

GeV to TeV Connections

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Multimessenger Data Analysis in the Era of CTA

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electror





Non-thermal Gamma-ray Emission





High Energy Gamma-rays (aka GeV or HE) Very High Energy Gamma-rays (aka TeV or VHE)



The Fermi Observatory

Gamma-ray Burst Monitor (GBM): 8 keV to 40 MeV entire unocculted sky Large Area Telescope (LAT): 20 MeV to >1 TeV 20% of the sky, entire sky every 3 hours

Complementarity with targeted/pointed observations: Full sky, broad energy range, and public data



Fermi Gamma-ray Burst Monitor (GBM)





Fermi Large Area Telescope (LAT)



Anti-Coincidence Detector

charged particle separation

Tracker charged particles cause conversion $\gamma \rightarrow e^+e^$ direction

> **Trigger** read out: ~2 kHz down-link: ~400 Hz

Calorimeter Energy measurement





Fermi-LAT γ-ray sky

Extragalactic Sources

Active Galactic Nuclei. Starburst Galaxies...

Solar Flares + Terrestrial Gamma-ray Flashes

Local Sources

>1 GeV,

9 year map

+Supernova Remnants +Pulsar Wind Nebulae +Globular Clusters



Exotic and Transient Astrophysics

Pulsars

>5000 sources





Fermi-LAT Catalogs

- 4th Fermi PointSource Catalog (4FGL)
- 2nd GRB Catalog
- 2nd LAT AGN Catalog (2LAC)
- 3rd Pulsar Catalog
- 3rd Fermi High Energy Catalog (3FHL)
- 1st Fermi-LAT Low Energy Catalog (1FLE)
- Fermi-LAT Extended Source Catalogs (FEGS,FHES)
- ...https://fermi.gsfc.nasa.gov/ssc/data/access/lat/
- In the works: Solar Flare Catalog, 3LAC, transient



Fermi Senior Review

Fermi Primary Mission Objectives:

- Detect additional sGRB-GW counterparts (Disentangle emission structure, dynamics and viewing geometry of NS mergers)
- Use sGRB-GW time delays as probes of cosmology, fundamental physics and neutron star physics
- Resolve emission mechanisms in blazers by finding gamma-ray flares in coincidence with high-energy neutrinos





Fermi Senior Review

Fermi Primary Mission Objectives:

- Resolve physical processes in gammaray novae and other galactic binaries by measuring simultaneous multiwavelength light curves.
- Explore and interpret the growing diversity of gamma-ray pulsar systems including adding new MSPs to PTAs
- Identify new temporal behavior only measurable with long baselines such as binary SMBHs.







CTA Key Science Programs

CTA key science programs (KSP) overlap completely with Fermi science **CTA Science book mentions Fermi 108 times**

> The full sky capability and overlapping energy range make Fermi-LAT and CTA unique observing partners

>10 GeV, 3FHL map

**L. Tibaldo counted 11



Galactic Program



Galactic Center Observations

High energy diffuse emission complementary to Fermi-LAT observations <100 GeV





Galactic Program

Galactic Plane Survey

Overlap with 2FHL sources: 70 potentially observable sources (SNR, PWNe)



arXiv:1709.07997 13





Transient Program

Gamma-ray Bursts

Fermi-LAT observations illustrate potential for CTA detections



Simulated CTA light curve of GRB 080916C

Extrapolation from Fermi-LAT light curve above 0.1 GeV R. Caputo , NASA/GSFC | CTA in Sesto 2019 Figure 9.1 (arXiv:1709.07997)



GRB190114c



First time detection of a GRB at sub-TeV energies; MAGIC detects the GRB 190114C

ATel #12390; Razmik Mirzoyan on behalf of the MAGIC Collaboration on 15 Jan 2019; 01:03 UT

Credential Certification: Razmik Mirzoyan (Razmik.Mirzoyan@mpp.mpg.de)

Subjects: Gamma Ray, >GeV, TeV, VHE, Request for Observations, Gamma-Ray Burst

Referred to by ATel #: 12395

The MAGIC telescopes performed a rapid follow-up observation of GRB 190114C (Gropp et al. GCN 23688; Tyurina et al., GCN 23690, de Ugarte Postigo et al., GCN 23692, Lipunov et al GCN 23693, Selsing et al. GCN 23695). This observation was triggered by the Swift-BAT alert we started observing at about 50s after Swift T0: 20:57:03.19. The MAGIC real-time analysis shows a significance >20 sigma in the first 20 min of observations (starting at T0+50s) for energies >300GeV. The relatively high detection threshold is due to the large zenith angle of observations (>60 degrees) and the presence of partial Moon. Given the brightness of the event MAGIC will continue the observation of GRB 190114C until it is observable tonight and also in the next days. We strongly encourage follow-up observations by other instruments. The MAGIC contact persons for these observations are R. Mirzoyan (Razmik Mirzoyan@mpp.mpg.de) and K Noda (nodak@icrr.u-tokyo.ac.jp). MAGIC is a system of two 17m-diameter Imaging Atmospheric Cherenkov Telescopes located at the Observatory Roque de los Muchachos on the Canary island La Palma, Spain, and designed to perform gamma-ray astronomy in the energy range from 50 GeV to greater than 50 TeV.

Cosmic Ray Spectra of Various Experiments





What is creating these ultra-high energy particles?

Multimessenger Program

Neutrinos are the smoking gun signature for hadronic acceleration.

proton.



Fermi Gamma-ray Space Telescope

> IceCube Neutrino Observatory

IACTs

Photon

Neutrino

Goal: Find neutrinos coming from an extra-galactic source

Courtesy: A. Franckowiak





Variable Fermi-LAT γ-ray sky

Fermi-LAT observes over 2000 blazars

>300 MeV over 3 months in 2008



Variable Fermi-LAT **y**-ray sky





IceCube starts issuing Transient Alerts



R. Caputo , NASA/GSFC | CTA in Sesto 2019 Fermi-LAT Coll., ApJ 846, 2017, Video

207 107 22012

102014

Fermi-LAT Coll., ApJ 846, 2017, Video credits: Matteo Giomi, Fermi-LAT Collaboration 20

1012011

AUD 2010





Spectra of TXS 0506+56 across wavelengths and messengers



Joint detection of TXS 0506+56 with Fermi-LAT and MAGIC in the flaring state made evidence of hadronic production in the particle jet possible



Leptonic, *Hadronic*... **Both**?



CTA can give insight on particle jet composition



Keivani et al. (2018)

Gao et al. (2018)

17 August 2017







Time from merger (seconds)





The Multimessenger Age

The unbound neutron star material undergoes r-process nucleosynthesis, expands, decays, and emits thermal radiation as a kilonova t = hours to weeks

The kilonova emission evolves from UV \rightarrow Optical \rightarrow IR on a descale of hours to weeks and reveals the details of the early merger and the ejecta velocity, mass, and composition







R. Capulo, NASAVOSEC | CIA III Sesio 2019



What did we learn?







Speed of Gravity

$$-3 \times 10^{-15} \leqslant rac{\Delta v}{v_{\mathrm{EM}}} \leqslant +7 \times 10^{-16}$$

Tested Equivalence Principle of Gravity: Ruled out many theories of modified gravity to explain dark matter







Start of O3: April 1 >10 mergers

BNS merger S190510g



Continual GWs: "Neasurements of pulsar glitches (Vela)









Electromagnetic Spectrum







Gravitational Wave Spectrum







Neutríno Spectrum

Mystery of Missing Mass



1930s- Zwicky, others

Coma cluster of galaxies: only small % mass from luminous matter

Dunkle Materie

31

F. Zwicky, Astrophys. J. 86, 217 (1937)





Mystery of Missing Mass



1970s- Rubin, Ford, Thonnard Galactic Rotation Curves

'ouTube: https://www.youtube.com/watch?v=Omjx3OKAGDo http://cse.ssl.berkeley.edu/bmendez/ay10/2002/notes/lec18.htm





Modified Gravity?



Test of Weak Equivalence principle



Boran et al., PRD 97, 041501 (2018),

Gravitons and photons travel in space-time in the same way

*SN1987a found the same thing for neutrinos and photons





What do we know?





stable

not p/n (baryons) Particle(s)

not hot neutral constraints from CMB, primordial nucleosynthesis





Portrait of a Candidate







Dark Matter Program

Searches for thermal WIMPs









Ongoing analyses to incorporate CTA observational strategy, better IRFs and more accurate diffuse emission





Dark Matter Program

Future sensitivities for thermal WIMPs





Fermi Tools



- Fermipy
 - Open-source python framework
 - High-level interface to Fermitools
 - Available on GitHub
 - See talk by F. De Palma Tuesday
- 3ML
 - Open-source python framework for multi-wavelength/multimessenger analysis
 - Available on GitHub
 - See talk by N. Omodei Tuesday
- GSPEC
 - Python based GUI for analyzing Fermi GBM data
 - Seamless interface to XPEC



Summary



- Quick tour of *some* of the Fermi/CTA Synergies
 - GeV to TeV... with some keV/MeV sprinkled in
 - Impossible to overstate the depth of connection between CTA and Fermi science
- Unprecedented understanding of fundamental processes producing the most extreme events



In Conclusion...



