

# *Direct neutrino mass measurements -*



Christoph Wiesinger (TUM, MPP), FPCapri, 12.06.2022

*Neutrino Mass | KATRIN | Project 8 | ECHo / HOLMES | Outlook*

“for the discovery of neutrino oscillations, which shows that

# Neutrinos have mass”

[Kajita, McDonald, Nobel Prize in Physics 2015]

- flavor eigenstates are **linear combinations** of mass eigenstates

$$\nu_l = \sum_i U_{li} \nu_i$$

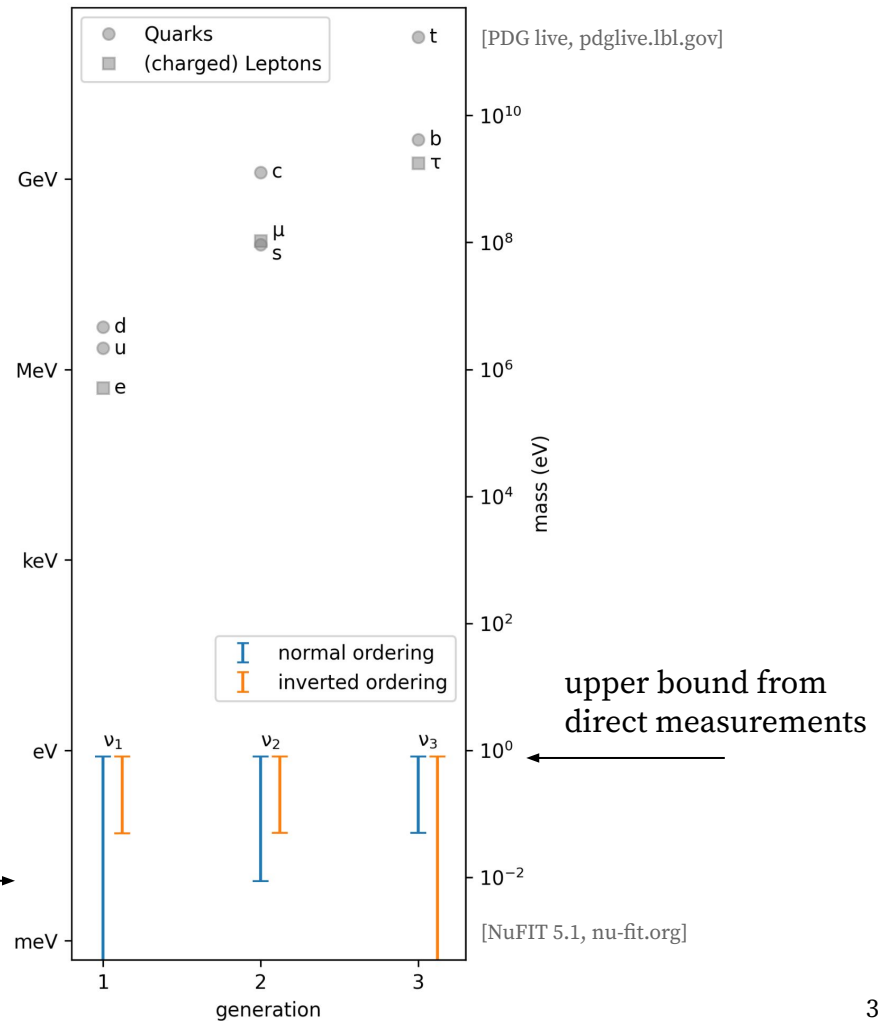
- mass squared differences

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

## → neutrino oscillations

- mass mechanism, **absolute mass** and mass ordering **unknown**

lower bounds from oscillation experiments

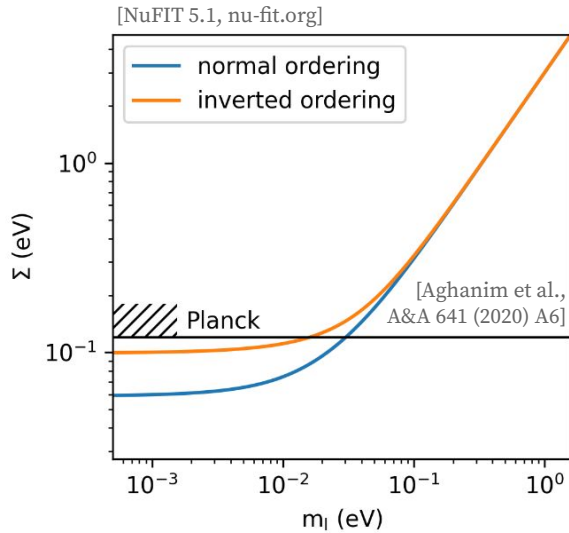


# Mass observables

laboratory-based

**cosmology**

$$\Sigma = \sum_i m_i$$



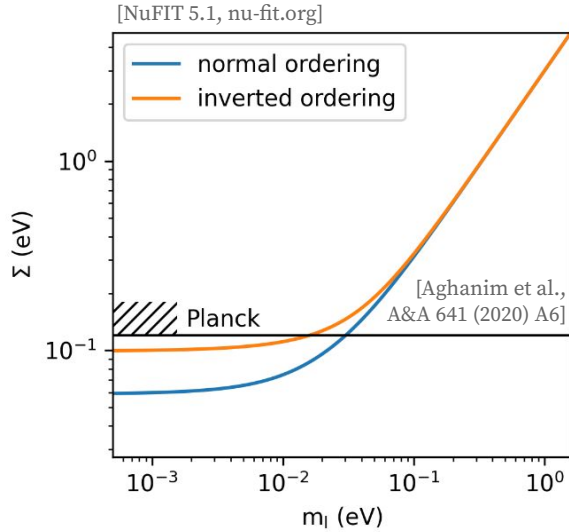
model-dependent

# Mass observables

laboratory-based

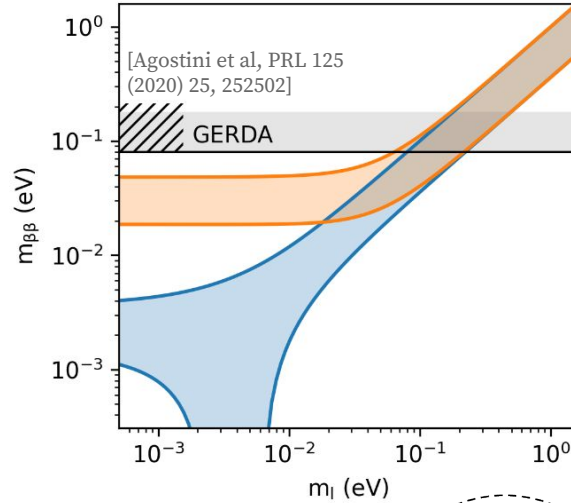
**cosmology**

$$\Sigma = \sum_i m_i$$



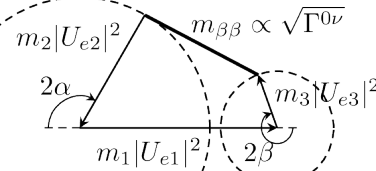
**neutrinoless  $\beta\beta$ -decay**

$$m_{\beta\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$



model-dependent

[Bilenky et al., PRD 64 (2001) 053010]

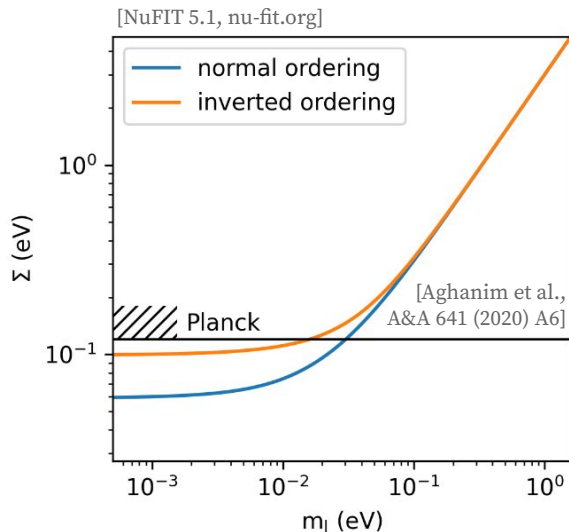


# Mass observables

laboratory-based

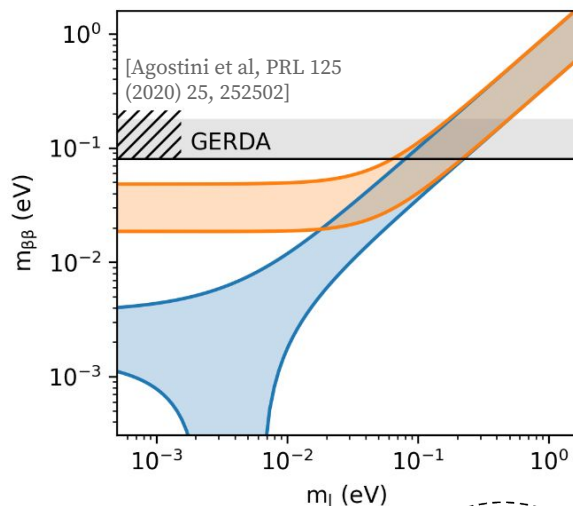
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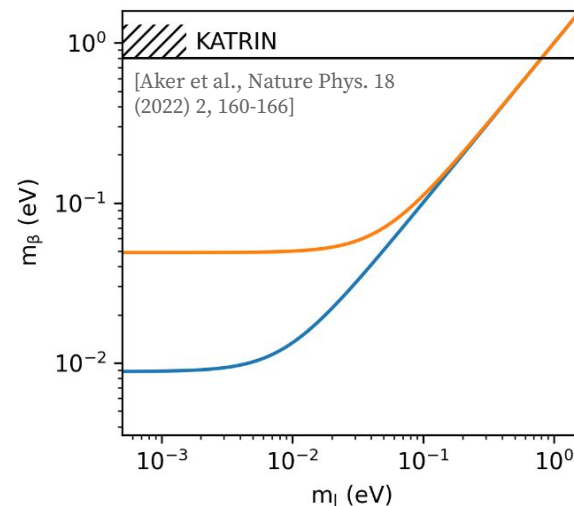
**neutrinoless  $\beta\beta$ -decay**

$$m_{\beta\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$



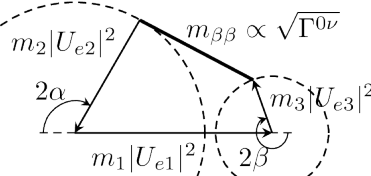
**$\beta$ -decay kinematics**

$$m_\beta = \sqrt{\sum_i |U_{ei}^2| m_i^2}$$



model-dependent

[Bilenky et al., PRD 64 (2001) 053010]



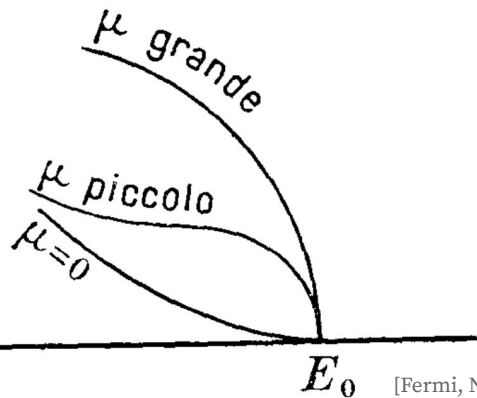
# $\beta$ -decay kinematics

- **spectral distortion**, maximal at **endpoint** energy  $E_0$

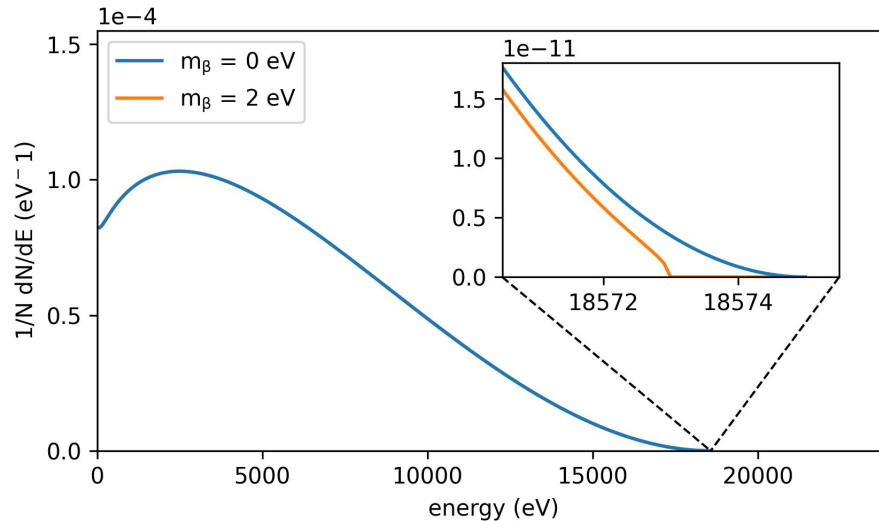
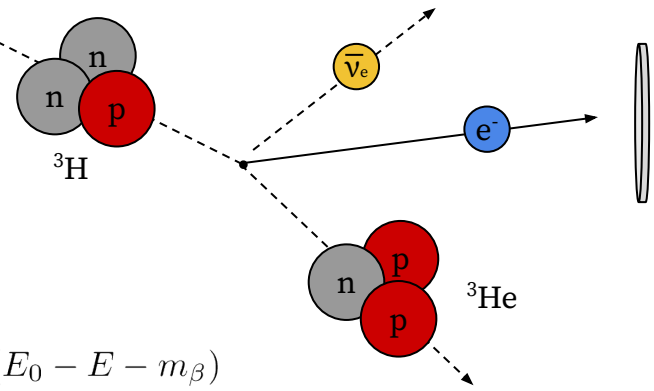
$$\frac{d\Gamma}{dE} \propto F(E, Z) \cdot p \cdot (E + m_e) \cdot (E_0 - E) \cdot \sqrt{(E_0 - E)^2 - m_\beta^2} \cdot \theta(E_0 - E - m_\beta)$$

- purely based on **kinematics** and **energy conservation**

→ independent on neutrino nature

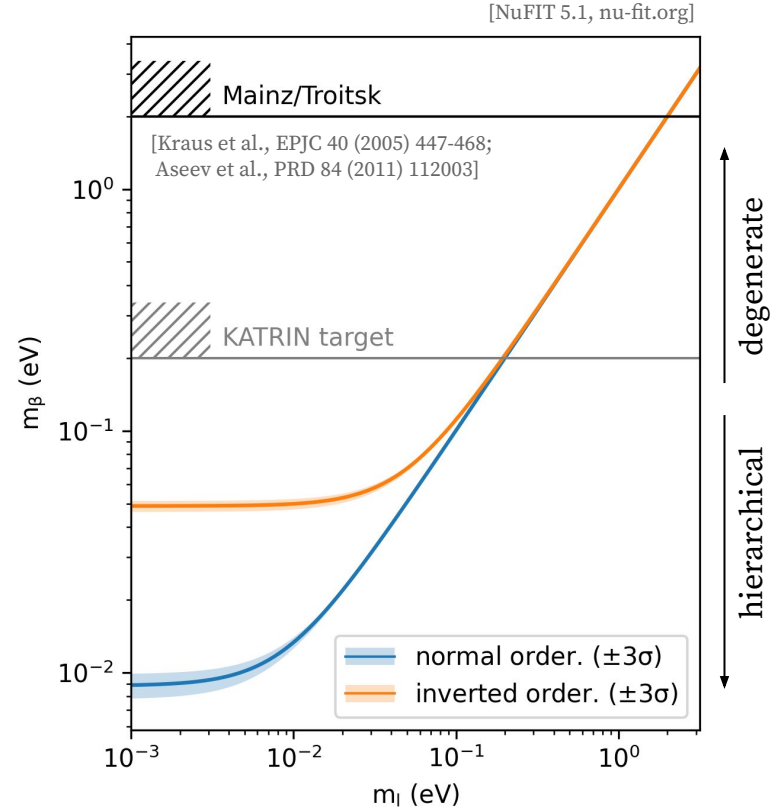


[Fermi, Nuovo Cim. 11 (1934) 1-19]

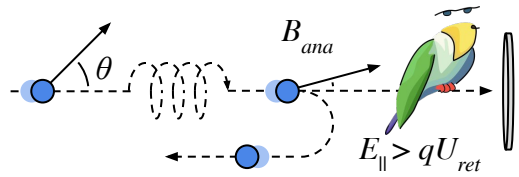


# Experimental challenges

- **high activity** radioactive source, **low Q-value**
- **tritium**  ${}^3\text{H}$  ( $T_{1/2} = 12.3$  yr,  $E_0 = 18.6$  keV),  
**holmium**  ${}^{163}\text{Ho}$  ( $T_{1/2} = 4570$  yr,  $E_0 = 2.8$  keV)
- excellent **energy resolution**,  $O(1)$  eV
- low **background**
- **high precision** understanding of theoretical spectrum and experimental response
- **current:** Karlsruhe Tritium Neutrino (**KATRIN**) experiment, probe **degenerate** scale
- **future:** resolve **normal** vs. **inverted** ordering





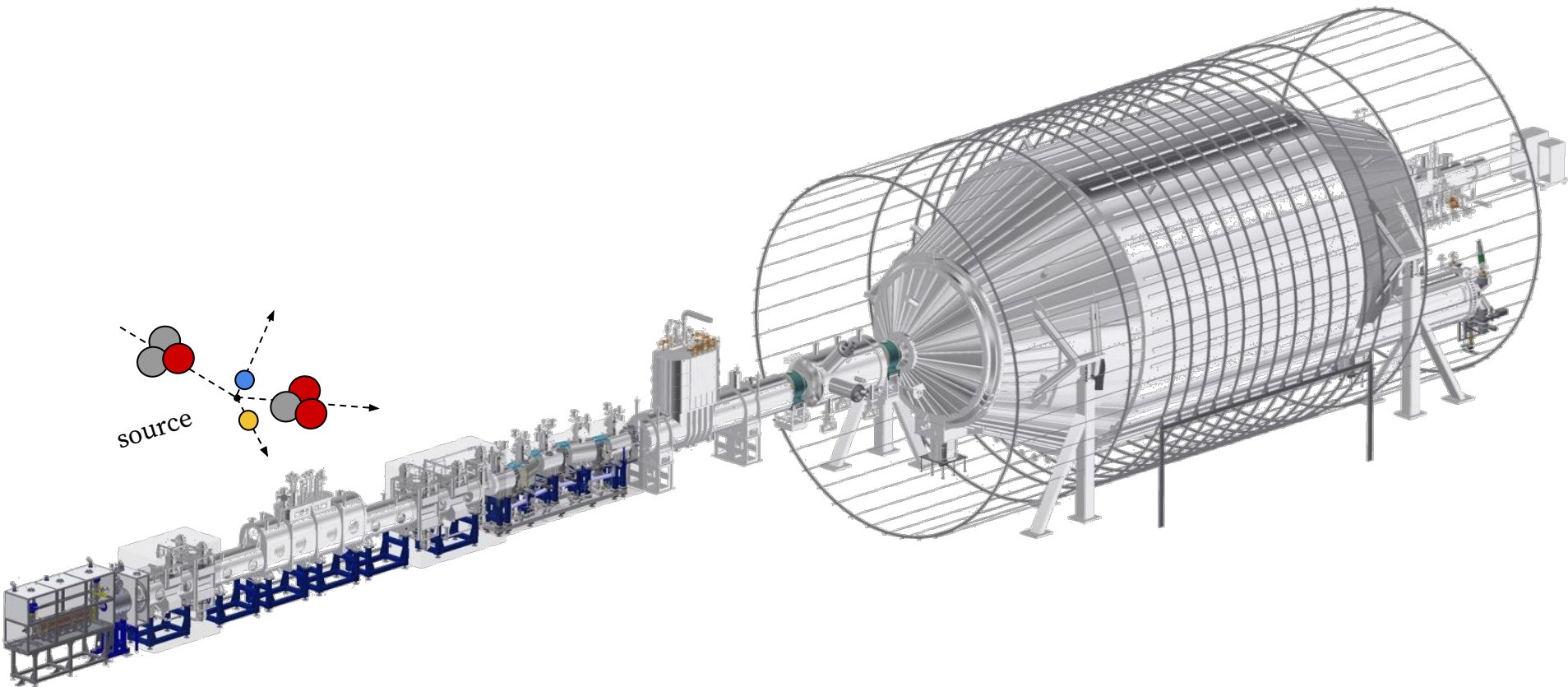


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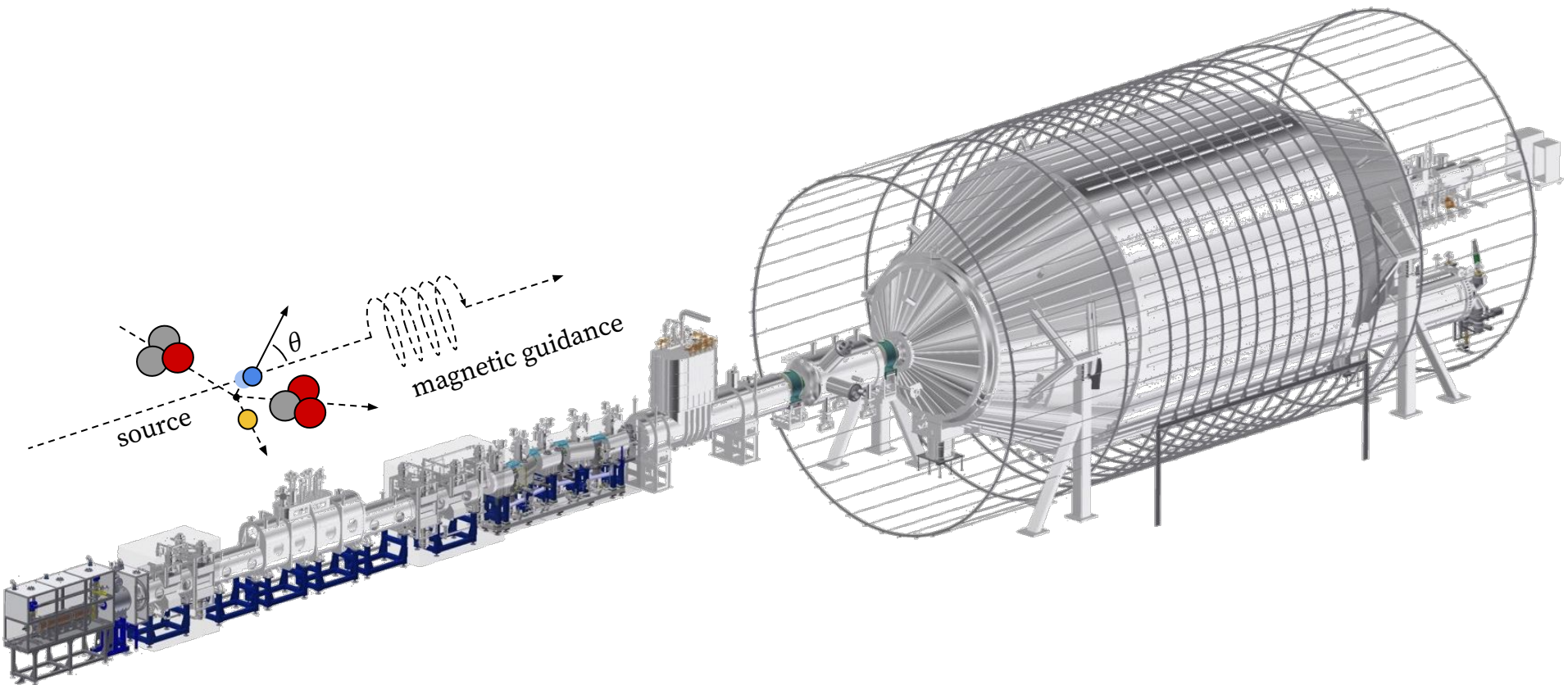
*Enter KATRIN*



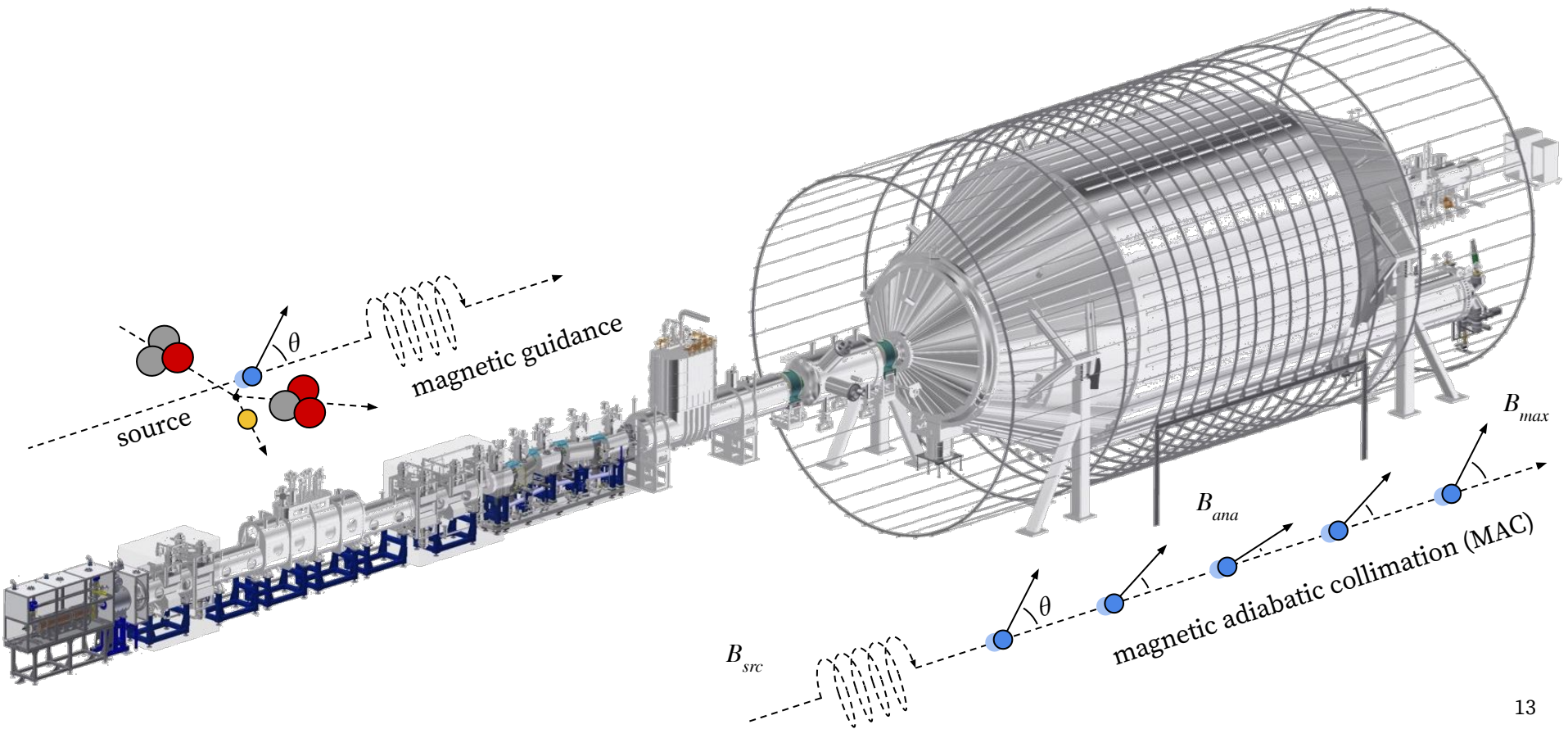
# *Working principle*



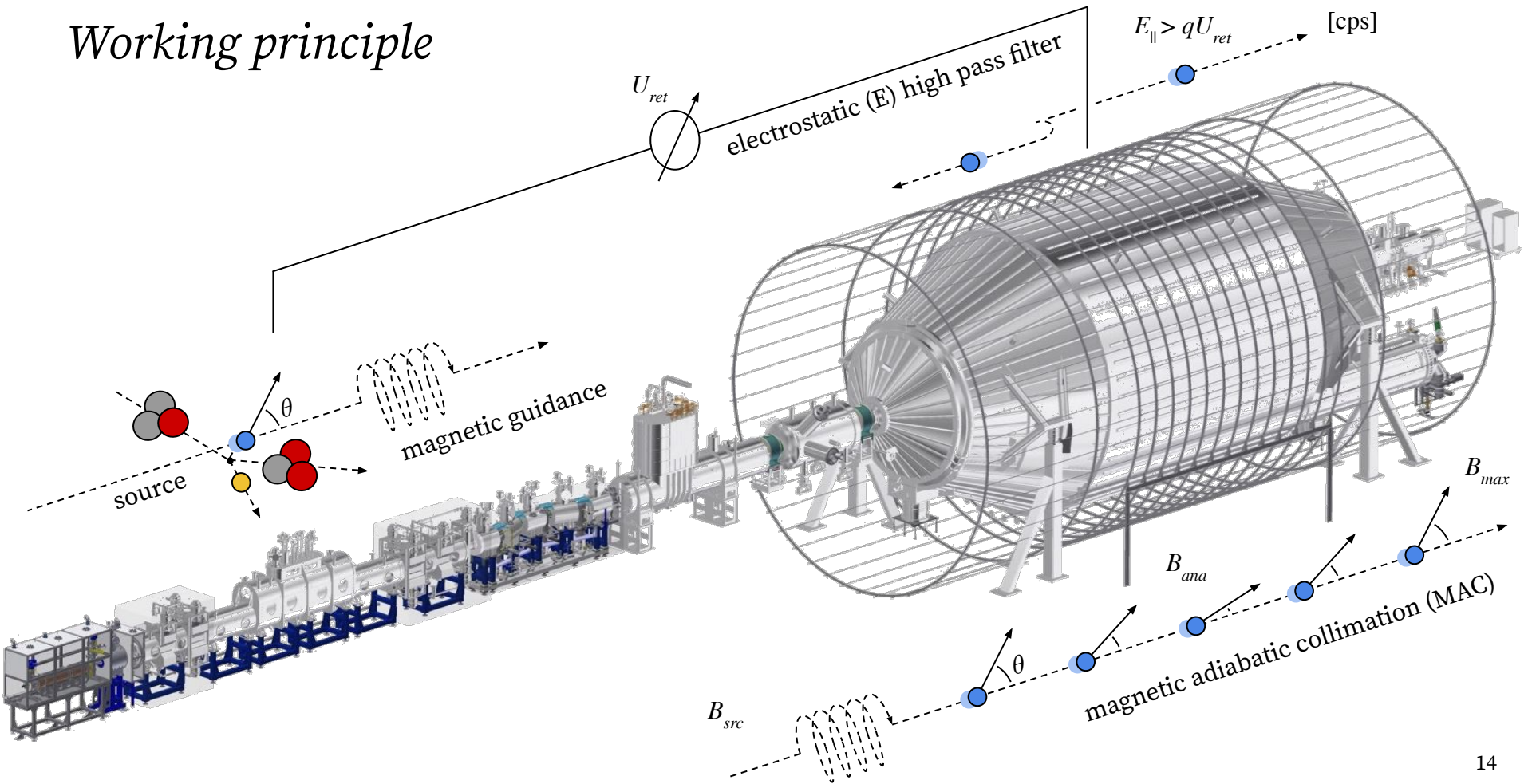
# *Working principle*



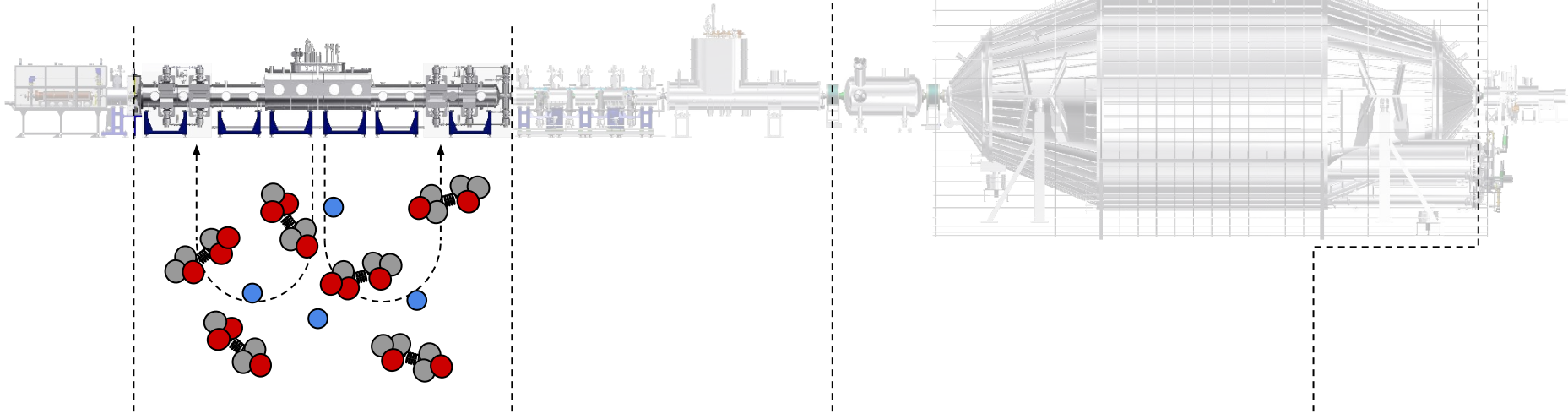
# Working principle



# Working principle



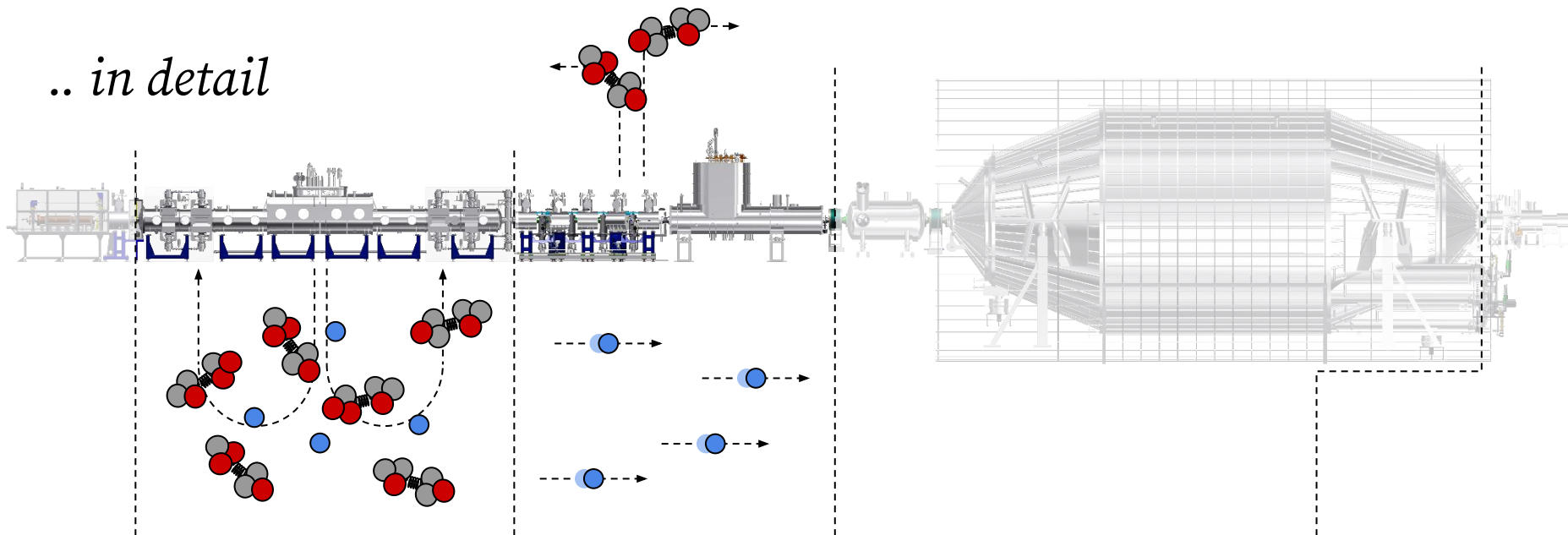
*.. in detail*



**windowless gaseous  
tritium source**

- molecular tritium  
in **closed loop**
- **100 GBq**

*.. in detail*



**windowless gaseous tritium source**

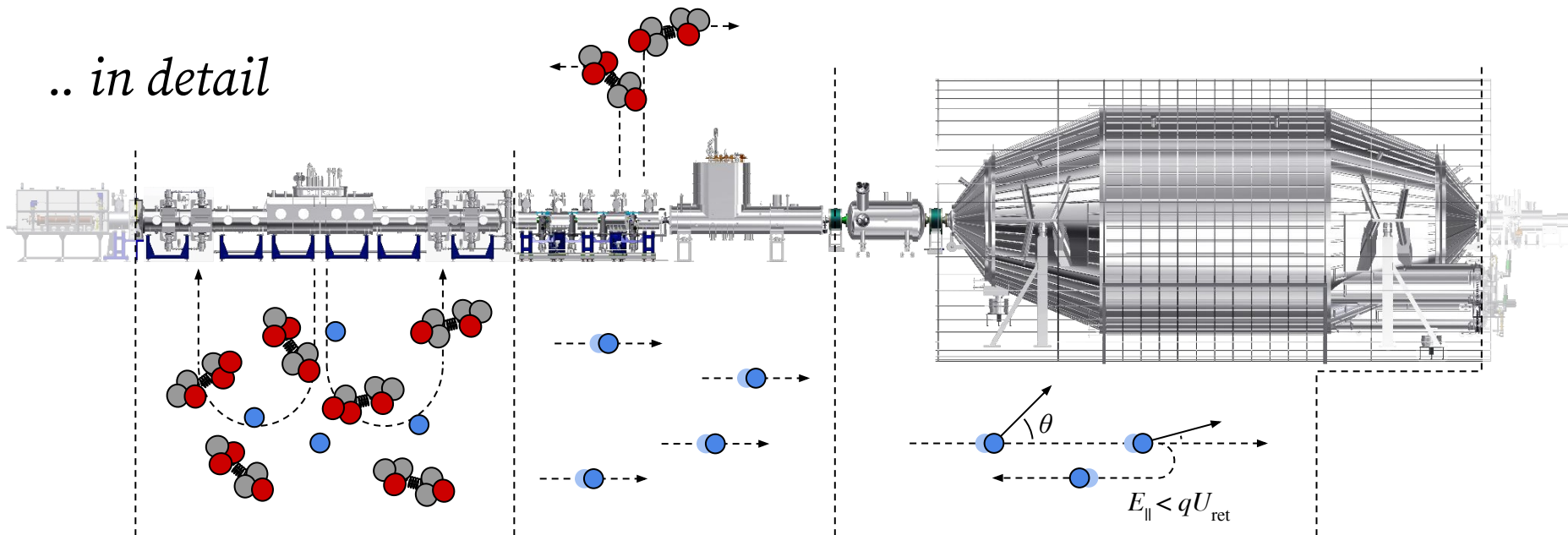
- molecular tritium in **closed loop**
- **100 GBq**

**transport section**

- tritium gas/ion removal
- reduction by **> 10<sup>14</sup>**



.. in detail



**windowless gaseous tritium source**

- molecular tritium in **closed loop**
- **100 GBq**

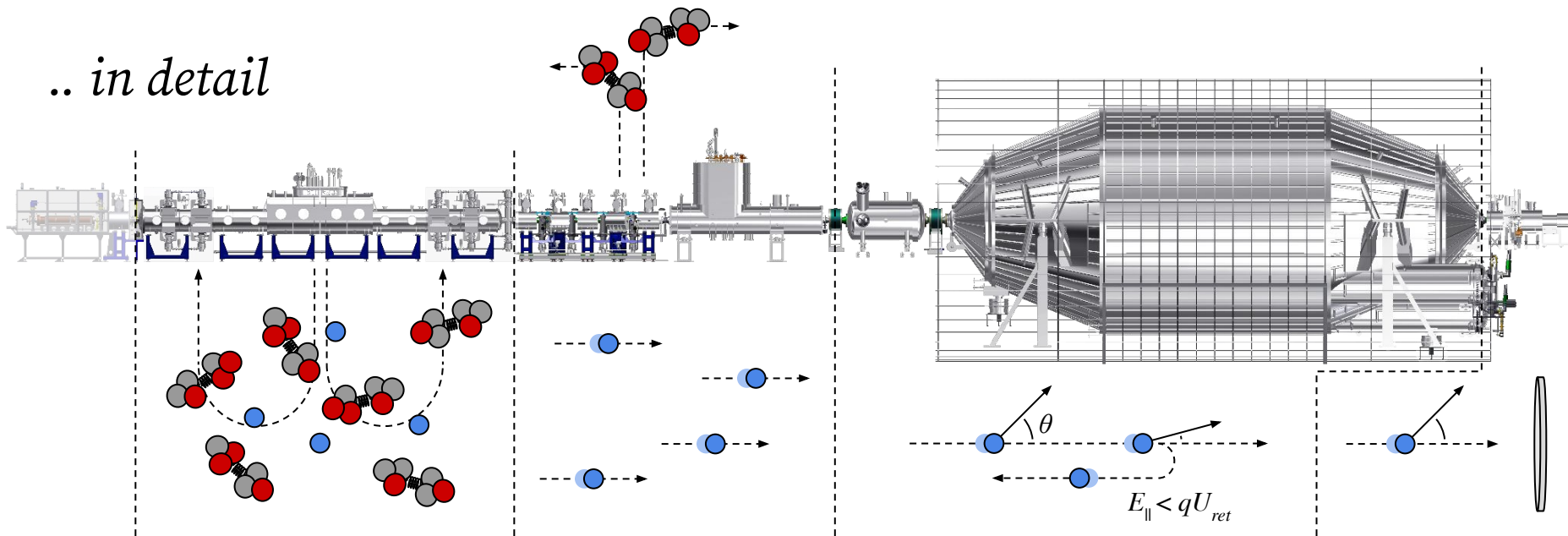
**transport section**

- tritium gas/ion removal
- reduction by  $> 10^{14}$

**spectrometer system**

- (pre-)/main-spectrometer
- high resolution, **0(1) eV**
- large acceptance angle, **0-51°**

.. in detail



**windowless gaseous tritium source**

- molecular tritium in **closed loop**
- **100 GBq**

**transport section**

- tritium gas/ion removal
- reduction by  $> 10^{14}$

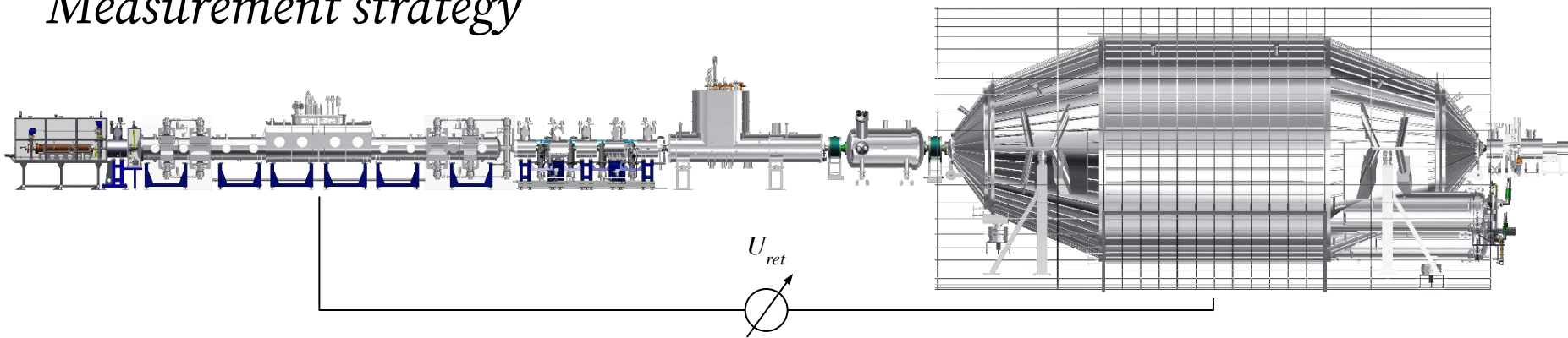
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- (pre-)/main-spectrometer
- high resolution, **0(1) eV**
- large acceptance angle, **0-51°**

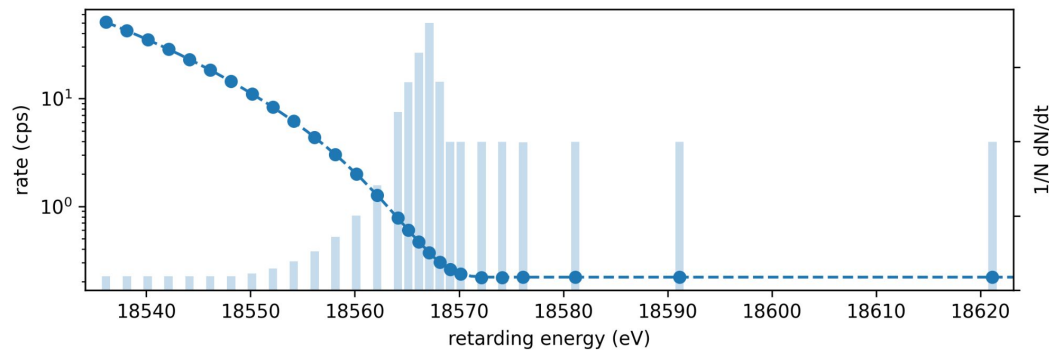
**detector section**

- focal plane detector, **148 pixel PIN-diode**

# Measurement strategy



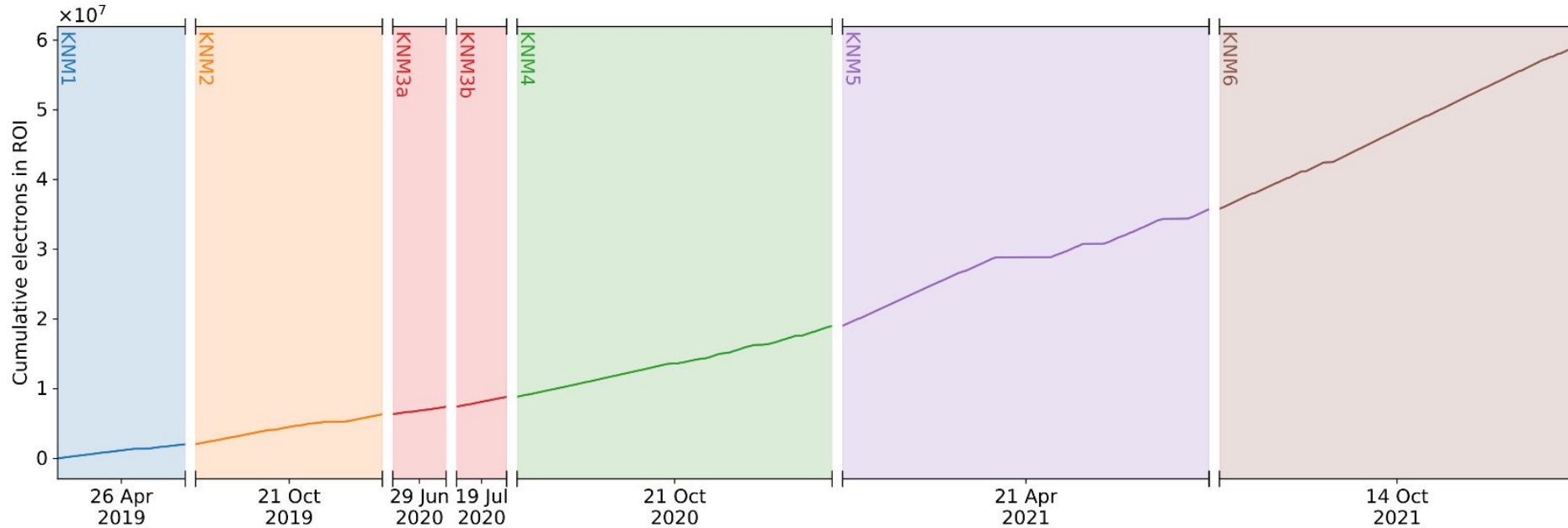
- ~**30 scan steps** with varying duration, measurement time distribution (MTD)
- ~**2 h scan** duration, up/down/random scans,  $O(100)$  scans per campaign
- several **campaigns** per year



# Data taking overview

**< 0.8 eV (90% CL)**

[Aker et al., Nature Phys. 18 (2022) 2, 160-166]



**< 1.1 eV (90% CL)**

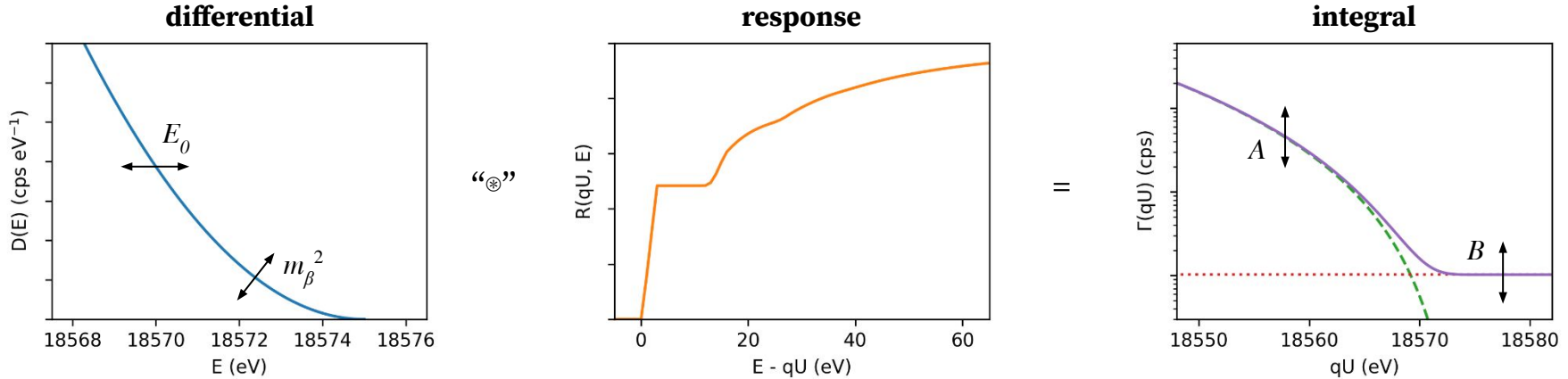
[Aker et al., PRL 123 (2019) 22, 221802]



# Analysis strategy

- maximum likelihood fit of **model**

$$\Gamma(qU) \propto A \int_{qU}^{E_0} D(E; m_\beta^2, E_0) R(qU, E) dE + B$$



with free **amplitude**  $A$ , **squared neutrino mass**  $m_\beta^2$ , **endpoint**  $E_0$  and **background**  $B$

- theoretical** (Fermi theory, molecular excitations) and **experimental** inputs (calibration measurements)

# Neutrino mass results

## 1<sup>st</sup> campaign, 2 million events (22 days)

[Aker et al., PRL 123 (2019) 22, 221802]

- best fit, **p-value = 0.6**

$$m_{\beta}^2 = (-1.0^{+0.9}_{-1.1}) \text{ eV}^2$$

### → upper limit

$$m_{\beta} < 1.1 \text{ eV (90\% CL)}$$

## 2<sup>nd</sup> campaign, 4 million events (31 days)

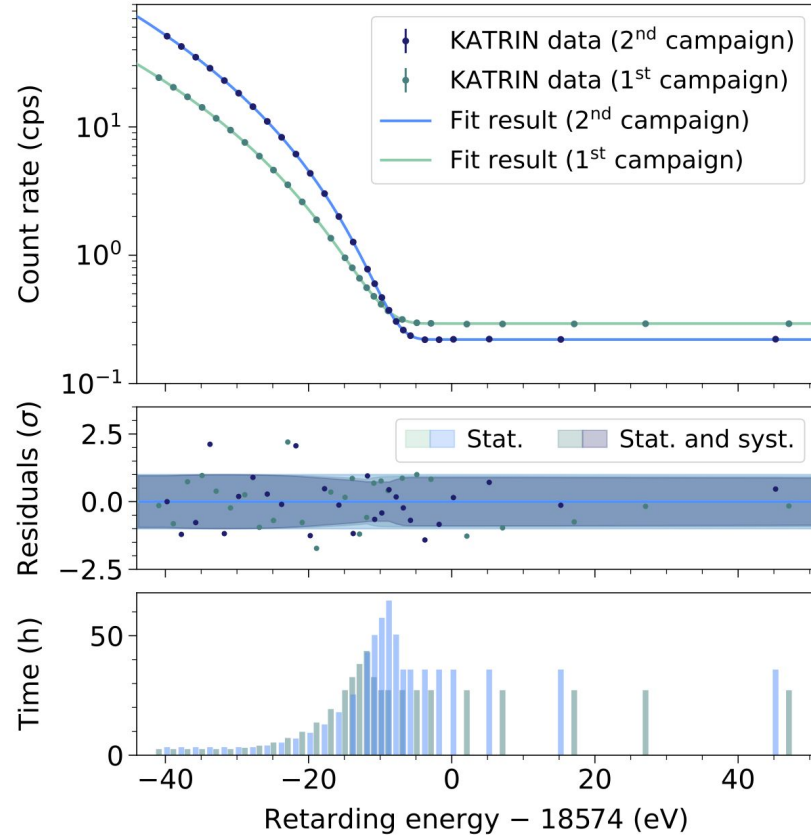
[Aker et al., Nature Phys. 18 (2022) 2, 160-166]

- best fit, **p-value = 0.8**

$$m_{\beta}^2 = (0.26 \pm 0.34) \text{ eV}^2$$

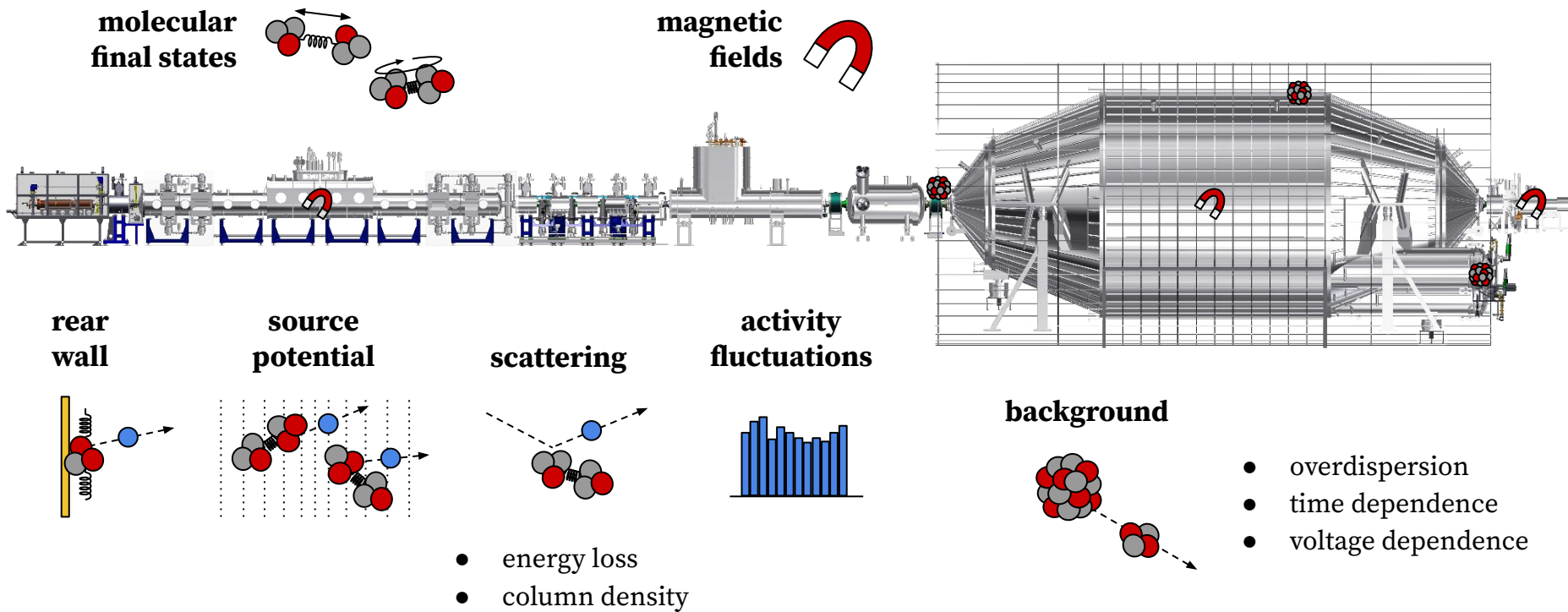
### → upper limit

$$m_{\beta} < 0.9 \text{ eV (90\% CL)}$$

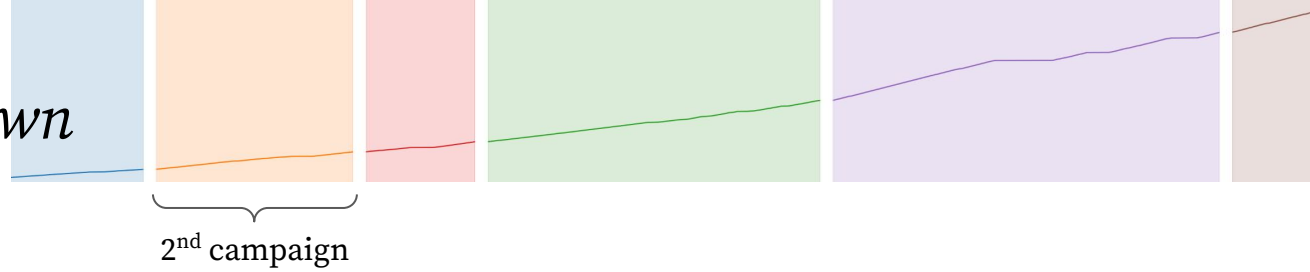


$$m_{\beta} < 0.8 \text{ eV (90\% CL)}$$

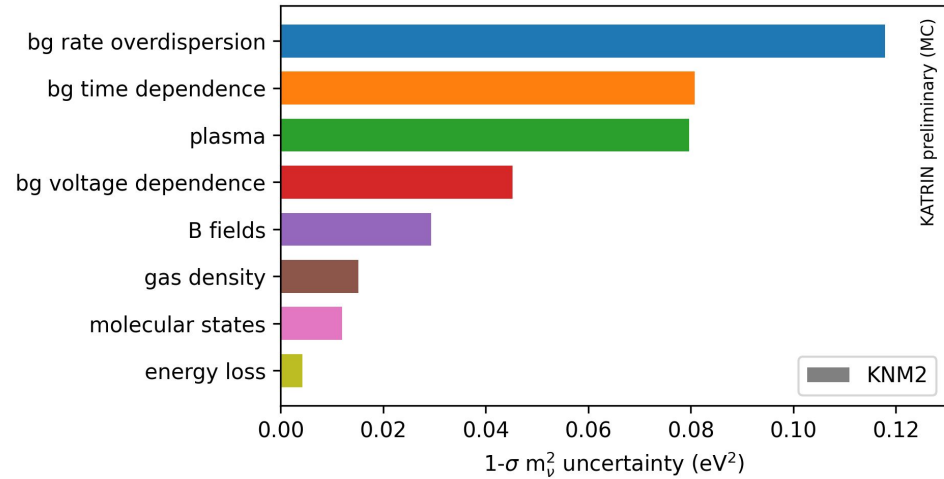
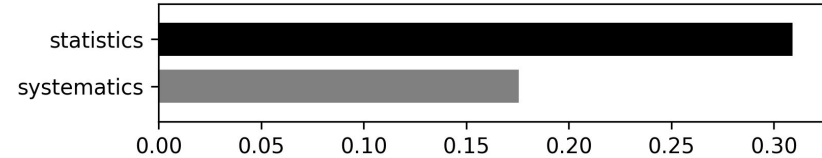
# Systematic uncertainties



# Uncertainty breakdown

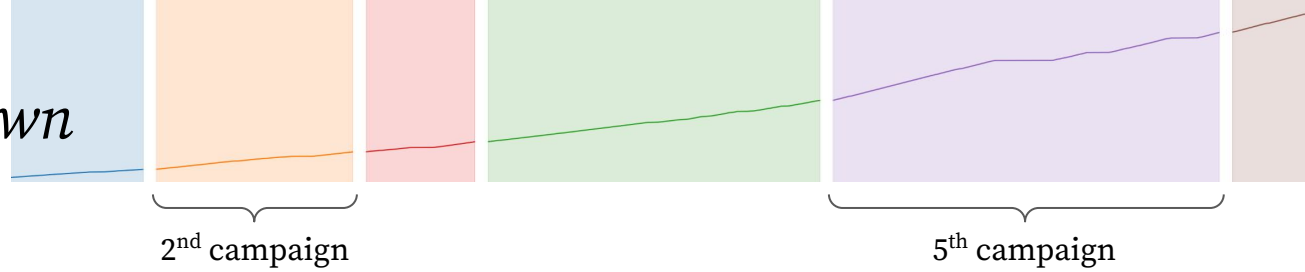


- **statistical uncertainty dominates**, systematics non-negligible
- **background**-related uncertainties dominate systematics budget
- significant **plasma** uncertainty





# Uncertainty breakdown



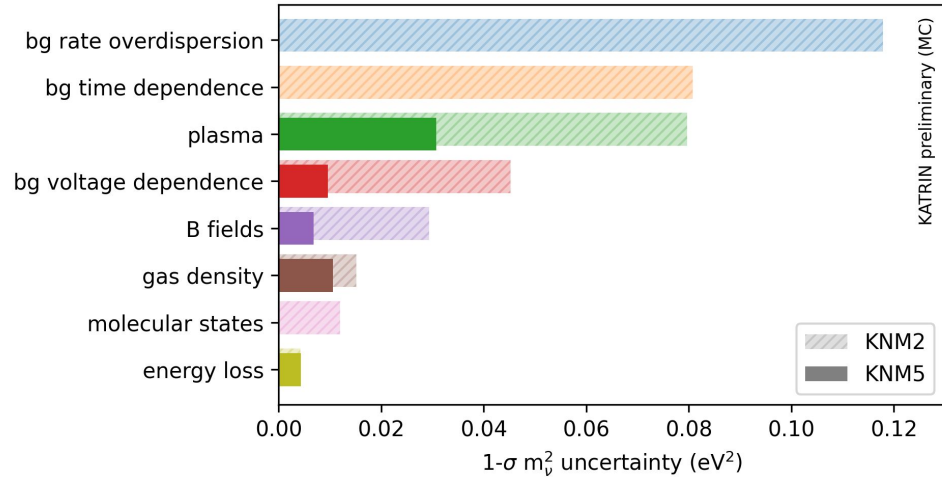
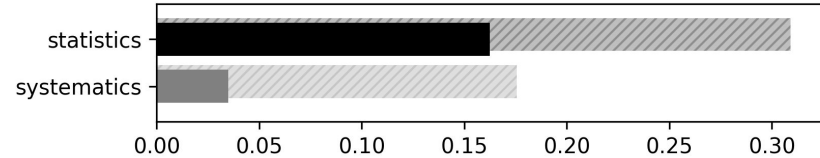
- **statistical uncertainty dominates**, systematics non-negligible
- still statistics dominated, **systematics largely improved**

- **background**-related uncertainties dominate systematics budget
- **mitigation** techniques, avoid Penning trap, shifted analyzing plane (SAP)

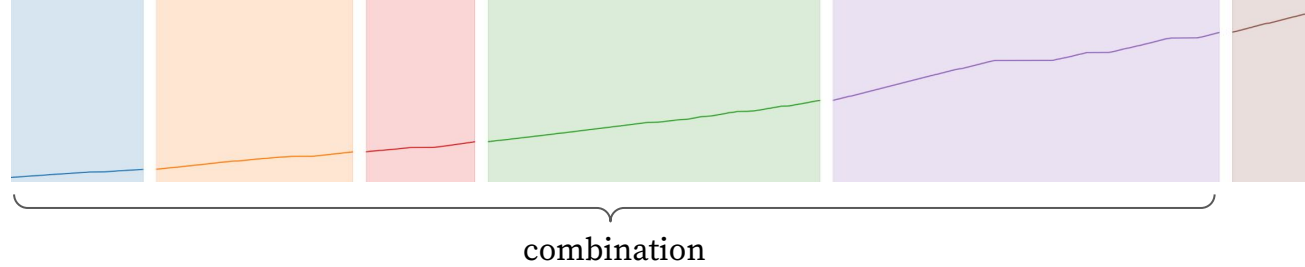
[Lokhov et al., EPJ C 82 (2022) 3, 258]

- significant **plasma** uncertainty
- high-statistics  $^{83m}\text{Kr}$  campaign, tritium scans at same temperature/gas density

[Altenmüller et al., J.Phys.G 47 (2020) 6, 065002]



# Outlook



- **combined analysis** of first 5 periods, significant increase of statistics
- model evaluation **computationally challenging**
- fast **neural network** interpolation

[Karl et al., EPJ C 82 (2022) 5, 439]

- **substantial improvement** of systematics and background
- sensitivity projection (in case of no signal)

$$m_{\beta} < 0.5 \text{ eV (90\% CL)}$$

- **data taking ongoing**
- final sensitivity goal

$$m_{\beta} < 0.2 \text{ eV (90\% CL)}$$

## beyond neutrino mass:

- search for eV-scale **sterile neutrinos**

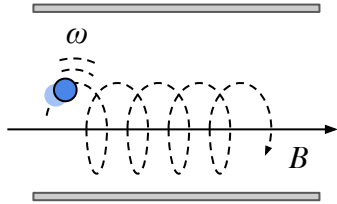
[Aker et al., PRD 105 (2022) 7, 072004]

- **relic neutrino** search

[Aker et al., arXiv:2202.04587]

- test of **Lorentz invariance**

[Lehnert, PLB 828 (2022) 137017]



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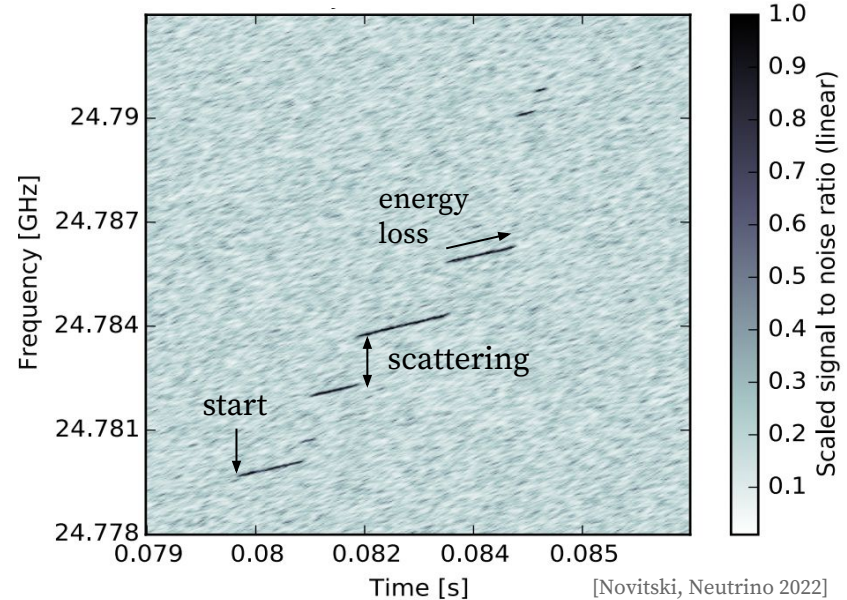
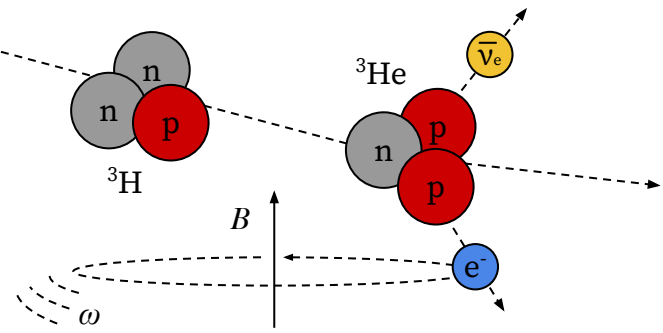
# Working principle

- **cyclotron radiation** emission spectroscopy (CRES) for tritium decay electrons

[Monreal, Formaggio, PRD 80 (2009) 051301]

$$\omega(\gamma) = \frac{\omega_0}{\gamma} = \frac{eB}{E + m_e}$$

- **source transparent** to microwave radiation, *source = detector* concept
- high precision **differential** frequency measurement, eV-scale resolution
- **low background**



# Project 8

- **proof-of-concept**, single electron spectroscopy
- (molecular) tritium endpoint measurement, first **neutrino mass limit**

$$m_{\beta} < 178 \text{ eV (90\% CL)}$$

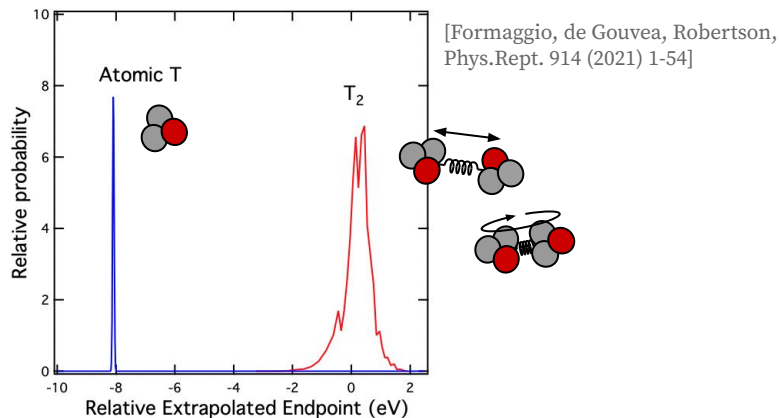
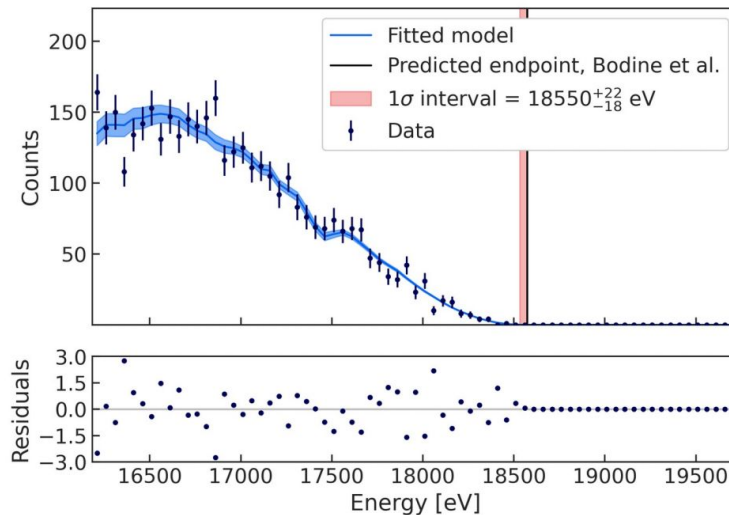
[Novitski, Neutrino 2022]

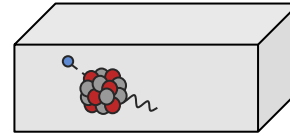
path beyond:

- **m<sup>3</sup>-scale trap** (antenna array or cavity resonator)
  - **atomic tritium** source
- sensitivity down to 40 meV

[Ashtari Esfahani et al., arXiv:2203.07349]

[Novitski, Neutrino 2022]





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# Working principle

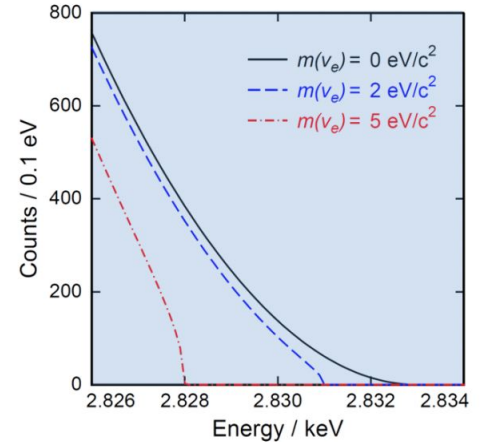
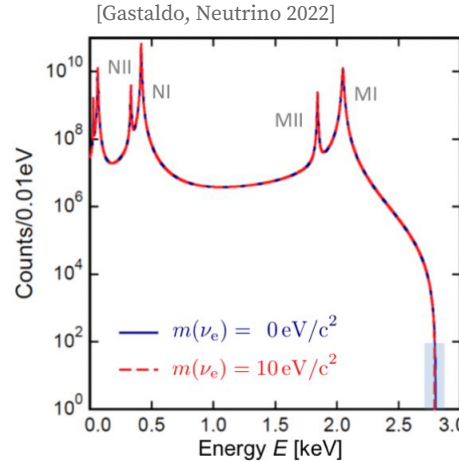
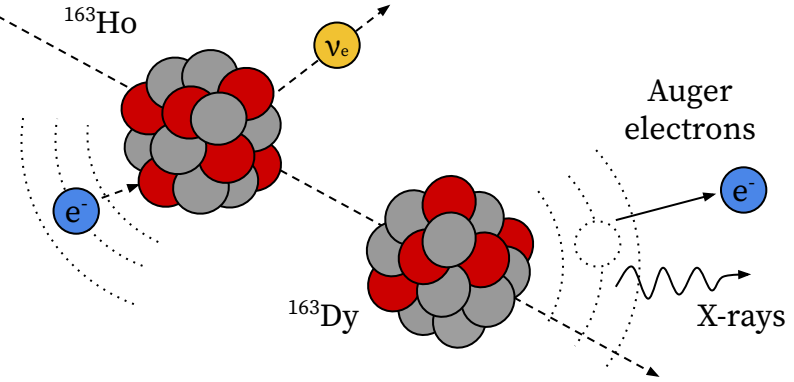
- calorimetric measurement of  $^{163}\text{Ho}$  **electron capture decay**

[De Rujula, Lusignoli, PLB 118 (1982) 429]

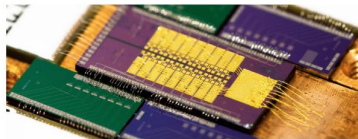
  - super-low **Q-value**

$$Q_{EC} = (2.833 \pm 0.034) \text{ keV}$$

[Eliseev et al., PRL 115 (2015) 062501]
  - **sub-eV** sensitivity requires **MBq-scale activity**
- $^{163}\text{Ho}$  **implanted** into **cryogenic micro-calorimeters**
  - eV-scale **differential** measurement
  - *source = detector* concept, pile-up limits pixel activity



# ECHo



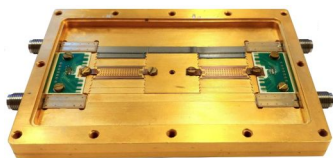
- array of **metallic magnetic calorimeters** (MMC) with  $^{163}\text{Ho}$ -implanted absorber, 10 Bq per pixel
- first **neutrino mass limit** (4 pixels with 0.2 Bq)

$$m_\nu < 150 \text{ eV (95\% CL)}$$

[Velte et al., EPJ C 79 (2019) 12, 1026]

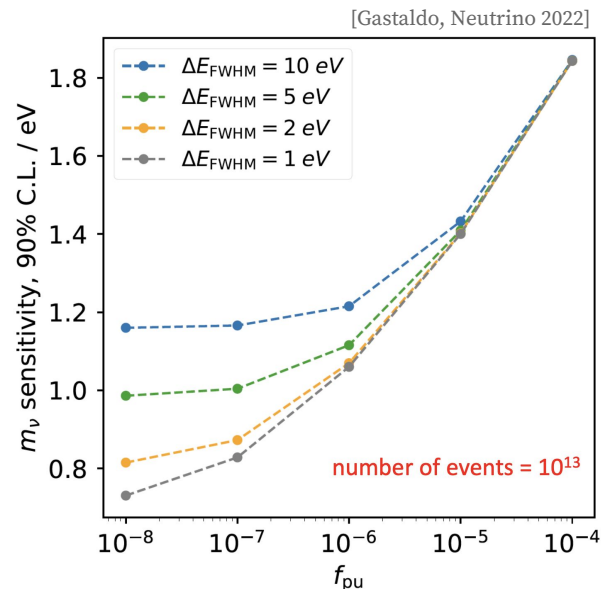
- analysis of **new data** ongoing (60 pixels with 1 Bq)  
sensitivity:  $m_\nu < 20 \text{ eV (95\% CL)}$

# HOLMES



- array of **transition edge sensors** (TES) coupled to  $^{163}\text{Ho}$ -implanted absorber, 300 Bq per pixel

sensitivity for **coming phases** of ECHo/HOLMES:

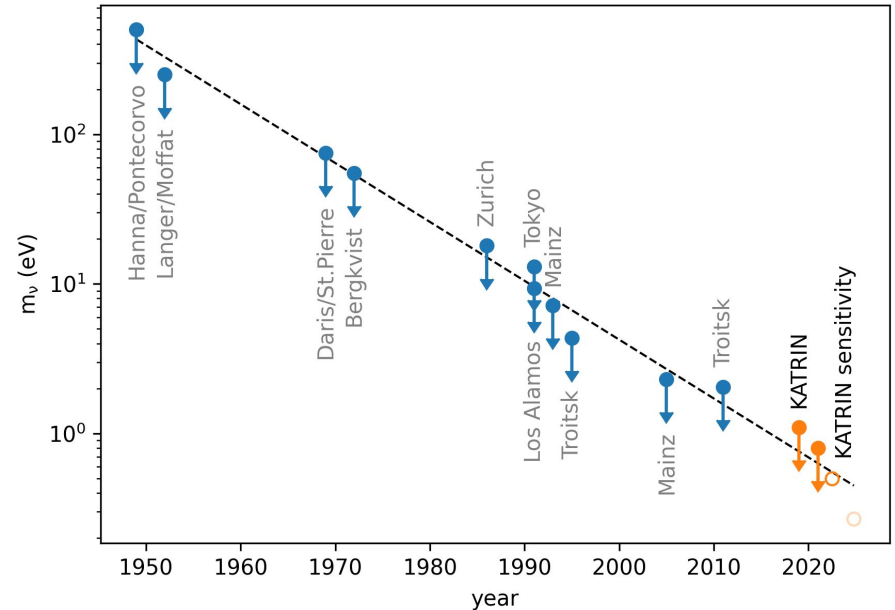




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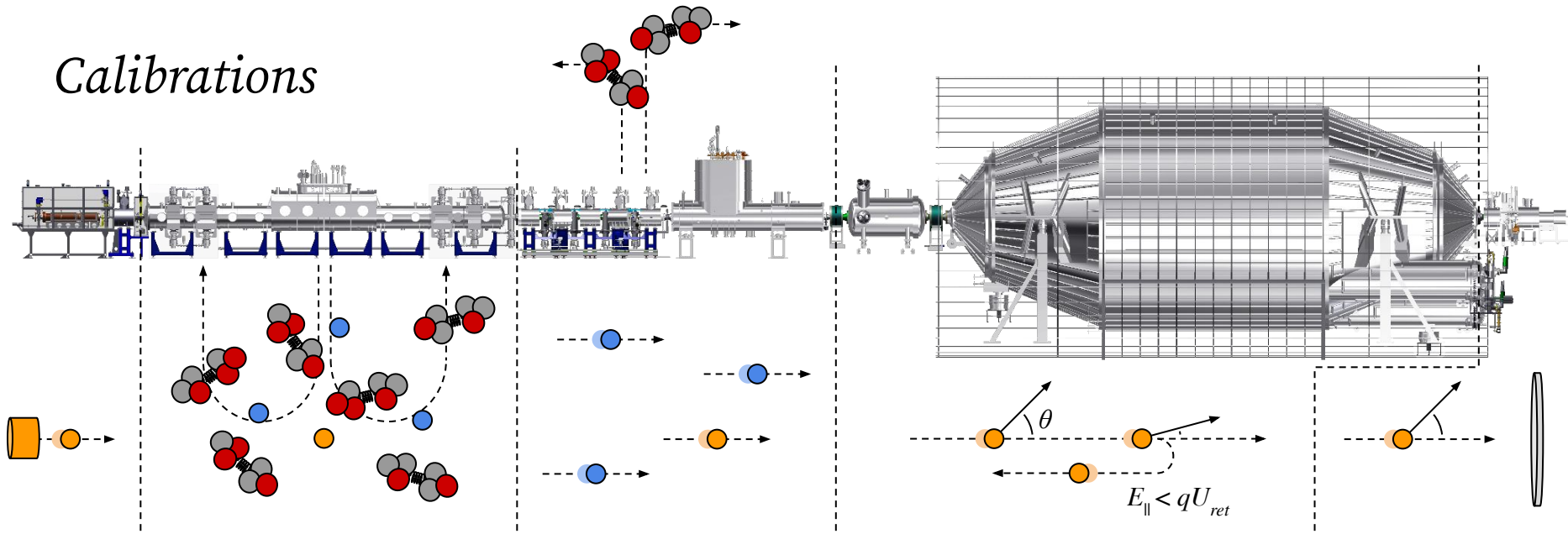
# Conclusions

- **KATRIN** measurement ongoing
  - first direct **sub-eV** neutrino mass limit  
[Aker et al., Nature Phys. 18 (2022) 2, 160-166]  
 $m_\beta < 0.8 \text{ eV}$  (90% CL)
  - substantial improvement of **systematics** and **background**, increased **statistics**
- **promising perspectives** to go beyond
  - tritium cyclotron radiation emission spectroscopy, **Project 8**
  - holmium-implanted cryogenic micro-calorimeters, **ECHO/HOLMES**
  - ...



*Backup*

# Calibrations



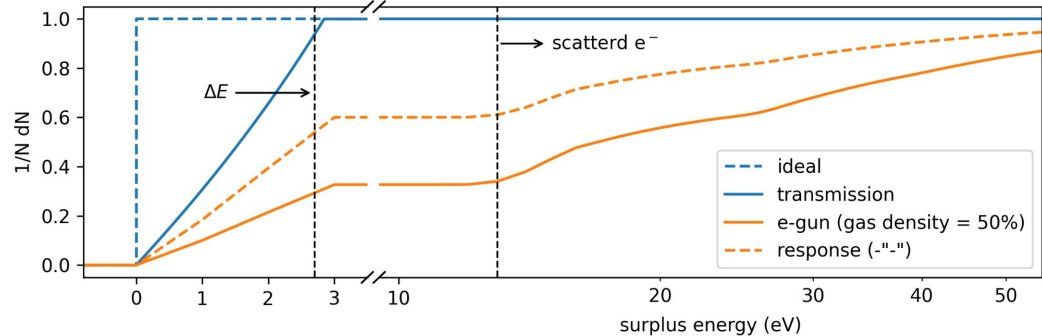
## rear section

- high intensity **e-gun**

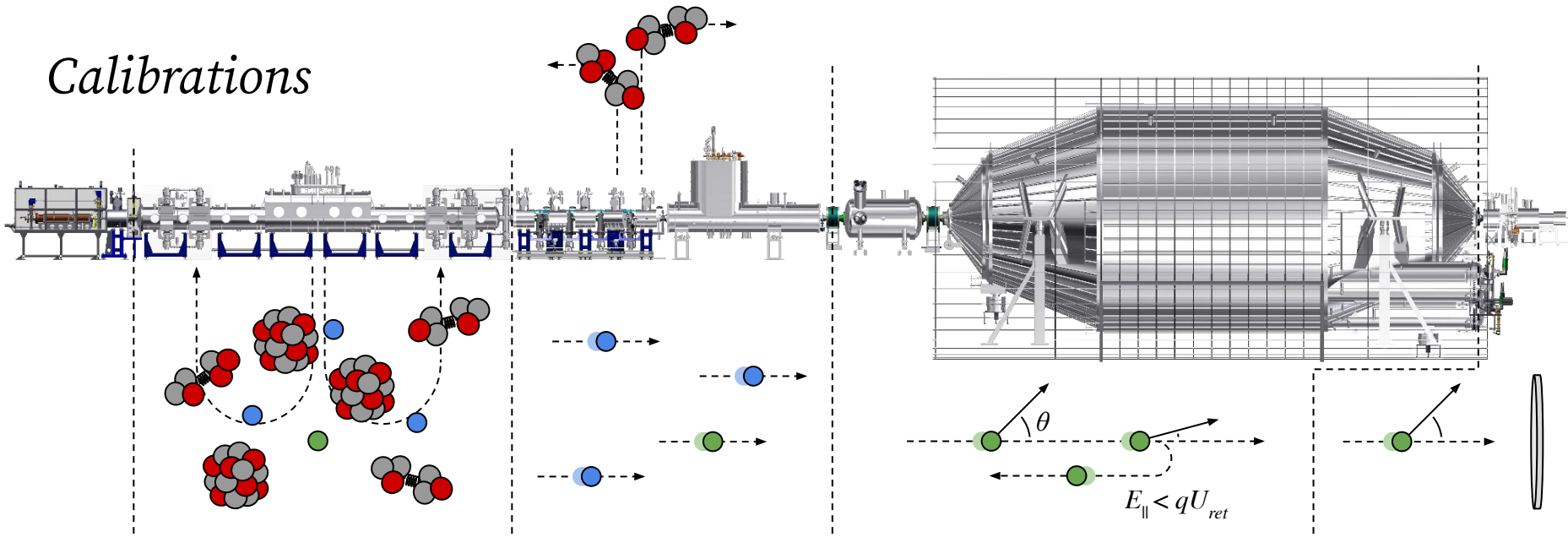
[Behrens et al., EPJ C 77 (2017) 6, 410]

- precise determination of **column density** and **energy-loss function**

[Hannen et al., Astropart.Phys. 89 (2017) 30-38;  
Aker et al., EPJ C 81 (2021) 7, 579]



# Calibrations



## windowless gaseous tritium source

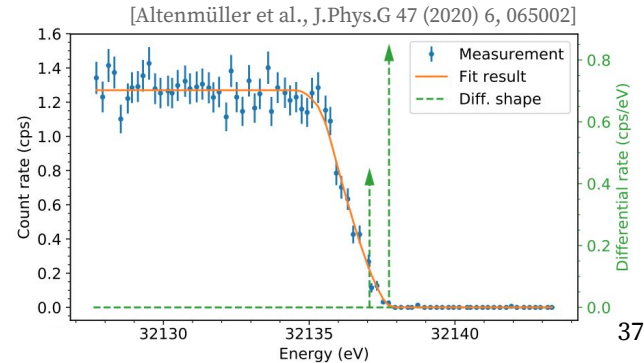
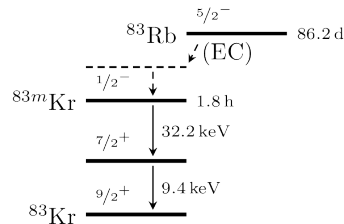
- $^{83m}\text{Kr}$  from  $^{83}\text{Rb}$ , conversion electrons

[Sentkerestiová et al., J.Phys.Conf.Ser. 888 (2017) 1, 012072;

Arenz et al., JINST 13 (2018) 04, P04020]

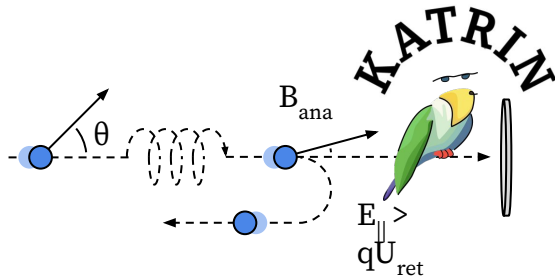
→ variations of **source potential** and **spectrometer fields**

[Arenz et al., EPJ C 78 (2018) 5, 368]

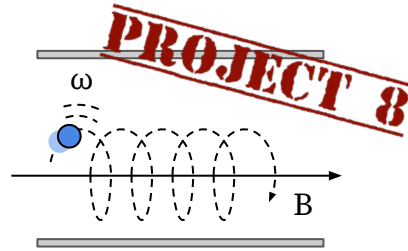


# Experimental efforts

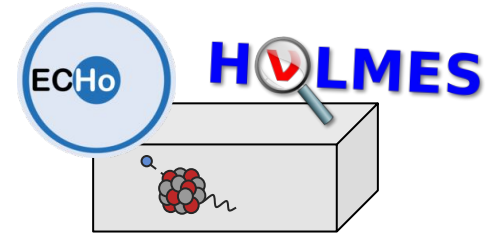
tritium



electrostatic **filter**  
(MAC-E)



**cyclotron radiation**  
emission spectroscopy  
(CRES)



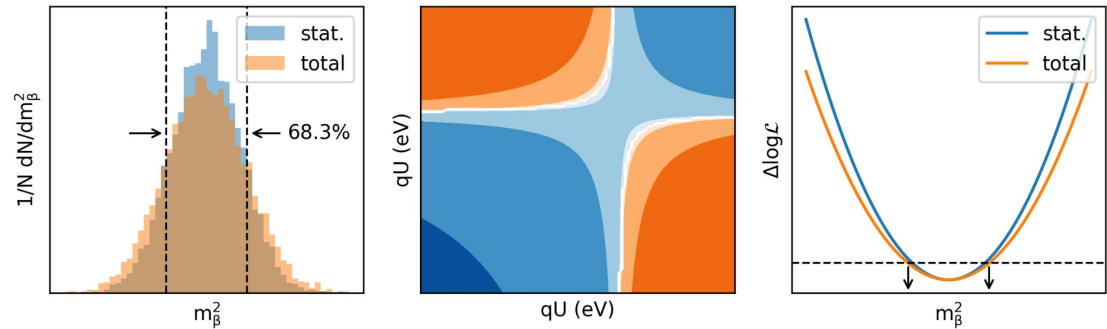
cryogenic  
**calorimeters**

R&D phase

# Analysis strategy

- **blind analysis**, procedures and inputs frozen on Monte Carlo data (MC twins) with  $m_\beta = 0$
- cross-checks between **3 independent analysis frameworks/groups**
- **2-step unblinding**
  - fit on data w/ **blinded molecular final state distribution (FSD)**
  - fit on data w/ correct FSD

- **uncertainty propagation**
  - Monte Carlo propagation (sampling)
  - covariance matrices
  - nuisance parameter, pull terms



# KATRIN neutrino mass (KNM) 2

[Aker et al., Nature Phys. 18 (2022) 2, 160-166]

- analysed **ring-wise** (consistent with uniform fit), excellent goodness-of-fit, **p-value = 0.8**,

$$m_{\beta}^2 = (0.26 \pm 0.34) \text{ eV}^2$$

compatible with zero

- **upper limit** using Likhov-Tkachov

[Likhov, Tkachov, Phys.Part.Nucl. 46 (2015) 3, 347-365]

(consistent with Feldman-Cousins)

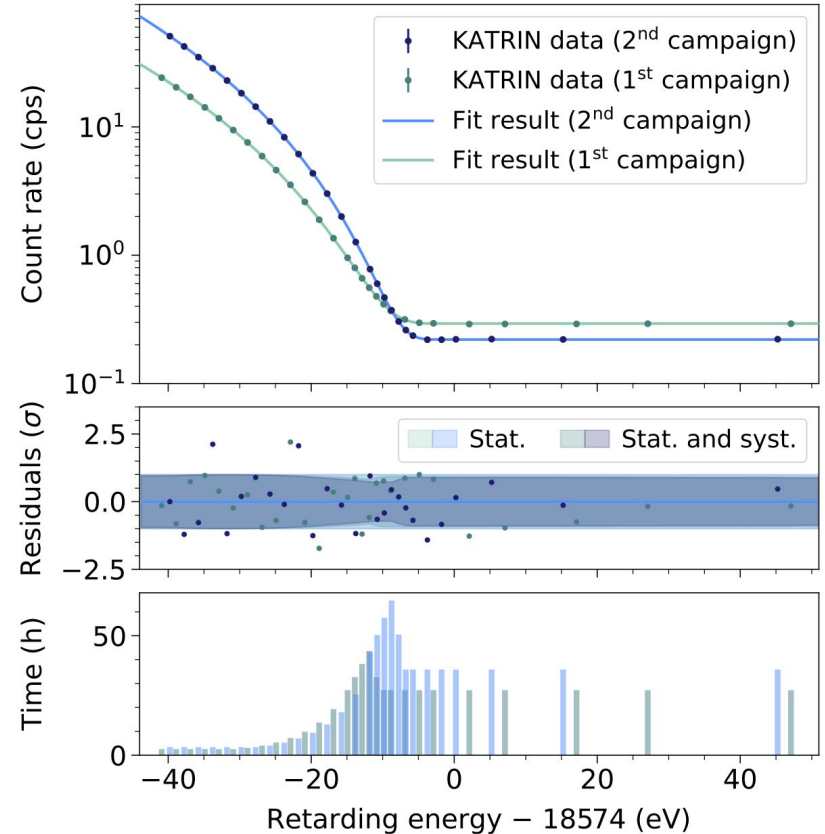
[Feldman, Cousins, PRD 57 (1998) 3873-3889]

$$m_{\beta} < 0.9 \text{ eV (90\% CL)}$$

sensitivity:  $m_{\beta} < 0.7 \text{ eV (90\% CL)}$

- combination with KNM1

$$m_{\beta} < 0.8 \text{ eV (90\% CL)}$$





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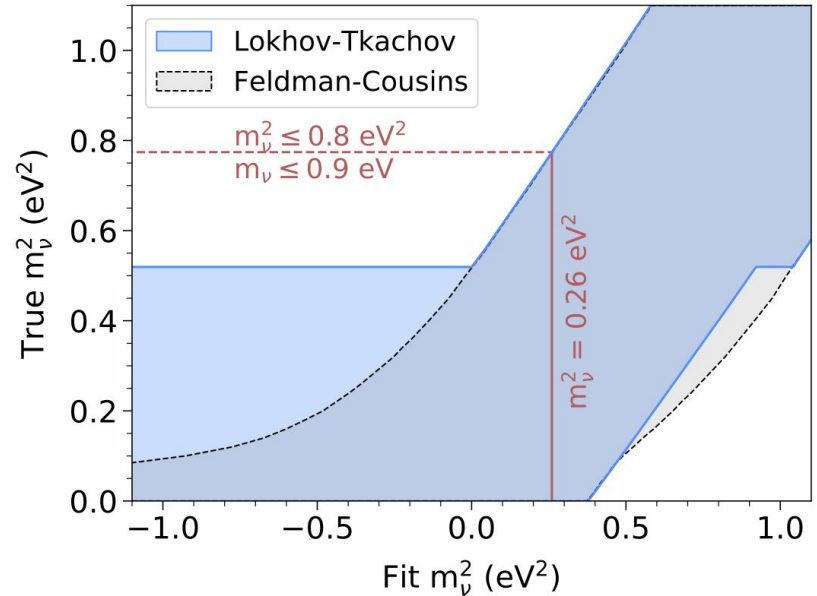
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# Beyond neutrino mass

- search for eV-scale **sterile neutrinos**

→ kink in spectrum

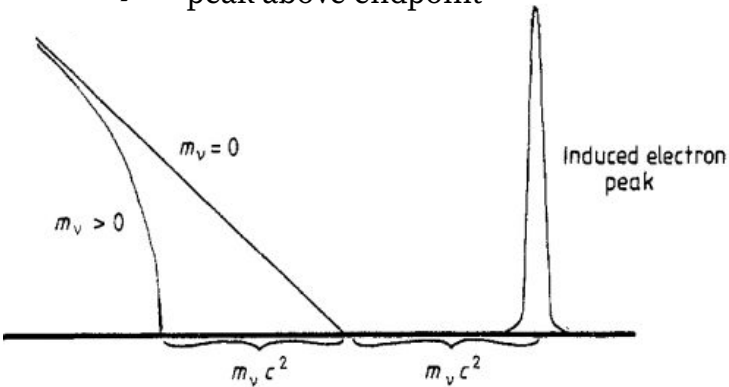
[Aker et al., PRD 105 (2022) 7, 072004]

- keV-scale steriles with **TRISTAN**

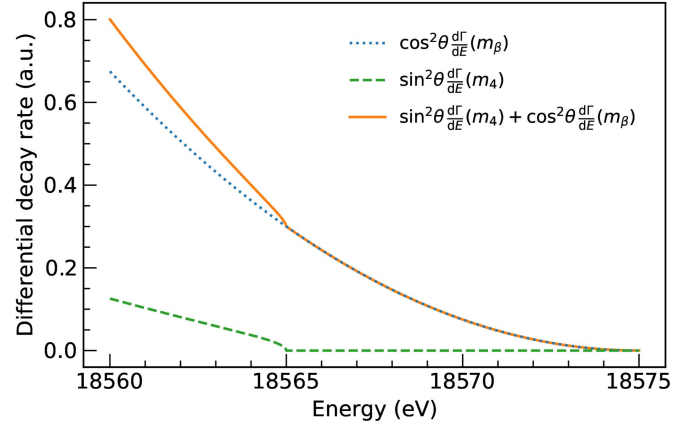
[Mertens et al., J.Phys.G 46 (2019) 6, 065203]

- **relic neutrino** overdensity

→ peak above endpoint



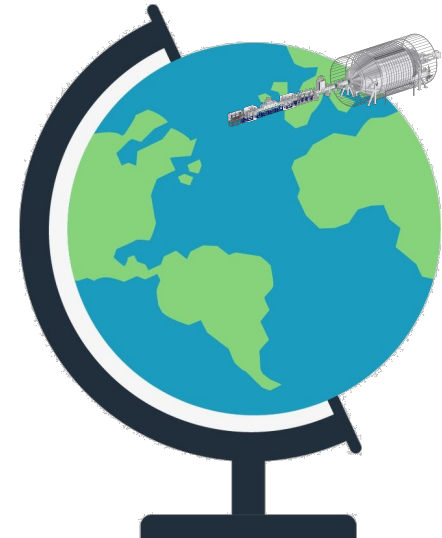
[Irvine, Humphreys, J. Phys. G 9 (1983) 847]



- **Lorentz invariance violation**

→ sidereal modulation

[Lehnert, PLB 828 (2022) 137017]



# Spectrometer backgrounds

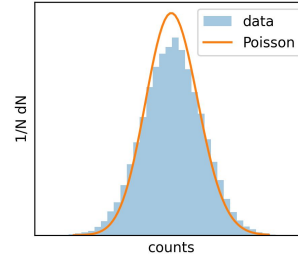
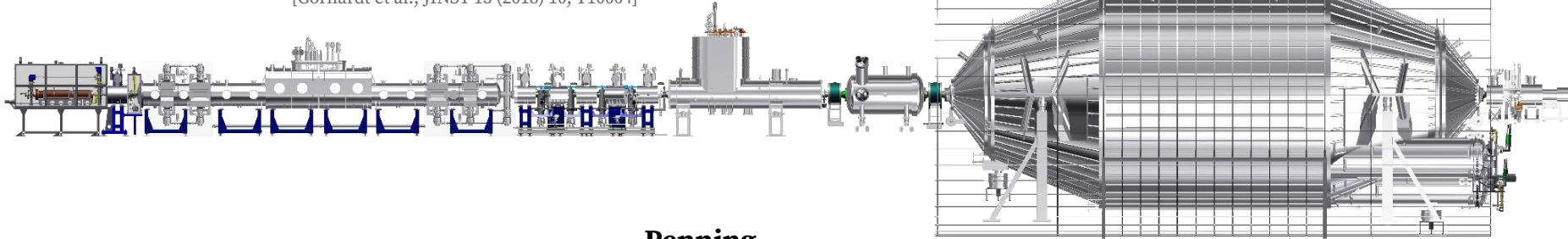
- $^{219}\text{Rn}$  decays ( $T_{1/2} = 4\text{s}$ ) in spectrometer

→ trapped electrons, **non-Poisson rate**

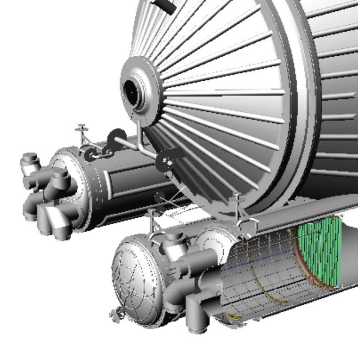
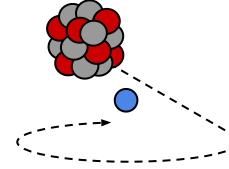
[Mertens et al., *Astropart.Phys.* 41 (2013) 52-62]

- improved nitrogen cooled baffles

[Görhardt et al., *JINST* 13 (2018) 10, T10004]



## Radon-219



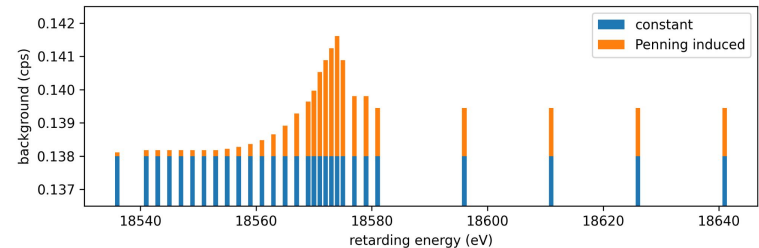
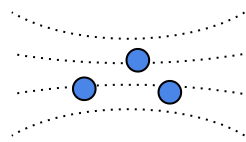
- emptied in between scan steps

→ **background increase** during measurement  $O(1) \mu\text{cps/s}$

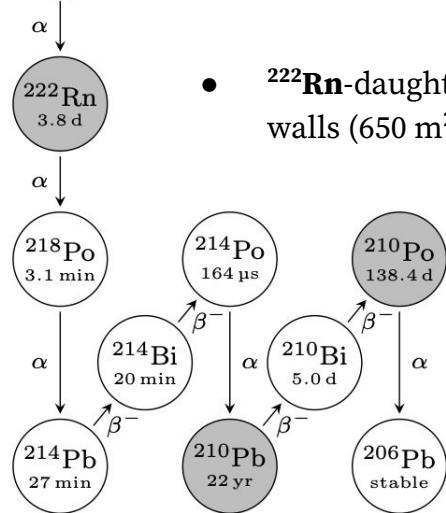
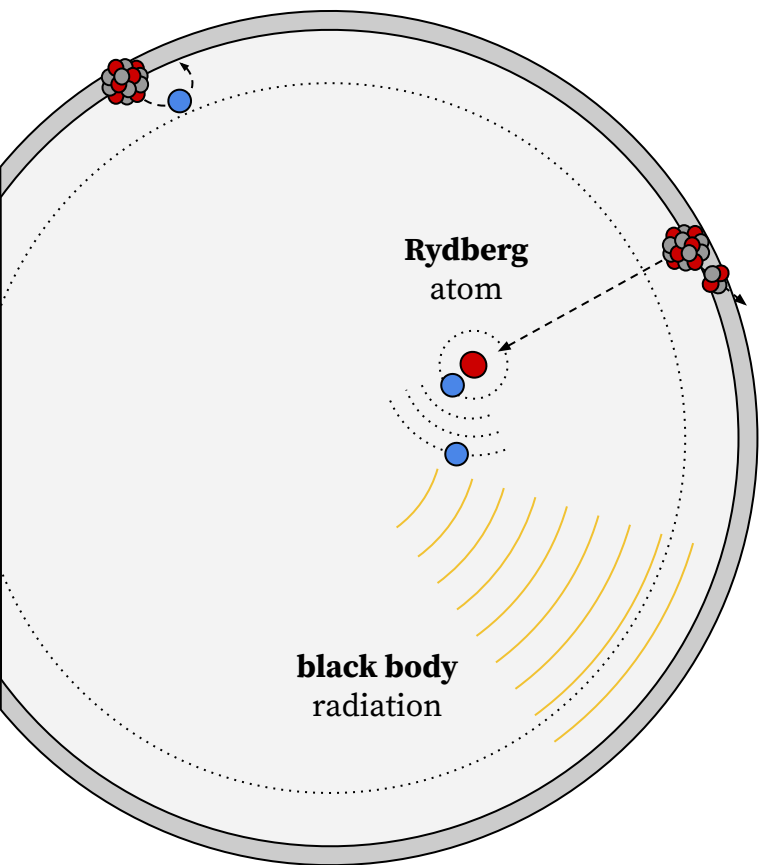
→ scan-step-**duration dependent**

- switch off pre-spectrometer

## Penning trap



# Spectrometer backgrounds



- $^{222}\text{Rn}$ -daughter **plate out** on spectrometer walls ( $650 \text{ m}^2$ )

→ sputtering of **Rydberg atoms** in  $^{210}\text{Po}$  decay

→ **ionized** by black body radiation

- **shifted-analyzing plane (SAP)**

