

# An overview of Leptogenesis

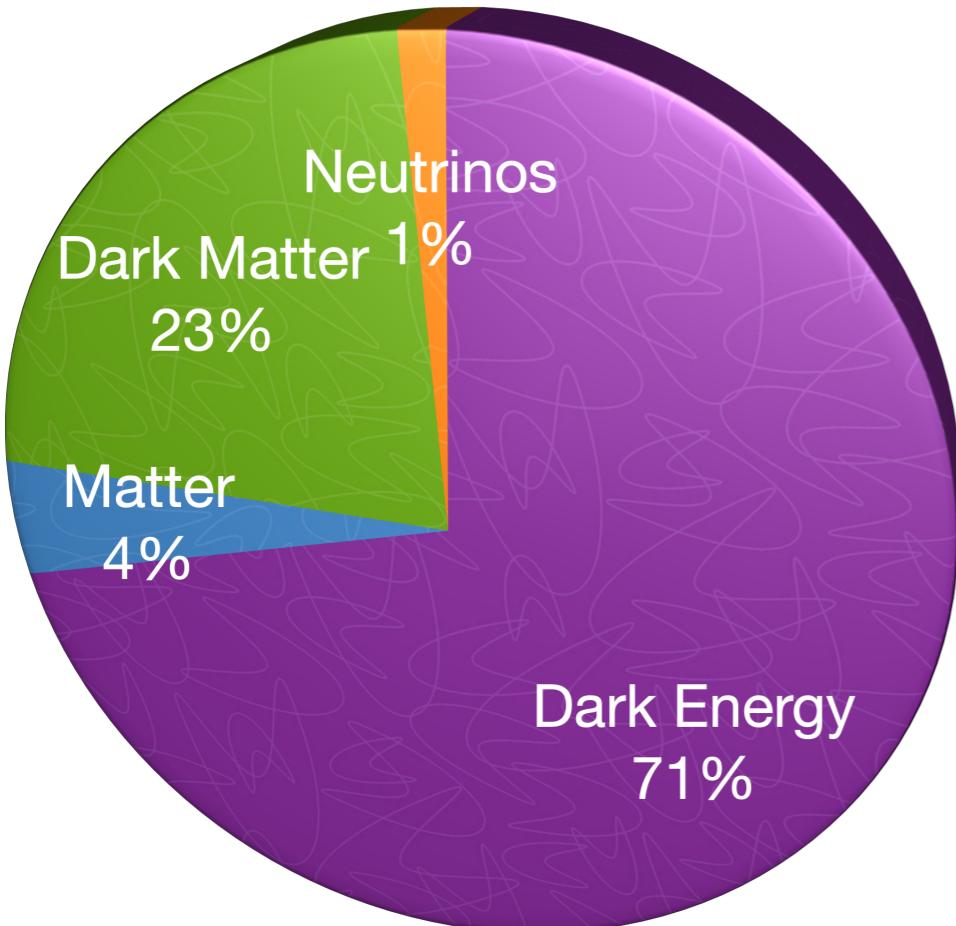
**FPCapri 2022**

**Jessica Turner, IPPP, Durham University**

**12 June 2022**



# Universe's Energy Budget



$$\eta_B = (6.02 - 6.18) \times 10^{-10}$$

Planck 1807.06209 (2018)

## Sakharov's Conditions



Baryon number violation

Kuzmin, Rubakov & Shaposhnikov  
*Phys.Lett.B* 155 (1985)



C & CP-violation

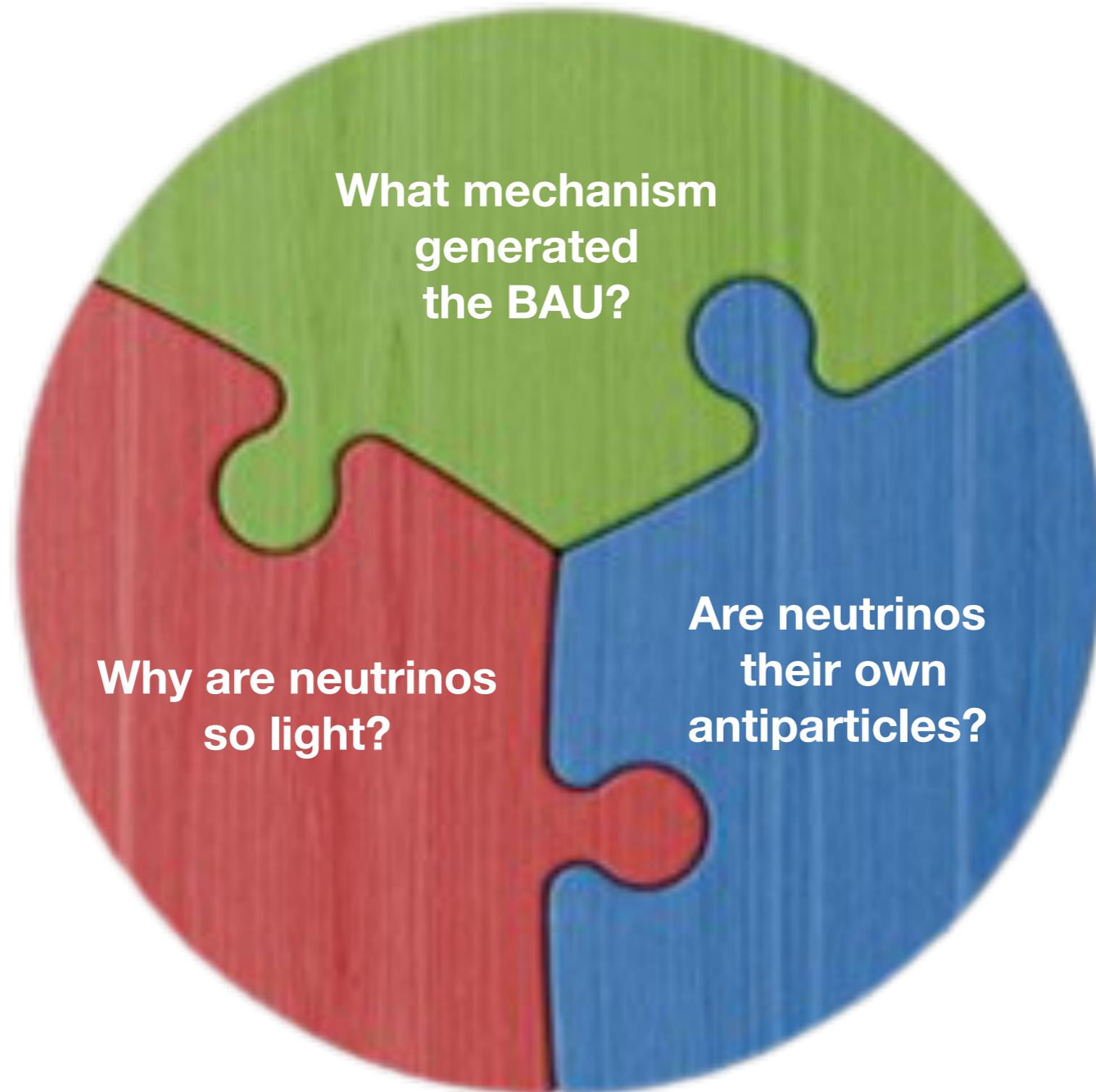
Gavela, Hernandez, Orloff & Pene *Mod.Phys.Lett.*  
A9 795-810 (1994) Huet & Sather *Phys.Rev. D* 51  
379-394 (1994)



Departure from thermal equilibrium

Kajantie, Laine, Rummukainen  
& Shaposhnikov *Phys.Rev.Lett.*  
77 2887-2890 (1996)

# Motivation for Leptogenesis

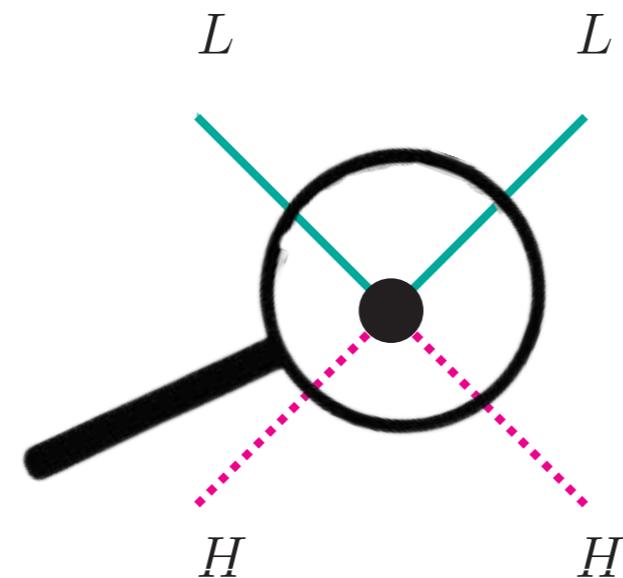


# The Seesaw Mechanism

The Standard Model is an effective theory which contains non-renormalisable operators

Weinberg, *Phys.Rev.Lett.* 43 (1979)

$$\mathcal{L} \supset -Y_{ij} \frac{L^i H L^j H}{2M} + \mathcal{O}\left(\frac{1}{M^2}\right) + \text{h.c}$$

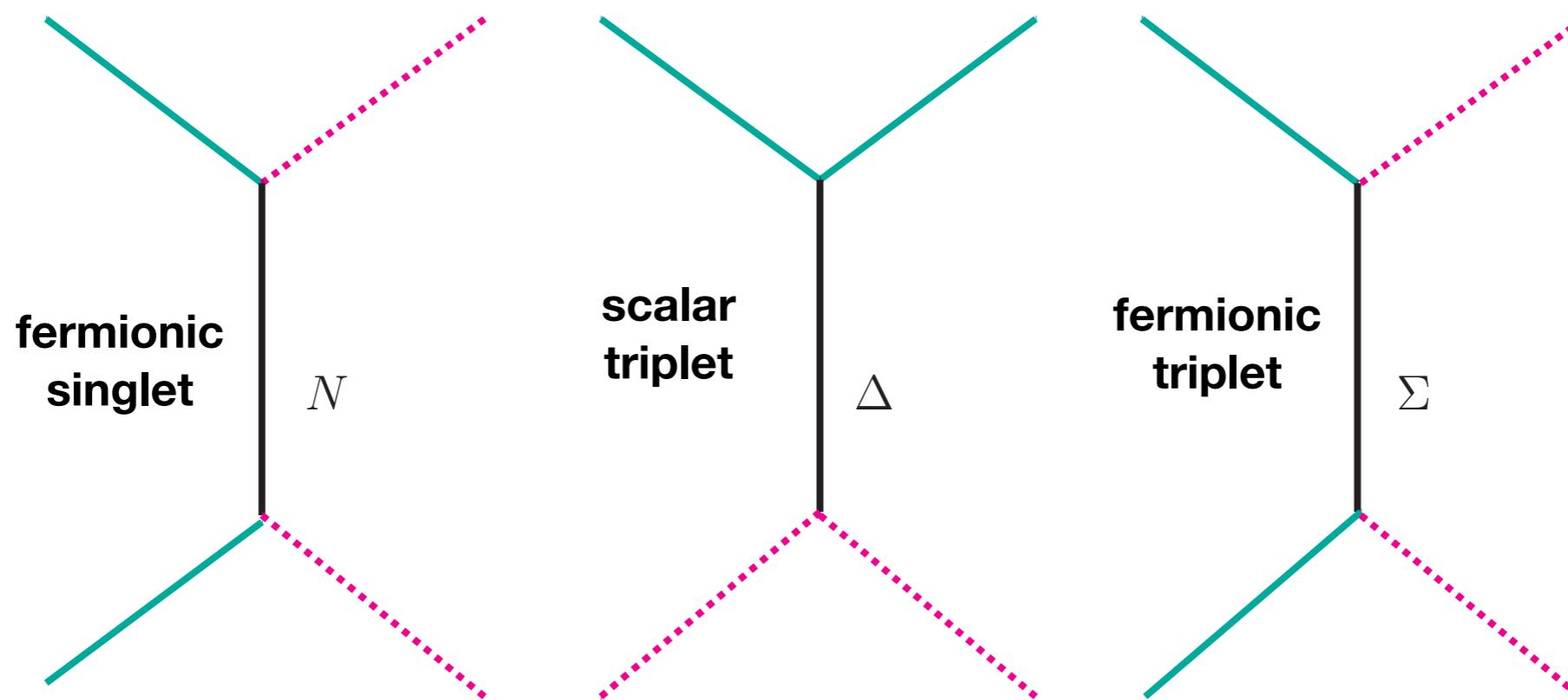


# The Seesaw Mechanism

After SSB a Majorana mass is produced for the active neutrinos

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Minkowski, Yanagida, Glashow, Gell-Mann, Ramond, Slansky, Mohapatra, Senjanovic

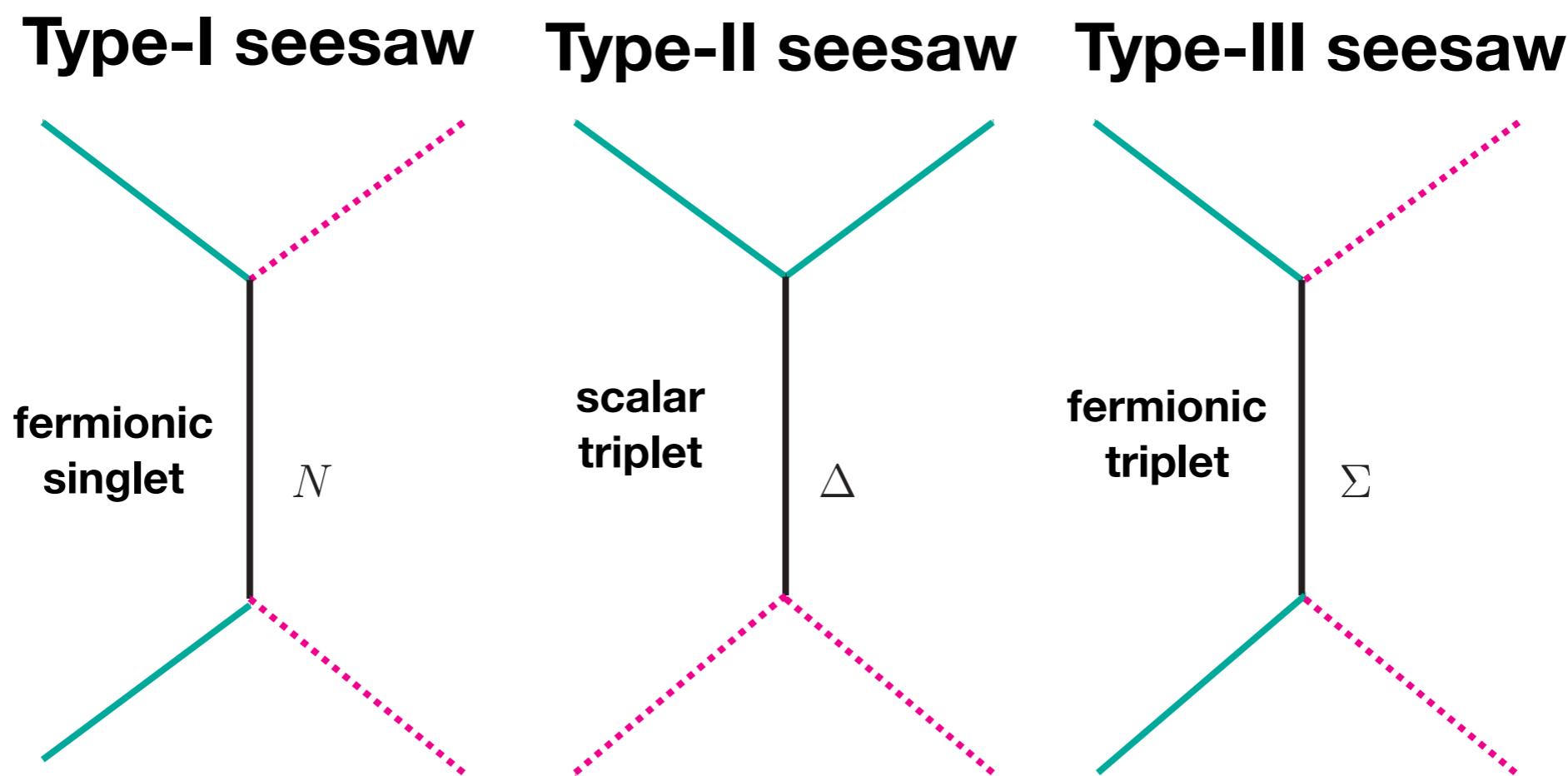
Magg, Wetterich, Lazarides, Shafi, Mohapatra, Senjanovic, Schecter, Valle

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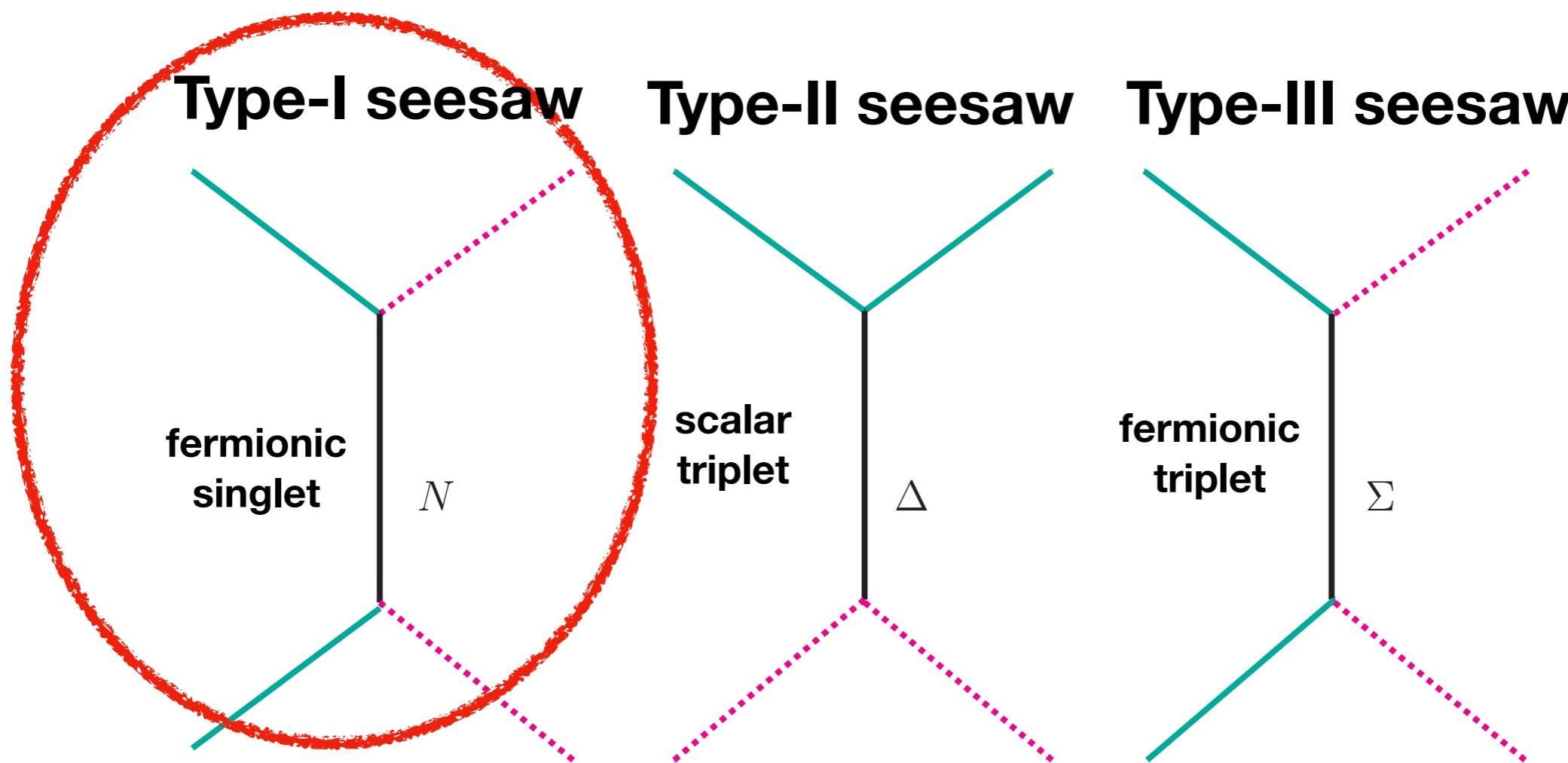
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# The Seesaw Mechanism

$$\mathcal{L} \supset -\overline{L}_\alpha Y_{\alpha i} N_i \tilde{H} - \frac{1}{2} \overline{N}_i^C M_{N_i} N_i + \text{h.c.}$$

After diagonalising the mass matrix

$$m_\nu \approx \frac{m_D m_D^T}{M_N} = \frac{Y^2 v^2}{M_N}$$

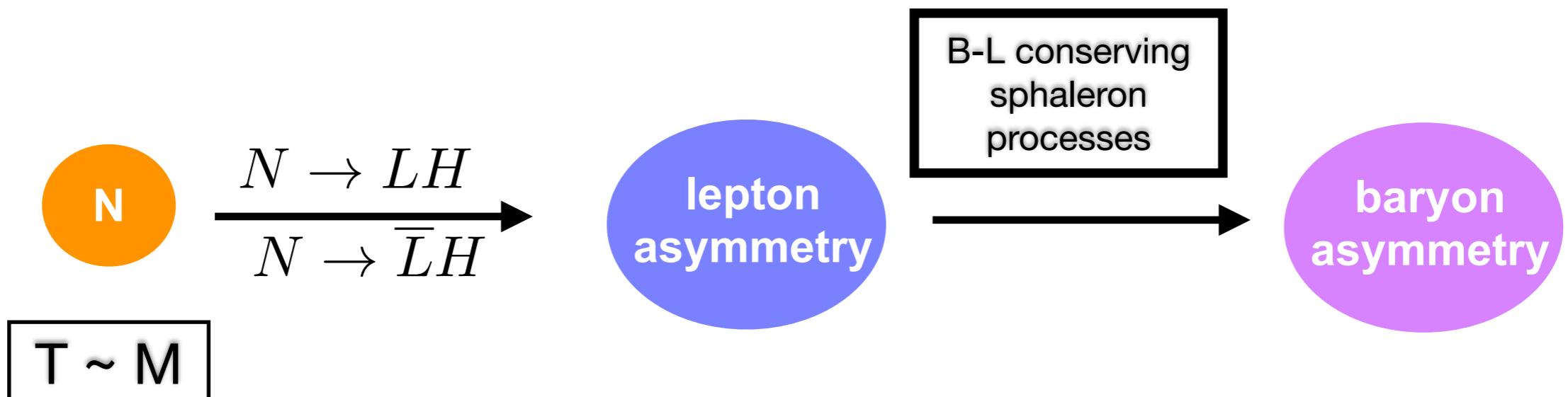
## Sakharov's Conditions



Image courtesy of Symmetry Magazine

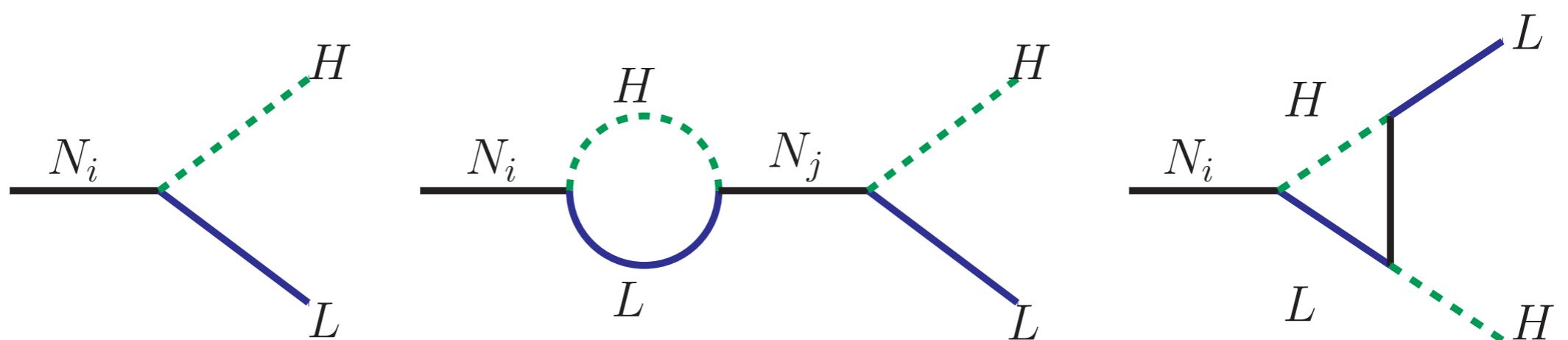
- Baryon number violation
- C & CP-violation
- Departure from thermal equilibrium

# Thermal Leptogenesis



**Decay asymmetry from interference between tree and loop level diagrams**

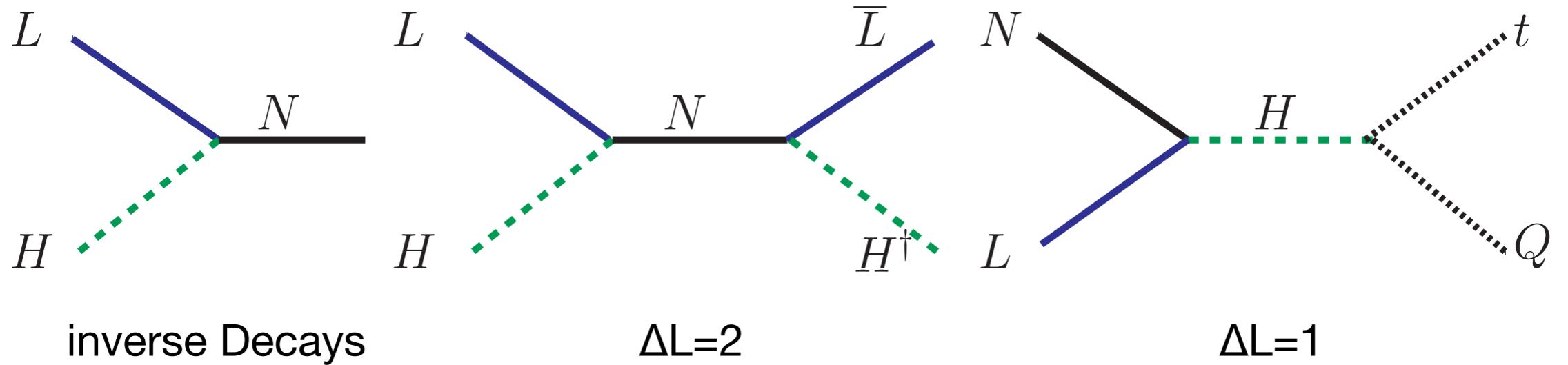
Covi, Roulet, Vissani



$$\epsilon_i = \frac{\Gamma_i - \overline{\Gamma}_i}{\Gamma_i + \overline{\Gamma}_i}$$

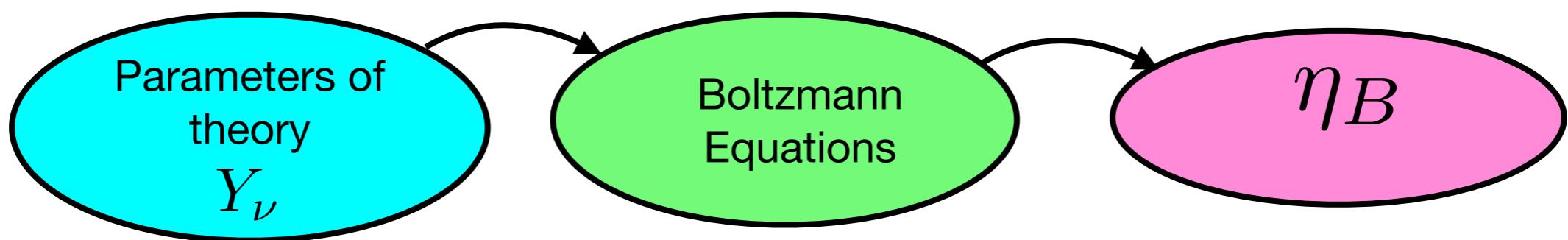
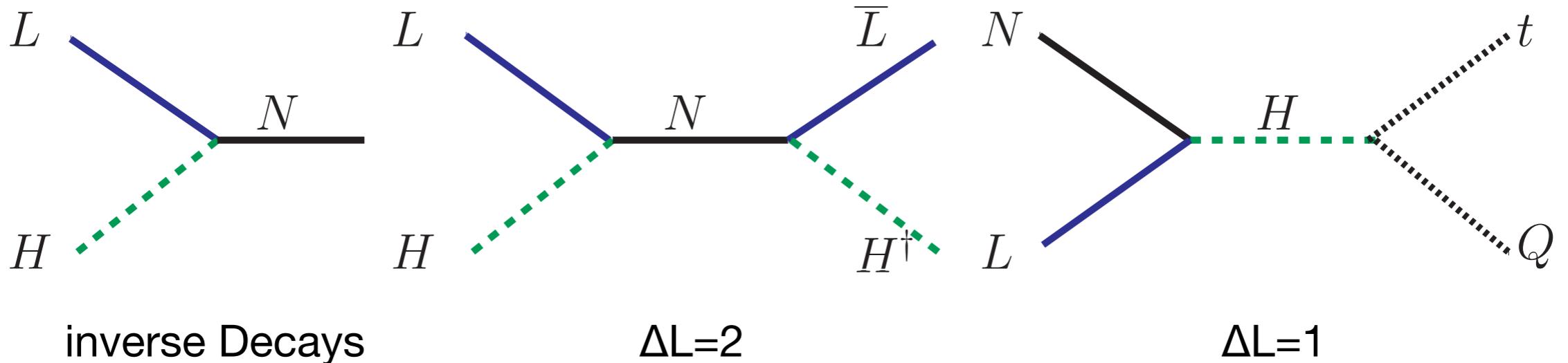
# Thermal Leptogenesis

Washout and scattering processes



# Thermal Leptogenesis

Washout and scattering processes



$$\frac{dn_{N_i}}{dz} = - D_i (n_{N_i} - n_{N_i}^{\text{eq}}),$$

**source**                    **sink**

$$\frac{dn_{B-L}}{dz} = \sum_{i=1}^3 \left( \epsilon^{(i)} D_i (n_{N_i} - n_{N_i}^{\text{eq}}) - W_i n_{B-L} \right).$$

RHN Mass

$\mathcal{O}(10^{12})$  GeV

Fukugida & Yanagida (1986)

$\mathcal{O}(10^6)$  GeV

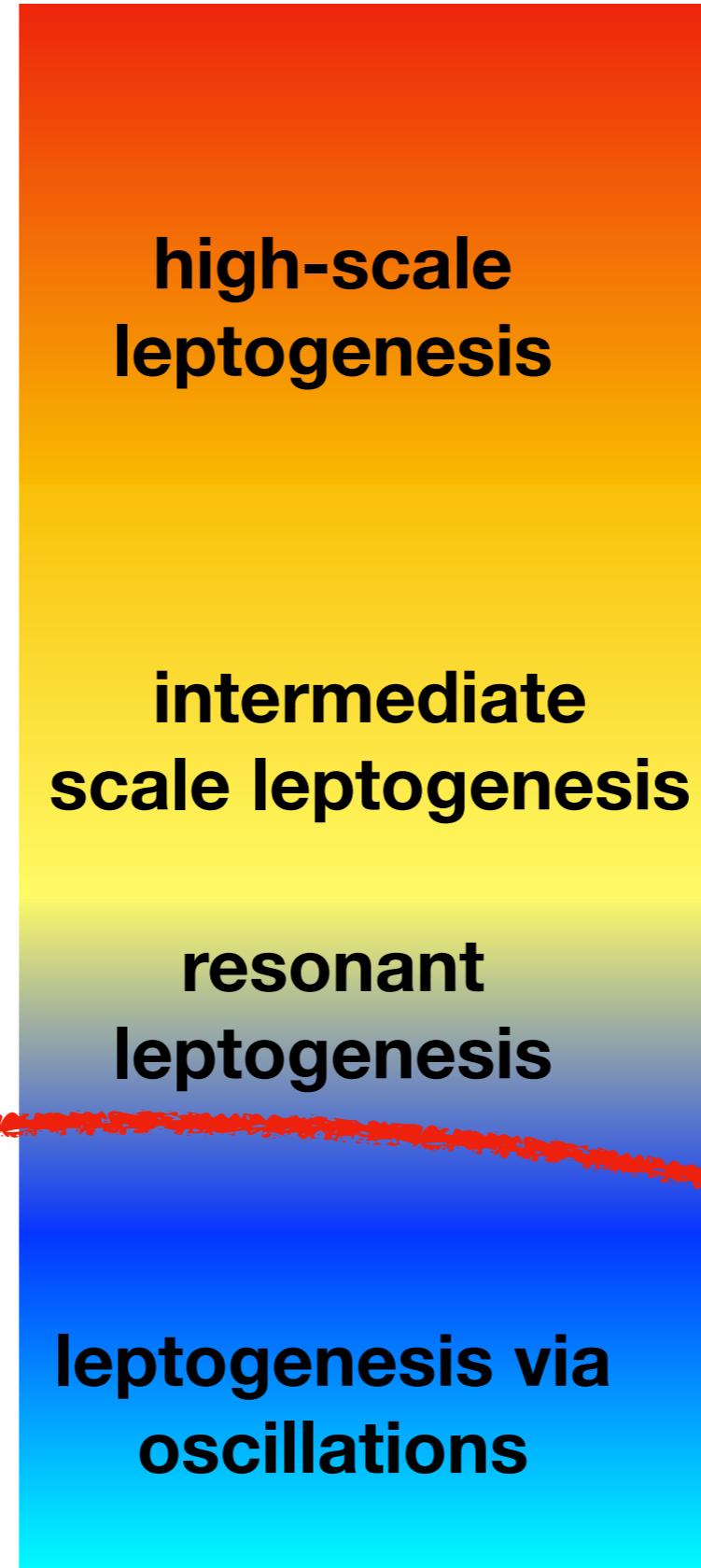
Moffat, Petcov, Pascoli, Schulz  
& JT (2018)

$\mathcal{O}(10^3)$  GeV

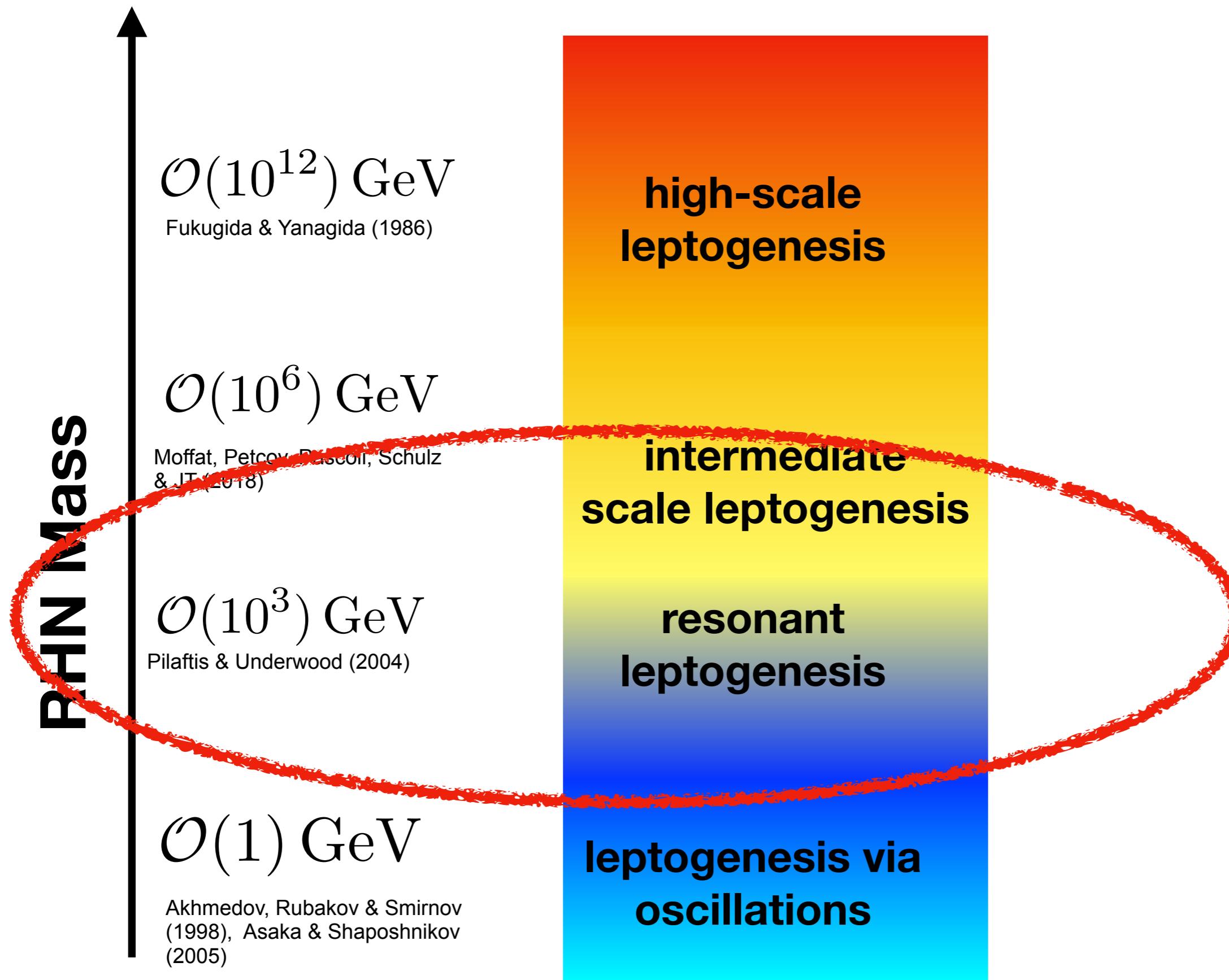
Pilaftsis & Underwood (2004)

$\mathcal{O}(1)$  GeV

Akhmedov, Rubakov & Smirnov  
(1998), Asaka & Shaposhnikov  
(2005)



Dedicated talk by Juraj Klaric

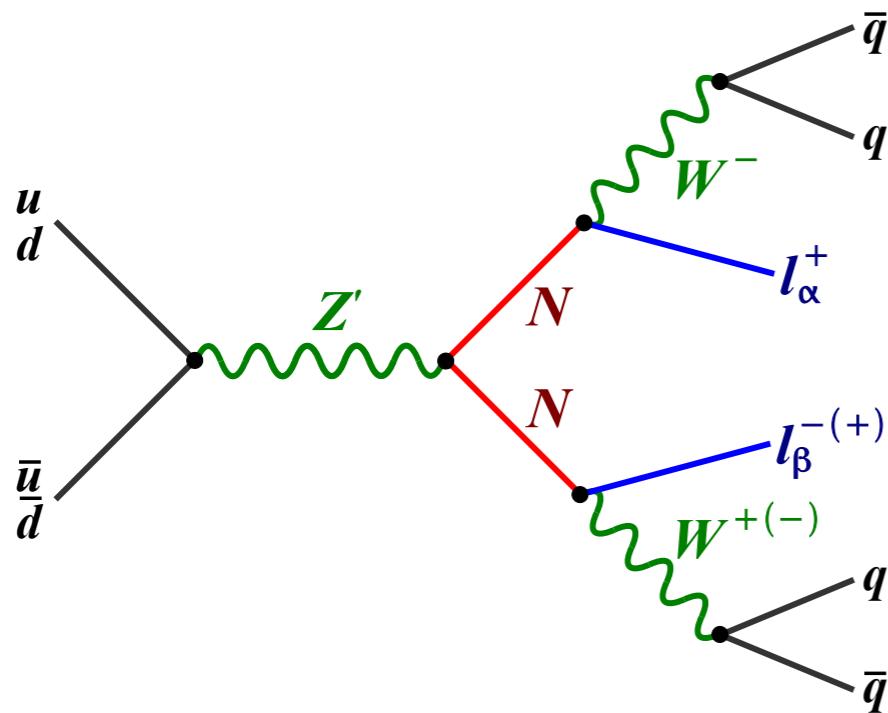


# Resonant Leptogenesis

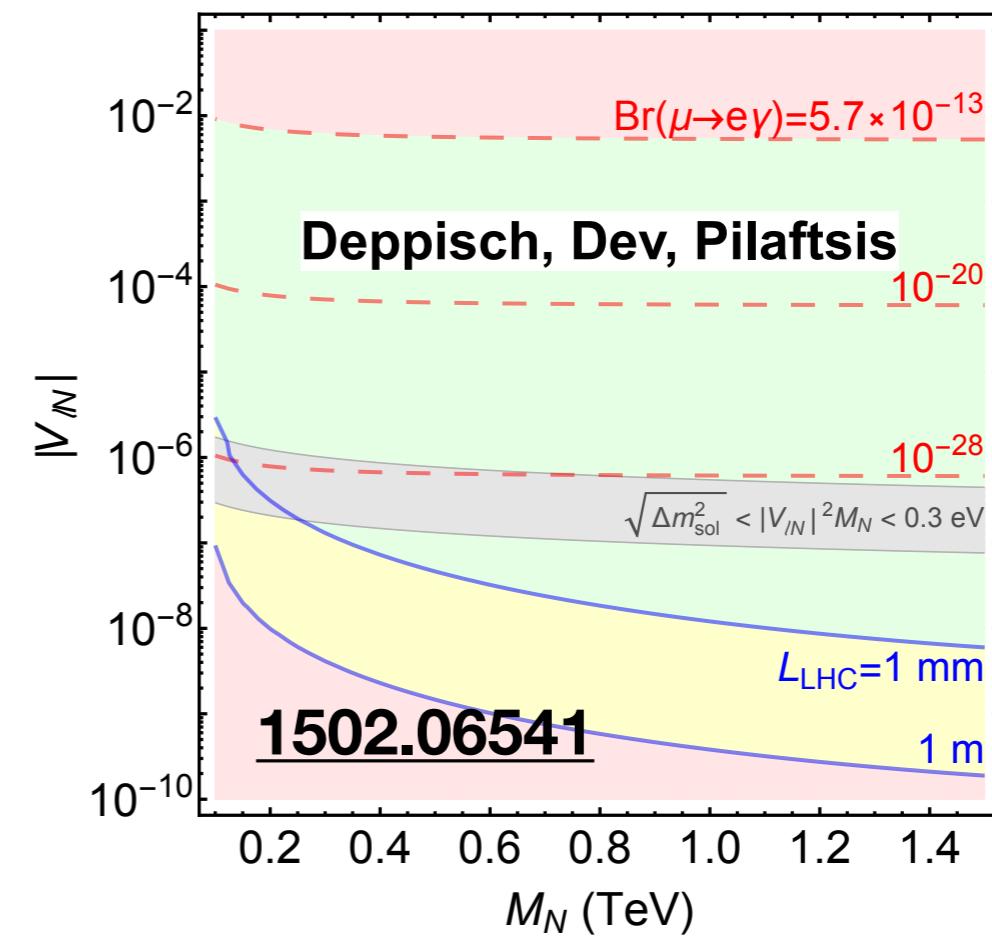
- Regime where RHNs decay width similar to their mass differences. Mass range  $\sim \text{TeV}$

Pilaftsis & Underwood (2004)  
Abada, Aissaoui, Losada (2005)

- RHN masses explained by additional  $U(1)_{B-L}$  symmetry and can be sufficiently long-lived  $\rightarrow$  displaced-vertex signature searched for at LHC, MATHUSLA or SHiP.



LNV, LFV at colliders  
talk by Stefan Antusch

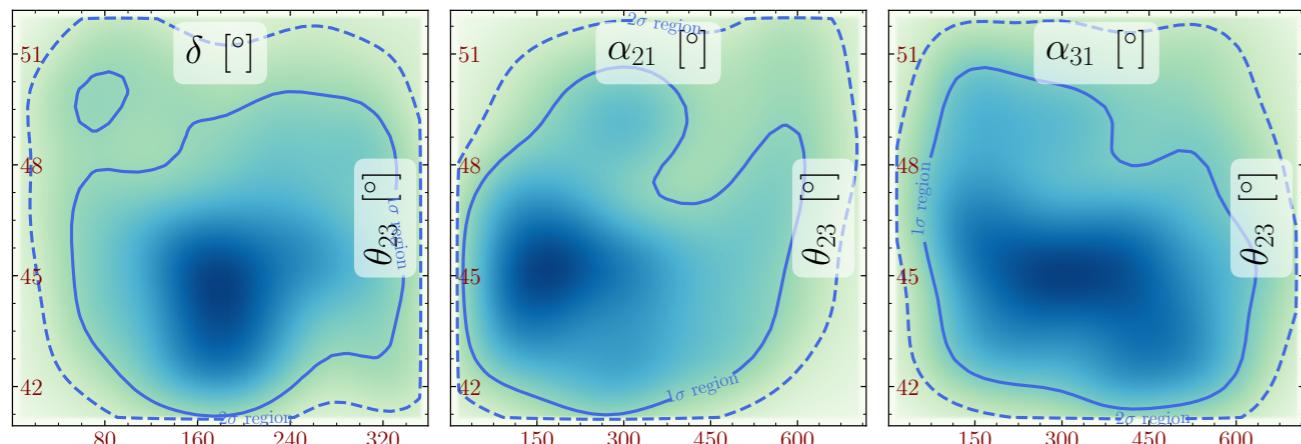


# Resonant leptogenesis in the Neutrino Option

- Assume scale invariance above  $M_N$
- Integrate out TeV scale RHN and RG evolve: Higgs potential produced for  $M_N \sim 1000$  TeV

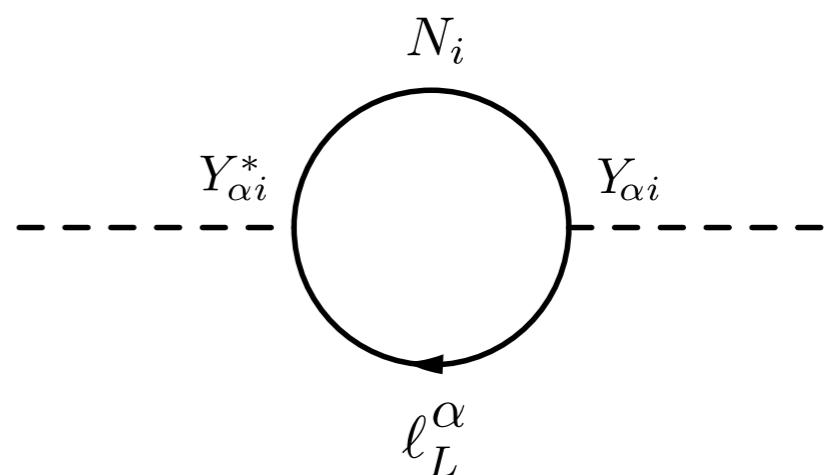
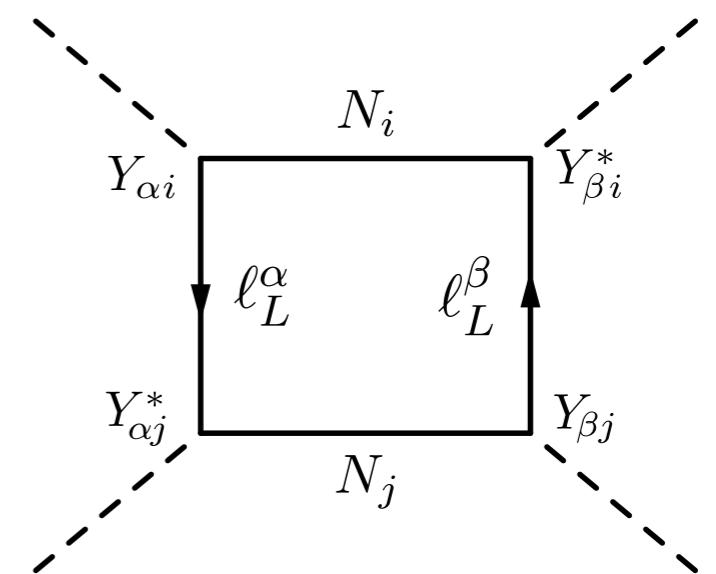
Brdar, Hemboldt, Iwamoto, Schmitz  
Brivio, Moffat, Pascoli, Petcov, Turner

## Normal Ordering



$$\frac{\Delta M}{M} \sim 10^{-8} \quad \overline{M} = 1.2 \times 10^6 \text{ GeV}$$

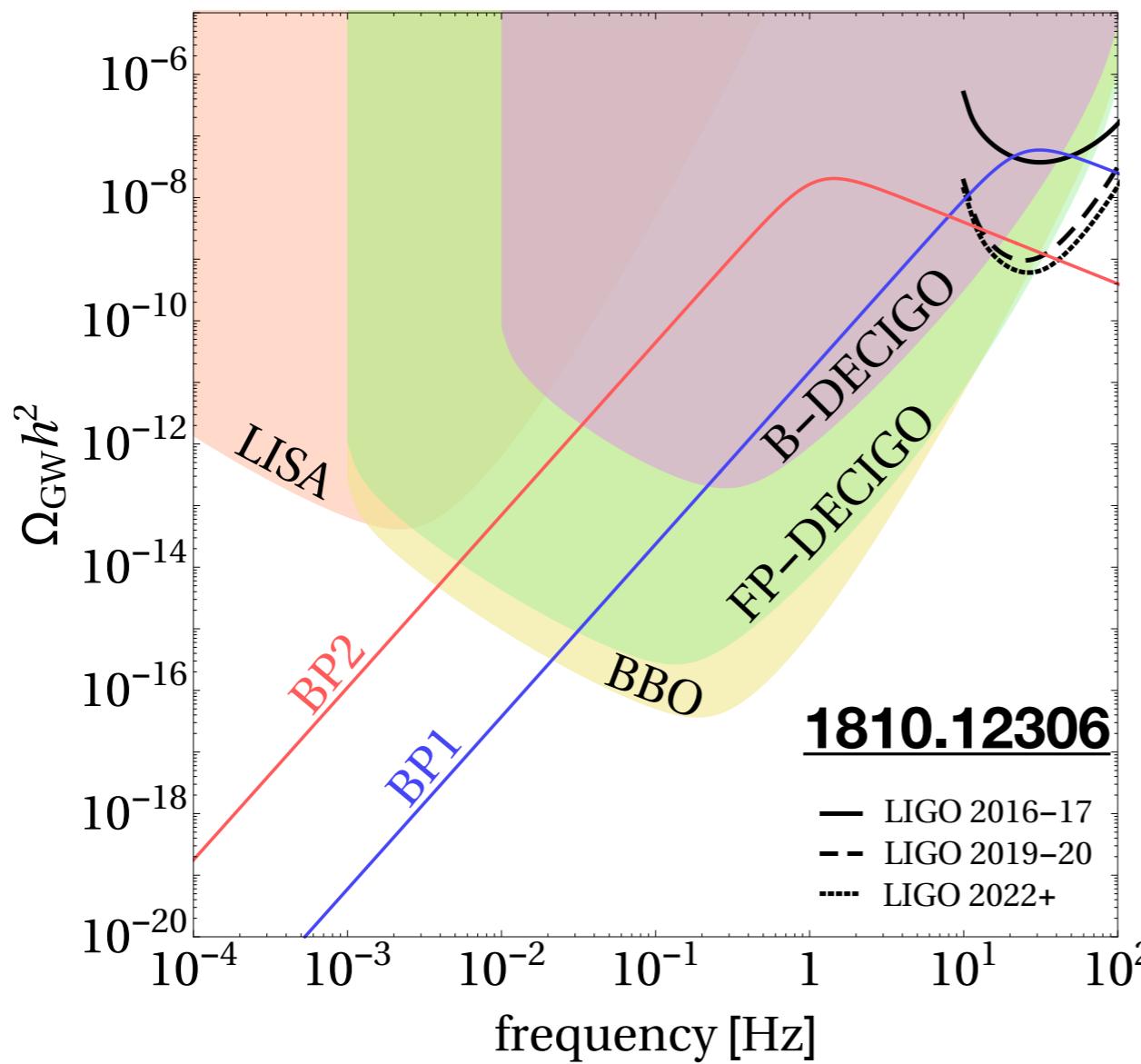
slight preference for  
 $\theta_{23} < 45^\circ$  and  $\delta \sim 180^\circ$



Scale invariance broken at  
quantum level

# Resonant leptogenesis in the Neutrino Option

- UV-completion of Neutrino Option (**Brdar**, Emonds, Helmboldt, Lindner) minimal renormalisable model based on classical scale invariance
- New scalar breaks scale-invariance → generates mass for RHNs and strong first order phase transition



**Brdar, Helmboldt, Kubo (2018)**

See also “Probing the seesaw scale with gravitational waves” Okada & Seto

RHN Mass

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(1998), Asaka & Shaposhnikov  
(2005)

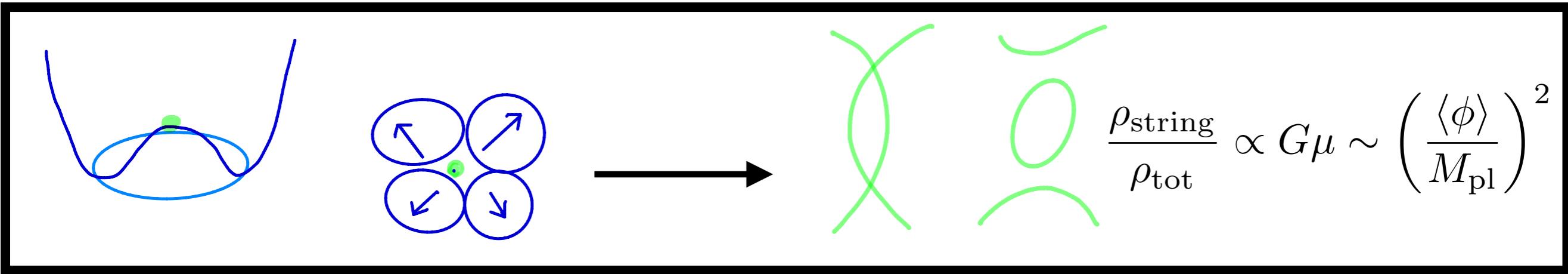
**high-scale  
leptogenesis**

**intermediate  
scale leptogenesis**

**resonant  
leptogenesis**

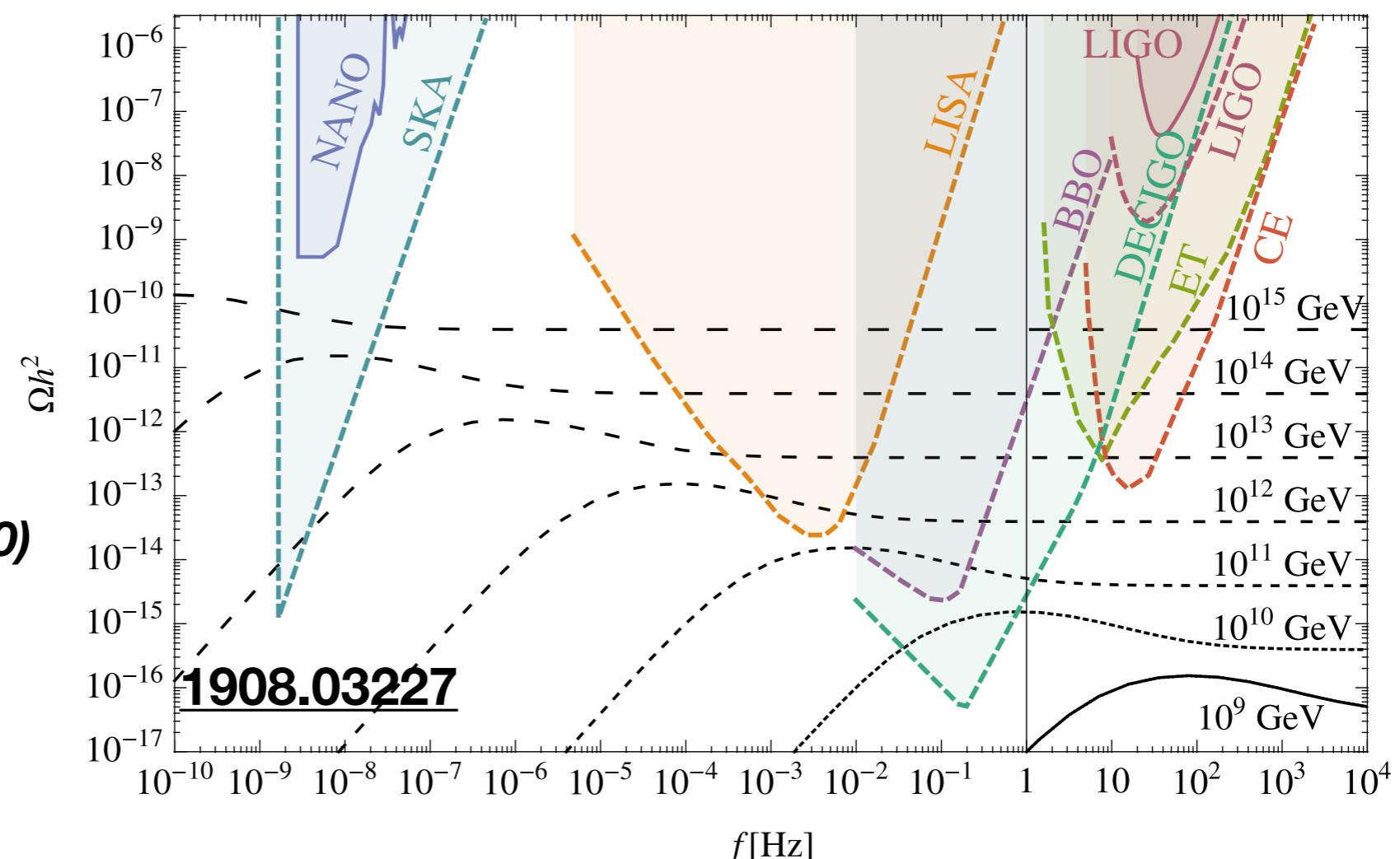
**leptogenesis via  
oscillations**

Difficult to test,  
however gravitational  
waves offer an additional  
telescope on high-scale  
leptogenesis



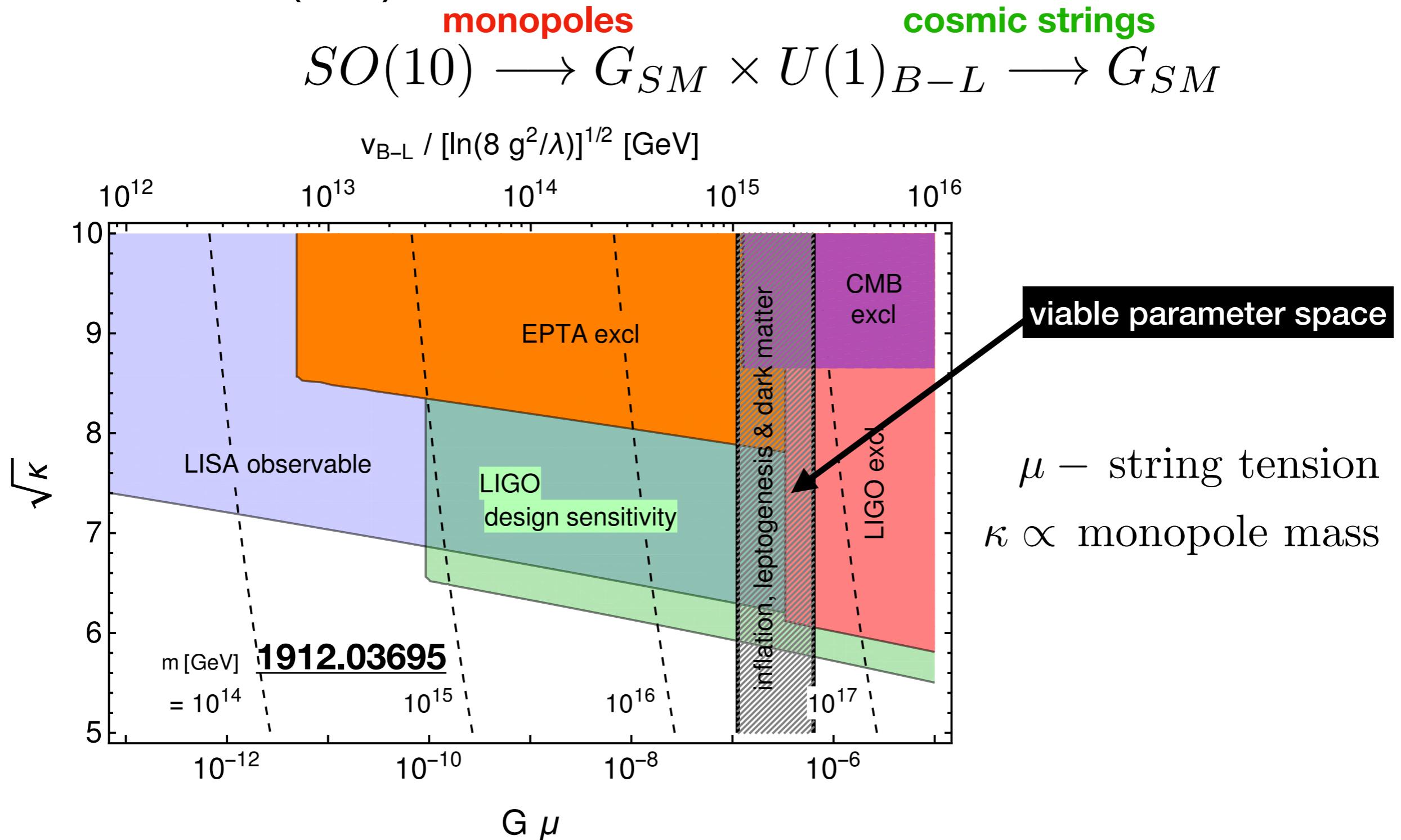
- Highlighted that GWs from cosmic string network generic prediction of seesaw mechanism

Dror, Hiramatsu, Kohri ,  
Murayama & White (2020)



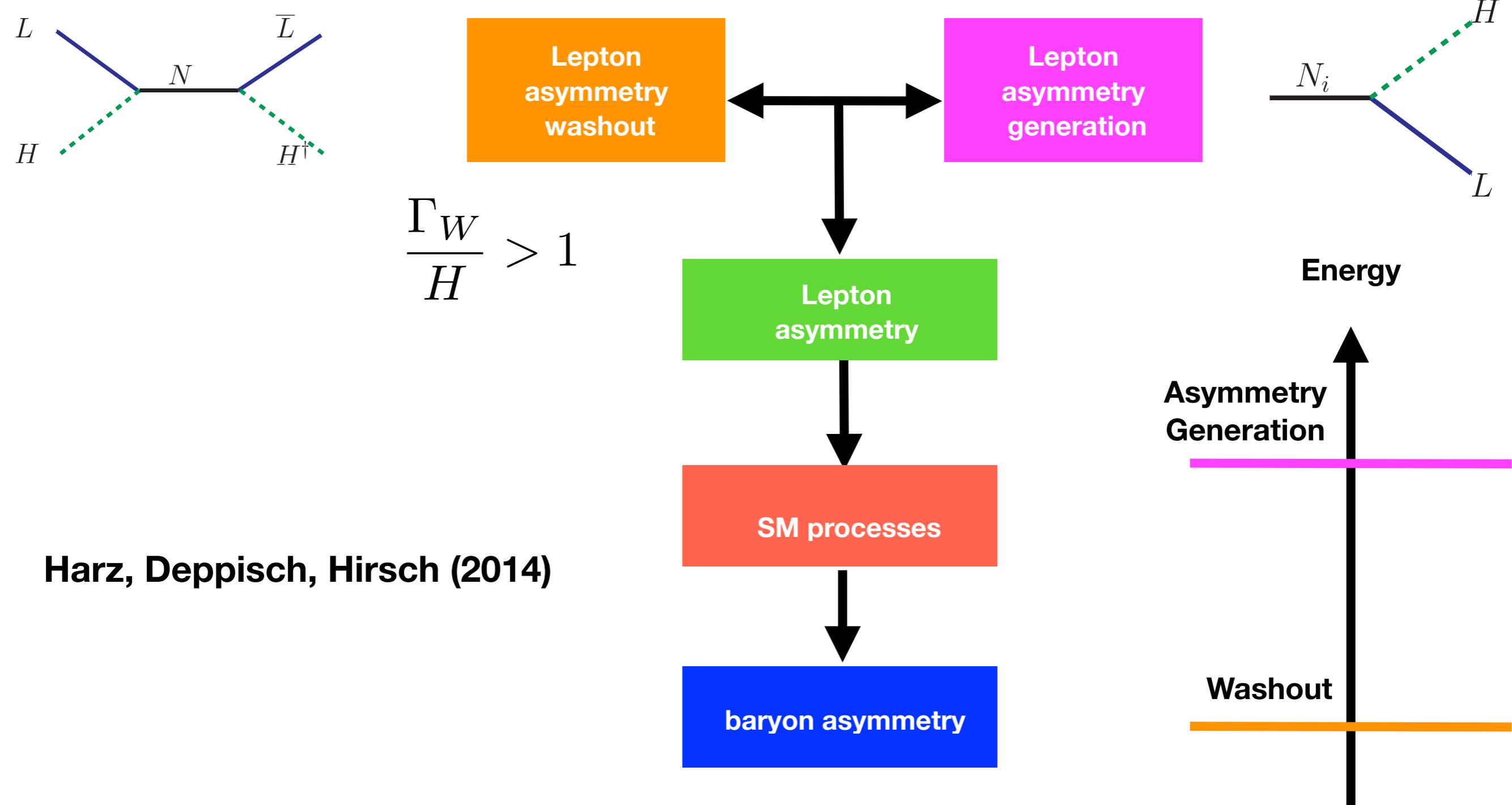
# Thermal leptogenesis

- $U(1)_{B-L}$  used to explain inflation, leptogenesis and neutralino (DM).



Buchmuller, Domcke, Murayama & Schmitz (2019)

# Falsifiability of High-Scale Leptogenesis



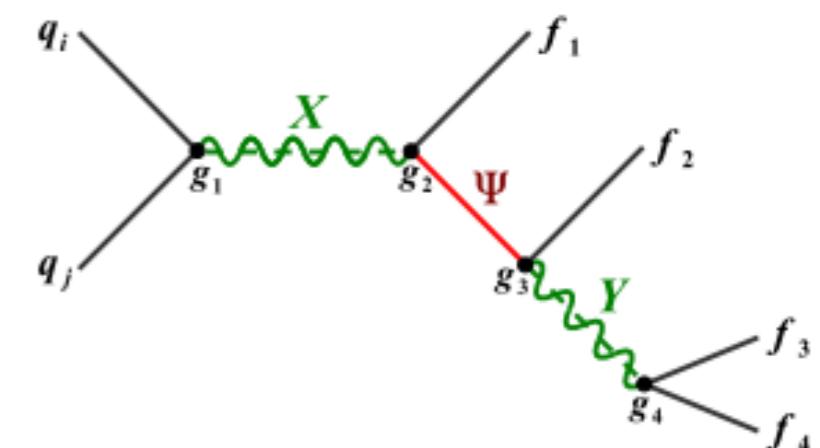
Harz, Deppisch, Hirsch (2014)

If washout processes are large they may be searched for and could possibly falsify leptogenesis

- Observation of LNV (TeV) washout processes at the LHC would falsify high-scale leptogenesis

Deppisch & Harz (2014)

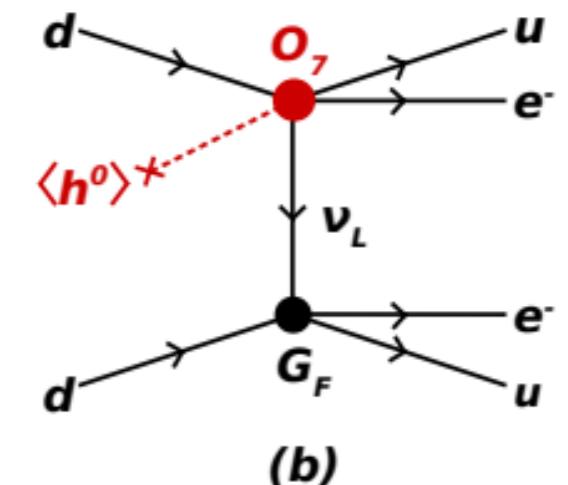
Deppisch, Graf, Harz, Huang (2017)



- Observation of NDBD with new physics from  $>$  dim-5 operators would falsify high-scale leptogenesis

- Caveats:

1. NDBD only probes electron flavour so including flavour effects can evade this
2. Dark U(1) symmetry



- Nonetheless, if one observes LNV at the TeV scale  $\implies$  high-scale leptogenesis can be falsified.

# Thermal leptogenesis and primordial black holes

- BH merger observed via GWs.
- Primordial BHs could have formed in early Universe:  
 $0.1 \leq M_{\text{PBH}} \text{ (g)} \leq 10^9$
- 2010.03565, Yuber Perez-Gonzalez & JT

## PBHs in a nutshell

- Lighter PBHs have shorter lives
- PBH temperature inverse proportional mass

$$T_{\text{PBH}} \approx 1.06 \text{ GeV} \left( \frac{10^{13} \text{ g}}{M_{\text{PBH}}} \right)$$

$$\frac{dM_{\text{PBH}}}{dt} = -\varepsilon(M_{\text{PBH}}) \frac{M_P^4}{M_{\text{PBH}}^2}$$

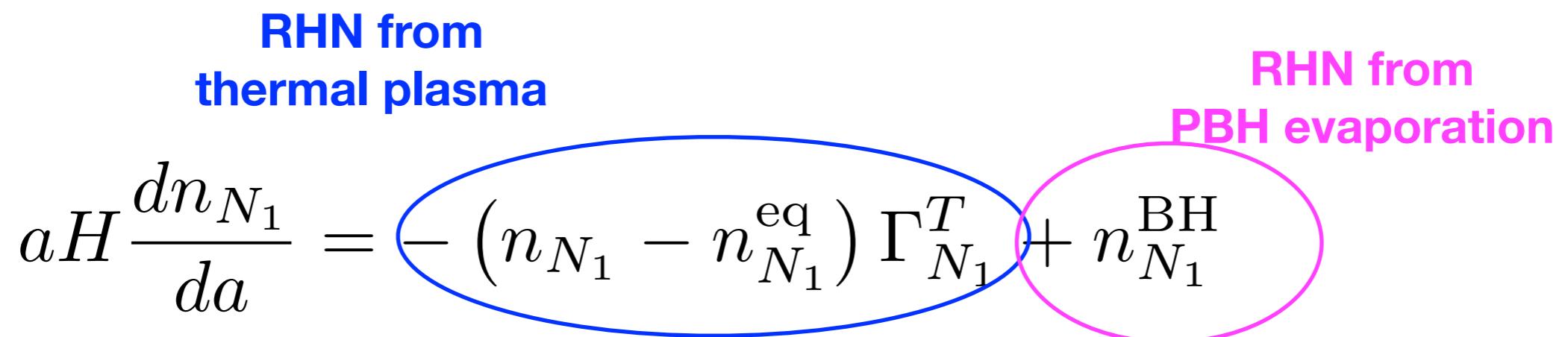
- PBHs produce particles indiscriminately  $T_{\text{PBH}} \gtrsim M$
- Assume PBH dominated EU & monochromatic mass spectrum

# Thermal leptogenesis and primordial black holes

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$$aH \frac{dn_{N_1}}{da} = - \left( n_{N_1} - n_{N_1}^{\text{eq}} \right) \Gamma_{N_1}^T + n_{N_1}^{\text{BH}}$$

**RHN from  
thermal plasma**                            **RHN from  
PBH evaporation**



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**RHN from thermal plasma**      **RHN from PBH evaporation**

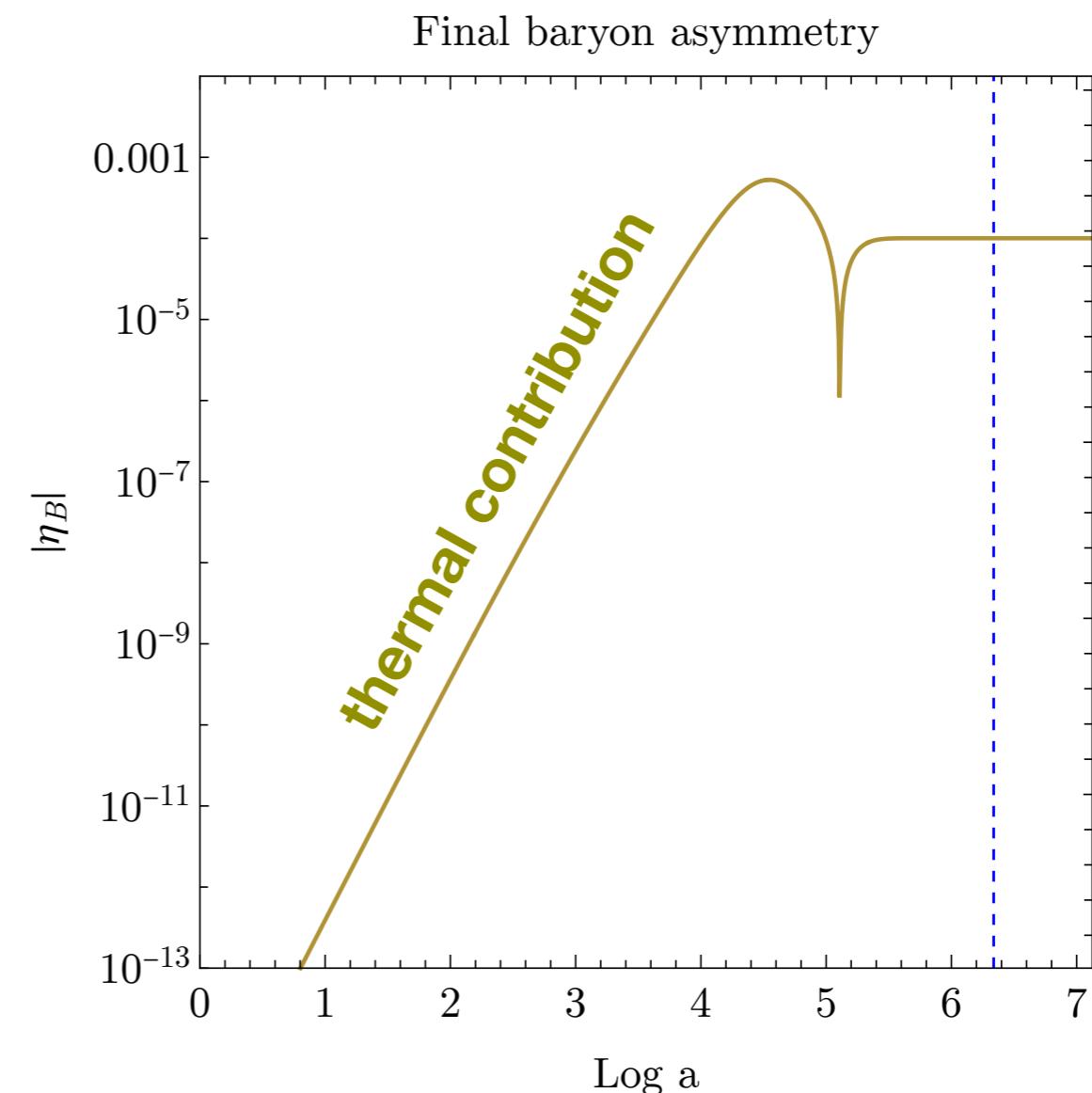
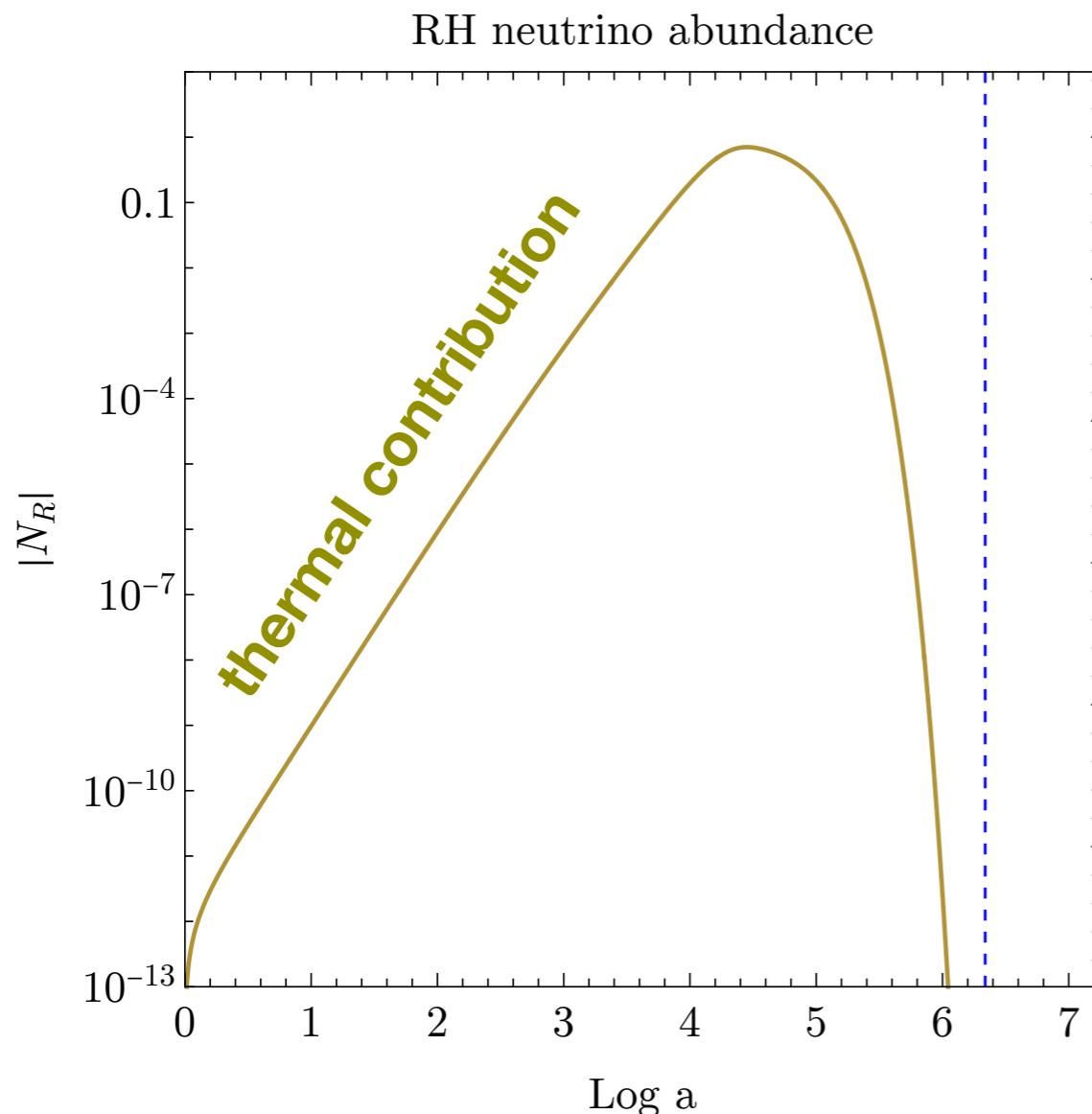
$$aH \frac{dn_{\alpha\beta}^{\text{B-L}}}{da} = \epsilon_{\alpha\beta}^{(1)} [(n_{N_1}^{\text{TH}} - n_{N_1}^{\text{eq}}) \Gamma_{N_1}^T + n_{N_1}^{\text{BH}} \Gamma_{N_1}^{\text{BH}}] - W n_{\alpha\beta}$$

**B-L from thermal leptogenesis**      **B-L asymmetry PBH leptogenesis Hawking radiation**

## A. PBH evaporate during/shortly after RHs

thermally produced  $\rightarrow$  PBHs creates initial condition which gets erased by fast interactions in the plasma

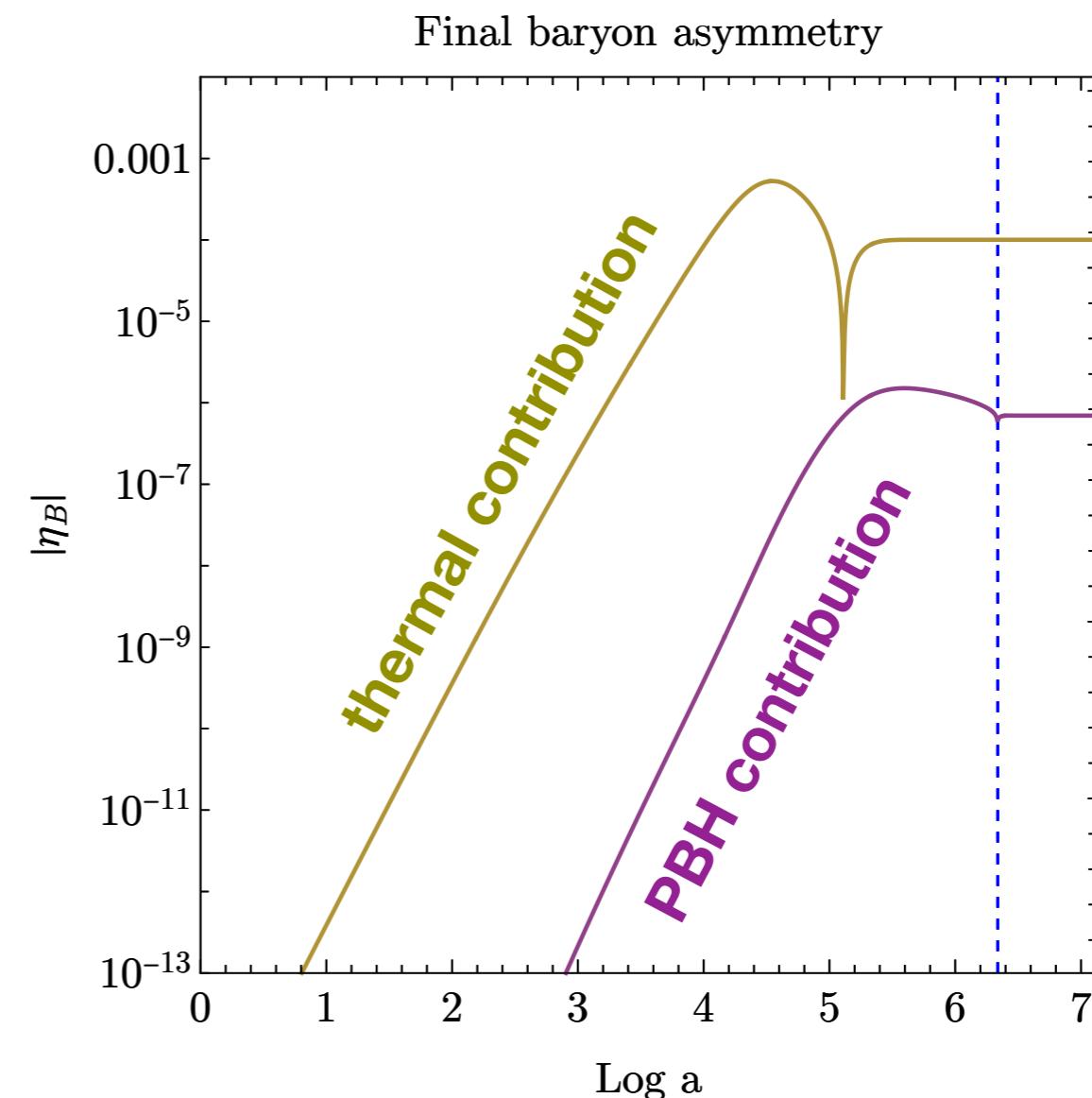
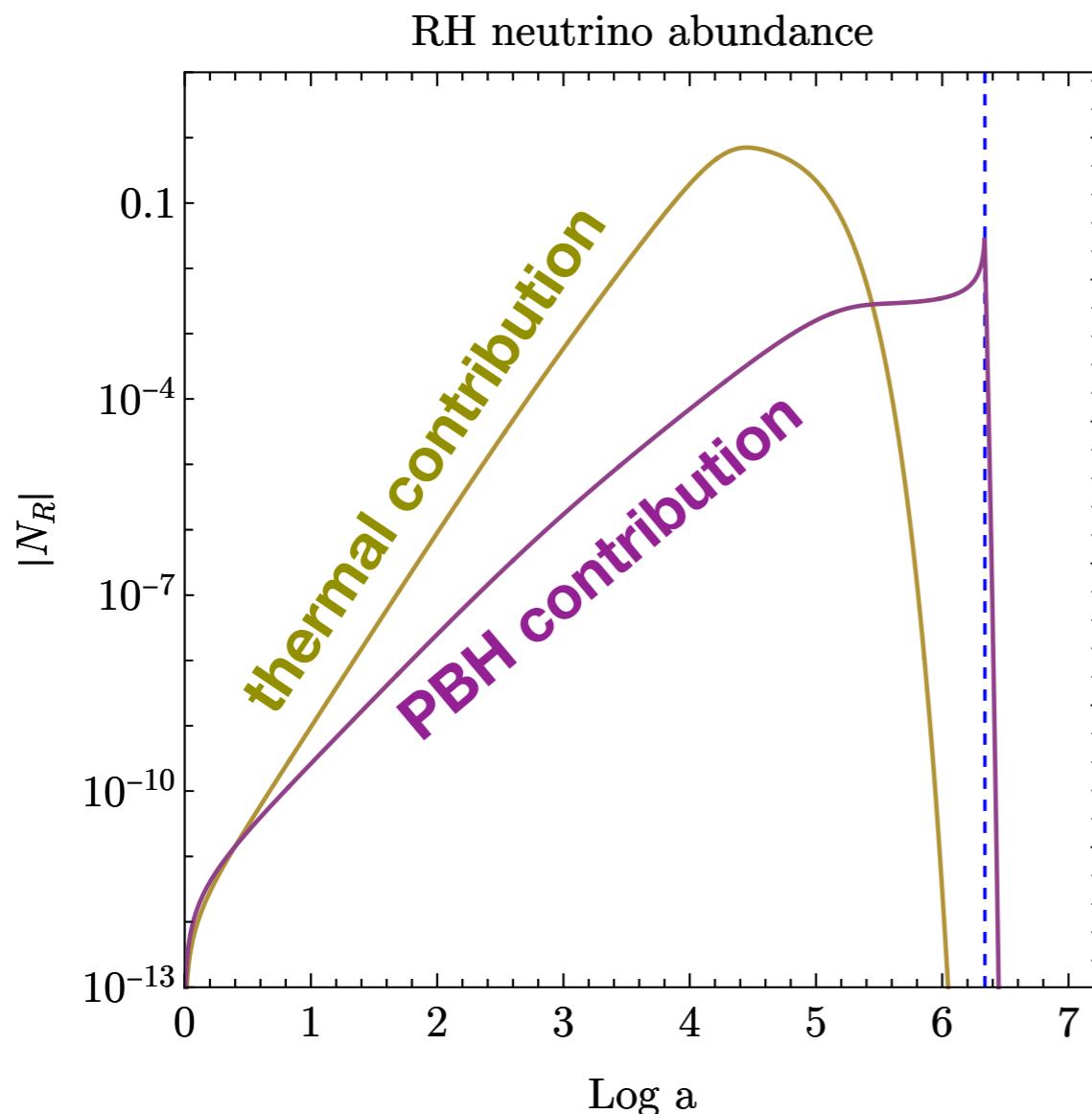
$$M_i = 1.7 \text{ g} \quad \beta' = 10^{-3} \quad M_N = 10^{11} \text{ GeV}$$



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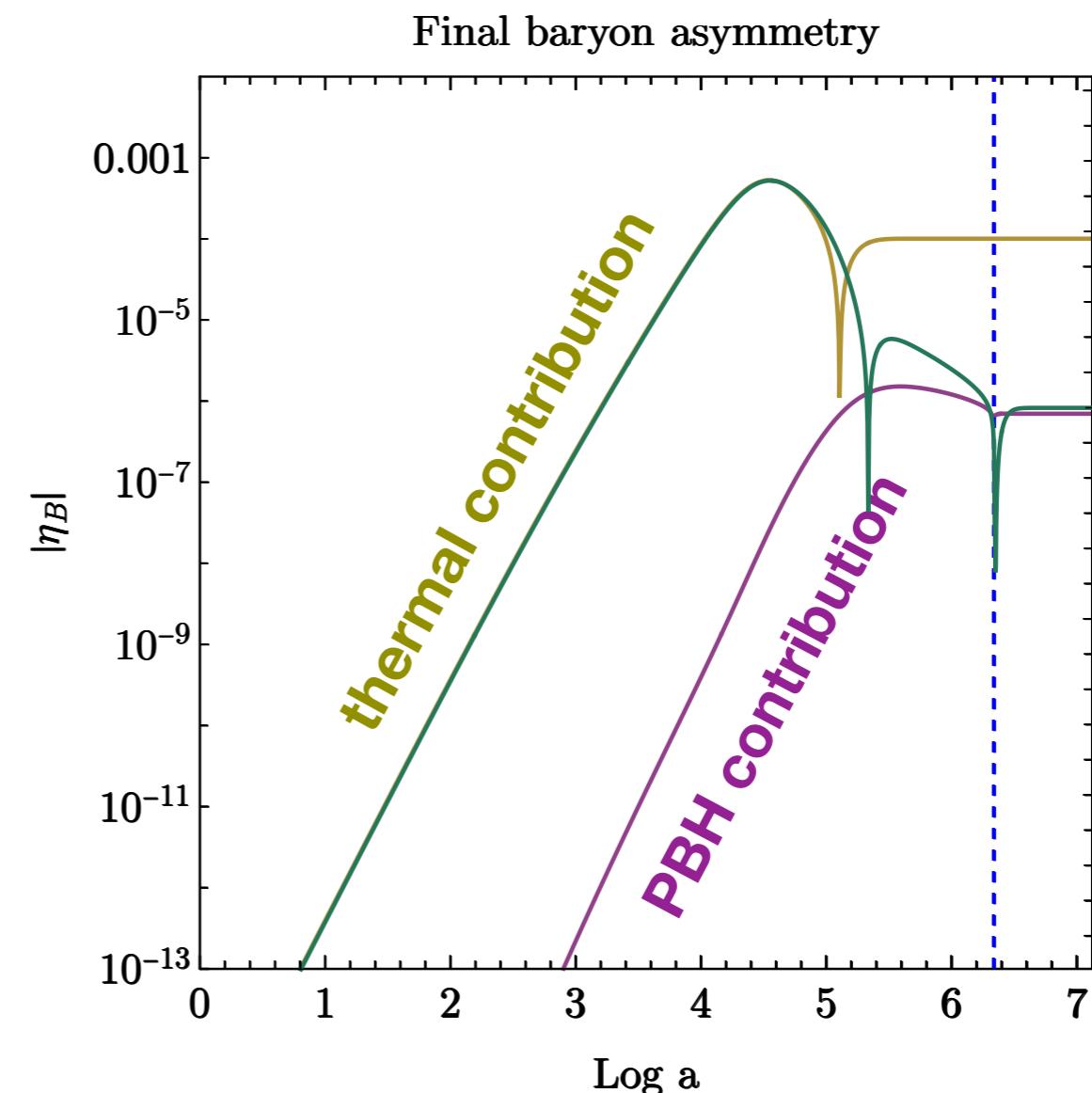
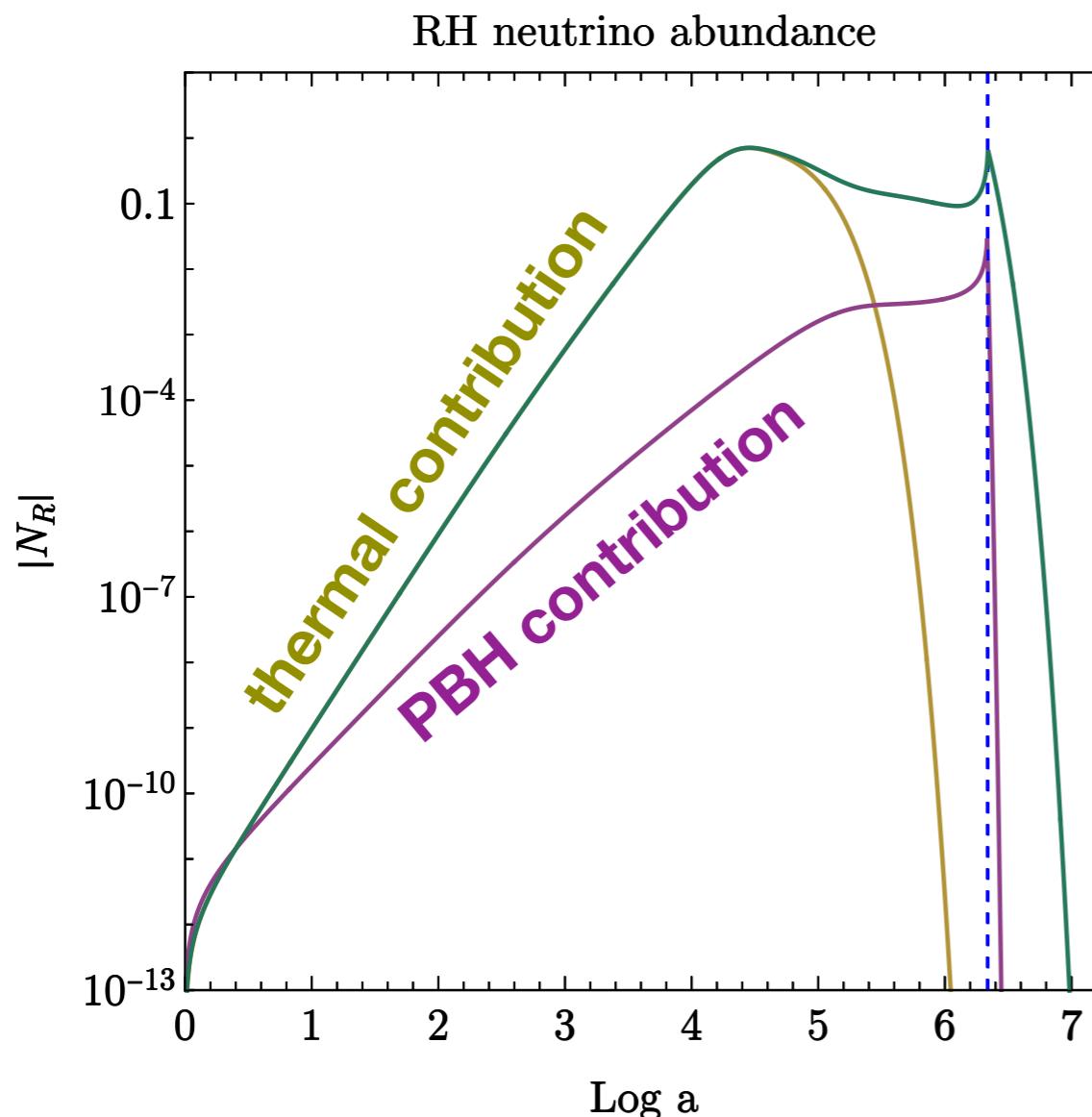
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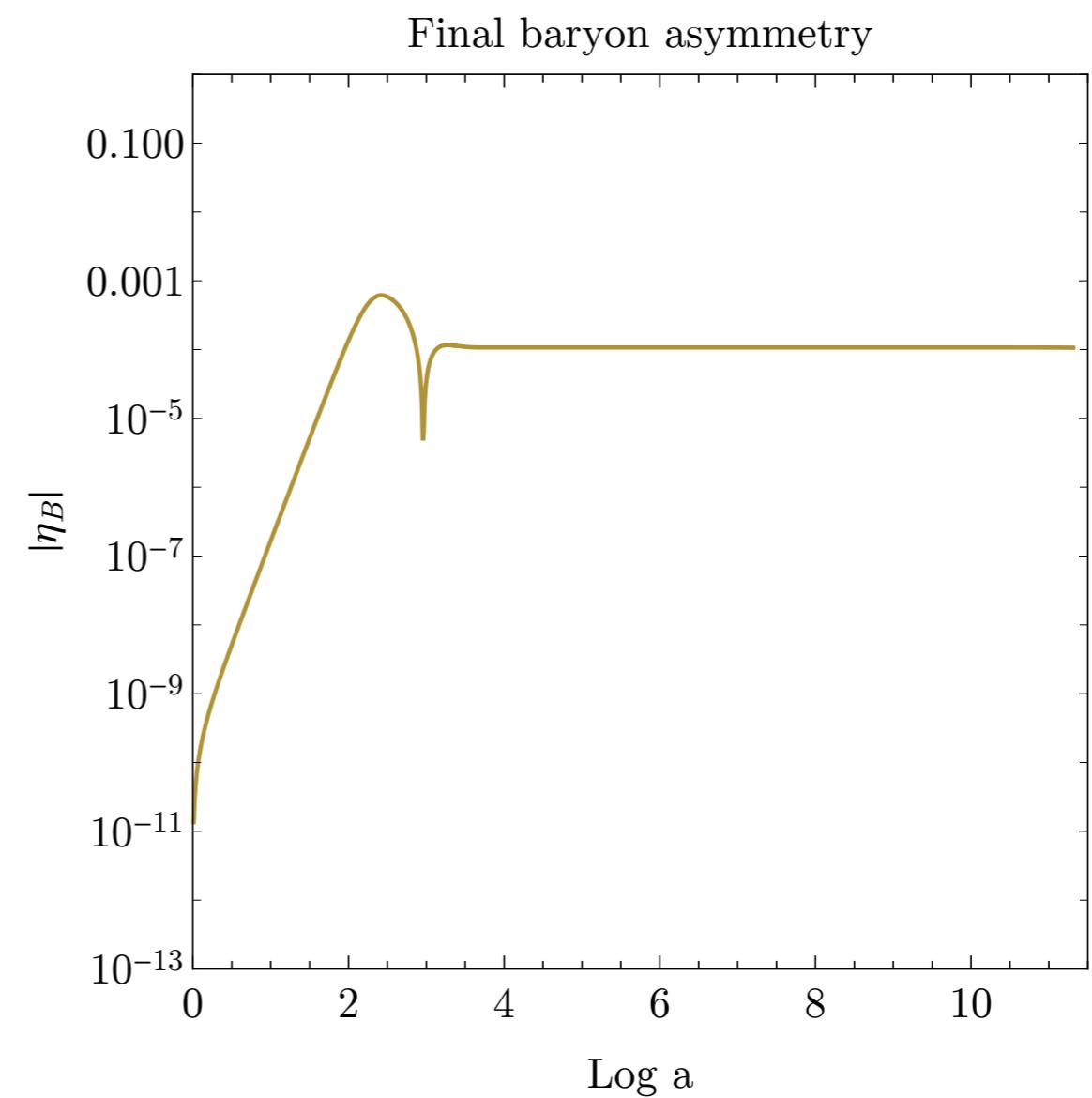
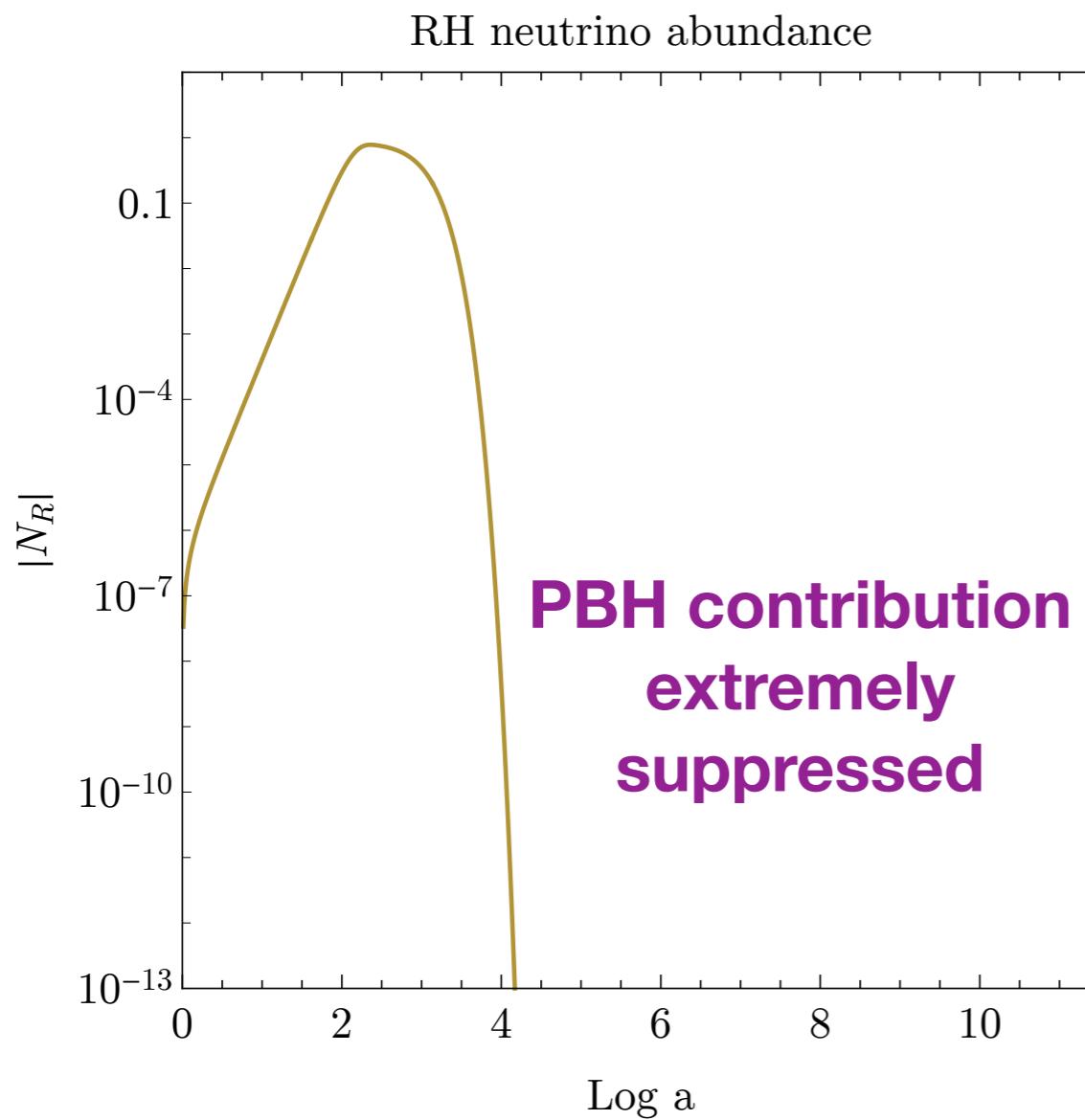
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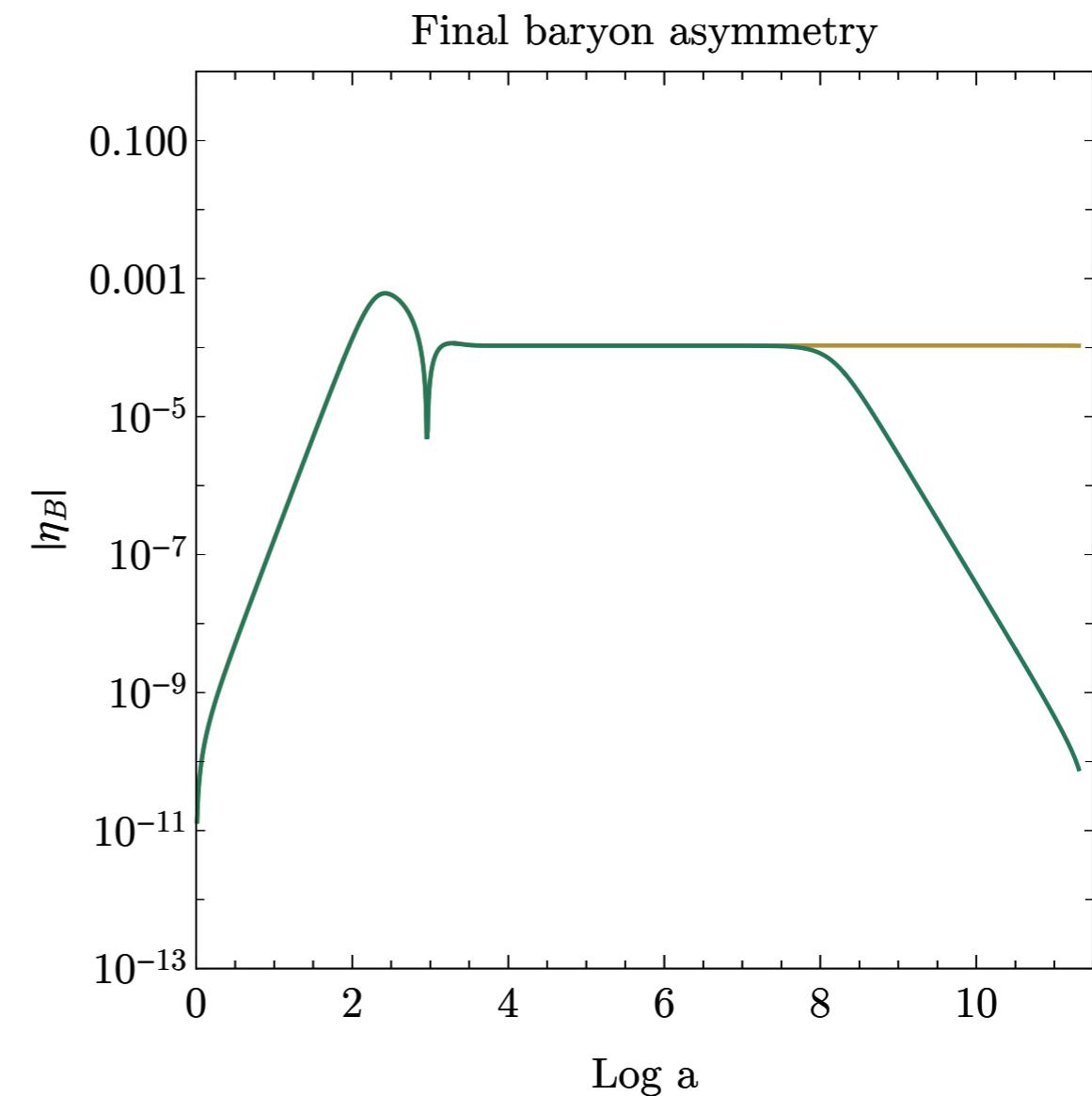
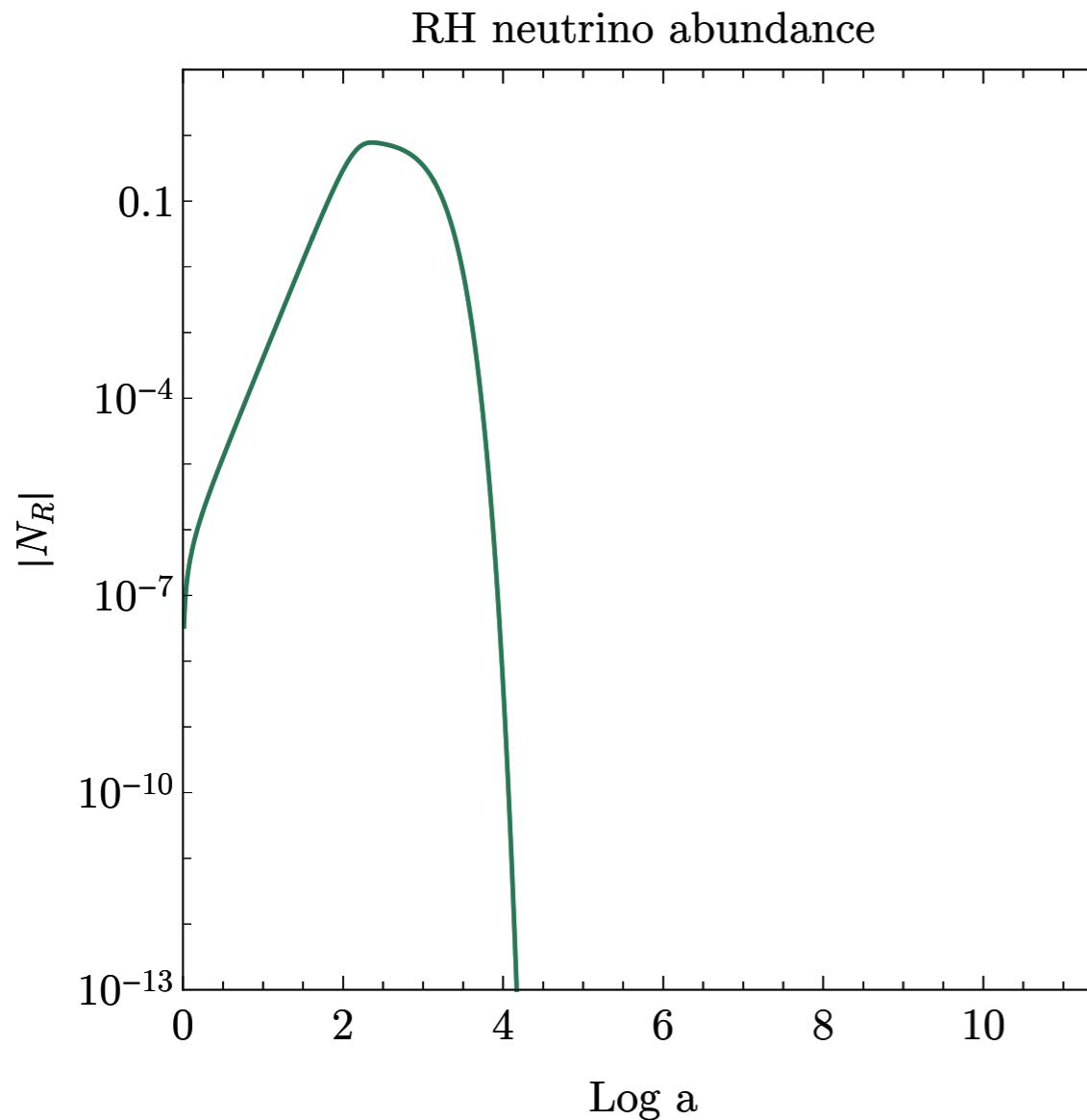
## D. PBH evaporation occurs **way after** thermal leptogenesis era

$$M_i = 10^4 \text{ g} \quad \beta' = 10^{-3} \quad M_N = 10^{11} \text{ GeV}$$



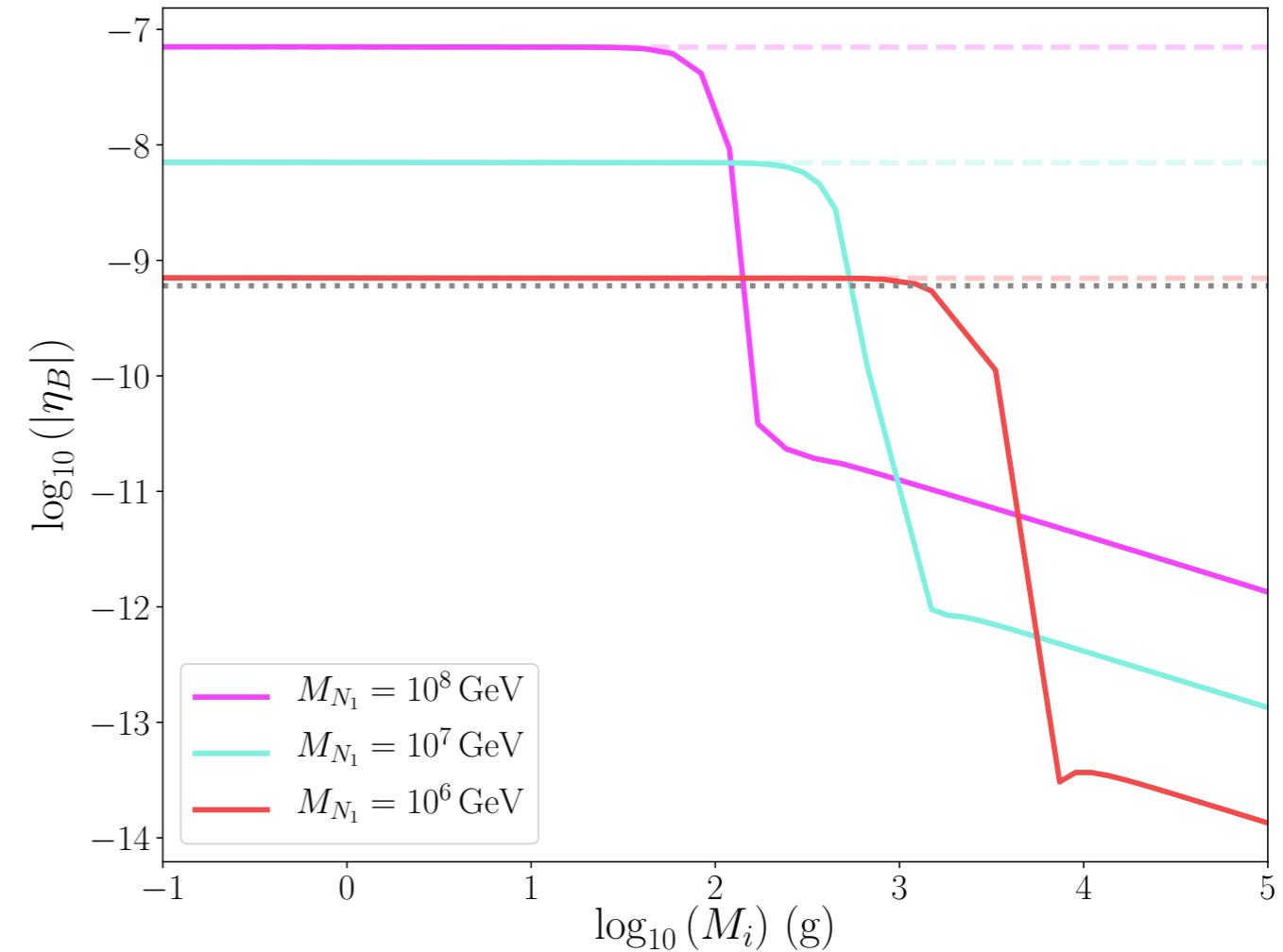
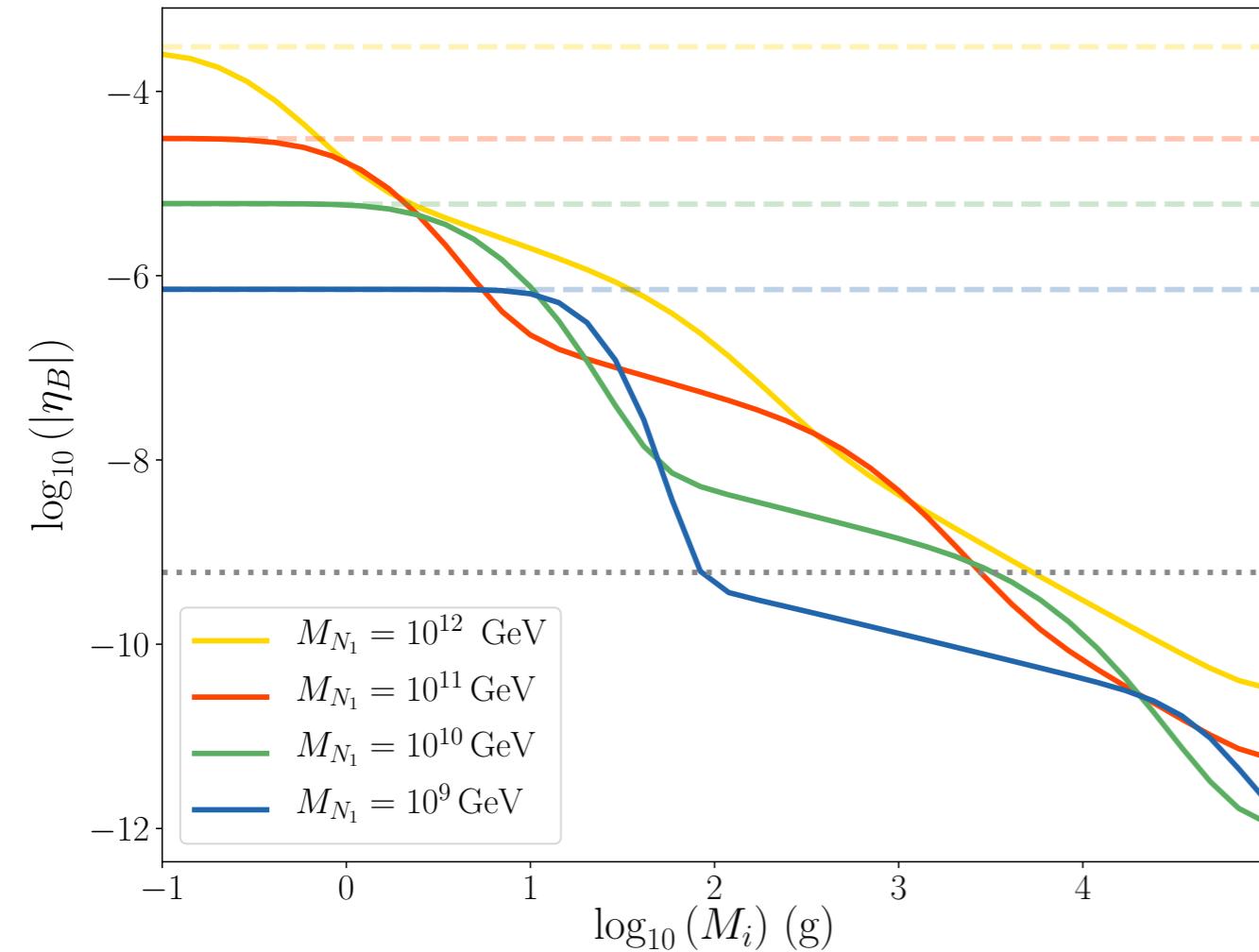
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# Thermal leptogenesis and primordial black holes

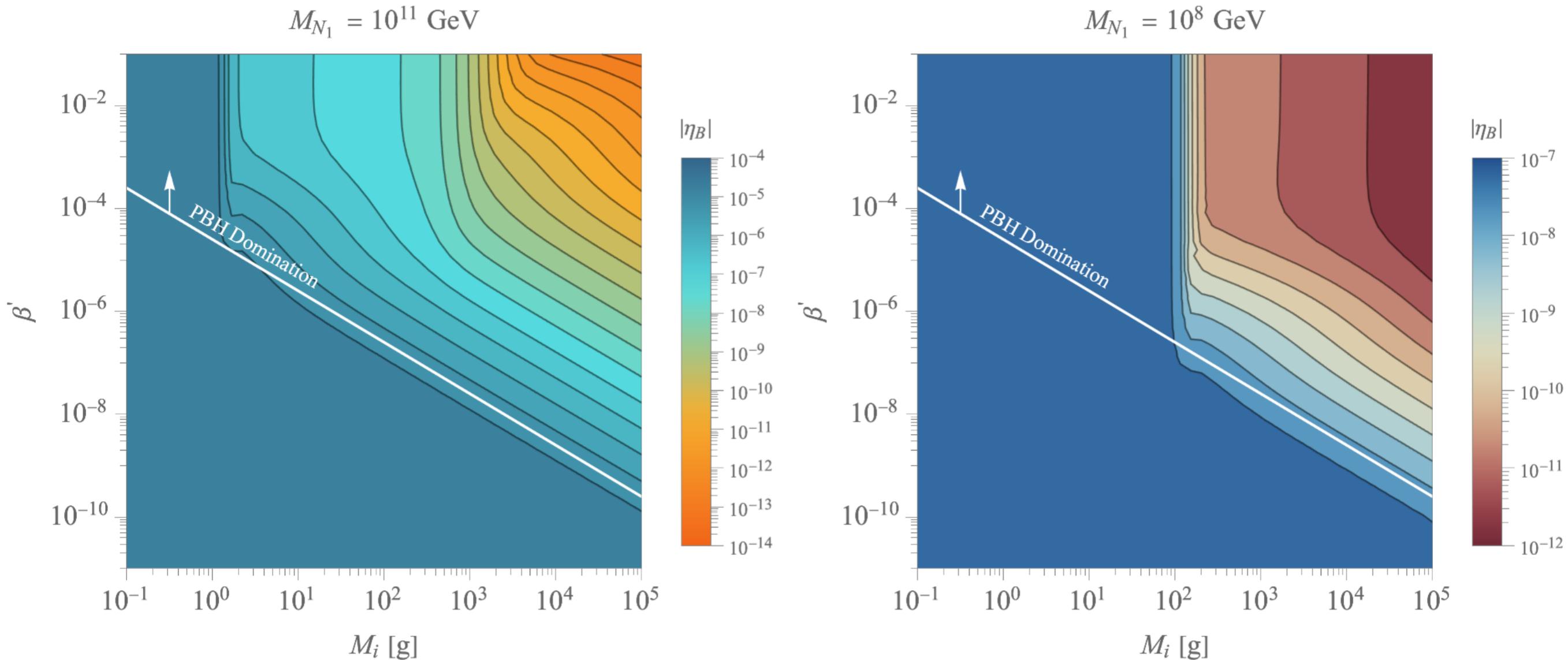
$$\beta' = 10^{-3}$$



Chose Yukawa matrix for maximal baryon asymmetry  
Lines indicates upper bound

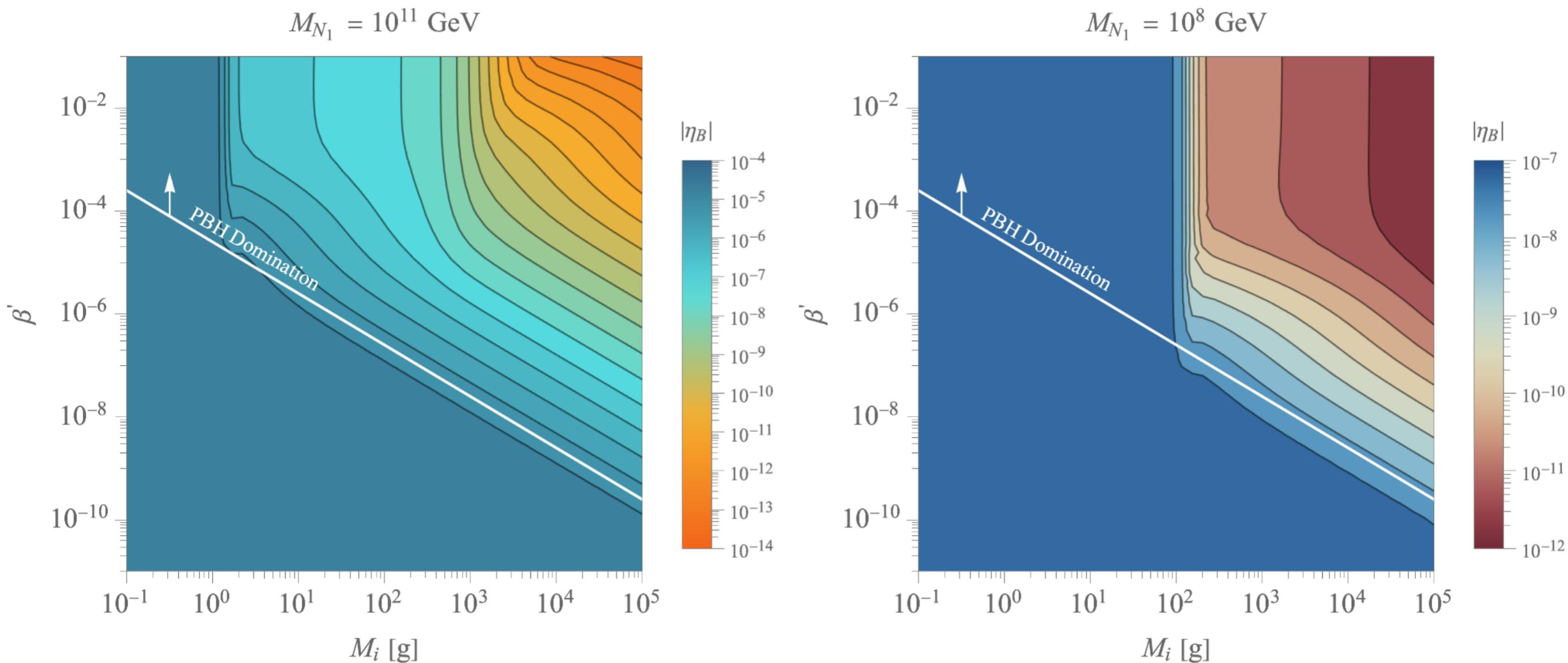
# Thermal leptogenesis and primordial black holes

Dilution effect present as long as there is PBH domination

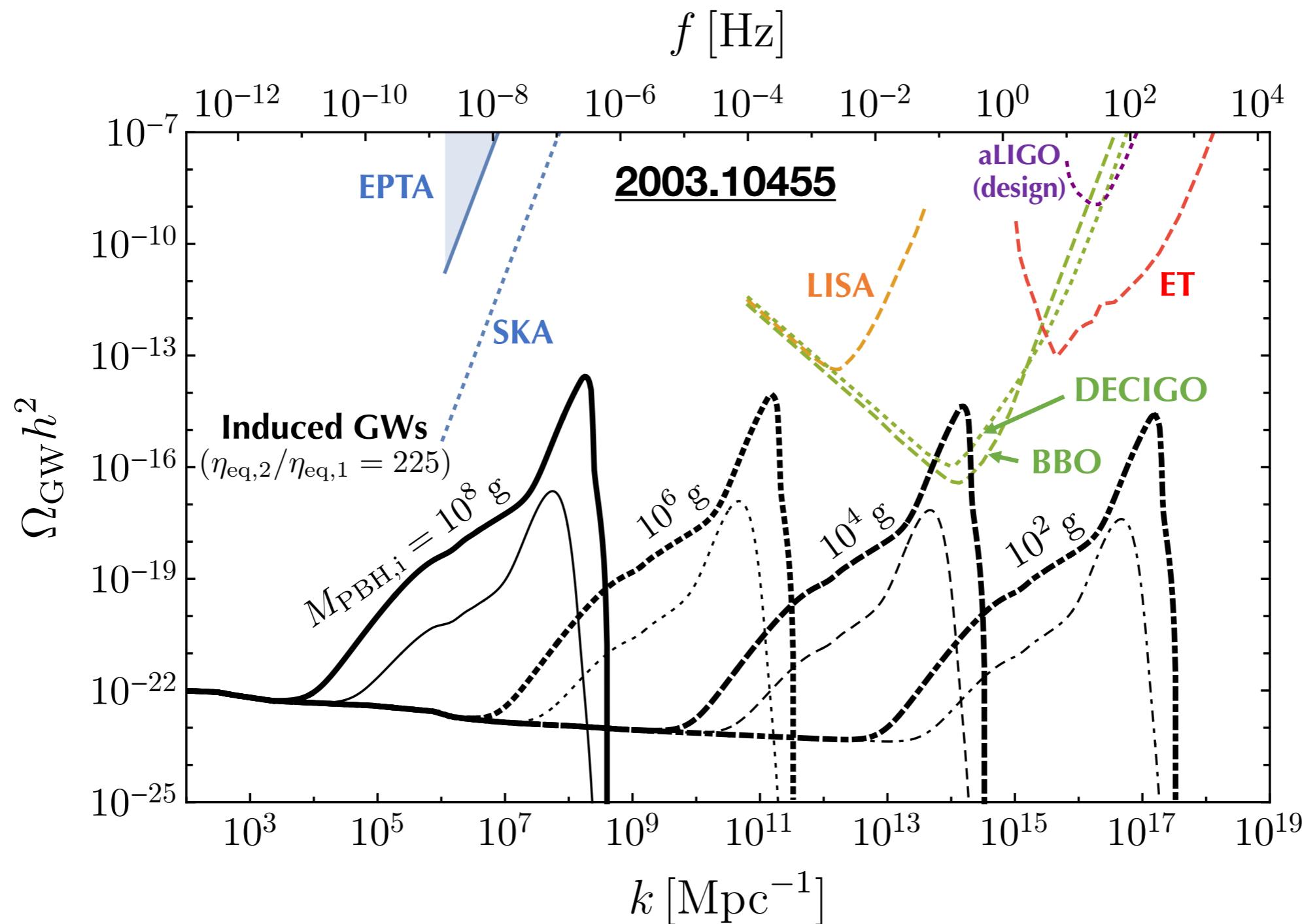


# Thermal leptogenesis and primordial black holes

Detection of PBHs in mass range  $> 0.1$  kg would place thermal leptogenesis under serious tension.



When PBHs evaporate  $\Rightarrow$  sudden transition  
 Matter to radiation domination  $\Rightarrow$  gravitational potential  
 oscillates  $\Rightarrow$  SGWB generated



# Summary

- Leptogenesis is a plausible explanation for the smallness of neutrino masses and the observed matter anti-matter asymmetry
- In the type-I seesaw framework for leptogenesis, the mass of the RHN can range from GeV -  $10^{13}$  GeV scale.
- Low-scale (and some regions of resonant) leptogenesis can be probed by a broad range of present and future experimental facilities (see Juraj's talk)
- Gravitational waves are a complementary probe of intermediate and high-scale leptogenesis and LNV & LFV searches will be key.