

# Il Sapore in SUSY: Scenario Gerarchico

Marco Nardecchia  
SISSA e INFN, Trieste

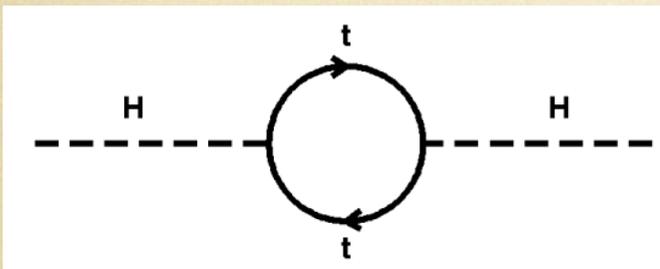
G. F. Giudice, M.N., A. Romanino  
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# Il Problema del Sapore oltre il MS

- Assumendo che il MS sia valido fino ad una scala di energia  $\Lambda$  :

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}}^{\text{gauge}} + \mathcal{L}_{\text{SM}}^{\text{Yukawa}} + \mathcal{L}_{\text{SM}}^{\text{Higgs}} + \frac{\mathcal{L}_5}{\Lambda} + \frac{\mathcal{L}_6}{\Lambda^2} + \dots$$

- Limite superiore:



$$\begin{aligned} (m_h^2)_{\text{phys}} &= (m_h^2)_{\text{tree}} + \frac{2}{\sqrt{2}\pi^2} G_F m_t^2 \Lambda^2 \\ &= (m_h^2)_{\text{tree}} + (0.3\Lambda)^2 \\ \Rightarrow \Lambda &\lesssim 1 \text{ TeV} \end{aligned}$$

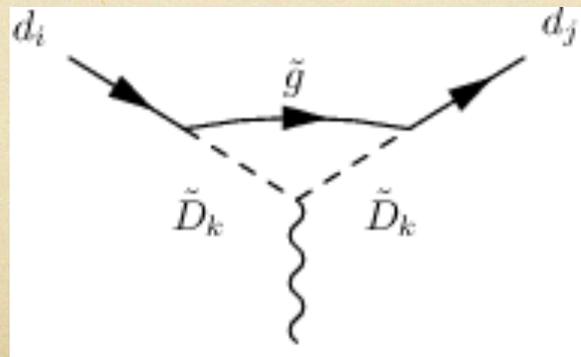
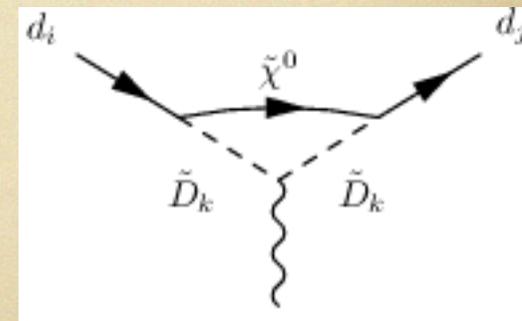
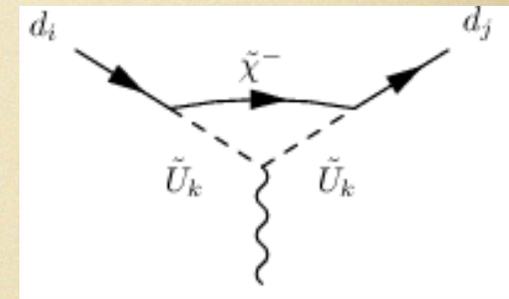
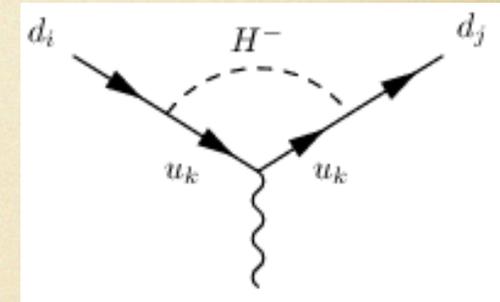
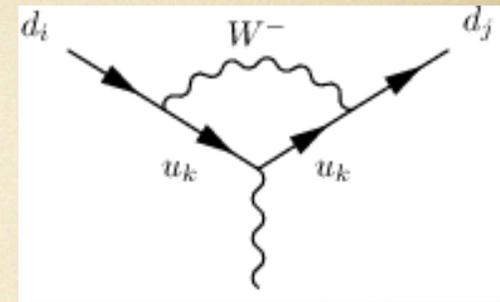
- Limiti inferiori:

broken symmetry	operators	scale $\Lambda$
$B, L$	$(QQQL)/\Lambda^2$	$10^{13}$ TeV
flavor (1,2 <sup>nd</sup> family), CP	$(\bar{d}s\bar{d}s)/\Lambda^2$	1000 TeV
flavor (2,3 <sup>rd</sup> family)	$m_b(\bar{s}\sigma_{\mu\nu}F^{\mu\nu}b)/\Lambda^2$	50 TeV

- Se  $\Lambda \lesssim 1 \text{ TeV}$ , la nuova fisica ha un struttura di sapore non generica

# Violazione del Sapore nel MSSM

- $W$  e quarks up
- $H$  e quarks up
- Chargino e squarks up
- Neutralino e squarks down
- Gluino e squarks down

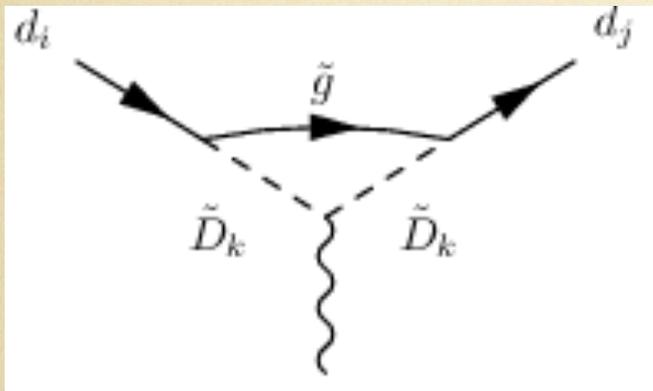


# Soppressione dei Contributi SUSY

Consideriamo per esempio, il contributo del gluino ad un processo  $\Delta F=1$ :  $d_i^A \rightarrow d_j^B$

L'ampiezza di Feynman assume la seguente forma:

$$\begin{aligned} A, B &= \{L, R\} \\ i, j &= \{1, 2, 3\} \\ i &\neq j \end{aligned}$$



$$\mathcal{W}_{d_i^A k} f(x_k) \mathcal{W}_{k d_j^B}^\dagger$$

$$x_k = \frac{m_{\tilde{q}_k}^2}{m_{\tilde{g}}^2} \quad k = \{1, \dots, 6\}$$

Abbiamo soppressione se:

- Le particelle nei loops sono pesanti:
- Gli squarks nei loops hanno massa degenera:
- La matrice unitaria di mixing è circa l'identità:

$$\lim_{x \rightarrow \infty} f(x) = \lim_{x \rightarrow 0} f(x) = 0$$

$$f(x) \mathcal{W}_{d_i^A k} \mathcal{W}_{k d_j^B}^\dagger = f(x) \delta_{AB} \delta_{ij} = 0$$

$$f(x_k) \delta_{A_i k} \delta_{k B_j} = f(x_k) \delta_{AB} \delta_{ij} = 0$$

# Un Approccio Bottom-Up

- Espansione in elementi fuori diagonale della matrice di massa:

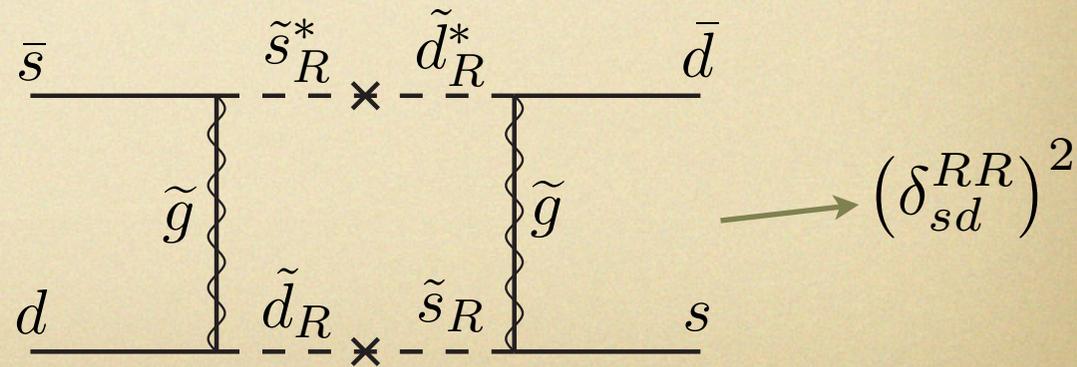
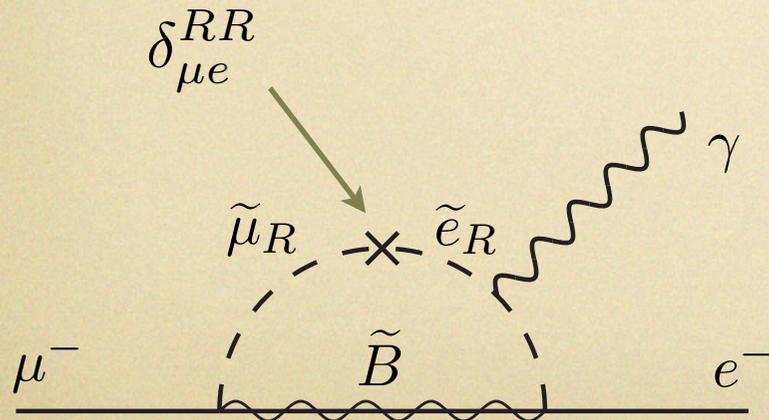
$$\mathcal{M}^2 = \mathcal{M}_0^2 + \delta\mathcal{M}^2$$

- Scenario degenere:  $\mathcal{M}_0^2 = \tilde{m}^2 I_{6 \times 6}$

$$A(\Delta F = 1) = x f^{(1)}(x) \delta_{ij} \quad \delta_{ij} \equiv \frac{(\delta\mathcal{M}^2)_{ij}}{\tilde{m}^2} \quad x = \frac{\tilde{m}^2}{M^2}$$

$$A(\Delta F = 2) = \frac{x^2}{3!} g^{(3)}(x) \delta_{ij}^2$$

- $f, g$ : funzioni di loop
- $\delta_{ij}$ : inserzioni di massa che violano il sapore



# Lo Scenario Gerarchico

Nello scenario gerarchico la matrice di massa degli sfermioni (blocco LL o RR), ha la seguente forma :

$$\tilde{m}^2 = \begin{pmatrix} h_{11} & h_{12} & a_1 \\ h_{21} & h_{22} & a_2 \\ \bar{a}_1 & \bar{a}_2 & l_3 \end{pmatrix} \quad \text{dove "h" è il blocco pesante, i rimanenti elementi sono leggeri}$$

In particolare abbiamo che:  $l \lesssim \text{TeV}^2, h_1, h_2 \gg l$

La diagonalizzazione di massa avviene mediante una matrice unitaria:

$$\mathcal{M}^2 = \mathcal{W} \text{Diag}(\tilde{m}_{L1}^2, \tilde{m}_{L2}^2, \tilde{m}_{R1}^2, \tilde{m}_{R2}^2, \tilde{m}_1^2, \tilde{m}_2^2) \mathcal{W}^\dagger$$

Decomponendo ora l'ampiezza, possiamo eliminare il contributo delle particelle pesanti:

$$f\left(\frac{\mathcal{M}^2}{M^2}\right)_{Ai, Bj} = \mathcal{W}_{Ai, Lk} f\left(\frac{\tilde{m}_{Lk}^2}{m_{\tilde{g}}^2}\right) \mathcal{W}_{Lk, Bj}^\dagger + \mathcal{W}_{Ai, Rk} f\left(\frac{\tilde{m}_{Rk}^2}{m_{\tilde{g}}^2}\right) \mathcal{W}_{Rk, Bj}^\dagger + \mathcal{W}_{Ai, p} f\left(\frac{\tilde{m}_p^2}{m_{\tilde{g}}^2}\right) \mathcal{W}_{p, Bj}^\dagger$$

Blocco pesante LL

Blocco pesante RR

Contributo dominante della terza famiglia leggera

# Lo Scenario Gerarchico

Motivazioni: [Effective SUSY, Choen Kaplan Lepeintre Nelson '97]

- Complementare all'assunzione con spettro degenere
- Allevia il problema del sapore, senza compromettere la naturalezza [Dimopoulos-Giudice '95]
- Anche se assumiamo uno spettro degenere ad una data scala di energia, gli effetti delle correzioni radiative (Yukawa) tendono a rendere la terza famiglia leggera

$$A(\Delta F = 1) = f(x) \hat{\delta}_{ij}$$

$$A(\Delta F = 2) = g^{(1)}(x) \hat{\delta}_{ij}^2$$

$$x = \frac{\tilde{m}_3^2}{M^2} \quad \hat{\delta}_{a3} = \frac{\mathcal{M}_{a3}^2}{\tilde{m}_h^2}$$

Ci sono solo 4 inserzioni che violano il sapore:

$$\hat{\delta}_{bd}^{LL}, \hat{\delta}_{bd}^{LL}, \hat{\delta}_{bd}^{RR}, \hat{\delta}_{bd}^{RR}$$

$$\hat{\delta}_{ij}^{LL} \equiv \hat{\delta}_{i3}^{LL} \hat{\delta}_{j3}^{LL*}$$

$$\hat{\delta}_{ij}^{LR} \equiv \frac{\mathcal{M}_{L3,R3}^2}{\tilde{m}^2} \hat{\delta}_{i3}^{LL} \hat{\delta}_{j3}^{RR*} \quad i, j = 1, 2$$

$$\hat{\delta}_{i3}^{LR} \equiv \frac{\mathcal{M}_{L3,R3}^2}{\tilde{m}^2} \hat{\delta}_{i3}^{LL}$$

Correlazioni tra le osservabili nello scenario degenere e gerarchico possono essere molto diverse:

$$\frac{A(\Delta F = 2)}{[A(\Delta F = 1)]^2} \Big|_{\text{degenerate}} = \frac{g^{(3)}}{6g^{(1)}} \left( \frac{f}{f^{(1)}} \right)^2 \frac{A(\Delta F = 2)}{[A(\Delta F = 1)]^2} \Big|_{\text{hierarchical}}$$

# Limiti sulle Inserzioni

x=1 Hierarchy

$D_0 - \bar{D}_0$  mixing

Degeneracy

$$\left| \hat{\delta}_{ut}^{LL} \hat{\delta}_{ct}^{LL*} \right| < 8.0 \times 10^{-3} \left( \frac{m_{\bar{t}}}{350 \text{ GeV}} \right) \quad \left| \delta_{uc}^{LL} \right| < 3.4 \times 10^{-2} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right)$$

$B \rightarrow X_s \gamma$

$$\begin{array}{|l} \left| \text{Re}(\hat{\delta}_{sb}^{LL}) \right| < 2.2 \times 10^{-2} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right)^2 \left( \frac{10}{\tan \beta} \right) \\ \left| \text{Im}(\hat{\delta}_{sb}^{LL}) \right| < 6.7 \times 10^{-2} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right)^2 \left( \frac{10}{\tan \beta} \right) \end{array} \quad \begin{array}{|l} \left| \text{Re}(\delta_{sb}^{LL}) \right| < 3.8 \times 10^{-2} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right)^2 \left( \frac{10}{\tan \beta} \right) \\ \left| \text{Im}(\delta_{sb}^{LL}) \right| < 1.1 \times 10^{-1} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right)^2 \left( \frac{10}{\tan \beta} \right) \end{array}$$

$\Delta m_{B_s}$

$$\begin{array}{|l} \left| \text{Re}(\hat{\delta}_{sb}^{LL}) \right| < 9.4 \times 10^{-2} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right) \\ \left| \text{Im}(\hat{\delta}_{sb}^{LL}) \right| < 7.2 \times 10^{-2} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right) \end{array} \quad \begin{array}{|l} \left| \text{Re}(\delta_{sb}^{LL}) \right| < 4.0 \times 10^{-1} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right) \\ \left| \text{Im}(\delta_{sb}^{LL}) \right| < 3.1 \times 10^{-1} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right) \end{array}$$

$$\left| \hat{\delta}_{sb} \right| \approx |V_{ts}^*| \approx 4 \times 10^{-2}$$

$B_d^0 - \bar{B}_d^0$  mixing

$$\begin{array}{|l} \left| \text{Re}(\hat{\delta}_{db}^{LL}) \right| < 4.3 \times 10^{-3} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right) \\ \left| \text{Im}(\hat{\delta}_{db}^{LL}) \right| < 7.3 \times 10^{-3} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right) \end{array} \quad \begin{array}{|l} \left| \text{Re}(\delta_{db}^{LL}) \right| < 1.8 \times 10^{-2} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right) \\ \left| \text{Im}(\delta_{db}^{LL}) \right| < 3.1 \times 10^{-2} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right) \end{array}$$

$$\left| \hat{\delta}_{db} \right| \approx |V_{td}^*| \approx 8 \times 10^{-3}$$

$\Delta m_K$

$$\sqrt{\left| \text{Re}(\hat{\delta}_{db}^{LL} \hat{\delta}_{sb}^{LL*})^2 \right|} < 1.0 \times 10^{-2} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right) \quad \sqrt{\left| \text{Re}(\delta_{ds}^{LL})^2 \right|} < 4.2 \times 10^{-2} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right)$$

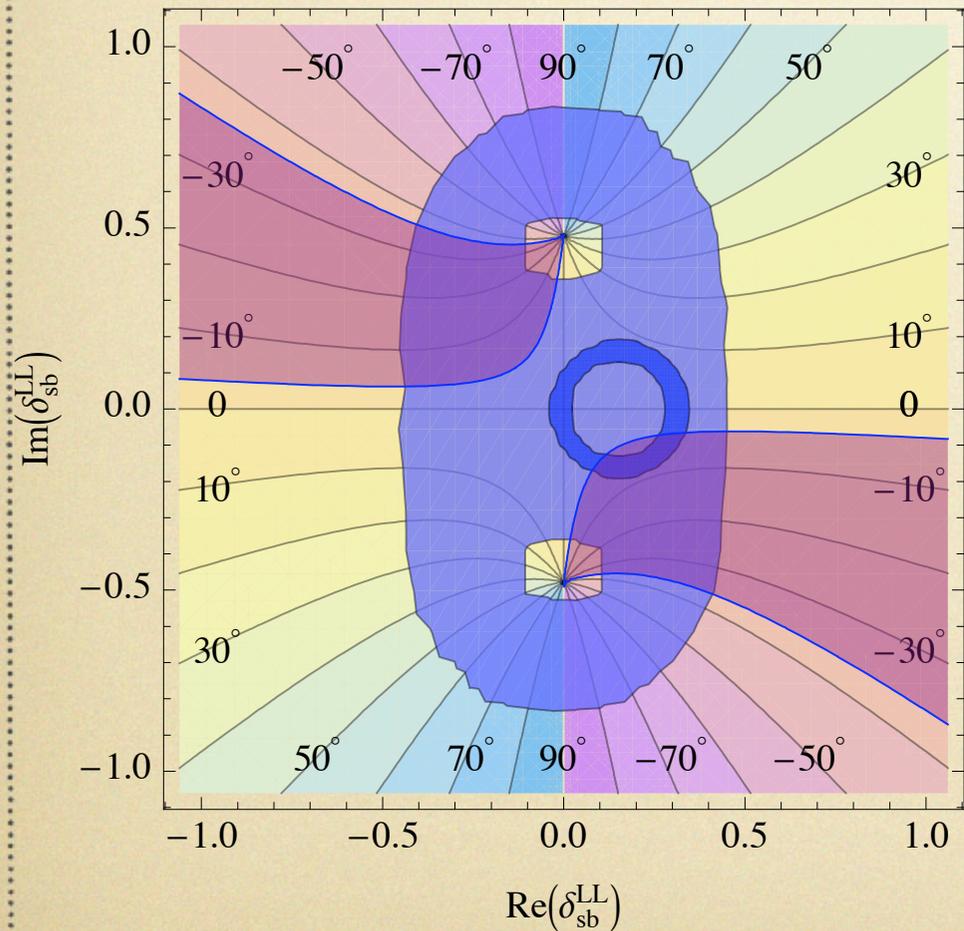
$$\left| \hat{\delta}_{db} \hat{\delta}_{bs} \right| \approx 3 \times 10^{-4}$$

$\epsilon_K$

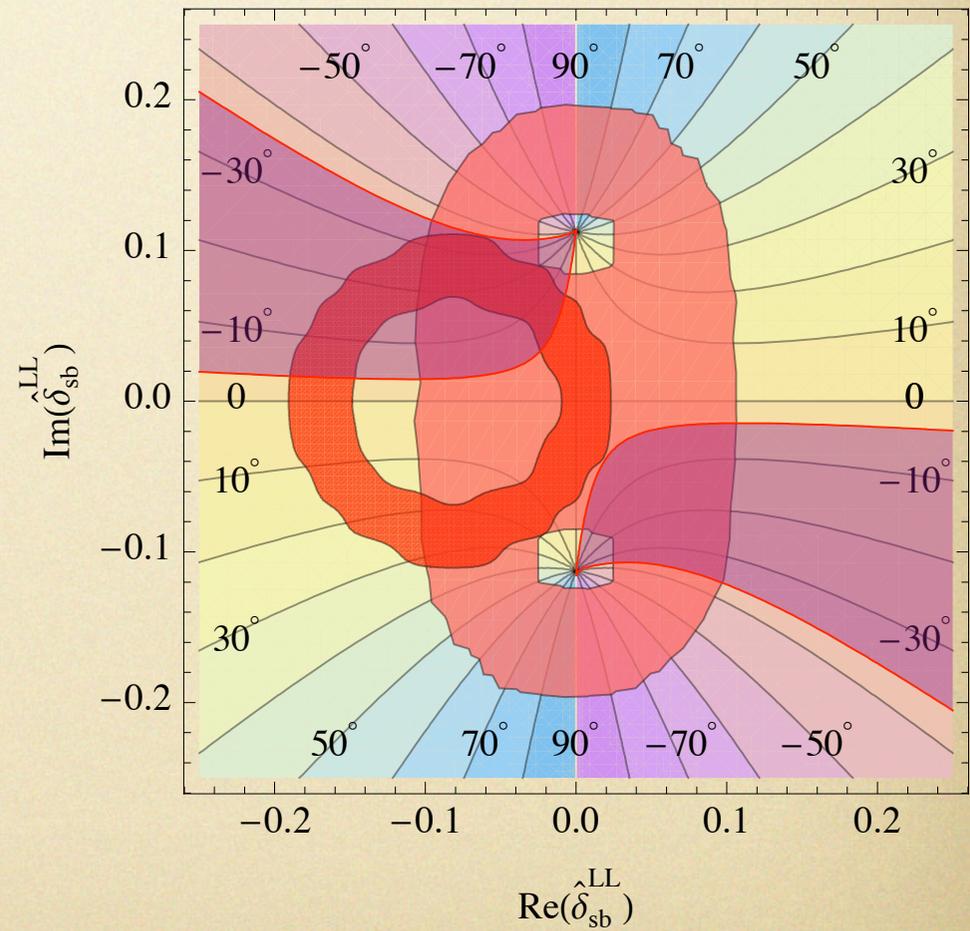
$$\sqrt{\left| \text{Im}(\hat{\delta}_{db}^{LL} \hat{\delta}_{sb}^{LL*})^2 \right|} < 4.4 \times 10^{-4} \left( \frac{m_{\bar{b}}}{350 \text{ GeV}} \right) \quad \sqrt{\left| \text{Im}(\delta_{ds}^{LL})^2 \right|} < 1.8 \times 10^{-3} \left( \frac{m_{\bar{q}}}{350 \text{ GeV}} \right)$$

$$\Delta M_{B_s}, \quad B \rightarrow X_s \gamma, \quad \phi_{B_s}$$

## Scenario Degenerere



## Scenario Gerarchico



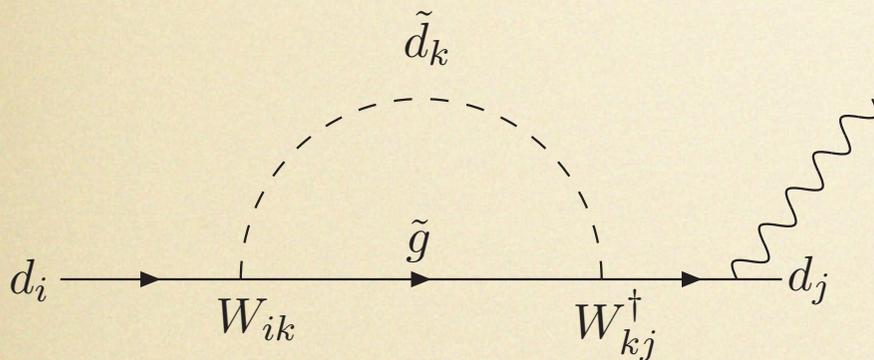
$$\tilde{m} = M_{\tilde{g}} = \mu = 350 \text{ GeV}, \quad \tan \beta = 10, \quad A = 0$$

# Scenario Gerarchico: Conclusioni

- E' possibile avere le prime due famiglie di sfermioni pesanti senza compromettere la naturalezza della massa dell'Higgs
- Si allevia il problema del sapore nel MSSM
- Utile scenario di confronto per modelli in cui l'assunzione di uno spettro degenere non è valida
- Offre predizioni peculiari e correlazioni tra diverse osservabili

# FCNC processes

- Consider the chirality conserving case  $\begin{cases} d_i \rightarrow d_j \\ i \neq j \end{cases}$



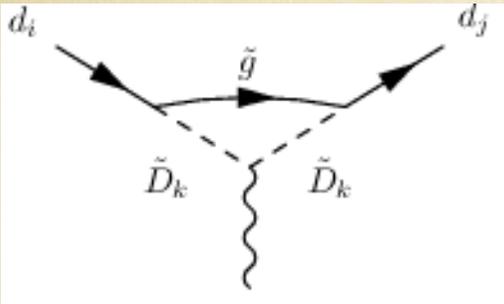
$$\sum_{k=1}^3 W_{ik} f(x_k) W_{kj}^\dagger \quad x_k \equiv \frac{m_{\tilde{d}_k}^2}{m_{\tilde{g}}^2}$$

- In the hierarchical case:  $\sum_{k=1}^3 W_{ik} f(x_k) W_{kj}^\dagger \approx f(x_3) W_{i3} W_{3j}^\dagger \equiv f(x_3) \hat{\delta}_{ij}$
- We can define the flavor violating parameters:

$$\begin{cases} \hat{\delta}_{13} = W_{13} W_{33}^\dagger \approx W_{13} \\ \hat{\delta}_{23} = W_{23} W_{33}^\dagger \approx W_{23} \\ \hat{\delta}_{12} = W_{13} W_{32}^\dagger \approx \hat{\delta}_{13} \hat{\delta}_{23}^* \end{cases}$$

We expect a suppression in the contribution to Kaon Physics and a correlation between the parameters

# Natural size of the insertions



- $\mathcal{W}_{A_i k}$  is the mixing between the fermion “Ai” and the k-th sfermion mass eigenstate
- In a general basis we have:

$$V = U_L^u U_L^{d\dagger} \quad \mathcal{W} = \begin{pmatrix} U_L & 0 \\ 0 & U_R \end{pmatrix} \mathcal{W}', \quad U_R M U_L^\dagger = \text{diagonal}, \quad \mathcal{W}'^\dagger \mathcal{M}^2 \mathcal{W}' = \text{diagonal}.$$

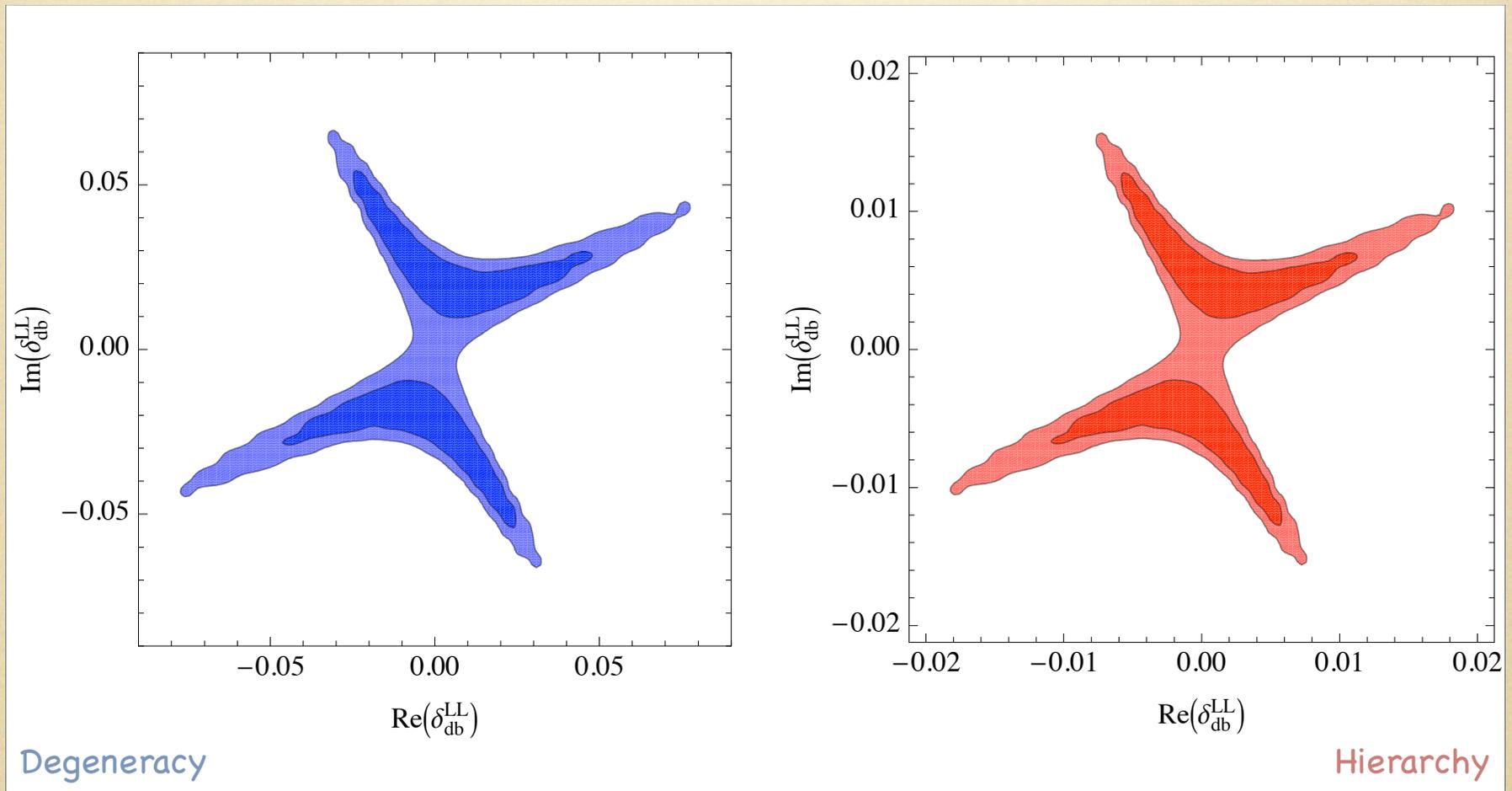
- We now focus on LL insertion in the down sector. In the hierarchical scenario we have that:  $\hat{\delta}_{ij}^{LL} = \mathcal{W}_{i3} \mathcal{W}_{3j}^\dagger$
- In particular  $\hat{\delta}_{i3}^{LL} = \mathcal{W}_{i3} \mathcal{W}_{33}^\dagger \approx \mathcal{W}_{i3} = (U_L^d \mathcal{W}'_L)_{i3}$
- Now we assume that the CKM matrix is dominated by the rotation in the down sector:

$$\hat{\delta}_{i3}^{LL} \approx (V^\dagger \mathcal{W}'_L)_{i3} = \sum_{k=1}^3 V_{ki}^* (\mathcal{W}'_L)_{k3}$$

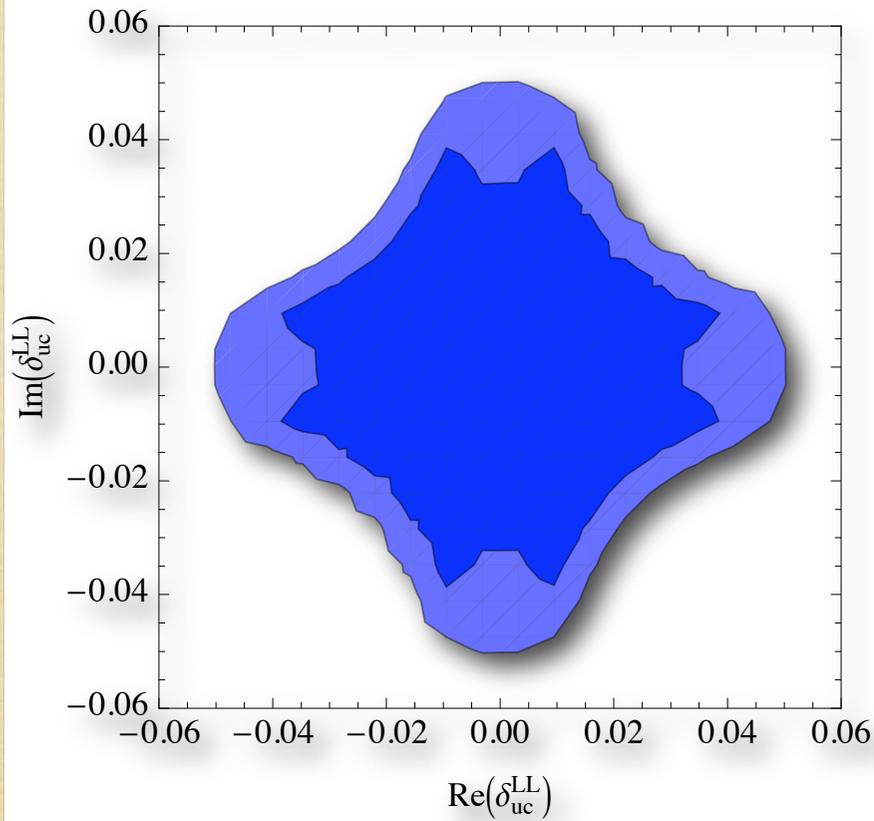
- Barring accidental cancellations:

$$\left| \hat{\delta}_{i3} \right| \geq |V_{3i}^* (\mathcal{W}'_L)_{33}| \approx |V_{3i}^*|$$

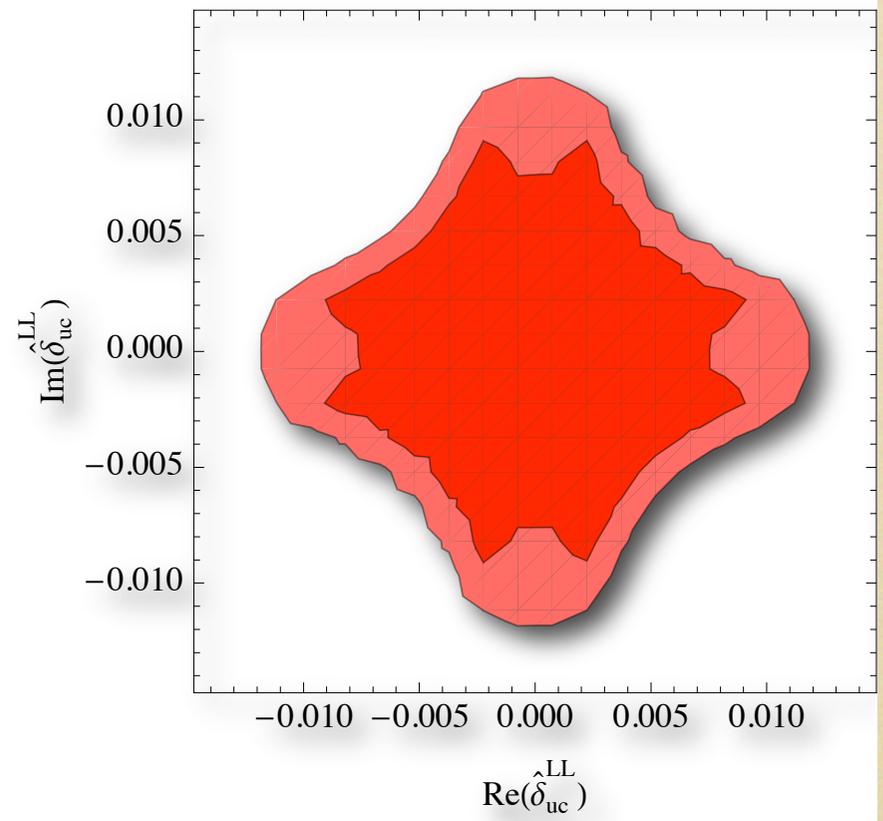
# B-d Transition



# D Mixing



Degeneracy

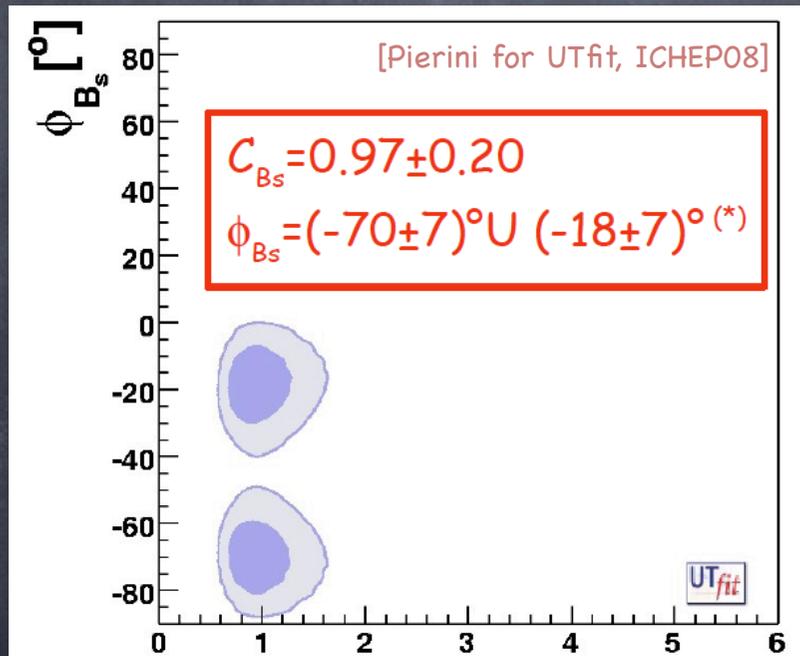


Hierarchy

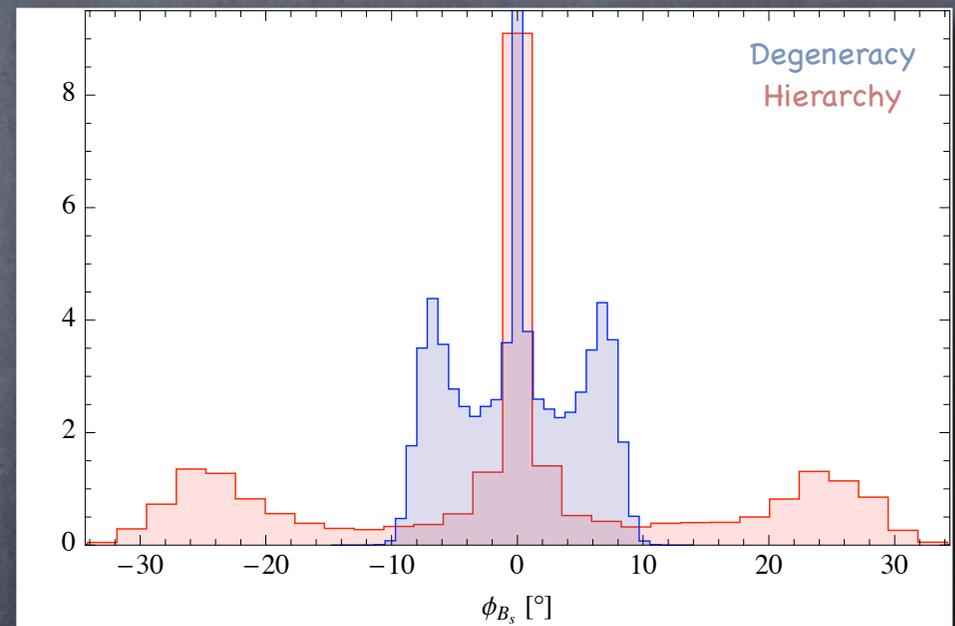
# $\phi_{B_s}$

$$\langle B_s | H_{\text{eff}}^{\text{full}} | \bar{B}_s \rangle = C_{B_s} e^{2i\phi_{B_s}} \langle B_s | H_{\text{eff}}^{\text{SM}} | \bar{B}_s \rangle$$

$$\langle B_s | H_{\text{eff}}^{\text{SM}} | \bar{B}_s \rangle = A_s^{\text{SM}} e^{-2i\beta_s} \quad \beta_s = \arg(-(V_{ts}V_{tb}^*)/(V_{cs}V_{cb}^*)) = 0.018 \pm 0.001$$



(\*) Significance reduced from  $\sigma \sim 3.0$   $C_{B_s}$  to  $\sigma \sim 2.5$  due to the correlation between  $\Delta\Gamma$ 's and  $\phi$ s in D0 likelihood



# Inputs

Parameter	Value	Gaussian ( $\sigma$ )	Uniform ( $\frac{\Delta}{2}$ )	Reference
$ \varepsilon_K $	$2.229 \times 10^{-3}$	$0.012 \times 10^{-3}$	—	[19]
$\Delta m_K$ (ps $^{-1}$ )	$5.292 \times 10^{-3}$	$0.009 \times 10^{-3}$	—	[19]
BR( $B \rightarrow X_s \gamma$ )	$3.55 \times 10^{-4}$	$0.26 \times 10^{-4}$	—	[20]
$\Delta m_{B_s}$ (ps $^{-1}$ )	17.77	0.12	—	[19]
$\Delta m_{B_d}$ (ps $^{-1}$ )	0.507	0.005	—	[19]
$ M_{12}^D $ (ps $^{-1}$ )	$7.7 \times 10^{-3}$	$2.5 \times 10^{-3}$	—	[21]
$\bar{\rho}$	0.167	0.051	—	[22]
$\bar{\eta}$	0.386	0.035	—	[22]
$\lambda$	0.2255	0.010	—	[19]
$ V_{cb} $	$41.2 \times 10^{-3}$	$1.1 \times 10^{-3}$	—	[19]
$F_K$ (GeV)	0.160	—	—	[19]
$F_{B_d}$ (MeV)	189	27	—	[23]
$F_{B_s} \sqrt{B_s}$ (MeV)	262	35	—	[23]
$F_D$ (MeV)	201	3	17	[24]
$\hat{B}_K$	0.79	0.04	0.08	[24]
$B_1^B$	0.88	0.04	0.10	[24]
$\eta_{cc}$	0.47	0.04	—	[25]
$\eta_{ct}$	0.5765	0.0065	—	[25]
$\eta_{tt}$	1.43	0.23	—	[25]
$\bar{m}_t$ (GeV)	161.2	1.7	—	[24]
$\bar{m}_b$ (GeV)	4.21	0.08	—	[24]
$\bar{m}_c$ (GeV)	1.224	0.057	—	[26]

# Soft Masses and RGE

$$\mathbf{m}_{\tilde{d}}^2 = \begin{pmatrix} \overset{\text{soft}}{\downarrow} m_Q^2 + m_d^2 + \Delta_{\tilde{d}_L} & v(a_d^* \sin \beta - \mu y_d \cos \beta) \\ v(a_d \sin \beta - \mu^* y_d \cos \beta) & m_d^2 + m_d^2 + \Delta_{\tilde{d}_R} \end{pmatrix}$$

$\swarrow$  soft

Roughly speaking, the sfermion mass matrix is dominated by the soft terms, because all the others are suppressed by the EW scale. For the soft terms the RGE are:

First and second family:  $16\pi^2 \frac{d}{dt} m_{\phi_i}^2 = - \sum_{a=1,2,3} 8C_a(i) g_a^2 |M_a|^2 + \frac{6}{5} Y_i g_1^2 S$

Third family Left:  $16\pi^2 \frac{d}{dt} m_{Q_3}^2 = X_t + X_b - \frac{32}{3} g_3^2 |M_3|^2 - 6g_2^2 |M_2|^2 - \frac{2}{15} g_1^2 |M_1|^2 + \frac{1}{5} g_1^2 S$

Third family Right:  $16\pi^2 \frac{d}{dt} m_{d_3}^2 = 2X_b - \frac{32}{3} g_3^2 |M_3|^2 - \frac{8}{15} g_1^2 |M_1|^2 + \frac{2}{5} g_1^2 S$

Also if we start with degenerate condition at very high energy, we end up with a split situation because of the Yukawa!!