



**New results
from
the CUORE experiment**



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Outlook

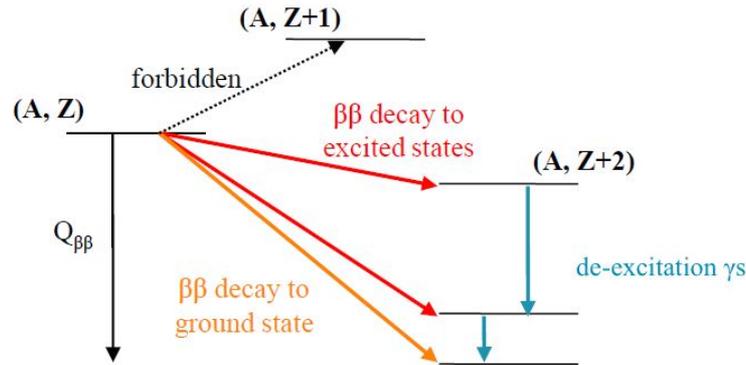
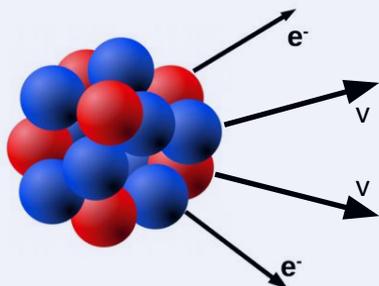
- Neutrinoless double beta decay
- The CUORE experiment
- CUORE initial operations & optimization
- CUORE physics data taking
 - ◆ Data processing
 - ◆ New results from CUORE on search for ^{130}Te $0\nu\beta\beta$ decay
- Conclusions

Double beta decay

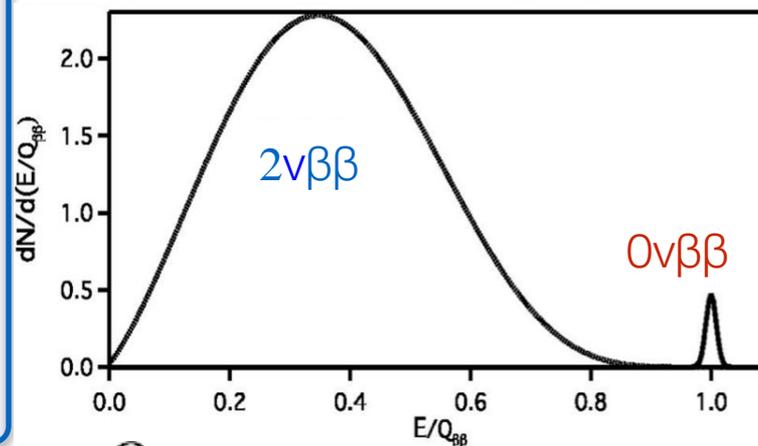
Double beta decay is a rare nuclear decay: $(N, Z) \rightarrow (N-2, Z+2)$

Two neutrino double beta decay ($2\nu\beta\beta$)

- 2nd order process allowed in SM ($\Delta L = 0$)
- Observed in several nuclei: $T^{1/2}_{2\nu\beta\beta} \sim 10^{18-24}$ yr



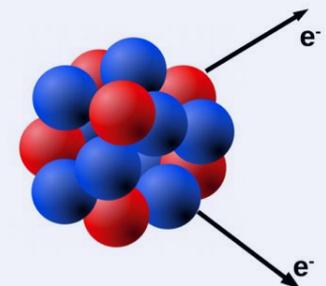
$\beta\beta$ summed e^- energy spectrum



$Q_{\beta\beta}$ Q-value of the $\beta\beta$ decay

Neutrinoless double beta decay ($0\nu\beta\beta$)

- Beyond SM process ($\Delta L = 2$)
- Not yet observed
- $T^{1/2}_{0\nu\beta\beta} > 10^{24-26}$ yr

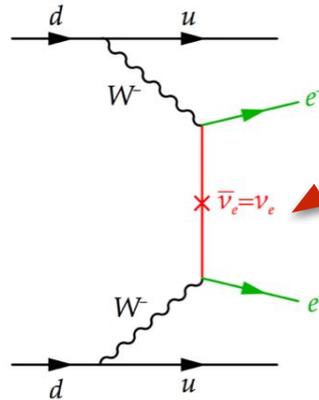


Neutrinoless double beta decay

Observation of $0\nu\beta\beta$ decay would imply:

- Lepton number violation
- Hint on origin of matter/anti-matter asymmetry
- Presence of a Majorana term for the neutrino mass
- Constraints on neutrino mass hierarchy and scale

$0\nu\beta\beta$ mechanism:
Light Majorana neutrino exchange



Majorana mass term
coupling
the two neutrinos

From $0\nu\beta\beta$ decay rate measurements one can infer the effective neutrino mass term

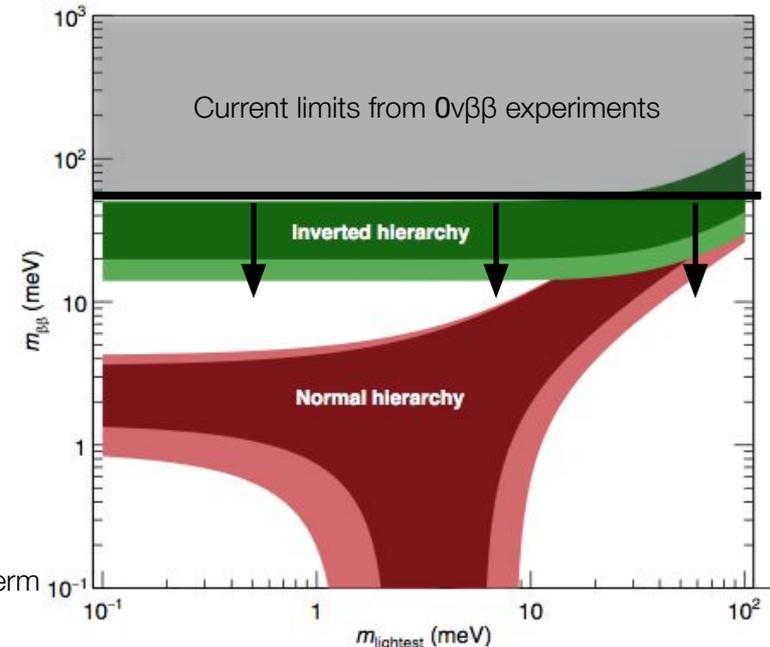
$$\frac{1}{T_{1/2}^{0\nu}} \propto G(Q_{\beta\beta}, Z) |M_{nucl}|^2 |m_{\beta\beta}|^2$$

Phase space integral

Nuclear matrix element
(NME)

Effective neutrino mass term

$$m_{\beta\beta} = \left| \sum_i m_{\nu_i} U_{ei}^2 \right|$$



Experimental $0\nu\beta\beta$ half-life sensitivity

Finite background

$$S_{0\nu} \propto \eta \cdot \epsilon \cdot \sqrt{\frac{M \cdot T}{\Delta \cdot B}}$$

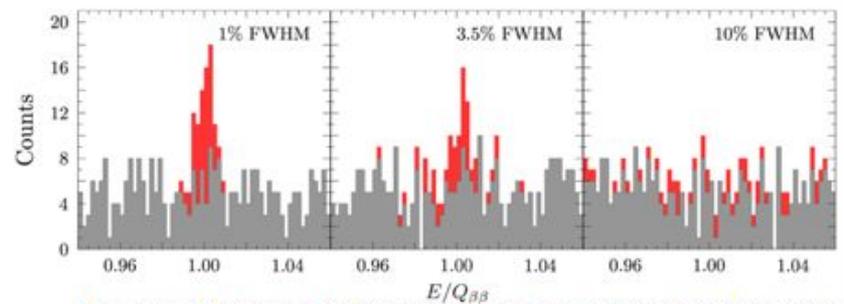
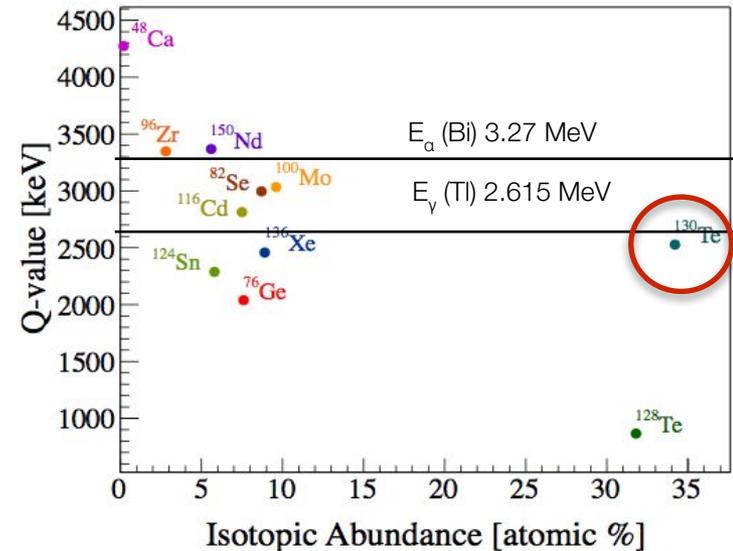
Suitable nuclei for double-beta decay are
 48Ca, 76Ge, 82Se, 100Mo, 116Cd,
 130Te, 128Te, 136Xe...

Isotope choice

- High isotopic natural abundance or enrichment
- High $Q_{\beta\beta}$
- Good detection efficiency:
 $\beta\beta$ source embedded into the absorber
- Excellent energy resolution

Exposure

- Large active mass (M) detector
- Long live-time



J. J. Gómez-Cadenas and J. Martín-Albo, Phenomenology of neutrinoless double beta decay, Proc. of Science (GSS14), 004 (2015)

Search for $0\nu\beta\beta$ decay of ^{130}Te

- ^{130}Te highest natural isotopic abundance $\eta(^{130}\text{Te}) = 34.167\%$,
no detector enrichment
- ^{130}Te within the detector absorber of TeO_2 (**containment efficiency ~ 90%**)
- $Q_{\beta\beta}(^{130}\text{Te}) = 2527.518$ keV, above most of the natural radioactivity
- TeO_2 operated as low temperature macro-calorimeters (~ 10 mK): good energy resolution Δ (**$\sim 0.2\%$ FWHM/E**), compared to other $\beta\beta$ experiments
- Reproducible growth of large number of high quality detectors; **large active mass detector**: 742 kg TeO_2 (~ 206 kg of ^{130}Te)

The CUORE experiment

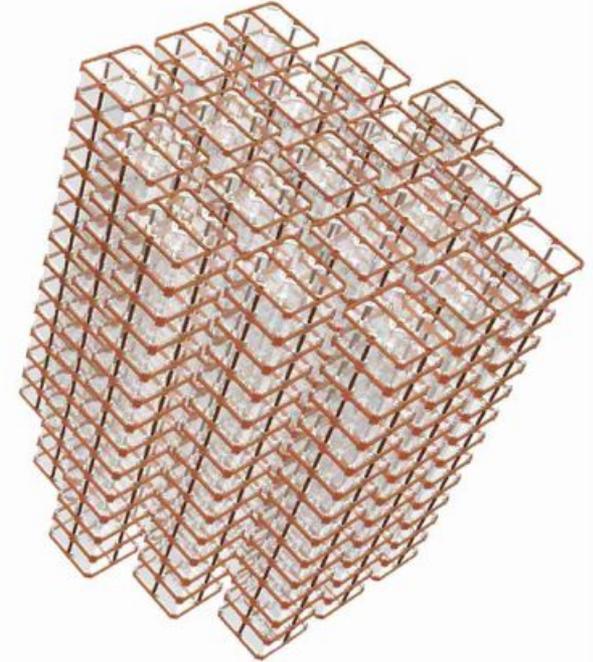
CUORE

Cryogenic Underground Observatory for Rare Events

Ton-scale cryogenic experiment
using 988 $^{(\text{nat})}\text{TeO}_2$ crystals operated at ~ 10 mK
- large mass and high granularity -

Primary goal:
Search for $0\nu\beta\beta$ decay in ^{130}Te

Test Majorana nature of neutrino $\longleftrightarrow 0\nu\beta\beta$ decay

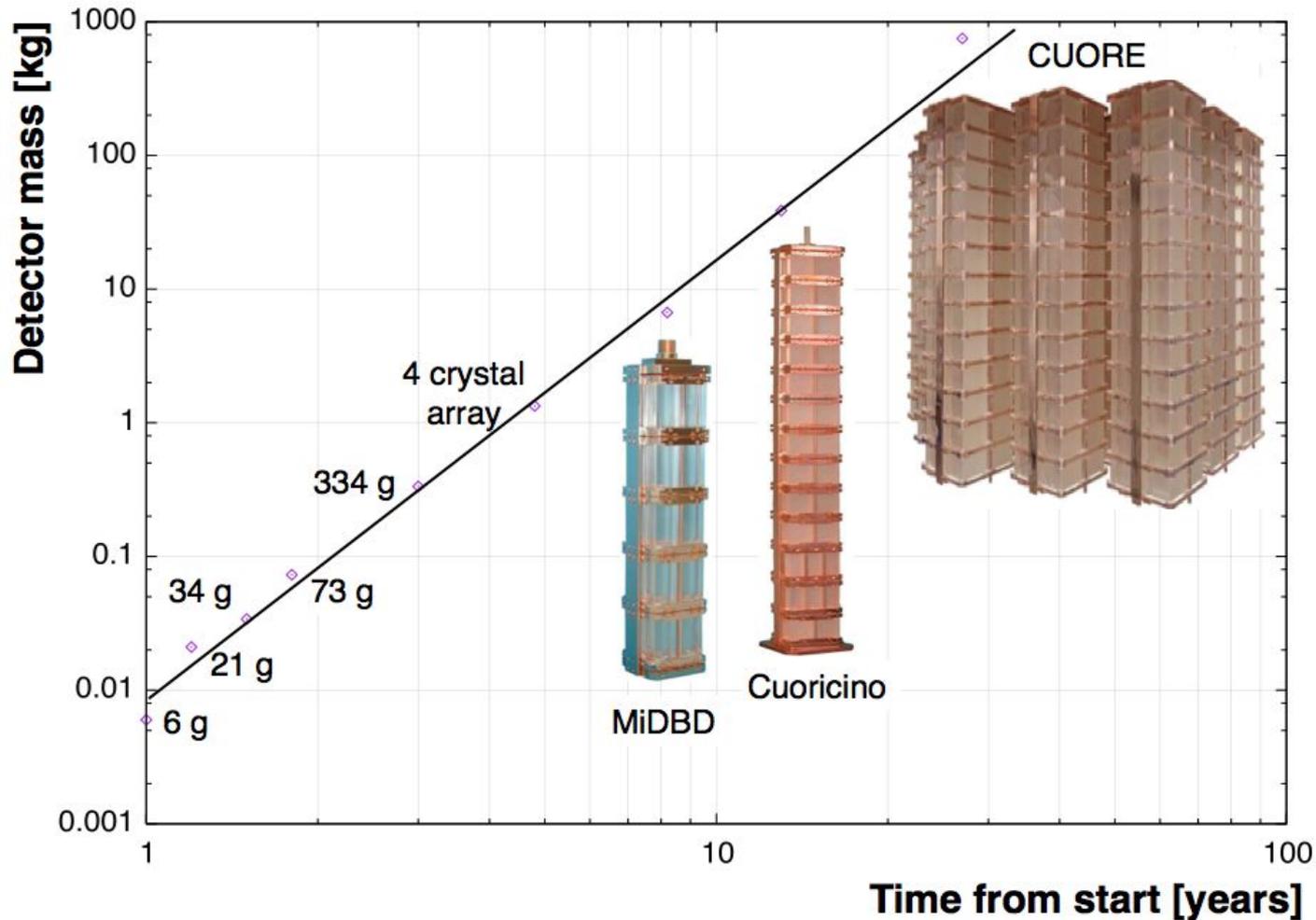


Experiment hosted at
Laboratori Nazionali del Gran Sasso (LNGS)

CUORE ^{130}Te $0\nu\beta\beta$ projected sensitivity:
 $T_{1/2}^{0\nu} \sim 9 \times 10^{25}$ yr (90% C.L.) in 5 years
 $m_{\beta\beta} < 50\text{-}130$ meV

History of TeO_2 macro-calorimeters for $0\nu\beta\beta$

From few g to 1 tonne TeO_2 cryogenic calorimeters for double beta decay search



The CUORE challenge

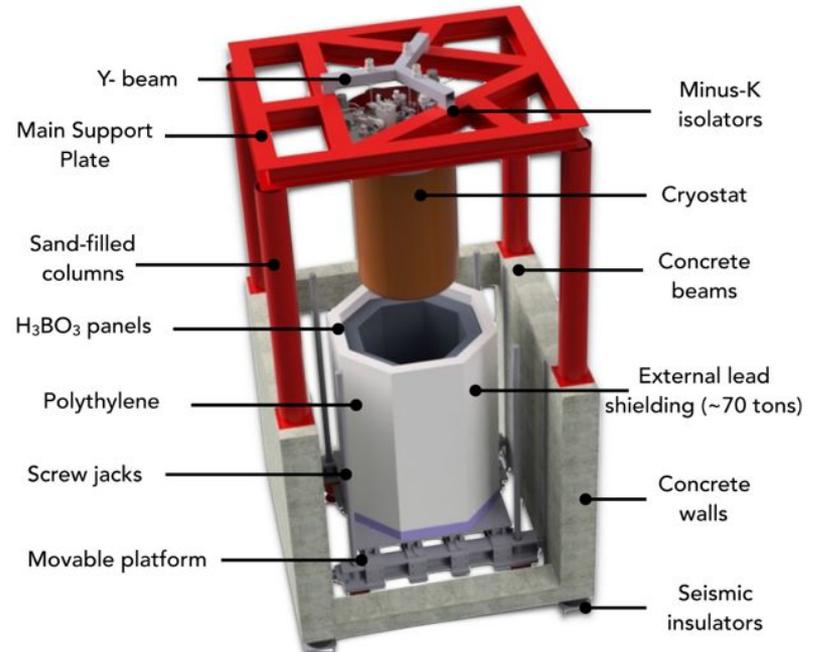
• Low background

- Deep underground location (LNGS)
 - Overburden: 1400 m calcareous rock (3600 m.w.e)
 - Cosmic ray rate reduction: 10^{-6} relative to the surface
- Strict radio-purity controls on materials and assembly
- Passive shields (Pb) from external and cryostat radioactivity
- Detector: high granularity and self-shielding

Background goal: 10^{-2} c/(keV · kg · yr)
in the Region Of Interest (ROI) around $Q_{\beta\beta}$



Roman lead shield
(^{210}Pb depleted)
@ 4K



The CUORE challenge

- **Low temperature and low vibrations**

TeO₂ detectors to be operated as bolometers at ~10 mK

- Multistage cryogen-free cryostat:

Nested vessels at decreasing temperature

Cooling systems: Pulse Tubes (PTs) and Dilution Unit (DU)

- Mass to be cooled < 4K: ~ 15 tons (IVC volume and Cu vessels, Roman Pb shield)

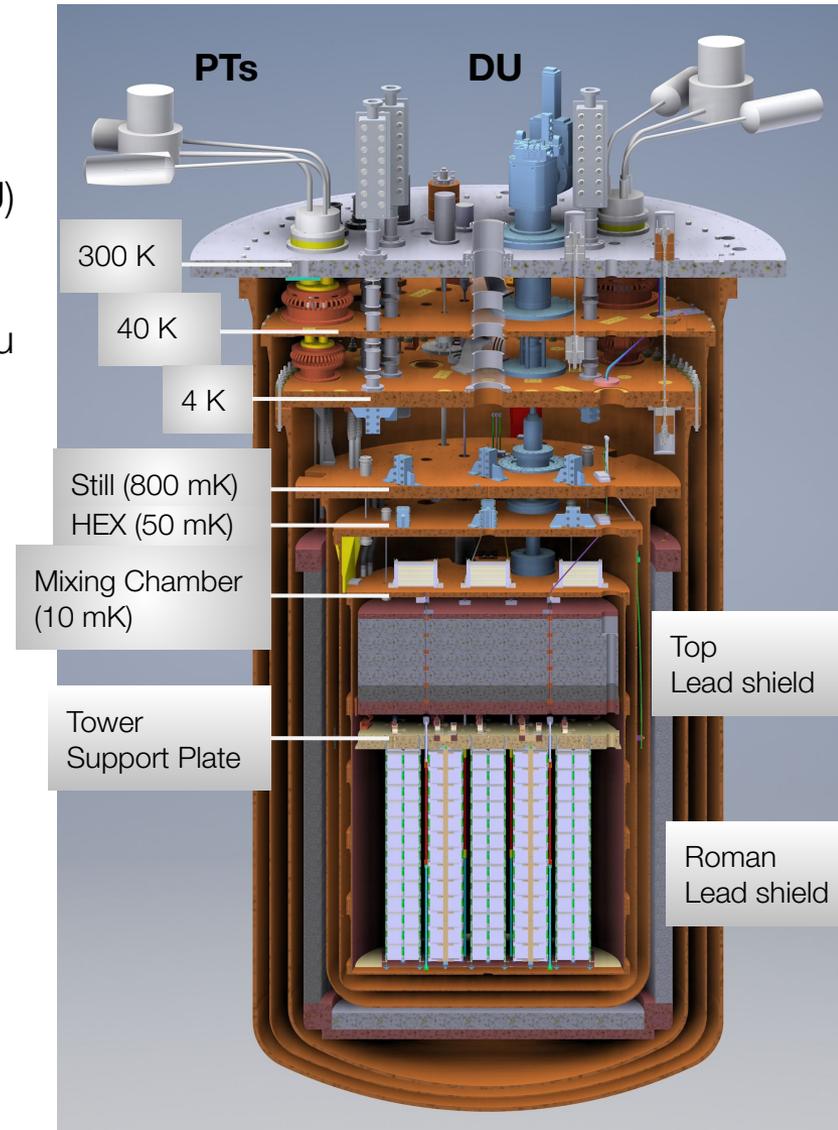
- Mass to be cooled < 50 mK: ~ 3 tons (Top Pb shield, Cu supports and TeO₂ detectors)

- Mechanical vibration isolation

Reduce energy dissipation by vibrations

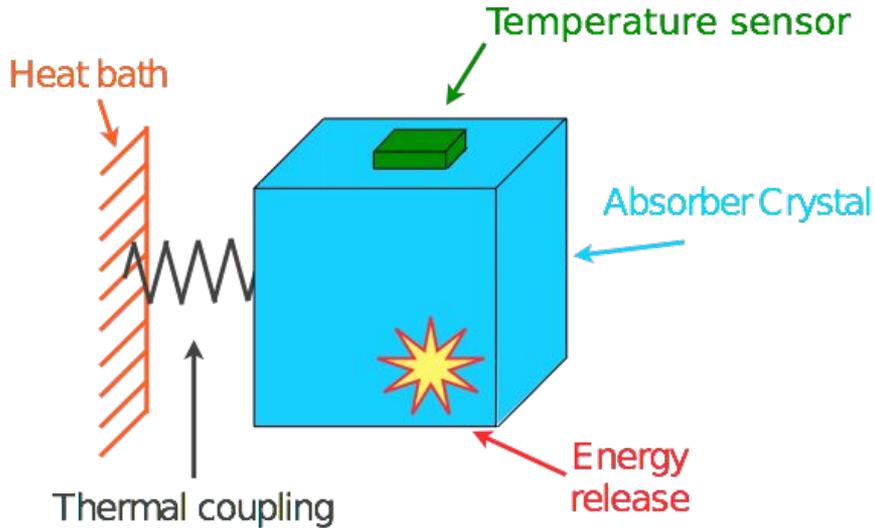
Target energy resolution: 5 keV FWHM

in the Region Of Interest (ROI) around Q_{ββ}



CUORE and the bolometric technique

CUORE TeO_2 detectors are operated as cryogenic bolometers (energy converted into phonons)



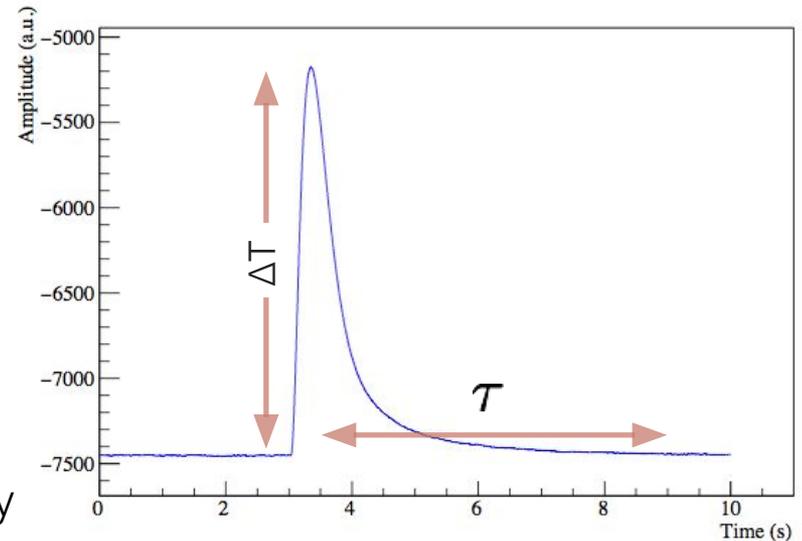
- Low heat capacity @ $T \sim 10$ mK
- Excellent energy resolution ($\sim 0.2\%$ FWHM)
- Same detector response for different particles (heat only)
- Slowness if coupled with NTDs (suitable only for rare event searches)

Simplified thermal model:

One thermal capacity C (crystal)
 One thermal link G (btw crystal/heat bath)

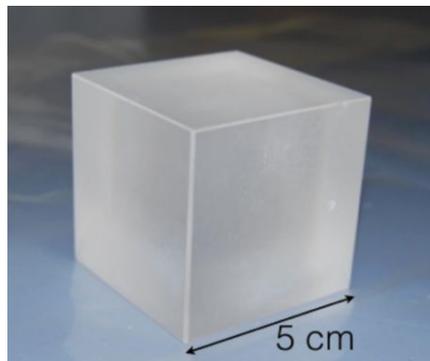
$$\Delta T \propto \frac{E_{dep}}{C} \quad \tau = \frac{C}{G}$$

Amplitude of the pulse $\propto \Delta T \propto$ Energy deposition



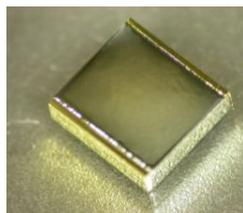
The CUORE detector

CUORE instrumented bolometers

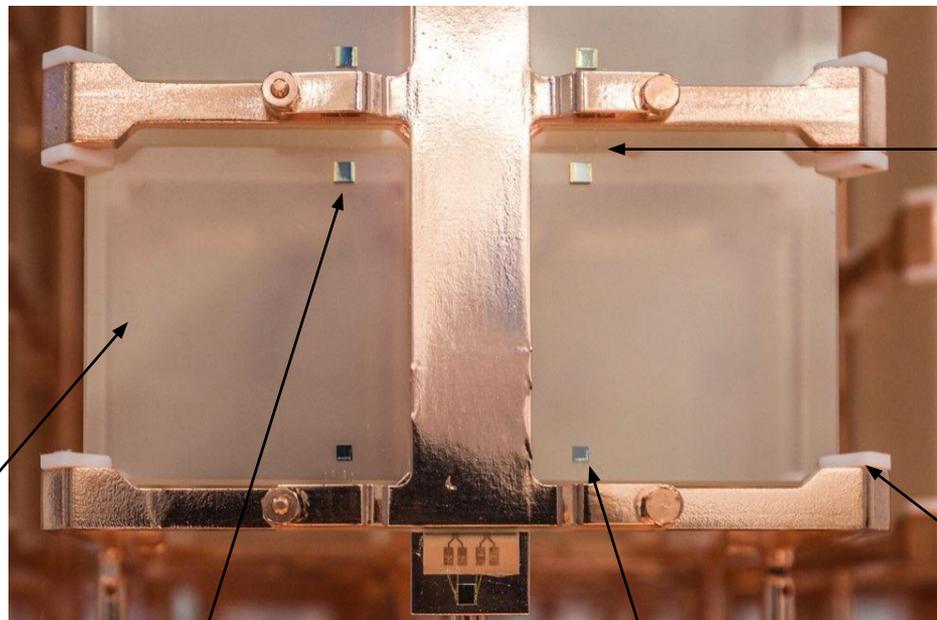


^(nat)TeO₂ crystal
 Absorber = 0νββ source
 5.0 × 5.0 × 5.0 cm³, 750 g mass
 C(T) ~ 2.3 × 10⁻⁹ J/K (@ 10 mK)
 → ΔT_{crystal} ~ 100 μK/MeV
 → τ ~ 0.1 - 1 s

TeO₂ absorber



Ge-NTD thermistor
 3.0 × 2.9 × 0.9 mm³
 Working impedance of the thermistors:
 R_{wp} ~ 100 MΩ - 1 GΩ
 → ΔV_{NTD} ~ 400 μV/MeV (@10 mK)

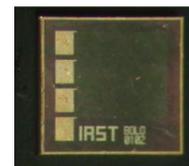


Au-wire bonding to Cu-PEN read-out strips

PTFE holders

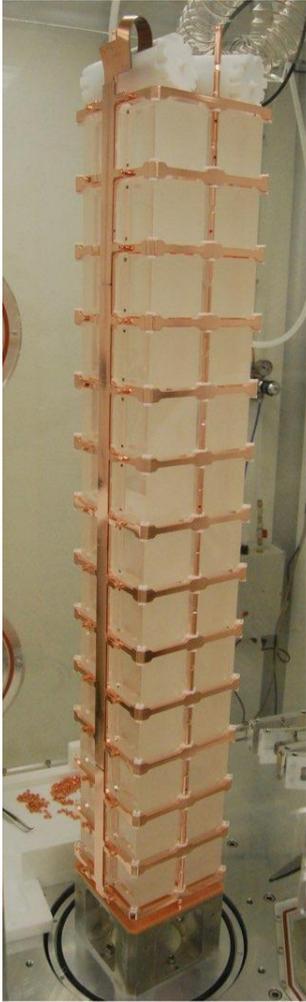
Ge-NTD

Si-heater



Si heater
 2.3 × 2.4 × 0.5 mm³
 Provides periodically a fixed energy for gain stabilization

The CUORE detector



CUORE detector
Array of closely packed 988
TeO₂ crystals
arranged in 19 towers
High Mass of TeO₂:
742 kg (206 kg of ¹³⁰Te)
and high granularity

CUORE tower:

52 crystals
arranged in 13 floors
of 4 crystals each

CUORE-0 experiment:
CUORE-tower demonstrator

CUORE array:

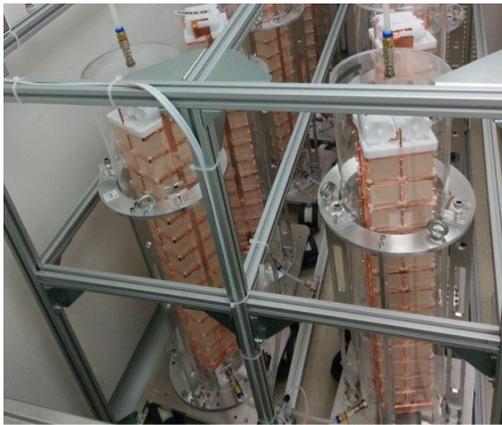
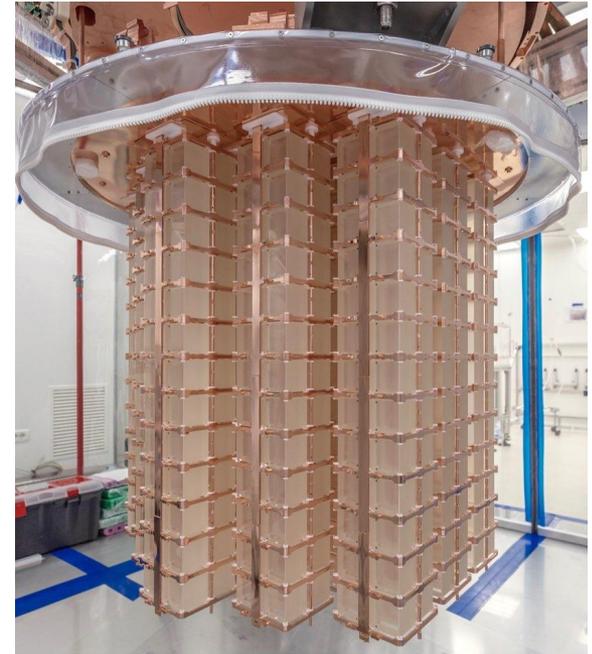
19 towers



CUORE commissioning



- Detector assembly: 2012-2014
- Cryogenic system commissioning: Completed in Feb.2016
- Detector installation: Completed at the end of Aug. 2016



CUORE data taking



From early 2017 up to now

- First operation of a large array of macro bolometers
- Completely different detector and cryogenic system

Initial activities:

- Reach low temperature for system and detectors
- Reduce vibrations
- Optimize detectors thermal response
- Collect first physics data (measure background and release first $0\nu\beta\beta$ results)
- Learn how to make effective data analysis on 1000 bolometers

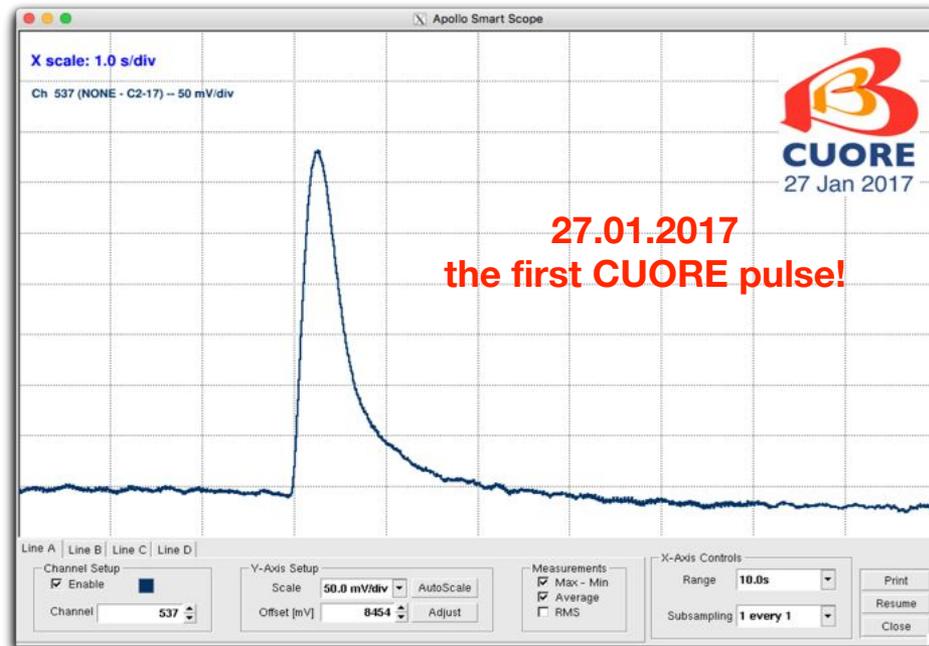
CUORE initial operations



- First Detector cool-down: Started in Dec. 2016
 - Cooldown to 4K: ~ 20 days
 - Pumping He exchange gas in IVC: ~ 10 days
 - Cooldown from 4K to 8 mK: ~ 3 days

Total detector cooldown time from 300K to 10 mK: ~ 1 month

- First CUORE data: Early 2017

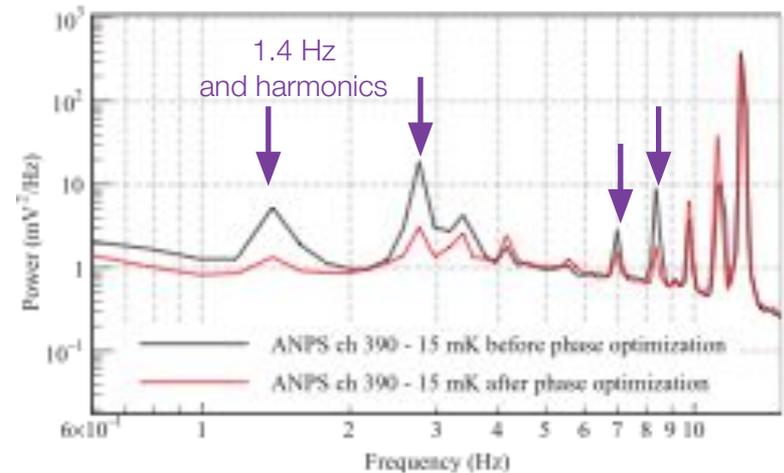
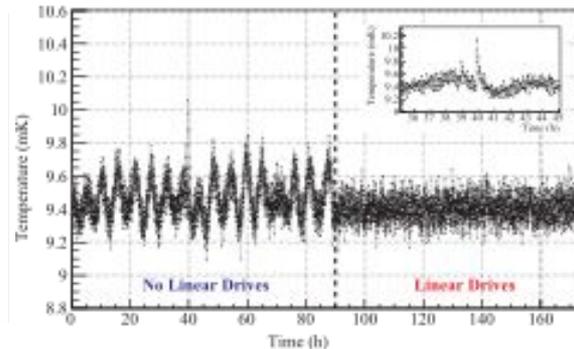
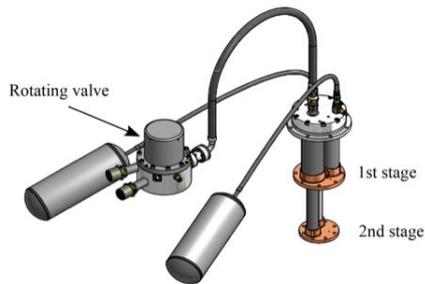


CUORE optimization

Noise reduction

Pulse Tubes induced vibrations: **Pulse Tube active noise cancellation**

- Linear Drives: precise control of the PTs motor-head rotation frequency
- Control the relative phases of the pressure oscillations in the Pulse Tubes and set the detectors minimum noise phase configuration



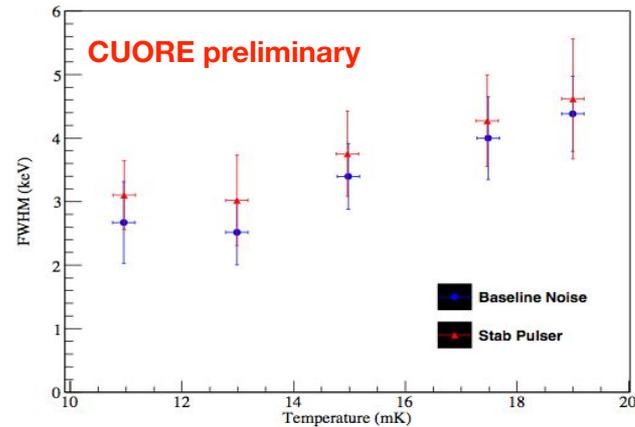
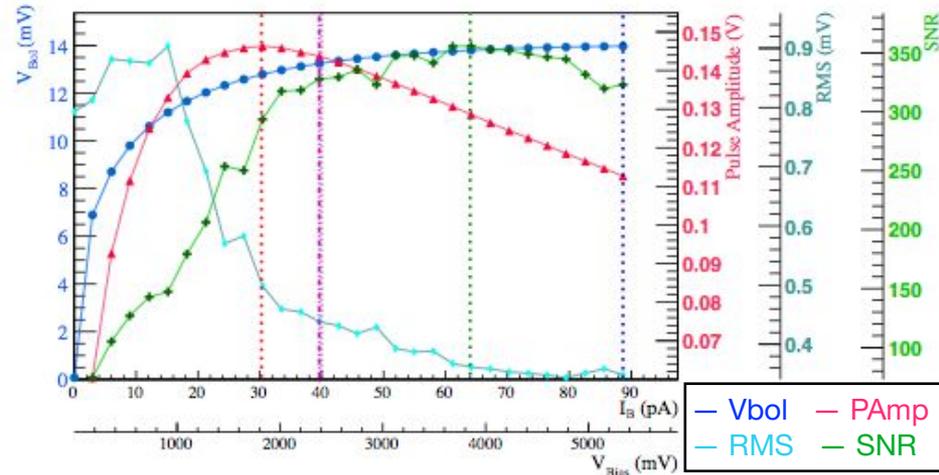
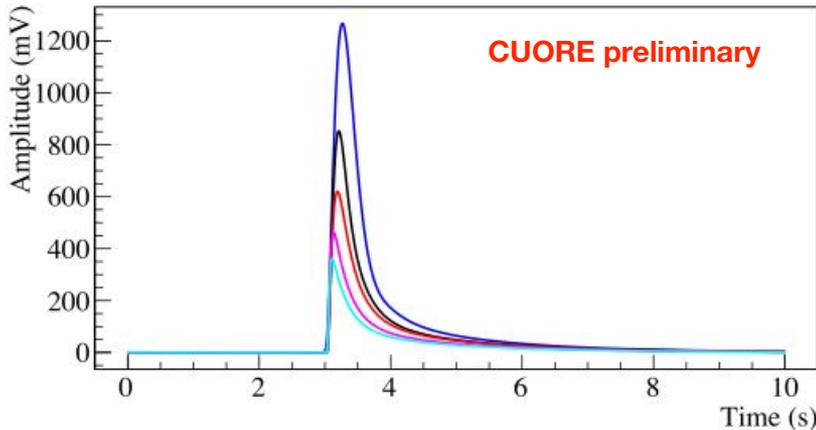
CUORE optimization

Load Curves, Working Points & Temperature scans

Achieve high quality detector readout with a good signal-to-noise ratio

Dedicated procedures and algorithms in CUORE to automate the load curve measurement and the working point identification at each T_{base} .

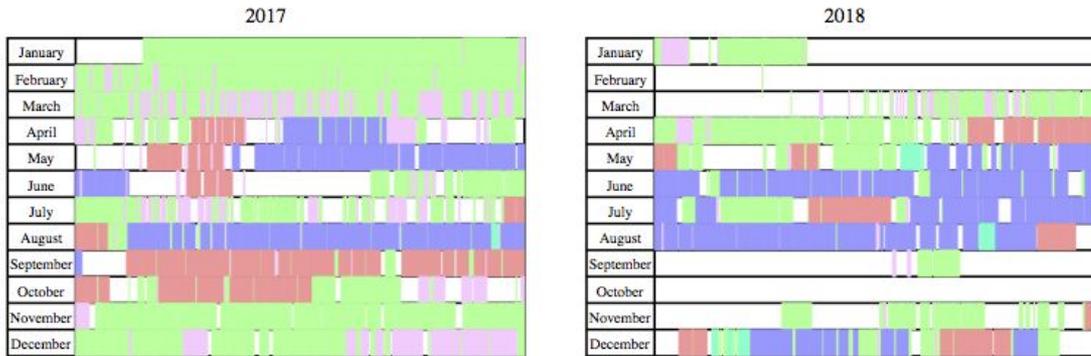
- T 11 mK, rise time (10%-90%) 0.148 s, fall time (90%-30%) 0.303 s
- T 13 mK, rise time (10%-90%) 0.102 s, fall time (90%-30%) 0.340 s
- T 15 mK, rise time (10%-90%) 0.090 s, fall time (90%-30%) 0.408 s
- T 17 mK, rise time (10%-90%) 0.055 s, fall time (90%-30%) 0.407 s
- T 19 mK, rise time (10%-90%) 0.039 s, fall time (90%-30%) 0.422 s



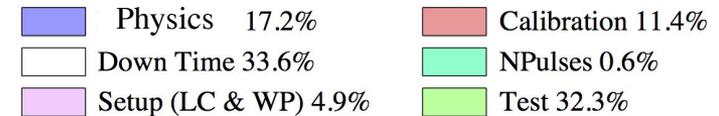
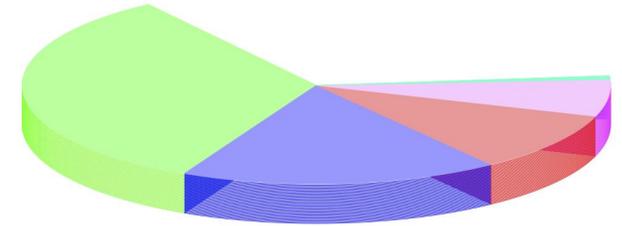
CUORE data taking



Data taking up to March 2019



CUORE Run Time Breakdown
Jan 2017 - Feb 2019



The duty cycle of the system was poor, dominated by down-time (short warm-ups and cool-downs) and technical runs.

We concentrated our efforts on improving the reliability of the overall system and the stability of the data-taking.

- Performed a major maintenance of the cryogenic system (early 2019)
- Performed an upgrade of the calibration system (2018)
- Improved the data processing and analysis techniques
- Focussed on physics data-taking

CUORE data taking: improvements

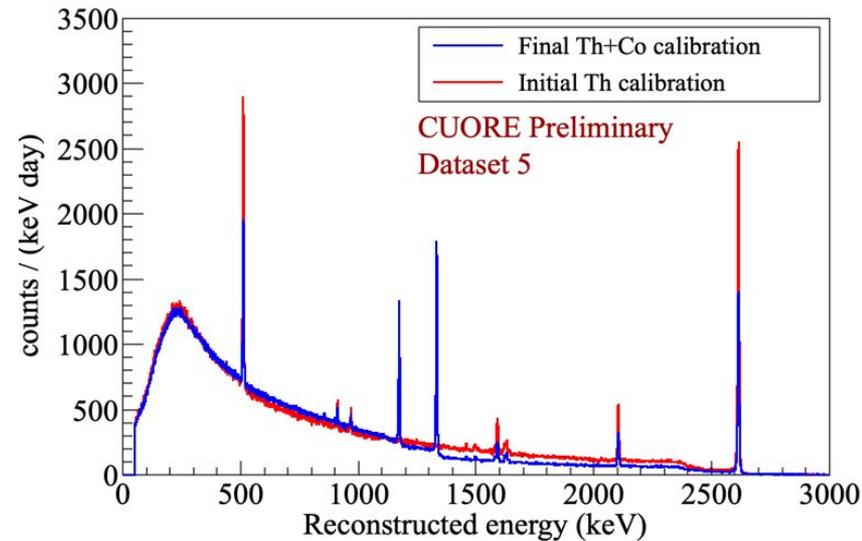
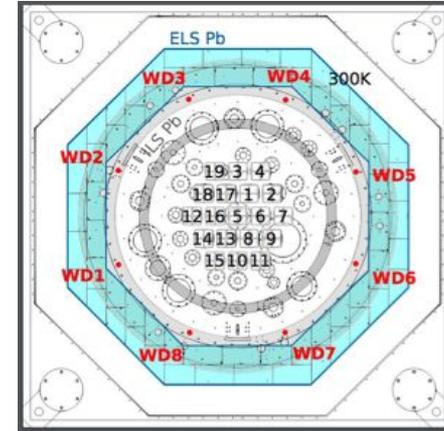


External Detector Calibration System (EDCS)

Internal Detector Calibration System: strings of ^{232}Th γ -ray sources reaching the detectors down to 10 mK, complex procedure for deployment and interference with the cryogenic system operation

→ Installation of External Detector Calibration System (EDCS)

EDCS: $^{232}\text{Th} + ^{60}\text{Co}$ γ -ray sources, strings positioned in between the 300K vessel and the external Lead Shield



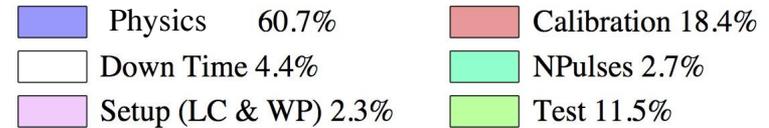
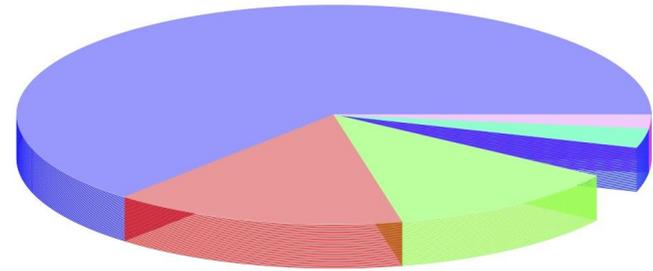
CUORE data taking



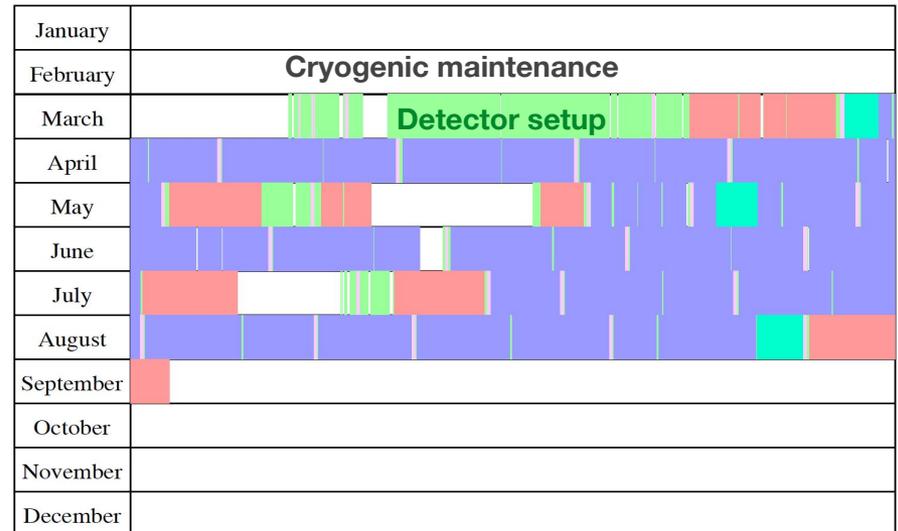
Data taking from March 2019 on

- **Reached stable data-taking with high duty cycle**
- Physics data bracketed by initial and final calibrations
- Calibration with EDCS
- Only short maintenance operations at dataset closure; checks for noise and thermal response consistency after the activities
- 3 complete consecutive datasets acquired, new dataset started in these days
- Priority is data taking - no further optimizations in the near future

CUORE Run Time Breakdown from March 2019



2019



CUORE data taking

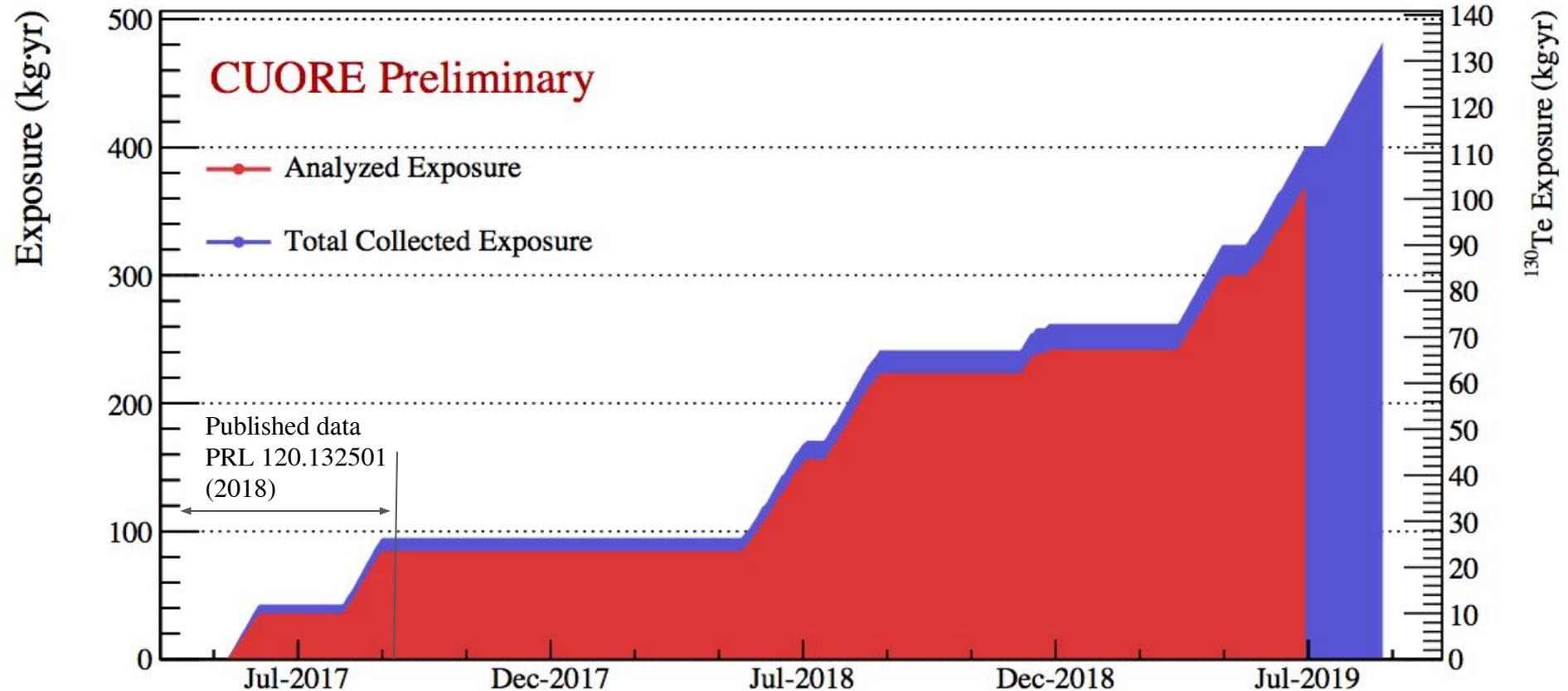


Current CUORE exposure

Collected Background exposure (TeO_2): 400.9 kg yr

Analyzed Background exposure (TeO_2): 369.9 kg yr

^{130}Te exposure: 102.9 kg yr



CUORE first results

CUORE published results - Data from 2017

984/988 operational channels

7 weeks of physics data taking (Dataset 1, Dataset 2)

$^{nat}\text{TeO}_2$ exposure: 86.3 kg yr,

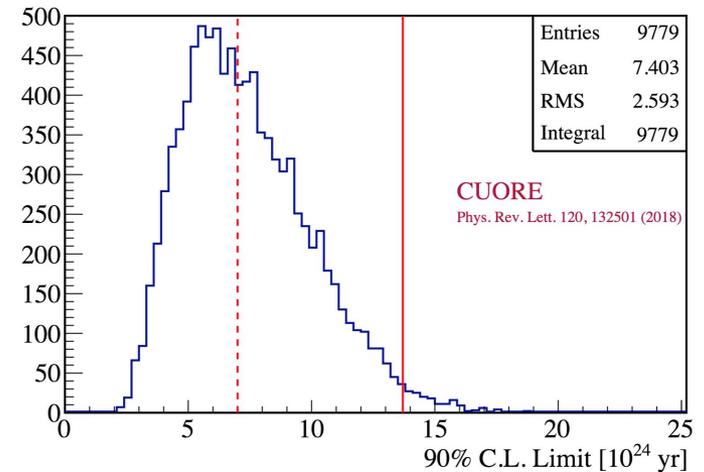
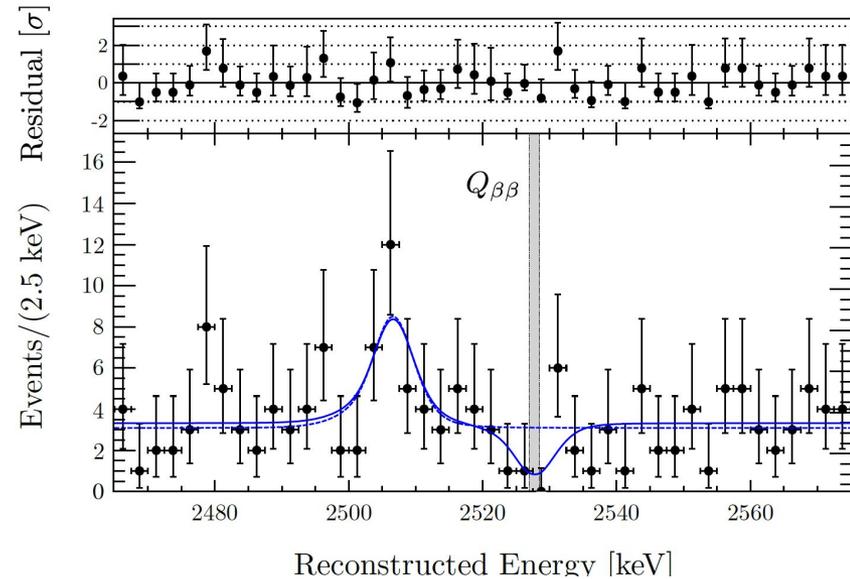
^{130}Te exposure: 24.0 kg yr

ROI background index (B) $\sim 1.4 \times 10^{-2}$ c/(keV · kg · yr)

UEML fit performed in the ROI [2465,2575 keV]
to evaluate best fit decay rate for $0\nu\beta\beta$ of ^{130}Te

No evidence of signal

Half-life limit for $0\nu\beta\beta$ in ^{130}Te (90%C.L including syst.)
 $T_{0\nu}$ (^{130}Te) $> 1.3 \times 10^{25}$ yr



CUORE trigger algorithm

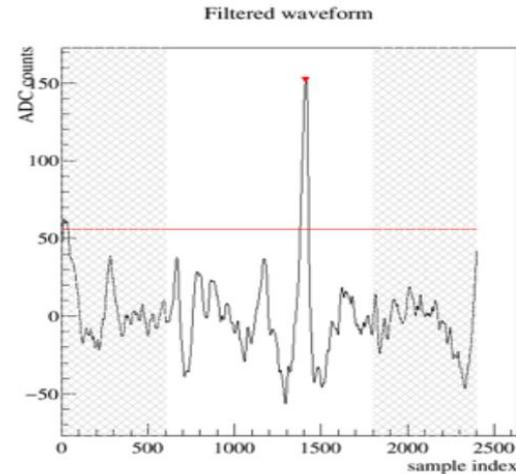
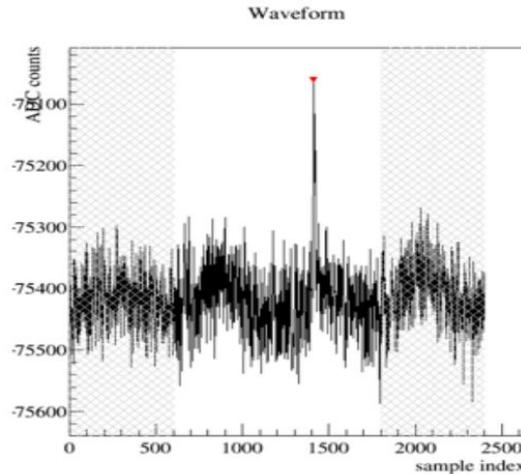
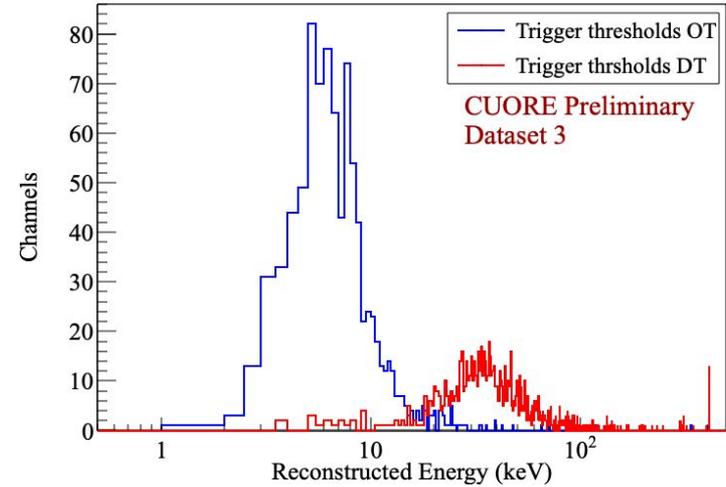


OT trigger

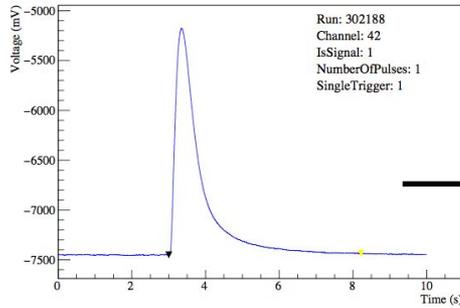
Disentangle low energy signals from fake signals produced by noise, lower the detectors trigger thresholds

→ Optimum Trigger (OT) algorithm: identifies a signal when the amplitude of the filtered signal waveform exceeds a configurable threshold

- OT trigger applied for offline re-triggering of the continuous recorded stream of data



CUORE data processing

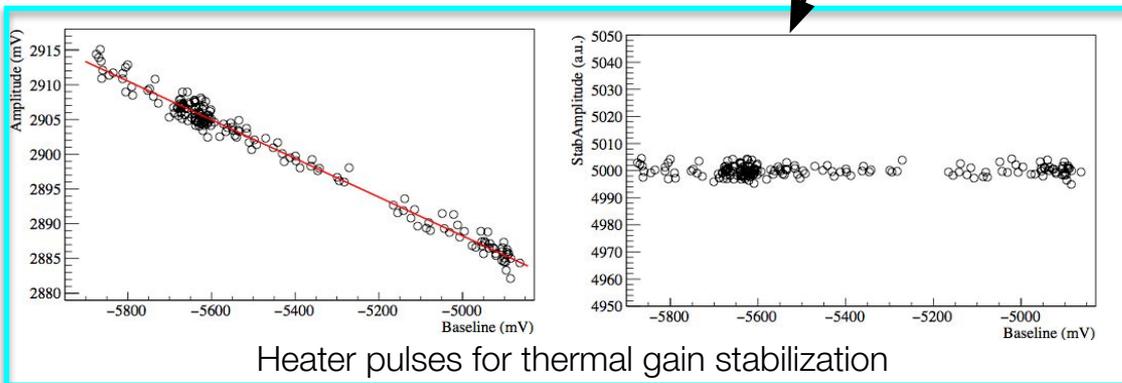
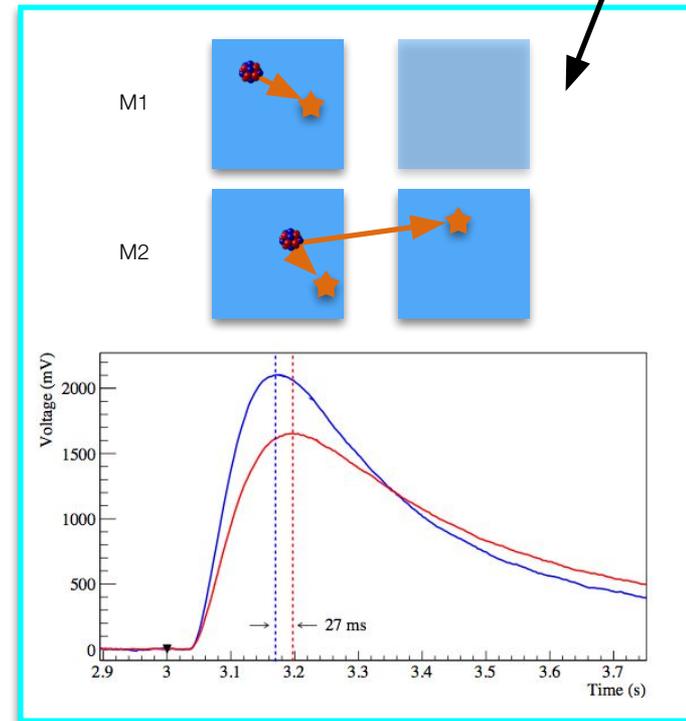
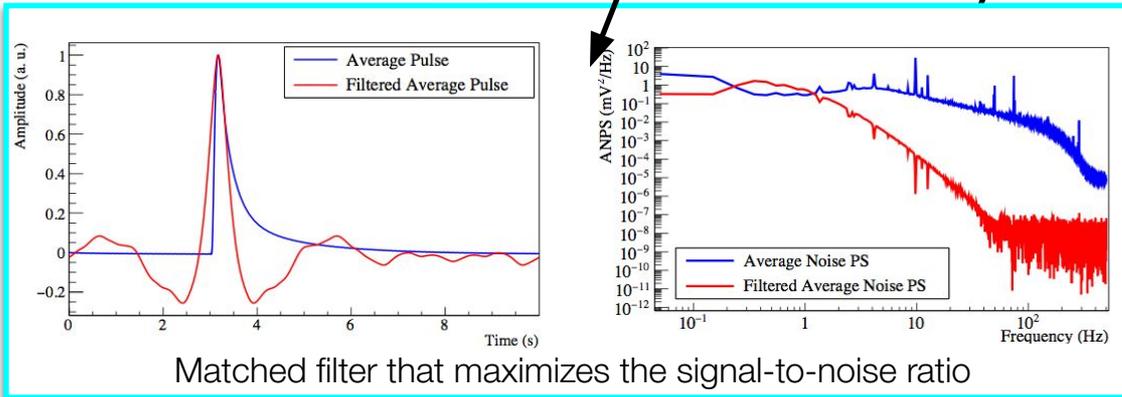


Optimum Filter

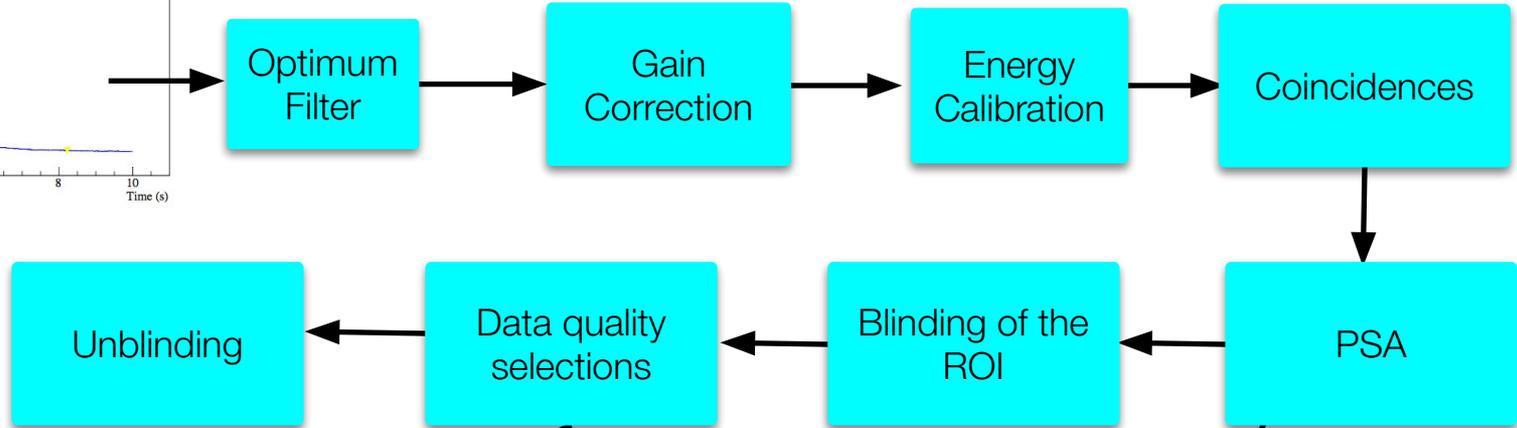
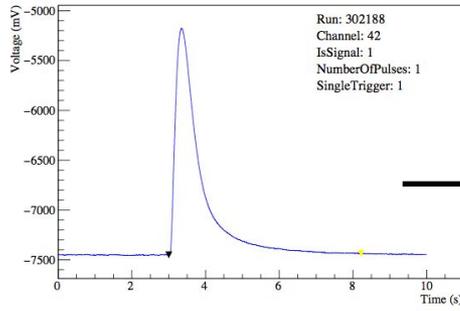
Gain Correction

Energy Calibration

Coincidences

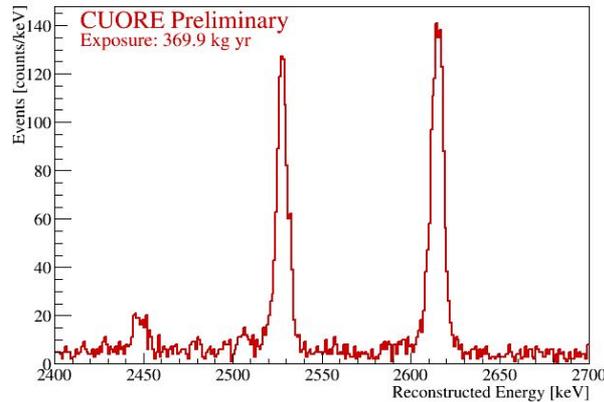


CUORE data processing

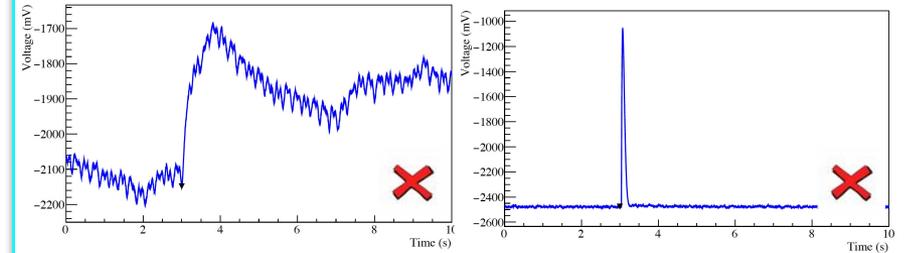


- Optimization of data quality cuts on Blinded data

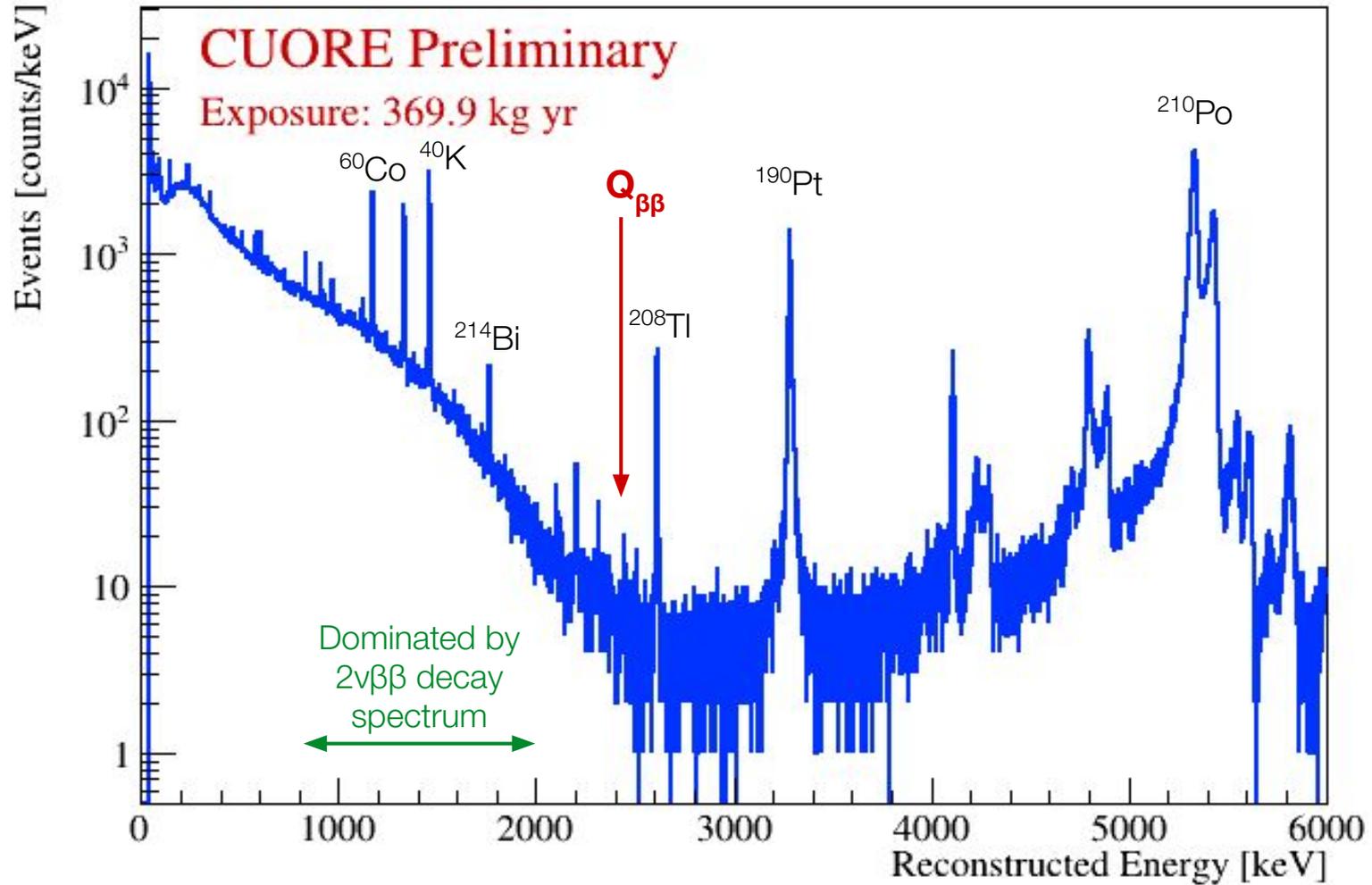
Summed Spectrum (Blinded)



Reject non-physical pulses



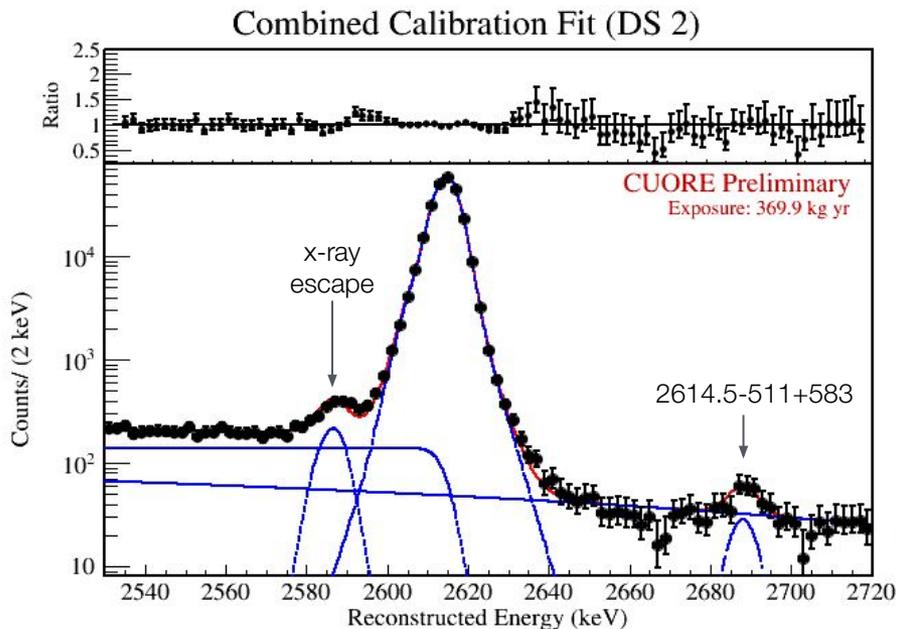
Summed Spectrum



Lineshape

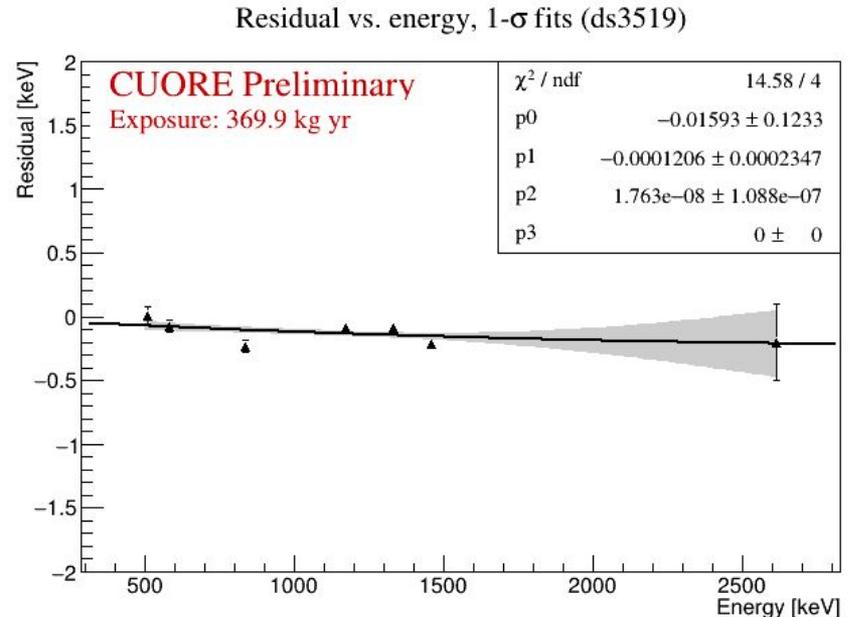
Lineshape fit

- Fit ^{208}Tl line at 2614.5 keV in calibration spectra to precisely evaluate lineshape
- Main peak parameterized with sum of 3 Gaussian
- Fit run on single tower simultaneously for multiple datasets



Bias of peak position in physics data

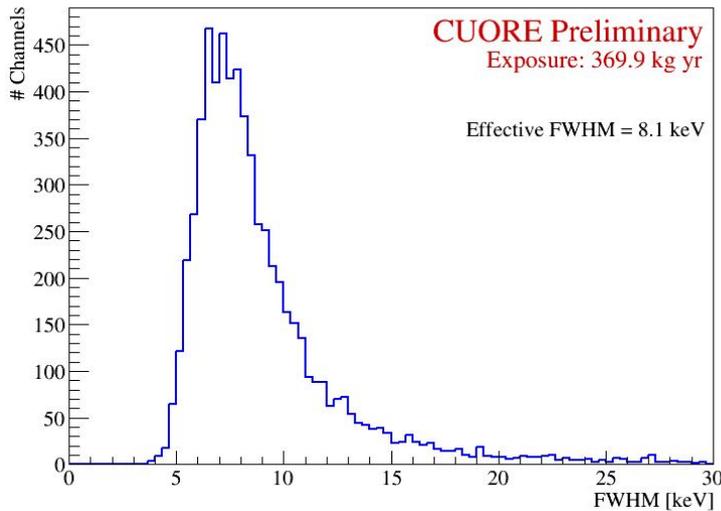
- Fit peaks in physics spectrum with 3-Gaussian
- Compute residuals from literature values
- Fit residuals vs energy with 2nd order polynomial → Systematics!



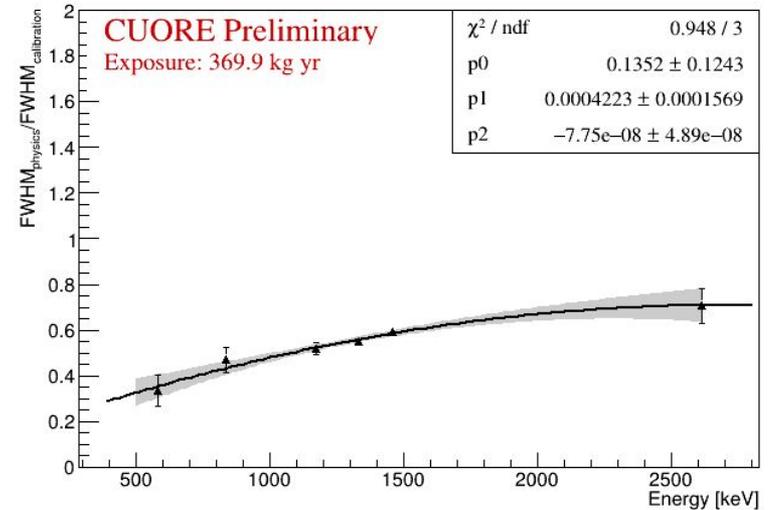
Scaling of resolution to physics data

- Extract FWHM from peak fit in physics spectra
- Fit ratio of FWHM in physics and calibration data as a function of energy
- Extrapolate to $Q_{\beta\beta}$

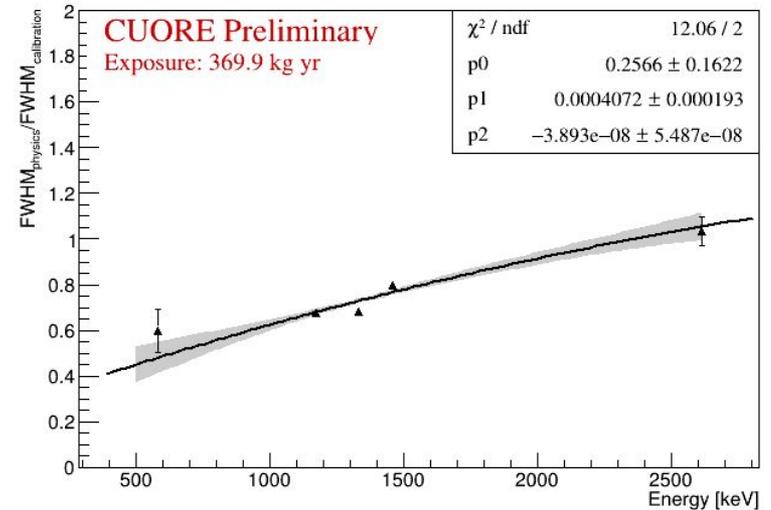
Calibration Resolution at 2615 keV



Background resolution vs. energy, 1- σ fits (ds3519)



Background resolution vs. energy, 1- σ fits (ds3567)



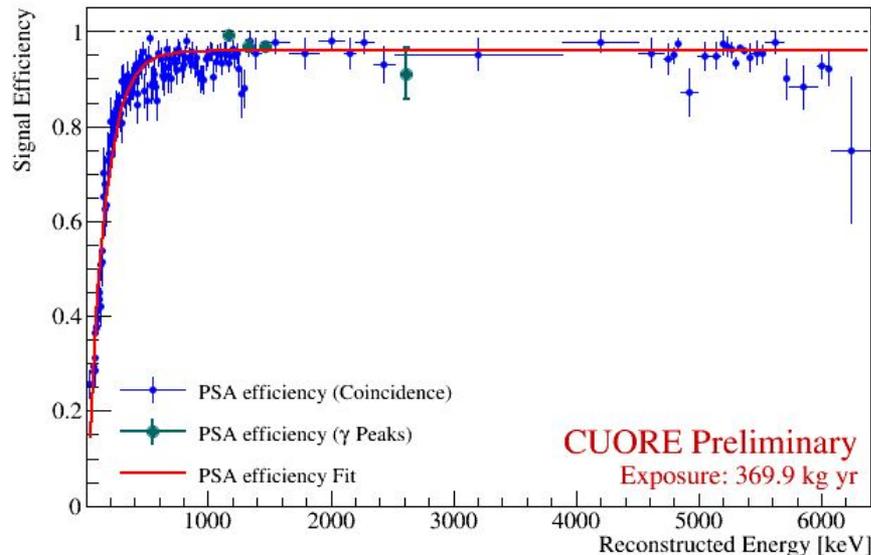
Efficiencies

Reconstruction efficiency Comprises:
→ trigger
→ event reconstruction
→ pile-up identification

Anti-coincidence efficiency Quantifies the probability of properly identifying a single-site event

Pulse-Shape Analysis efficiency Fraction of events passing a multi-dimensional cut on 6 pulse-shape variables

PSA Efficiency (DS 2)



Parameters for the analysis



Parameter	Value
Number of datasets	7
Number of channels	842-948 (depending on dataset)
TeO₂ exposure	369.94 kg·yr
FWHM at 2615 keV in calibration data	8.1 keV
FWHM at Q_{ββ} in physics data	8.7 keV
Reconstruction efficiency	(95.9578 ± 0.0033)%
Anti-coincidence efficiency	(98.95 ^{+0.15} _{-0.16})%
PSA efficiency	(92.04 ± 0.11)%
Tot. analysis efficiency	(87.41 ± 0.18)%
Syst. on analysis efficiency	±1.9%
Containment efficiency	(88.350 ± 0.090)%

Statistical approach

Bayes theorem:
$$P(\vec{\theta}|\vec{E}) = \frac{\mathcal{L}(\vec{E}|\vec{\theta}) \cdot \pi(\vec{\theta})}{\int_{\Omega} \mathcal{L}(\vec{E}|\vec{\theta}) \cdot \pi(\vec{\theta}) d\vec{\theta}}$$

Likelihood:
$$\mathcal{L}(\vec{E}|\vec{\theta}) = \prod_{\text{dataset}} \prod_{\text{channel}} \left[\frac{e^{-\lambda} \lambda^n}{n!} \prod_{\text{event } i} \left(\frac{s}{\lambda} \text{pdf}_{0\nu\beta\beta}(E_i|\vec{\theta}) + \frac{c}{\lambda} \text{pdf}_{60Co}(E_i|\vec{\theta}) + \frac{b}{\lambda} \frac{1}{\Delta E} \right) \right]$$

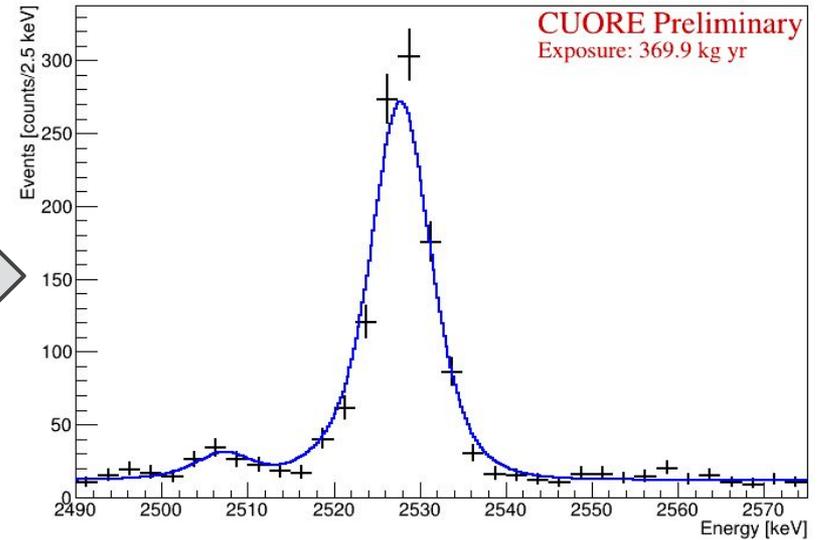
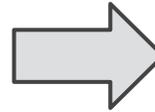
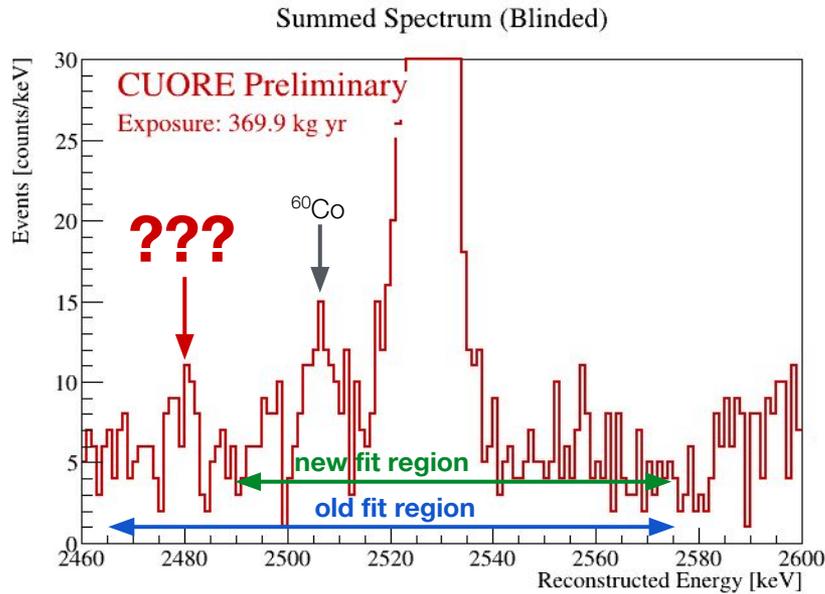
Expectation value: $\lambda=s+c+b$

Width of fit region: ΔE

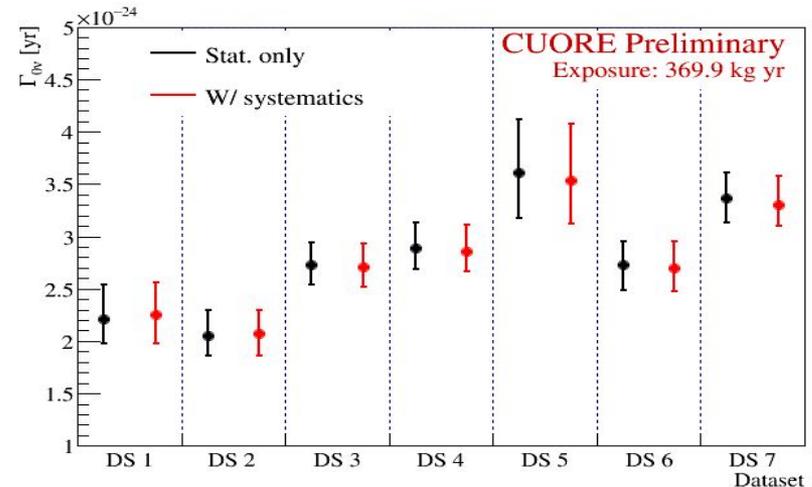
Systematics implemented as nuisance parameters

Parameter	Dependence	Method
Analysis efficiency I	Dataset	Gaussian
Analysis efficiency II	Global	Flat in [0.981, 1.021] range
Energy bias	Dataset	Fit residuals of peaks in physics spectrum from literature values with 2 nd order polynomial
Energy resolution	Dataset	Fit ratio of FWHM in physics and calibration data with 2 nd order polynomial
Q_{ββ}	Global	Gaussian, 2527.518(13) keV
¹³⁰Te isotopic fraction	Global	Gaussian, 34.1668(16)%

Blinded fit

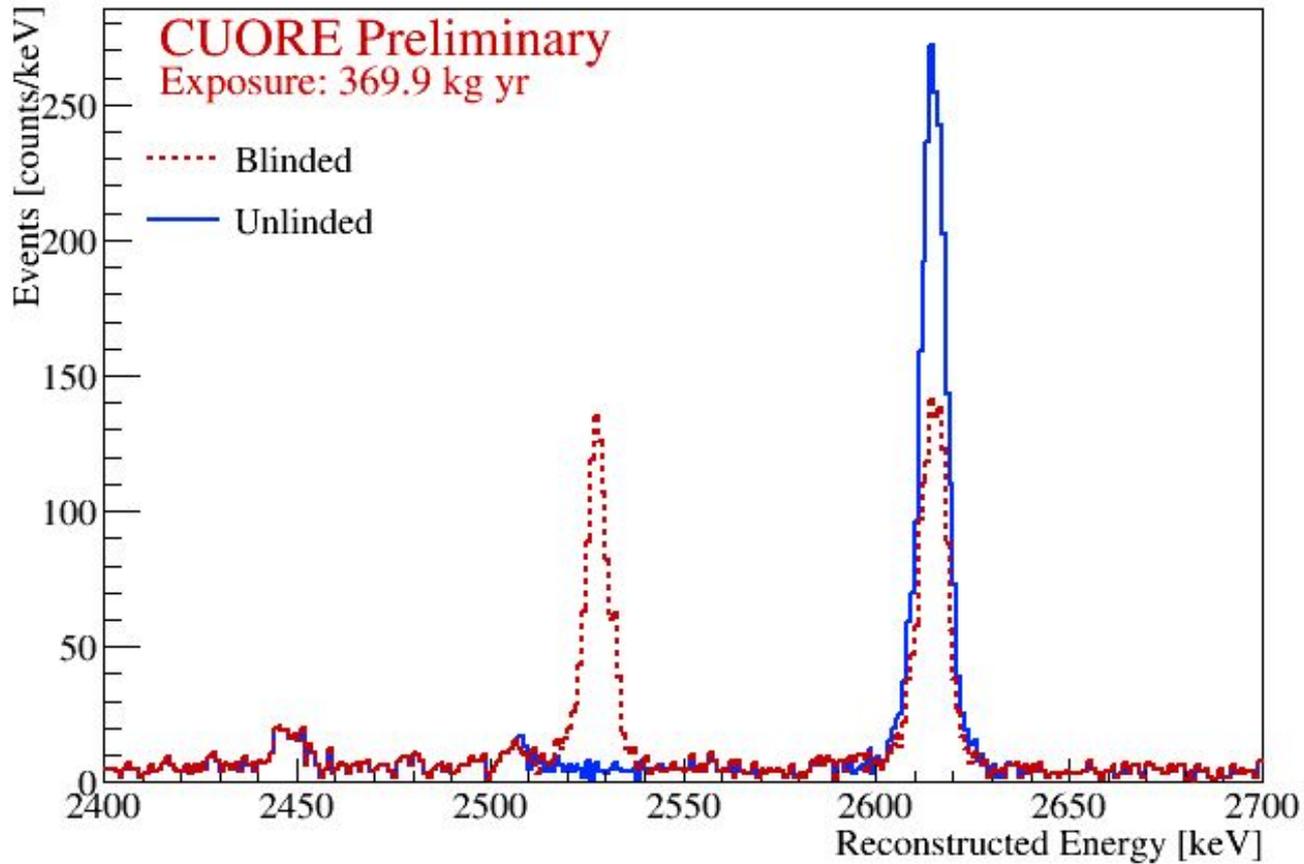


- Fit range: [2490, 2575] keV
- Components: flat background, ^{60}Co peak, $0\nu\beta\beta$ peak
- Fit run independently for each dataset and tested w/ and w/o additional nuisance parameters for systematics
→ $0\nu\beta\beta$ consistent in both cases for all datasets



Unblinding

Summed Spectrum



Background in α region and at $Q_{\beta\beta}$

α region

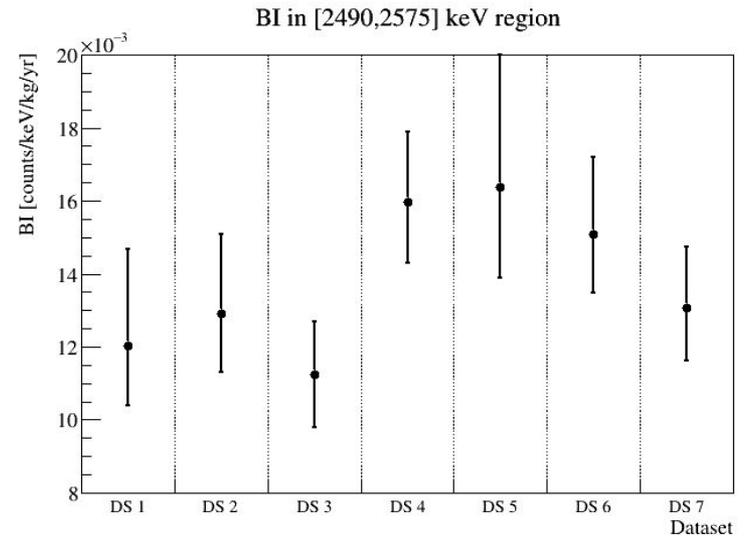
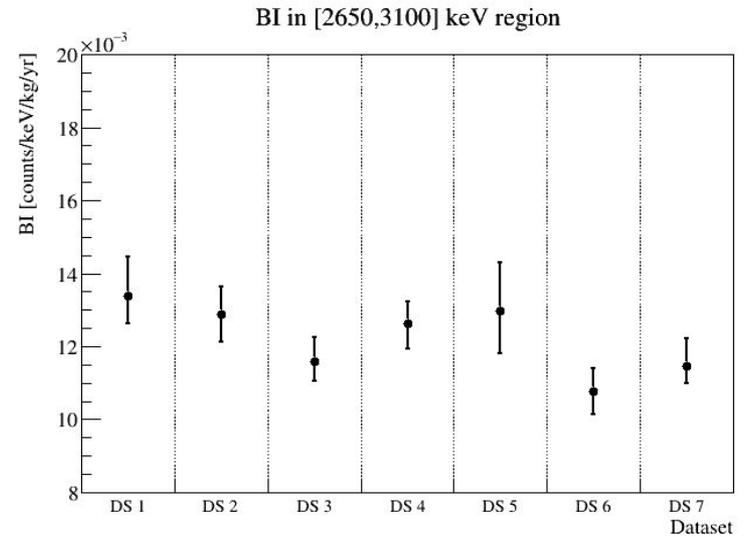
- Fit flat background in [2650,3100] keV region
- Average α background: $1.210(28) \cdot 10^{-2}$

$Q_{\beta\beta}$ region

- Fit with flat background + ^{60}Co peak in [2490,2575] keV region
- Average BI: $1.369(69) \cdot 10^{-2}$ cts/keV/kg/yr

Dataset **BI [$\times 10^{-2}$ counts/keV/kg/yr]**

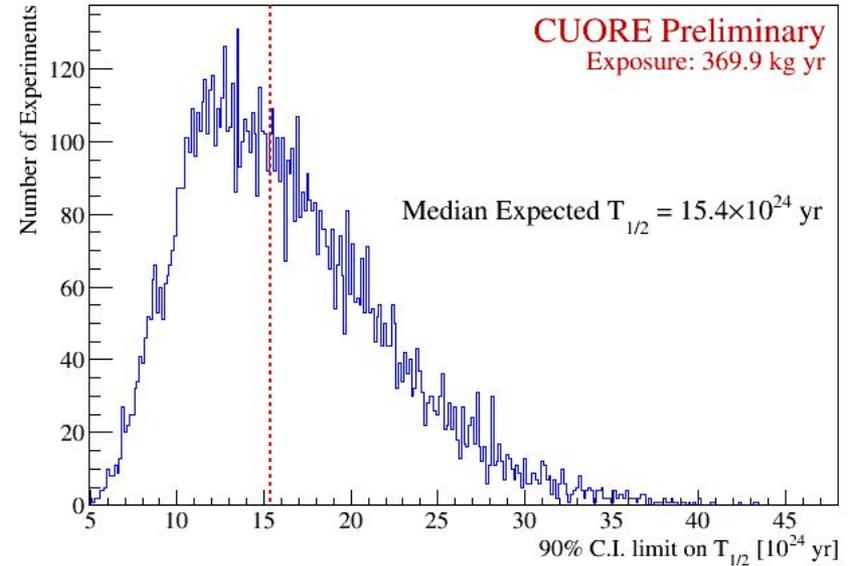
1	$1.20^{+0.26}_{-0.17}$
2	$1.29^{+0.22}_{-0.16}$
3	$1.11^{+0.15}_{0.14}$
4	$1.60^{+0.19}_{-0.17}$
5	$1.64^{+0.44}_{-0.25}$
6	$1.51^{-0.21}_{-0.16}$
7	$1.13^{+0.17}_{-0.14}$



Method

- Fit unblinded data with flat background and ^{60}Co components only (no $0\nu\beta\beta$)
- Generate set of 10^4 toy-MC data-sets according to bkg-only model
- Fit toy-MC with signal+bkg model
- Extract distribution of 90% C.I. halflives limits

Projected Sensitivity



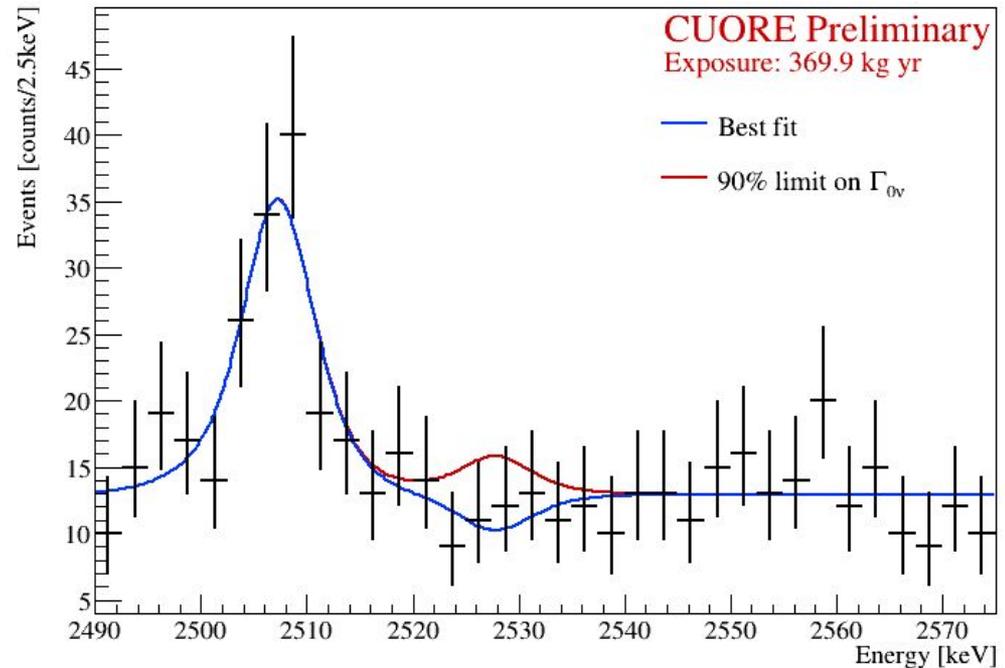
Details for NERDs

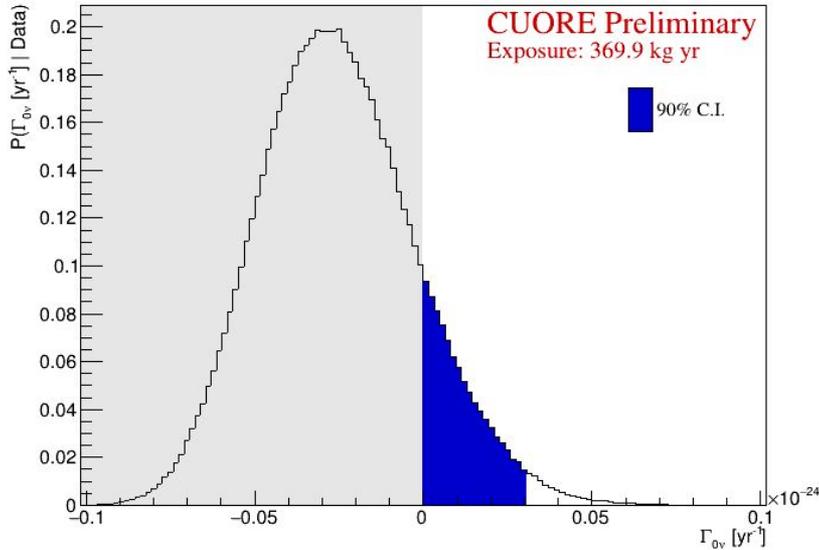
- Posterior for background index (B) and ^{60}Co rate from bkg-only fit have uncertainties larger than corresponding Poisson uncertainty
→ Sample from enlarged distribution: $\text{TRandom3}::\text{PoissonD}(B^2/\sigma^2) \cdot \sigma^2/B$
- ^{60}Co peak generated using same 3-Gaussian lineshape used for the fit

Unblinded fit

Method

- Bayesian analysis based on [BAT](#)
- Allow negative non-physical range for $\Gamma_{0\nu}$ to evaluate the amplitude of possible background under-fluctuations
- Repeat fit on physical range only
→ Results on $\Gamma_{0\nu}$ obtained from this!
- Free params: ^{60}Co peak rate & position, $\Gamma_{0\nu}$ rate, background
- Repeat fit with additional nuisance parameters to account for systematics

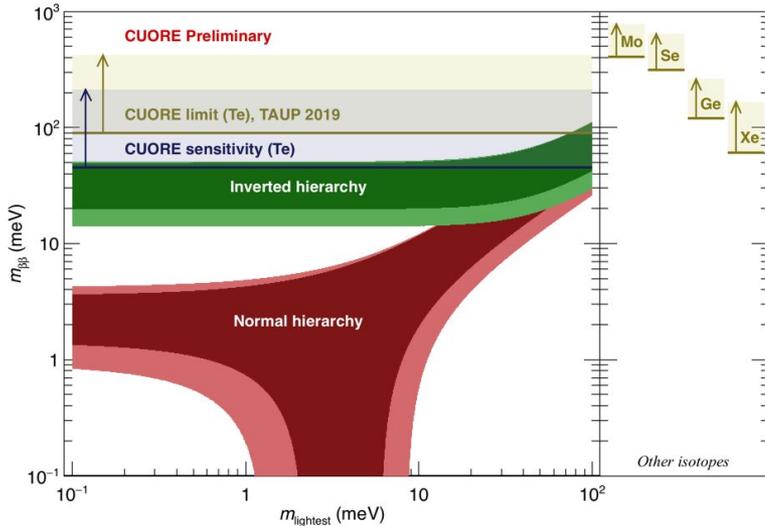




PRELIMINARY Results

- No evidence for $0\nu\beta\beta$ decay 😞
- Background under-fluctuation yields a best-fit value of:
 $\Gamma_{0\nu} = -3.0^{+2.8}_{-1.8} \cdot 10^{-26} \text{ yr}^{-1}$
- Marginalized limit computed on physical range: $\Gamma_{0\nu} < 3.0 \cdot 10^{-26} \text{ yr}^{-1}$

$$T_{1/2}^{0\nu} > 2.3 \cdot 10^{25} \text{ yr at 90\% C.I.}$$



- Systematics affect the limit by $\sim 1\%$
- Probability of getting a stronger limit: 13%
- Assuming the light neutrino exchange:

$$m\beta\beta < 0.09-0.42 \text{ eV at 90\% C.I.}$$

- Starting to work on a publication

Conclusions

CUORE is the first tonne-scale operating bolometric $0\nu\beta\beta$ detector.

- New CUORE physics results of $T_{0\nu}$ in ^{130}Te with increased exposure (370 kg yr TeO_2)
- The CUORE data taking is currently underway to collect 5 years of run time
- Investigating the potential of the CUORE experiment for the search for rare events and/or for physics beyond the Standard Model other than the $2\nu\beta\beta$ decay of ^{130}Te

Thank you on behalf of The CUORE collaboration



Yale



CAL POLY
SAN LUIS OBISPO



UCLA



SAPIENZA
UNIVERSITÀ DI ROMA



Backup

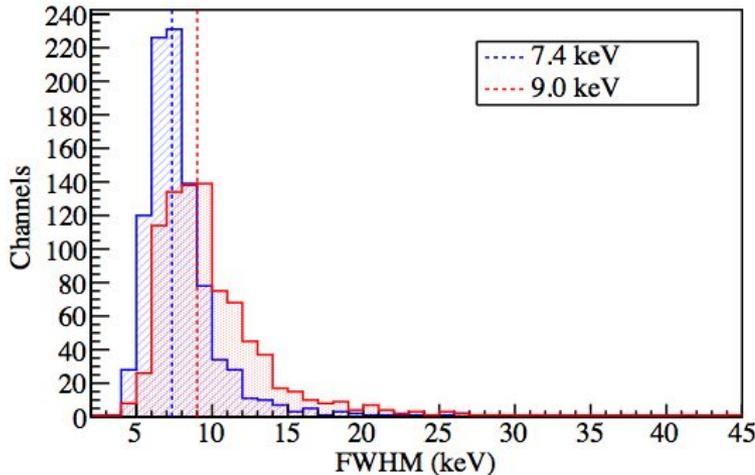
CUORE optimization & first physics data

Improved Detector performance during optimization in 2017

- Energy resolution in calibration runs @²⁰⁸Tl decay gamma-peak (2615 keV)

Dataset 1: 9.0 keV FWHM

Dataset 2: 7.4 keV FWHM



Improved resolution from Dataset 1 to Dataset 2 due to:

- Investigation and upgrades to the electronics grounding
- **Active cancellation of the PT-induced noise**
- **Optimization of the operating temperature and detector working points**
- Software and analysis upgrades

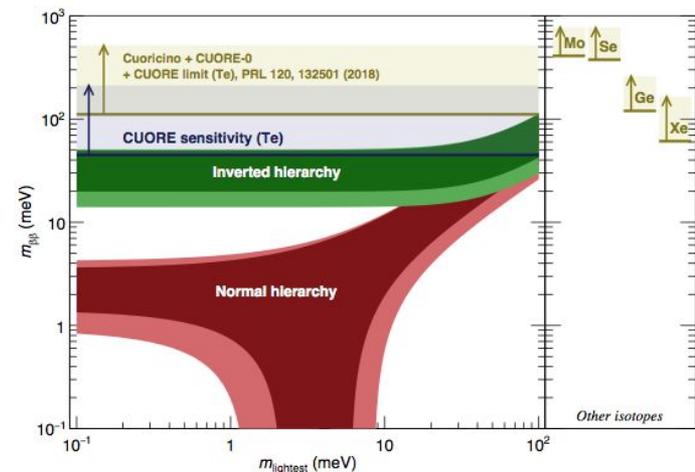
First CUORE physics data

984/988 operational channels

7 weeks of physics data taking (Dataset 1, Dataset 2)

^{nat}TeO₂ exposure: 86.3 kg yr, ¹³⁰Te exposure: 24.0 kg yr

ROI background index (B) $\sim 1.4 \times 10^{-2}$ c/(keV · kg · yr)

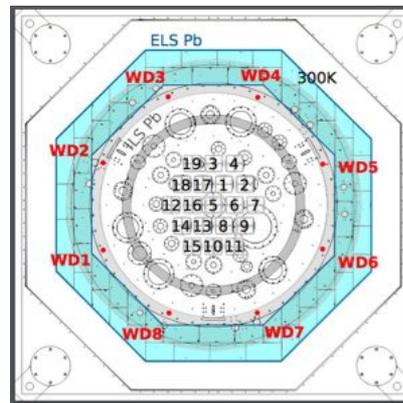
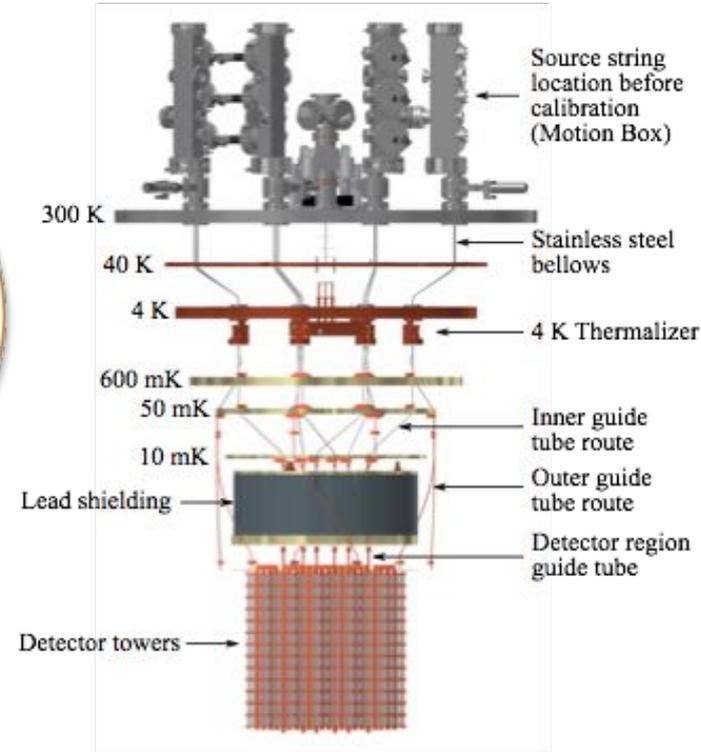
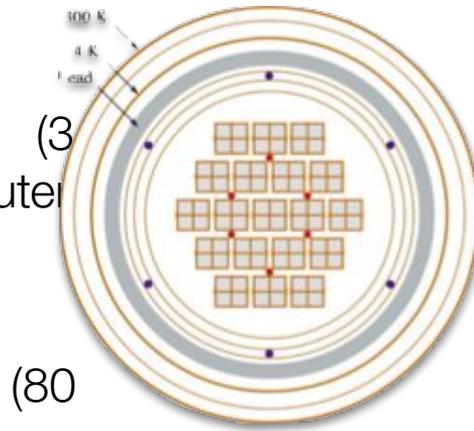


CUORE auxiliary systems

Detector Calibration Systems

^{232}Th γ -ray sources

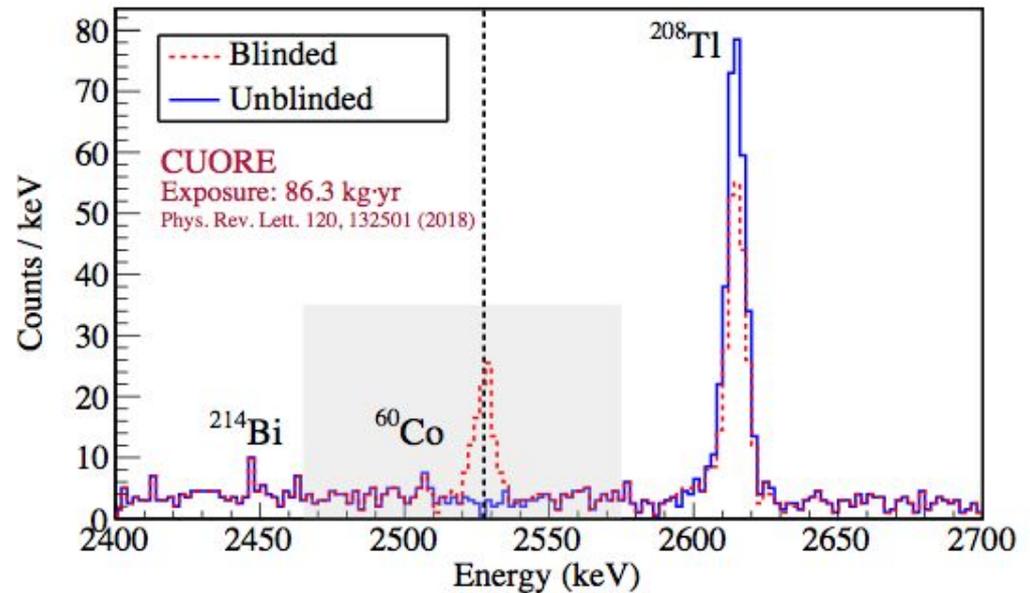
- Internal DCS: 12 strings
Bq for 6 inner str, 19.4 Bq for 6 outer str)
- External DCS: 8 strings
kBq total)



Blinding of the background spectrum

Blinding procedure:

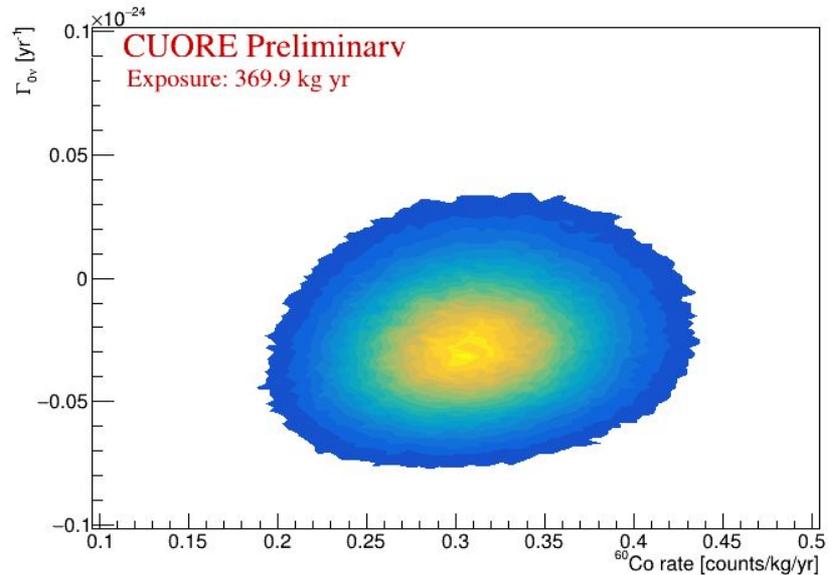
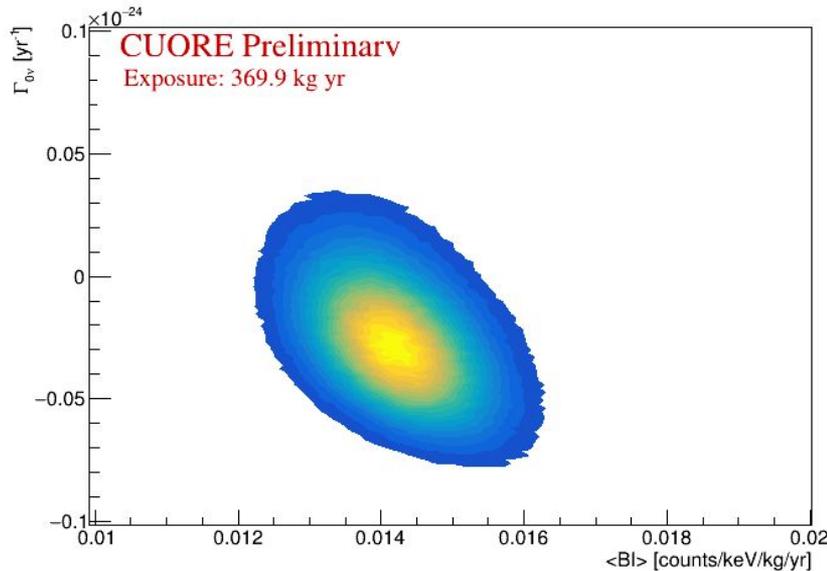
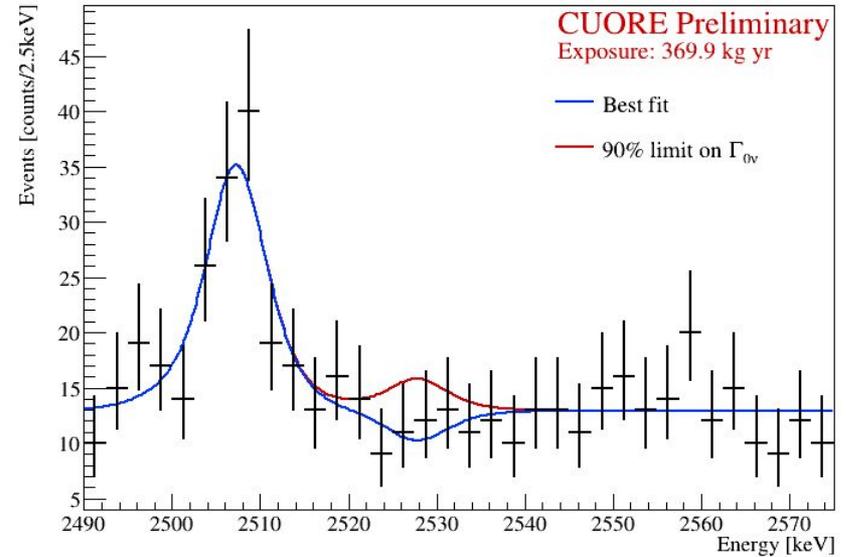
- Choose a random fraction and move events from ± 20 keV of 2615 keV to the $Q_{\beta\beta}$ and vice versa
 - The blinding algorithm produces an artificial peak around the $0\nu\beta\beta$ Q-value and blinds the real $0\nu\beta\beta$ rate of ^{130}Te
 - When all data analysis procedures are fixed the data are eventually unblinded



Unblinded fit

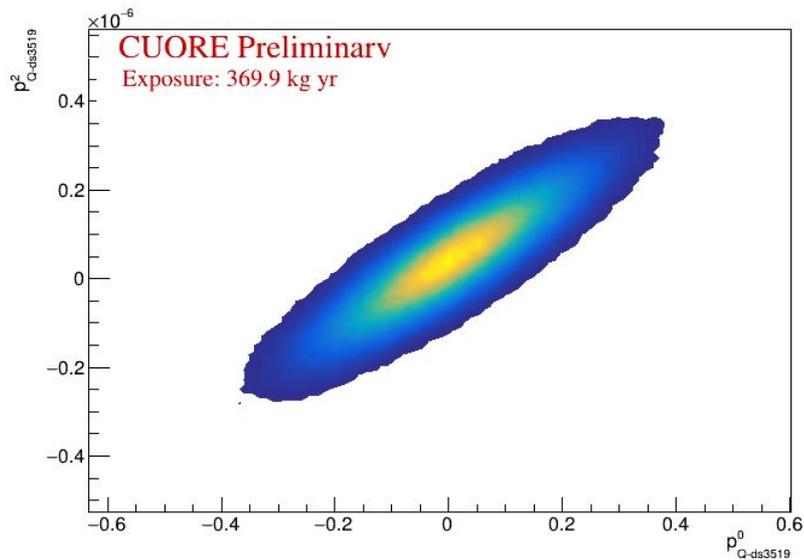
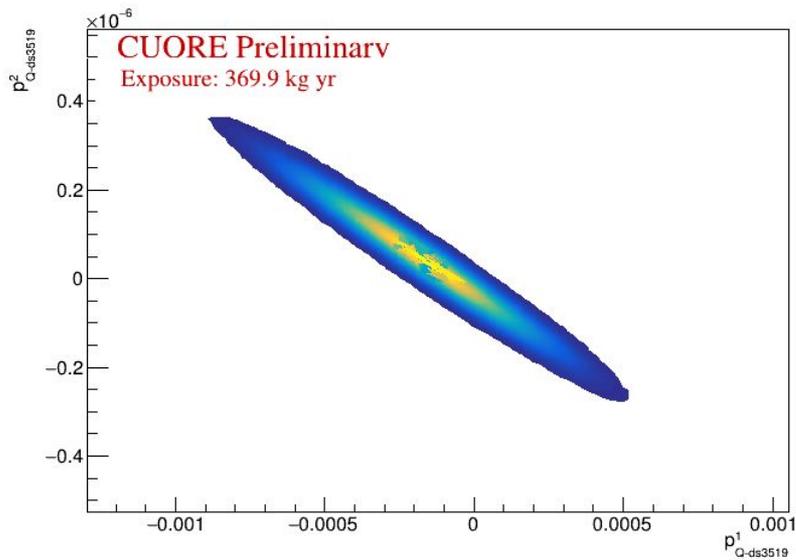
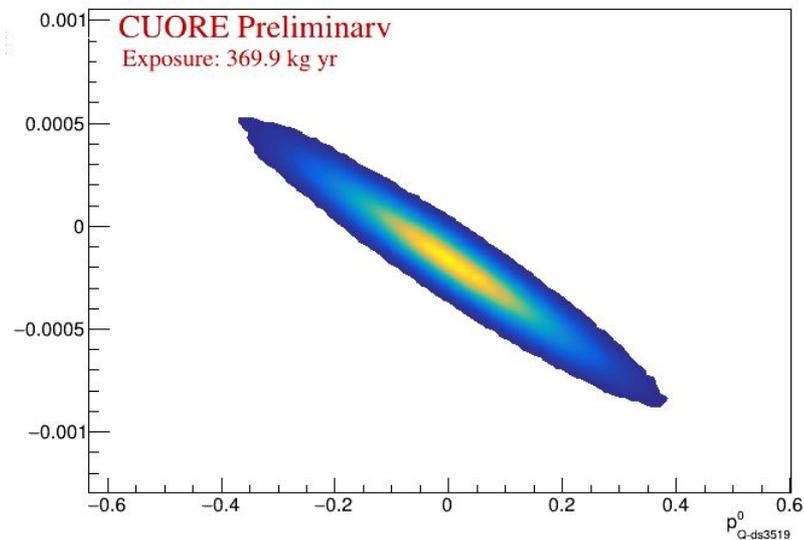
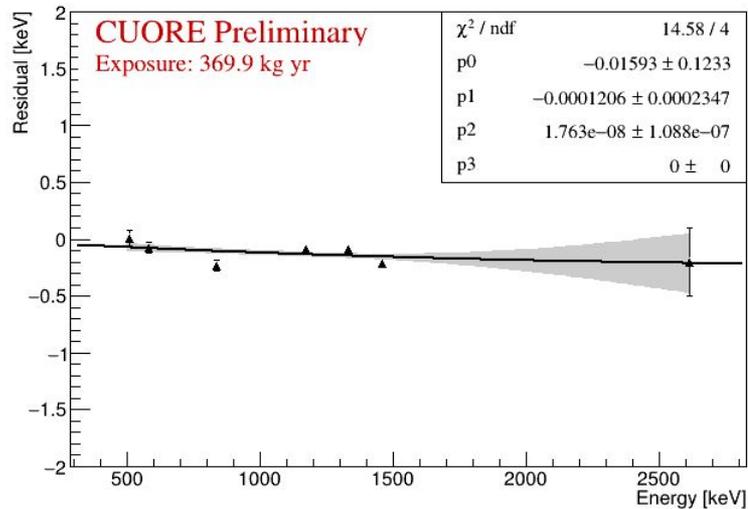
Method

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- Allow negative non-physical range for $\Gamma_{0\nu}$ to evaluate the amplitude of possible background under-fluctuations
- Repeat fit on physical range only
→ Results on $\Gamma_{0\nu}$ obtained from this!
- Repeat fit with additional nuisance parameters to account for systematics



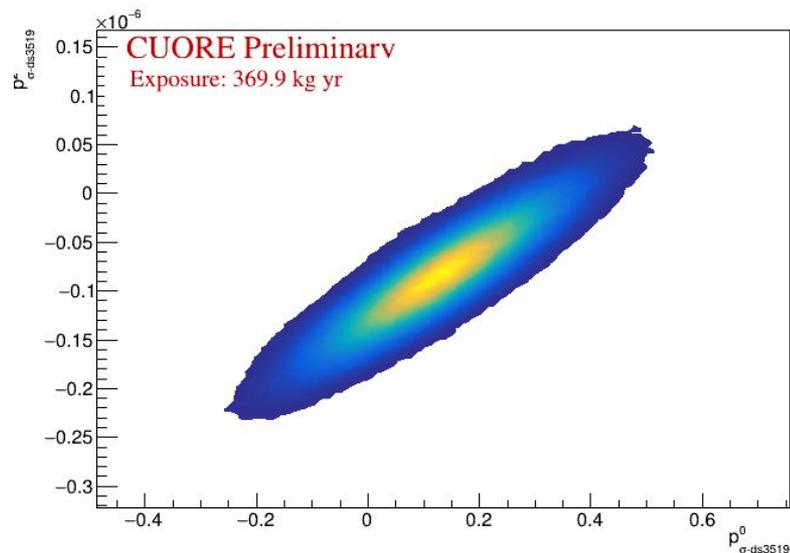
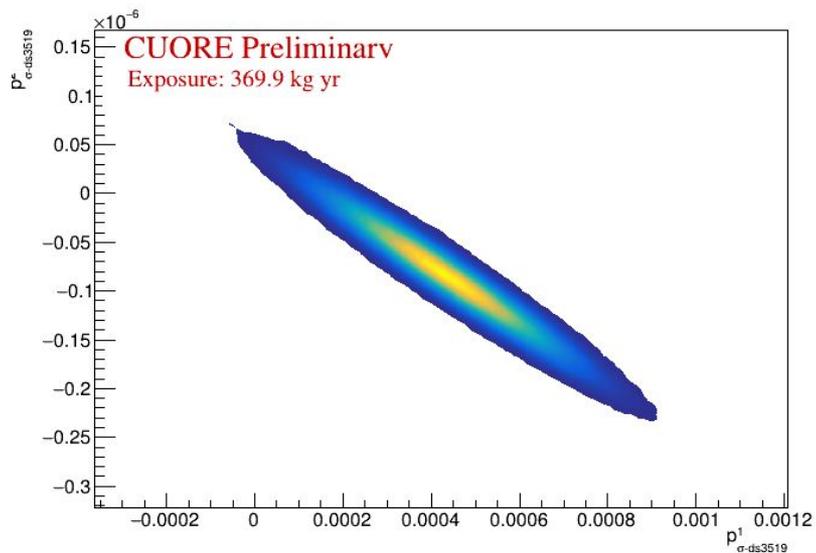
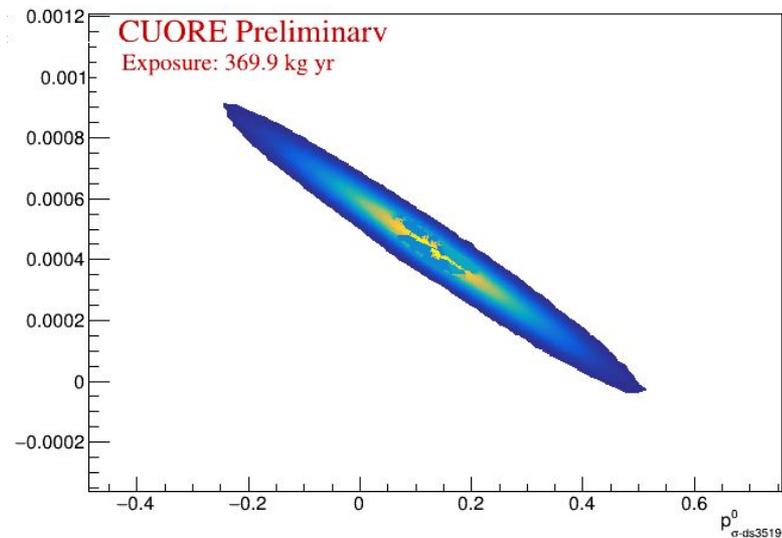
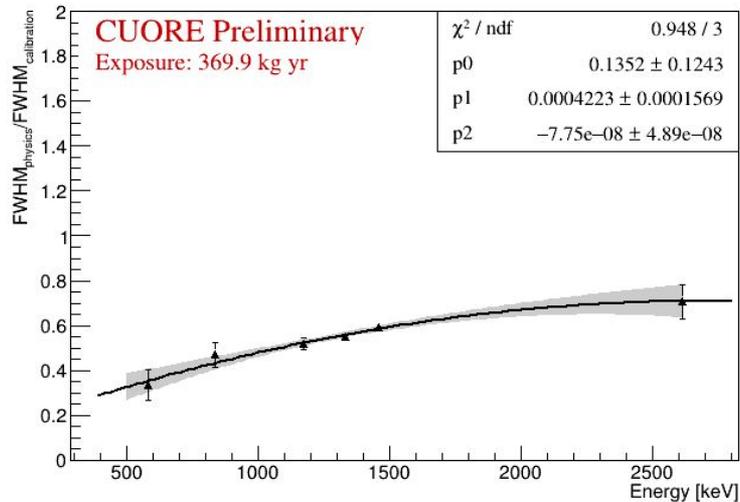
Unblinded fit: systematic on lineshape

Residual vs. energy, 1- σ fits (ds3519)

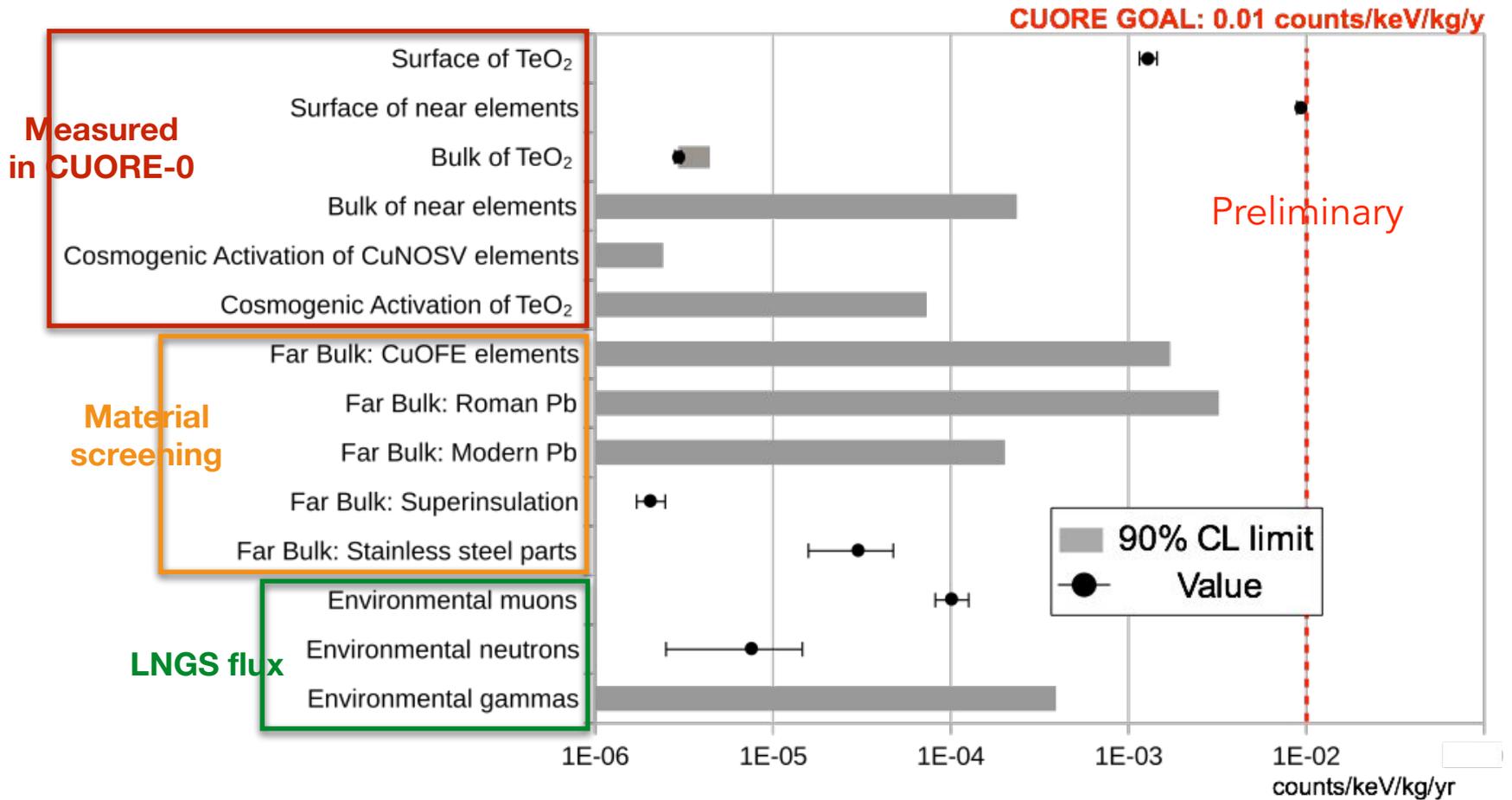


Unblinded fit: systematic on lineshape

Background resolution vs. energy, 1- σ fits (ds3519)



CUORE background budget



CUORE sensitivity and perspectives

CUORE $0\nu\beta\beta$ exclusion sensitivity in 5 years (90% C.L.):

$$S_{0\nu} \sim 9 \times 10^{25} \text{ yr}$$

with

nominal background: $10^{-2} \text{ c}/(\text{keV} \cdot \text{kg} \cdot \text{yr})$

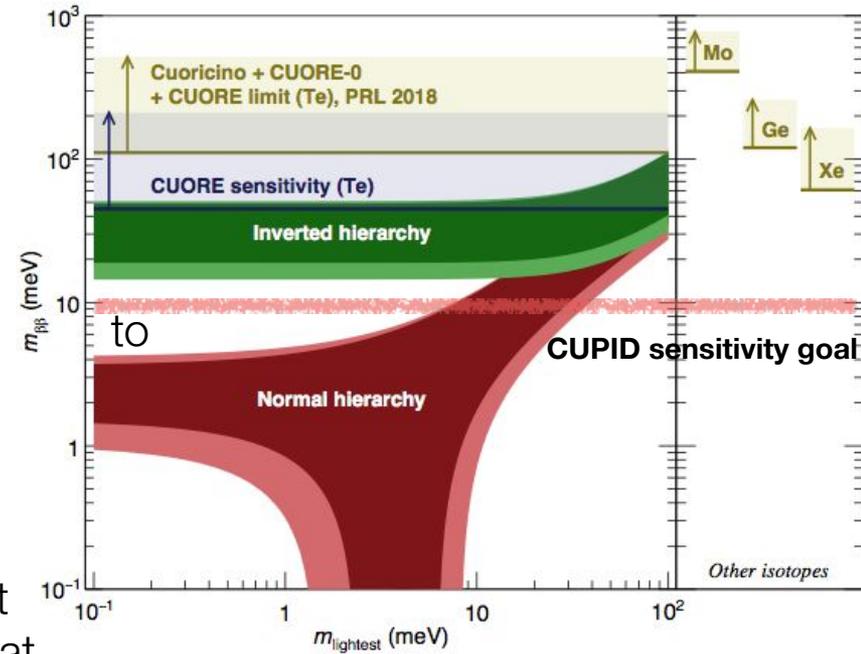
and

nominal energy resolution : 5 keV FWHM
in the Region Of Interest (ROI)

Next generation of $0\nu\beta\beta$ decay experiments seek
be sensitive to the full Inverted Hierarchy region:

$$\text{Sensitivity } S_{0\nu} \sim 10^{27} \text{ yr}, m_{\beta\beta} \sim 6 - 20 \text{ meV}$$

CUPID (CUORE Upgrade with Particle ID) project:
build a future experiment with ~ 1500 enriched light
emitting bolometers mounted in the CUORE cryostat,
reaching nearly zero background goal,
 $< 10^{-4} \text{ c}/(\text{keV} \cdot \text{kg} \cdot \text{yr})$



Bkg

The CUORE challenge

•Low temperature and low vibrations

TeO₂ detectors to be operated as bolometers at temperature ~10 mK: need for cryogenic infrastructure

•Multistage cryogen-free cryostat:

Nested vessels at decreasing temperature

Cooling systems: Pulse Tubes (PTs) and Dilution Unit (DU)

-Mass to be cooled < 4K: ~ 15 tons (IVC volume and Cu vessels, Roman Pb shield)

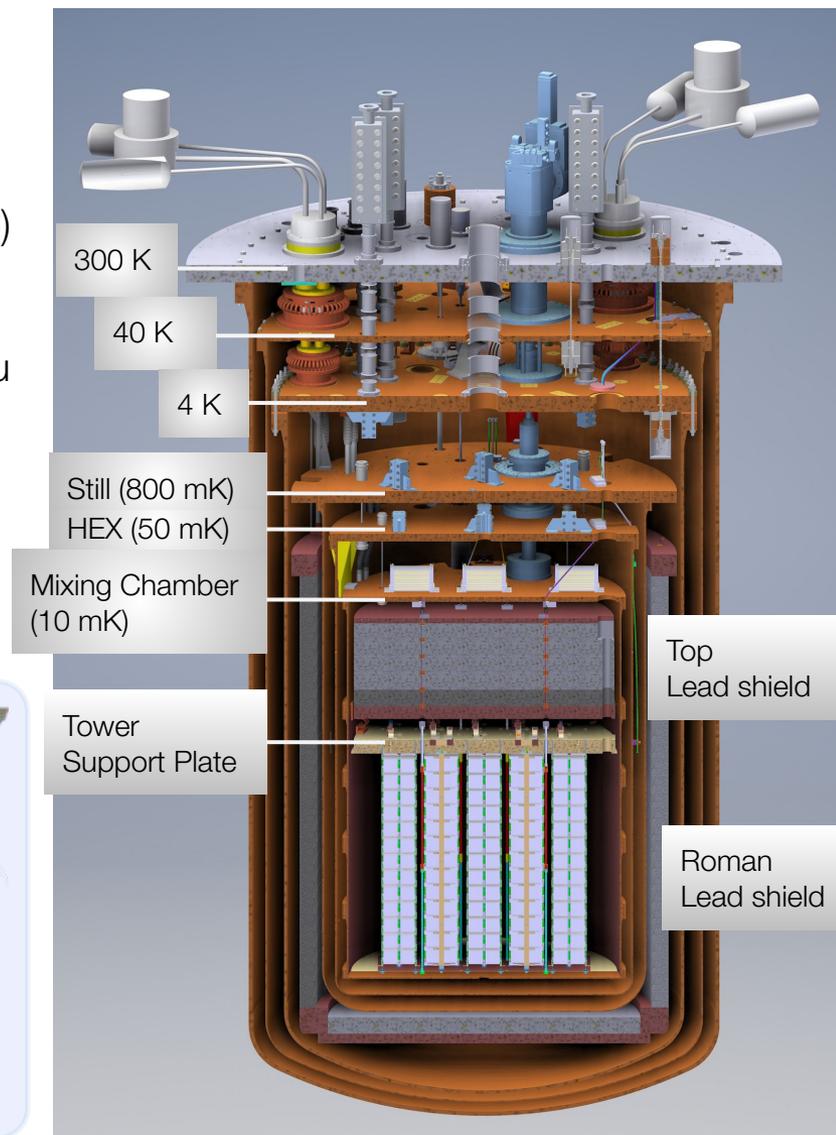
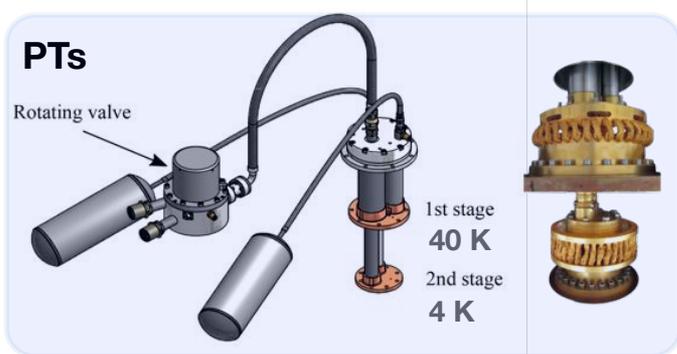
-Mass to be cooled < 50 mK: ~ 3 tons (Top Pb shield, Cu supports and TeO₂ detectors)

• Mechanical vibration isolation

Reduce energy dissipation by vibrations

Target energy resolution: 5 keV FWHM

in the Region Of Interest (ROI) around Q_{ββ}



Dell'Oro S. et al., Cryogenics 102, 9, (2019)
<https://doi.org/10.1016/j.cryogenics.2019.06.011>

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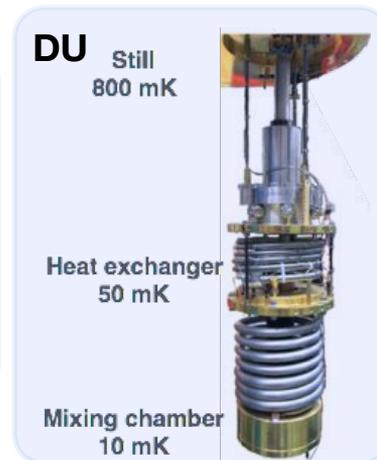
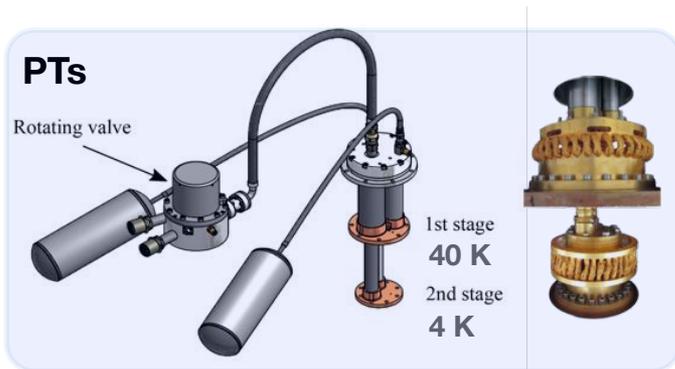
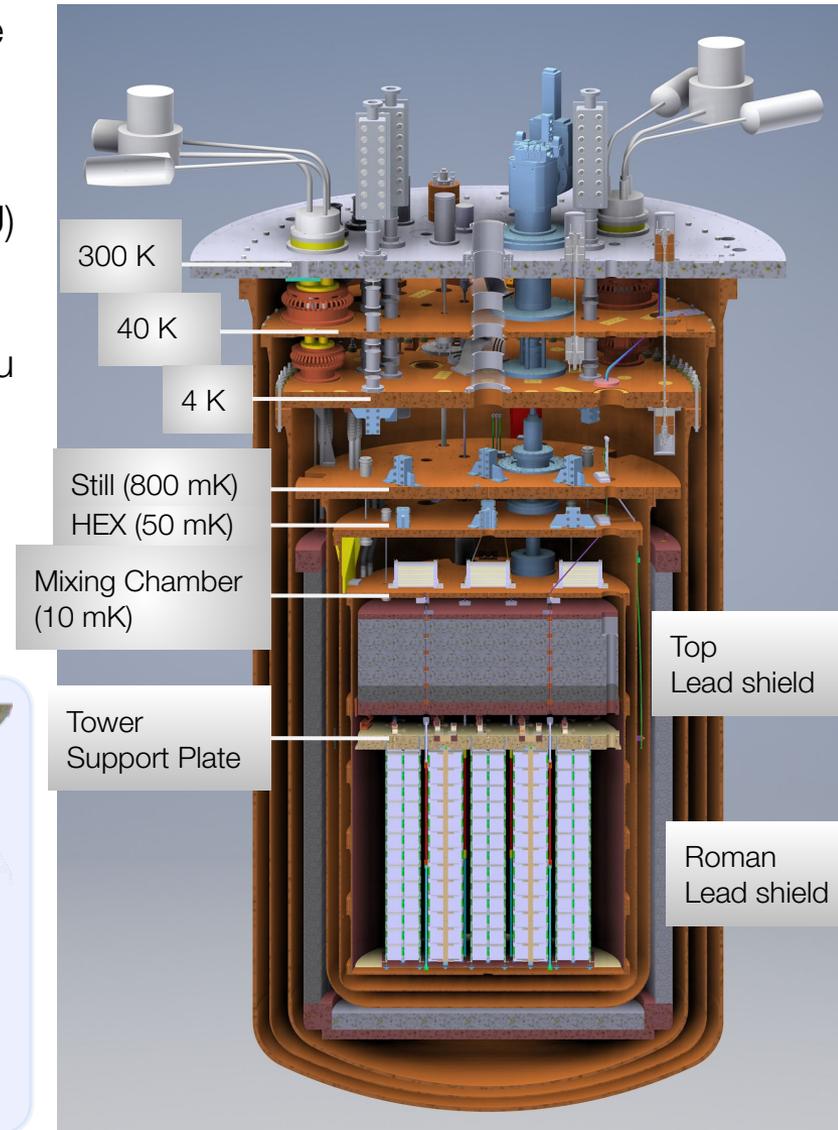
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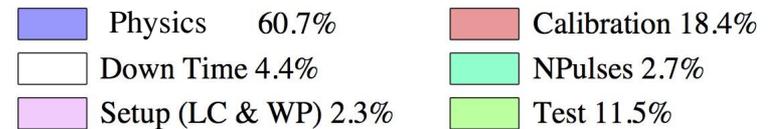
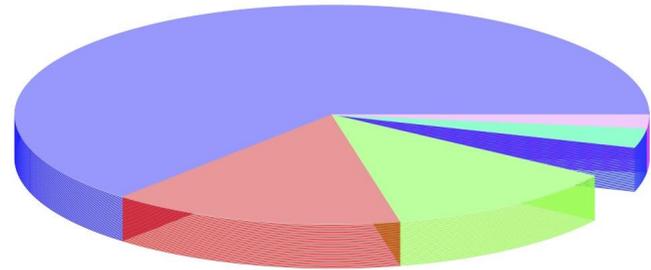
CUORE data taking



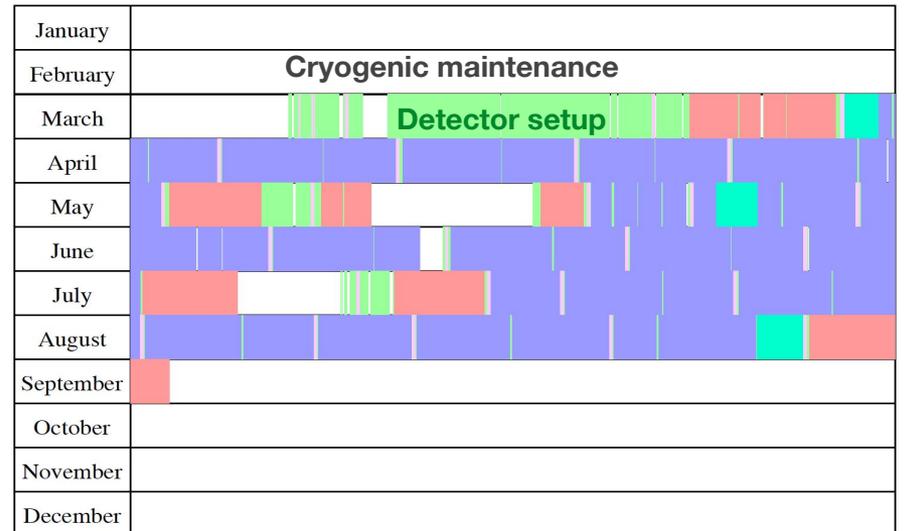
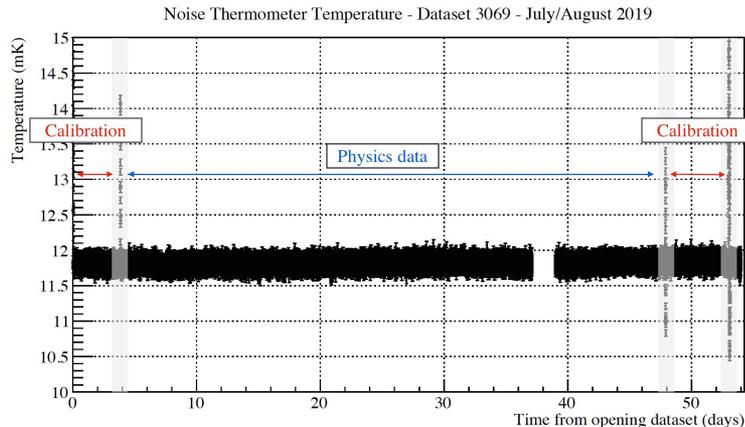
Data taking from March 2019 on

- Reached stable data-taking with high duty cycle
- Physics data bracketed by initial and final calibrations
- Calibration with EDCS
- Only short maintenance operations at dataset closure; checks for noise and thermal response consistency after the activities
- 3 complete consecutive datasets acquired, new dataset started in these days
- Priority is data taking - no further optimizations in the near future

CUORE Run Time Breakdown from March 2019



2019



Open questions about neutrinos



- ❖ The Minimal Standard Model (MSM) assumes massless neutrinos; need to extend it to accommodate for massive neutrinos
- ❖ Neutrino mass hierarchy
- ❖ Neutrino mass absolute values
- ❖ Why neutrino mass is so small?
- ❖ Neutrinos are Dirac or Majorana particles?

Experimental searches:

- Neutrino flavor oscillation parameters precision measurements: Super-K, SNO, KamLAND, Daya Bay, T2K, NoVA
- Sum of neutrinos masses from cosmological and astrophysical data: PLANCK,...
- Direct neutrino mass measurement from β -decay spectral shape: KATRIN (3H), HOLMES and ECHO (163Ho)
- Majorana nature of neutrino via neutrinoless double beta decay: CUORE, GERDA, EXO-200, NEMO...; CUPID, LEGEND, Kamland-Zen, SNO+,...

Neutrinoless double beta decay: unique tool to probe the Majorana nature of neutrino