



Review of current and future neutrinoless double beta decay experiments

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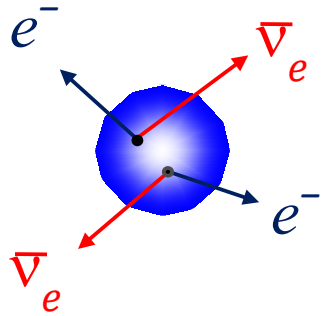
**8th Workshop on Theory, Phenomenology and Experiments in Flavour Physics
Jun 11 – 13, 2022 - Villa Orlandi, Anacapri, Capri Island, Italy**

Double Beta Decay

Rare Nuclear Decay

$$(A, Z) \rightarrow (A, Z+2) + 2e^- + [...]$$

occurs in a number of even-even nuclei in
A even multiplets



$\beta\beta$ -2 ν : two neutrino mode

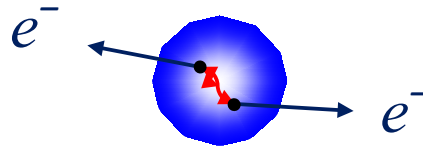
$$(A, Z) \rightarrow (A, Z+2) + 2e^- + 2\bar{\nu}_e$$

- allowed in Standard Model
- second order weak transition

$\beta\beta$ -0 ν : neutrinoless mode

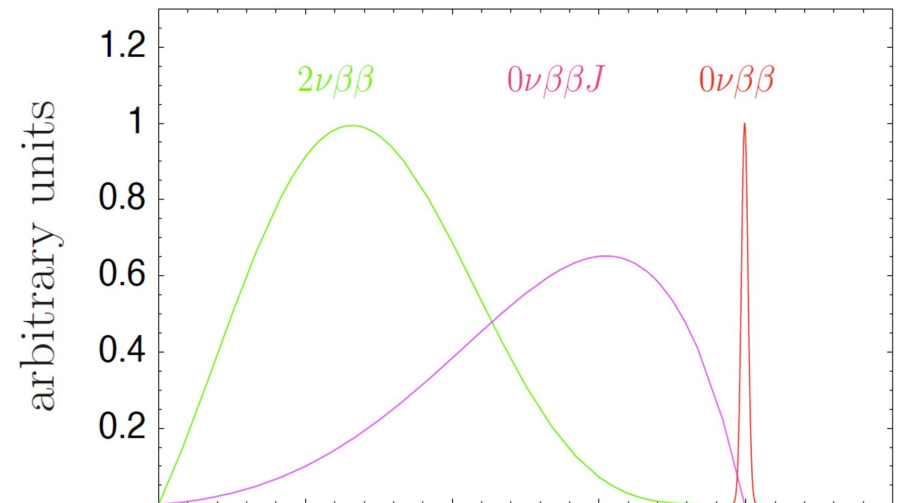
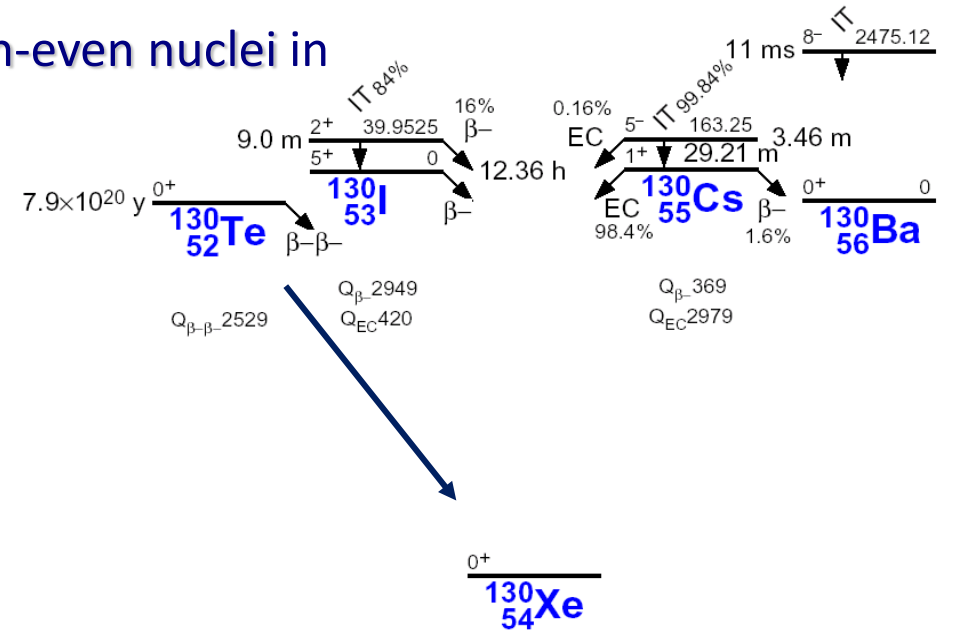
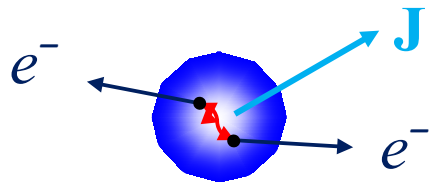
$$(A, Z) \rightarrow (A, Z+2) + 2e^-$$

- not allowed in SM ($\Delta L=2$, B-L violation)
- neutrino is a massive Majorana particle
- Majorana phases



$\beta\beta$ -0 ν : "exotic" modes

$$(A, Z) \rightarrow (A, Z+2) + \text{exotic particles}$$



Neutrinoless Double Beta Decay Process

With a specific approach and introducing some important approximations:

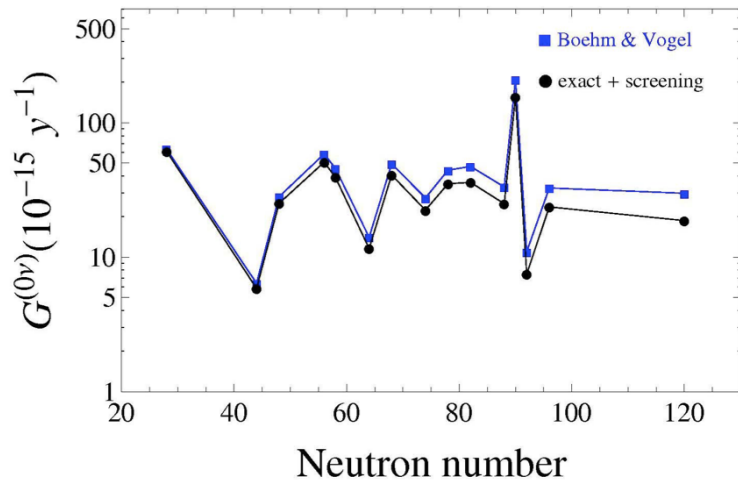
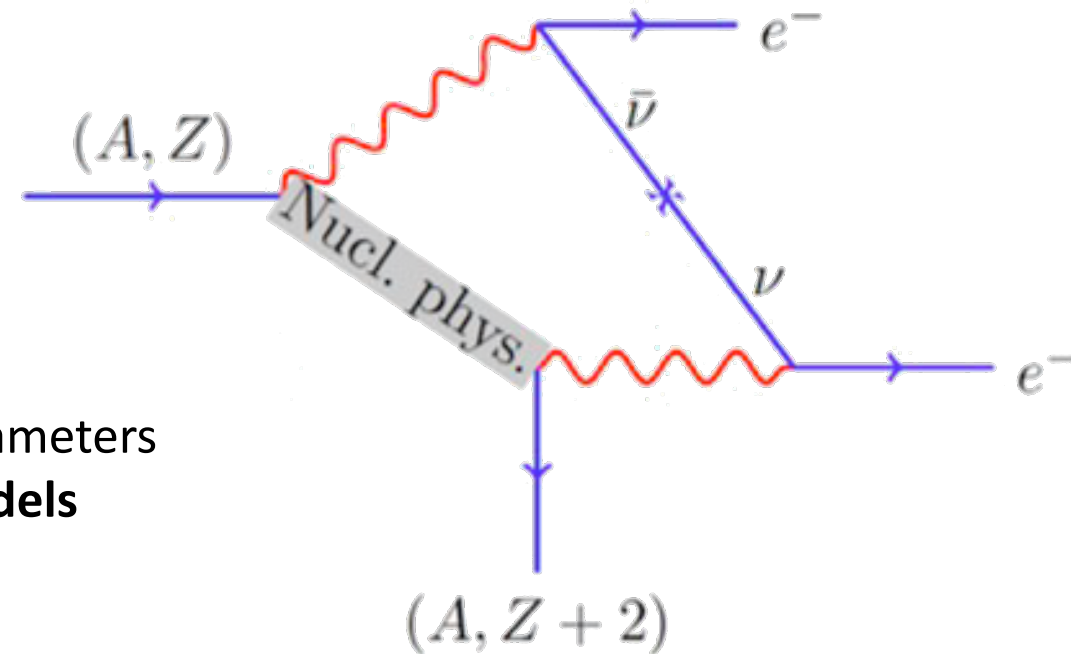
$$\Gamma^{0\nu} = G_x(Q, Z) |M_x(A, Z)|^2 |\eta_x|^2$$

phase space factor
calculable precisely

nuclear matrix element
not well established

particle physics parameters
depends on models

- massive Majorana neutrinos
- GUT's
- SUSY
- ...



F. Böhm and P. Vogel, *loc. cit.*

J. Kotila and F. Iachello, Phys. Rev. C 85, 034316 (2012)

Nuclear Matrix Elements

$M_{0\nu}$ values estimated from different Nuclear Matrix calculation methods

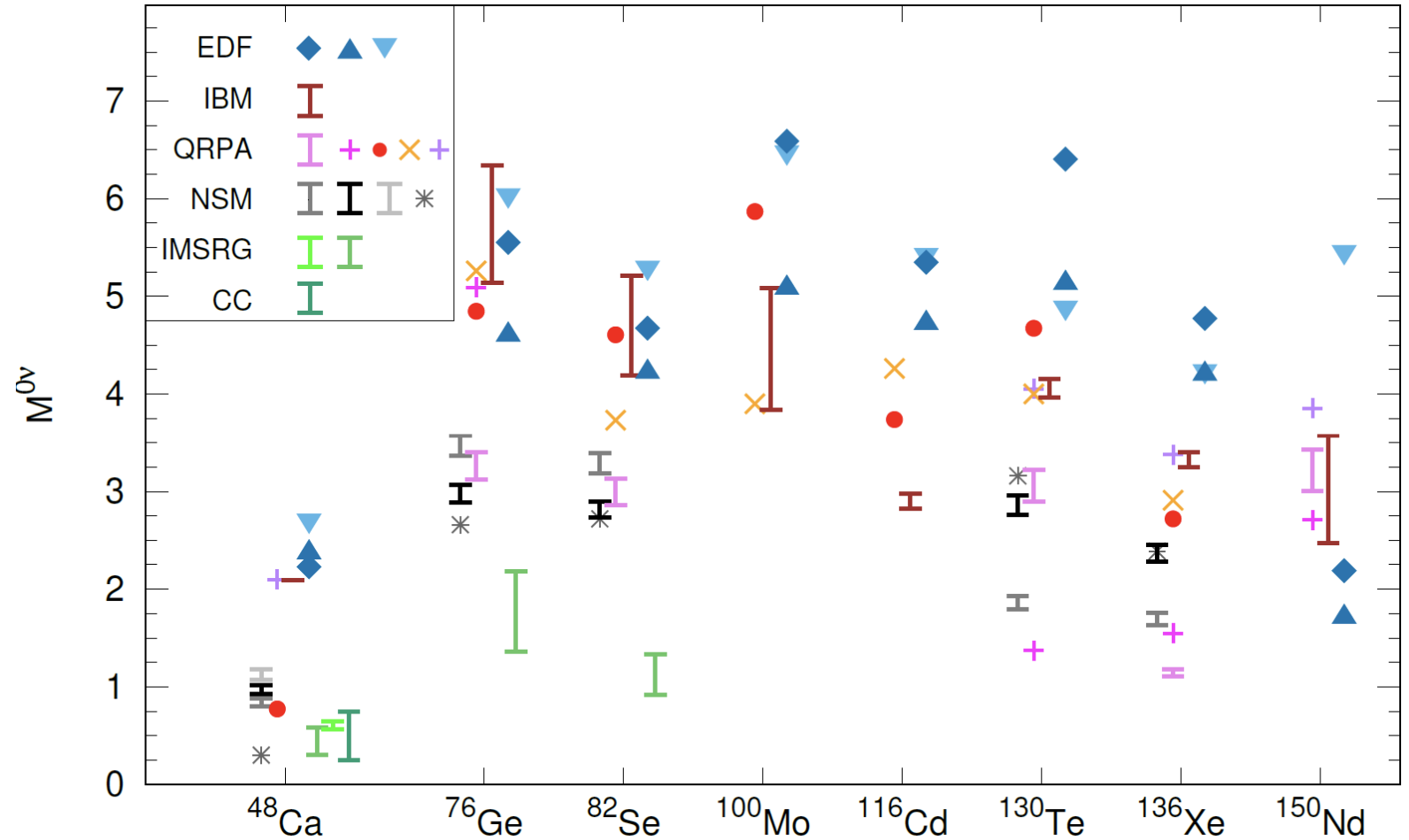
NME calculations:

- **EDF**: large
- **QRPA**: larger spread
- **NSM**: small
- **IMSRG** ^{48}Ca ab initio: very small



- Significant differences
- Results inaccuracy

g_A quenching?



Light Majorana Neutrinos Mechanism

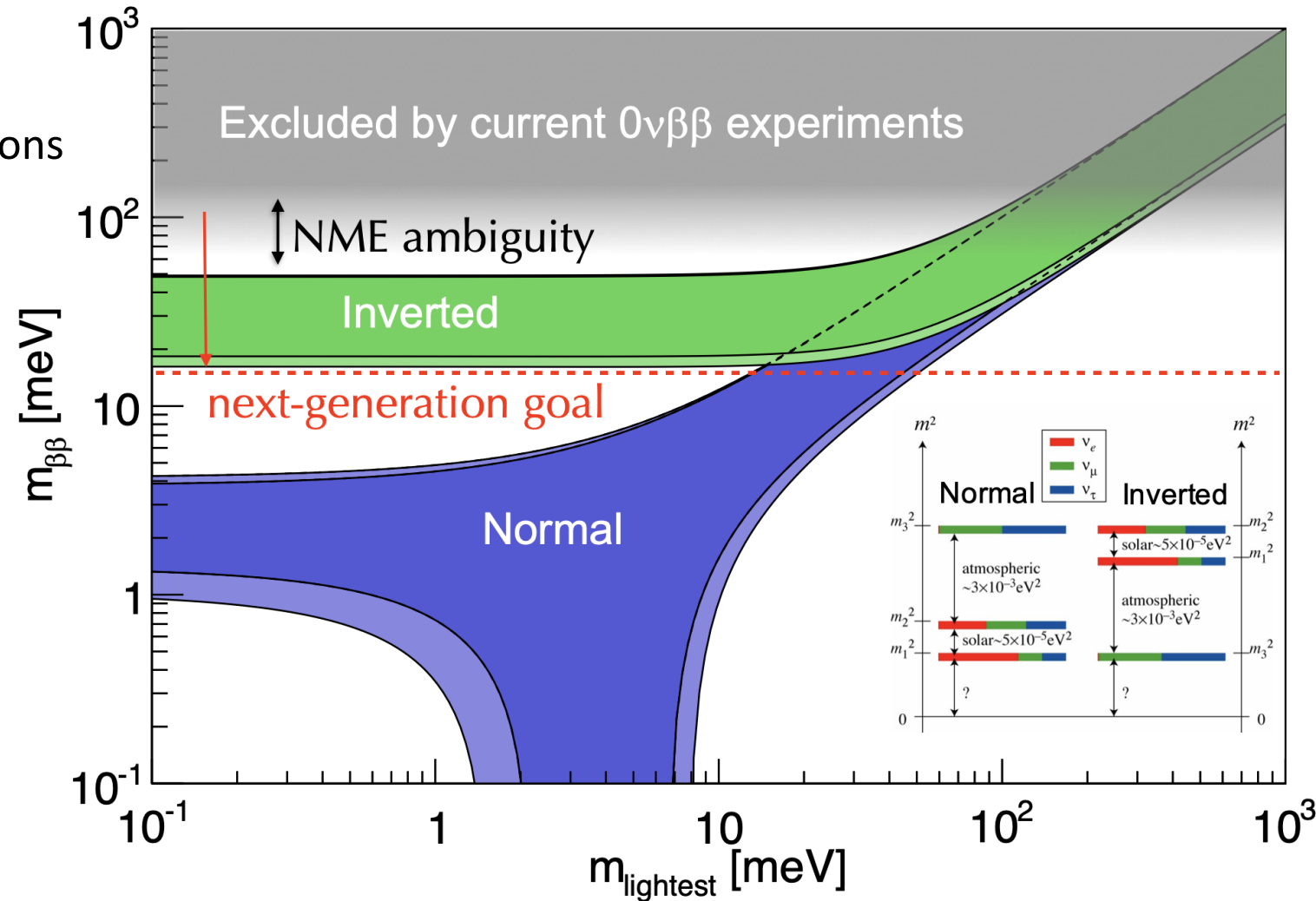
$\beta\beta 0\nu$ mediated by exchange of light Majorana neutrinos other mechanisms give no contributions

$$\eta_x = \langle m_{ee} \rangle = \sum_k U_{ek}^2 m_k$$

$$= c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 e^{i\alpha} m_2 + s_{13}^2 e^{i\beta} m_3$$

m_{ee} depends on three masses and two Majorana phases

- Transition amplitude is proportional to the coherent sum of neutrino masses
- Majorana phases play a crucial role: possible cancellations
- Oscillation experiments could identify only neutrino mass differences



Experimental Sensitivity

$$\tau_{1/2}^{0\nu} = \ln 2 \frac{\epsilon N_{\text{nuclei}} t_{\text{meas}}}{N_{\beta\beta}}$$

$$N_{\beta\beta} \leq \sqrt{\text{bkg} \cdot \Delta E \cdot M \cdot t_{\text{meas}}}$$

$$N_{\text{bkg}} \gg 1 \quad S_{1/2}^{0\nu} \propto \epsilon \frac{i.a.}{A} \sqrt{\frac{M \cdot t_{\text{meas}}}{\text{bkg} \cdot \Delta E}}$$

$$N_{\text{bkg}} \sim 1 \quad S_{1/2}^{0\nu} \propto \epsilon \frac{i.a.}{A} M \cdot t_{\text{meas}}$$

Many important Parameters

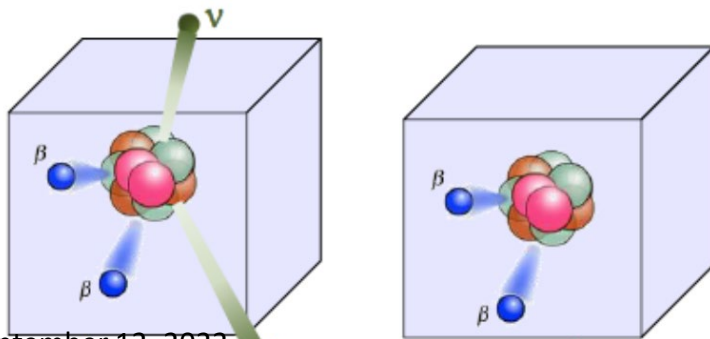
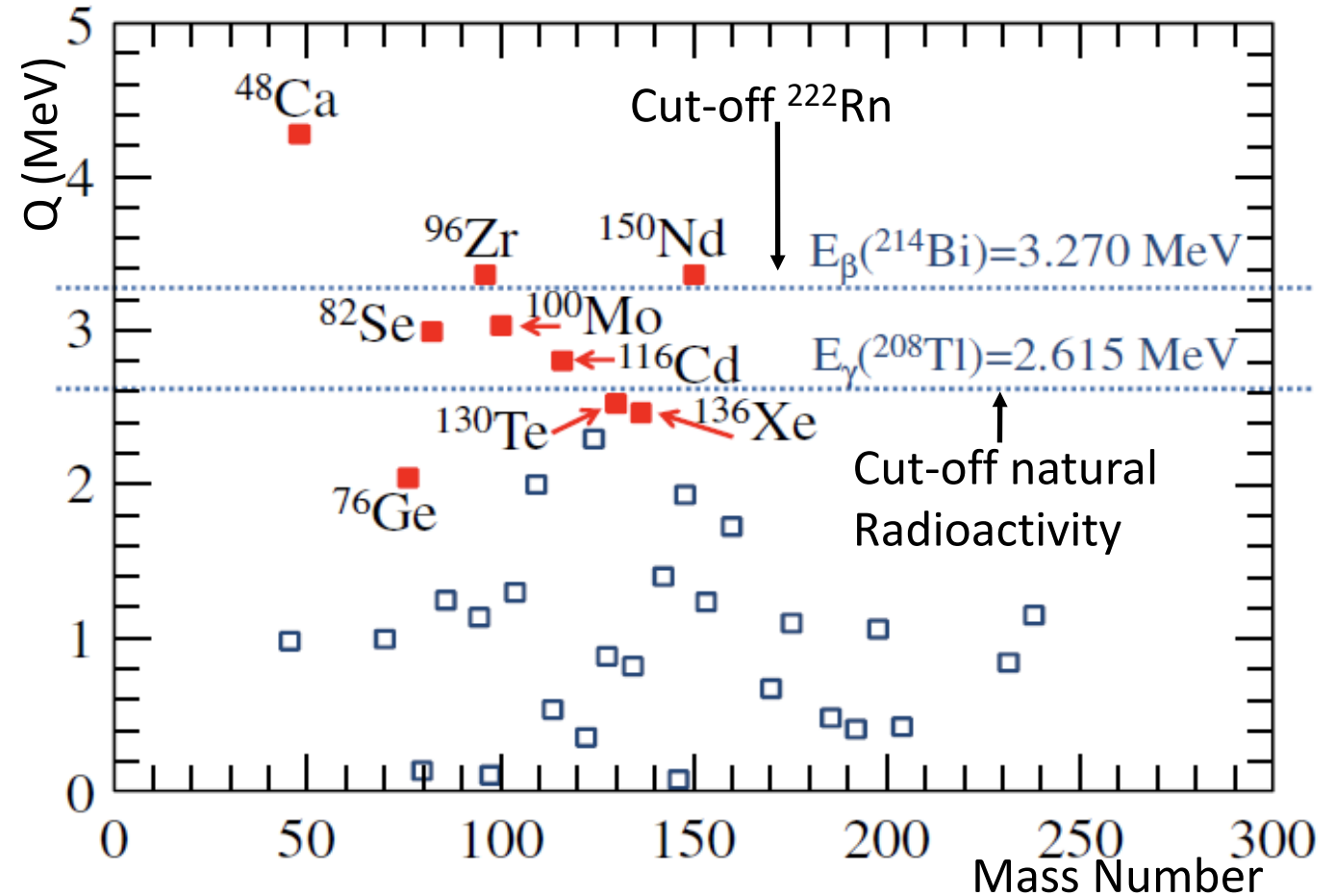
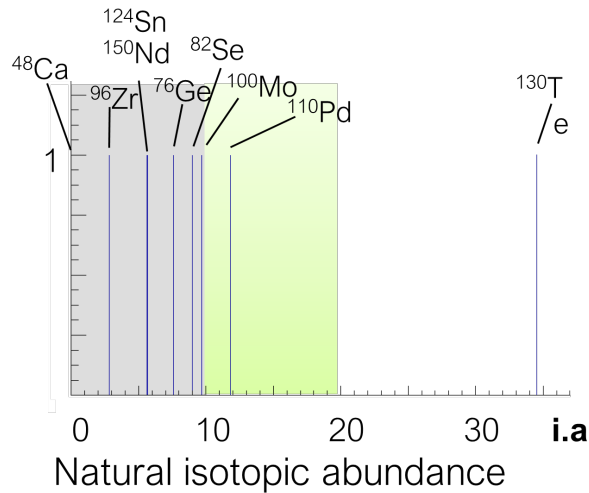
N_{nuclei}	number of active nuclei
t_{meas}	measurement time [y]
M	mass of the detector [kg]
ϵ	detection efficiency
$i.a.$	isotopic abundance
A	atomic number
ΔE	energetic resolution [keV]
bkg	background index [c/keV/y/kg]

- isotope choice/enrichment
- experimental approach
- scalability
- stability
- cost

Isotope selection

From the Table of Isotopes

- 35 isotopes with double beta decay transitions
- 9 promising for sensitive measurements
- most promising candidates: $Q_{\beta\beta} > 2-3$ MeV
- isotope enrichments are needed



Considering a calorimetric approach (**Source == Detector**)

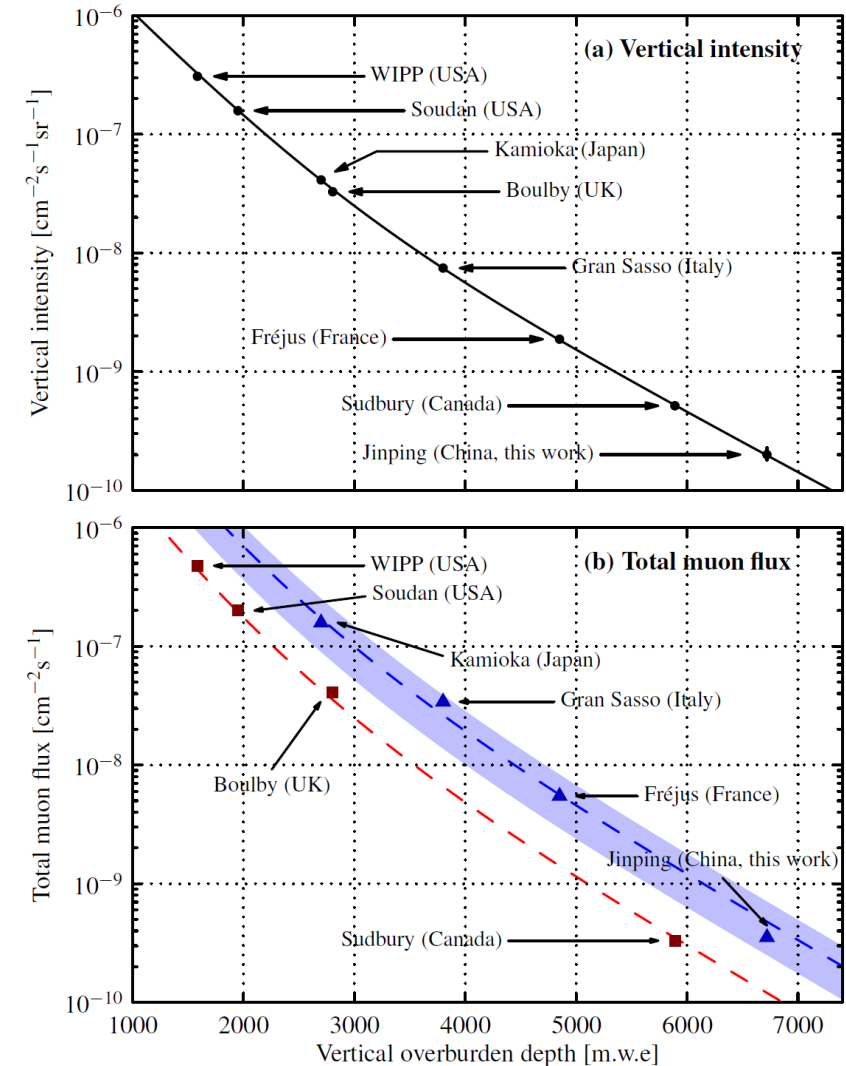
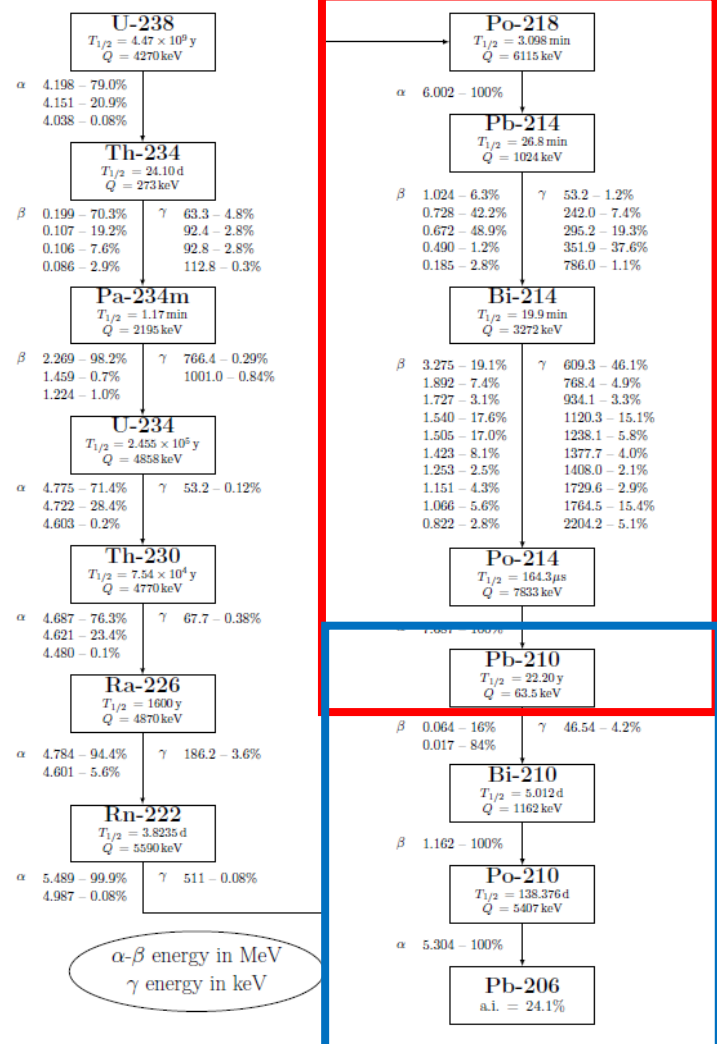
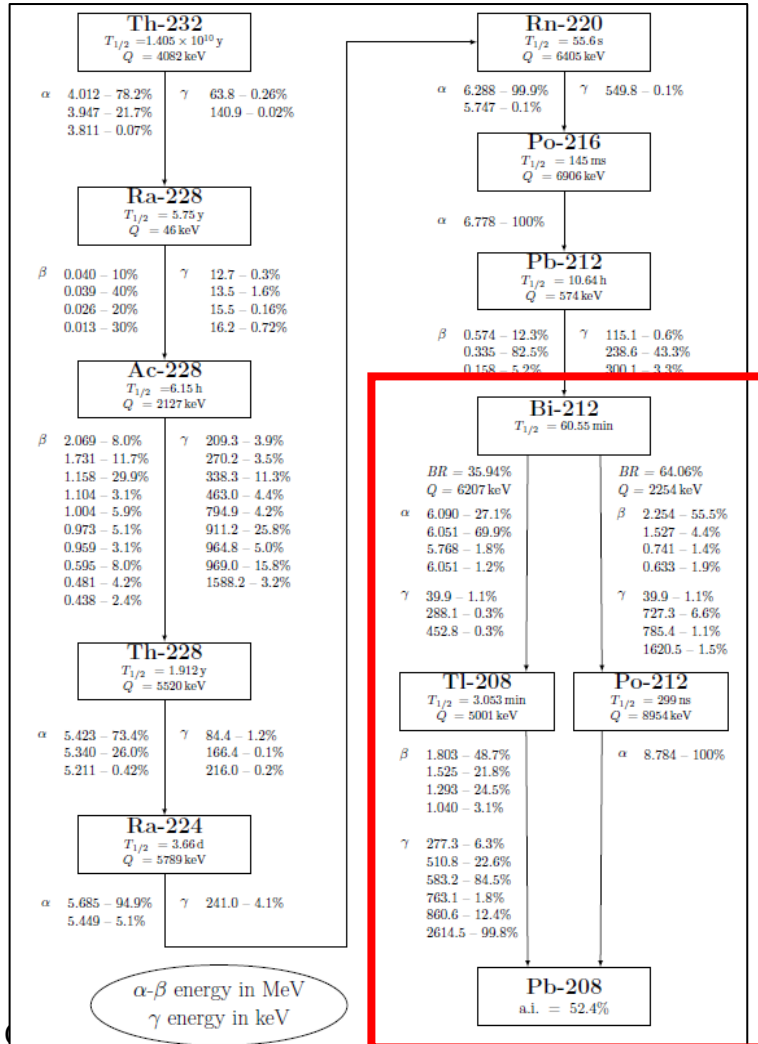
- **isotope enrichments** are needed
- **very clean materials** have to be identified

Background interferences

It is necessary to strongly reduce background events:

- Cosmic Rays
- Radioactivity

Guo et al., Chinese Physics **C45** (2021) 025001, arXiv:2007.15925v2



Possible Calorimetric Approaches

Fluid or diluted in a fluid

Source

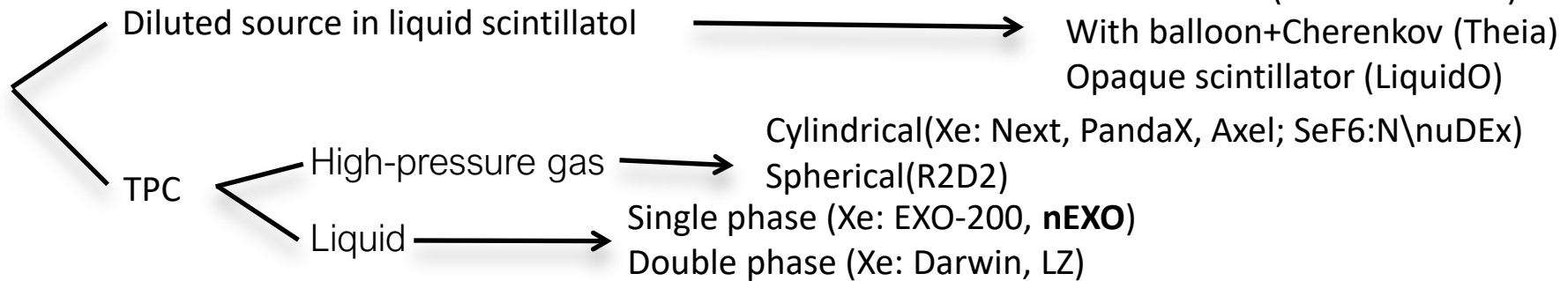
Integral part of a crystal

- Container size
- Concentration of the source

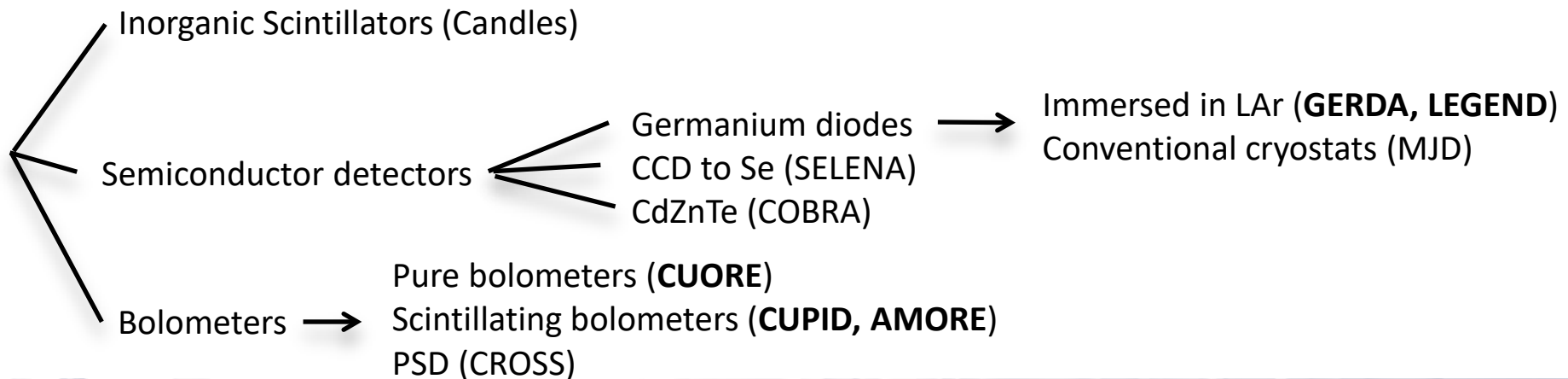
Scalability

- Size of a single crystal
- Number of crystals

Large volumes of fluid



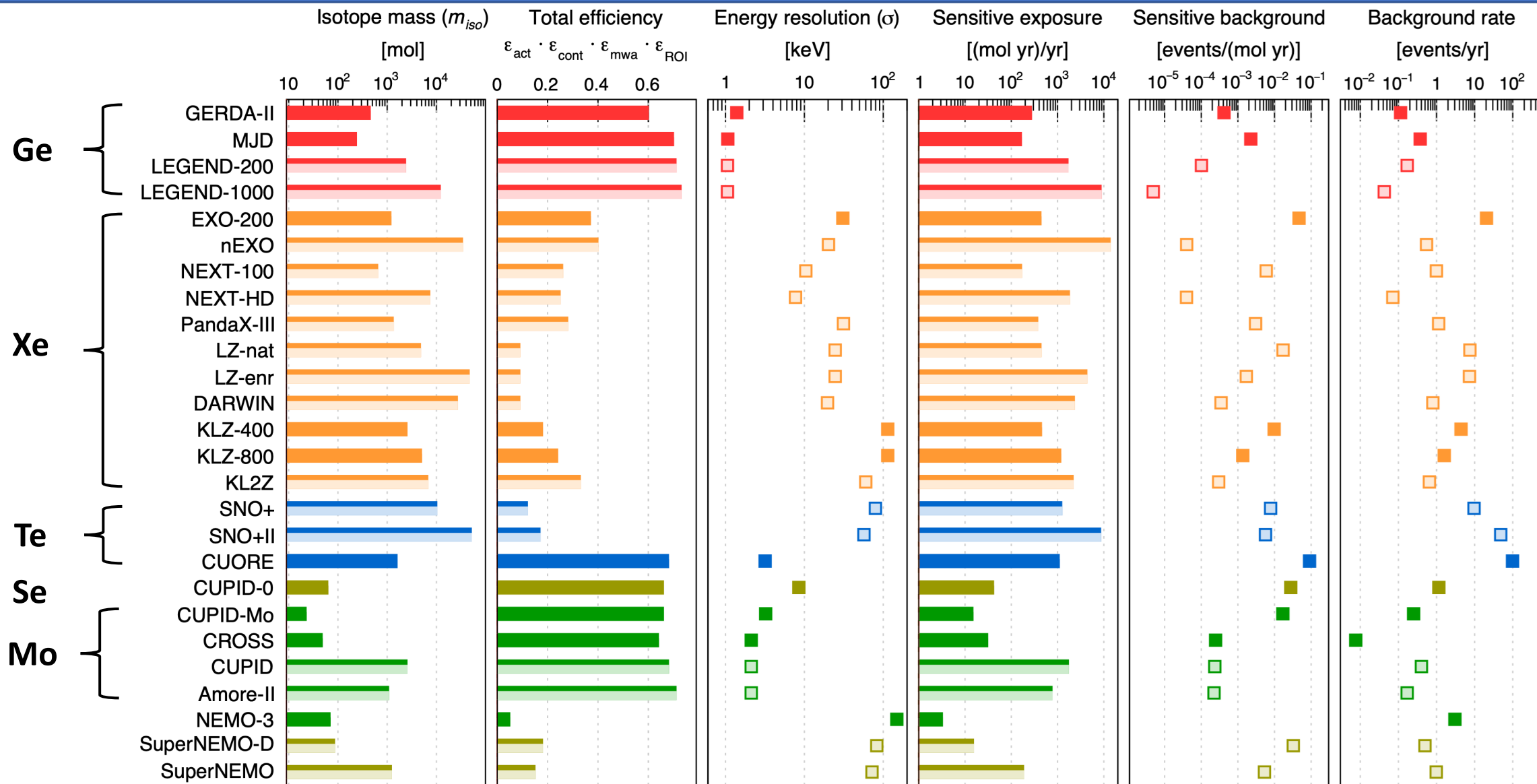
Crystal arrays



Present and Future experiments

	Experiment	Isotope	Status	Laboratory	Moles of isotope	Isotopic abundance	B (events/kg _{is} o/keV/yr)	Fiducial isotope mass (kg)	Active fraction (%)	Efficiency	FWHM (keV)
High-purity Ge detectors	GERDA-II	⁷⁶ Ge	completed	LNGS	450	0.87	6.00E-04	30	88	0.7189	3.29
	MJD	⁷⁶ Ge	completed	SURF	240	0.88	4.74E-03	16	90	0.8099	2.585
	LEGEND-200	⁷⁶ Ge	construction		2400	0.9	2.10E-04	166	91	0.819	2.585
	LEGEND-1000	⁷⁶ Ge	proposed		12000	0.91	1.00E-05	839	92	0.828	2.585
Xenon TPC's	EXO-200	¹³⁶ Xe	completed	WIPP	1200	0.81	1.99E-03	75	46	0.84	72.85
	nEXO	¹³⁶ Xe	proposed	SNOLAB	34000	0.9	2.10E-06	2959	64	0.66	47
	NEXT-100	¹³⁶ Xe	construction	LSC	640	0.9	4.90E-04	77	88	0.3724	23.5
	NEXT-HD	¹³⁶ Xe	proposed		7400	0.9	5.80E-06	956	95	0.3916	18.095
	PandaX-III-200	¹³⁶ Xe	construction	CJPL	1300	0.9	1.20E-04	136	77	0.481	72.85
	LZ-nat	¹³⁶ Xe	construction	SURF	4700	0.09	1.22E-03	90	14	0.8	58.75
	LZ-enr	¹³⁶ Xe	proposed	SURF	46000	0.9	1.20E-04	876	14	0.8	58.75
	Darwin	¹³⁶ Xe	proposed		27000	0.09	4.10E-05	477	13	0.9	47
	Large liquid scintillators	KLZ-400	¹³⁶ Xe	completed	Kamioka	2500	0.91	3.80E-04	150	44	0.97
KLZ-800		¹³⁶ Xe	data taking	Kamioka	5000	0.91	5.30E-05	394	58	0.97	267.9
KL2Z		¹³⁶ Xe	proposed	Kamioka	6700	0.91	2.20E-05	729	80	0.97	141
SNO-I		¹³⁰ Te	construction	SNOLAB	10000	0.348	3.00E-04	272	20	0.97	188
SNO-II		¹³⁰ Te	proposed	SNOLAB	51000	0.348	3.00E-04	1873	27	0.97	133.95
Cryogenic calorimeters	CUORE	¹³⁰ Te	data taking	LNGS	1585	0.348	5.38E-02	206	100	0.8096	7.52
	CUPID-0	⁸² Se	completed	LNGS	62	0.96	5.95E-03	5	100	0.6966	19.975
	CUPID-Mo	¹⁰⁰ Mo	completed	LSM	23	0.97	7.82E-03	2.3	100	0.6916	7.52
	CROSS	¹⁰⁰ Mo	construction	LSC	48	0.96	1.71E-02	5	100	0.675	4.935
	CUPID	¹⁰⁰ Mo	proposed	LNGS	2500	0.96	1.70E-04	250	100	0.711	4.935
	AMORE	¹⁰⁰ Mo	proposed	Yemilab	1100	0.96	1.70E-04	110	100	0.7462	4.935
Tracking calorimeters	NEMO-3	¹⁰⁰ Mo	completed	LSM	690	0.9	2.18E-03	69	100	0.11	347.8
	SuperNEMO-D	⁸² Se	construction	LSM	850	0.9	1.10E-04	70	100	0.28	195.05
	SuperNEMO	⁸² Se	proposed	LSM	1200	0.9	3.20E-05	98	100	0.28	169.2

Present and Future Experiments



Matteo Agostini, Giovanni Benato, Jason A. Detwiler, Javier Menéndez, Francesco Vissani - arXiv:2202.01787



E. Previtali

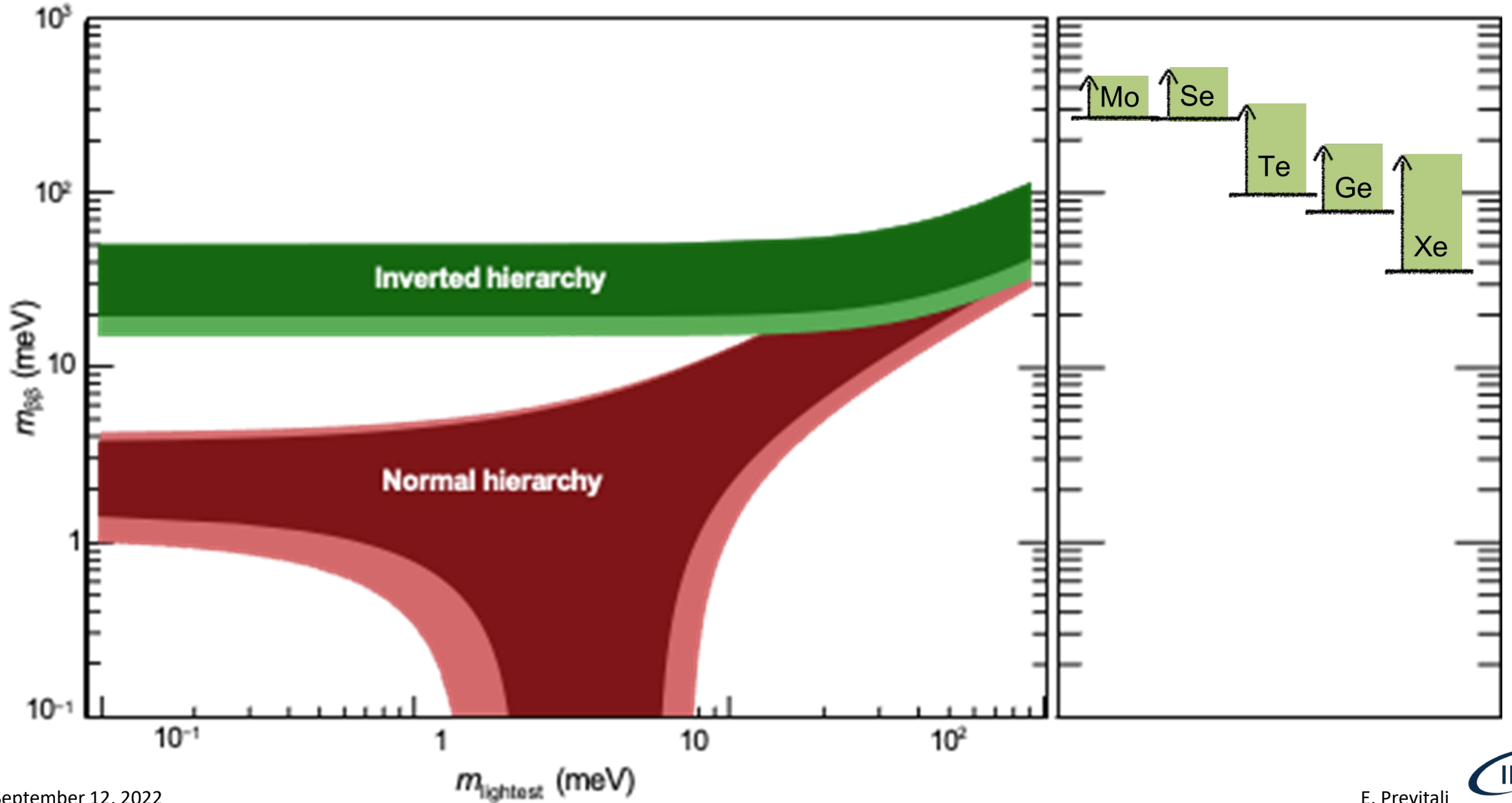
Present Experimental Sensitivities

Isotope	Daughter	Q	f _{nat}	f _{enr}	T(2)	T(0)
⁴⁸ Ca	⁴⁸ Ti	4267.98(32)	0.187(21)	16	(6.4 ^{+0.7} _{-0.6} (stat) ^{+1.2} _{-0.9} (syst)) 10 ¹⁹	> 5.8 10 ²²
⁷⁶ Ge	⁷⁶ Se	2039.061(7)	7.75(12)	92	(1.92±94) 10 ²¹	> 1.8 10 ²⁶
⁸² Se	⁸² Kr	2997.9(3)	8.82(15)	96.3	(8.60±0.03 (stat) ^{+0.19} _{-0.13} (syst)) 10 ¹⁹	> 3.5 10 ²⁴
⁹⁶ Zr	⁹⁶ Mo	3356.097(86)	2.80(2)	86	(2.35±0.14 (stat) ± 0.16 (syst))10 ¹⁹	> 9.2 10 ²¹
¹⁰⁰ Mo	¹⁰⁰ Ru	3034.40(17)	9.744(65)	99.5	(7.12 ^{+0.18} _{-0.14} (stat) ± 0.10 (syst)) 10 ¹⁸	> 1.5 10 ²⁴
¹¹⁶ Cd	¹¹⁶ Sn	2813.50(13)	7.512(54)	82	2.63 ^{+0.11} _{-0.12} 10 ¹⁹	> 2.2 10 ²³
¹³⁰ Te	¹³⁰ Xe	2527.518(13)	34.08(62)	92	(7.71 ^{+0.08} _{-0.06} (stat) ^{+0.12} _{-0.15} (syst)) 10 ²⁰	> 2.2 10 ²⁵
¹³⁶ Xe	¹³⁶ Ba	2457.83(37)	8.857(72)	90	(2.165 ±0.016 (stat) ±0.059 (syst))10 ²¹	> 1.1 10 ²⁶
¹⁵⁰ Nd	¹⁵⁰ Sm	3371.38(20)	5.638(28)	91	(9.34±0.22(stat) ^{+0.62} _{-0.60} (syst)) 10 ¹⁸	> 2.0 10 ²²

NEUTRINO 2022
updates

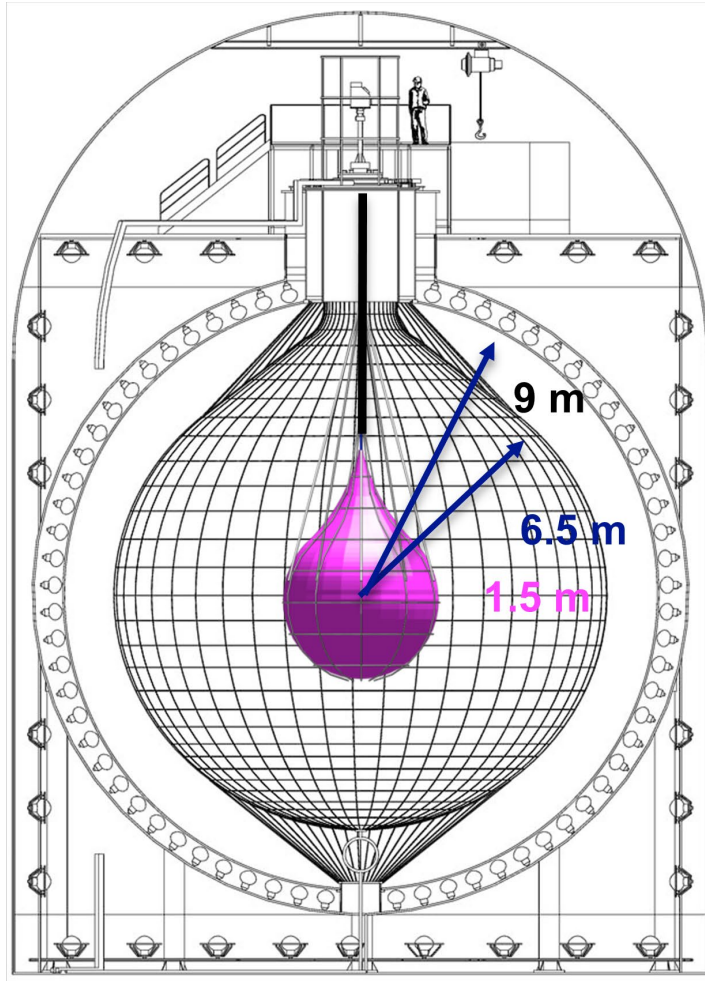
Experiment	Isotope	Limit T _{1/2} (yr)	BI (10 ⁻³ ckk _y)	BI (10 ⁻³ ckk _{iso} yr)	FWHM (keV)	Exposure (kg _{iso} ·yr)	m _{ββ} (meV)	
AMORE-I	¹⁰⁰ Mo	1.2·10 ²⁴	39	72.8	10	1.67	343-600	Y. Oh
GERDA	⁷⁶ Ge	1.8·10 ²⁶	0.5	6.8	3	103.7	79-180	J. Gruszko
MJD	⁷⁶ Ge	8.3·10 ²⁵	6	6.8	2.85*	64.5	113-269	J. Gruszko
CUORE	¹³⁰ Te	2.2·10 ²⁵	14.9	52.7	7.8	288.6	90-305	I.Nutini
CUPID-0	⁸² Se	4.7·10 ²⁴	3.5	5.4	20	8.8	276-570	A. Zolotarova
CUPID-Mo	¹⁰⁰ Mo	1.8·10 ²⁴	3.0	5.6	7.4	2.7	280-490	A.Zolotarova
KL-ZEN-800	¹³⁶ Xe	2.3·10 ²⁶ *		0.06	268	605	36-156	A. Nando

Sensitivities on Light Neutrino Majorana Mass



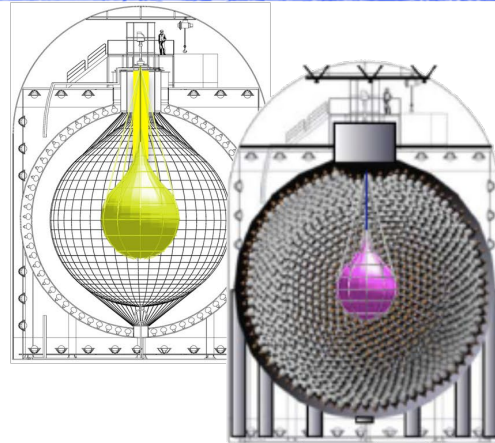
KamLAND-Zen

Large Liquid Scintillator Detector



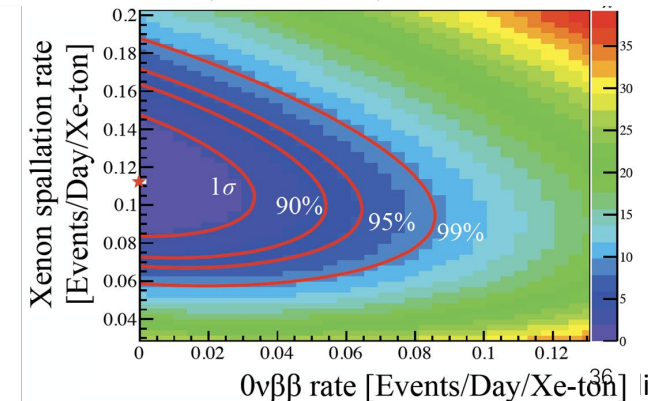
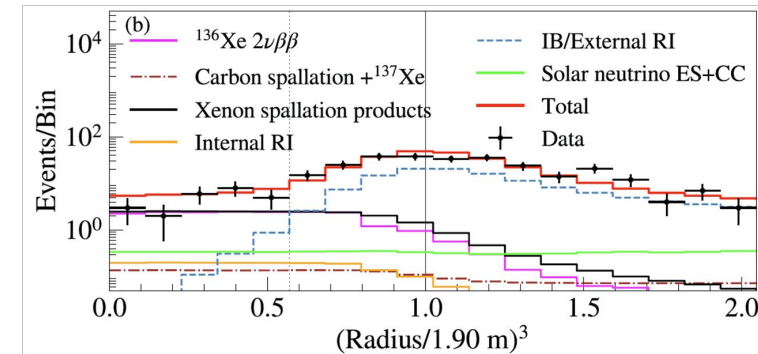
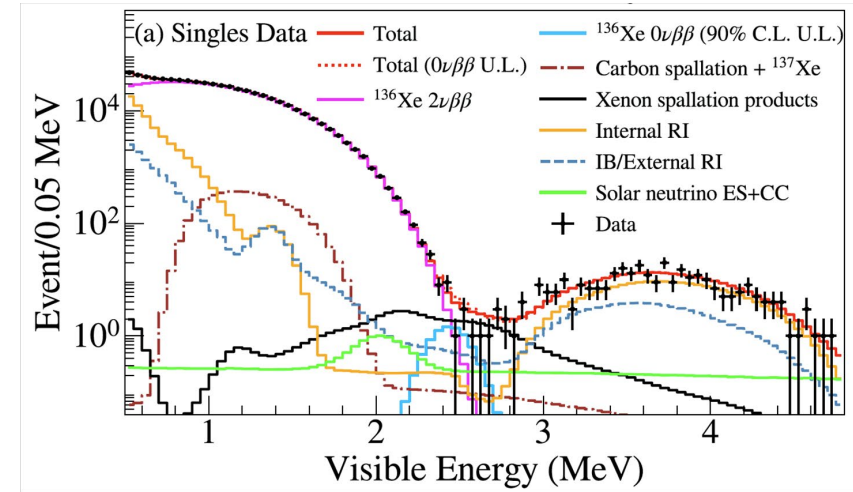
- Sensitivity:
 $> 5.6 \cdot 10^{25}$ yr (90% C.L.)
- Fit:
 $> 9.2 \cdot 10^{25}$ yr (90% C.L.)
- Phase I + II:
 $> 1.07 \cdot 10^{26}$ yr (90% C.L.)

• KL-Zen-800:
 data taking started in
January 2019 with 750 kg
 $enrXe$ (new balloon)



Phase-2: 2013/12/11 - 2015/10/27: **534.5 days (504 kg-yr)**

- KamLAND2-Zen with 1000 kg (proposed)



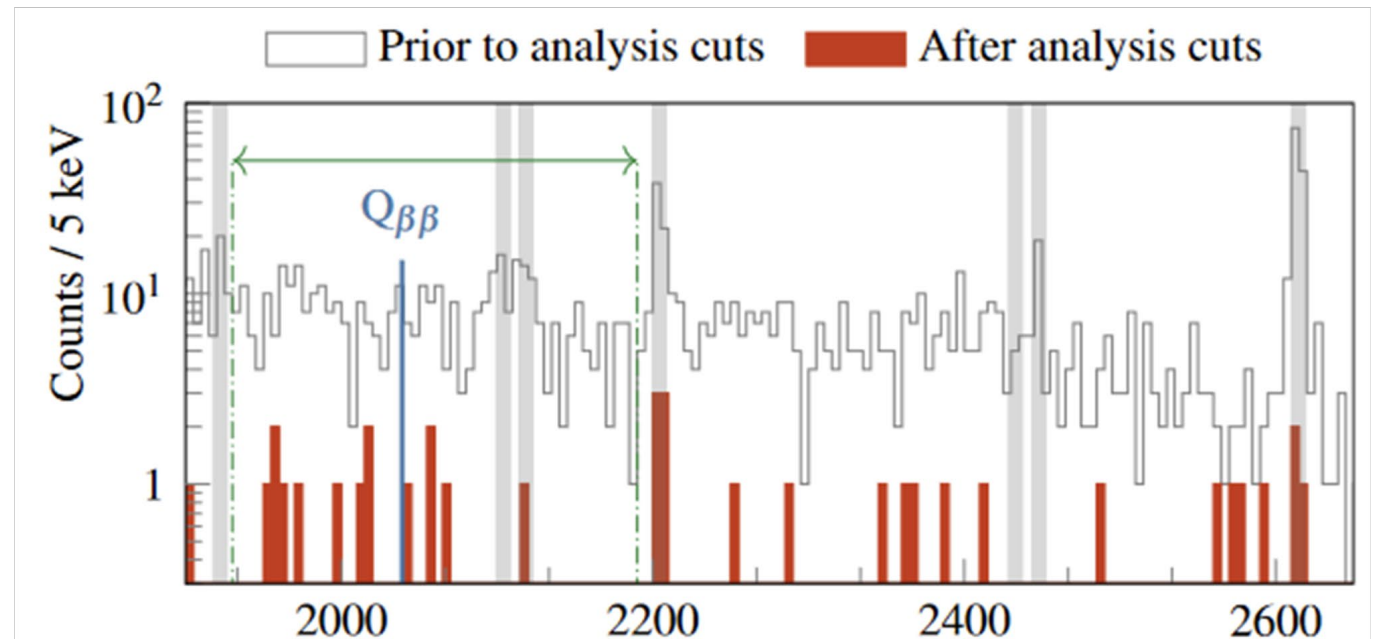
GERDA

High purity germanium diodes immersed in LAr

- $\Delta E < 3$ Kev FWHM @ $Q_{\beta\beta}$
- Pulse shape analysis: multi/single site vs.
- anticoincidence with LAr
 - scintillating fibers (WLS) coupled to SIPMS
 - PMT above and below the detector



- 37 germanium HP detectors
- 35 kg of ^{76}Ge
- LNGS: Hall A
- Exceptional background level
 $5.2_{-1.3}^{+1.6} \times 10^{-4}$ cont/(keV kg yr)
- Best limit on average lifetime
 $T_{1/2} > 1.8 \times 10^{26}$ yr
 $m_{\beta\beta} < 79\text{-}180$ meV

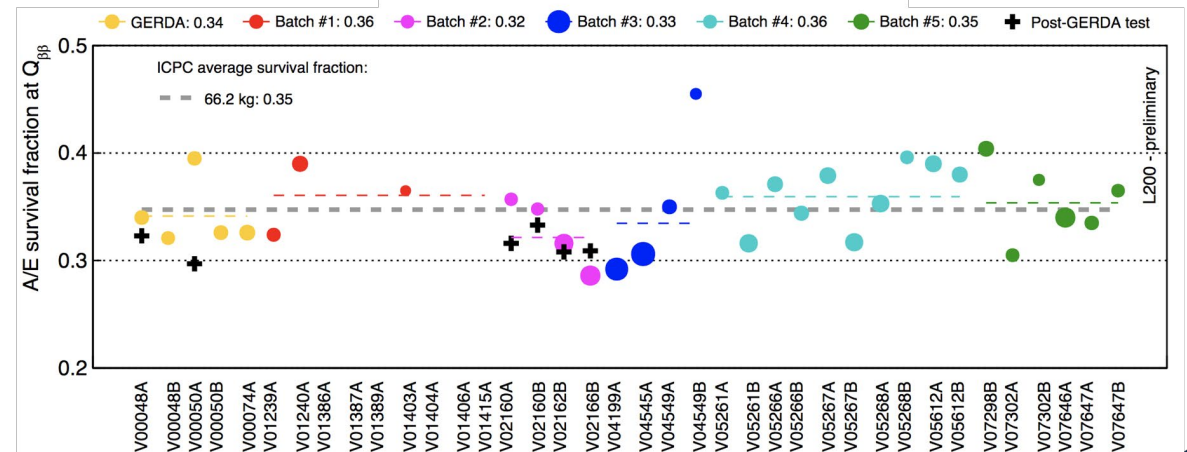
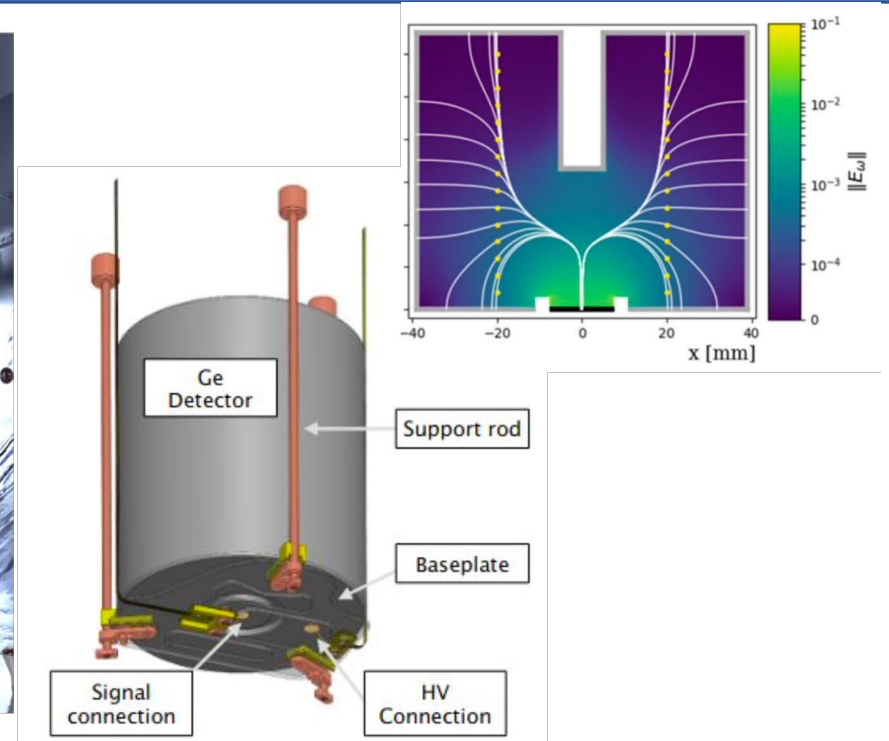
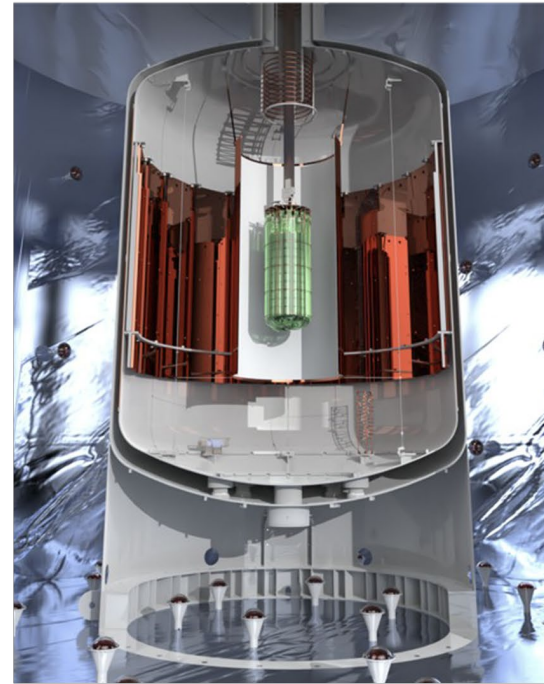


Phys. Rev. Lett. 125, 252502 (2020)

LEGEND-200

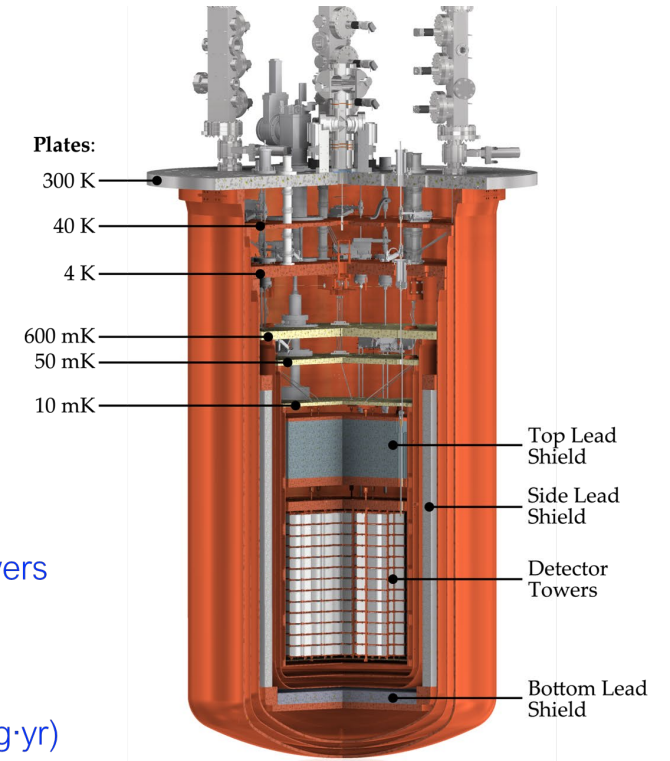
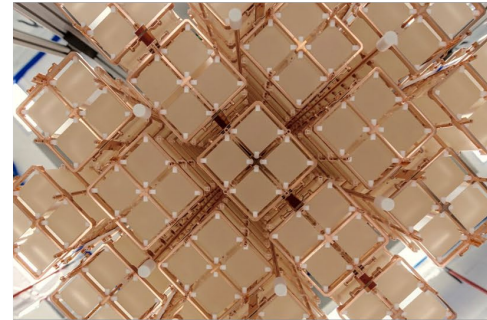
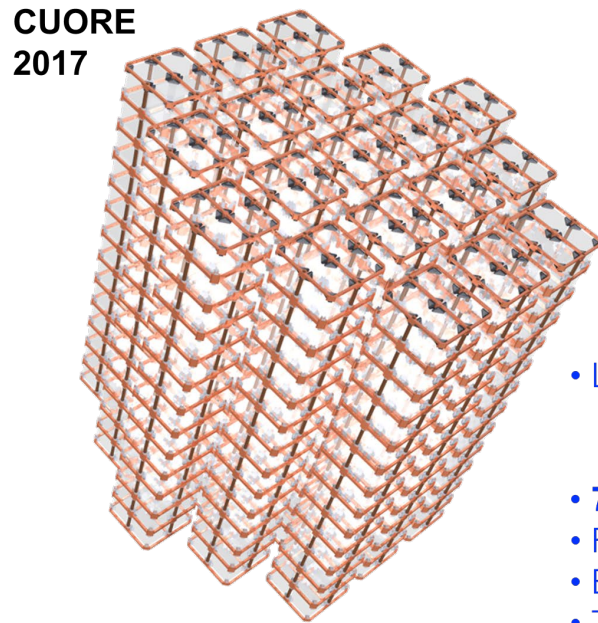
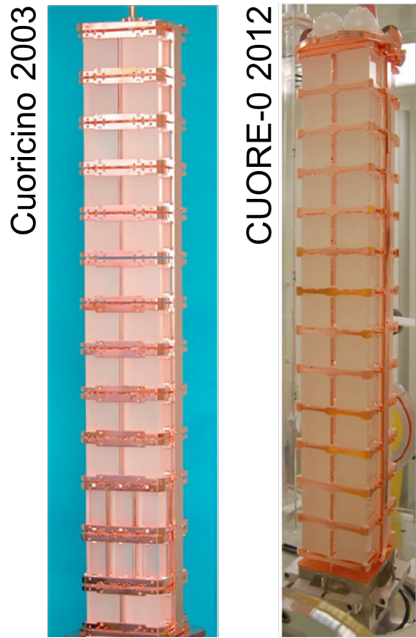
- Natural evolution of the GERDA principle
- Combines the best of Gerda and MJD
 - ▶ from GERDA:
 - detector configuration
 - infrastructure@ LNGS
 - system improvements
 - ▶ from MJD
 - selection of radio-pure materials
 - electronics
 - low threshold¹
- ⁷⁶Ge:
 - ▶ 35 kg from GERDA
 - ▶ 30 kg from MJD
 - ▶ 140 kg new materials
- New type of detector, already tested in GERDA
 - ▶ ICPC m > 2 kg (0.7-0.9 kg previously)
 - ▶ same energy resolution and PSD capability

- Infrastructure upgrade: February 2020
- Detector commissioning: September 2021
- Data taking: start in 2022

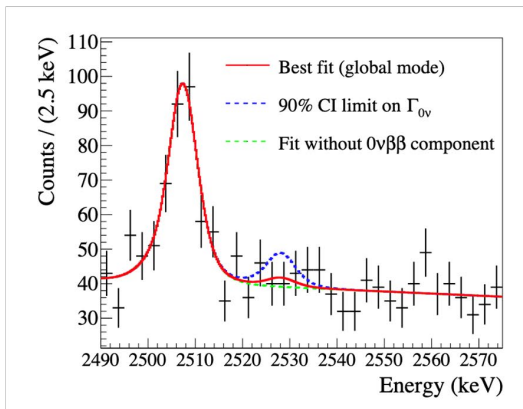


CUORE

Largest cryogenic particle detector ever realized



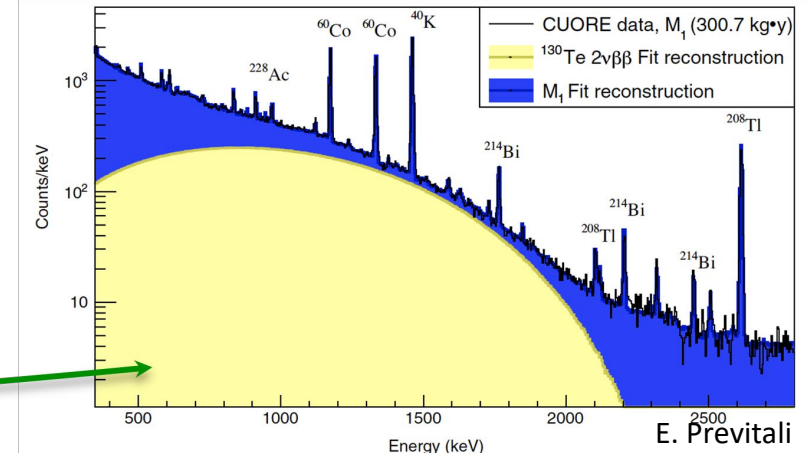
- LNGS
 - 988 TeO₂ crystals arranged on 19 towers
 - dedicated "dry" cryostat
- **742 kg** di natural TeO₂ (206 kg of ¹³⁰Te)
- FWHM: **7.7 keV** FWHM (ROI)
- Background: **(1.4 ± 0.2) · 10⁻²** cnts/(keV·kg·yr)
- Total exposure: 1.8 tonne (TeO₂)·yr



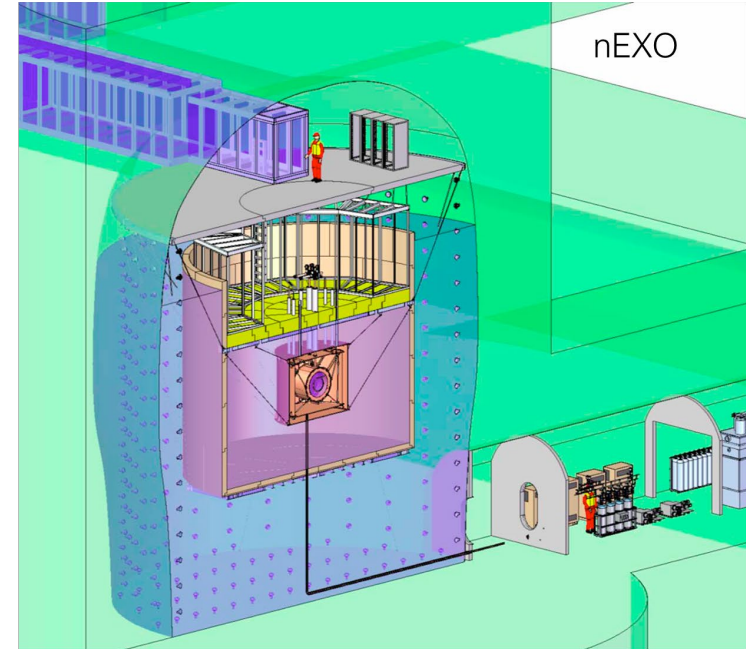
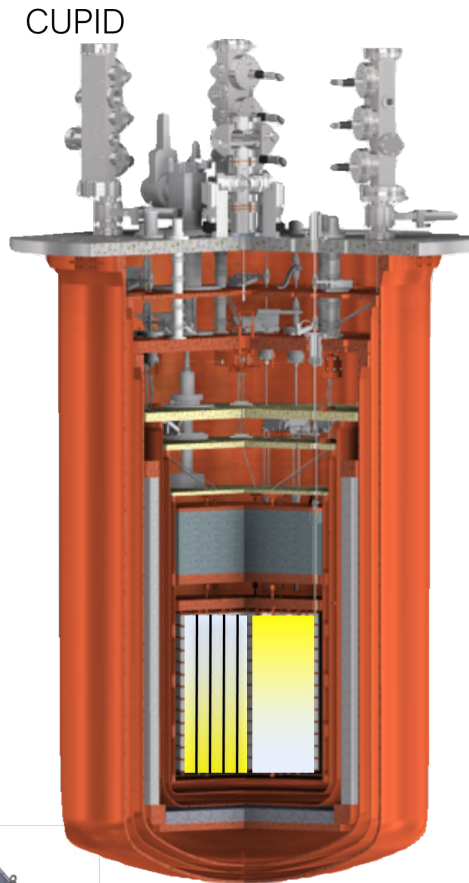
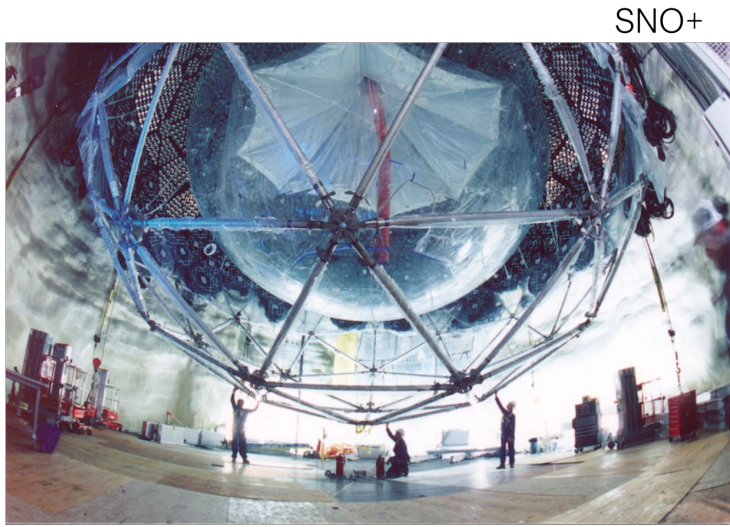
- $\beta\beta 0\nu$ result (1038.4 kg yr TeO₂, 288 kg yr ¹³⁰Te): :
 - $T_{1/2}(0\nu) > 2.2 \cdot 10^{25}$ yr (90% C.L.)
 - $m_{\beta\beta} < 90-305$ meV

Adams D. et al. (CUORE collaboration), *Nature* 604 (2022) 7904, 53-58
<https://www.nature.com/articles/s41586-022-04497-4>

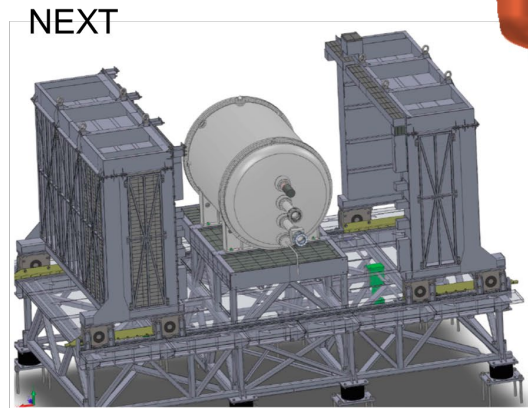
- $T_{1/2}(2\nu) = [7.71^{(+0.08}_{-0.06(\text{stat.})} + 0.12_{-0.15(\text{syst.})}] \cdot 10^{20}$ yr



Next Generation Experiments



KamLAND-ZEN



LEGEND

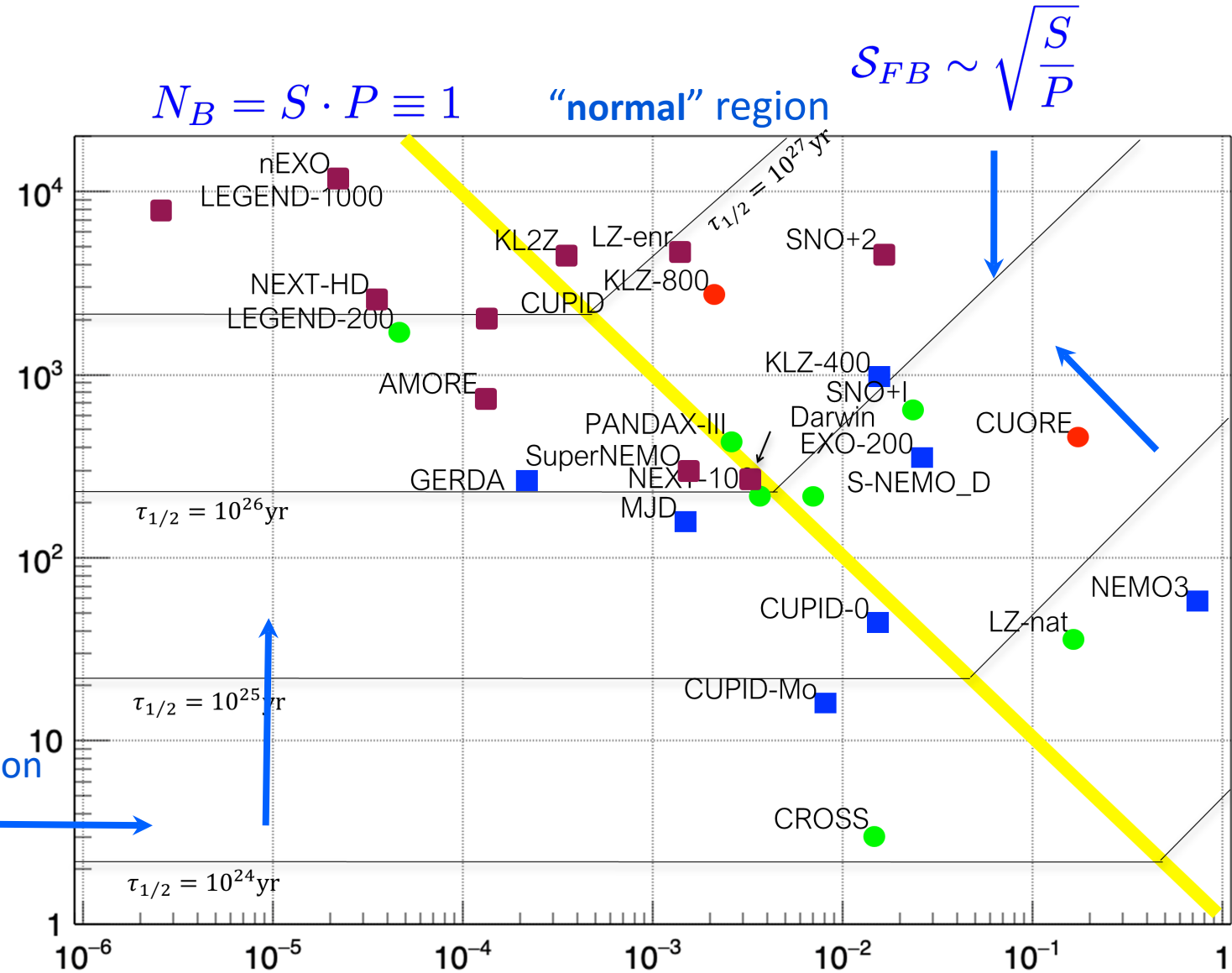


Sensitivity Plot: Present and Future

$$N_B = S \cdot P \equiv 1$$

$$S_{FB} \sim \sqrt{\frac{S}{P}}$$

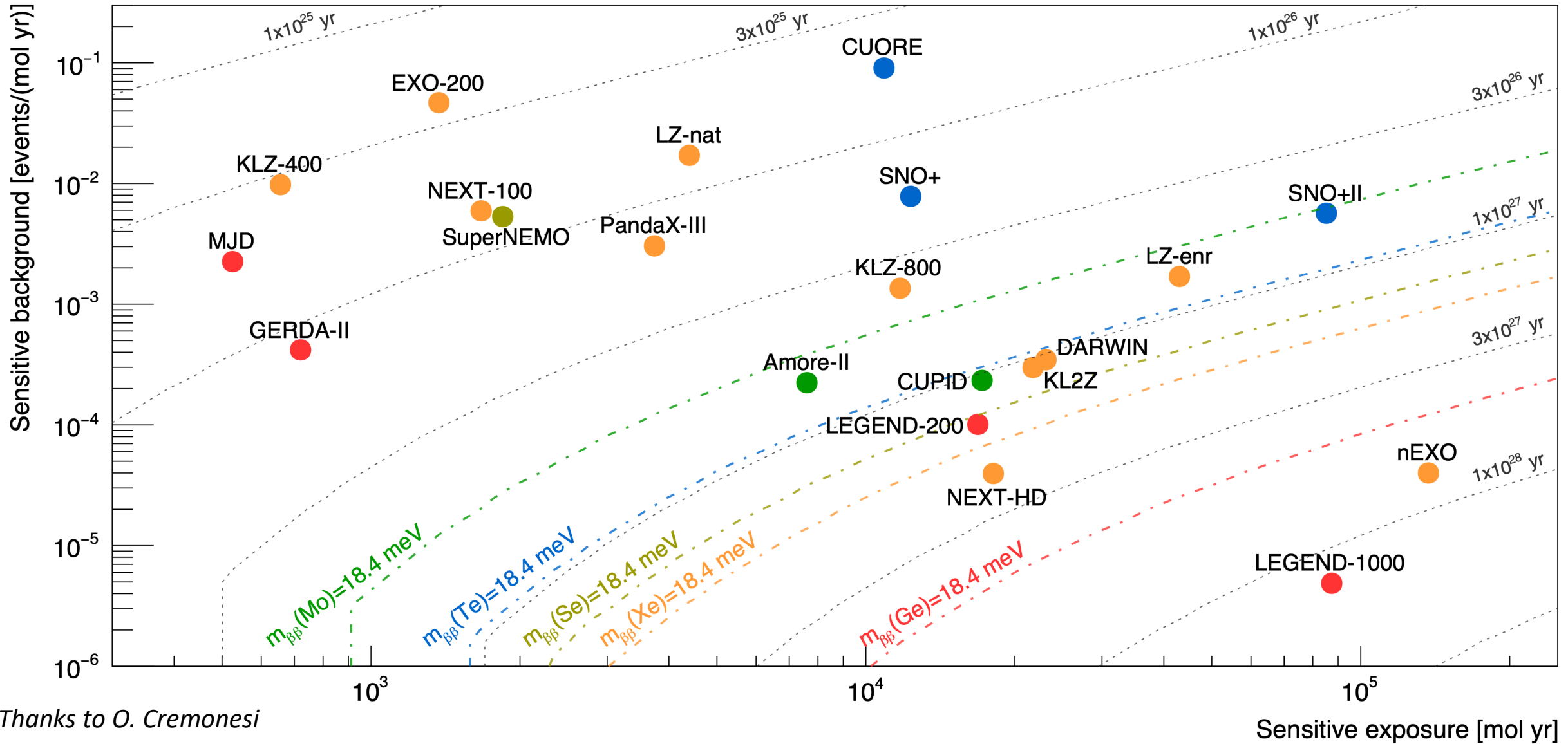
$$S_{ZB} \sim S$$



Bianconi e Cremonesi, arXiv:1310.3870 - Prog. Part. Nucl. Phys. 114 (2020) 103803



Sensitivities: a different view

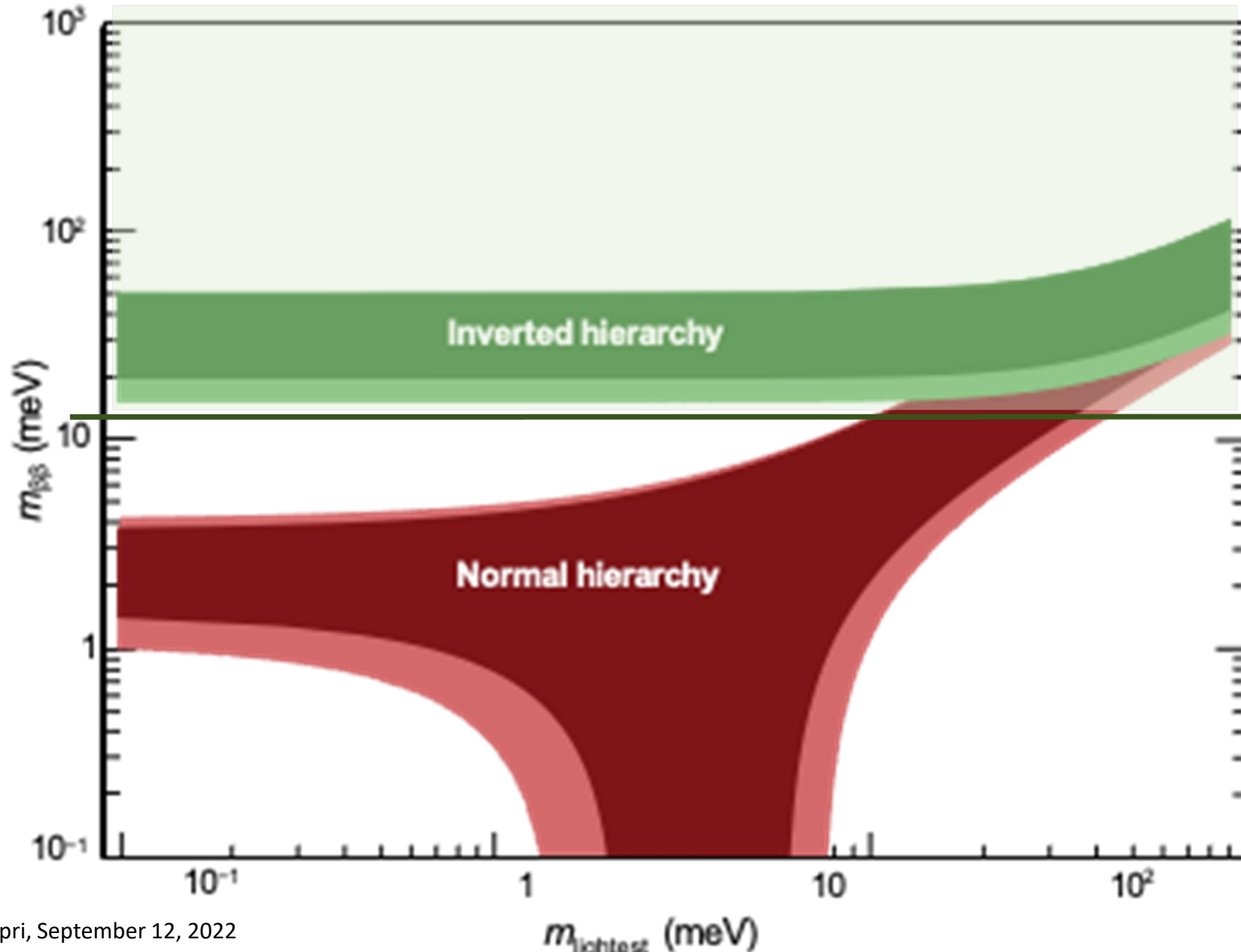


Thanks to O. Cremonesi

Sensitive exposure [mol yr]



Future Sensitivity



CUPID, LEGEND, KamLAND-ZEN, nEXO, ...

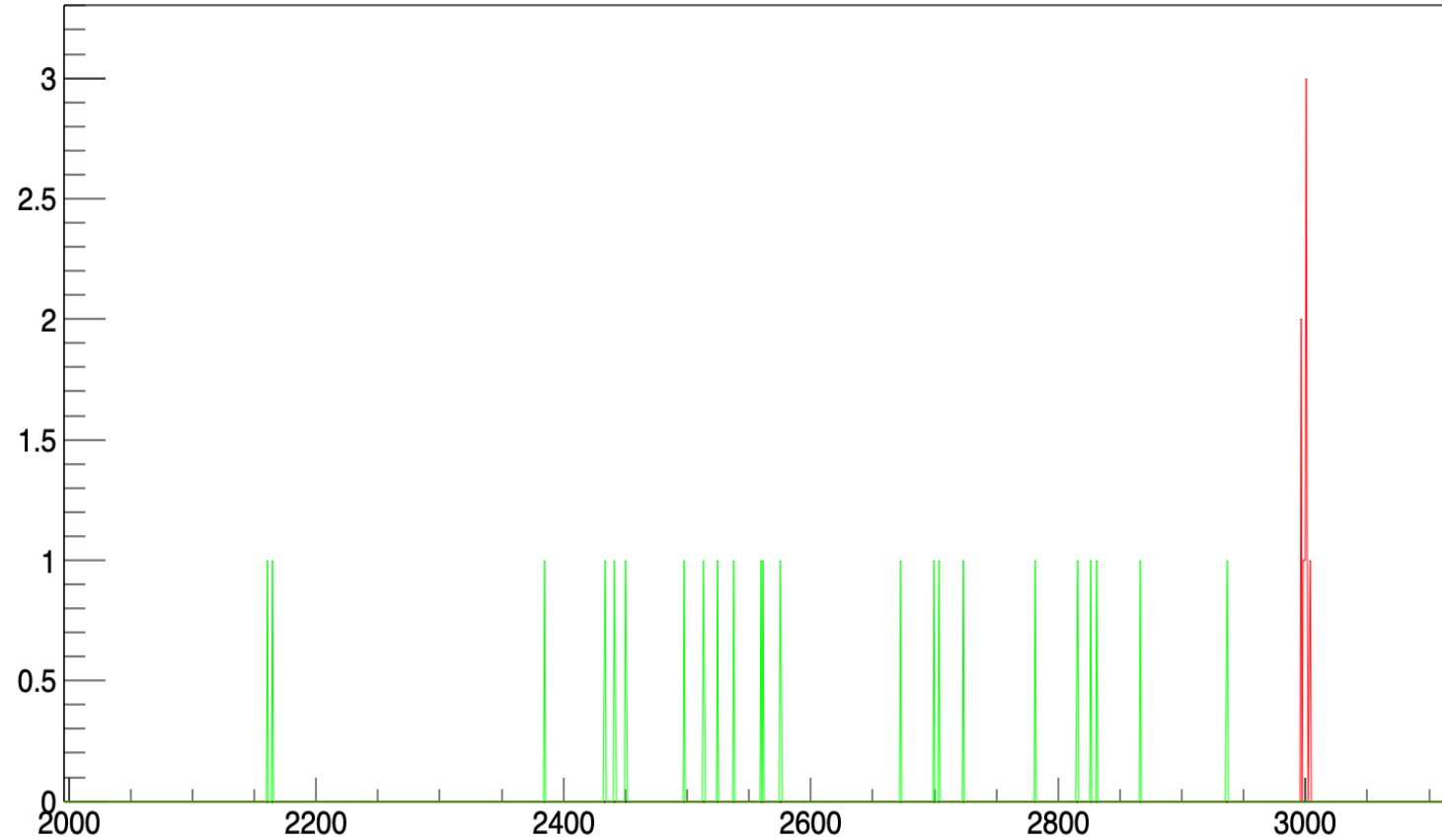
$$\tau_{0\nu} > 10^{27} - 10^{28} \text{ y}$$

Discovery Potential

If will we see a peak @ $\beta\beta$ decay energy?

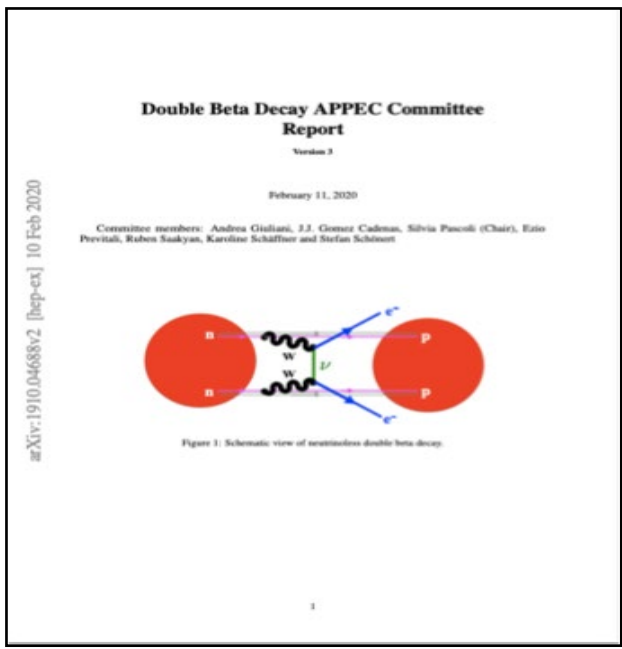
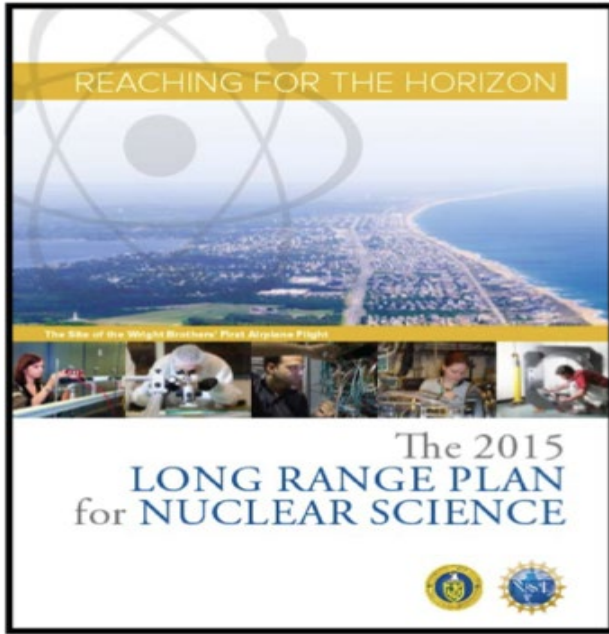
There are some important aspects:

- Few isotopes are needed
- High energy resolution detector crucial
- Nuclear matrix element determined
- Clear evaluation of possible interferences
-



In that sense a discovery potential sensitivity for each experiment is important

Complicated Strategy for the Future



DOE NP Portfolio Review
 July 2021
 CUPID
 LEGEND-1000
 nEXO



<https://science.osti.gov/np/nsac>

“We recommend the timely development and deployment of a U.S.-led ton-scale neutrinoless double beta decay experiment.”

<https://arxiv.org/abs/1910.04688>

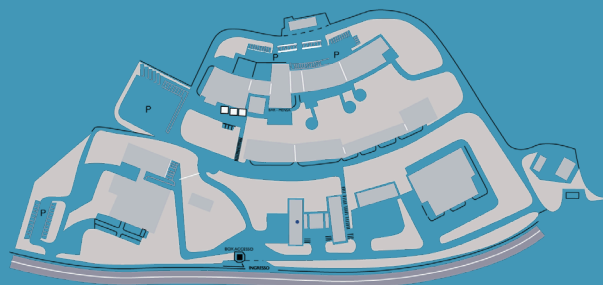
- Oct 2019: Roadmap document for the APPEC SAC on the future $0\nu\beta\beta$ decay experimental program in Europe
- $0\nu\beta\beta$ town meeting London
- Roadmap update 2022, town meeting in Berlin, June 2022

- Outcome: Realize international portfolio **LEGEND-1000, nEXO and CUPID with European partners**
- Open discussion in september 2021 @ LNGS.

“The international stakeholders in neutrino-less double beta decay research do agree in principle that the best chance for success is an international campaign **with more than one large ton-scale experiment** implemented in the next decade, **with one ton scale experiment in Europe and the other in North America.** “



INFN
LNGS



Thank you

