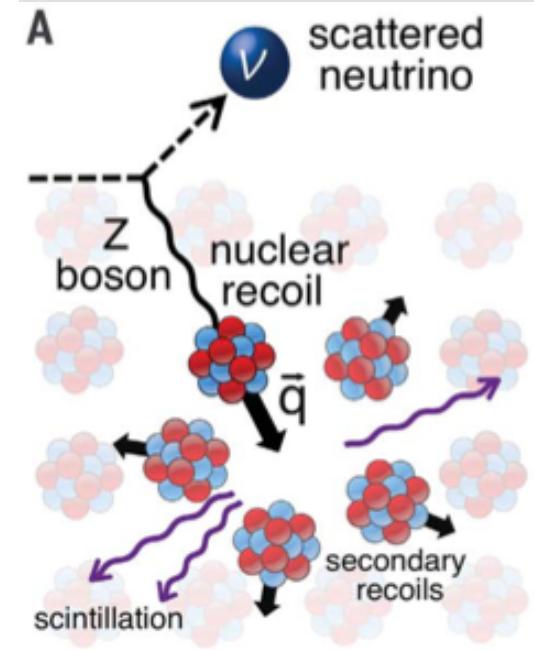


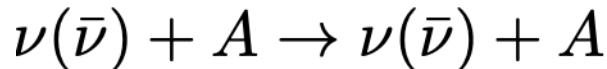
# nucleus: coherent elastic neutrino- nucleus scattering at CHOOZ



Riccardo Cerulli  
Luglio 2019

# Neutrino cross-sections and detectable energy

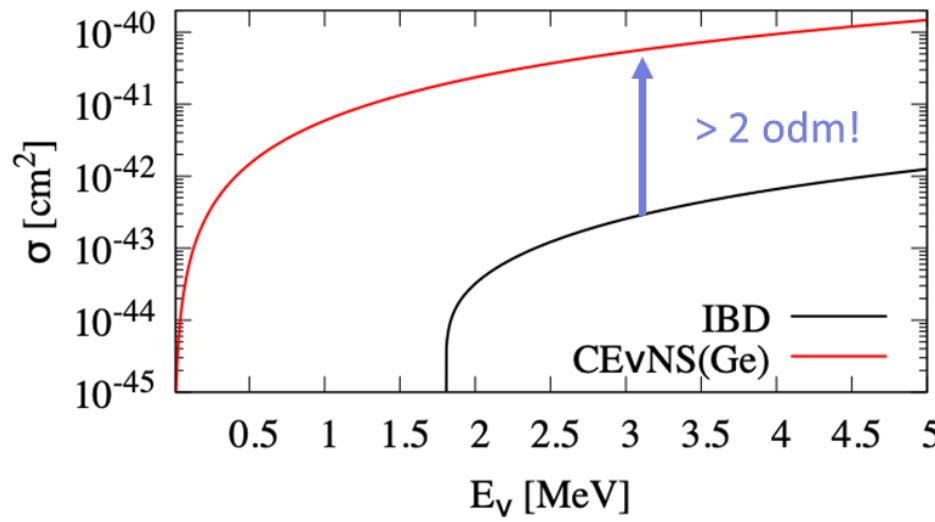
## Coherent Elastic $\nu$ -Nucleus Scattering



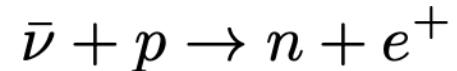
$$\sigma_{\text{CE}\nu\text{NS}} = \frac{G_F^2}{4\pi} F^2(q^2) Q_W^2 E_\nu^2$$

$$Q_W = N - Z(1 - 4 \sin^2 \theta_W) \sim N$$

$$E_\nu < qR \quad \quad \langle T \rangle = \frac{2}{3} \frac{E_\nu^2}{M_A}$$



## Inverse Beta Decay

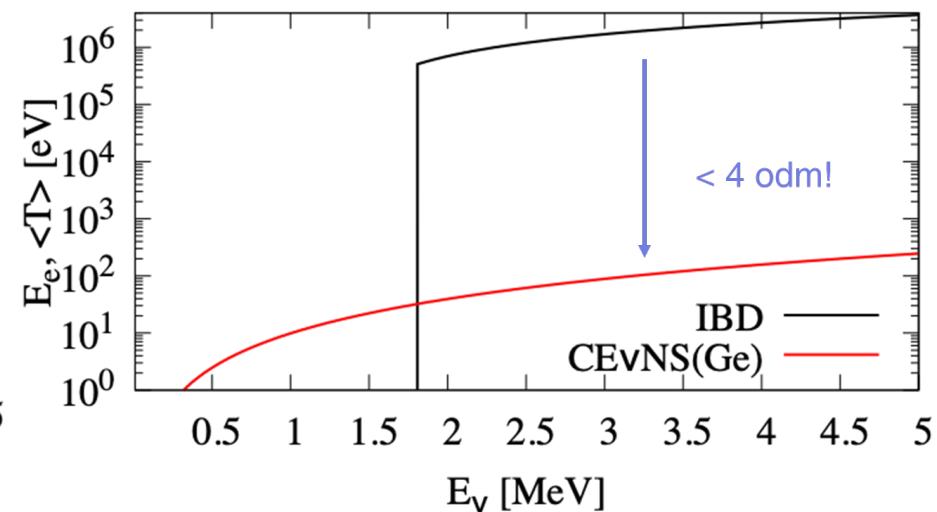


$$\sigma_{\text{IBD}}^0 = \frac{G_F^2 \cos^2 \theta_C}{\pi} (f^2 + 3g^2) E_e p_e$$

$$E_e = E_\nu - (M_N - M_p)$$

$$E_\nu > 1.806 \text{ MeV}$$

$$E_e = E_\nu - (M_N - M_p)$$



# 43 years to be discovered

PHYSICAL REVIEW D

VOLUME 9, NUMBER 5

1 MARCH 1974

RESEARCH

## Coherent effects of a weak neutral current

Daniel Z. Freedman<sup>†</sup>

National Accelerator Laboratory, Batavia, Illinois 60510

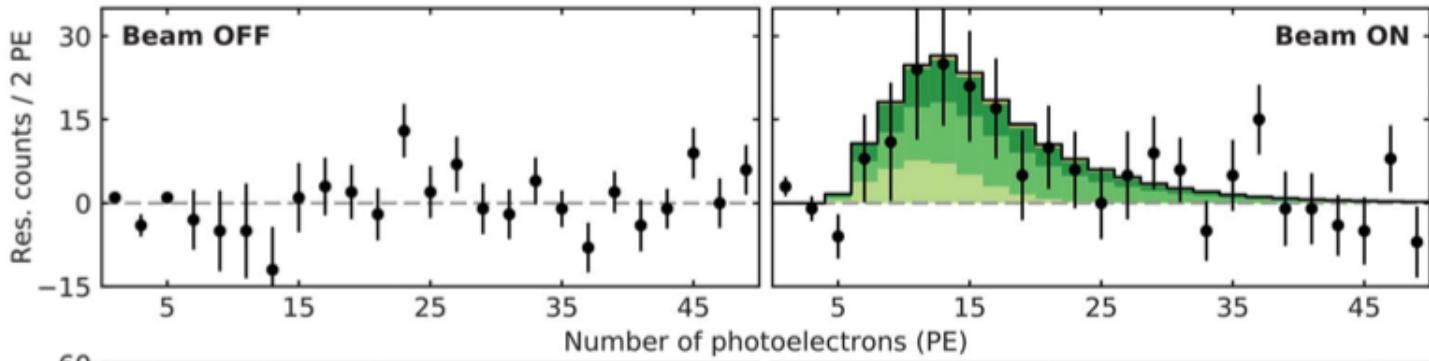
and Institute for Theoretical Physics, State University of New York, Stony Brook, New York 11790

(Received 15 October 1973; revised manuscript received 19 November 1973)

If there is a weak neutral current, then the elastic scattering process  $\nu + A \rightarrow \nu + A$  should have a sharp coherent forward peak just as  $e + A \rightarrow e + A$  does. Experiments to observe this peak can give important information on the isospin structure of the neutral current. The experiments are very difficult, although the estimated cross sections (about  $10^{-38} \text{ cm}^2$  on carbon) are favorable. The coherent cross sections (in contrast to incoherent) are almost energy-independent. Therefore, energies as low as 100 MeV may be suitable. Quasi-coherent nuclear excitation processes  $\nu + A \rightarrow \nu + A^*$  provide possible tests of the conservation of the weak neutral current. Because of strong coherent effects at very low energies, the nuclear elastic scattering process may be important in inhibiting cooling by neutrino emission in stellar collapse and neutron stars.

14.6 kg CsI[Na],

153 days beam off 308 days beam on



NEUTRINO PHYSICS

Science 357, 1123–1126 (2017)

## Observation of coherent elastic neutrino-nucleus scattering

D. Akimov,<sup>1,2</sup> J. B. Albert,<sup>3</sup> P. An,<sup>4</sup> C. Awe,<sup>4,5</sup> P. S. Barbeau,<sup>4,5</sup> B. Becker,<sup>6</sup> V. Belov,<sup>1,2</sup> A. Brown,<sup>4,7</sup> A. Bolozdynya,<sup>2</sup> B. Cabrera-Palmer,<sup>8</sup> M. Cervantes,<sup>5</sup> J. I. Collar,<sup>9\*</sup> R. J. Cooper,<sup>10</sup> R. L. Cooper,<sup>11,12</sup> C. Cuesta,<sup>13†</sup> D. J. Dean,<sup>14</sup> J. A. Detwiler,<sup>13</sup> A. Eberhardt,<sup>13</sup> Y. Efremenko,<sup>6,14</sup> S. R. Elliott,<sup>12</sup> E. M. Erkela,<sup>13</sup> L. Fabris,<sup>14</sup> M. Febbraro,<sup>14</sup> N. E. Fields,<sup>9‡</sup> W. Fox,<sup>3</sup> Z. Fu,<sup>13</sup> A. Galindo-Uribarri,<sup>14</sup> M. P. Green,<sup>4,14,15</sup> M. Hai,<sup>9§</sup> M. R. Heath,<sup>3</sup> S. Hedges,<sup>4,5</sup> D. Hornback,<sup>14</sup> T. W. Hossbach,<sup>16</sup> E. B. Iverson,<sup>14</sup> L. J. Kaufman,<sup>3||</sup> S. Ki,<sup>4,5</sup> S. R. Klein,<sup>10</sup> A. Khromov,<sup>2</sup> A. Konovalov,<sup>1,2,17</sup> M. Kremer,<sup>4</sup> A. Kumpan,<sup>2</sup> C. Leadbetter,<sup>4</sup> L. Li,<sup>4,5</sup> W. Lu,<sup>14</sup> K. Mann,<sup>4,15</sup> D. M. Markoff,<sup>4,7</sup> K. Miller,<sup>4,5</sup> H. Moreno,<sup>11</sup> P. E. Mueller,<sup>14</sup> J. Newby,<sup>14</sup> J. L. Orrell,<sup>16</sup> C. T. Overman,<sup>16</sup> D. S. Parno,<sup>13¶</sup> S. Penttila,<sup>14</sup> G. Perumpilly,<sup>9</sup> H. Ray,<sup>18</sup> J. Raybern,<sup>5</sup> D. Reyna,<sup>8</sup> G. C. Rich,<sup>4,14,19</sup> D. Rimal,<sup>18</sup> D. Rudik,<sup>1,2</sup> K. Scholberg,<sup>5</sup> B. J. Scholz,<sup>9</sup> G. Sinev,<sup>5</sup> W. M. Snow,<sup>3</sup> V. Sosnovtsev,<sup>2</sup> A. Shakirov,<sup>2</sup> S. Suchtya,<sup>10</sup> B. Suh,<sup>4,5,14</sup> R. Tayloe,<sup>3</sup> R. T. Thornton,<sup>3</sup> I. Tolstukhin,<sup>3</sup> J. Vanderwerp,<sup>3</sup> R. L. Varner,<sup>14</sup> C. J. Virtue,<sup>20</sup> Z. Wan,<sup>4</sup> J. Yoo,<sup>21</sup> C.-H. Yu,<sup>14</sup> A. Zawada,<sup>4</sup> J. Zettlemoyer,<sup>3</sup> A. M. Zderic,<sup>13</sup> COHERENT Collaboration#

The coherent elastic scattering of neutrinos off nuclei has eluded detection for four decades, even though its predicted cross section is by far the largest of all low-energy neutrino couplings. This mode of interaction offers new opportunities to study neutrino properties and leads to a miniaturization of detector size, with potential technological applications. We observed this process at a 6.7 $\sigma$  confidence level, using a low-background, 14.6-kilogram CsI[Na] scintillator exposed to the neutrino emissions from the Spallation Neutron Source at Oak Ridge National Laboratory. Characteristic signatures in energy and time, predicted by the standard model for this process, were observed in high signal-to-background conditions. Improved constraints on nonstandard neutrino interactions with quarks are derived from this initial data set.

Spallation neutron source at Oakridge

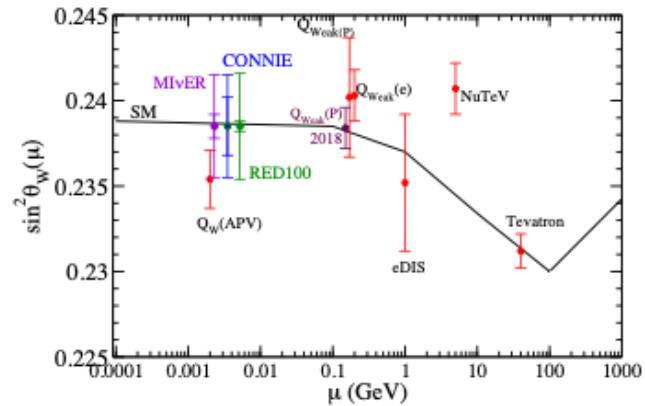
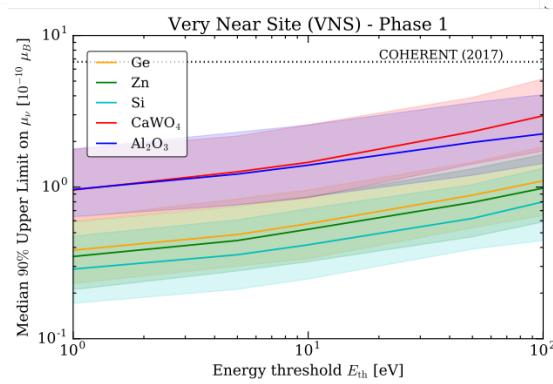
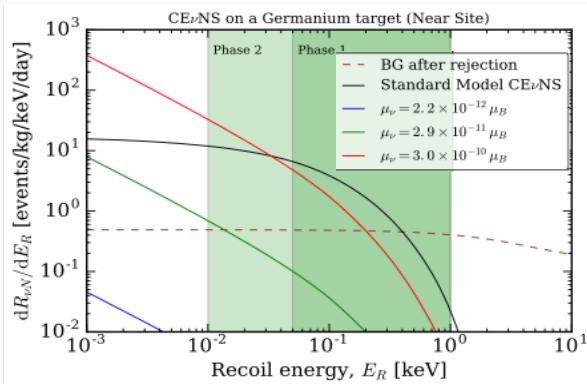
# Physics with CE $\nu$ NS and more

- $\sin^2\theta_W$  at low Q
- Non-standard interactions

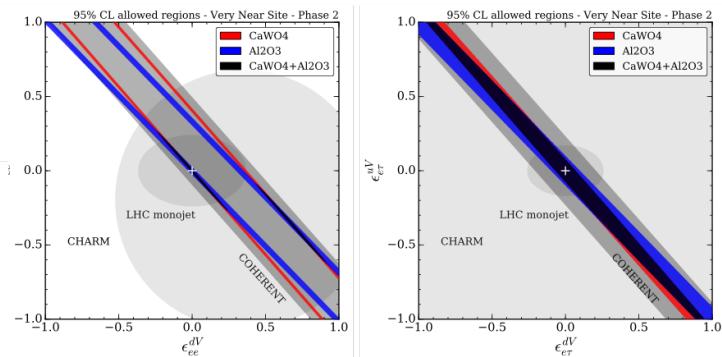
$$\mathcal{L}^{\text{NSI}} = -\epsilon_{\alpha\beta}^{fP} 2\sqrt{2}G_F(\bar{\nu}_\alpha \gamma_\rho L \nu_\beta)(\bar{f} \gamma^\rho P f),$$

- Neutrino magnetic moment

$$\frac{d\sigma_{\nu-N}^{\text{mag.}}}{dE_R} = \frac{\pi\alpha^2\mu_\nu^2 Z^2}{m_e^2} \left( \frac{1}{E_R} - \frac{1}{E_\nu} + \frac{E_R}{4E_\nu^2} \right) F^2(E_R).$$



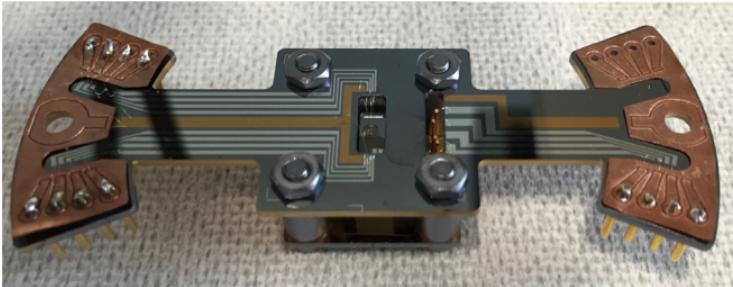
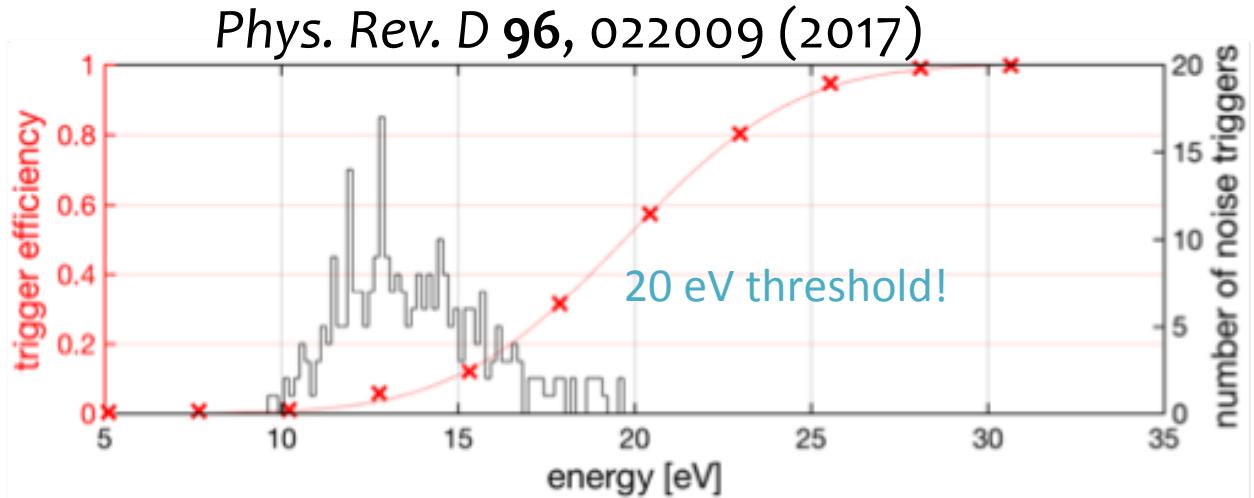
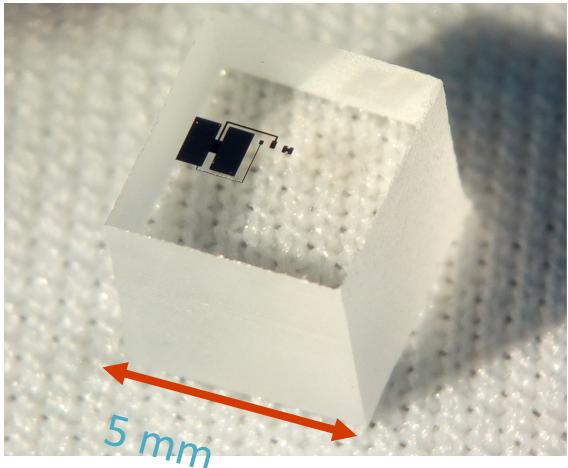
JCAP 1811 (2018) 016



- Non-proliferation: detectors can be used to find out fuel diversion (e.g. Plutonium removal from core, 1811.04737)

# NUCLEUS: detector

*CRESST technology: TES on  $\text{CaWO}_4$  or  $\text{Al}_2\text{O}_3$  at 10 mK*

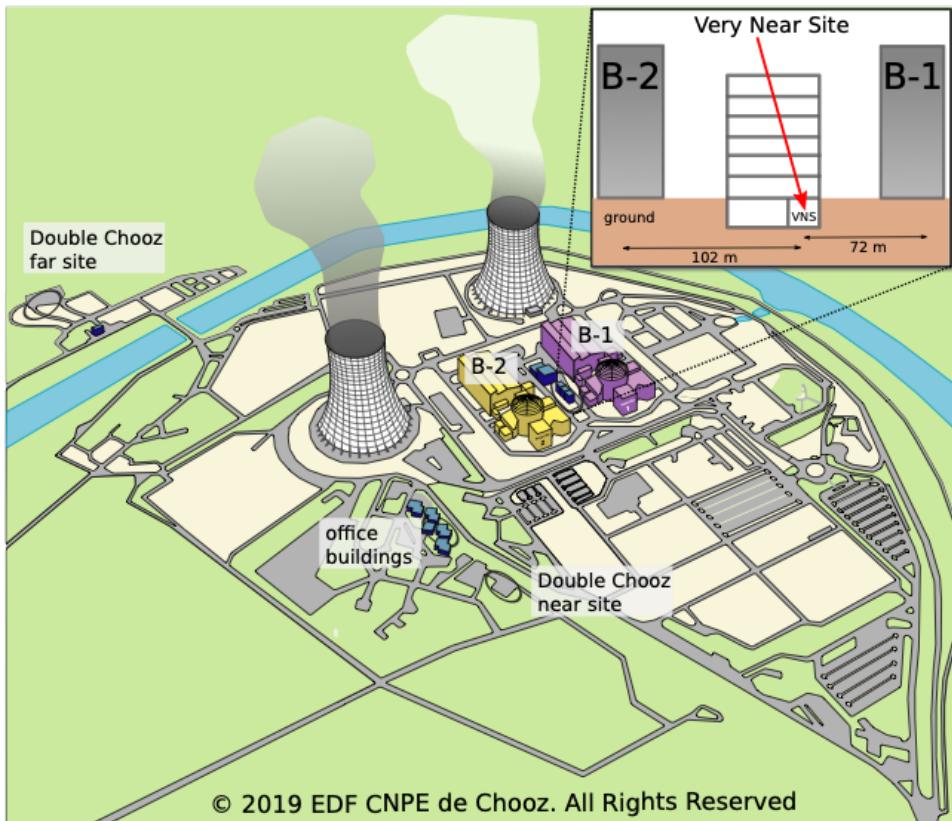


- Sapphire and  $\text{CaWO}_4$  crystals
- Flexible Si wafers as inner veto detectors

Si outer veto prototype

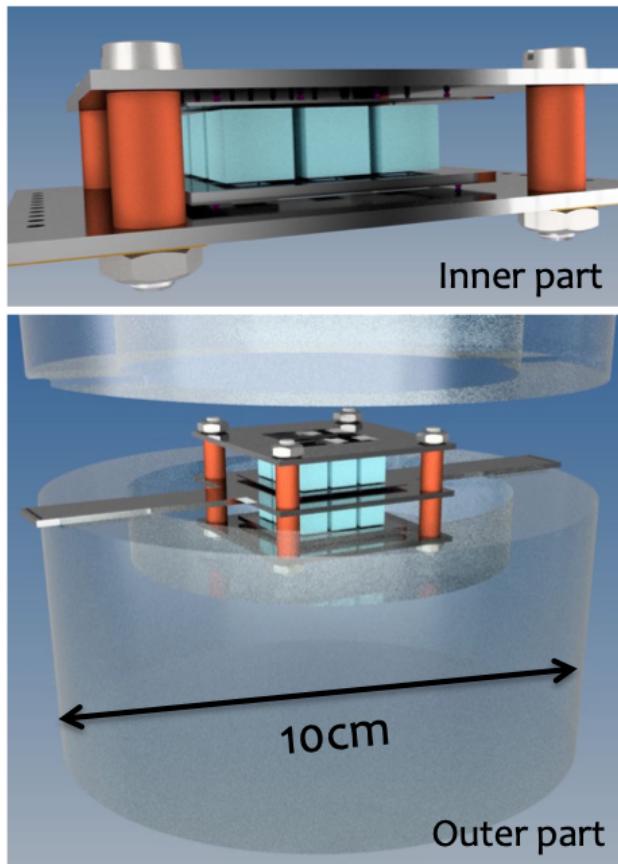
# NUCLEUS: site

- 10 g of CaWO<sub>4</sub> / Al<sub>2</sub>O<sub>3</sub> Target
- 18 detectors with 20 eV threshold
- CHOOZ very near site (VNS) (72 and 102 m from 4.5 GWth cores)



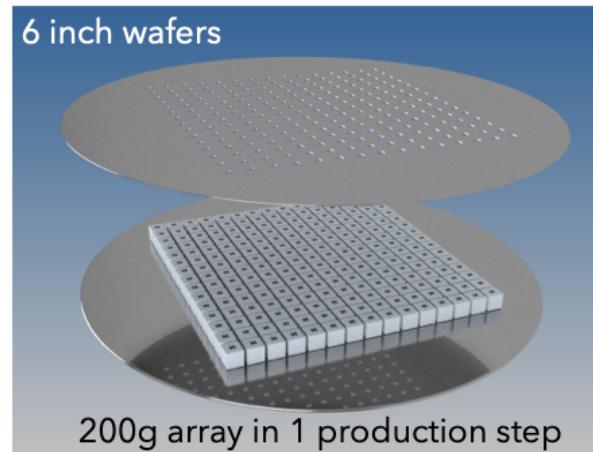
# NUCLEUS: 2-phase approach

NU-CLEUS 10g



NU-CLEUS 1kg

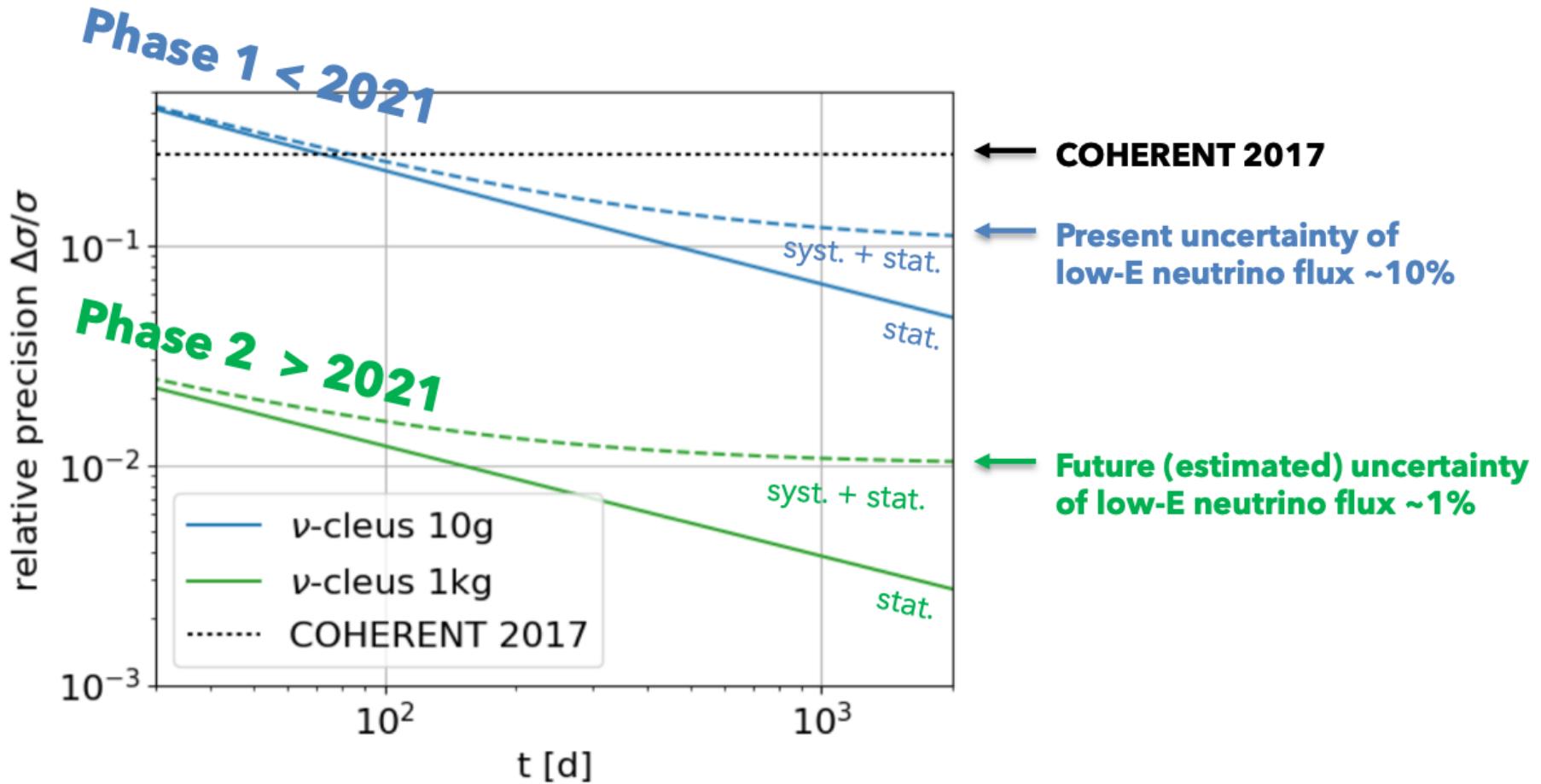
A scalable cryogenic detector



- Exploit semiconductor technology
- SQUID multiplexing

Breakthrough for the NU-CLEUS physics program!  
Impact for other cryogenic experiments!

# NUCLEUS: sensitivity



# NUCLEUS: collaboration



4.7 FTE + 0.4

**Roma1 - 2.8**

F. Cappella

A. Cruciani

L. Cardani

N. Casali

I. Colantoni

M. Vignati



Stefan Schönert  
Raimund Strauss  
Luca Pattavina  
Alexander Langenkämper  
Angelina Kinast  
Tobias Ortmann  
Elizabeth Mondragon-Cortes  
Lothar Oberauer  
Franz v. Feilitzsch



PD  
PD  
PhD  
PhD

PD  
(PhD)  
PhD with CEA (50/50)



Thierry Lasserre  
Matthieu Vivier  
Victoria Wagner  
Florence Adellier-Desage  
Loris Scola

Claudia Nones

+PhDs...



Federica Petricca  
Johannes Rothe  
Franz Pröbst  
Dieter Hauff  
Michele Mancuso  
Lucia Canonica  
Antonio Bento  
Leo Stodolsky  
Gode Angloher

Interested institutions:

- INFN, Sapienza
- IBS Korea??
- TU Vienna



Jochen Schieck  
Holger Kluck  
Florian Reindl  
Christoph Schwertner  
Vasile Ghete

**LNGS - 0.4**

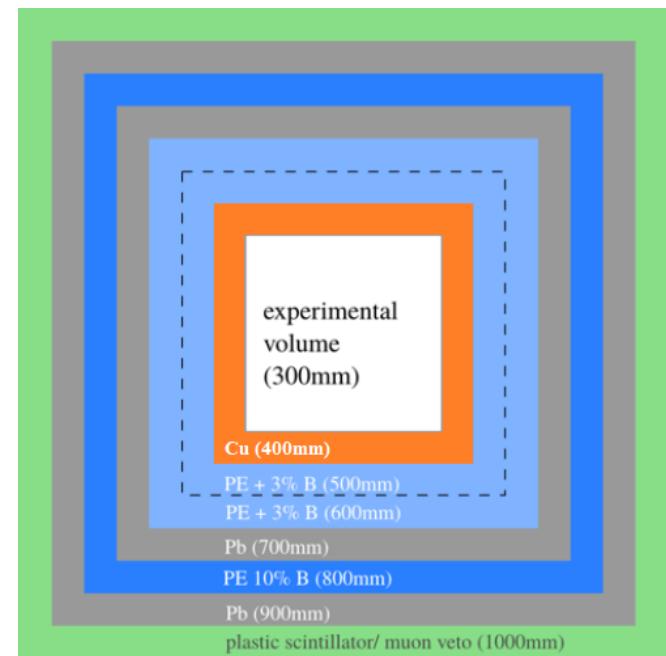
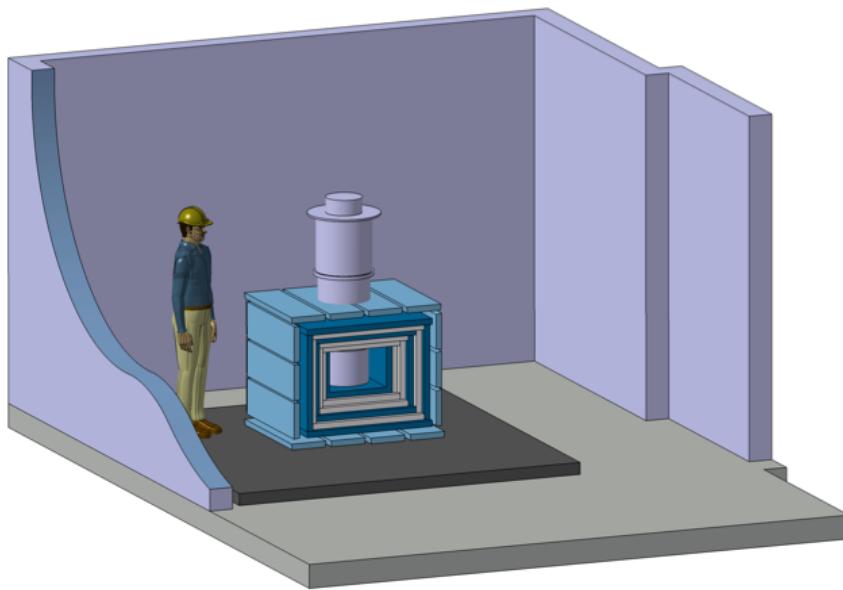
L. Pattavina

*on leave at TUM  
until 04/19*

Activity			2020	2021	2022	Cost
			semesters			k€
<b>NUCLEUS Phase 1 Summary schedule</b>	Procurement and construction					-
	Commissioning at TUM					-
	Installation and data taking at CHOOZ					-
<b>INFN contribution to NUCLEUS Phase 1</b>	3.1 Calibration system	RM1				100
	3.2 Inner veto	FE				70
	3.3 Shielding	RM2				50
	3.4 Material radioassay	RM2/ LNGS				5
	3.5 VNS bkg. characterization	RM1/2 LNGS				25
	3.6 Simulations	RM1/2				5
	3.7 Data analysis	RM1/2				
	Travels					108
<b>INFN R&amp;D</b>	4.1 Study of recoil yield	RM1/2 LNGS				n/a
	4.2 Study for <sup>51</sup> Cr (travels)	misc				16
<b>Total cost</b>						379

# External shield (Roma2)

- Cube of 1.1 m side external.
- Layers of 5 cm:
  - 500 bricks of lead 100 Bq/kg.
  - internal layers of (B?)-Polyethylene.
  - innermost layer of Cu, provided by TUM.



# MC Simulations (RM1, RM2)

- Already contributing to the simulation for the shield design (due by 12/19)
- Next contributions:
  - Evaluate the reliability of Monte Carlo codes at very low energy (Geant4 vs MCNP comparison).
  - Simulation for detector shielding design.
  - Simulation for the Veto system design and response.
  - Background model.

# Radioassay (RM2 - *LNGS*)

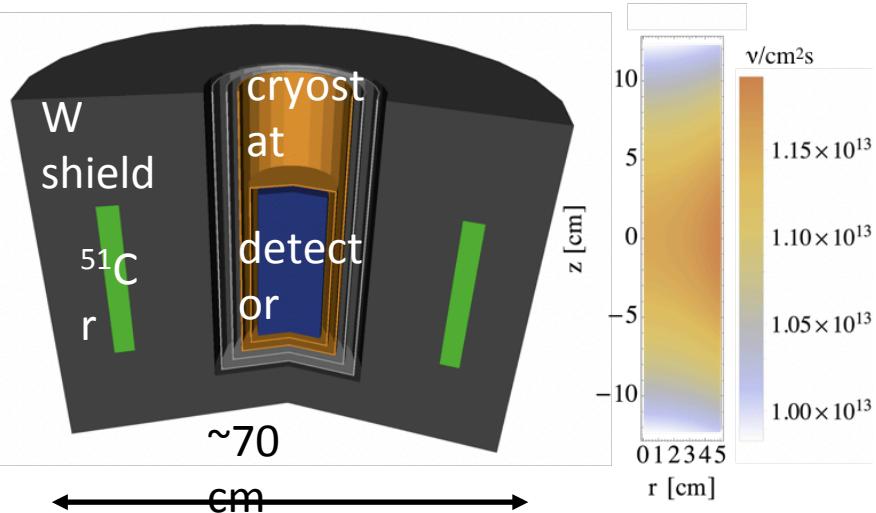
- Energy range 0-100 eV unexplored. Background not known a priori, estimation based on MC (reliable at such low energies?).
- Control of all possible background sources is mandatory  $O(\text{mBq/kg})$ .
- Measurements at MiB - Laboratorio di Radioattività
  - Detector components.
  - SQUIDS read-out system in proximity of the cryo-detector.
  - Shielding components.
  - Muon veto components

# R&Ds for Phase II

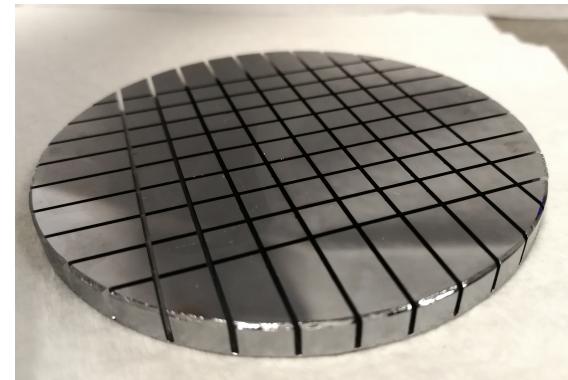
$^{51}\text{Cr}$ : Reuse of the 36 kg of  $^{50}\text{Cr}$  from GALLEX (INFN)

1% precision in 2 months on a 2 dm<sup>3</sup> Ge target

- Need neutron activation to  $^{51}\text{Cr}$
- Need detailed bkg. simulations



**BULLKID** (CSN5 2019-20): Array of KIDs to reach large targets



- 60 voxels of 5x5x5 mm<sup>3</sup>
- Grooved in a 5 mm thick wafer
- KID lithography side opposite to grooves

## Recoil yield

- Phase II target: 1% precision.
- The “no quenching” statement of phonon detectors need dedicated measurements at low energies (< 1 keV).

# NUCLEUS @ Tor Vergata

## Anagrafica

R. Cerulli              0.3

## Richieste

- Consumo: 5 kE
- Apparati
  - ✓ 30 kE Schermo di Piombo di bassa raioattività
  - ✓ 10 kE Polietilene
- Inventario
  - ✓ 5 kE Server per simulazioni e analisi dati

# Conclusions

- Coherent scattering of neutrino is a new field of research
- NUCLEUS has a well-defined scientific program (2020-22)
  - ▶ Could lead to world-best precision measurement of the x-sec.
  - ▶ Significant INFN contribution
- Next generation experiments (2023-):
  - ▶ Precision measurements
  - ▶ (%). Opportunity for leading INFN contribution.

