# Session Summary Neutrino, GW, Dark Matter, Dark Sector, Flavor and CPV



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Francesco Forti, INFN and University, Pisa





# Disclaimer and topics

- Lots of material, it is the summary of the summaries
  - No way to get in any detail (also because of limited knowledge)
  - Try to list interesting directions
- Three sessions:
  - Neutrino (accelerator and not accelerator), gravitational waves
  - Dark matter and dark session
  - Flavour physics and CP violation
- Comments on instrumentation and computing session
- https://cafpe.ugr.es/eppsu2019/



Neutrino Physics (accelerator and non-accelerator) summary of the session

**NOTE: including Gravitational Waves** 

### Conveners: Stan Bentvelsen, Marco Zito





# **Big questions**

What is the origin of the neutrino masses ? And of the leptonic mixing ?

What is the optimal strategy towards a complete set of measurements of neutrino oscillation parameters and towards a precision global fit of the PMNS matrix ?

Is the existing experimental program (reactor, SBL) sufficient to confirm or exclude the existence of sterile neutrino states with masses in the eV/c<sup>2</sup> range ? How to search for heavy neutral leptons with present and future facilities ? Is gravity described by the Einstein theory of general relativity? How do gravitational waves help to understand Dark Sector of the universe?

What is the proton-proton cross section at ultra-high energies?

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How can cosmic neutrino's help to pin-down their properties - oscillations and

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### Study of neutrino properties

### Today : few % precision on most of the parameters



Oscillation parameter		Best-fit		"1σ"	
		(NO)	(10)	error	
∆m <sup>2</sup>	/10 <sup>-3</sup> eV <sup>2</sup>	2.49	2.47	1.3 %	
δm <sup>2</sup>	/10 <sup>-5</sup> eV <sup>2</sup>	7.34	7.34	2.2 %	
sin <sup>2</sup> θ <sub>13</sub>	/10-2	2.23	2.24	3.0 %	
sin <sup>2</sup> $\theta_{12}$	/10-1	3.04	3.03	4.4 %	
sin <sup>2</sup> $\theta_{22}$		0.56	0.56	~ 5 %	

But there are still major unknowns : Dirac or Majorana, CP violation phase  $\delta$ , mass ordering and  $\theta_{23}$  octant, absolute mass

- Masses and oscillations
- Dirac or Majorana
- CP Violation
- New states



### Neutrino as a window to new physics

- The minimal extension is to have new elementary neutral fermions
- Neutrinos could have new interactions



### A very diverse experimental approach



### New long baseline projects

- Golden physics case for LBL: CP violation in PMNS would be a new source since 1964 discovery, possible link with leptogenesis, CP and MO linked to model building
- Since 2013 progress in the field had been faster than "rapid" convincing a growing community (>2000 worlwide) of physicists and the funding agencies to invest in these experiments
- Major players : DUNE, Hyper-Kamiokande, JUNO, ORCA



### Precision program in Europe

- Squeezing every bit of information out of the future experiments requires a complementary program (special rôle for Europe) to
  - Measure hadroproduction for the neutrino flux prediction (NA61)
  - Understand the neutrino-nucleus cross-section at the % level, both theoretically and with new facilities (Enubet, Nustorm)
  - Collaboration to be developed with nuclear physicists
- Next-to-next generation facilities (ESSnuSB, ...) are also under study











- Demonstration of membrane cryostats for LAr TPC
- ProtoDUNE-SP exposed to the SPS beam in 2018 and fully validated (purity, S/N, HV)
- ProtoDUNE-DP to be tested this summer

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Also T2K ND Upgrade, ENUBET, BabyMIND...

A game changer for the field worlwide !



### Measuring the neutrino masses

- Future cosmology missions (Euclid, DESI ...) target the neutrino mass detection ! (only parameter in common between particle physics and cosmology)
- Neutrinoless double beta decay : key to determine the neutrino nature. IO covered with the next generation of experiments (strong program in Europe: LEGEND, CUPID, NEXT ...)
- Single beta decay : KATRIN taking data. New exp. methods (Project8, ECHo) under development.

### Hunting new states (HNL)

- New neutrino states (Heavy Neutral Leptons) possibly related to neutrino masses and baryon asymmetry
- Their masses could be anywhere between eV and GUT scales !
- Very active area : searches based on LBL near detectors, beam dump, colliders, « symbiotic » expt.
- Unique sensitivity at low masses for the beam dump facility

### Are there sterile neutrinos?

- Motivated by long standing anomalies (LSND, MiniBoone, Gallium, reactor anomaly)
- No successful model and strong tension with other results (MINOS+, IceCube)
- A very rich program developing with SBL reactor neutrino experiments (Prospect, Solid, Stereo, Neutrino4, DANSS, NEOS). The results are ambiguous.
- Fermilab SBN three detectors program (SBND, MicroBoone, Icarus). 2020 results from MicroBoone
- We will know a lot more in the next couple of years





### Cosmic neutrinos

- Cosmic neutrinos are complementary to accelerator-based experiments : mass ordering with KM3NeT/ORCA and IceCube/PINGU.
- But also neutrino properties and interactions, tau neutrino, PMNS unitarity...

### **UHE Cosmic Rays**

- Flux suppression observed at ~10<sup>20</sup> eV
- Dipole distribution observed above 8 EeV : extragalactic origin
- Most models of UHECR from exotic sources are ruled out (monopoles, cosmic strings...)
- Proton-proton cross-sections
- The first EeV-neutrino should be 'just around the corner'

### **Multimessengers**

- Multimessenger : combining information from photons, neutrinos, CR, GW.
- Recent highlights : Neutron star merger (GW+photon), Blazar (neutrino's +photon)
- Many physics topics : dark matter searches, expansion history of the Universe, Lorentz symmetry, ...





### **Gravitational waves**

- Gravitational waves open a new field of research
  - Testing General Relativity in extreme conditions
  - Probing Black hole as Dark Matter
  - Equation of state in neutron stars
  - Serendipity : exotic objects, new fields and phenomena
- Einstein Telescope : 3rd generation of interferometers. Ultimate facility to probe the entire universe in the 1-100 Hz frequency range
- ET has put on the table a request for CERN to help with
  - Underground infrastructure, cryogenics
  - Vacuum, material and surface science
  - Electronics and data acquisition, computing



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### Neutrino session take-away message

- Neutrino oscillations, mass ordering, CPV phase:
  - Ongoing program DUNE, Hyper-Kamiokande, JUNO, ORCA
  - Need cutting-edge detectors and % precision on the flux and cross-sections: NA61, Neutrino Platform. New facilities under study.
  - Long term future for high precision LBL measurements with new techniques.
- Neutrino masses
  - Absolute mass is still unknown (m<~0.1 eV!). Laboratory (Katrin) and cosmology missions ongoing.
  - Dirac or a Majorana fermion ? Neutrinoless double beta decay searches
  - The neutrino mass might require new neutral fermions with masses between eV and GUT scale (!)
    - Clarify eV scale anomalies (SBL program at FNAL, reactor neutrino experiments)
    - Search for new states at the GeV-EW scale (beam dump and colliders)
- Astroparticle physics
  - Gravitational waves and multimessenger physics open up a new window on the Universe. Very strong physics case.
  - There is a very high impact on the field of particle physics (and fundamental interactions) (eg dark matter, neutrinos, general relativity, ...)
  - There is clearly an opportunity for the particle physics community and laboratories to expand their involvement in this program





# Dark Matter - Dark Sector Summary

Shoji Asai and Marcela Carena

on behalf of the Dark Matter–Dark Sector Group

B.Dobrich, J.Jaeckel, G.Krnjaic, K.Petridis and K Zurek



# Dark Matter/Dark Sector Introductory talk [H. Murayama]

### Dark Matter exists, awaiting for discovery













- How do we search for DM, depending on its properties? What are the main differences between light Hidden Sector DM and WIMPs? How broad is the parameter space for the QCD axion?
- 2) What are the most promising experimental programs, approved or proposed, to probe the different DM possibilities in a compelling manner?
- 3) How to compare results of different experiments in a more modelindependent way?
- 4) How will direct and indirect DM Detection experiments inform/guide accelerator searches and vice-versa?





### Very little clues on mass scale





### **Dark Sectors**

What is meant by a dark sector ?

A Hidden sector, with Dark matter, that talks to us through a Portal



### Indirect detection

No conclusive signals from indirect DM searches so far

But, slow and steady progress being made on indirect searches in many fronts

- Diffuse gamma rays, e.g Galactic Center GeV Excess
- 3.5 keV "Sterile Neutrino decay"
- Antiproton excess from cosmic rays
- Neutrinos from DM annihilation in the Sun



It is possible that in the future it will be a convincing signal from one or more indirect DM searches

This will have large impact on Direct Detection and Accelerator based DM searches











# WIMP Standard Candles

- Still a viable solution for Thermal DM (e.g. in many SUSY extensions/ regions)
- Being broadly probed by Direct and Indirect detection as well as Collider experiments

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#### **Pure Wino DM**

- Thermal abundance requires Wino mass of about 2.9 TeV
- DD: just above the neutrino floor. Ballpark of DarkSide 20k-200t-yr, DARWIN 200t-yr and Argo 3000t-yr.
- ID: Wino only constitutes all the DM for density profiles not generically produced in simulations of Milky Way-like galaxies
- @ Hadron Colliders: Disappearing tracks
- @Lepton Colliders: Reach close to kinematic limit plus
   precision measurements extended reach

#### **Pure Higgsino DM**

- Thermal abundance requires Higgsino mass of about 1.1 TeV
- DD: Suppressed. Deep in neutrino floor region
  - ID: Bounds strongly dependent on halo morphology.
  - @ Hadron Colliders: Disappearing tracks
  - @Lepton Colliders: Reach close to kinematic limit plus
     precision measurements extended reach





# Accelerator-based searches

- Dark photons/scalars to SM particles
  - Dark portals at beam dumps (NA62++, SHIP)
  - Dark portals at fixed target (NA64++, LDMX) – missed energy
- Heavy neutral leptons • Neutrino masses ?
- Invisible dark photon decays
  Several techniques also at colliders



DM Scatters in detector

LLP DECAY



Invisible decay





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# Axions/ALPs

### • QCD Axions as DM ?

• Search for axion-photon coupling

#### QCD axion

- Solves Strong CP problem
- Natural Dark Matter candidate Important target for experimental searches

Many Beyond the Standard Model scenarios contain

- Axion-like particles
- General light dark sector particles

**Probe wider parameter regions including different scenarios** 

Motivates a variety of non-accelerator experiments

#### **Helioscopes**



 Proposed BabyIAXO, leads to IAXO, with large discovery potential

#### Haloscopes



Mature Key Tec

- ADMX (US) is leading the field
- In Europe, MadMax is new key player
- Smaller efforts developing new techniques

#### Light-shining-through-walls

- ALPS II is well underway
- STAX is a new idea RF based
- JURA is long term plan



Lindnor and Iractorza's talks









#### Need for better coordination

- Consensus emerged on the need for more coordination between accelerator based, direct detection and indirect detection dark sector searches, for common interpretation of results.
- This will also be of fundamental importance to validate, through different channels, a possible dark matter discovery.
- To address this issue, it was recommended to make profitable use of the initiative of APPEC on the EuCAPT Astroparticle Theory Centre.
- This offers a strong opportunity to collaborate with working groups such as the LHC DM and Physics Beyond Colliders and the many recognized dark sector experiments using different approaches

# Comments from discussion

#### Need for technology support and exchange between communities

- Technology challenges are shared between and beyond the communities engaged in dark matter searches.
- CERN and other large European National labs has relevant expertise and infrastructure for most/many of the big challenges, including vacuum over large volume, cryogenics, photosensors, liquid argon detectors, design and operation of complex experiments, software and data processing.
- Expanded support for dark matter research at CERN would stimulate knowledge transfer, increase coordination and synergies between experiments, and add guidance and coherence to the overall program.



# Dark sector session take away message

- Change of paradigm:
  - OLD: a good BSM theory may provide an explanation for DM
  - NEW: understanding DM may provide a clue to BSM theory
- Enormous range of possibilities
  - Use all available experimental techniques
  - Complementarity of different approaches
  - Evaluate correlated impact of different measurements

- Dark Matter exists, awaiting for discovery
- In general, Dark Sector may exist, too
- Very little clue on mass scales now
- WIMP still main paradigm, reach  $\vee$  floor
- Many new ideas on lighter dark matter
- Colliders, beam dump, underground, cosmic rays, cavity, new technologies
- Vibrant area and need more data!





# Summary of the flavor session Antonio Zoccoli and Belen Gavela

### Flavor Physics → BSM

- EW Hierarchy... driven by the top in SM
- Strong CP problem
- Origin of weak CP and matter-antimatter asymmetry
- Flavour puzzle (quarks, charged leptons, neutrinos)

Flavour is the usual graveyard of BSM electroweak theories





### 1) The light sector $(u,d,s + e, \mu)$

Three "clear" cases calling for diversity in the short/mid-term:

1. EDMs [d <sub>e</sub> , d <sub>n</sub> , d <sub>N</sub> ,.]	Strong CP EW CP	new: Storage rings
2. μ→e processes	1 → 2 Gen Lep. Mix.	Intense µ beams
3. Rare K decays	1 → 2 Gen. Quark Mix.	Intense K beams

 Outstanding physics goals (fundamental & unique) Difficult experiments, but on a small scale

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beamline to measure hadronic



Rare K



- Forbidden/LFV decay searches can probe unreachable mass scales, will mostly come as byproducts
- Ultra-rare K decays are the most-promising driving goal in the field: experiments are hard and different shots must be taken at them
- After fully exploiting the existing p machines (CERN) SPS, J-PARC) with suitable sharing to fixed-target intensity frontier experiments, a very intense (MW) slowly-extracted proton machine would allow ultimate NP searches

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 $K \rightarrow \pi v v$  possible future

Now using 50% of protons

Extension of Hadron **Experimental Facility:** considered among IP priority projects (not yet approved nor funded), O(140) M\$: 2 new targets, 4 new lines



## 2) The heavy sector (b, c, t + $\tau$ + h)

Bright near-term future [~ 10 yrs] with LHCb (I) & **Belle-II** 

This is likely to be the most exciting frontier of particle physics in the next years, with large discovery potential (wide parameter range still to be explored)

Further "enriched":

→ lattice will keep pace with experimental precision  $\rightarrow$  present anomalies





# Rich flavor program



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### Belle II+1= Belle III

Just started within Belle II

Goal: x5 increase in peak luminosity

- Doable from a machine perspective?
- Detectors issues running at 4 10<sup>36</sup> cm<sup>-2</sup> s<sup>-1</sup>

Physics case

Under study, more before the end of 2019

### Z<sup>0</sup> factories

Goal: 10<sup>11</sup> – 10<sup>12</sup> Z<sup>0</sup> (CEPC) 5. 1012 Z<sup>0</sup> (FCCee)  $BR(Z^0 \rightarrow b\overline{b}) = 15\%$ 

ILD-like detector + charged hadron PID.

FCC-pp a dedicated experiment (à la LHCb)

e+e-Super Charm-Tau Factories: SCT (BINP, Novosibirsk) and STCF/HIEPA (China) E: 2 to 6 GeV Peak Luminosity (> 4 GeV) 1035 cm-2 s-1



## $\tau$ physics topics

- Searches for Lepton Flavor Violation in tau decays:
- clean & effective search for "natural" NP processes extremely suppressed in SM
- upper limits are effective constraints on NP models
- Lepton Universality tests:
- EW precision tests
- effective constraints on NP models
   Many topics where τ is a decay product (H, Z, W, B etc.) (→ EW group)

# $\tau$ physics plans of relevant factors

Present with many LFV studies:

- Belle-II
- LHC: (LHCb, ATLAS, CMS)

### Possible future facilities:

- HL-LHC, HE-LHC many improvements
- Super Charm Tau Factories
- TauFV [@ SHIP beam line]:





# 2) The heavy sector (b, c, t + $\tau$ + h)

Bright near-term future [~ 10 yrs] :

• LHCb (I) & Belle-II

Clear mid-term future [~ 10-20 yrs] with an interesting case:

 LHCb-II [@ HL-LHC] and Belle-III (?): ultimate precision on "all- visible" B and D decay modes [B<sub>s,d</sub> → μμ, R<sub>K,K\*</sub>, CKM, charm CPV, ...]

possibly complemented by specific initiatives such as:

- TauFV [@ SHIP beam line]: τ→3μ @ 10<sup>-10</sup>, outstanding goal, competitiveness with Belle-II to be clarified
- STC @ Novosibirsk or HIEPA [competitiveness after Belle-II less clear...]



# 2) The heavy sector (b, $c + \tau$ )

A strong case for the long-term future [ > 20 yrs]: Flavor physics @ FCC-ee [running at the Z-pole , with  $5 \times 10^{12}$  Z's !]

Unique potential on b and  $\tau$  decays with missing energy, from Z  $\rightarrow$  bb,  $\tau\tau$ 

Just one example: B→K\*ττ : ~1000 SM events @ FCCee vs. ~10 @ Belle-II

And of course FCC-ee would allow major improvements on EWPO & Higgs

→ ideal set-up to optimize indirect NP searches





# Higgs and flavor



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# Flavor session take away message

- Flavor physics crucial for BSM search:
  - Outstanding BSM scale reach à  $\Lambda>10^2$   $10^5$  TeV, especially for lepton flavor violation and universality.
  - Complementarity of low-energy, HE frontier and feebly interacting searches
- Many possibility in small, medium, large scale facilities
  - Precision EDM, g-2
  - e+e- colliders: B-factories, tau-charm factories
  - Several dedicated experiments
  - Vibrant upgrade program for the near and mid-term future
- Flavour is also a major legacy of LHC:
  - Many results from all experiments
  - Charged hadron PID is mandatory for a full physics program.
  - Essential that HE future experiments follow this same path
- Important to have experiments in very different environments (pp and e+e-), and with PID
- In the longer term: Z-factory is a fantastic tool for Flavor Physics



Instrumentation and Computing Session Xinchou Lou, Brigitte Vachon

### Instrumentation:

(a) What areas of instrumentation R&D should be supported, and how, in order to meet the needs of future experimental programs?

(b) How to preserve knowledge, technical expertise and train the future generation of experts in detector R&D?

### Computing:

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(a) How should **HEP computing evolve** in order to support future scientific programs and their specific needs?

(b) What **R&D activities** must be supported, and how, in order to enable this computing evolution?

# Instr/Comp General comments

- Detector/computing R&D needs strong support.
- Maintain/develop proper mixture of activities related to "now, next, horizon" (yes, even for computing)
- General opinion that R&D should not be centralized exclusively in large-scale facilities and/or in major labs.
- Current R&D collaborations (eg. RDx, AIDA2020, CALICE, etc.) seen to be effective models of collaboration.
- Need support to strengthen R&D consortia and establish new ones while preserving expertise and working relationships
- Provide opportunities to build on, and benefit from, synergies with non-HEP and industry
  - (e.g. through technology-centered R&D platforms across disciplines providing)



# Human factor

- Training: Students/postdocs often lack basic knowledge (computing skills and instrumentation).
  - University courses are often insufficiently oriented towards these technical aspects.
- Career recognition/opportunities: Instrumentation and Software & Computing activities need to be recognized as fundamental to research activities and bearing a large impact on the final physics results.
- Special considerations in computing: Teams working on S&C need to be well-rounded, including physicists and developers.





# THE FUN BEGINS...

# DIR ROAD AHEAD WHERE THE **PAVEMENT ENDS**

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