



Precise Tracking of Cosmic Muons Using Time-over-Threshold Property of NINO ASICs

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Motivations

- Natural Cosmic Ray source.
- Other applications like imaging large structures, monitoring volcanoes and all.
- Based on Coulomb Scattering. (High-Z/ Low-Z discrimination)
- Accurate measurement of scattering angle requires precise tracking.





Strategy of Muon eutries 400 **Scattering Tomography** 350 Scattering Angle 300 Distribution $\frac{dN}{d\theta_x} = \frac{1}{\sqrt{2\pi\theta_0}} e^{-\theta_x^2/2\theta_0^2}$ 250 200 150 100 $\theta_0 \approx \frac{13.6}{(pc\beta)} \times \sqrt{\frac{x}{X_0}} \times \left[1 + 0.038 \ln\left(\frac{x}{X_0}\right)\right]$ 50 10 scattering angle (deg) $X = \frac{716.4 \,\mathrm{g/cm^2}A}{\rho Z (Z+1) \ln \left(287 \,/ \,\sqrt{Z} \right)}$ х, у detectors muon tracks p = Momentum $\beta = V/C$ mcs X= Radiation Length L = ThicknessLateral Deviation A = Atomic Mass Angular Deviation *Z*= Atomic Number $(\Delta \theta)$ $\rho = \text{Density}$ 3



Effect of Resolution on Image reconstruction



1 cm Resolution

Ideal Resolution

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NINO Readout and TOT

|--|--|--|--|

- ✤ 8 channel readout
- Individual drivers for differential input
- Input threshold control 10fc to 100 fc
- Placement of Readout close to the detector



Time Width of o/p proportional to ToT of i/p.
Charge content of signal proportional to pulse area
Width of o/p of NINO can predict the center of multiple strips

Small i/p pulse



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Larger i/p pulse



The Experimental Setup





- Freon: Isobutane 95:5
- The muon triggered is obtained from coincidence of 3 plastic scintillators.
- Strip width is **1 cm.**
- Only one strip is triggered. 3 strips from either side have been studied.
- Altera Max 10 FPGA board.



The FPGA Based DAQ



The RTL diagram of the FPGA program



Results

Comparison of Oscilloscope and FPGA pulse widths

A 60 ns pulse (from a pulse generator) was measured by both MSO-4104B Oscilloscope and 500 MHz clock of FPGA.
2 ns offset was added to the

difference.



Difference of pulse widths measured in Oscilloscope and FPGA 1

Pulse Width at different Working Voltage



Width (ns)



 Signal width obtained from 1000 events.
The Center strip is triggered.
Tracking is better in avalanche mode.

Width Fraction of the n^{th} strip = $\frac{1}{\sum}$

Width of individual strips

Width of the nth Strip



Summary

- TOT measurement property of NINO has been exploited to find the center of the hits.
- □ A simple FPGA based DAQ for the cosmic muon imaging system.
- Use of PLL to use 2 different clks, one for event selection(50 MHz) and other for pulse width measurement (500 MHz).
- Pulse widths of the detector at avalanche and nearly streamer mode voltages have been studied.

Future Targets

- □ Calibration of o/p pulse and i/p charge has to be done.
- □ Use of thinner strips for better resolution
- Better FPGA clk for precise measurement,
- Utilize in the muon imaging system.

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Thank You

Back Ups

- Total signal width obtained from 1000 events.
- The Center strip is triggered.
- Tracking is better in avalanche mode.

Comparison of FPGA and Oscilloscope

VI curve and Efficiency

Different Resolutions

Lead on left Iron on right

- Altera/Intel[®] MAX[®] 10 FPGA 10M02SCE144C8G
 - 2000 Logic elements (LEs)
 - 108 Embedded memory (Kbits)
 - 12 User flash memory (KBytes)
 - Single Internal Configuration memory
 - 16 Embedded 18 x 18 multipliers
 - 2 General-purpose PLLs
 - 101 Maximum FPGA I/O pins
- Configuration Status and Set-Up Elements
 - On-board USB-Blaster compatible circuit for FPGA programming and debugging
- Expansion Header
 - Two 40-pin headers (GPIOs) provide 68 3.3V I/O pins
 - Two 5V power pins, two 3.3V power pins and four ground pins
 - One 6-pin header provides four 3.3V digital I/O pins one 3.3V power pin and one ground pin, compatible with Digilent Pmod[™] connector.
 - General User Interface Peripherals
 - 2 Yellow LEDs
 - o 2 Push-buttons
 - 2 7-Sement LED Displays
 - Clock System
 - On-board 50MHz clock oscillator
 - USB to UART Interface
 - CH340G based USB-UART converter for interfacing to a PC, including TxD and RxD LEDs