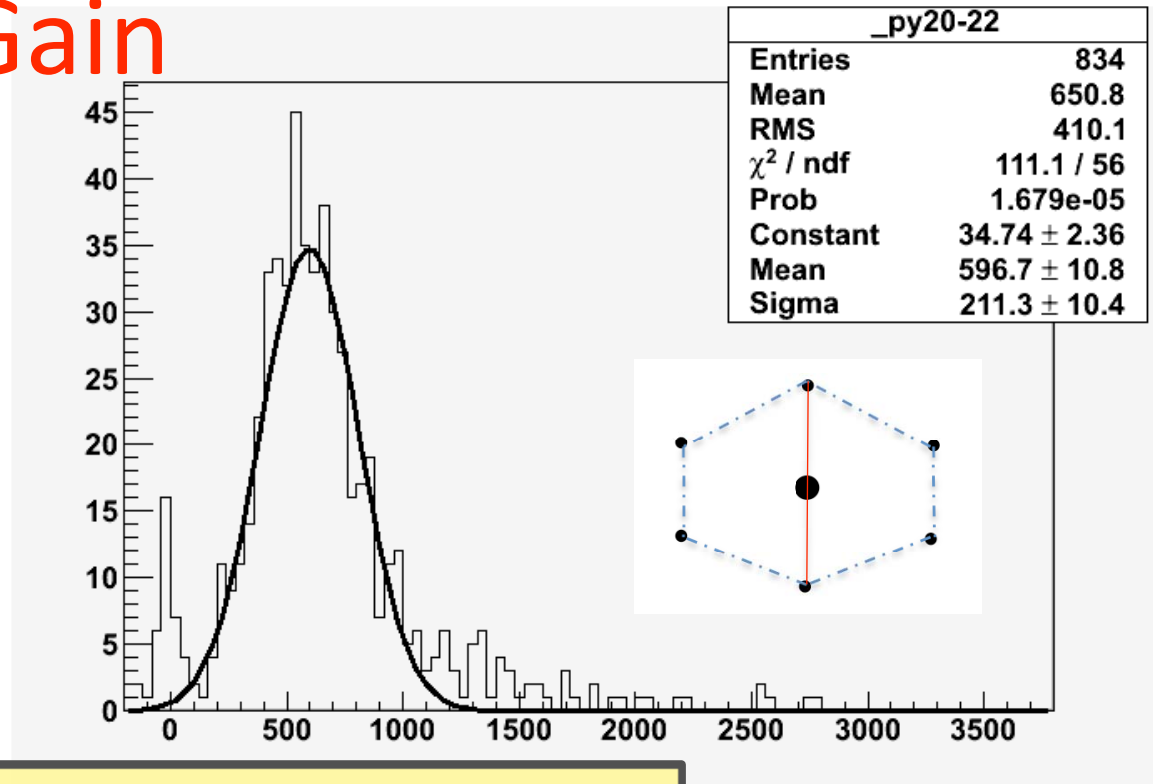
80%He-20% $i\text{C}_4\text{H}_{10}$  @ HV=1960V

# Charge profile in the next layer



80%He-20%iC<sub>4</sub>H<sub>10</sub> @ HV=1960V

# Gas Gain



Charge spectrum (ADC counts) for tracks passing around the center of a cell ( $L \sim 1.2\text{cm}$ ) -

80%He-20% $i\text{C}_4\text{H}_{10}$	@ HV=1960V	$G_{\text{gas}} \sim 3 \cdot 10^{-4}$
90%He-10% $i\text{C}_4\text{H}_{10}$	@ HV=1710V	$G_{\text{gas}} \sim 6 \cdot 10^{-4}$
63%He-37% $\text{CH}_4$	@ HV=2410V	$G_{\text{gas}} \sim 8 \cdot 10^{-4}$
63%He-37% $\text{CH}_4$	@ HV=2510V	$G_{\text{gas}} \sim 15 \cdot 10^{-4}$
79%He-21% $\text{CH}_4$	@ HV=2110V	$G_{\text{gas}} \sim 7 \cdot 10^{-4}$
79%He-21% $\text{CH}_4$	@ HV=2210V	$G_{\text{gas}} \sim 14 \cdot 10^{-4}$

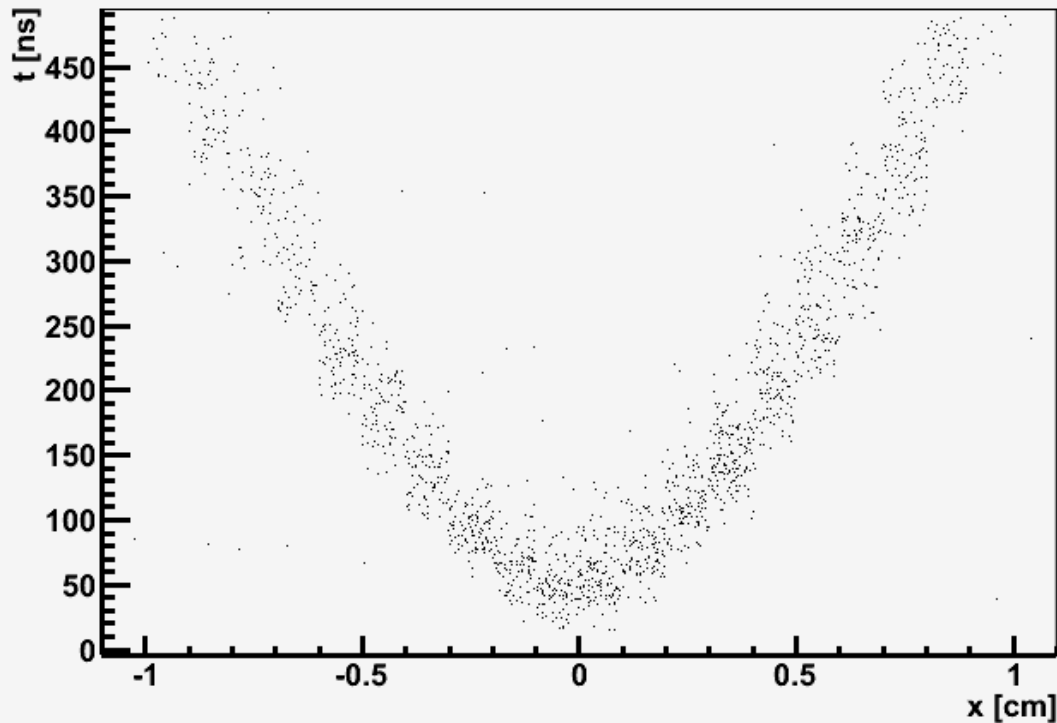


# Space-Time Relations & Spatial Resolution

Disclaimer:

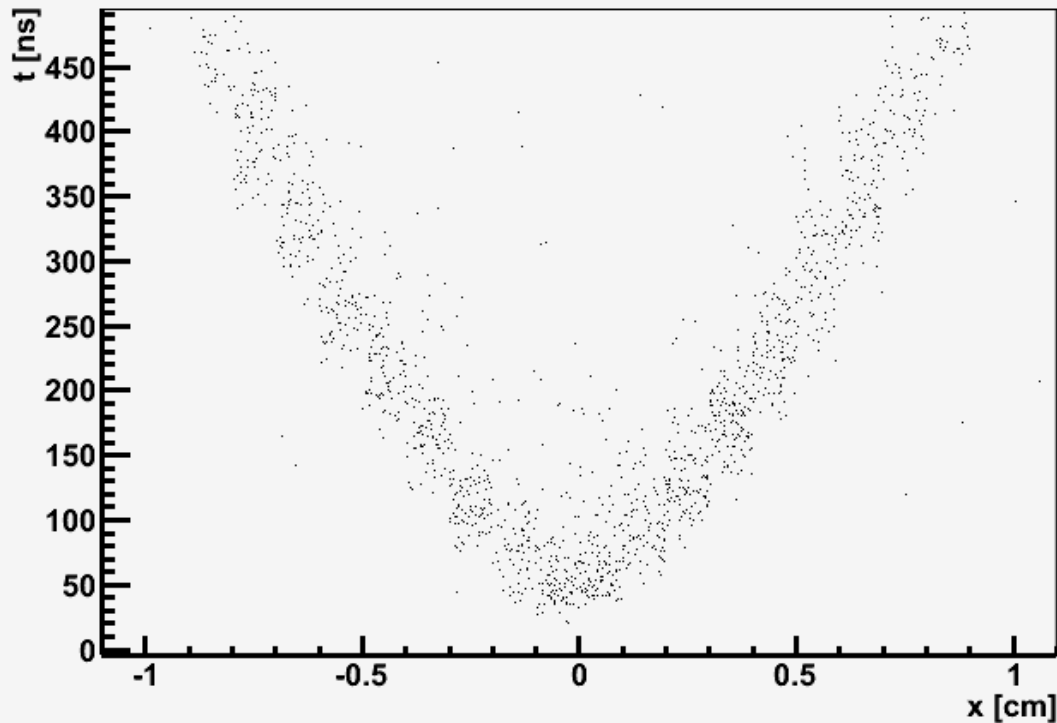
- Results shown on pages 10-19 to be considered VERY preliminary
- Still limited statistics from cosmic-ray runs
- Still approximate time-to-space relations
- Thresholds could be optimized...

# Space-Time Relations



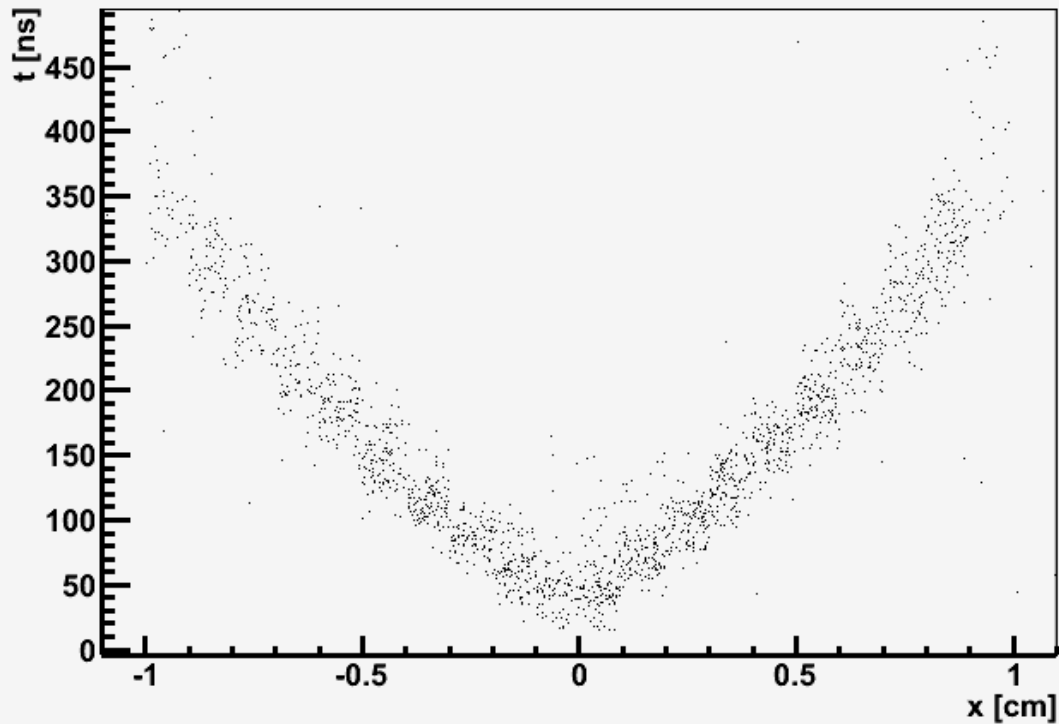
80%He-20% $i$ C<sub>4</sub>H<sub>10</sub> @ HV=1960V

# Space-Time Relations



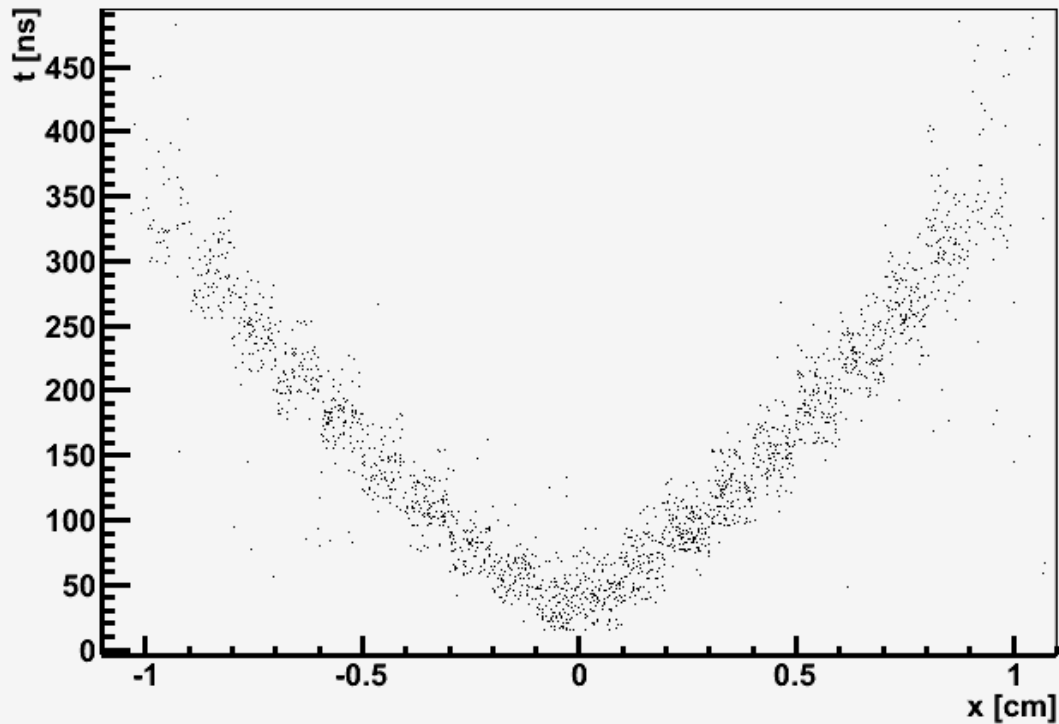
90%He-10% $i$ C<sub>4</sub>H<sub>10</sub> @ HV=1710V

# Space-Time Relations



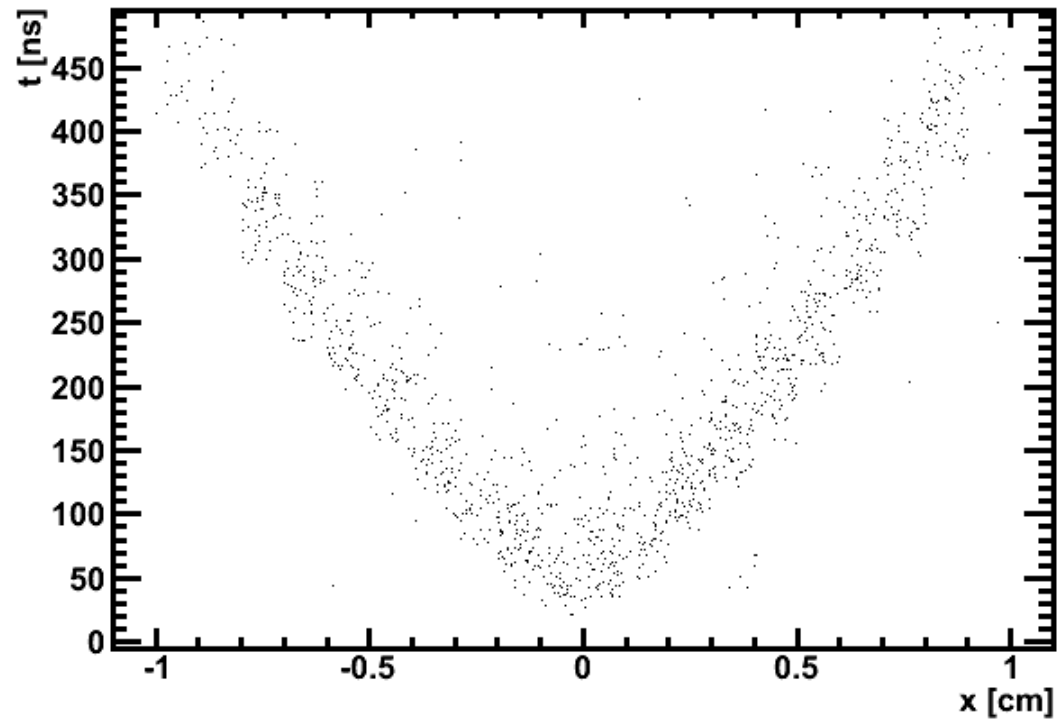
63%He-37%CH<sub>4</sub> @ HV=2410V

# Space-Time Relations



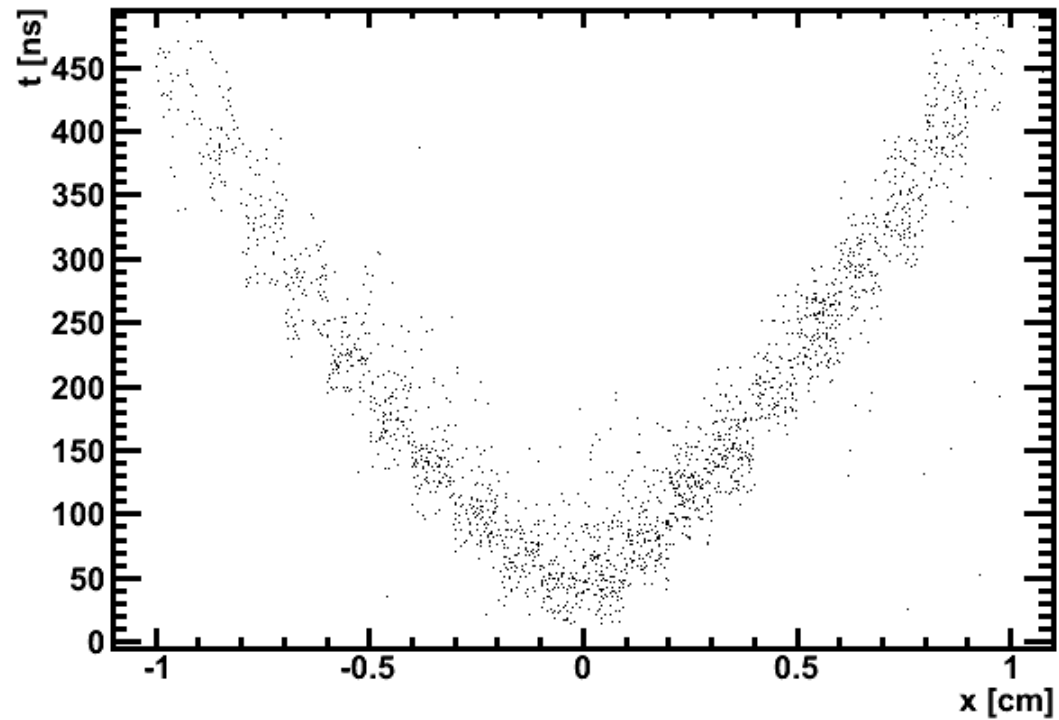
63%He-37%CH<sub>4</sub> @ HV=2510V

# Space-Time Relations



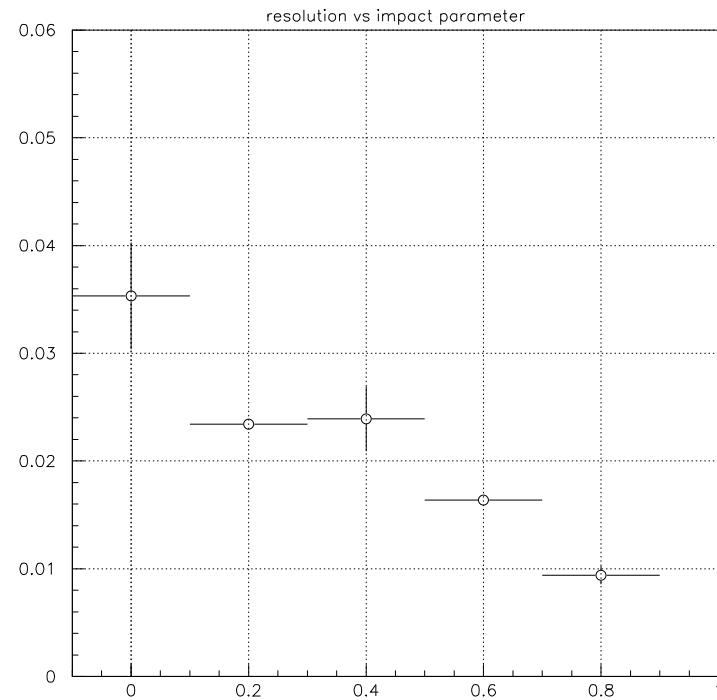
79%He-21%CH<sub>4</sub> @ HV=2110V

# Space-Time Relations



79%He-21%CH<sub>4</sub> @ HV=2210V

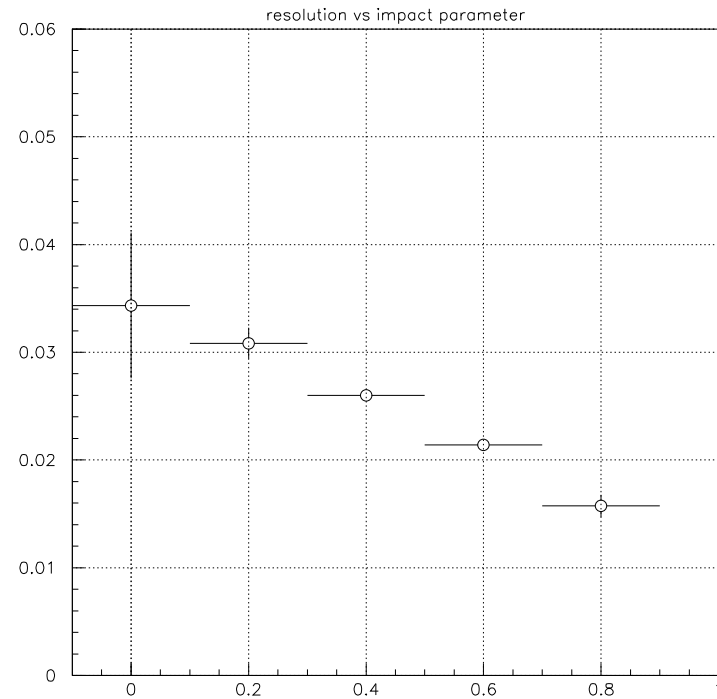
# Spatial Resolution



**80%He-20% $i$ C<sub>4</sub>H<sub>10</sub> @ HV=1960V**

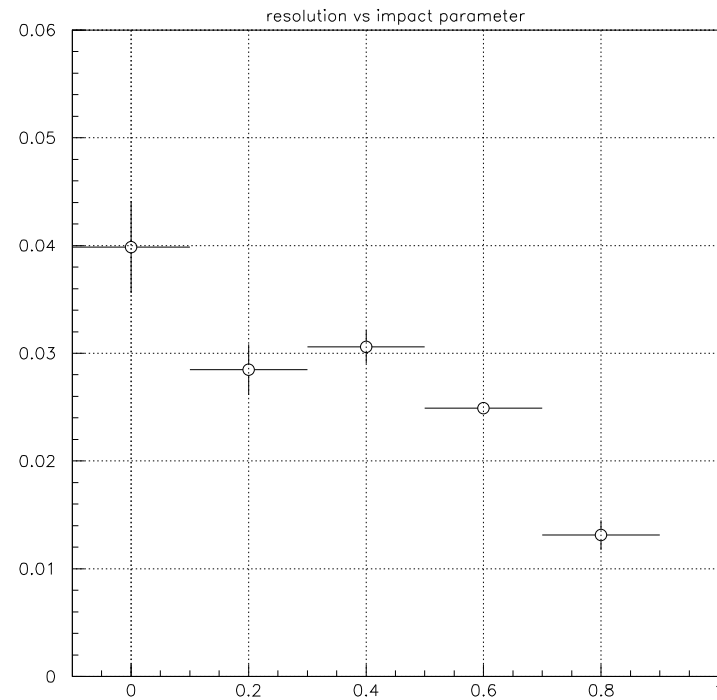


# Spatial Resolution



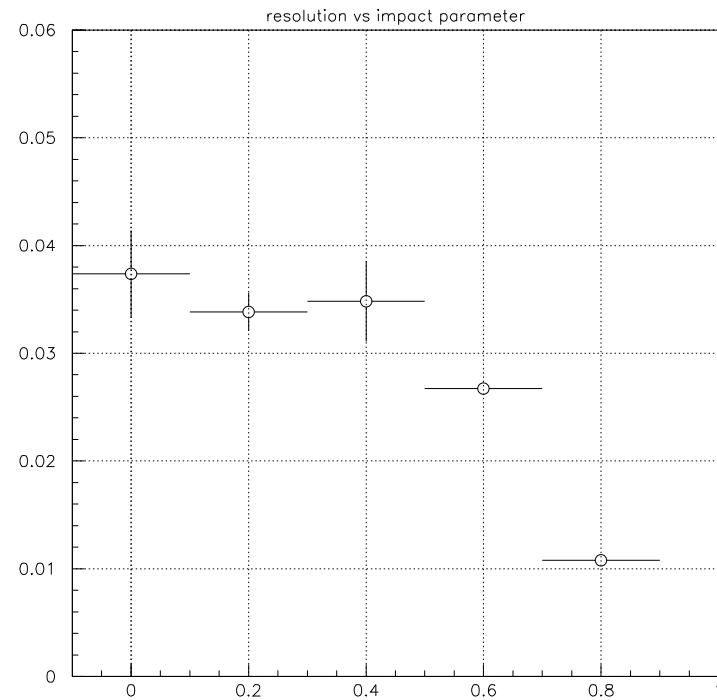
**63%He-37%CH<sub>4</sub> @ HV=2410V**

# Spatial Resolution



**63%He-37%CH<sub>4</sub> @ HV=2510V**

# Spatial Resolution



**79%He-21%CH<sub>4</sub> @ HV=2210V**

# Summary

- Collecting cosmic ray data to study several gas mixtures
- First results appearing
  - Consistent with *BABAR* experience
  - Mixtures with very low Methane content could be well-suited for Cluster Counting (next planned step)