

# Preventivi 2020 : Gruppo II

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Alberto Garfagnini

Consiglio di Sezione, Padova, 10 luglio 2019



**Istituto Nazionale di Fisica Nucleare**



## Padova CTA/MAGIC group (2019- 2020 ) anagrafica



Magic = 10 firme, CTA = 14 firme Tot\_FTE = 12

**Mose' Mariotti**  
group leader



**Alessandro de Angelis** PO



**Michele Doro**  
PA



**Eugenio Bottacini**  
RTDB



**Elsa Prandini**  
RTDA



**Riccardo Rando**  
RC



**Manuela Mallamaci**  
postdoc



**Ruben Lopez**  
postdoc



**Simona Palano**  
Assegno



**Denis Bastieri**  
PA



**Giovanni Busetto**  
PO



**Giampiero Naletto**  
PA



**Alessia Spolon**  
PhD



**Cerdc Perennes**  
Pos doc



**Luca Foffano**  
PhD



**Elsa bernardini**  
PA



**Adriano Pepato**  
Engineer



**Sandro Ventura**  
Ricercaore



**Daniele Corti**  
Electronic

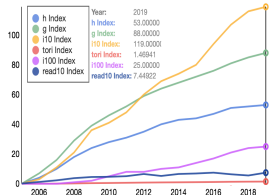
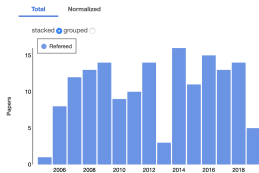




## MAGIC recent scientific results and highlights

- MAGIC gode di ottima salute, al momento, grazie alla manutenzione e comprensione dello strumento, possiede la migliore sensibilità ed affidabilità di sempre!
- Possiede una rilevante ed eccellente produzione scientifica testimoniata da indicatori bibliometrici,

	Totale	Referred
Number of papers	158	158
Normalized paper count	1.0	1.0



**h-index**  
The H-index is the largest number H such that H publications have at least H citations.

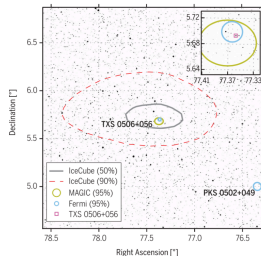
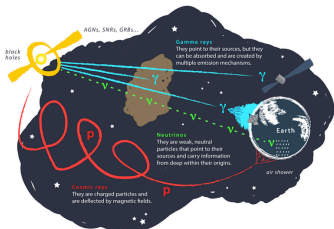
**i100-index**  
The i100-index is the number of publications with at least 100 citations.



## MAGIC recent scientific results and highlights

Nel 2018 rilevante la misura della controparte gamma al TeV in un follow-up di allerta neutrino da icecube, che ha sancito la prima probabile “sorgente” di neutrini cosmici

### MAGIC 2018: TRANSIENTS #1 – THE BLAZAR TXS0506+056





## MAGIC recent scientific results and highlights

Nel 2019, per la prima volta osservato un GRB da terra: dati di altissima qualità in regime quasi “background free”. Molte sorprese: press release e in progress (nature paper results under embargo)

### MAGIC 2018: TRANSIENTS #2 – GRB 190114C

[ [Previous](#) | [Next](#) | [ADS](#) ]

#### First time detection of a GRB at sub-TeV energies; MAGIC detects the GRB 190114C

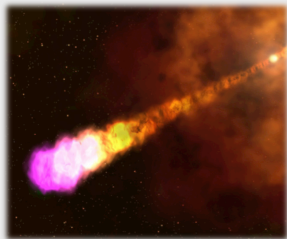
ATel #12390: [Razmik Mirzoyan on behalf of the MAGIC Collaboration](#)  
on 15 Jan 2019; 01:03 UT  
Credential Certification: Razmik Mirzoyan (Razmik.Mirzoyan@mpp.mpg.de)

Subjects: Gamma Ray, >GeV, TeV, VHE, Request for Observations, Gamma-Ray Burst

Referred to by ATel #: 12395, 12475



The MAGIC telescopes performed a rapid follow-up observation of GRB 190114C (Gropp et al., GCN 23688; Tyurina et al., GCN 23690; de Ugarte Postigo et al., GCN 23692; Lipunov et al., GCN 23693; Seising et al., GCN 23695). This observation was triggered by the Swift-BAT alert; we started observing at about 50s after Swift T0: 20:57:03.19. The MAGIC real-time analysis shows a significance >20 sigma in the first 20 min of observations (starting at T0+50s) for energies >300GeV. The relatively high detection threshold is due to the large zenith angle of observations (>60 degrees) and the presence of partial Moon. Given the brightness of the event, MAGIC will continue the observation of GRB 190114C until it is observable tonight and also in the next days. We strongly encourage follow-up observations by other instruments. The MAGIC contact persons for these observations are R. Mirzoyan (Razmik.Mirzoyan@mpp.mpg.de) and K. Noda (nodak@icrr.a.tokyo.ac.jp). MAGIC is a system of two 17m-diameter Imaging Atmospheric Cherenkov Telescopes located at the Observatorio Roque de los Muchachos on the Canary island La Palma, Spain, and designed to perform gamma-ray astronomy in the energy range from 50 GeV to greater than 50 TeV.





## Attività 2020: MAGIC

### Presa Dati

- 2+1 turni di presa dati



### Analisi dati partecipazione ai Gruppi scientifici

- AGN Physics working group



- Astroparticle and fundamental



- Galactic



CONVENOR



- Transient: neutrino, GRB, GW



Neutrino PI



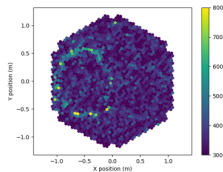
CB member



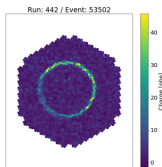
## CTA-LST Status/activity

- LST in commissioning: turni di messa a punto e inizio presa dati
- Workshop sul software ed analisi di LST a Padova organizzato dalla sezione presso LNL. Rubèn Lopez responsabile e coordinatore software ed analisi di LST.
- Graduale aumento qualità dei dati: presa dati sistematica con shifts: nella prima fase seguendo le osservazioni MAGIC

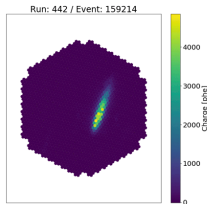
Dicembre 2018



Aprile 2019



Maggio 2019





## CTA-LST activity 2020



- In preparazione gare per i prossimi 3 telescopi: INFN PD meccanica funi, INFN TS-UD Calibration box,
- Richieste economiche PD 2020: Funi e meccanica ancillare per funi, completamento del cluster R&D a SiPm, Presa dati nel sito ed analisi combinata con dati MAGIC



**Design of a SiPM-based cluster for the Large Size Telescope camera of CTA**

Massimo Belloni<sup>1</sup>, Denise Costi<sup>2</sup>, Luigi Lestini<sup>3</sup>, Muel Marotti<sup>4</sup>, Roberto Pecher<sup>5</sup>, Eugenio Belloni<sup>6</sup>, Giovanni Basso<sup>7</sup>, Riccardo De Angelis<sup>1,4</sup>, Paolo Di Pietro<sup>8</sup>, Michele Dotti<sup>9</sup>, Elisa Pirelli<sup>10</sup>, Piero Valania<sup>11</sup>, Carlo Francesco Ugolini<sup>12</sup>

<sup>1</sup>INFN Padova, <sup>2</sup>INFN Padova, <sup>3</sup>Università di Padova, <sup>4</sup>Università di Udine, <sup>5</sup>IST and I.P. Labor, <sup>6</sup>INFN Torino, <sup>7</sup>INFN GNS, <sup>8</sup>INFN GNS, <sup>9</sup>Università di Torino, <sup>10</sup>INFN GNS, <sup>11</sup>INFN GNS, <sup>12</sup>INFN GNS

A Silicon Photomultiplier (SiPM)-based photodetector will be built to be possibly used in the Large Size Telescope (LST) camera of the Cherenkov Telescope Array (CTA). It has been designed to match the size of the standard Photomultiplier Tube (PMT) cluster unit and to be compatible with mechanical, electronics and front plane optics of the first LST camera. Here, we describe the overall SiPM cluster design along with the main differences with respect to the currently used PMT cluster unit. The first electronics of the SiPM panel and its layout are also presented. In order to derive the best working condition for the final unit, we measured the SiPM performance in terms of gain, photodetector efficiency and cross-talk. A panel, a unit of 14 SiPMs, has been built. We will discuss also some preliminary results regarding this device and we will highlight the future steps of this project.



## Richieste finanziarie servizi di Sezione 2020

Struttura	missioni	consumo	altri_cons	trasporti	manutenzione	inventario	licenze-SW	apparati	spservizi	TOTALI
BA										
CT										
GSGC										
NA										
PD	66.00	24.00		4.00		12.00		170.00	31.00	307.00

### RICHIESTE Servizi di Sezione CTA 2020

#### LE 2 mU

- Sviluppo e produzione schede per interfaccia per cluster SiPm 2,0 mese/uomo

#### Bellato 4 mU

- Nuova versione PCB sensori SiPm LST (sviluppo produzione e test) 2,0 mese/uomo
- Montaggio dei pixel nel cluster LST e test funzionali del sistema 2,0 mese/uomo

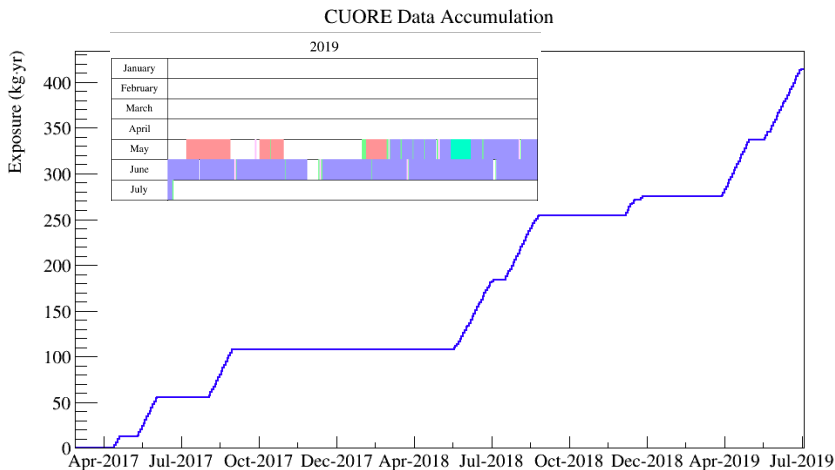
#### OM 3.5 mU

- Supporto manutenzione specchi MAGIC 1,5 mese/uomo
- Supporto per realizzazione meccanica dei cluster SiPm LST 1,0 mese/uomo
- Supporto realizzazione meccanica "ancellare" Funi 2,0 mese/uomo

#### UT 3 mU

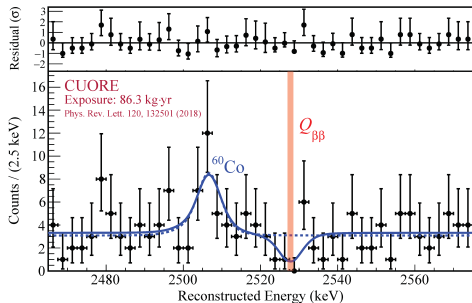
- Supporto Gara acquisto funi, produzione meccanica "ancellare" Funi 2,0 mese/uomo

**Cuore2019:** finalmente un buon duty time, si parte per i 5 anni di data taking da specifiche di progetto (tranne 7keV) Risolti i molteplici problemi della DCS che limitavano il duty time e mettevano in pericolo il criostato.



CUORE nuovi risultati sotto embargo TAUP 2019, da aprire la scatola «unblind», ci sono però i vecchi.

BACKGROUND Model **EXCELLENT!!**

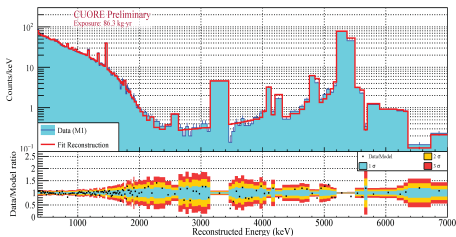


**Up:** reconstructed  $^{208}\text{Tl}$  2516 keV line in the calibration spectrum with residuals with data points (black) and the reconstructed line-shape (red). The dotted blue lines are the components of the line shape corresponding to different physical processes.

**Down:** the best fit in the ROI. The value of the  $Q_{\beta\beta}$  is marked in

**Top:** M10 spectrum (blue) and the fit to the spectrum, built out of different simulated background components (red). The variable binning has been chosen to contain peaks into a single bin to avoid dependence on details of the peak shape while having good resolution of the model in the continuum region.

**Bottom:** ratio of the data and reconstructed model with 1s, 2s and 3s error bars. The continuum is well described while we have moderate disagreement in the heights of a few peaks.



- Optimum Trigger (OT) has been successfully implemented in CUORE; Antonio Branca 30% september 1)talk CUORE EPS-hep 2019 july 2)Poster OT Taup 2019
- The algorithm can be used either online or on data already on tape;
- The software tools needed to apply the trigger offline (retriggering software) have been updated to allow retriggering of a data-set in few hours (6/7 weeks of data can be retriggered in about 8 hours);
- The OT is applied for the data production for TAUP results:
  - All data-sets that were acquired with the standard derivative trigger have been retriggered in few days;
  - New data will be retriggered at the closing of each data-set;
- The OT data allow to lower the energy threshold:
  - Detection efficiency is obtained from pulsar events, but also a new method based on fake injected pulses has been implemented;
  - Studies are ongoing to determine the best analysis energy thresholds exploiting pulses shape parameters;
  - neutrino-less double beta decay analysis:
    - Threshold from 150 keV, used for PRL-2018 result, down to 40 keV;
    - Background contribution from Compton scattering in more than one crystal will be discarded;
  - Dark-Matter and low-energy nuclear processes: require thresholds of at least 10 keV;

Cupid-RD finito, alle prossime CNS2 si chiede apertura sigla, che in qualche anno si fonderà con CUORE, di seguito informazioni dal pre-CDR di luglio 2019, in corso di review finale

**CUPID** (CUORE Upgrade with Particle IDentification) starts from the CUORE technical expertise, material selection, background model, and excellent energy resolution, and combines it with the capability to eliminate the backgrounds from  $\alpha$  particles using scintillating crystal. In addition, the high Q-value of  $0\nu\beta\beta$  in  $^{100}\text{Mo}$  naturally suppresses the gamma backgrounds by over an order of magnitude, compared to CUORE. We demonstrate that the scintillating bolometer technology based on  $\text{Li}_2\text{MoO}_4$  crystals highly enriched in  $^{100}\text{Mo}$  is scalable to ton-scale isotopic masses. An experiment deployed in the existing CUORE cryostat would have a discovery potential covering the entire inverted hierarchy region of neutrino masses.

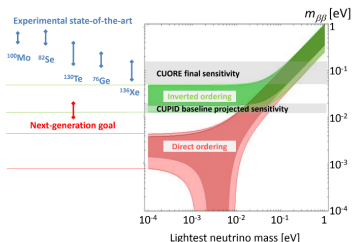


Figure 2: The effective Majorana mass  $m_{\beta\beta}$  as a function of the lightest neutrino mass provides the parameter space typically used to compare  $0\nu\beta\beta$  decay experiments. The experimental state-of-the-art and the goal of next-generation experiments are shown on the left. The final sensitivity of CUORE and the projected sensitivity of CUPID baseline are also reported.

Table 1: Main parameters of the conservative baseline CUPID detector design.

Parameter	Baseline
Crystal	$\text{Li}_2\text{MoO}_4$
Crystal size	$\varnothing 50 \text{ mm} \times \text{h} 50 \text{ mm}$
Crystal mass (g)	308
Number of crystals	1534
Number of light detectors	1652
Detector mass (kg)	472
$^{100}\text{Mo}$ mass (kg)	253
Energy resolution FWHM (keV)	5
Background index (counts/(keV·kg·yr))	$10^{-4}$
Containment efficiency	79%
Selection efficiency	90%
Lifetime	10 years
Half-life limit sensitivity (90% C.L.)	$1.5 \times 10^{27} \text{ y}$
Half-life discovery sensitivity ( $3\sigma$ )	$1.1 \times 10^{27} \text{ y}$
$m_{\beta\beta}$ limit sensitivity (90% C.L.)	10 – 17 meV
$m_{\beta\beta}$ discovery sensitivity ( $3\sigma$ )	12 – 20 meV

## Richieste economiche Cuore-Pd 2020

Anagrafica: Ric. 0,5/2 FTE: AB e ML(non firmatario) Tecn. 0.6/1 FTE: LT

a Novembre 2018 AB fa finito il suo biennio nel Vetting Board

Capitolo	Descrizione	Parziali		Totale
		Richiesta	SJ	Richieste
	1. Turni presso LNGS 2 FTE	8.00		
	2. 2 meeting di collaborazione in Italia + riunioni PI	2.00		
	3. Meeting di analisi dati	2.00		
	4. 1 meeting di collaborazione all'estero (USA)	2.00		
<b>MISSIONI</b>	5. Conferenza estero	2.00		
	6. Attività sperimentale sala C e studio di mercato di nuovi refrigeratori x CUPID sala A (Leiden, Oxford, Blueforce)	5.00		
				<b>21.50</b>
<b>CONSUMO</b>	1. Flussostato per He3, gadget per pannello DU	2.00		
				<b>2.00</b>
	Non si richiede attività di servizio			
				Totale richiesta 23,5 k€

- Dune** • 1069 collaborators from 177 institutions in 31 countries  
• 578 faculty/scientists, 184 postdocs, 109 engineers, 198 PhD students

Major achievement in 2018: ProtoDUNE performances

Major milestone in 2019: **FD TDR completed (including ND CDR summary)**

Planning for the first detector module to be SP and the second to be DP ('2+1+1 model' with 3<sup>rd</sup> module SP and 4<sup>th</sup> 'module of opportunity').

Start of Detector 1 Installation: August 2024

Start of Detector 2 Installation: August 2025

Start beam operations in 2026

*Sigla INFN: Nu\_at\_FNAL, resp. naz. S. Bertolucci*

- Impegno per il Near Detector (utilizzo di KLOE, magnet plus EM)
- Impegno per il Far Detector (Photodetection)
- Impegno per SBN (responsabilità del top CRT, Cosmic Ray Tagger, analysis contribution)

**➔ APPROVED by DUNE-DOE-INFN in the last days !!!**

*Preventivi in corso di elaborazione.*

*Richieste ai servizi – Padova : 1 mese off.mecc. per installazione CRT a FNAL*



# ENUBET

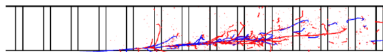


This project has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme (grant agreement No 681647).

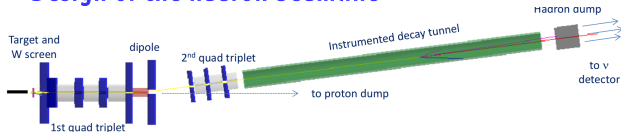
<http://enubet.pd.infn.it>

- Instrumented decay tunnel
- $K^+ \rightarrow e^+ \nu_e n^0 \rightarrow$  large angle  $e^+$
- $\nu_e$  flux prediction =  $e^+$  counting
- % level precision on  $\nu_e$  flux

R&D for the instrumentation of the decay tunnel



Design of the hadron beamline



**ERC Cons. Grant (2016-2021) [ENUBET, PI: A. Longhin, UNIPD (host)+INFN]**  
**Grant MIUR – Bando FARE (2017-2021) [NuTech]**  
**NEW! Approvazione come esp. CERN “Neutrino Platform” [NP06/ENUBET]**

## The ENUBET collaboration

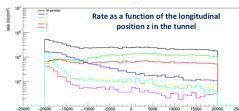


60 physicists,  
12 institutions

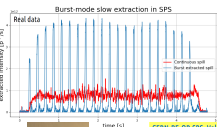


### INFN-PD

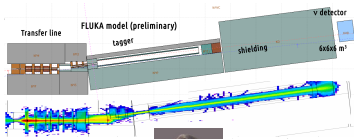
Contributo **cruciale**: coordinamento (PI), simulazione del tunnel instrumentato, della beamline (2 art. 36), sviluppo di sistemi di estrazione dei protoni al CERN/SPS e sviluppo di digitizer custom (1 PhD).



F. Pupilli



M. Pari



G. Brunetti

# ENUBET-Padova, anagrafica



- Il progetto e' **interamente** finanziato da **ERC**
- sigla di supporto in **CSN2 (ENUBET\_2)** per piccole spese
  - 1) non previste nel Grant Agreement (i.e. missioni laureandi, PhD)
  - 2) inventariabili per cui EU rifonde una frazione del costo corretta per ammortamento a fine progetto (difficile da rendicontare).
- **Gruppo e anagrafiche:**
  - **EU (time-sheets):** **A. Longhin** (70%), **F. Pupilli** (100% art. 36 ENUBET), **G. Brunetti** (100% art. 36 ENUBET), **G. Collazuol** (10%).
  - **ENUBET\_2:** **E. Conti** (40%), **F. Dal Corso** (30%), **M. Mezzetto** (10%), **M. Pari** (PhD, 70%)

2019 highlights

# 2019 highlights /4



## Riconoscimento nella CERN Neutrino Platform

Proposta sottomessa alla Neutrino Platform Call nel Dicembre 2018:

- ✓ Sviluppo della tecnologia dei fasci monitorati (ENUBET) e taggati (NUTECH) grazie all'expertise della divisione acceleratori CERN (proton extraction, transfer line, beam dump, target area)
- ✓ Valutare implementazione al CERN con ProtoDUNE-Single/Double Phase
- ✓ Coagulare l'interesse della comunità (soprattutto T2K/HK e DUNE) per un esperimento short baseline dedicato alle sezioni d'urto e alla riduzione delle sistematiche nei fasci long-baseline
  
- ✓ **Proposta:**  
CERN-SPSC-2018-034 ; SPSC-I-248 (31/10/18).
- ✓ **Approvazione SPSC e Research Board:**
  - 1/2019 e 3/2019
- ✓ **Memorandum of Understanding:**
  - in progress

The physics case of the **ENUBET** project and the exciting possibilities of a tagged neutrino beam are recognized by the SPSC. The committee recognizes the technological development for a neutrino beam without a horn using a quadrupole-based solution, and appreciates the close collaboration of the ENUBET collaboration with the CERN accelerator sector. The SPSC supports the proposed programme, and welcomes the opportunity to continue reviewing the experiment; test-beam requests will be considered via the standard annual procedure. **The Research Board approved the participation of ENUBET in the Neutrino Platform, with reference NP06, on the understanding that**

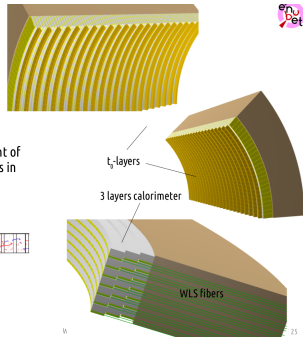
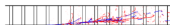


# Attività / richieste 2020

- Completare il **design della beamline**
  - **Update delle physics performance** includendo il **monitoraggio dei muoni** e la risoluzione energetica intrinseca (10%) dei fasci narrow band
  - **Prototipo ingegnerizzato** del dimostratore. Test beam a **FNAL**.
  - **Procurement** del materiale per il dimostratore e **design meccanico**
- 
- **Servizio progettazione** per il dimostratore: **3-6 m.u.** (da contatti preliminari di persona)
  - **Officina meccanica** (prototipo intermedio per test beam): **2 m.u.**
  - **Preventivi CSN-II 2019**  
Missioni: 3 KEUR (PhD)

## The tagger demonstrator

- Length ~ 3 m
- allows the containment of shallow angle particles in realistic conditions
- Fraction of  $\phi$
- Due by 2021



## ET\_Italia: Sezione di Padova

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**Attività ET\_Italia:** le attività hardware per il 2020 saranno prevalentemente concentrate sulla caratterizzazione e per sostenere la candidatura del sito **Sos Enattos** (Sardegna). Padova in questo non è coinvolta. La scelta del sito si dovrebbe fare nel 2021-22

### Attività Sezione di Padova

Studio del caso scientifico

- Cosmologia
- Popolazione e modelli

Enabling technologies

- Misura dell'assorbimento (a 1550 nm) dei substrati in Si a temperature criogeniche.

### Richiesta economica 2019

**5 keuro:** mono cristalli di silicio ad alta purezza ed ottiche per 1550 nm

### Richiesta alla sezione

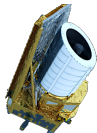
- Officina meccanica: 1 mesi uomo

## ET\_Italia: Anagrafica Sezione di Padova

Cosmologia	S. Matarrese	PO-Unipd	10%	M. Mapelli	PA-Unipd	10%
	N. Bartolo	PA-Unipd	10%	U. Di Carlo	PhD-Unipd	10%
	M. Liguori	Ric-Unipd	10%	N. Giacobbo	PhD-Unipd	10%
	D. Bertacca	RTDA	10%	M. Pasquato	Astrofit-fellow	10%
	M. Peloso	PA-Unipd	10%	M. Spera	MC-fellow	10%
	G. Orlando	PhD-Unipd	10%	A. Ballone	Post Doc	10%
	A. Ganz	PhD-Unipd	10%			
Hardware	J.P. Zendri (**)	PRic-INFN	40%	L. Conti	Ric-INFN	20%
	G. Ciani	RTDB	20%	M. Bazzan	Ric-Unipd	20%
	C. Lazzaro	RTDA	20%	M. Pegoraro	Tecn. INFN	10%

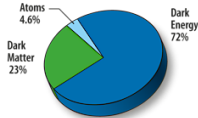
(\*\*) Resp locale

**TOTALE 2.6 FTE**



# Euclid Space Mission

## Dark Energy Study



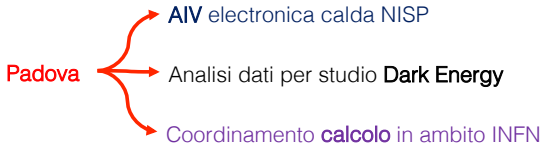
**Euclid** è una missione spaziale ESA di classe-M dedicata allo studio:

- origine dell'espansione accelerata dell'Universo
- Dark Energy e Gravità su larga scala

Status missione:

- completamento degli strumenti NISP e VIS, consegna ritardata a feb. 2020
- **lancio alla fine del 2022**, critico per fine contratto del lanciatore Soyuz
- 6 anni di presa dati

Rinnovato **contratto ASI 2019-2022** (2 RTD x 3 anni) per attività di integrazione e validazione (AIV) elettronica calda dello strumento NISP





## NISP: Integrazione e validazione Warm Electronics

[E.Borsato, S.Dusini, F.Laudisio, A.Renzi, C.Sirignano, S.Ventura]

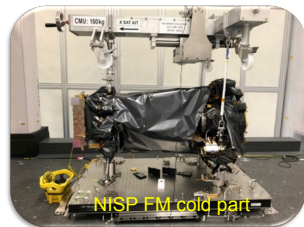
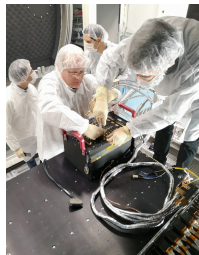
**Integrazione di NISP-EM** (Engineering Model) con test a LAM (Marsiglia) in Thermo-Vuoto (TV). Aprile

- Importante contributo Italiano (INFN e INAF) per integrazione elettrica, Test-scripts, esecuzione test
- Problemi durante EMC → Test aggiuntivo sett. prossima

### Integrazione e Validazione NISP Warm Electronics FM

- Due ulteriori mesi di ritardo nella consegna di DPU FM
- Per assorbire parte del ritardo molta dell'attività è spostata a LAM
- 3 campagne di test
  - Validazione WE ASW FM [resp. INFN]
  - NISP quasi-FM TV per la caratterizzazione dell'ottica e dei rivelatori (Ott. 2019) [supporto]
  - NISP FM TV test finali di performance dello strumento (Gen.-Feb. 2020) [supporto]

Molta pressione per il mantenere la scadenza Feb. 2020 (era Dic 2019)



# Euclid analysis: Dark Energy

## Padova

- **Co-leadership** del work package **Euclid WP3 del Scientific Working Group** di cross-correlazione con CMB [ **Alessandro Renzi** ]
- Validazione algoritmi per il calcolo del Power Spectrum Galassie [ **Alessandro Renzi** ]
- Studio dei vuoti cosmici come tracer per Dark Energy, gravità modificata, masse dei neutrini, **Giovanni Verza** PhD student, relatori **Chiara Sirignano e Alessandro Renzi**

## Coordinamento tra sezioni

- WG di fisica tra le sezioni INFN (Bo, Ge, Mi, Pd) che partecipano a Euclid (coord. **Luca Stanco**)
- Riunioni bi-settimanali:
  - Voids-lensing cross correlation,
  - Voids and Large Scale Structure (LSS)
  - LSS-CMB cross correlation.
  - Supporto a OU-SPE (ricostruzione spettri NISP)
- Collaborazione con SISSA sui Vuoti Cosmici

## Calcolo per analisi dati Euclid

Risorse calcolo HPC/MPI per le attività di analisi

- A. **Cluster ZEFIRO (INFN-Pi): small/medium size HPC** cluster con High-speed low-latency communication per **sviluppo e ottimizzazione dei codici ibridi MPI/OpenMP**
- Development and validation of the **Power Spectrum** code for galaxy clustering.
  - Development of the cross-correlation estimator and likelihood using Planck CMB and Euclid galaxies for neutrino parameters estimation.
  - Development of a full 3D power spectrum including Redshift Space Distortions and likelihood of the neutrino parameters estimation.
  - 128 CPU con 2048 computational cores, 8 GB RAM/core, ~ 300 TB disk
- B. **CINECA: simulazioni N-body** cosmologiche di medie dimensioni
- Almeno 1000 cores, ~8 GB RAM/core, high-speed low-latency communication
  - 3.5 M core x hour/year
  - ~500 TB disk storage per 3 – 4 anni (post processing e analisi)

## Anagrafica e Richieste

Anagrafica Padova stabile a **4.8 FTE**: PHD student Fulvio Laudision → Giovanni Verza

Nome	Contratto	Qualifica	Aff. CSN	Percentuale
Bartolo Nicola	Associato	Prof. Associato	4	10
Benevento Gianpaolo	Associato	Dottorando	4	10
Bertacca Daniele	Associato	Ricercatore DFA	4	10
Dusini Stefano	Dipendente	Ricercatore INFN	2	70
Karagiannis Dionysios	Associato	Assegnista	4	10
Liguori Michele	Associato	Ricercatore DFA	4	10
Matarrese Sabino	Associato	Prof. Ordinario	4	10
Naletto Giampiero	Associato	Prof. Associato	2	10
Renzi Alessandro	Dipendente	Ricercatore INFN	2	100
Sirignano Chiara	Associato	Ricercatore DFA	2	70
Stanco Luca	Dipendente	Dirigente Ricerca	2	60
Ventura Sandro	Dipendente	Primo Tecnologo	2	10
Verza Giovanni	Associato	Dottorando	2	100
<b>Totale</b>				<b>4,8</b>

L.Stanco: resp. Nazionale

A.Renzi: Co-lead Euclid WP3 (SWG)

C.Sirignano: Euclid Diversity Committee

S.Dusini: NISP Deputy AIV Manager e  
Deputy NISP Test Manager

A **livello nazionale Euclid cresce** anche quest'anno con ingresso del gruppo di **Ferrara e Torino**

Richieste finanziarie:

- Calo attività NISP anche se coda test NISP-FM (2020), supporto integrazione e test satellite (fino a 2022), commissioning in volo (2023).
- Cresce il peso delle attività di analisi (workshop, comunicazioni a conferenze)

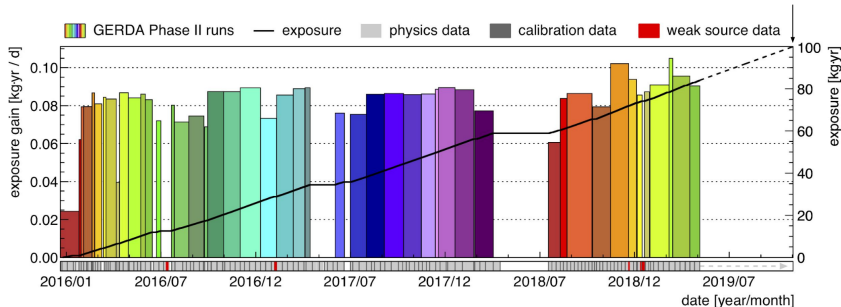
# GERDA Fase II



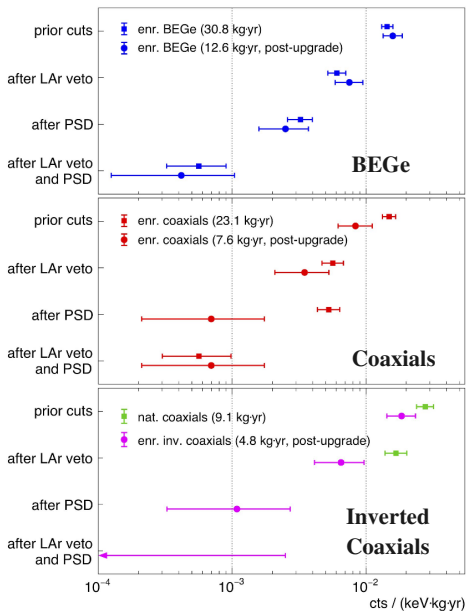
## data taking

02-Nov-2019

11:31:32



- La presa dati sta procedendo in maniera regolare
- ◆ Si prevede di raggiungere l'esposizione prevista di 100 kg·yr ai primi di novembre di quest'anno
- A partire da gennaio 2020 l'infrastruttura di GERDA passerà a LEGEND-200



## GERDA Fase II

- BEGe e Coassiali:  
BI <  $10^{-3}$  cts/keV·kg·yr
- I nuovi rivelatori (Inverted Coaxial) mostrano BI in linea con i BEGe
- La pubblicazione finale dei dati di GERDA verrà fatta in primavera del prossimo anno
- Fra breve verranno pubblicati i dati mostrati a NEUTRINO 2018

## LEGEND-200



- **Inizio commissioning ottobre 2020**
- La quantità di Ge arricchita nell'isotopo  $^{76}\text{Ge}$  necessaria per raggiungere i 200 kg è già acquisita
- La produzione dei cristalli (Inverted Coaxial) è iniziata
- I primi diodi verranno prodotti in agosto 2019
- Definiti il FE e il preamplificatore per i diodi
- finalizzato il LAr veto
- In gennaio 2020 primi lavori per l' integrazione tra rivelatori-FE-preamplificatore-DAQ all' interno del setup di GERDA (cioè in condizioni reali).

## **Attività di Padova nel 2020**

### GERDA

- analisi dati (background modelling, majoroni, ....)
- vari duties all'interno della Collaborazione

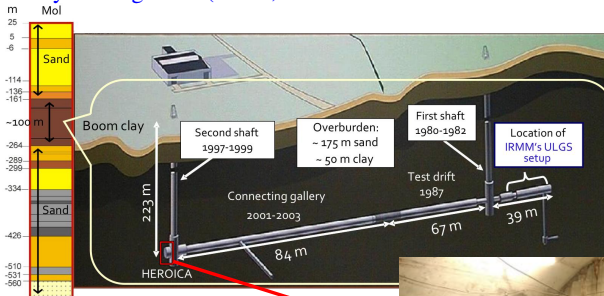
### LEGEND-200

- lavori sul DAQ, slow control
- partecipazione ai tests sui nuovi rivelatori (Inverted Coaxial) ad HADES (Belgio)
- sviluppo di MC
- vari duties all' interno della Collaborazione



**HADES (High Activity Disposal Experimental Site)**

Facility Underground (a Mol, vicino al sito della Canberra di Olen).



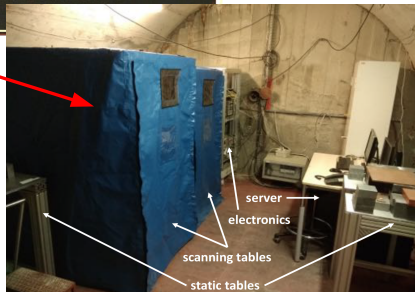
HADES è sotto la direzione dell'ente SCK-EN

**Setup sperimentale in HADES:**

2 tavoli per lo scanning della superficie laterale dei rivelatori

+

2 tavoli per lo studio delle prestazioni in termini di pulse shape discrimination



## Richieste finanziarie

## Richieste ai Servizi della Sezione

## Anagrafica

Missioni	31.0 keuro
Consumo	4.0 keuro
Inventario	10.0 keuro
Apparati	250.0 keuro

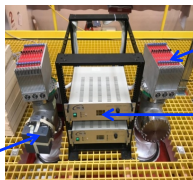


Officina Meccanica	0.5 m.u.
Progettazione Mecc.	0.5 m.u.
Officina elettronica	0.5 m.u.
Calcolo e reti	2 m.u.

Bettini A.	PO	0%
Brugnera R.	PA	70%
Garfagnini A.	PA	30%
Lippi I.	INFN	60%
Pertoldi	dottorando	100%
Sada C.	PA	60%
Von Sturm K.	Assegnista	100%

## ICARUS: Technical status -1

- Progress since January 2019:
  - ✓ All flanged connections, ports, valves, feedthrough flanges on top of two cold vessels were installed and tightened.
  - ✓ All internal connections of wire chambers, PMTs, HV system, optical fibers, and slow controls verified before the tightening
  - ✓ After closing all the apertures, when the internal volumes became dark, all PMT turned on & checked for functionality.
  - ✓ All TPC FE readout electronics (mini-crates, CAEN boards and power supplies) installed and verified. Internal connections were verified for a third time together with the external ones (preliminary noise measurements).



*PMT Optical fibers feedthrough  
(covered for protection)*

*TPC Mini-crate  
with 9 read-out  
boards*

*2 power  
supplies*

Slide# : 1

## Technical status -2

- ✓ North part of side CRT was installed.
- ✓ All cryogenic equipment & transfer lines installed, welded and pressure tested, with exception of collector lines from the chimneys to gas recirculation units (under construction).
- ✓ Argon vent lines from magnetic safety valves from overpressure automatic valves installed;
- ✓ Strain gauges located on sides of cold vessels connected; readout system installed and operated.
- ✓ After installation of flanged connections, ports, valves and Ar cryogenic & purification equipment, the two cold vessels tested at 350 mbar overpressure.
- ✓ Both cold vessels are under vacuum pumping since 3 weeks. Leak detection & repair is in progress.

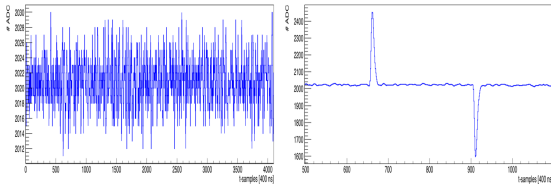


*Behavior of cold vessels, as measured by strain gauges, in agreement with finite element analysis calculations.*

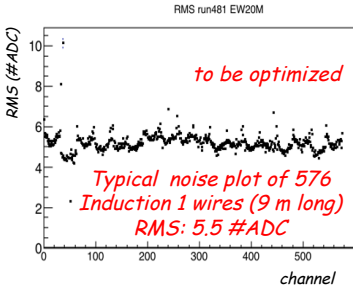
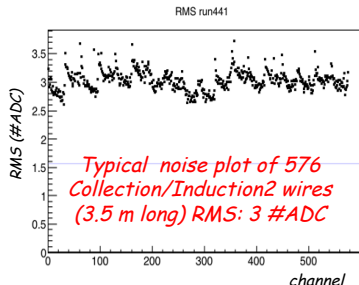
Slide# : 2

## TPC electronics: Vertical Slice test

- A full TPC vertical slice performed on all 96 feed-through flanges injecting test pulse at far end of chamber wires and reading out the signals by A2795 on the other end to evaluate overall performance.



*Typical baseline noise and test pulse shape  
(board 1, ch 37, Collection wire): RMS: 3.1 #ADC*

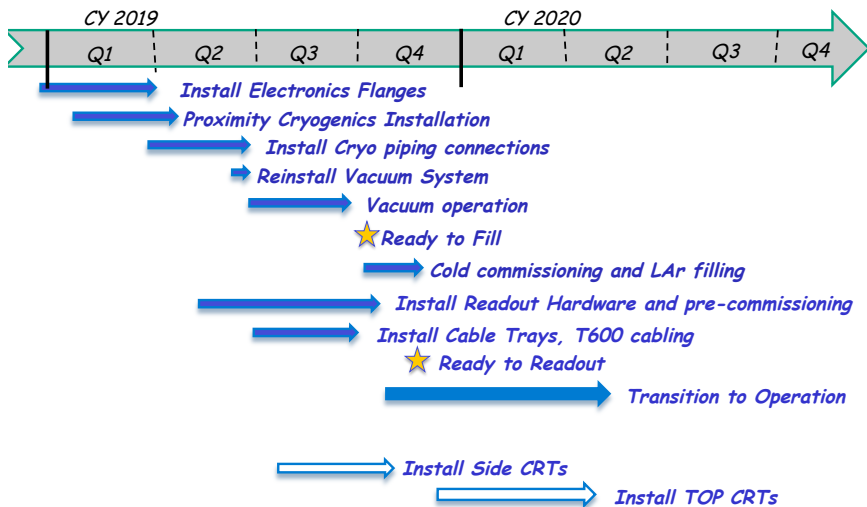


## DAQ/Trigger electronics: technical status

- ICARUS Trigger/DAQ at FNAL rely on architecture already deployed at LNGS: a waveform recording of TPCs/PMTs signals, triggered by scintillation light in coincidence with Early Warning of p beam extraction, synch by White Rabbit.
- *DAQ read-out based on CAEN A3818 optical links and PCI express:*
  - TPC + PMT+ CRT signal read by 30 PCs, each hosting 2 CAEN A3818;
  - 54 + 6 spare A3818 boards for TPC /PMT delivered to FNAL;
  - Delivery of 30+5 spare PCs HP ProLiant DL 380 model and associated connecting fibers expected by July.

Slice tests already performed with one development server with 1 A3818 to receive data via optical links: the on-line already supports 5 Hz data acquisition!
- *Trigger system based on NI PXIe instrumentation:*
  - PXIe-8135 CPU, a SPEXI board and 3 NI PXIe-7820R boards with FPGA for Trigger Logic implementation: under test at CERN/FNAL;
  - VHDL codes for generating TT-Link, clock and gate opening, tested at FNAL;
  - LabVIEW driver for SPEXI hardware is in preparation/test.

## ICARUS-T600: detailed plan-timeline



## ICARUS financial requests for 2020

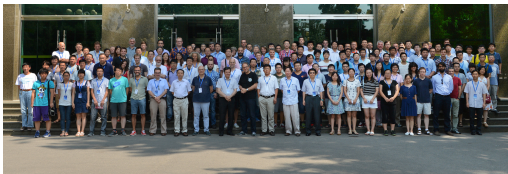
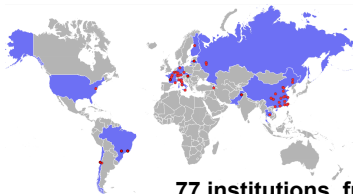
Phys. + Techs:  
10 people, 7.6 FTE

OM: 2 MU at FNAL  
OE: 2 MU at FNAL, 3 MU at Pd  
Calcolo: 2 MU (test data transfer)

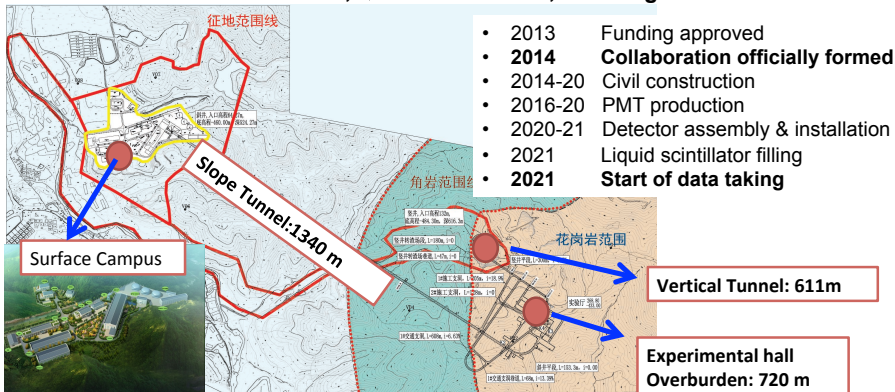
items	
Consumo: manutenzione elettronica	20 k€
Liquidi criogenici LNL	15 k€
Trasporti materiali FNAL via CERN	15 k€
Costruzione apparati: 700 preamp boards (8 ch)	18 k€
Inventario: 2 PCs local analysis	5 k€
SP Servizi: mechanics workshop	2 k€
<b>Total</b>	<b>75 k€</b>

Missions at FNAL	
T600 post-commissioning/ calibration 2 MU	11 k€
On site expert, T600 maintenance 4 MU	22 k€
110 shifts 5.5 MU	30 k€
Participation to: SBN Oversight Board, Dir. Review Coll. Meetings,	10 k€
<b>Total</b>	<b>73 k€</b>





**77 institutions, from 17 countries, counting 629 scientists**



# JUNO TAO

**Goal : to measure the reactor neutrino spectrum**

(NMH measurement should not be affected by fine structures of the spectrum)

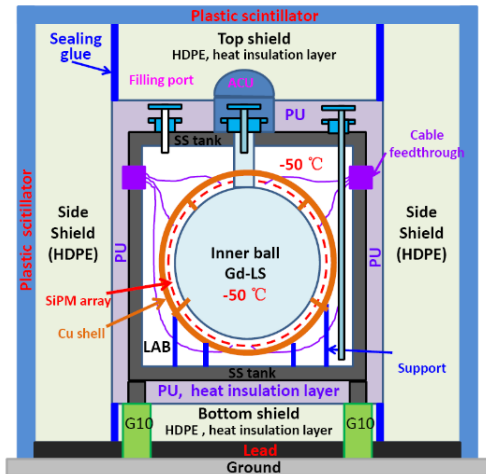
## Detector concept :

Gd-loaded LS @ -50°C+ SiPM  
 700k/year @ 40m from Taishan  
 20x JUNO 6-year data in 3 year  
 Energy resolution:  $1.5\%/\sqrt{E}$

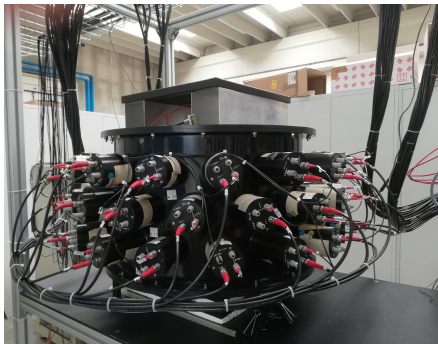
## Status:

Design and R&D on the way:

- LS works in -50°C
- SiPM & its readout electronics
- Mechanical design
- Measured onsite muon/neutron flux
- Prototype a low temperature LS detector



- first integration tests (full electronics chain: PMT → DAQ) in Legnaro, beginning of July 2019
- 8 PMT connected and continuously read out
- the system will be used for verifying the long term stability of the electronics



8 PMT connected to the 3 GCU | Trigger on threshold, BEC provides the OR of the 8 channels  
Data acquired and stored on 5<sup>th</sup> of July (data taking 20 minutes)

## Attività 2020

- completamento dei test verifica elettronica e di integrazione
- test di lunga durata: misura delle prestazioni e affidabilità dell'elettronica
- sviluppo firmware della GCU (L3 a Padova: A. Bergnoli)

## Anagrafica : 6.6 FTE (Preliminare)

A. Bergnoli	70%	R. Brugnera	30%	S. Dusini	30%
A. Garfagnini	70%	A. Giaz	100%	I. Lippi	40%
F. Marini	100%	M. Mezzetto	10%	F. H. Sawy	100%
C. Sirignano	30%	L. Stanco	30%		
M. Bellato	10%	F. Dal Corso	20%	C. Fanin	20%

## Richieste 2020 - CSNII

Missioni	80 k	Consumo	10 k
Trasporti	10 k	Inventariabile	10 k
Costruzione Apparat	2 M		
[Gara GCU]			

## Richieste 2019 - Servizi Sezione

O. Meccanica	4 mu	realizzazione test facility elettronica
S. Elettronica	24.0 mu	sviluppo firmware elettronica e sistema di test

## Quantum Science and Technology



PAPER

## Towards quantum communication from global navigation satellite system

RECEIVED  
2 August 2018REVISED  
19 October 2018ACCEPTED FOR PUBLICATION  
9 November 2018PUBLISHED  
20 December 2018

Luca Calderaro<sup>1,2</sup>, Costantino Agnesi<sup>1,2</sup>, Daniele Dequal<sup>3</sup>, Francesco Vedovato<sup>1,2</sup>, Matteo Schiavon<sup>1,2</sup>,  
Alberto Santamato<sup>1</sup>, Vincenza Luceri<sup>4</sup>, Giuseppe Bianco<sup>3</sup>, Giuseppe Vallone<sup>1,2</sup> and Paolo Villorosi<sup>1,2</sup>

<sup>1</sup> Dipartimento di Ingegneria dell'Informazione, Università di Padova, via Gradenigo 6B, I-35131 Padova, Italy

<sup>2</sup> Istituto Nazionale di Fisica Nucleare (INFN)—sezione di Padova, Italy

<sup>3</sup> Matera Laser Ranging Observatory, Agenzia Spaziale Italiana, Matera, Italy

<sup>4</sup> e-GEOS SpA, Matera, Italy



Our research was partially funded by the MoonLIGHT-2 project of INFN.

National coordinator: Dr. Simone Dell'Agnello

## *GNSS orbit tasted at 20000km: GLONASS spacecrafts*

two GLONASS terminals equipped with an array of corner-cube retroreflectos (CCRs), namely Glonass-134 and Glonass-131 (Space Vehicle Number: 802 and 747, respectively)

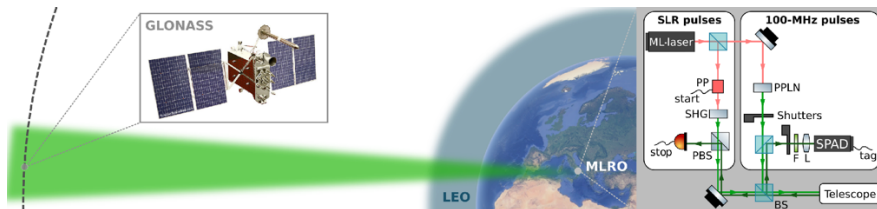
The targeted GNSS satellites are part of different generations, GLONASS-K1 for Glonass-134 and GLONASS-M for Glonass-131, both equipped with a planar array of CCRs, with circular and rectangular shape respectively

Their CCRs are characterized by the absence of coating on the reflecting faces, such that the light is back reflected by total internal reflection (TIR). This implies a far field diffraction pattern (FFDP) which is quite different from the simple Airy disk given by a circular aperture



## GLONASS returns summary

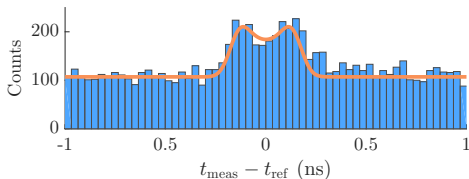
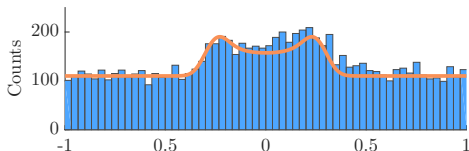
Satellite passage	Slant distance (km)	Detector	$\bar{R}_{\text{det}}$ (Hz)	SNR	$\bar{\mu}_{\text{sat}}$	$l_{\text{down}}$ (dB)	$l_{\text{rec}}$ (dB)
Glonass-134	19,500	SPAD	58	0.53	15	62.1	11.8
	20,200	SPAD	59	0.41	16	62.5	11.8
Glonass-131	20,250	SPAD	27	0.43	15	62.6	14.8
		PMT	6	0.21	16	62.6	21.8



L. Calderaro et al. *Towards Quantum Communication from Global Navigation Satellite System*, *Quantum Sci. Technol.* **4** 015012 (2019).

## GLONASS temporal footprint

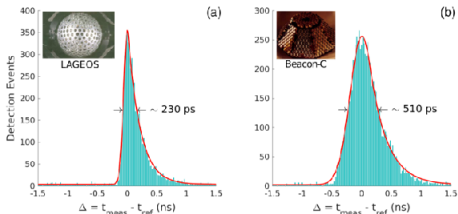
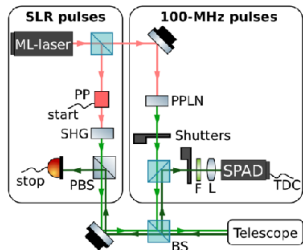
The incident angle of the beam on the array: 9 deg and 5 deg



L. Calderaro et al. *Towards Quantum Communication from Global Navigation Satellite System*, **Quantum Sci. Technol.** **4** 015012 (2019).



## Temporal resolution in the single photon time tagging reduced to 230 ps over 7000km



A 100-MHz pulse train is used to test the detector system temporal resolution. The **silicon single photon avalanche detector (SPAD - Micro-Photon-Devices Srl)**, with  $\approx 50\%$  quantum efficiency,  $\approx 400$  Hz dark count rate and 40 ps of jitter.

The time of arrival is tagged with 1 ps resolution (quTAG TDC from qutools GmbH)

The **evident tail** is the actual limitation for the temporal resolution, and is due to the design of the active part of the detector.

A novel design is available by MPD.

Research Article

Vol. 36, No. 3 / March 2019 / Journal of the Optical Society of America B 599

Journal of the  
Optical Society  
of America

B

OPTICAL PHYSICS

### Sub-ns timing accuracy for satellite quantum communications

COSTANTINO AGNESI,<sup>1,2</sup> LUCA CALDERARO,<sup>1,2,3</sup> DANIELE DEQUAL,<sup>4</sup> FRANCESCO VEDOVATO,<sup>1,2,3</sup> MATTEO SCHIAVON,<sup>1,2</sup> ALBERTO SANTAMATO,<sup>1</sup> VINCENZA LUCERI,<sup>5</sup> GIUSEPPE BIANCO,<sup>4</sup> GIUSEPPE VALLONE,<sup>1,2</sup> AND PAOLO VILLORESI<sup>1,2,\*</sup>

<sup>1</sup>Dipartimento di Ingegneria dell'Informazione, Università degli Studi di Padova, via Gradenigo 6B, 35131 Padova, Italy

<sup>2</sup>Istituto Nazionale di Fisica Nucleare - sezione di Padova, Padova, Italy

<sup>3</sup>Centro di Ateneo di Studi e Attività Scientifiche "G. Colombo," Università degli Studi di Padova, via Venezia 15, 35131 Padova, Italy

<sup>4</sup>Matera Laser Ranging Observatory, Agenzia Spaziale Italiana, Matera, Italy

<sup>5</sup>Ve-GEOSS SpA, Matera, Italy

- Detection of **cosmological axions** through their **coupling to electrons or photons**
- **Electron coupling:** Due to the **motion of the solar system** in the galaxy, the axion DM cloud acts as an **effective RF magnetic field** on electron spin exciting magnetic transitions in a magnetized sample and **producing rf photons**

Expected rf power

$$P_{\text{out}} = \frac{P_{\text{in}}}{2} = 8 \times 10^{-26} \left( \frac{m_a}{2 \cdot 10^{-4} \text{ eV}} \right)^3 \left( \frac{V_s}{1 \text{ liter}} \right) \left( \frac{n_S}{10^{28} / \text{m}^3} \right) \left( \frac{\tau_{\text{min}}}{10^{-6} \text{ s}} \right) \text{ W}$$

- **Photon coupling:** DM axion are converted into **rf photons** inside a **resonant cavity** immersed in a **strong magnetic field**

Expected rf power

$$P_a = 1.85 \times 10^{-25} \text{ W} \left( \frac{V}{0.036 \text{ l}} \right) \left( \frac{B}{2 \text{ T}} \right)^2 \left( \frac{g_f}{-0.97} \right)^2 \times \left( \frac{C}{0.589} \right) \left( \frac{\rho_a}{0.45 \text{ GeV cm}^{-3}} \right) \left( \frac{\nu_c}{9.067 \text{ GHz}} \right) \left( \frac{Q_L}{201000} \right)$$

# Activity in the laboratory

Refurbishing of a **Low Power Dilution Refrigerator** completed

First tests of a **Josephson Parametric amplifier (JPA)** @ 100 mK

- Expected  $T_{\text{noise}} \sim 0.5$  K

New **in-house procedure for production of YIG spheres** up to 2.5 mm diameter

Coupling of a **superconducting cavity loaded with YIG sphere** achieved

New **photonic cavities** on the way

Calibration scheme with laser input

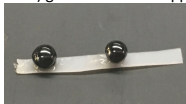
Refurbishing of a **High Power Dilution Refrigerator** under way

New high field (8 T) magnet to be delivered



Dilution system


Home made YIG spheres ( $\phi = 2$  mm) glued on teflon support



Dilution insert with rf electronics and cavity

# QUAX limit on axion electron coupling

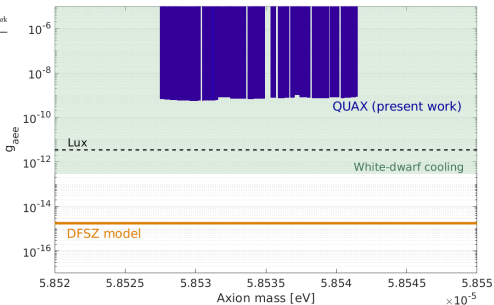
Eur. Phys. J. C (2018) 78:705  
<https://doi.org/10.1140/epjc/s10052-018-6163-8>

THE EUROPEAN  
 PHYSICAL JOURNAL C 

Regular Article - Experimental Physics

Operation of a ferromagnetic axion haloscope at  $m_a = 58 \mu\text{eV}$

This is the first limit in the parameter space  $\{m_a, g_{aee}\}$  obtained from an experiment searching for axions as the main Dark Matter component (Haloscope)



Limit is still poor but:

- Material volume 2.6 mm<sup>3</sup>
- System total noise temp. 15 K
- Relaxation time 0.1 ms

QUAX expected 2019

- Material volume 42 mm<sup>3</sup> / 10<sup>5</sup> mm<sup>3</sup>
- System total noise temp. 0.5 K
- Relaxation time 0.3 ms / 2 ms

QUAX (final)

- System total noise temp. counter ( $T_{\text{eff}} < 1 \text{ mK}$ )

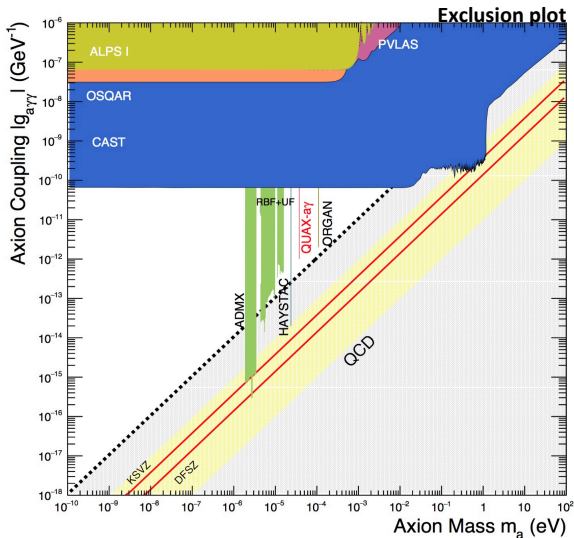
The set-up with 10 YIG spheres giving 42 mm<sup>3</sup> of volume already mounted and tested at 4 K

# First results on axion-photon coupling

Results obtained by the QUAX- $a\gamma$  set-up are competitive with other experiments

Just appeared on  
**Phys Rev D rapid comm.**

Starting from 2020 QUAX- $a\gamma$  will start a **search with mass scanning** employing the **Josephson Parametric Amplifier** and a **higher Q cavity**



# QUAX 2020

## PD activity for 2020:

- New Set-up for axion photon coupling on **Large Cooling Power Dilution Unit**
- Sapphire Copper Cavity with tuning and scanning device [  $Q = 10^6$  ]
- Integration of the **Josephson Parametric Amplifier in the ultra - low temperature (<50 mK) cryogenic system** [  $T_{\text{noise}} @ \text{Quantum Limit}$  ]
- Test with High Magnetic Field [ **8 Tesla** ]

Sezioni INFN partecipanti all'esperimento: **LNL, Padova, LNF, Napoli, Trento**

Resp. Naz.: **Giovanni Carugno**

Partecipazione PD : G.Carugno 50%, C.Braggio 40%, R.Di Vora 100%, A.Pepato 40%,  
L.Taffarello 40 %, M.Pegoraro 20%

**Totale: 2.9 FTE**

Richieste 2020 PD: **90 KE**

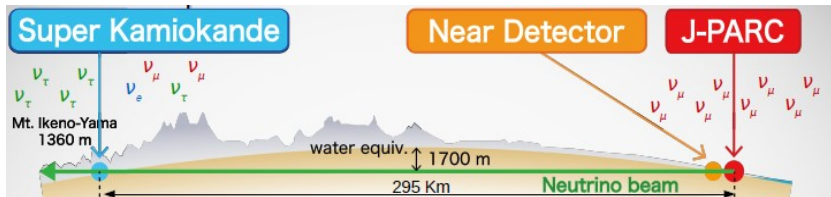
Richieste Servizi: 12 M.U. O.M. , 8 M.U. O.E. , 2 M.U. U.T.



# Prospettive esperimenti T2K & SuperKamiokande



G.Collazuol – CdS INFN PD - 2019/7/10

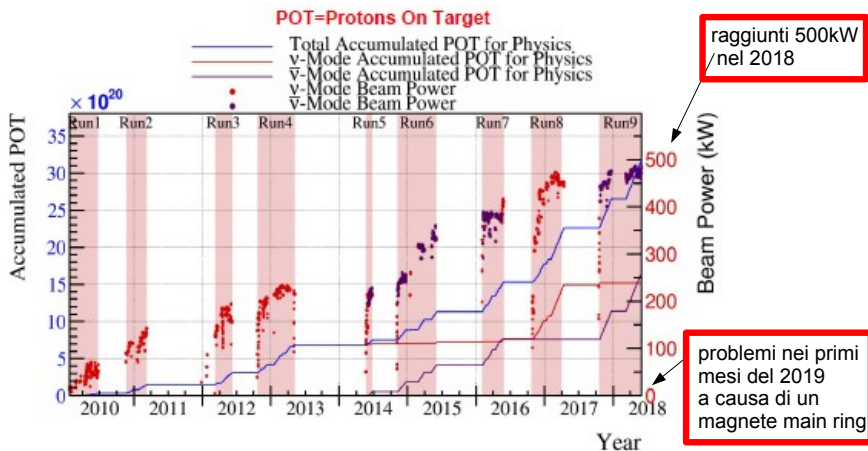


## L'attuale gruppo locale

Staff: **G.C.** (T2K&SK), **M.Laveder**(T2K), **A.Longhin**(T2K), **M.Mezzetto** (T2K)  
 Post-doc: **M.Lamoreux** (INFN Fellini, T2K&SK), **N.Ospina** (Assegnista J, SK)  
 Dottorandi: **F.Iacob** (T2K&SK), **M.Pari** (T2K) / Laureandi: **M.Menotti** (T2K upgrade)  
 Supporto Tecnico per T2K-upgrade: **G.Cogo**  
 In stretta collaborazione per T2K-upgrade @ LNL: **M.Cicerchia**, **F.Gramegna**, **T.Marchi**



# Protons on Target



23 Jan. 2010 – 31 May 2018

POT total:  $3.16 \times 10^{21}$

$\nu$ -mode  $1.51 \times 10^{21}$  (47.83%)

$\bar{\nu}$ -mode  $1.65 \times 10^{21}$  (52.17%)





Prospettive → T2K-II

Beam and Near Detector upgrade

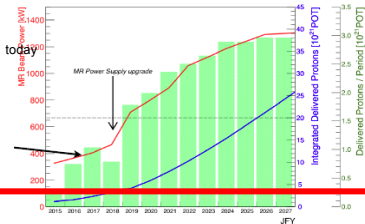


T2K-II extension proposal (arXiv:1609.04111):

- Aim at  $3\sigma$  CPV sensitivity or CPC exclusion (w/ favorable param.)
- Accumulate  $20 \times 10^{21}$  pot by 2026 (x3 original T2K program)

Mandatory (for high stat)

- High Intensity Neutrino beam
- **Accelerator/Beam-line Upgrade** approved
- gradual increase of power to 1MW (1.3MW) by 2021 (2026)



Mandatory (at high stat)

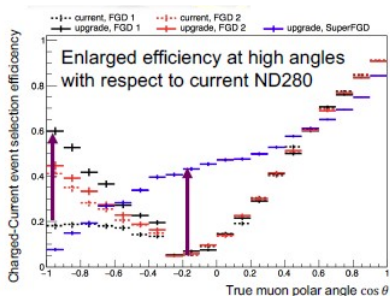
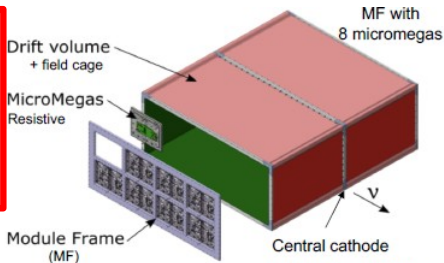
- Reduce detector systematics to 4% & cross-section model dependence
- **Near Detector Upgrade** → TDR 2018 → Approved by JPARC and SPSC



# Near Detector Upgrade - HA TPC

Importante coinvolgimento INFN Pd in costruzione delle **TPC orizzontali** (High Angle TPC):

- **coordinamento progetto HA-TPC** (G.Collazuol)
- disegno e **costruzione Field Cage** con INFN-LNL e INFN-Ba



HA-TPC V-TPC

Parameter	Value	
Overall $x \times y \times z$ (m)	$2.0 \times 0.8 \times 1.8$	$0.85 \times 2.2 \times 1.8$
Drift distance (cm)	90	
Magnetic Field (T)	0.2	
Electric field (V/cm)	275	
Gas Ar-CF <sub>4</sub> -iC <sub>4</sub> H <sub>10</sub> (%)	95 - 3 - 2	
Drift Velocity $cm/\mu s$	7.8	
Transverse diffusion ( $\mu m/\sqrt{cm}$ )	265	
Micromegas gain	1000	
Micromegas dim. $z \times y$ (mm)	$340 \times 420$	$340 \times 360$
Pad $z \times y$ (mm)	$10 \times 11$	$7 \times 10$
N pads	36864	124272
el. noise (ENC)	800	
S/N	100	
Sampling frequency (MHz)	25	
N time samples	511	

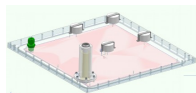
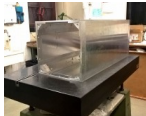
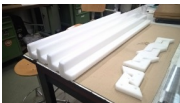
# T2K HATPC prototipo

(1) Preparazione MOLD → INFN Bari

(3) Produzione @ NEXUS  
(Barcellona), supporto IFAE



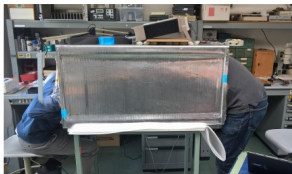
(2) Produzione parti strutturali in POM-C, catodo e rifinitura MOLD @ officina meccanica INFN Pd



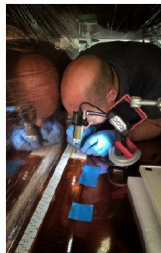
(4) Fresatura di precisione  
Flange @ officina LNL



(5) Assemblaggio, test elettrici e gas in laboratorio a LNL  
(gruppi T2K PD e LNL)



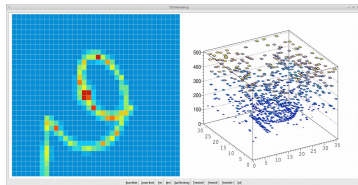
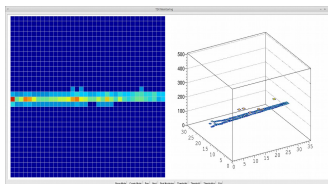
No relevant issues in soldering + gluing → 1.5 days



# T2K HATPC attivita` 2019/20

## Test Beam @ DESY June 2019

→ new RMM Resistive type Micromegas (charge sharing) → OK



## Test Run @ CERN w/ Cosmics, since July 2019

→ new prototype Field Cage + new RMM = complete HA-TPC validation

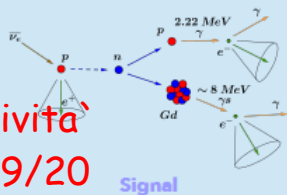
→ Attivita` 2020 → costruzione e test 2 HATPC

## Attivita` 2019/20

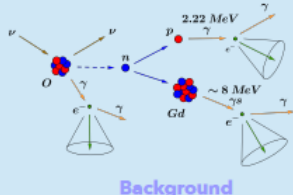
# NCQE neutron multiplicity

- **SK-Gd project:** load SK with Gadolinium sulphate to produce detectable 8 MeV gamma cascade by neutron capture.
- **Signal:** Inverse Beta Decay (IBD), detected by prompt positron signal + delayed radiative neutron capture on Gd. Analysis target: Supernova Relic Neutrinos (SRN).
- **Background:** Neutrino-Oxygen neutral current quasi-elastic (NCQE) interactions, induced by atmospheric neutrinos, mimic the prompt+delayed signal pattern of IBD induced by SRN.

Attività  
2019/20



Signal



Background

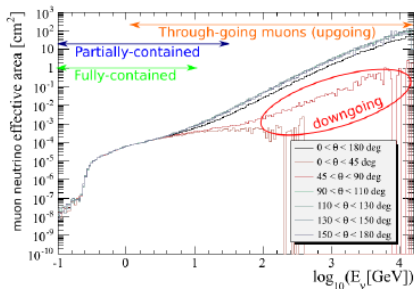
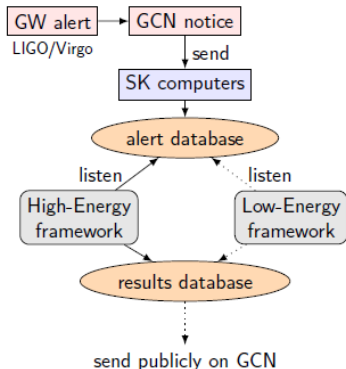
PhD student currently working on **NCQE-events neutron multiplicity** analysis (T2K-SK group):

- Well controlled NCQE events induced by triggered neutrino beam of T2K experiment.
- Nowadays neutron capture on **Hydrogen** (2.2 MeV hardly detectable gamma, advanced hit PMTs pattern recognition with **neural network**).
- Future perspective: adapt neutron tagging analysis and techniques developed for Hydrogen-capture to the **Gadolinium**-capture case, as soon as SK-Gd starts.

## Astrophysics with Super-Kamiokande

Search for neutrinos in Super-Kamiokande in coincidence with transient astrophysical events (GW,  $\nu$ ,  $\gamma$ ...)

- One postdoc (FELLINI fellow) currently working on the development of the automatic process for follow-up of alerts with SK High-Energy samples ( $E = 10^{-1} - 10^6$  GeV). Ongoing internal tests with public GW O3 alerts.



Attività` 2019/20



## Anagrafica e Richieste 2020



### Anagrafica 2019

G.Collazuol 60% (+10% ENUBET)  
 N.Ospina (PD) 60%  
 F.Iacob (PhD) 90%  
 M.Pari (PhD) 30% (+70% ENUBET)  
 G.Cogo (Assegno Tecn.) 80 %

M.Laveder 60%  
 A.Longhin 30% (+70% ENUBET)  
 M.Mezzetto 80% (+10% ENUBET)  
 M.Lamoreux 100% FELLINI

Totale FTE = 4.9 T2K-SK + 1.0 T2K-FELLINI + 1.6 ENUBET

### Richieste 2020

Missioni interne 4k€

Missioni estero 80k€

- shifts e meetings T2K & SK
- test prima HA-TPC at DESY

Costruzione apparati 90k€

- HA-TPC Field Cage (FC) finali

Consumabile 10k€

- RMM ed elettronica TPC

### Richieste Servizi Sezione 2020

- U.Tecnico (principalmente G.Cogo) 9 mu per disegno & controllo produzione FC
- O.Meccanica: 2 mu per parti strutturali in termoplastica FC
- S.Elettronica: 2 mu per test parte elettrica FC

# Virgo: alcuni risultati scientifici

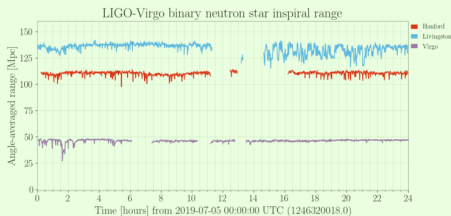
**GWTC-1** (risultati di O1&O2 su Compact Binary Mergers):

- 11 confirmed detections (FAR < 1/30 days,  $p_{\text{astro}} > 50\%$ )
- 10 BBHs (4 previously unpublished)
- 1 BNS

arXiv:1811.12907

- Began on April 1st 2019
- Duration : 1 year
- First run with open public alerts :  
<https://gracedb.ligo.org/>
- Alerts sent out if a detection meets the FAR threshold of < 1/2 months

Ad oggi durante O3:  
21 public alerts  
3 retractions



During O3, about 3dB squeezing routinely injected in Adv:  
Squeezing level > 2.5 dB: 72.5% of the total time  
determinante il contributo di Virgo Pd



# Virgo – Pd : attività sperimentale 2020

- R&D per EPR squeezing :
  - coherent control loops: PLLs
  - Sensors for Mode matching (lente elettro-ottica)
  - Actuators for Mode matching (telescopio motorizzato)
  - Sensors for auto-alignment: deflettore elettro-ottico
- R&D coating
  - dettagli nella slide successiva
- R&D nonequilibrium thermal noise
- commissioning di Frequency Dependent Squeezing
- attività sperimentali per Frequency Dependent Squeezing saranno (forse) finanziate da EGO come AdV+ project (pending decision by EGO council 18-19 Jul)

## R&D on coating

- **Coating Research and Development (CRD).** Research on new materials with low thermal noise for fabrication of new mirrors for Virgo. Preparatory work for O5.
  - **Mode Matching (QNR)** Research on new techniques for measurement and control of the mode mismatch between optical beams and cavities. Crucial for SQZ.
- Study of nanocrystals formation in SiC (nanoSiC)** The idea is to determine a protocol to produce dispersed crystals with nanometric size in an amorphous SiC matrix, a material which could be potentially be used for test masses coatings.

per questa ricerca il gruppo di Virgo-Pd cresce ulteriormente grazie alla partecipazione di G. Maggioni (0.1FTE) di UniPd

# Analisi Pd/Tn - Attività per il 2020

- **Run O3**
  - **Burst All-Sky Short Duration Generic Transients Signals**
    - Contributo all'analisi con cWB e alla scrittura articolo
  - **Burst All-Sky Long Duration Generic Transients Signals**
    - Analisi con cWB e contributo alla scrittura articolo
  - **IMBBH (Intermediate Mass Binary Black Holes)**
    - Contributo all'analisi con cWB
  - **Stellar Mass BBH & eBBH**
    - Analisi burst cWB e contributo alla scrittura degli articoli
  - **Analisi con cWB per il follow-up segnali GW**
    - Supporto low-latency per follow-up elettromagnetico
    - Analisi post-merger di segnali gravitazionali da BNS
- **Supporto e Sviluppo pipeline cWB per Run O3/O4**
  - **Supporto ai gruppi che utilizzano cWB per ricerche del Run O3**
  - **Supporto Detector Characterization per Virgo**
    - Shifts per Run O3
  - **Aggiornamento della pipeline per Run O4**
    - miglioramento degli algoritmi di ricerca e infrastruttura della pipeline (CNAF)

**ASTRO AND POPULATIONS @ VIRGO PD:**

**Staff:** Michela Mapelli (40%)  
**Postdocs:** Yann Bouffanais (40%), Mario Spera (50%)

**TOPICS in 2020:**

- \* CONTRIBUTION TO MASS SPECTRUM AND RATE ESTIMATION for O3 IN THE FRAME OF THE POP&RATE LIGO-Virgo GROUP (Spera, Mapelli, Bouffanais)
- \* CONTRIBUTION TO THE STUDY AND ASTROPHYSICAL CHARACTERIZATION OF S190521g AND OTHER POSSIBLE MASSIVE BBH CANDIDATES (Mapelli)
- \* DEVELOPMENT OF A NEW BAYESIAN ANALYSIS FRAMEWORK FOR RATE DEPENDENCE WITH REDSHIFT (Bouffanais, Mapelli)

**REQUESTS in 2020:**

- \* travel expenses to participate in Virgo weeks (Bouffanais, Mapelli)
- \* 3 M CPU hours at CINECA (to be requested to Punturo – Cosmai)

# Richiesta

richiesta economica in via di definizione, anche in coordinazione con Virgo-Italia.

Le richieste ai servizi della sezione dipendono dal programma scientifico e quindi anche dalla richiesta economica e sono in fase di valutazione

ricercatori	FTE
Bazzan	60
Bouffanais	40
Ciani	80
Ciolfi	50
Conti	80
De Pietri	50
Lazzaro	80
Mapelli	40
Spera	50
Vardaro	80
Vedovato	100
Zendri	60

**+1PhD student**

Potrebbero esserci alcune modifiche sugli FTE indicati.

tecnologi	FTE
Pegoraro	20

servizi	mesi persona
off. meccanica	in valutazione
off. elettronica	in valutazione
uff. tecnico	in valutazione