Detector Geometry WG

M. Rama and A. Stocchi for the DGWG SuperB general meeting, Annecy 18 March 2010

Studies before this meeting

Summary of the main studies performed so far

System	Recent studies
SVT	 Time-dependent measurements as a function of the layer0 Tracking performance as a function of the SVT outer radius Time-dep meas. with B→KsKs as a function of the SVT outer radius Tracking performance as a function of the number of layers Degradation of sin2beta error when the boost goes from 0.28 to 0.238
DCH	 Tracking performance as a function of the DCH inner radius Tracking and dE/dx as a function of the DCH length
forward PID	 B→K(*) vv SL tag with/without TOF B→K(*) vv HAD tag with/without TOF (in progress)
EMC	 B→τν with/without backward EMC E resolution of fwd EMC as a function of material in front of it (prel)
IFR	 Optimization of the muon selection

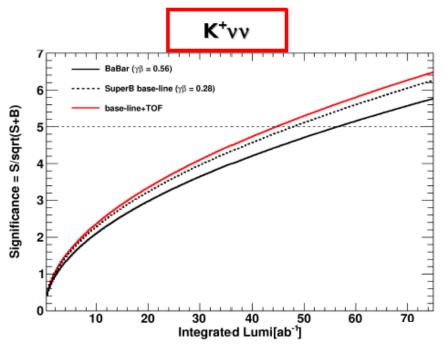
See also the DGWG wiki page:

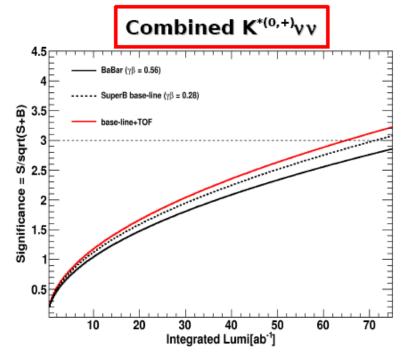


Shown in Frascati: no BhaBha+radBhBha included

fwd PID (TOF) with $B \rightarrow K^{(*)} vv$: reach

A. Perez **CONCLUSIONS:**





5σ significance (stat-only): ~55ab⁻¹ BaBar: SuperB-base line: ~48ab⁻¹ +TOF: ~44ab⁻¹

Gain on significance: boost ~ 7-8%

fwd PID ~ 5%

3σ significance (stat-only): BaBar: > 75ab⁻¹ SuperB-base line: ~71ab⁻¹ +TOF: ~64ab⁻¹

Shown in Frascati: no BhaBha+radBhBha included

backward EMC for B→tv study

A. Rakitin

CONCLUSIONS:

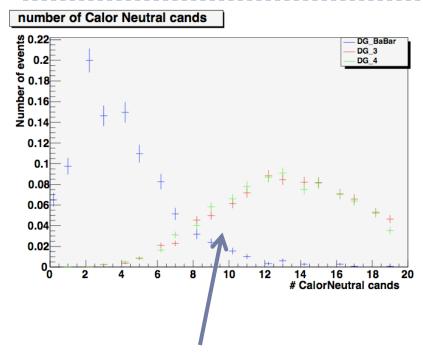
$\tau \to \ell \nu \nu, \pi \nu$			$\tau \to \rho(\pi\pi^0)\nu$	$\tau \to \pi 2\pi^0 \nu, a_1(3\pi)\nu, 3\pi\pi^0 \nu$		
Mimicking decay	Rel. BF		Mimicking decay	Rel. BF	Mimicking decay	Rel. BF
$B \to D^{(*(*))0} \ell \nu$	$\mathcal{O}(10^2 - 10^3)$		$B \to D^{(*(*))0}\rho$	$\mathcal{O}(10^2 - 10^3)$	Anything with final state	???
$B \to D^{(*)} \pi \ell \nu$	$O(10^2)$	direct	$B \rightarrow J/\psi \rho$	1.39	$\pi 2\pi^{0}, 3\pi \text{ or } 3\pi\pi^{0}$	
$B o$ non-charm ℓu	$\mathcal{O}(1)$	Ė	B o non-charm $ ho$	$\mathcal{O}(1)$		
		indirect ρ	$B o$ stuff, stuff $ o ho + X_{missed}$???		
		_ =	Anything with final state $\pi^0 + \operatorname{track} + X_{missed}$???		
		π ⁰ +	Special case: $B o\pi^0\ell u$ (nothing is missed)	0.55		
Done, $\sim 8\%$ improvement Partially done, $\sim 10\%$ imp. for direct $\sim 10\%$					Definitely need	
in $S/\sqrt{S+B}$ due to	o bwd EMC	need generic MC for the rest			generic MC	

Next steps:

- a) Include other signal modes and bkg decays
- b) Include the main B tag modes (HAD+SL)

results with $B \rightarrow D^0 \pi$ tag

Effect of radBhabha+Bhabha on B>K+vv

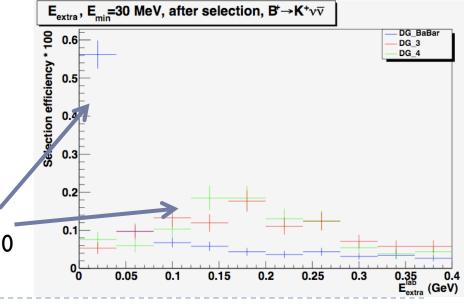


increased number of neutrals

signal Eextra no longer peaks at 0 when (rad)Bhabha is added

E. Manoni
this meeting
(see also A. Perez
@DGWG session)

Note: bkg overestimated by a factor ~4



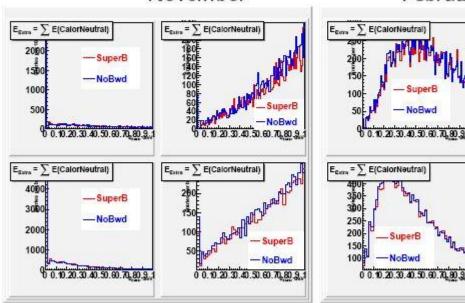
Effect of radBhabha+Bhabha on B >τν

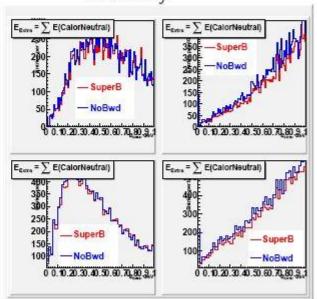


E_{extra}

A. Rakitin this meeting

no bkg November with bkg February





• Top: All B_{sig} , bottom: best B_{sig}

• Left: sig MC, right: generic MC

Note: bkg overestimated by a factor ~4

Need to evaluate how the machine bkg changes the conclusions we've reached so far (see sl. 2)

- Note: pairs background were not included in the FastSim Feb. production. We must include it in next productions.
- Studies concerning the position of the L0 are using the bkg rates as input (especially pairs).
- Results concerning tracking (SVT/DCH transition radius, #SVT layers, effect of boost on T.D. analyses) should still be valid at first order
 - At some point we wanted to evaluate the effect of bkg on reconstruction. It was already in the to-do list but not done yet.
- Can we trust the previous estimate of the fwd PID impact on $B \rightarrow K^{(*)}vv$? At first order our guess is yes
- ▶ The physics reach of $B \rightarrow \tau v$ vs bwd EMC must be re-evaluated

Development work needed. Some preliminary thoughts

- Consolidate the current analyses
 - remove n_clust<15 from B->Knunu SL tag, characterize the bkg properties
- Optimize the analysis according to the SuperB environment, different from Babar
- Understand what margins there are at the detector level
 - for example, how much the EMC time windows can be reduced, and what would be the impact on cluster reconstruction?
- Organize ad hoc (mini?)productions before the Summer prod
 - Can we speed up the current MC event rate, for example with ad hoc pre-selection cuts or by simulating subsamples of the events, still preserving the prediction power of the analyses? The statistics of the Feb production was not large enough to study the bkg properties of rare decays.

