

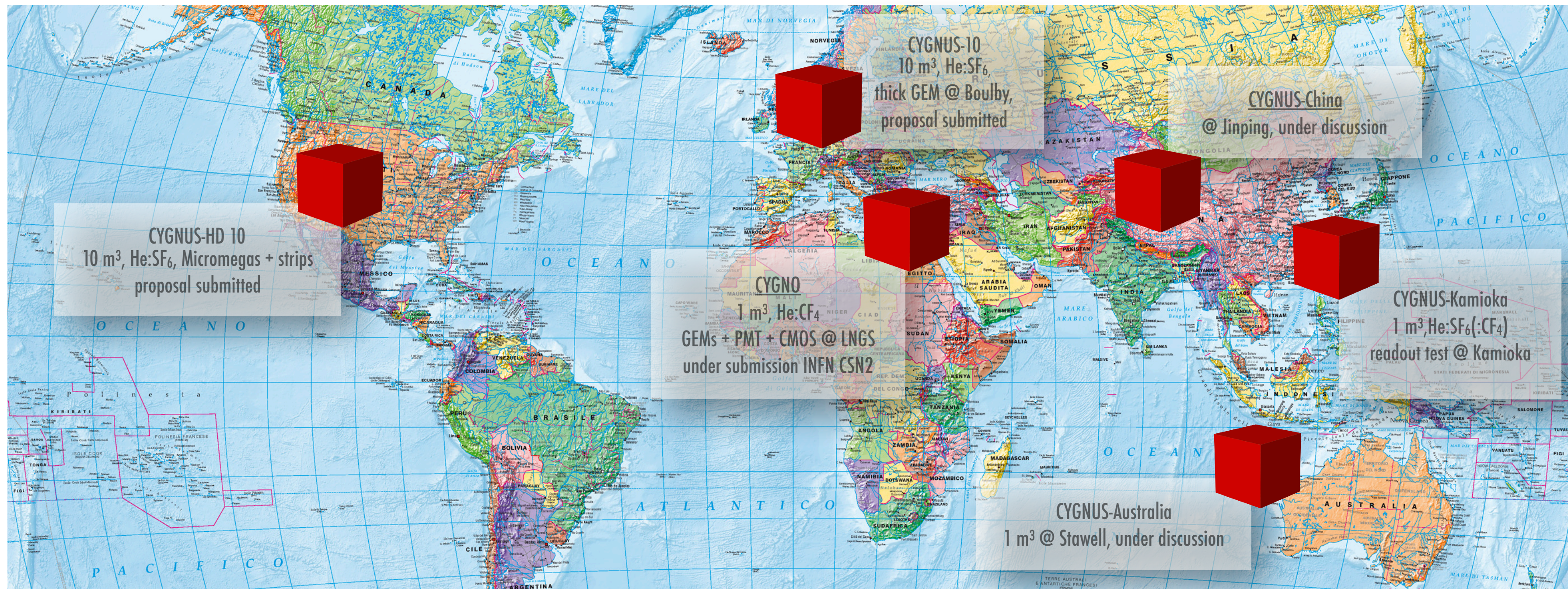
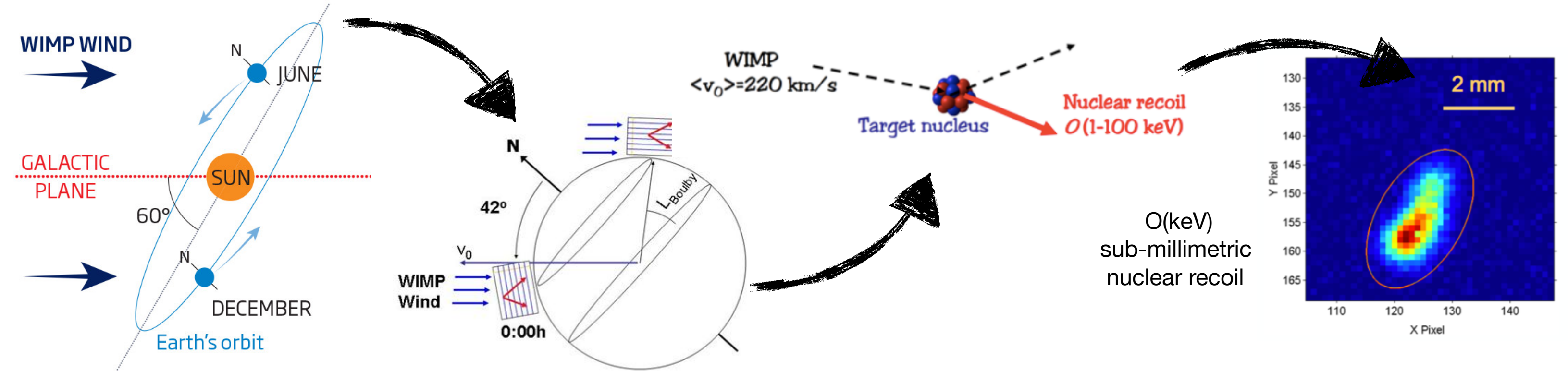
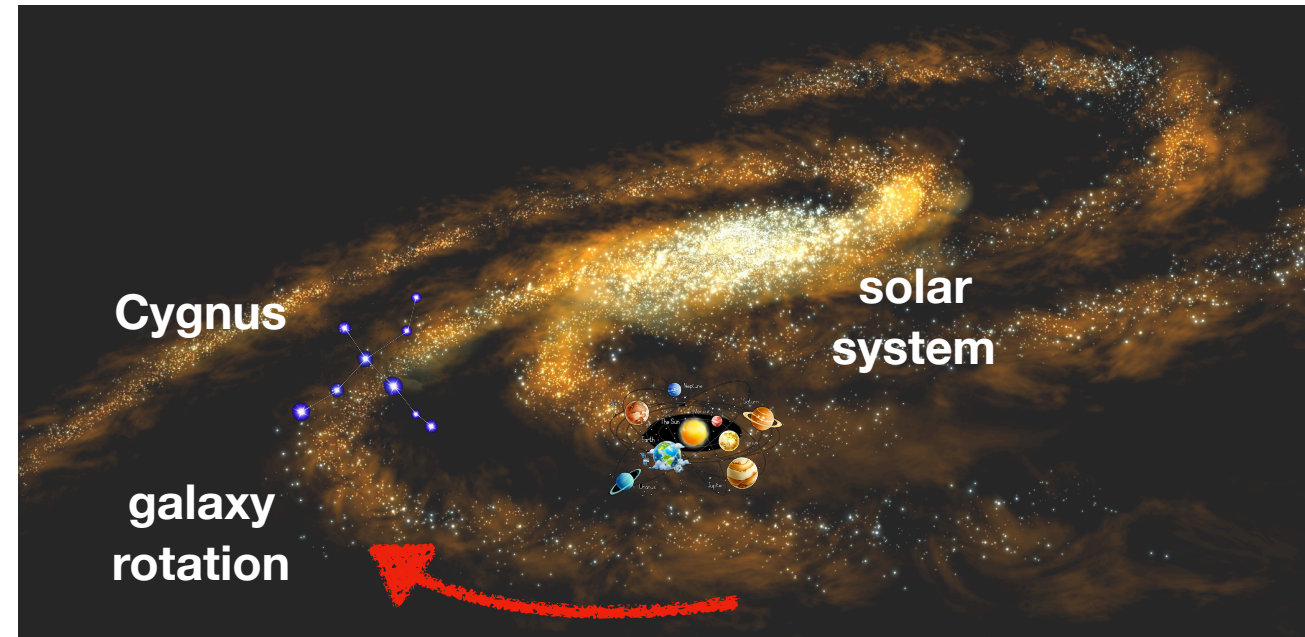


CYGNO/INITIUM project

a CYGNus TPC module with
Optical readout

G. Mazzitelli, E. Baracchini, R. Bedogni, F. Bellini, L. Benussi, S. Bianco, L. Bignell,
M. Caponero, G. Cavoto, E. Di Marco, C. Eldridge, A. Ezeribe, R. Gargana, T. Gamble,
R. Gregorio, G. Lane, D. Loomba, W. Lynch, G. Maccarrone, M. Marafini, A. Messina,
A. Mills, K. Miuchi, F. Petrucci, D. Piccolo, D. Pinci, N. Phan, F. Renga, G. Saviano,
N. Spooner, T. Thorpe, S. Tomassini, S. Vahsen.

CYGNUS collaboration objective



Galactic Nuclear Recoil Observatory

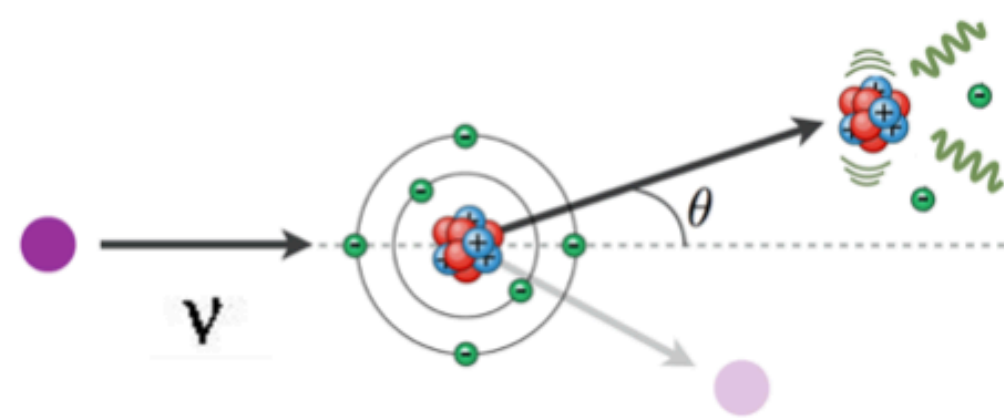
- Energy threshold 1 keV_{NR}
- Target mass 100-1000 kg (F, He)
- Zero neutron background
 - no steel (vacuum) vessel (acrylic?)
 - ceramics; almost no internal electronics
- x, y, z fiducialisation and radon rejection
 - either negative ion drift or other technique
 - material selection and scrubbing is not enough
- Gamma discrimination below 10 keV_{NR}
- Directional sensitivity

Physics motivation

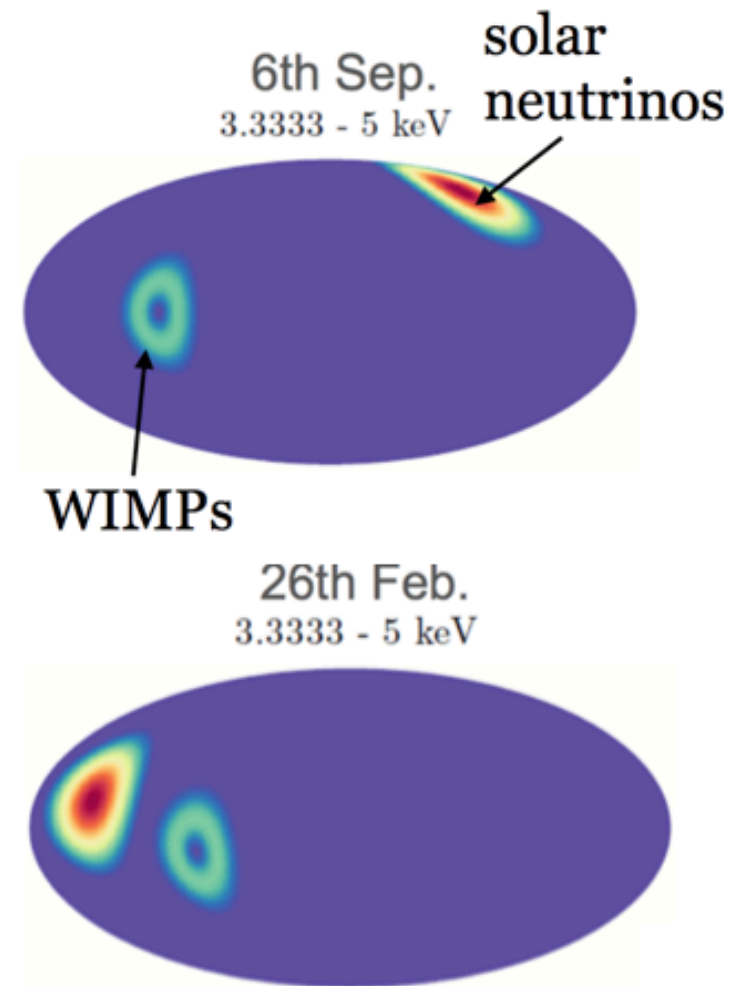
- low energy nuclear recoil $O(10 \text{ GeV})$, dark matter
- low energy electrons recoil $O(10 \text{ keV})$, light dark matter
- low energy neutrino scattering, solar physics

NOTE: only a directional detector can distinguish from WIMP signal

Coherent Neutrino-Nucleus scattering

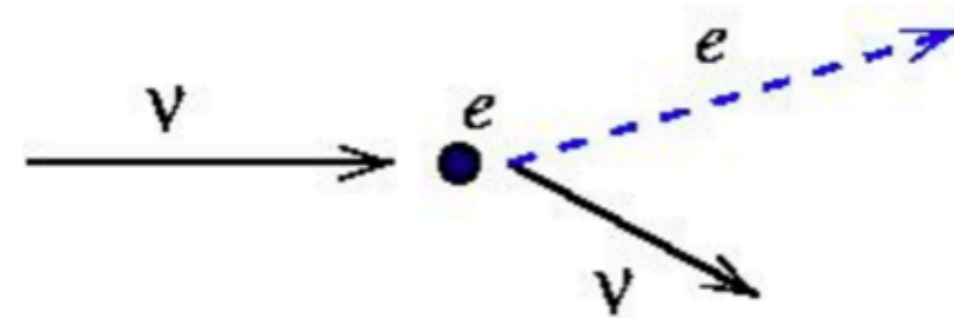


C. O'Hare et al, Phys. Rev. D 92 063518 (2015)

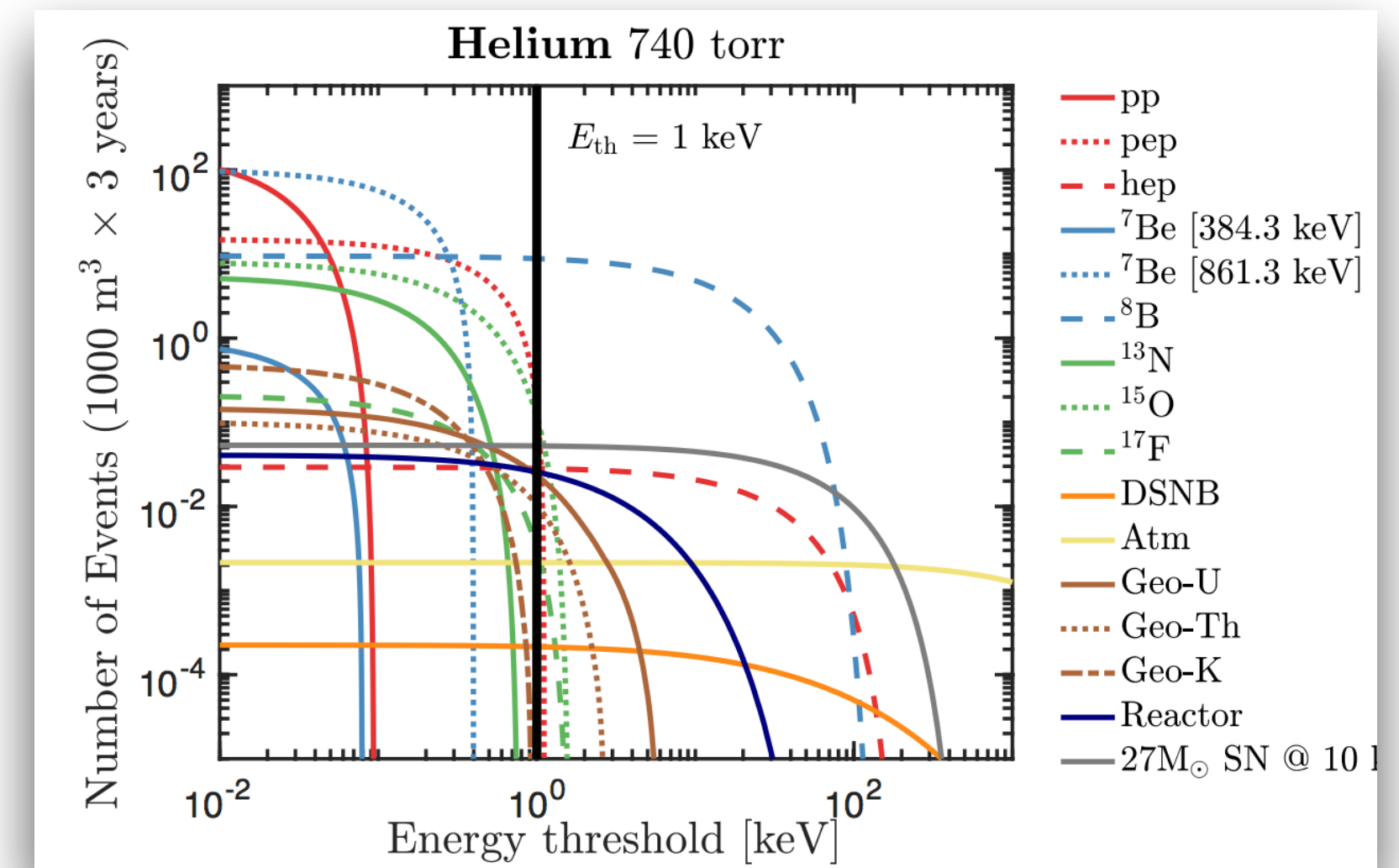


NOTE: only a directional detector can distinguish from ER background

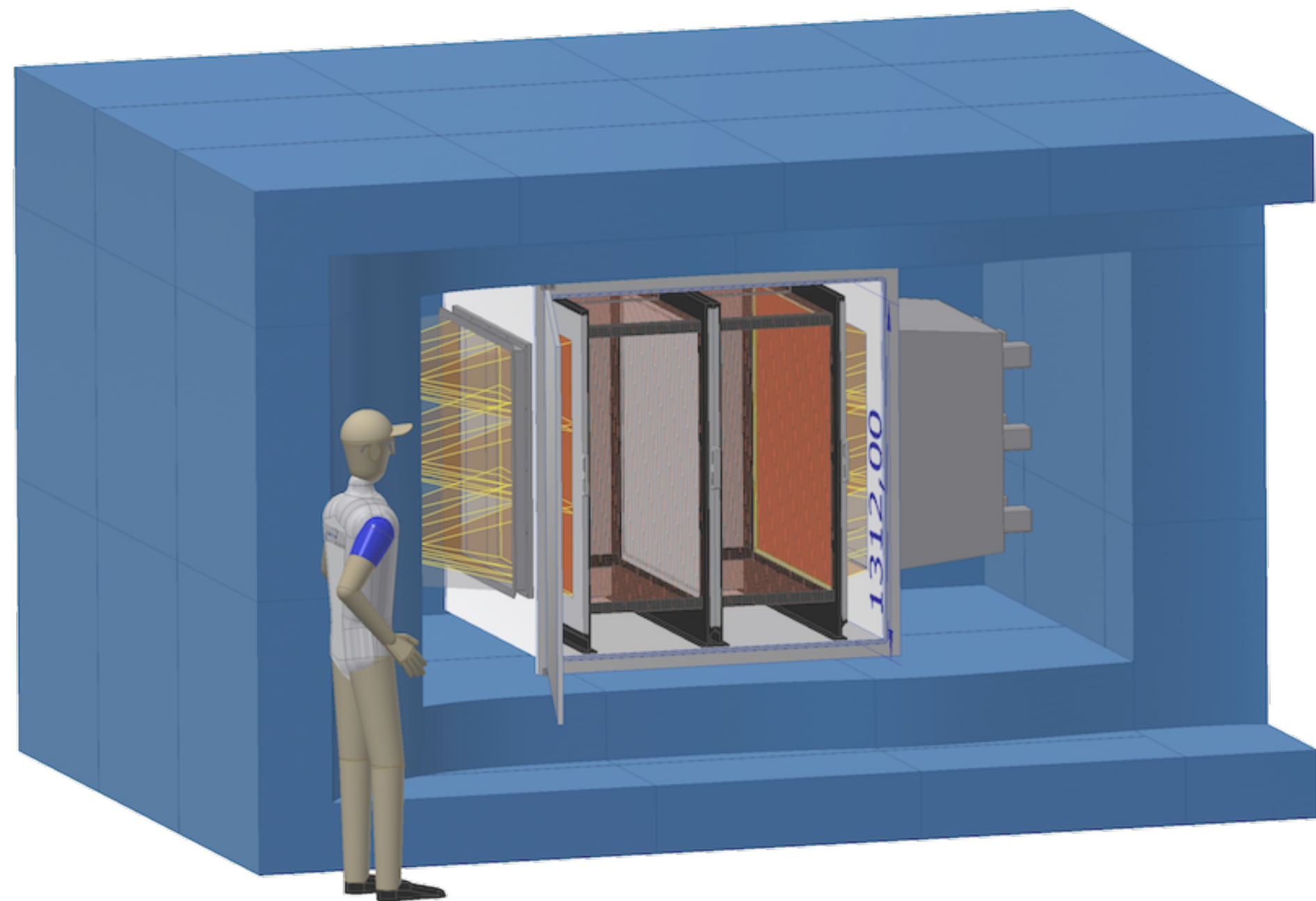
Elastic Neutrino-Electron scattering



E. Baracchini et al.



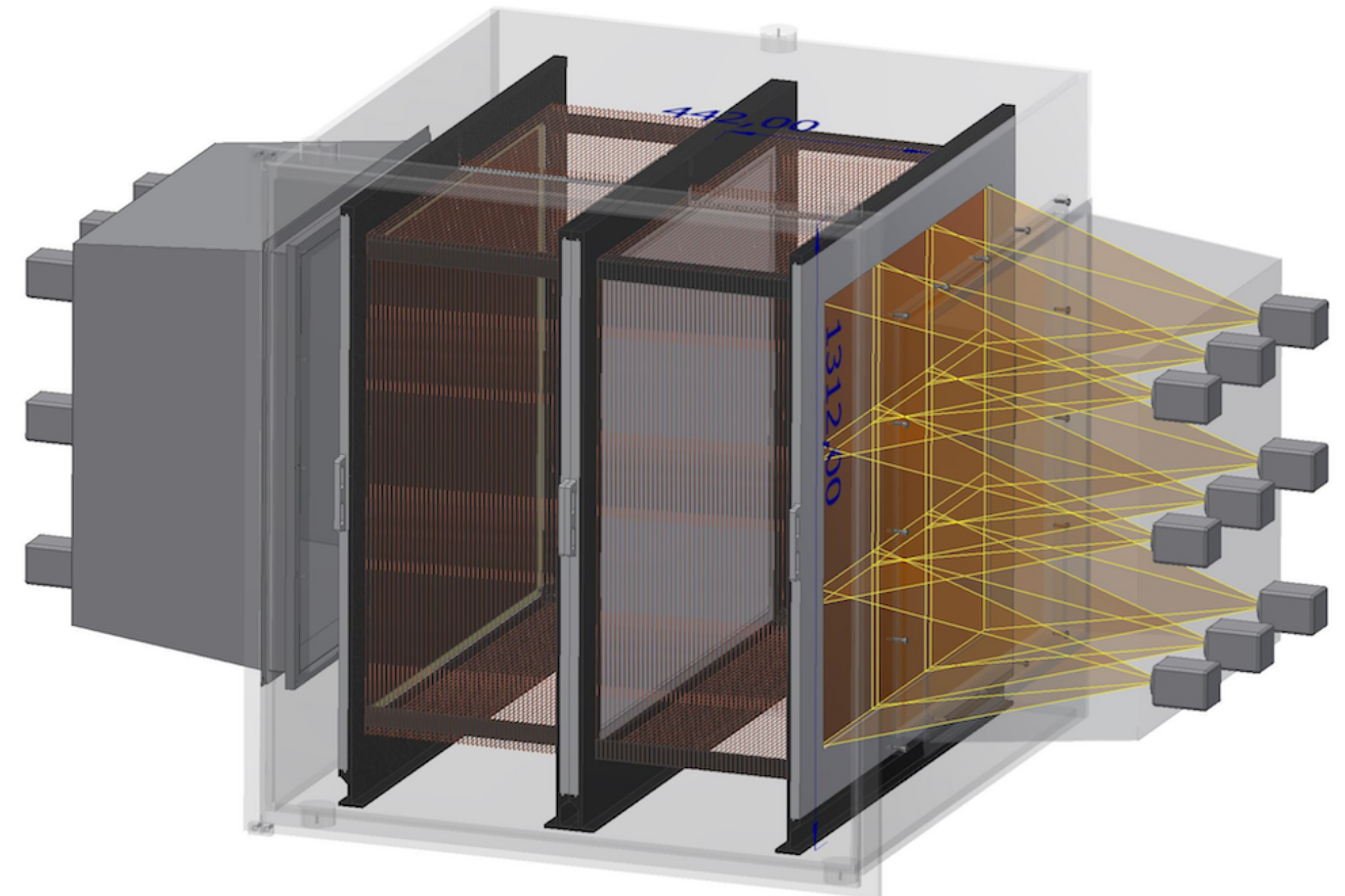
The CYGNO/INITIUM project



18 cameras monitoring
330*330 mm each
with **150 μ** resolution and
a sensitivity of **$\sim 1 \text{ ph} / 20 \text{ eV}$** released in gas

CYGNO is a demonstrator exploiting large gas TPC, GEM based charge amplification, high granularity and sensitivity of optically readout at atmospheric pressure in HeCF_4 based gas mixture

INITIUM-ERC is an R&D for testing possibility to improve nuclear recoil threshold and directionality by means of negative ions gas mixture in CYGNO demonstrator



**A total of 95 10^6 readout
15 x 15 mm^2 pixels**

The demonstrator strategy

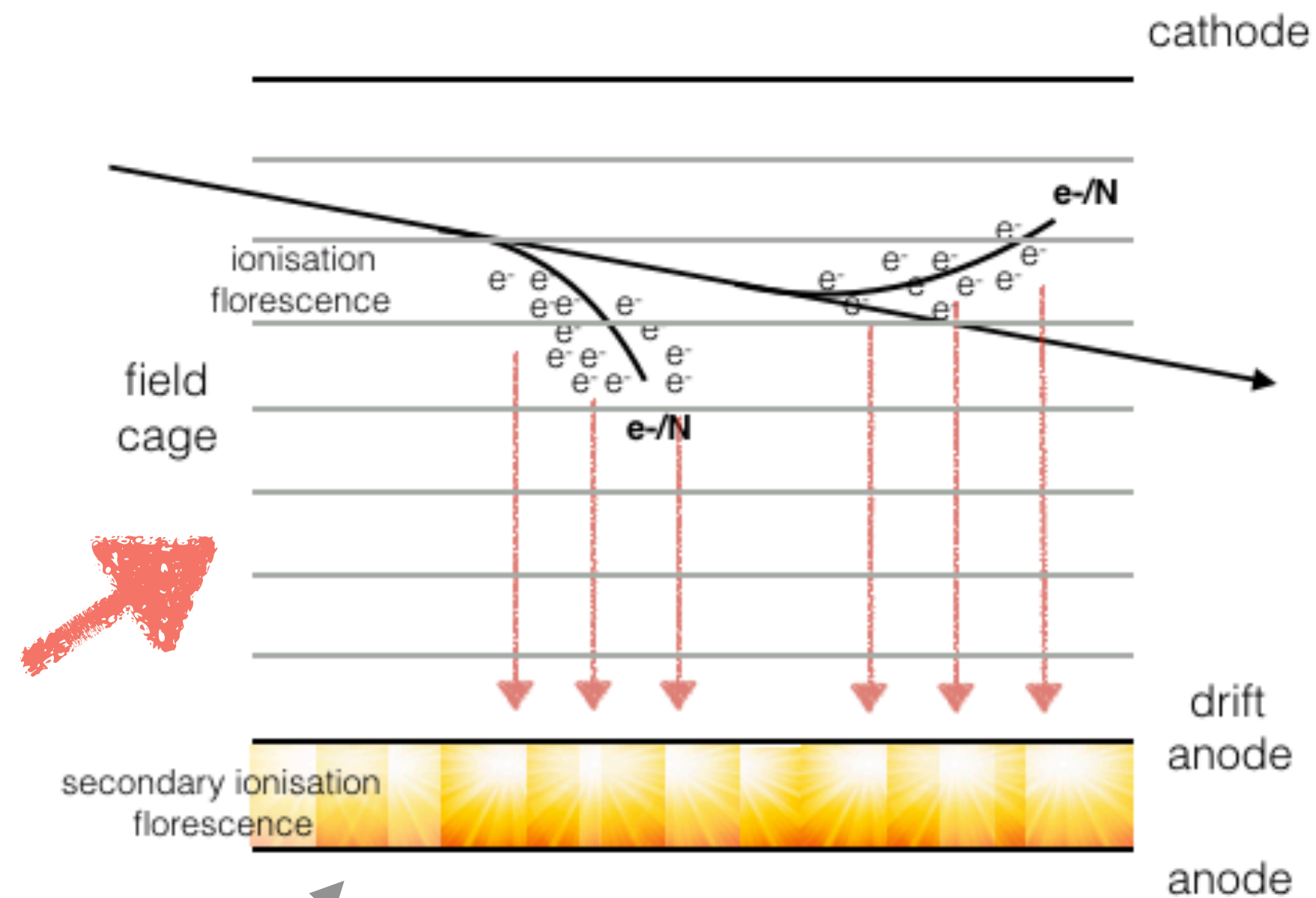
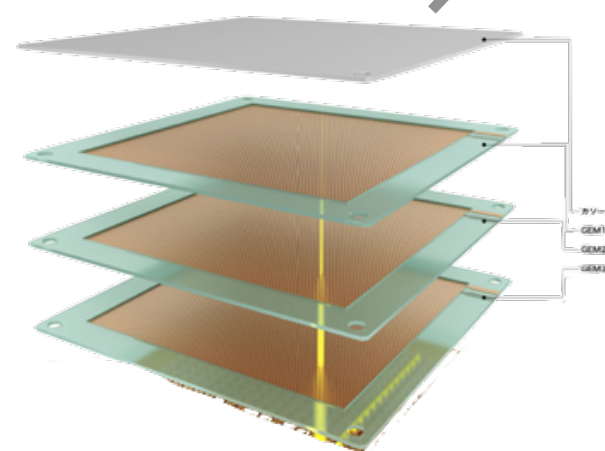
Time Projection Chambers provide:

- **3D** tracking (position and direction);
- total **released energy** measurement;
- **dE/dx** profile (pid, head-tail);
- reduced readout channel number;

gas represents an interesting target:

- nuclei free path can be **long** enough to be reconstructed;
- **low mass** gases allow an efficient momentum transfer from light DM;

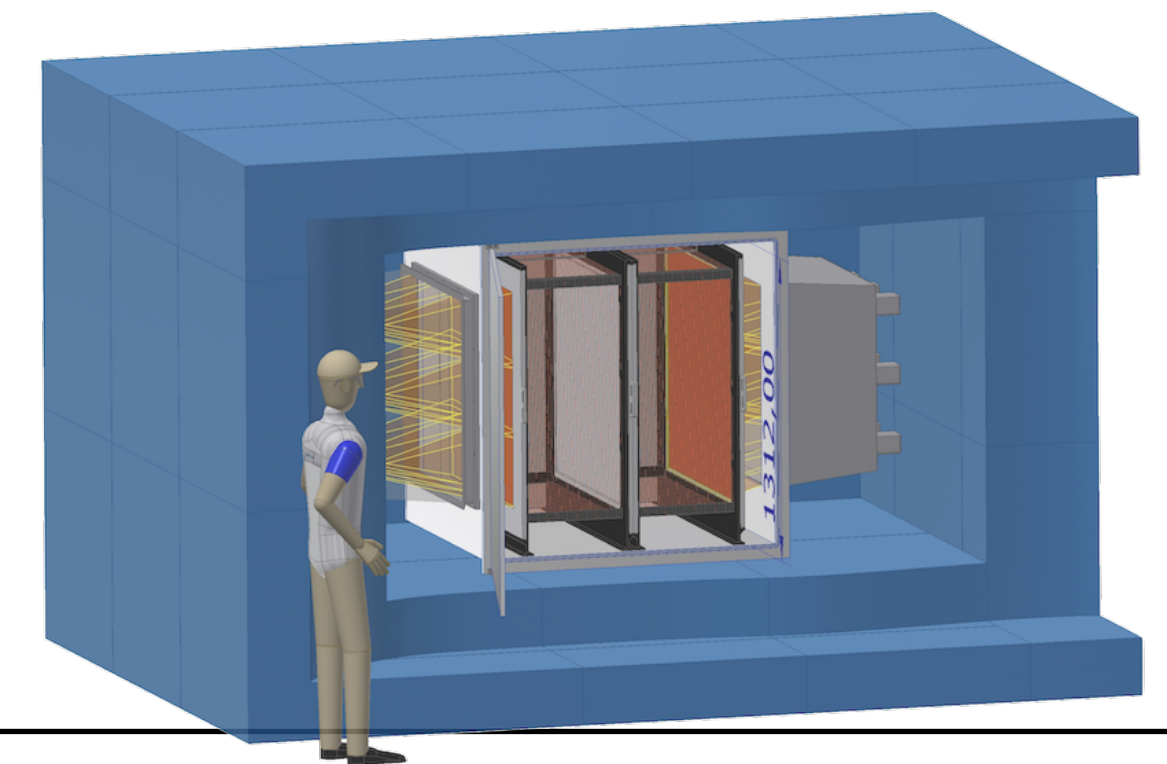
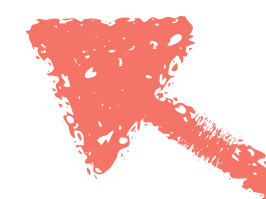
multiple GEM structures is used to obtain **gain** and **stable** detectors.



PMT



CMOS camera



high granularity (CMOS+PMT)
optical read out:

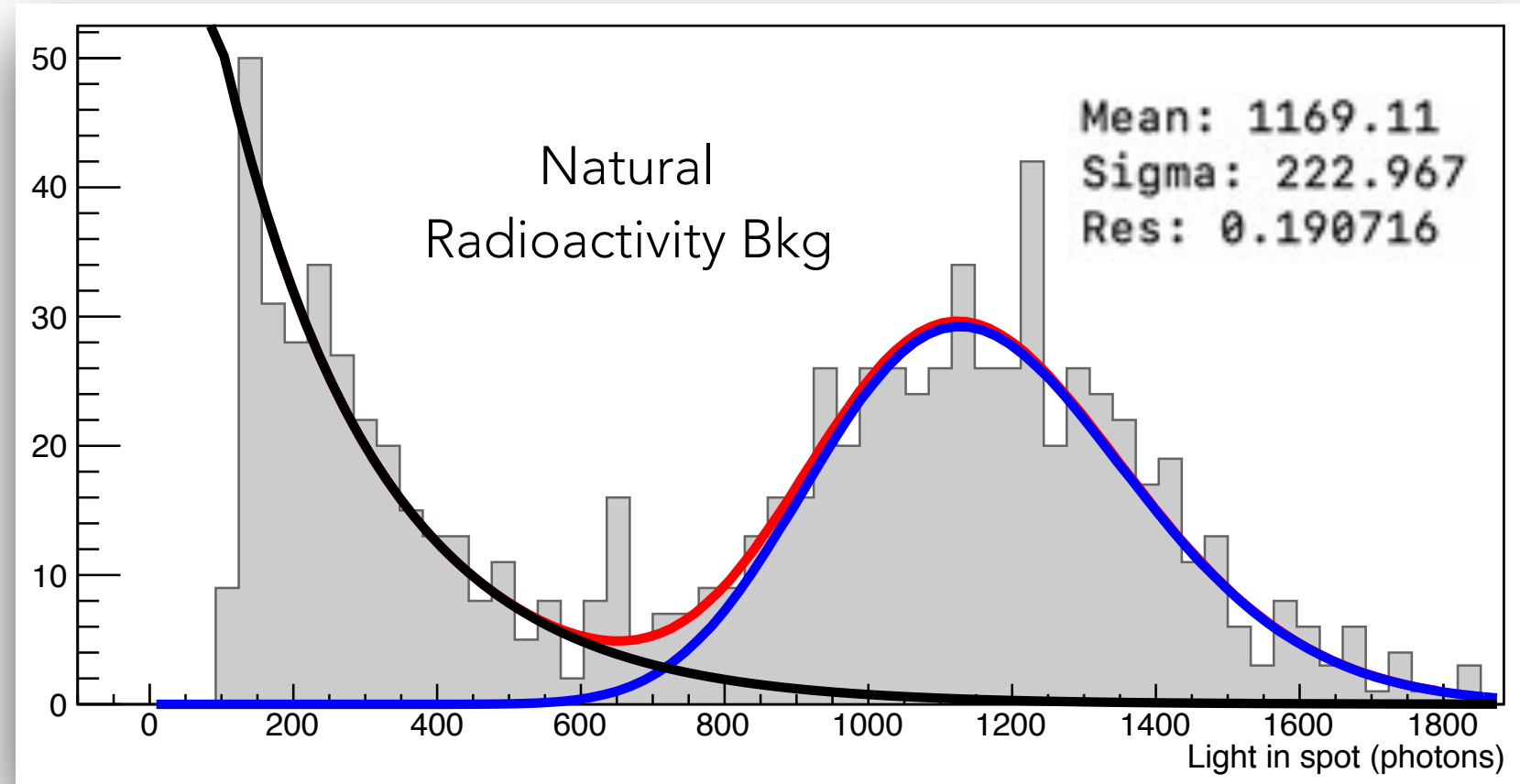
- threshold
- discrimination
- directionality;
- x, y, (z) fiducialisation
- electronics decoupling

atmospheric pressure He gas mixture:

- high target density (low threshold)

Energy threshold and sensitivity

JINST_024P_0519



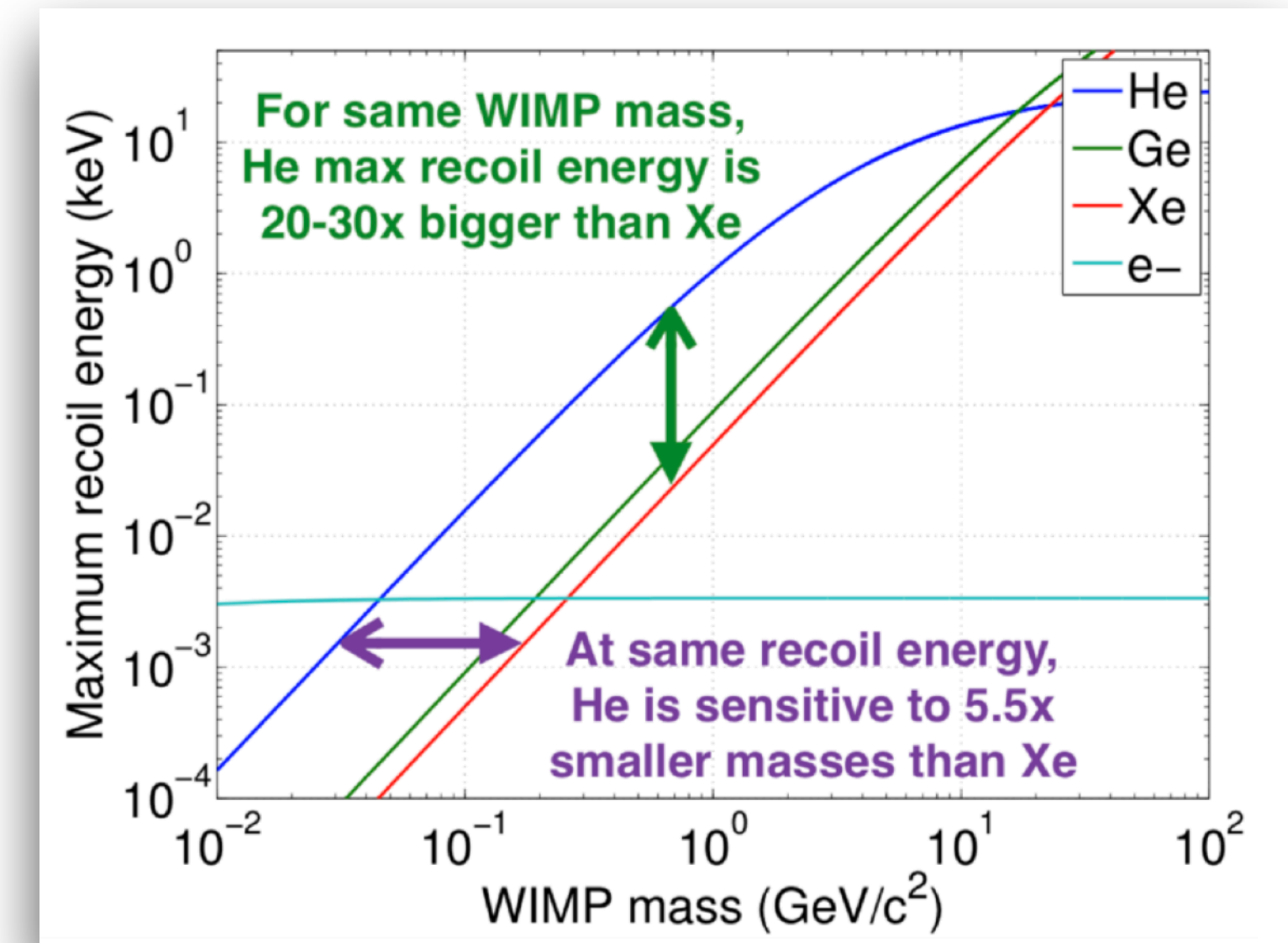
About 1170 photons are detected:
i.e. 1 photon each 5 eV released.

The **operative threshold** value is one of the parameter to evaluate the ultimate sensitivity to WIMPS. Moreover, crucial values is represented by the **mass of the target nucleus** (m_N), given the ratio between the m_N and the m_{WIMP} , the maximum transferred energy fraction is given by:

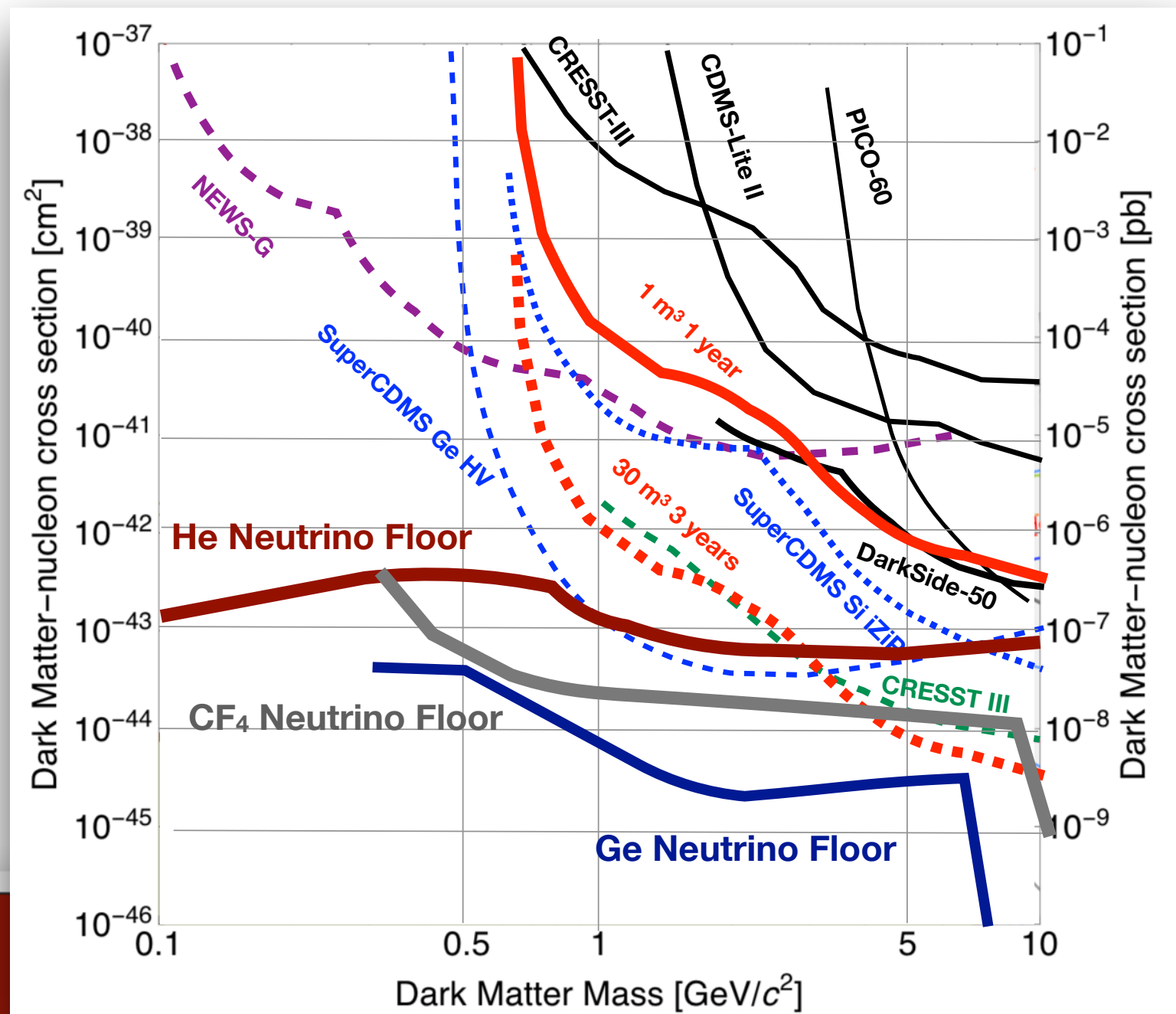
$$\epsilon = \frac{4\rho}{(\rho + 1)^2} \quad \text{where} \quad \rho = \frac{m_N}{m_{WIMP}}$$

Therefore, a WIMP with a 1 GeV mass, would transfer different energy to different target nuclei:

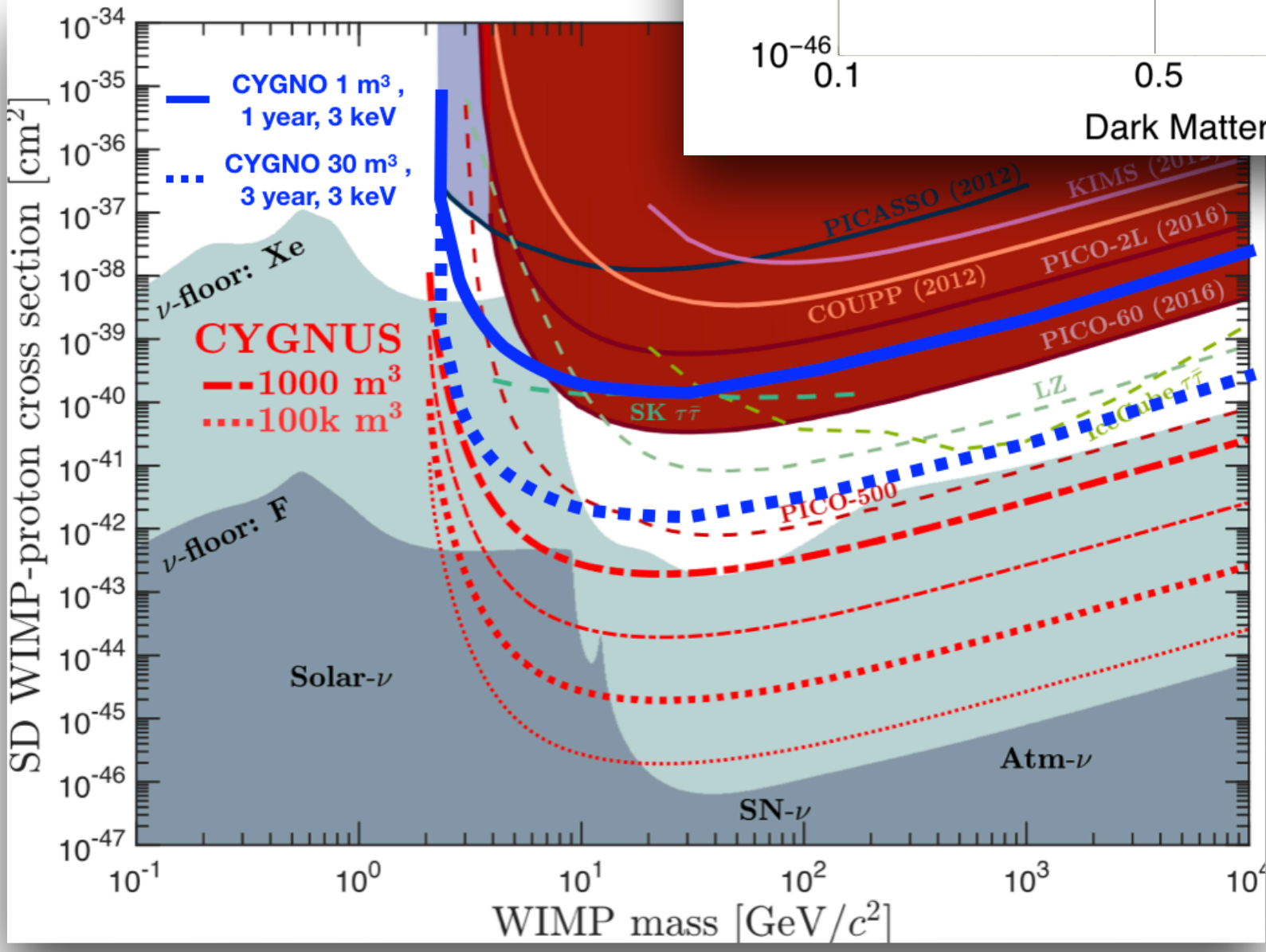
Target	Max Energy Transf. (keV)
He	1,2
Ar	0,2
Xe	0,06



WIMP energy sensitivity



- **Energy threshold 1 keV_{NR}**
- **Target mass 30-100 kg (F, He)**
- **Zero neutron background**
 - no steel (vacuum) vessel (acrylic?)
 - ceramics; almost no internal electronics
- **x, y, z fiducialisation and radon rejection**
 - either negative ion drift or other technique
 - material selection and scrubbing is not enough
- **Gamma discrimination below 10 keV_{NR}**
- **Directional sensitivity**



estimate with zero background

Directionality, head tail & PID

Time Projection Chambers provide:

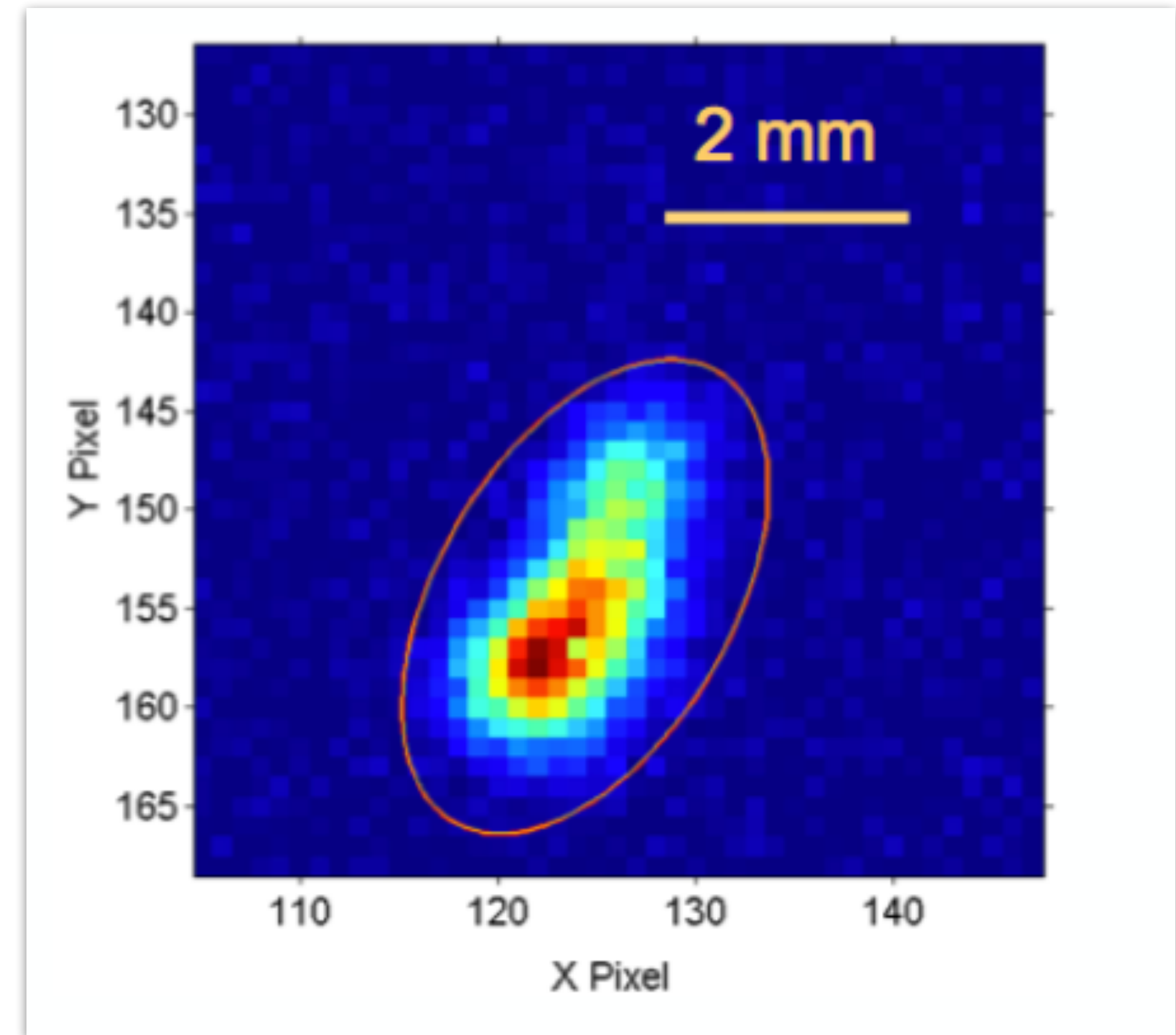
- 3D** tracking (position and direction);
- total **released energy** measurement;
- dE/dx** profile (pid, head-tail);
- reduced readout channel number;

gas represents an interesting target:

- nuclei free path can be **long** enough to be reconstructed;
- low mass** gases allow an efficient momentum transfer from light DM;

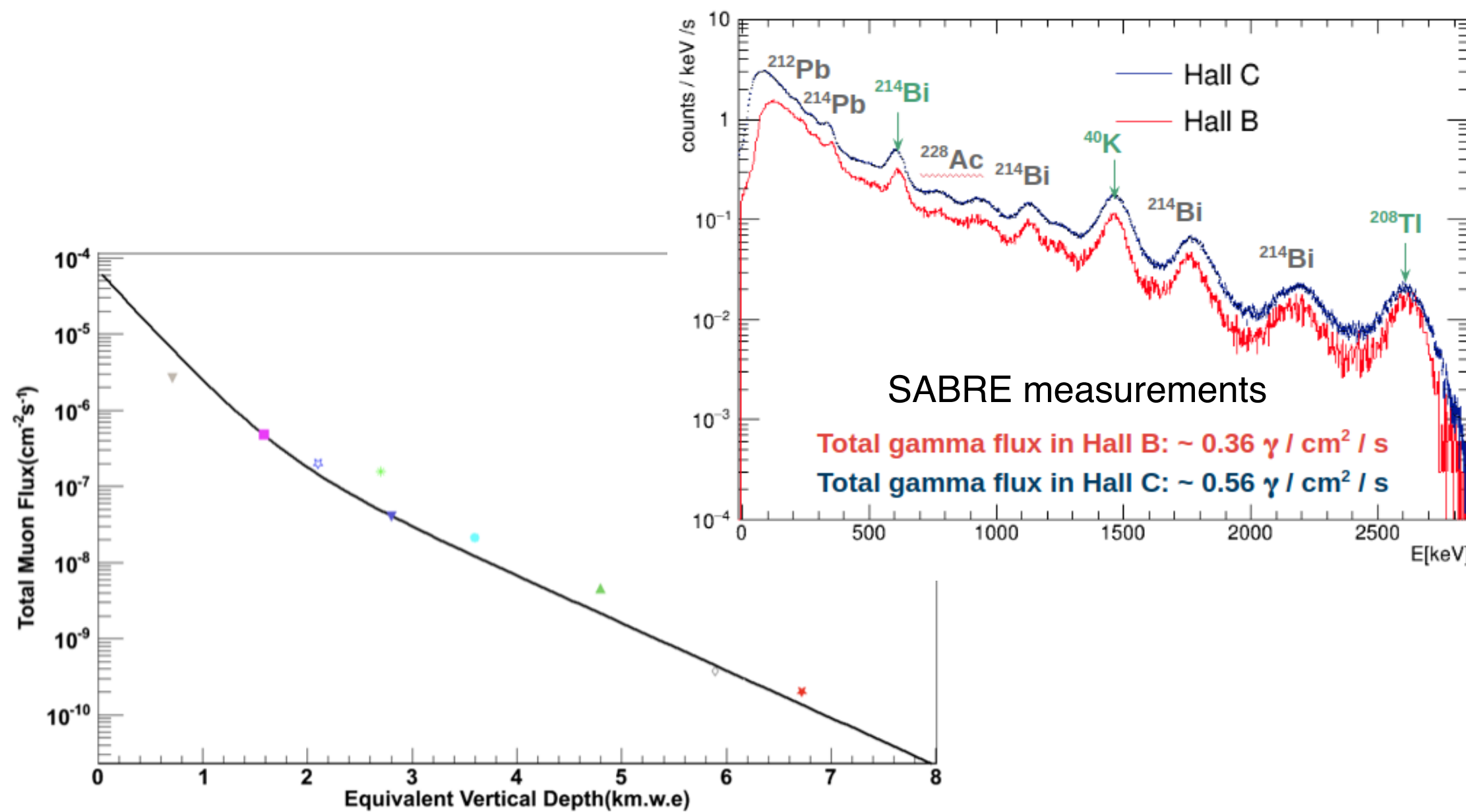
avalanche mechanism allows a **sensitivity to single primary electrons** (i.e. energy release of 30-40 eV);

Nuclear recoil in gas

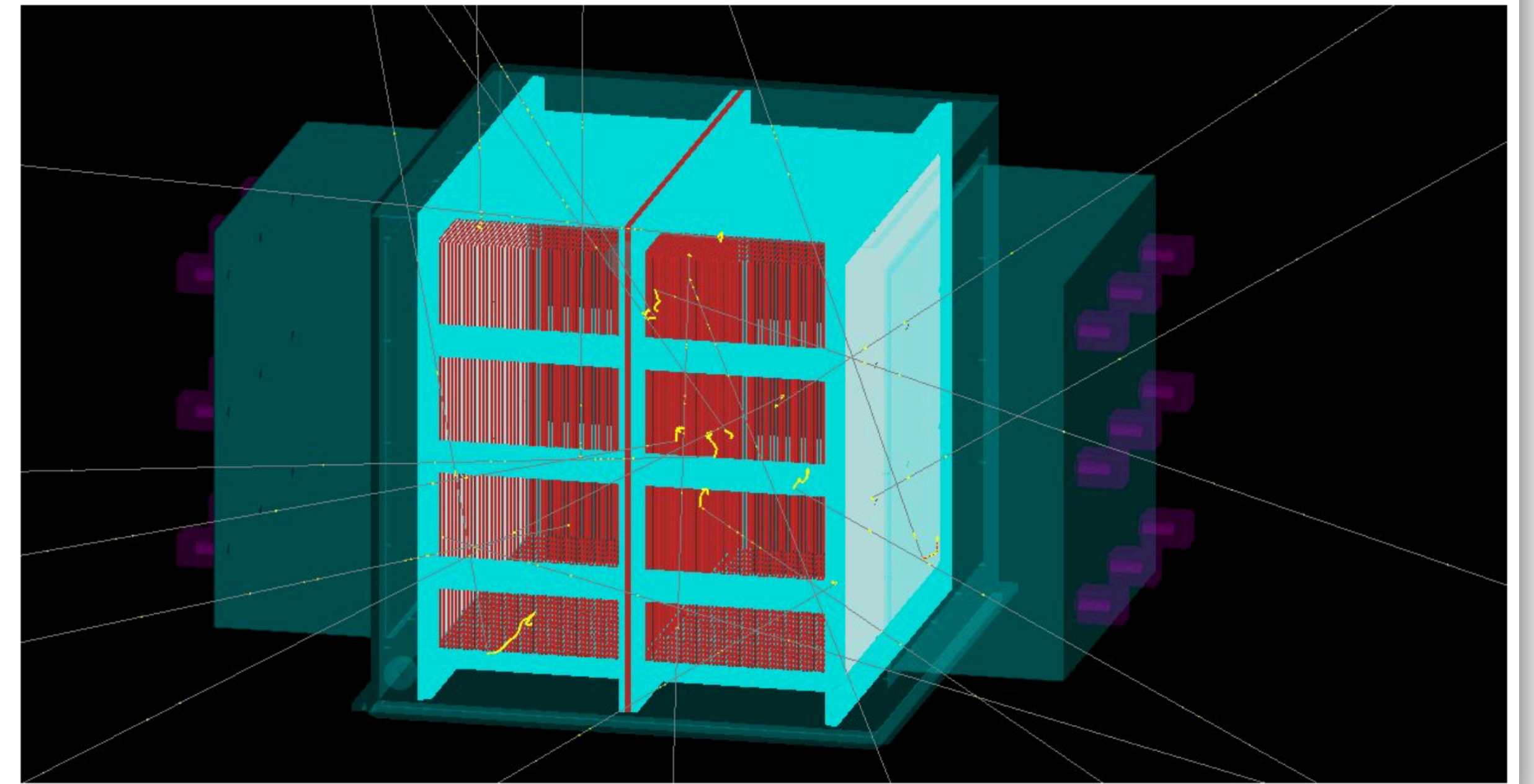


Background & Full Sim

- **internal background:**
gas radioactivity and materials
➔ materials choice and gas purification
- **external background:**
gamma, neutrons, and comics
➔ shielding (water+Cu+Pb?+...)



Example: ^{14}C decays in the gas



G. Cavoto, F. Bellini, A. Messina, G. D'Imperio

G. Mazzitelli for CYGNO/INITIUM Collaboration

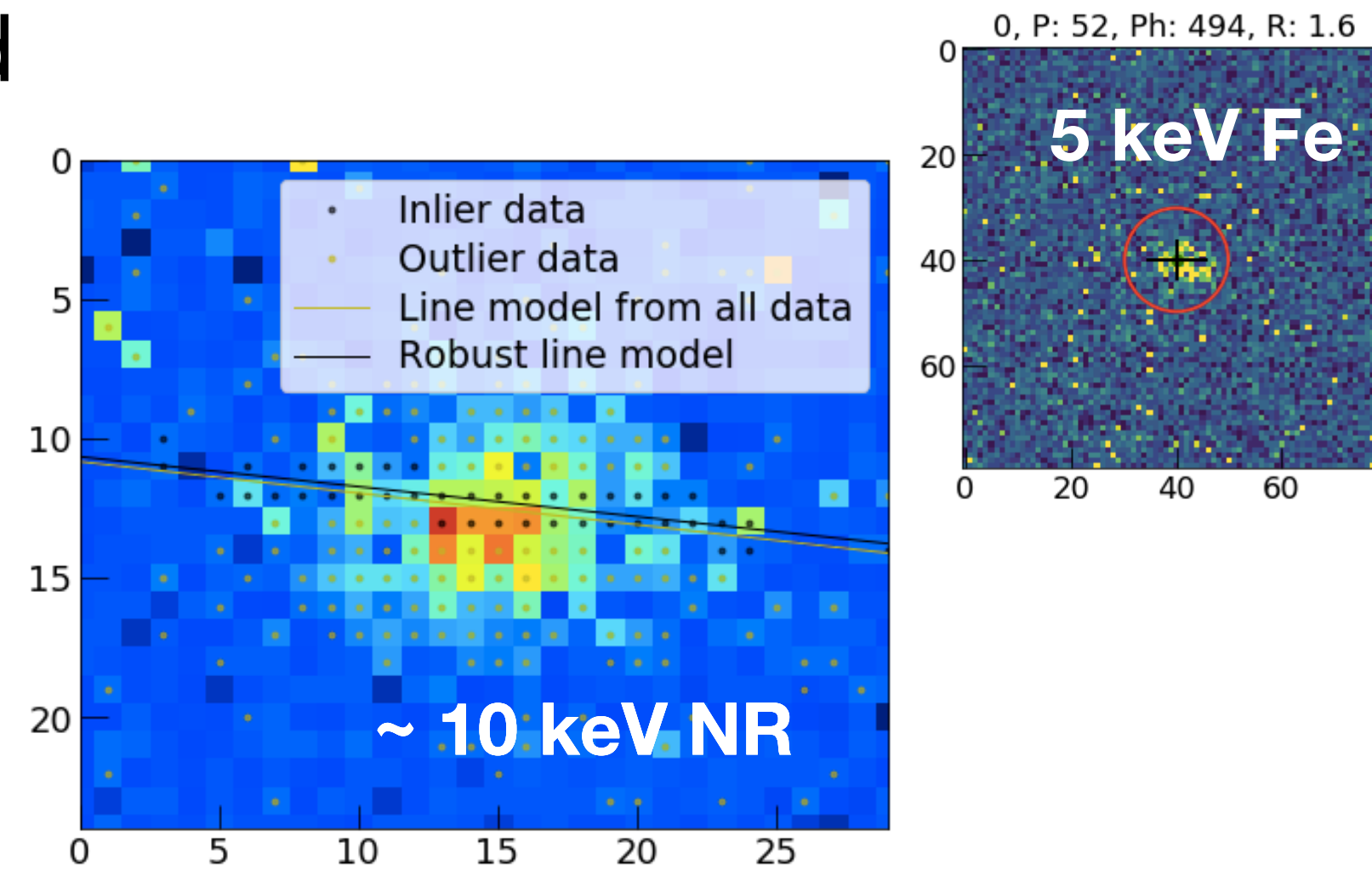
Data analysis

The effective energy threshold is determinate by the ability to identify candidate over background

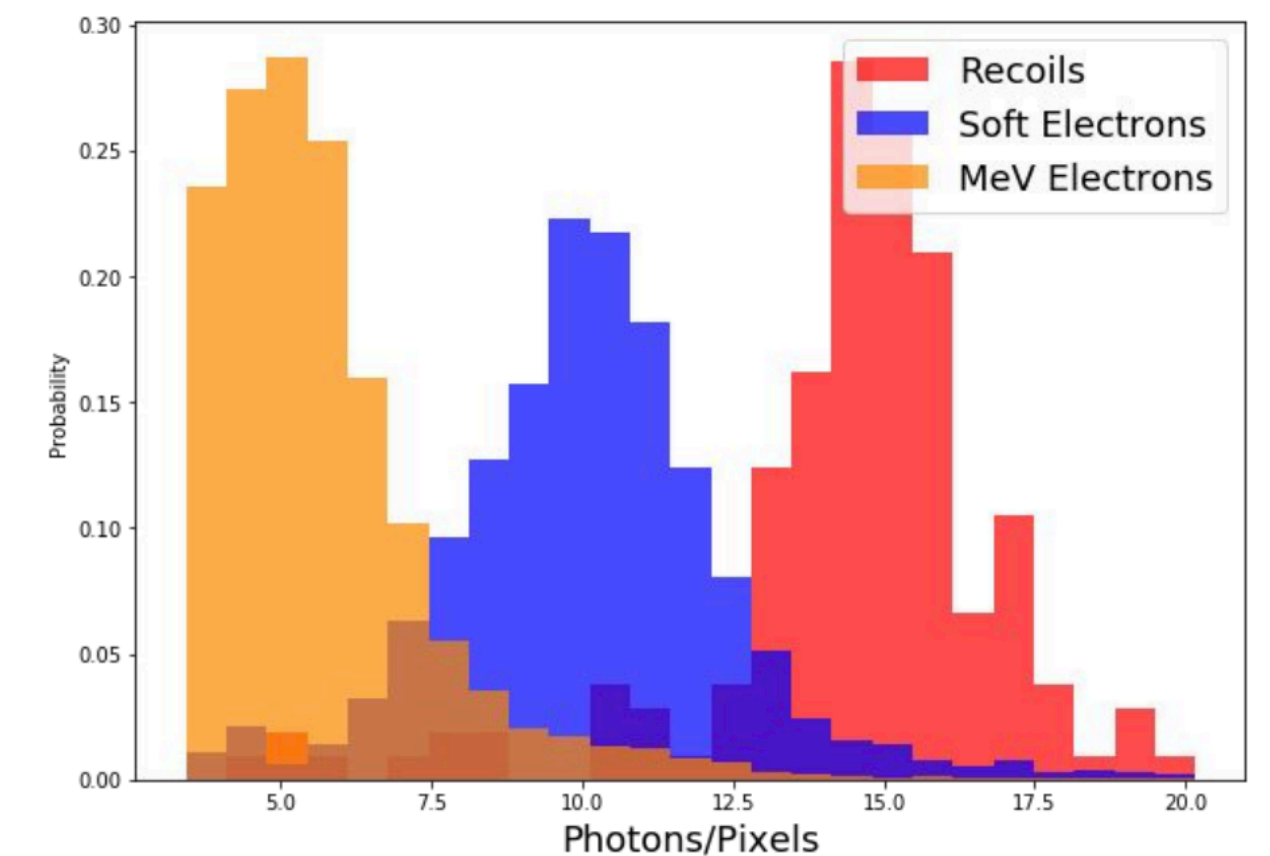
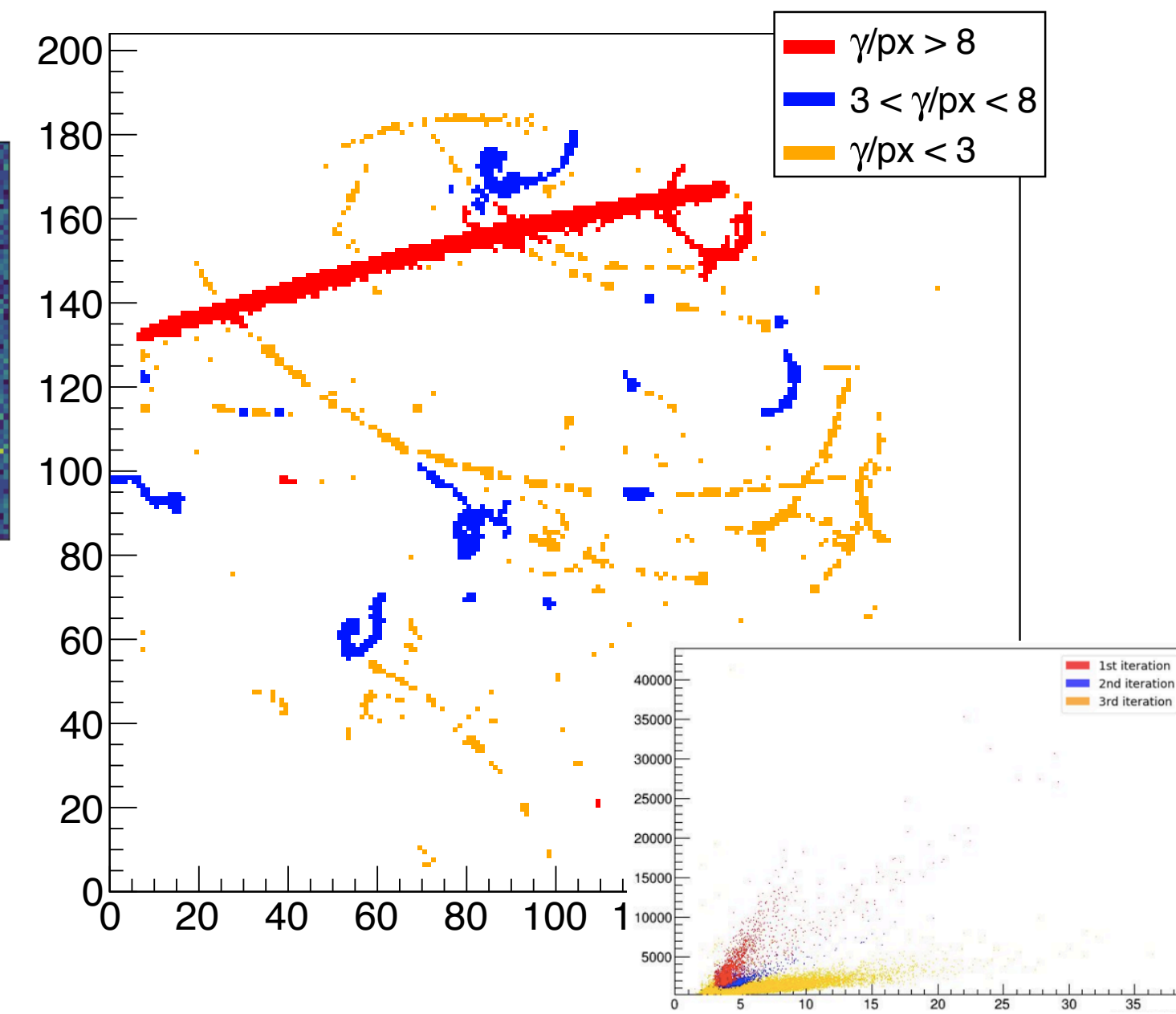
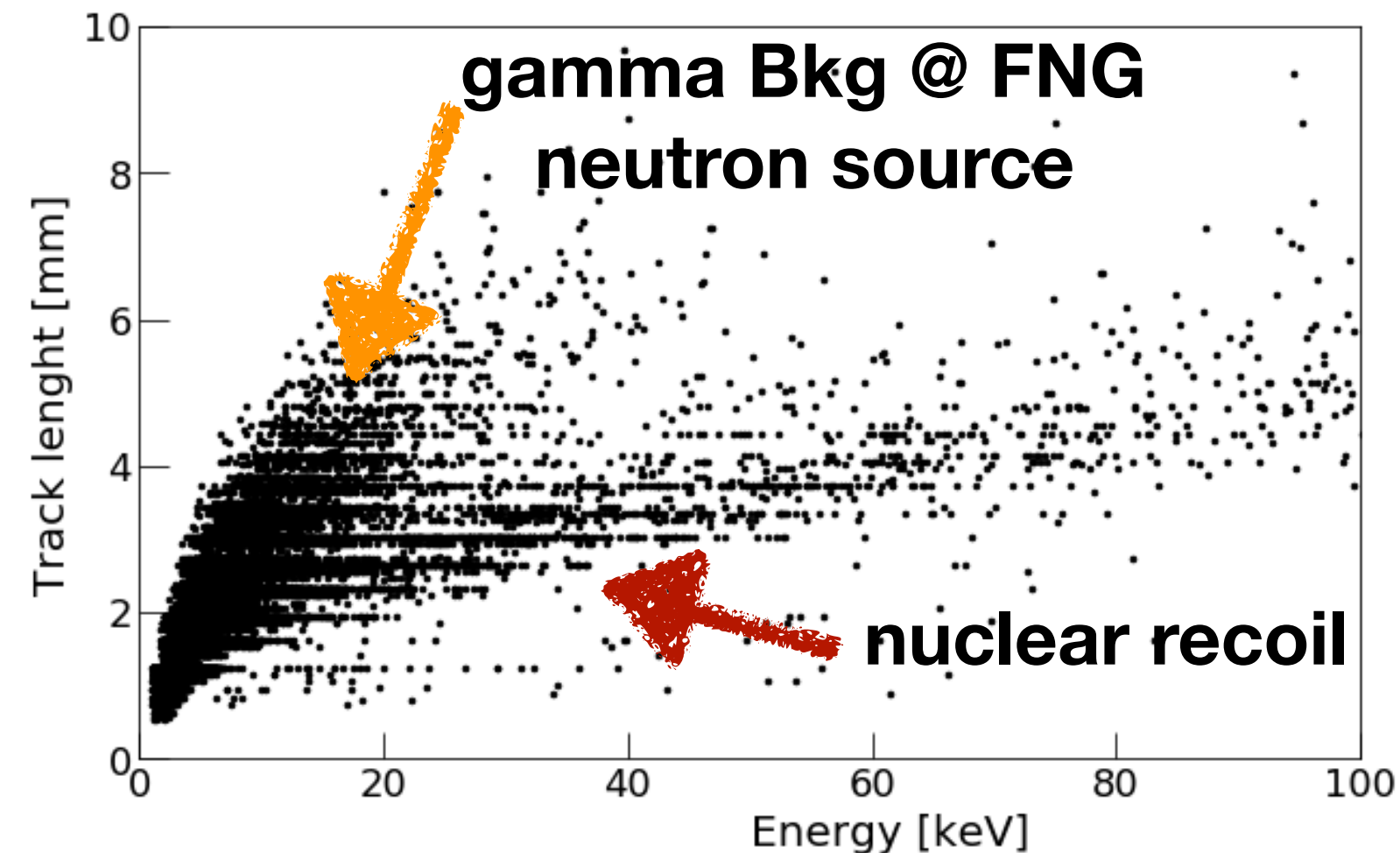
- Particle Identification (PID)
- directionality & head tail
- topology (sparsity, curly, etc)

Moreover, a throughput of \sim GB/s (strongly dependent on underground background condition) is foreseen and a first level real time analysis is need in order downsample data

- front end farm, GPU/FPGI
- machine learning tools

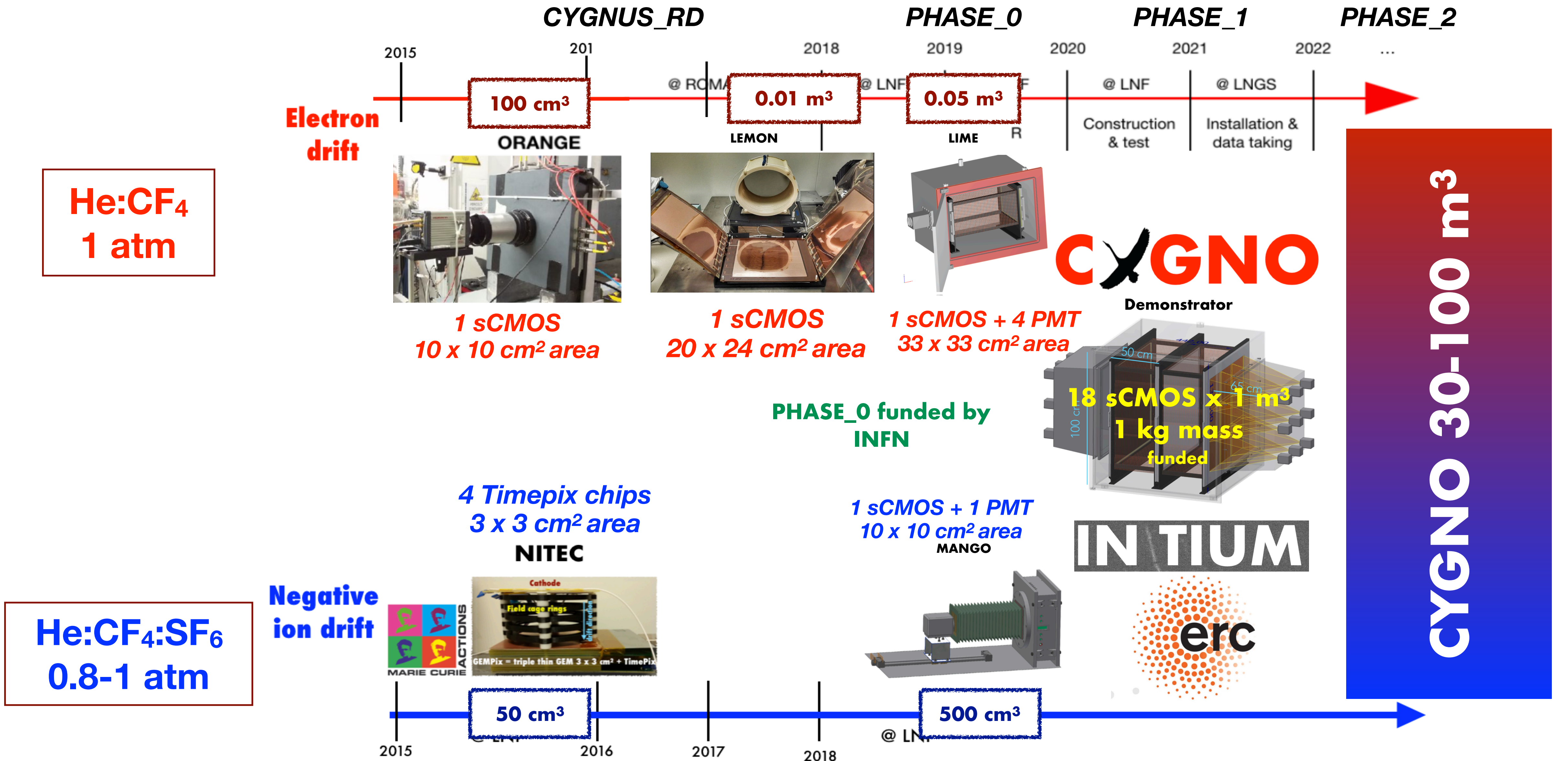


clustering_run831_Nsig_1_Mcut_350_Pcut_0_scale_4_close_2_nccs.txt



i2DBSCAN classification

Roadmap R&D and project Phase



Phase0 - TDR status and working plan

Technical Design Report

Esperimento XXX

In questo documento sono descritte le linee guida principali che necessariamente devono essere presenti nella redazione di un Technical Design Report (TDR).

Questo documento è derivato dal template redatto e approvato dal Gruppo di Lavoro "Project Management" dell'Istituto Nazionale di Fisica Nucleare (INFN) ed è declinato tenendo conto delle peculiarità dei Laboratori Nazionali del Gran Sasso (LNGS).

Autore	Verificato da	Approvato da
--------	---------------	--------------

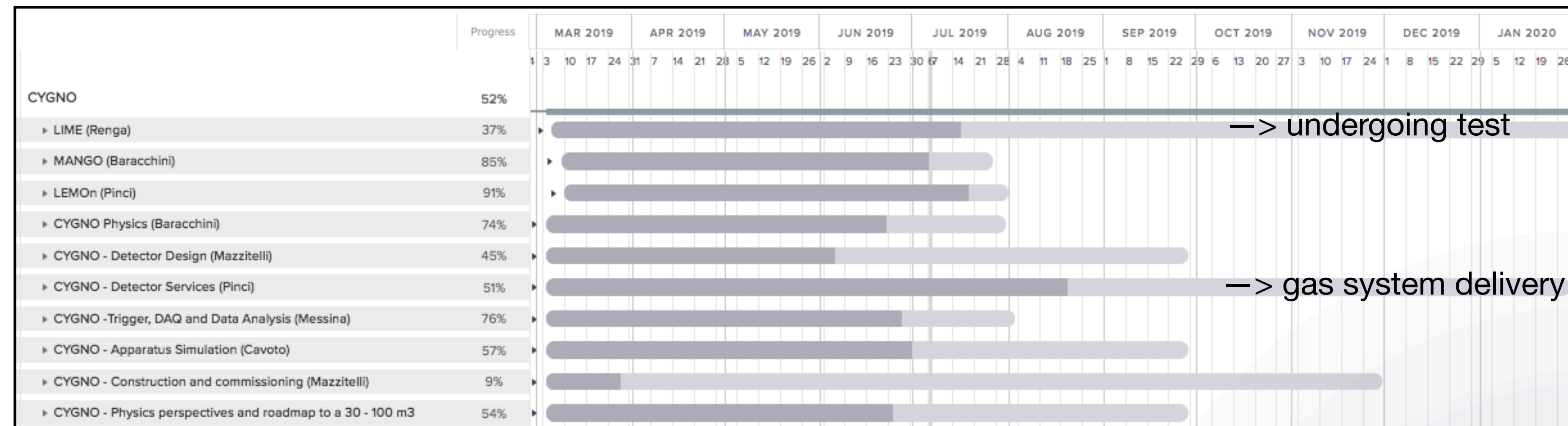
Summary

Lista degli autori.....	4
Executive Summary	4
1. Motivazioni scientifiche.....	4
2. Risultati dall'R&D	4
3. Organigramma Esperimento	4
4. Specifiche e parametri - Overview	4
5. Descrizione Tecnica	5
6. Validazione	6
7. Installazione e Commissioning	6
8. Aspetti HSE	6
9. Radioprotezione	7
10. Gestione del Progetto	7
Fasi del Progetto.....	7
Work Breakdown Structure	7
Cronoprogramma	8
Budget	8
Risks	9
Organizational Breakdown Structure	9

- Spokesperson: INITIUM, E. Baracchini (GSSI), CYGNO, D. Pinci (INFN-ROMA1)
- Technical Coordinator: G. Mazzitelli (INFN-LNF)
- Engineering Coordinator: S. Tomassini (INFN-LNF)
- Services Coordinator: D. Pinci (INFN-ROMA1)
- Read Out Coordinator: L. Benussi (INFN-LNF)
- Physics Coordinator: E. Baracchini (INFN-ROMA1)
- Simulation Coordinator: G. Cavoto (INFN-ROMA1)
- DAQ & Analysis Coordinator: A. Messina (INFN-ROMA1)
- Local Responsible: to be define
- Site Manager: to be define
- Funds Responsible: E. Baracchini (GSSI), F. Renga (INFN-ROMA1) , D. Pinci (INFN-ROMA1), G. Mazzitelli (INFN-LNF)
- GLIMO-S&E: to be define



TDR and R&D 2019 GANTT



S. Gazzana (management), M. Tobia (sicurezza),
R. Adinolfi (ambiente), DT LNGS

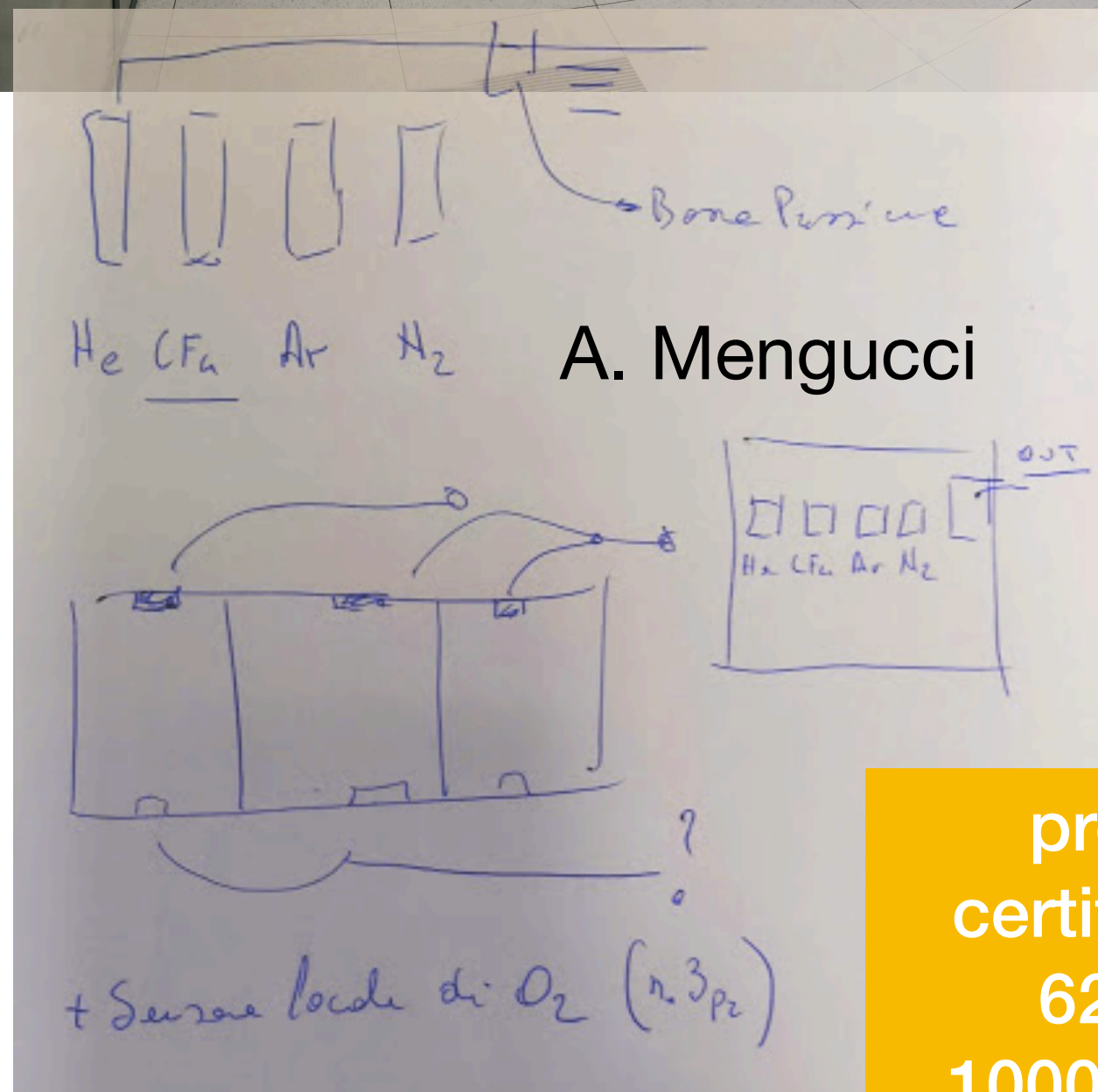
Phase0 LNF infrastructure



many thanks to MU2E group and in particular to E. Paoletti and F. Marino

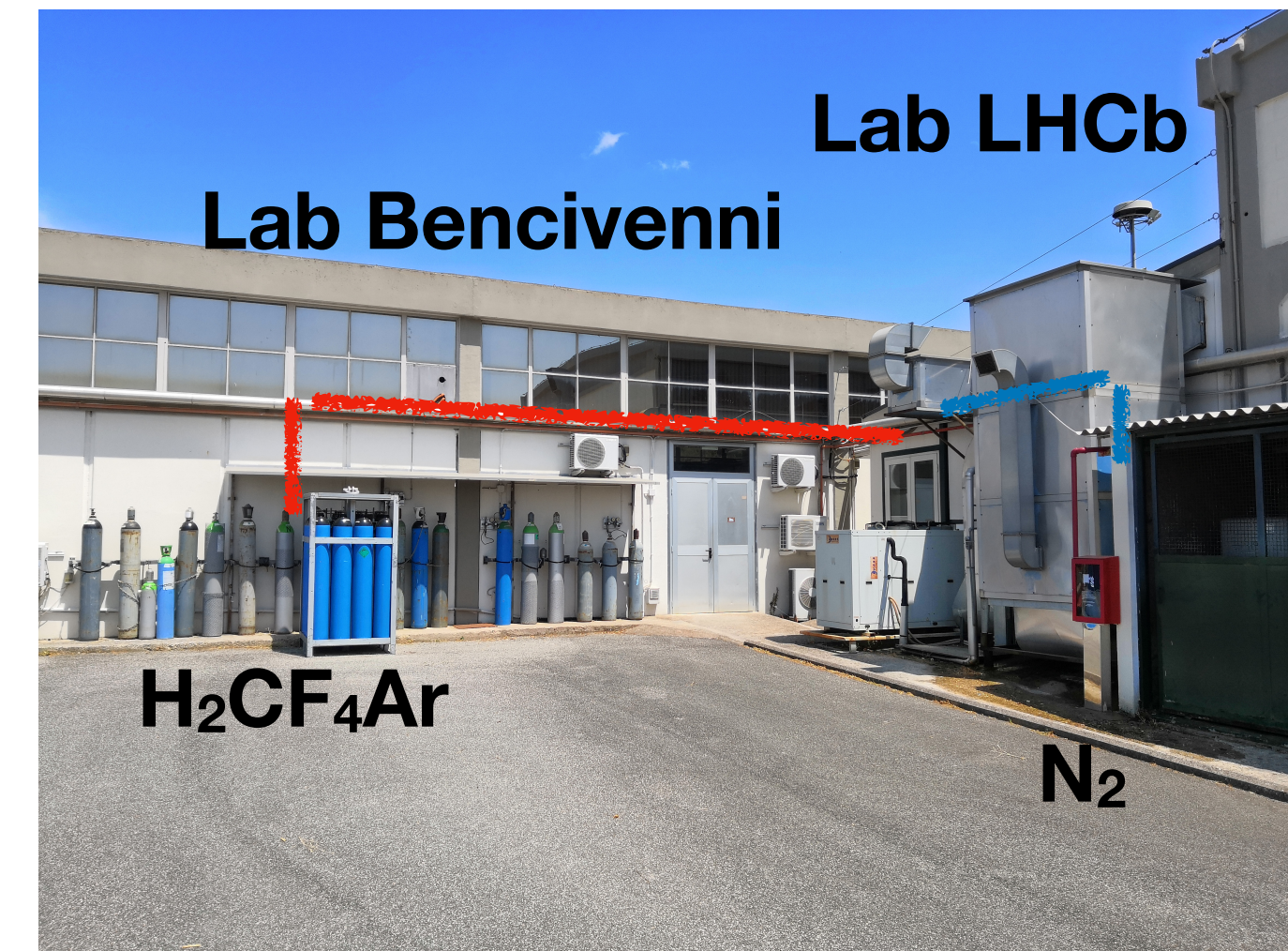


Refuse, Reduce, Reuse, Repurpose, Recycle!

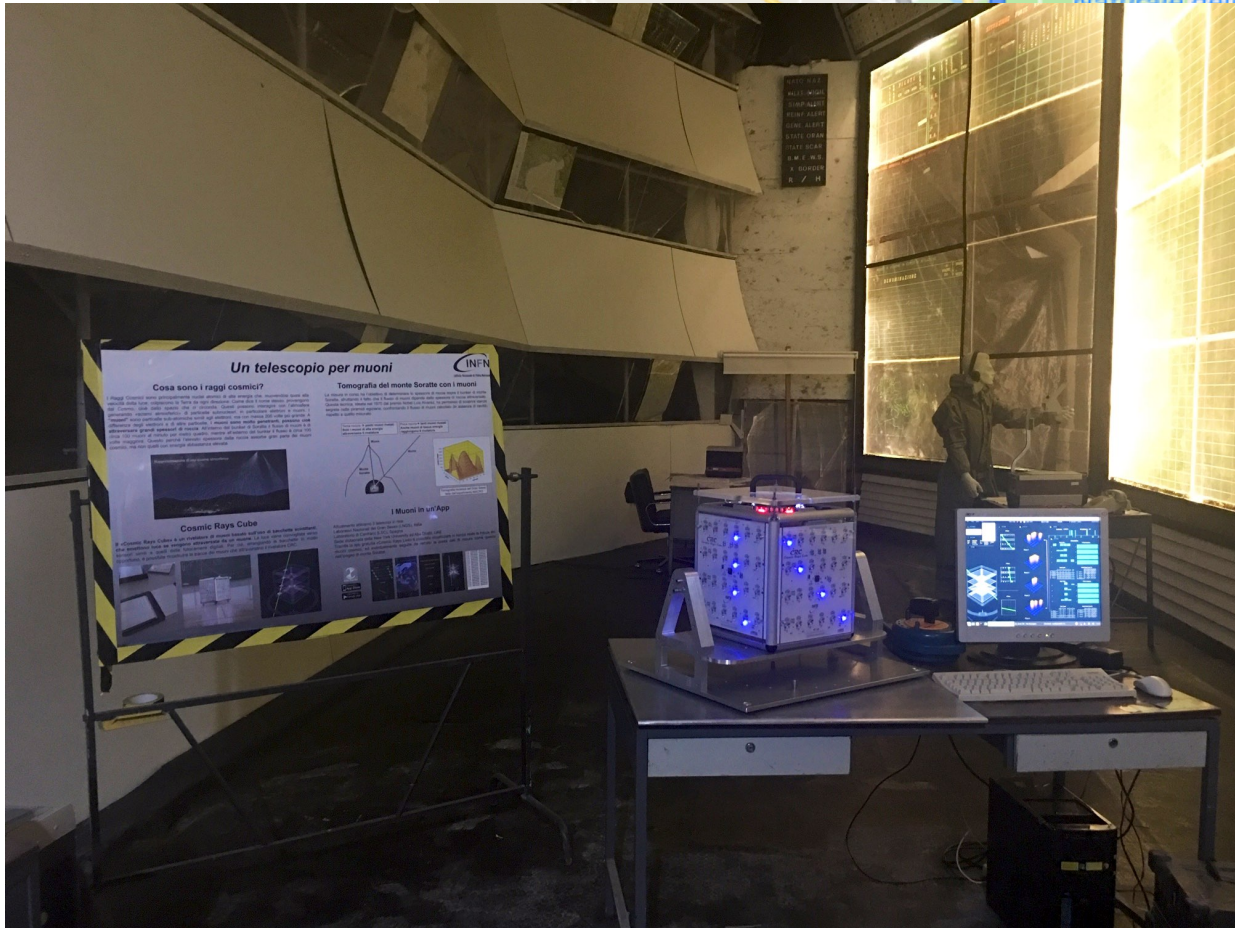
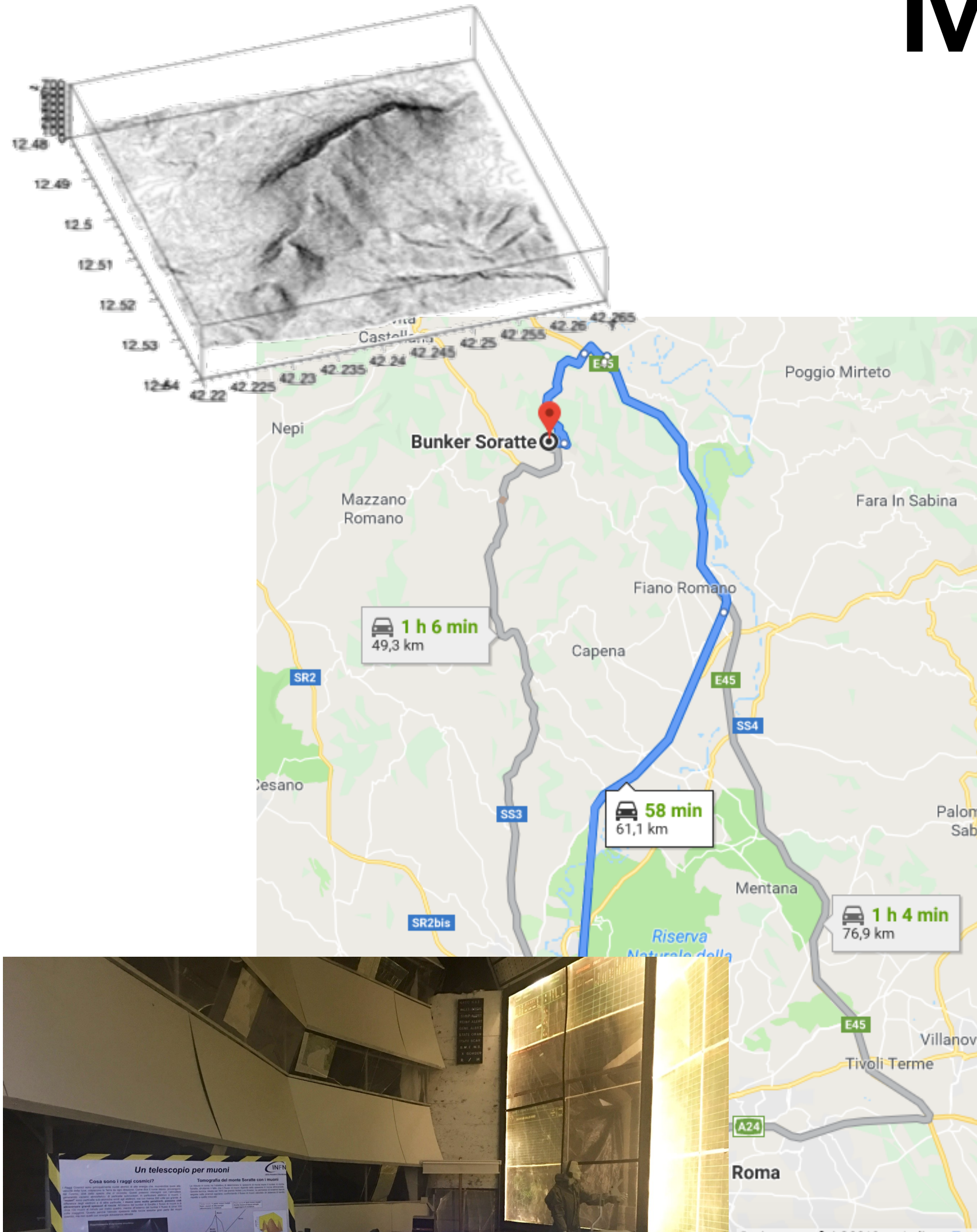


basamento

preventivo ditta,
certificazioni incluse,
6200 euro + IVA
1000 euro monitor O₂



Monte Soratte site



- Under M. Soratte, a dismissed bunker partially used as a museum
- Some free galleries could be used as a site for tests under **reduced radioactivity conditions**:
- 200 - 400 m of rock (limestone) in vertical direction, few 10 m in horizontal direction
- cosmic ray measurements on going (LNGS + C. Gustavino), ~ 1/100 w.r.t. outside

- Identified as a possible site for the PTOLEMY experiment
- There is an interest by the CYGNO/INITIUM group for tests of prototypes
- What about building **a facility for tests under reduced environmental radiation** (cosmics + natural radioactivity)?
- Possible short-term plan:
 - site characterisation (cosmics, gamma, neutrons, radon,...) in collaboration with LNGS and **LNF**
 - evaluation of safety issues
 - evaluation of potential interest of other groups (**multidisciplinary** and **interdisciplinary**)
 - evaluation of possibile **public engagement** impact
- Initial costs could be borne by the PTOLEMY & CYGNO/INITIUM group, then?

Phase0 - The GWP (CF₄, SF₆) gases issue

Tests of **eco-friendly** gas mixtures in GEM based detectors with optical readout (D. Piccolo et al)

The European Community has prohibited the production and use of gas mixtures with Global Warming Power GWP > 150

- This is valid mainly for industrial (refrigerator plants) applications
- **Scientific laboratories are excluded today**

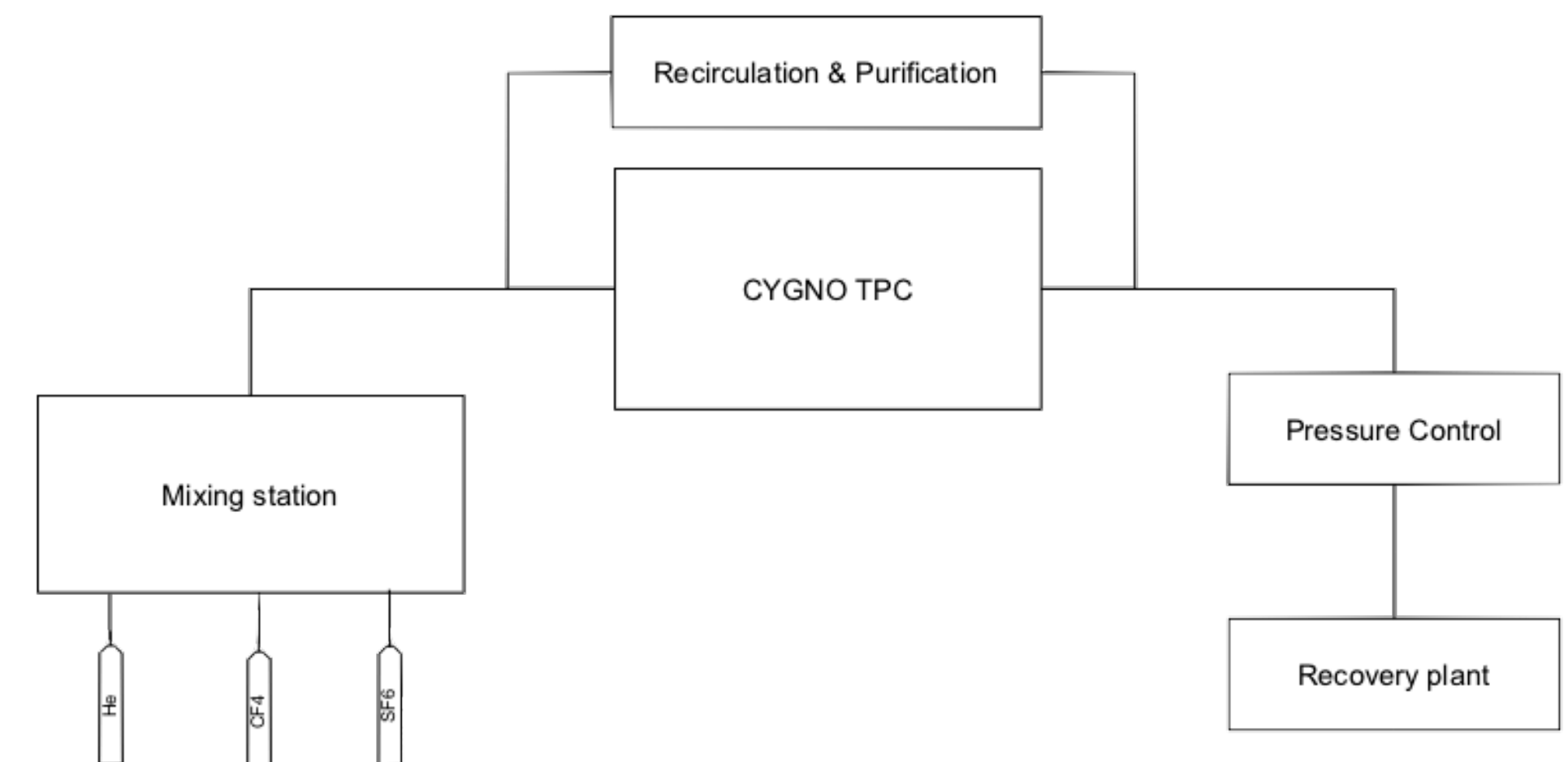
GWP(CO₂) = 1, GWP(CF₄) = 6500; GWP(SF₆)=23,900

Many GEM based applications uses or plan to use tetrafluoromethane (CF₄) in the mixture LHCb, Cygno (He-CF₄) etc.

Although scientific laboratories could still use CF₄ a recovery system is needed to not put CF₄ in the environment, moreover, prices of banned gases could became more expansive in the next years

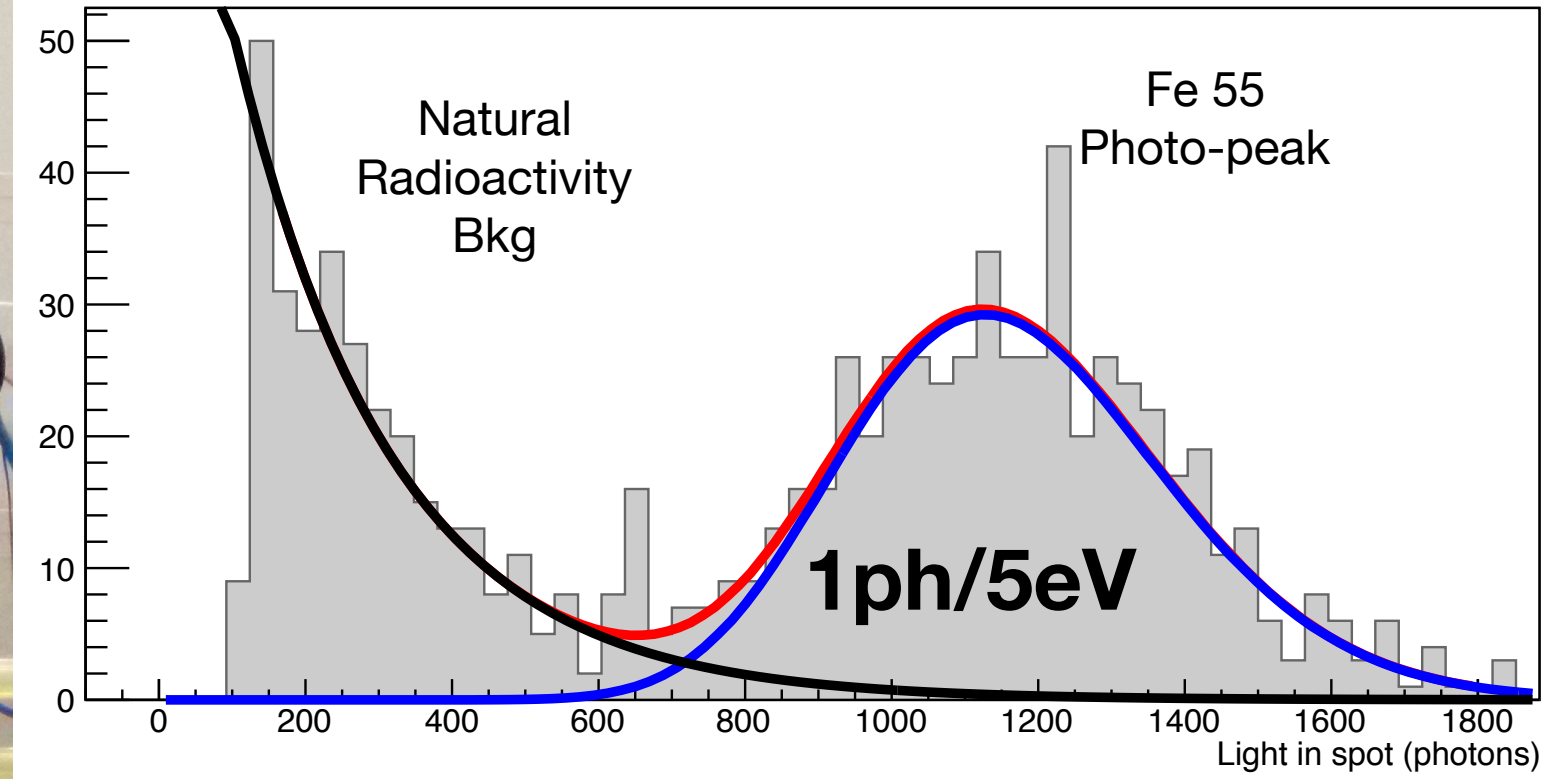
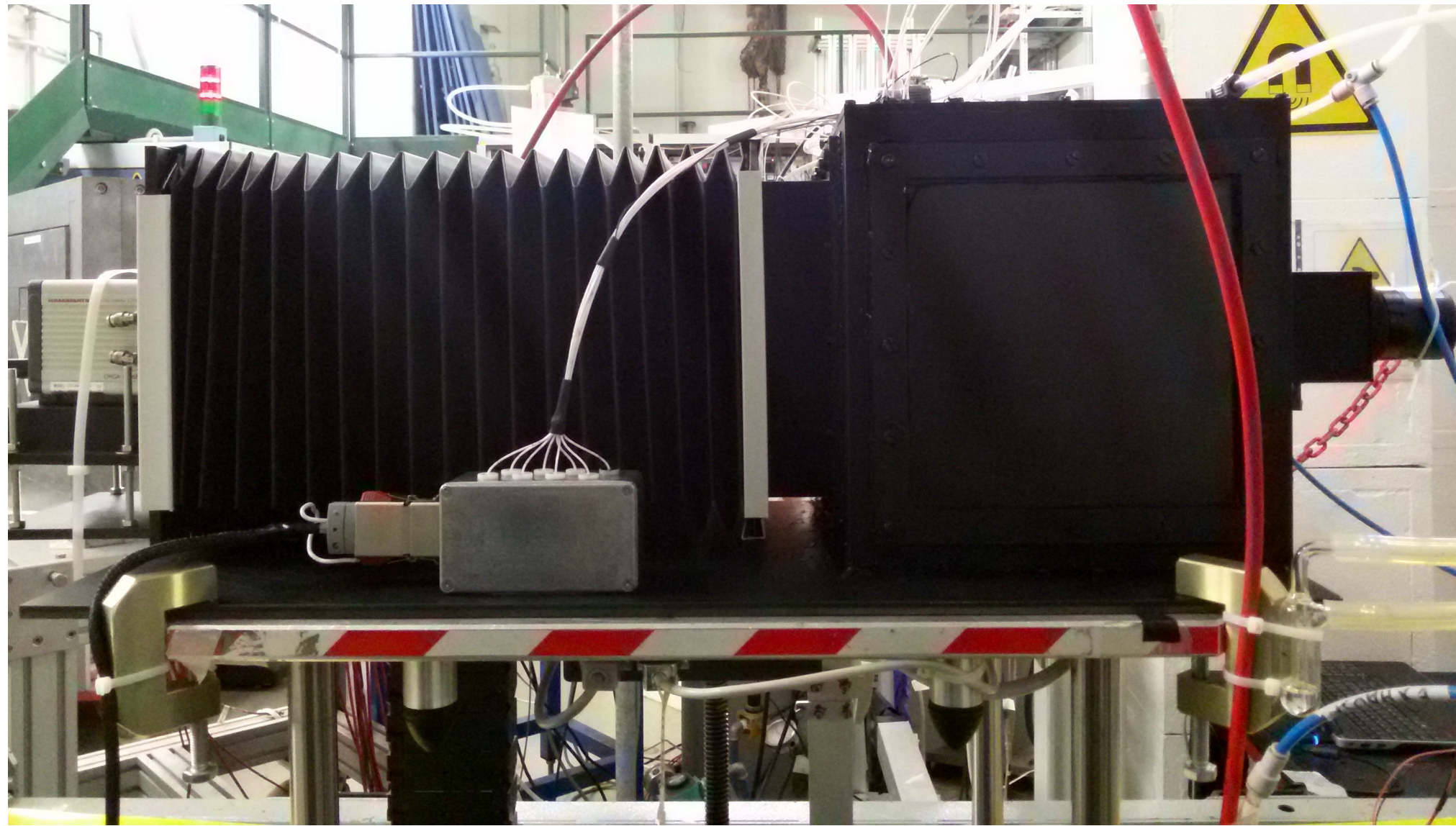
use of HFO to replace CF₄ is not straightforward.
10% of HFO reduce light emission ~10

Mix	HV GEM	<Qgem> (pC)	G (10 ⁴)
He-CF ₄ (60-40)	340	-2.1	8
He-CF ₄ (60-40)	350	-2.9	10
He-CF ₄ (70-30)	320	-2.4	8.5
He-CF ₄ (70-30)	330	-3.2	11
He-CF ₄ -HFO(70-30-10)	360	-1.4	5
He-CF ₄ -HFO(70-30-10)	370	-2.2	7
He-CF₄ (80-20)	380	~0.6	~ 1.7



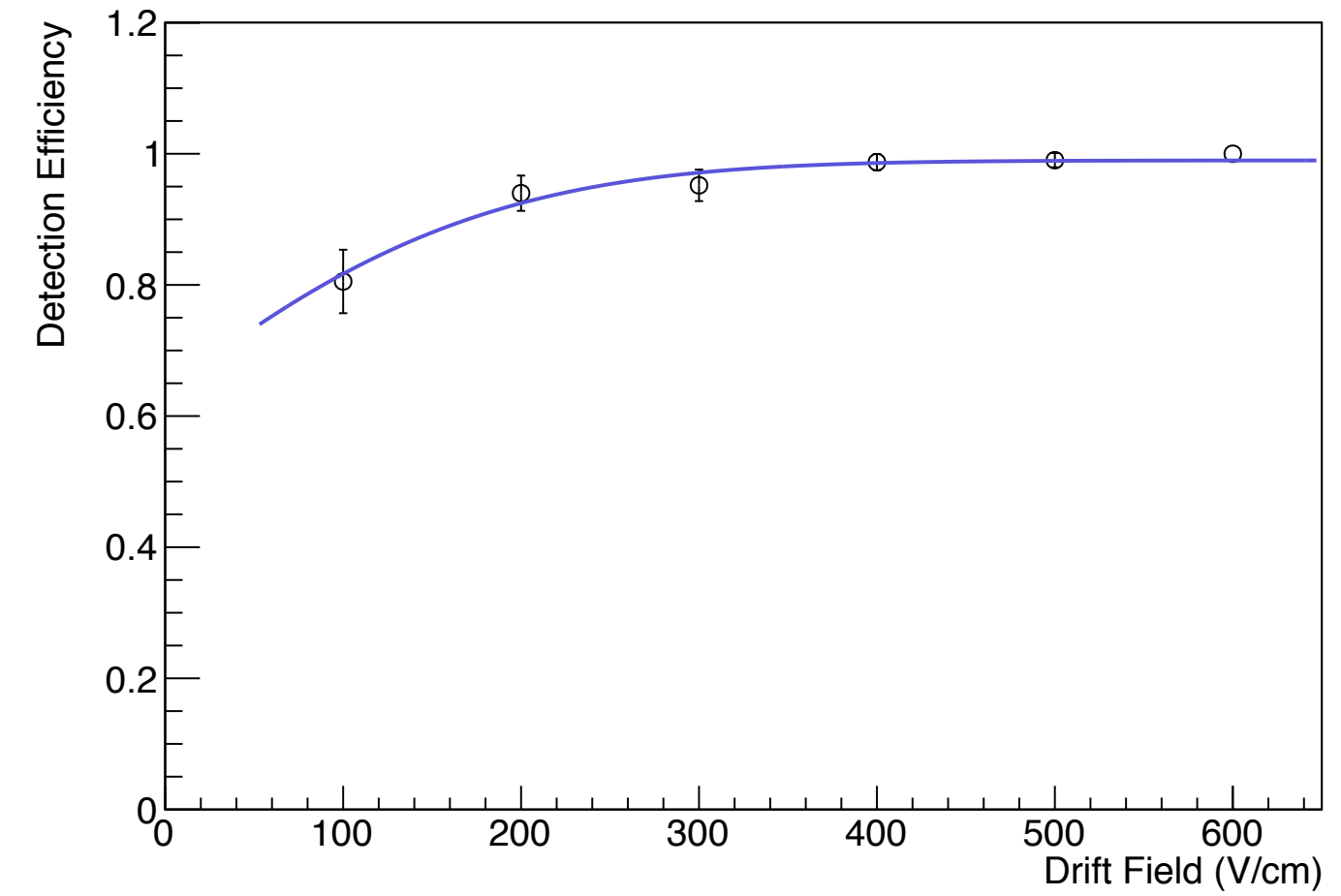
CYGNO gas system (He, CF₄, SF₆) block diagram executive design and construction under evaluation

Phase0 - Prototype



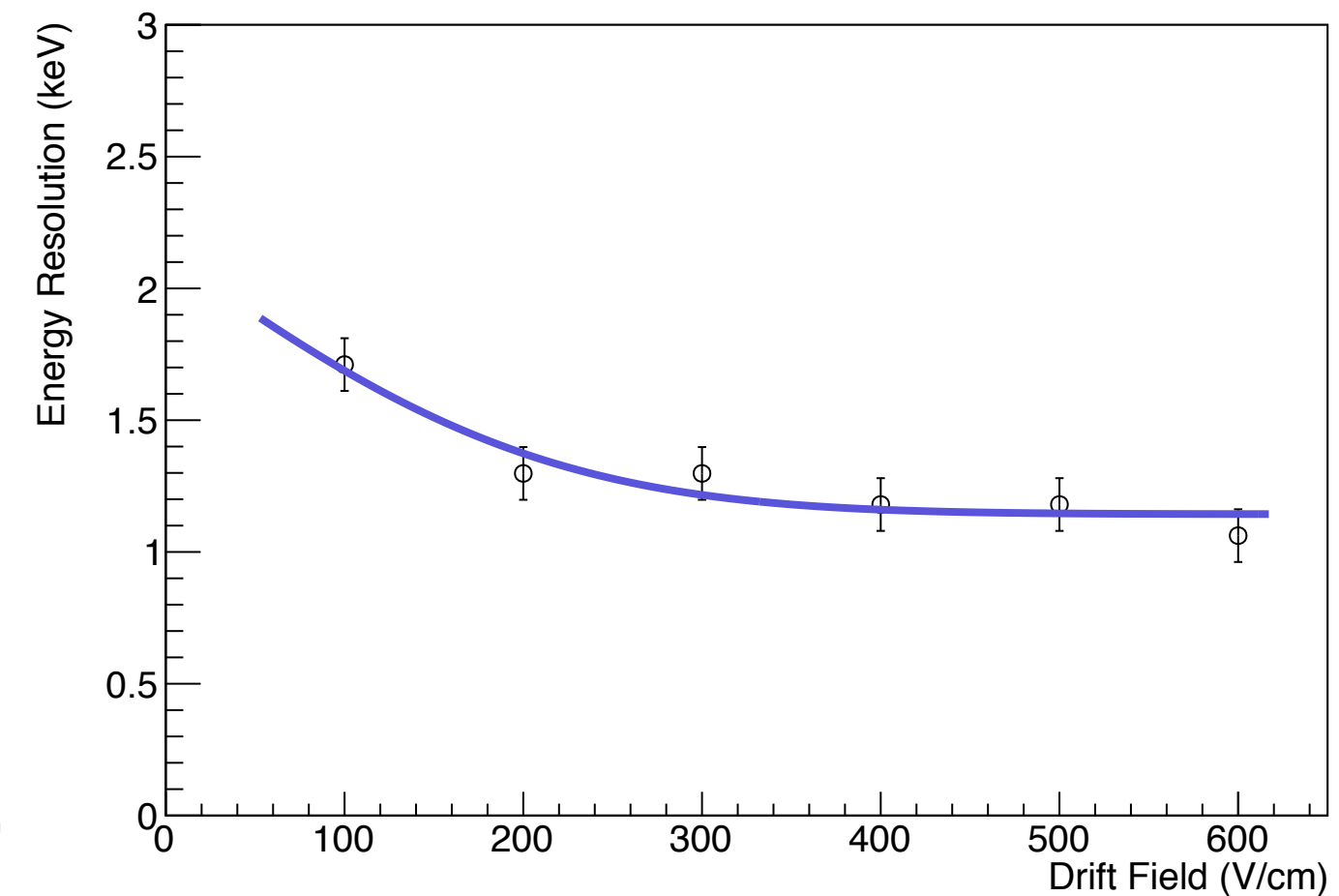
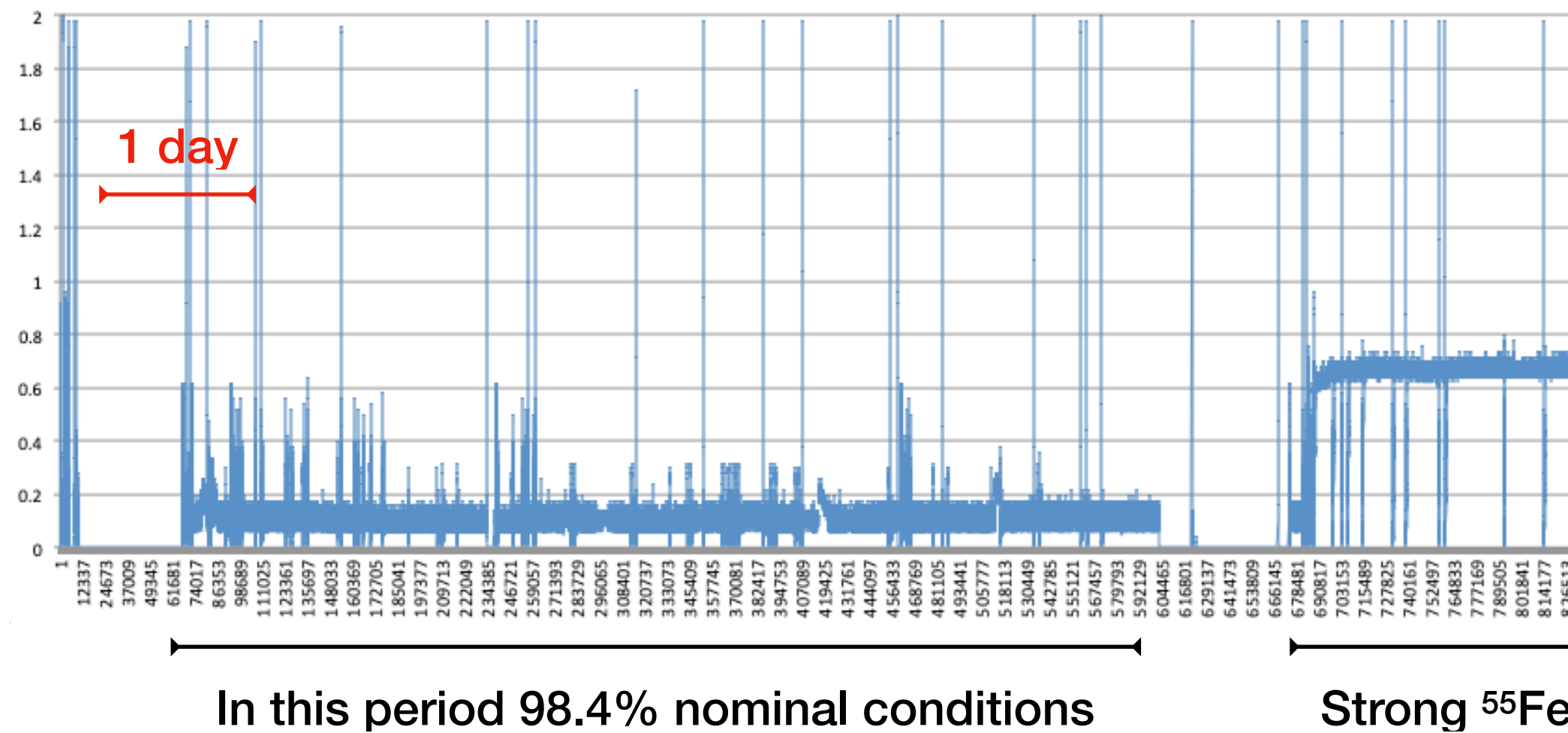
**2 keV energy threshold
(conservative) with 18% energy
resolution @ 5.9 keV for events at 20
cm drift distance**

JINST_024P_0519

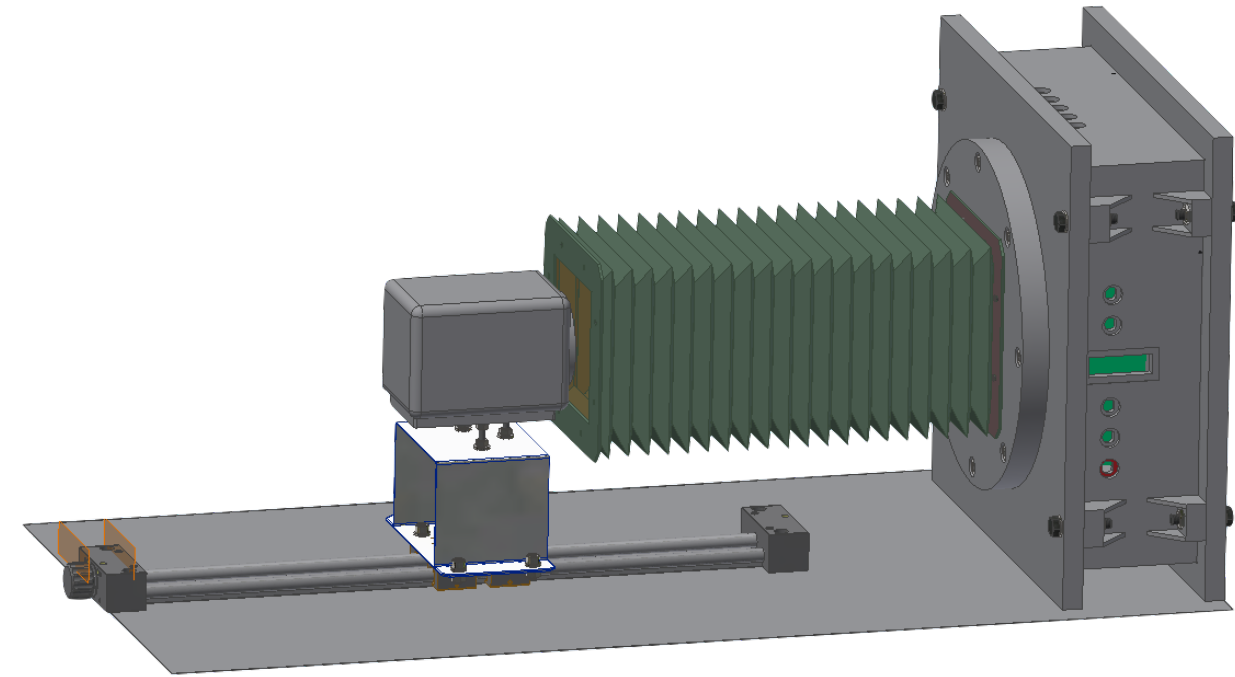
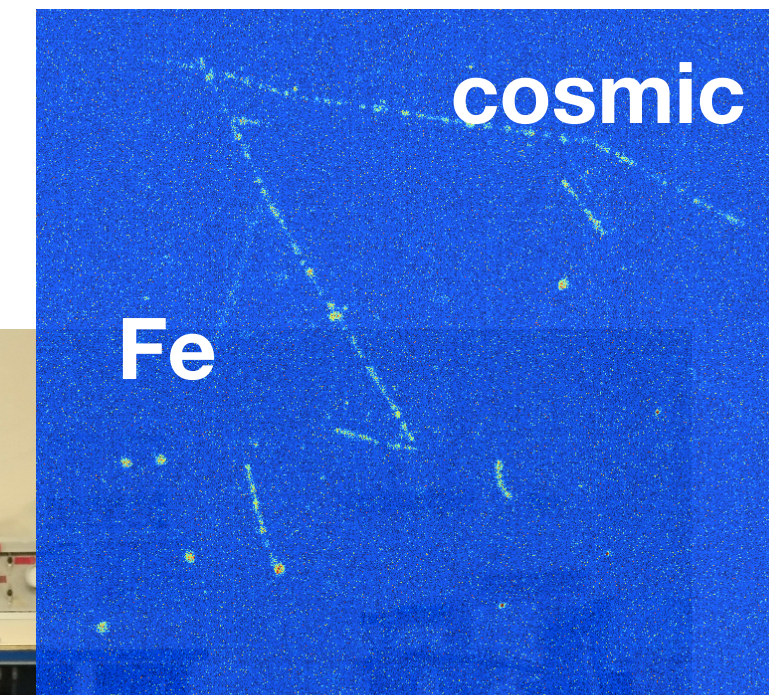


**LEMOn: Large Elliptical Module Optically
readout**

- 7 litre sensitive volume
- 25 cm drift
- 20*24 GEM
- 3D printed
- semi-transparent cathode



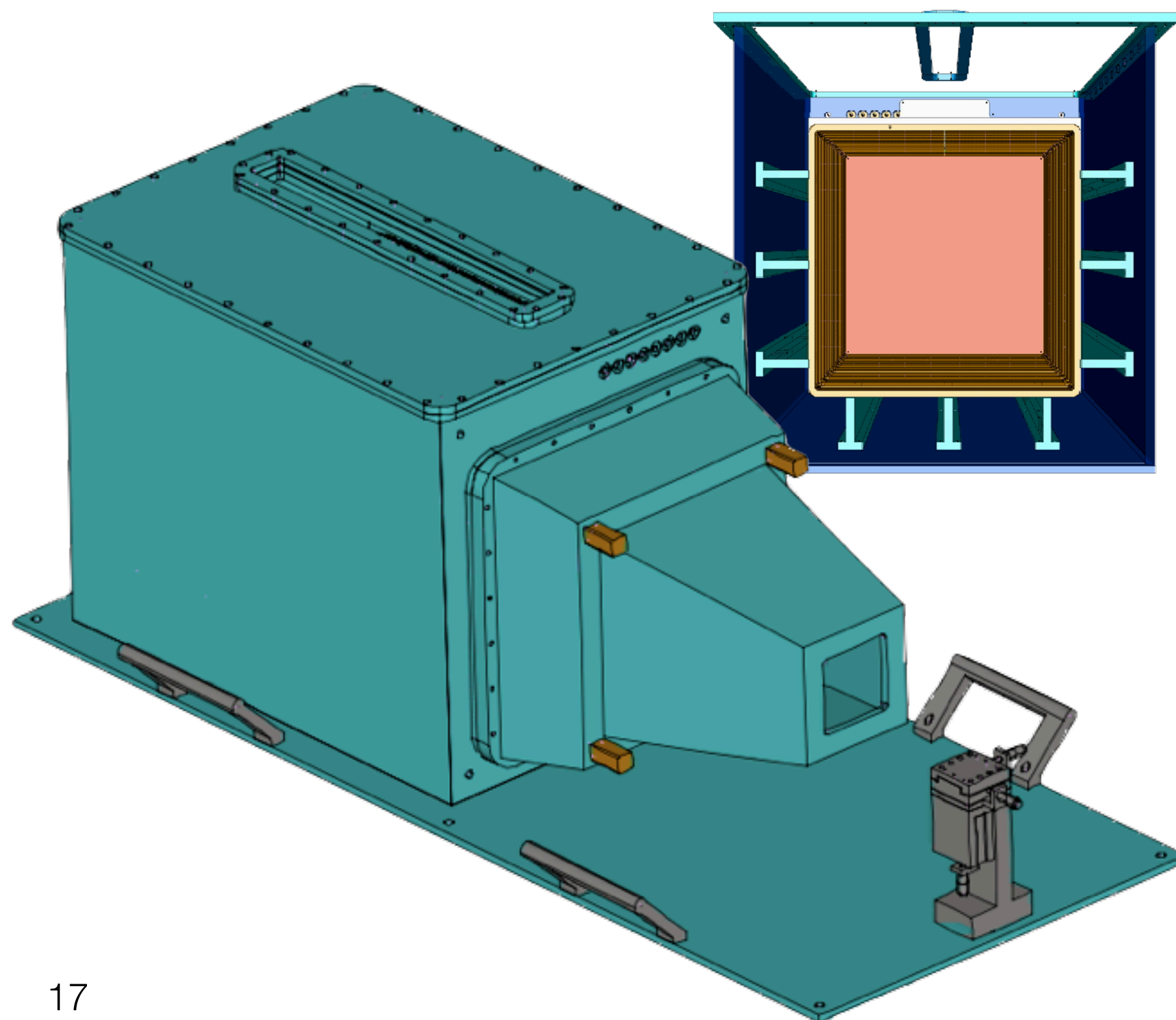
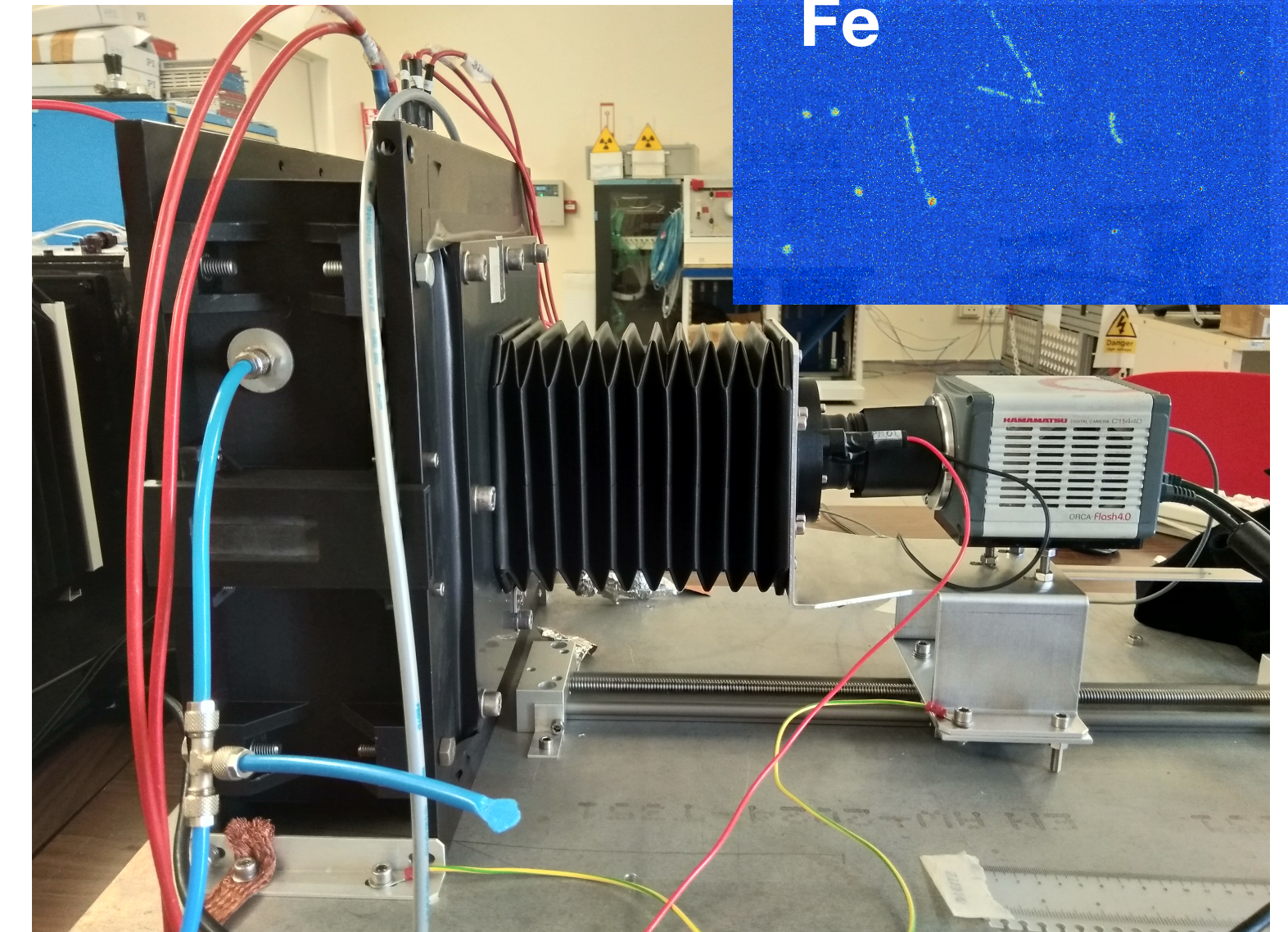
Phase0 - Prototype (con't)



C. Capoccia, A. Pelosi, F. Rosatelli, S. Tomassini

MANGO - Multipurpose Apparatus for Negative ion studies with GEM and Optical readout

- 5 cm drift gap
- THGEM test
- 4 GEM test
- Negative Ion test



LIME: Long Imaging Module

- 50 cm long drift gap
- studying materials
- performing a detailed study, minimisation and simulation of radioactive background;
- gas re-circulation and purification.
- optimisation of PMT/SiPM readout and trigger.
- HV Test

**50-liter prototype
the delivery is foreseen for half of
July!**

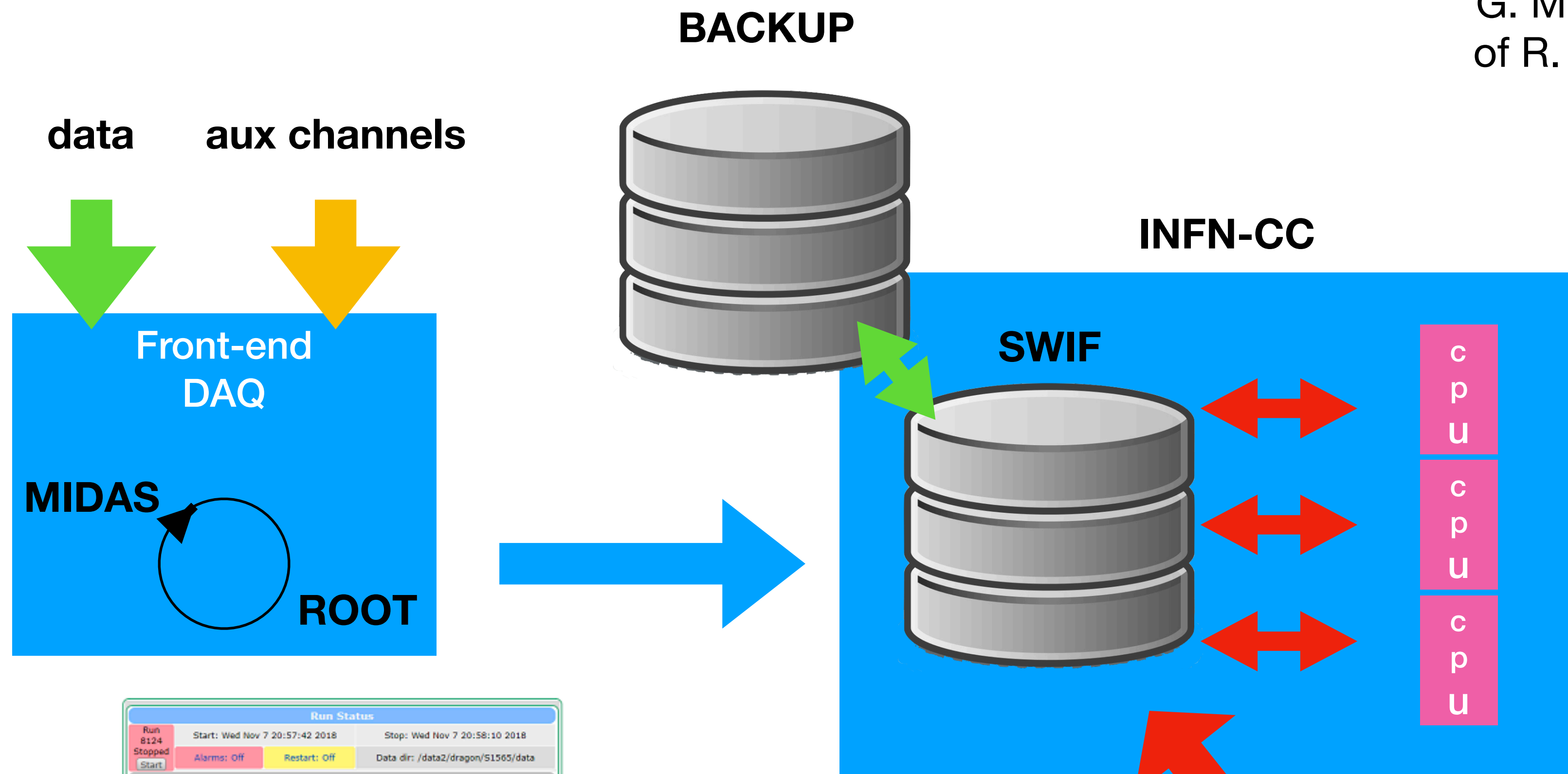
**Tests expected in fall 2019
@ BTF and the in 2020 at LNGS**

A. Orlandi, E. Paoletti, L. Passamonti, D. Pierluigi, A. Russo

G. Mazzitelli for CYGNO/INITIUM Collaboration

INFN-CC @ LNGS-LNF

G. Mazzitelli thanks to the support of R. Gargana, D. Maselli, S. Stalio



Jupyter notebook: Python3, PyROOT, ROOT

```

57 ellipse.set_transform(transf + ax.transData)
58 return ax.add_patch(ellipse)
59 def get_correlated_dataset(x, dependency, mu, scale):
60     latent = np.random.randn(n, 2)
61     dependent = latent.dot(dependency)
62     scaled = dependent * scale
63     scaled_with_offset = scaled + mu
64     # return x and y of the new, correlated dataset
65     return scaled_with_offset[:, 0], scaled_with_offset[:, 1]
66 np.random.seed(0)
67
68 PARAMETERS = {
69     "Positive correlation": np.array([0.85, 0.25],
70                                     [0.15, -0.65]),
71     "Negative correlation": np.array([0.9, -0.4],
72                                     [0.1, -0.6]),
73     "Weak correlation": np.array([1, 0],
74                                  [0, 1]),
75 }
76
77 mu = 2, 4
78 scale = 3, 5
79
80 fig, axs = plt.subplots(1, 3, figsize=(9, 3))
81 for ax, (title, dependency) in zip(axs, PARAMETERS.items()):
82     x, y = get_correlated_dataset(800, dependency, mu, scale)
83     ax.scatter(x, y, s=2)
84
85     ax.axhline(c='grey', lw=1)
86     ax.axvline(c='grey', lw=1)
87
88     confidence_ellipse(x, y, ax, edgecolor='red')
89
90 ax.scatter(mu[0], mu[1], c='red', s=3)
91 ax.set_title(title)
92
93 plt.show()
  
```

AAI users authentication

Run Status				
Run 8124	Start: Wed Nov 7 20:57:42 2018	Stop: Wed Nov 7 20:58:10 2018		
Stopped	Alarms: Off	Restart: Off	Data dir: /data2/dragon/51565/data	
1552340687 14:44:47.821 2019/03/11 [thresh,INFO] Program thresh on host smaug stopped				
Equipment				
Equipment +	Status	Events	Events[/s]	Data[MB/s]
HeadVME	Idle	0	0.0	0.000
HeadScaler	Idle	2.754M	1.0	0.000
TailVME	Idle	0	0.0	0.000
TailScaler	Idle	0	0.0	0.000
Epics	Frontend stopped	0	0.0	0.000
Wiener	OK	0	0.0	0.000
Logging Channels				
Channel	Events	MB written	Compr.	Disk Level
#0: run8119.mid	0	0.000	0.0%	77.8%
Lazy Label	Progress	File Name	# Files	Total
g-drive	0%	run6895.mid	34	0.0%
Clients				
fe_head [ixdragon01.triumf.ca]	mhttpd [smaug.triumf.ca]	mserver [smaug.triumf.ca]		
Logger [smaug.triumf.ca]	fewiener [smaug.triumf.ca]	fevScaler [smaug.triumf.ca]		
fe_tail [ixdragon02.triumf.ca]				

MIDAS run control and slow control

User HTTP queries

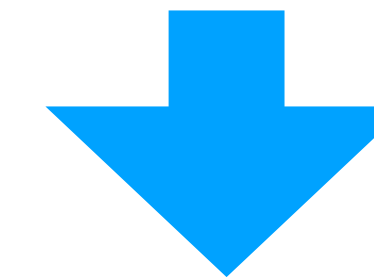


CYGNO/INITIUM @ LNF 2019-2020

CYGNO/INITIUM prevede attività che avranno il loro apice fra il 2020 e il 2021 con la costruzione del rivelatore da installare presso i LNGS nel 2022. Per il secondo semestre del 2019, oltre alla progettazione del rivelatore principale, si prevede l'assemblaggio dei prototipi LIME e MANGO e la preparazione/realizzazione delle attività di test ai LNGS/Soratte con tali prototipi

Richieste/assegnazioni ultimo CIF

- 5 mu/0.7 FTE servizio Servizio Meccanica DR per la progettazione di CYGNO/INITIUM e supporto all'assemblaggio di LIME/MANGO
- 1 mu servizio SPCM supporto lavorazioni prototipo LIME/MANGO
- 0.25 mu servizio SPCM per la stampa 3D prototipo LIME/MANGO
- 6 mu/0.8 FTE servizio Servizio Costruzione Rivelatori DR supporto al montaggio dei prototipi LIME/MANGO e supporto alle attività di preparazione/realizzazione delle misure presso i LNGS.
- 3 mu Servizio Elettronico e Automazione



Richieste LNF 2010:

Il maggiore impatto sui servizi LNF sarà nella seconda metà del 2020
 Servizio Costruzione Rivelatori DR, Servizio Meccanica DR in piena con l'attuale, probabile un maggiore coinvolgimento SEA e SPCM

Richieste CSNII 2020: supporto alle attività di R&D per CYGNO 30-100 e cofinanziamento ad INITIUM per costi di costruzione del detector non rendicontabili

Vai alla sezione: [LNF](#) [LNGS](#) [RM1](#)

SEZIONE	NOME COGNOME	TIPO	CONTRATTO	QUALIFICA	RICERCATORI	TECNOLOGI	TOT. PERS.	FTE	FTE / PERS.		
LNF	Bedogni Roberto				x			15	-10		
	Benussi Luigi				x			10			
	Bianco Stefano				x			20			
	Maccarrone Giovanni				x			30	+10		
	Mazzitelli Giovanni				x			60	+ 40 INITIUM		
	Piccolo Davide				x			20			
	Tomassini Sandro						x	10	+ 10 INITIUM		
LNF					1.55 fte	6 pers.	0.1 fte	1 pers.	7	2.2	0.314

Giovanna Saviano + 20

Michele Caponero + 20

Un articolo 15 bandito + 100

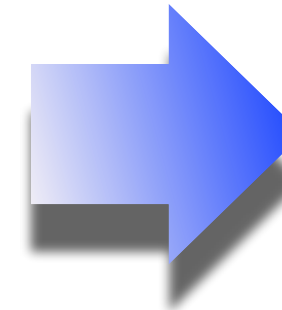
Un assegno di ricerca + 100

$$FTE = 2.45 + 1 + 0.8 + 0.7 \rightarrow 4.95 + 1...$$

Effort Phase1 2020

anagrafica 2020

anagrafica 2019



Vai alla sezione: LNF LNGS RM1

SEZIONE	NOME COGNOME	TIPO	CONTRATTO	QUALIFICA	RICERCATORI	TECNOLOGI	TOT. PERS.	FTE	FTE / PERS.		
LNF	Bedogni Roberto				x			15			
	Benussi Luigi				x			10			
	Bianco Stefano				x			20			
	Maccarrone Giovanni				x			30			
	Mazzitelli Giovanni				x			60	+ 40 INITIUM		
	Piccolo Davide				x			20			
	Tomassini Sandro					x		10	+ 10 INITIUM		
LNF					1.55 fte	6 pers.	0.1 fte	1 pers.	7	1.7	0.236
LNGS	Baracchini Elisabetta				x			20	+ 80 INITIUM		
LNGS					0.2 fte	1 pers.	0 fte	pers.	1	0.2	0.200
RM1	Cavoto Gianluca				x			20	+ 10 INITIUM		
	D'Imperio Giulia				x			50			
	Di Marco Emanuele				x			10	+ 10 INITIUM		
	Marafini Michela				x			20			
	Messina Andrea				x			30			
	Pinci Davide				x			40	+ 10 INITIUM		
	Renga Francesco				x			30	+ 10 INITIUM		
RM1					2 fte	7 pers.	0 fte	pers.	7	2.0	0.286
TOTALE					3.75 FTE	14 PERS.	0.1 FTE	1 PERS.	15	3.85	0.257

1.7 INITIUM

activity partially founded by European Research Council (ERC) grant agreement No 818744

	Appartenenza	Qualifica	FTE CYGNO	FTE INITIUM
Baracchini E.	GSSI-LNGS	Professore	0,20	0,80
Dho G.	GSSI-LNGS	PhD		1,00
PhD 1	GSSI-LNGS	PhD		1,00
PhD 2	GSSI-LNGS	PhD		1,00
Postdoc	GSSI-LNGS	Postdoc		1,00
Bedogni R.	LNF	Ricercatore	0,05	
Benussi L.	LNF	Ricercatore	0,10	
Bianco S.	LNF	Primo Ricercatore	0,20	
Caponero M.	LNF	Primo Ricercatore	0,20	
Maccarone G.	LNF	Primo Ricercatore	0,40	
Mazzitelli G.	LNF	Primo Ricercatore	0,60	0,40
Piccolo D.	LNF	Primo Ricercatore	0,20	
Saviano G.	LNF	Ricercatore	0,20	
Tomassini S.	LNF	Tecnologo	0,10	0,10
Cavoto G.	Roma1	Ricercatore	0,20	0,10
D'Imperio G.	Roma1	Assegnista	0,50	
Di Marco E.	Roma1	Ricercatore	0,10	0,10
Marafini M.	Roma1	Ricercatore	0,20	
Messina A.	Roma1	Ricercatore	0,30	
Pinci D.	Roma1	Ricercatore	0,40	0,10
Renga F.	Roma1	Ricercatore	0,30	0,10
Iacoangeli F.	Roma1	Tecnologo	0,20	
Petrucci F.	Roma3	Ricercatore	0,20	
Totale			4,65	5,70

Budget Phase1 2020

Richieste CSNII 2020

RM1	Richieste k€
DAQ (sviluppo schede DAQ)	8
Missioni (LNGS, LNF, CERN, Conferenze)	15
TOTALE	23

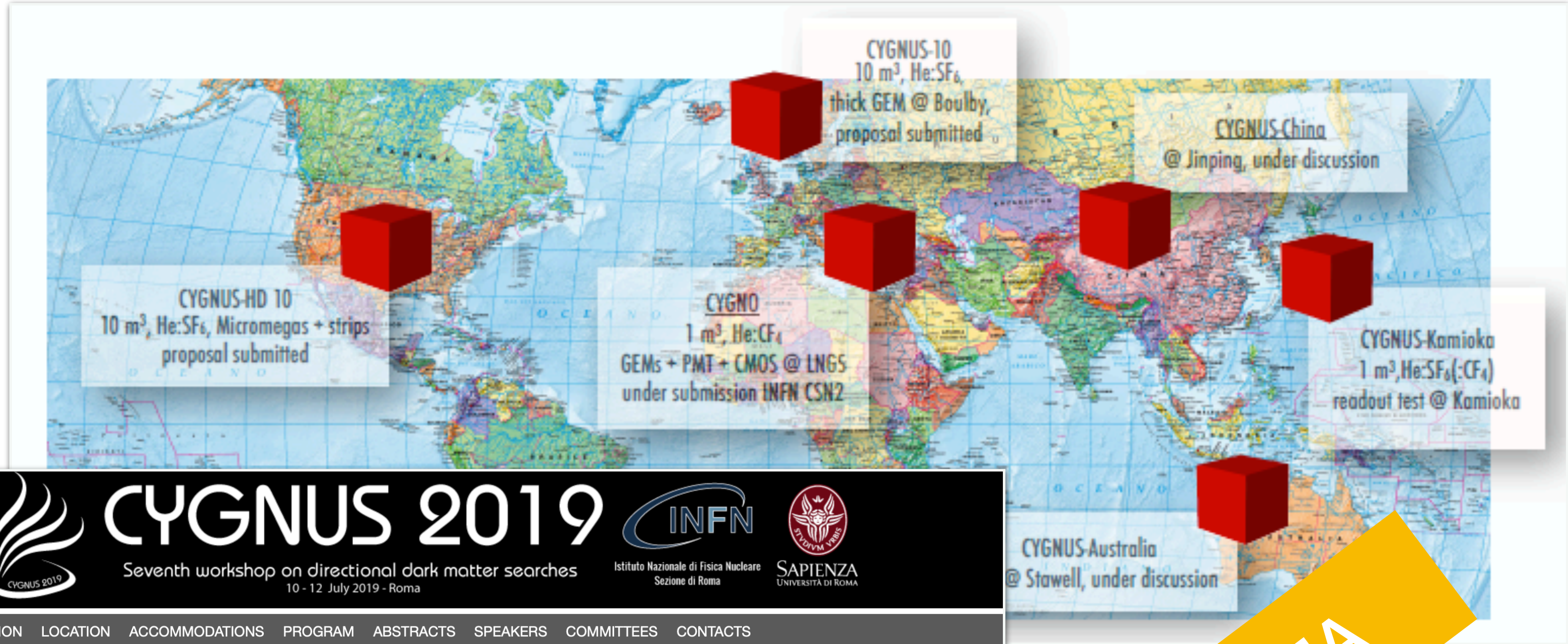
LNGS	Richieste k€
GAS (LIMF test @ LNGS)	3
Meccanica	2
Missioni (LNGS, LNF, CERN, Conferenze)	6
TOTALE	11

RM3	Richieste k€
Missioni (LNGS, LNF, Conferenze)	1
TOTALE	1

LNF	Richieste k€
PRA, etc	7
Test materiali (GEM a bassa radioattività)	15
Missioni (LNGS, RM1, CERN, Conferenze)	17
GAS (HeCF4, no aresol)	10
TRASPORTI	3
R&D sistema monitoraggio strutturale	5
TOTALE	57

Budget INITUM (su 5 anni)

- 700 k€ al netto dell'overhead INFN/RM1
- 200 k€ fondi NRC: personale LNF (1 art15 per 2 anni) + (1 ass ricerca per 2 anni) + 50 k€
- 500 k€ per costruzione apparato
- equivalente budget al GSSI per dottorandi e assegnasti



 **CYGNUS 2019**  
Seventh workshop on directional dark matter searches
10 - 12 July 2019 - Roma
Istituto Nazionale di Fisica Nucleare
Sezione di Roma
SAPIENZA
UNIVERSITÀ DI ROMA

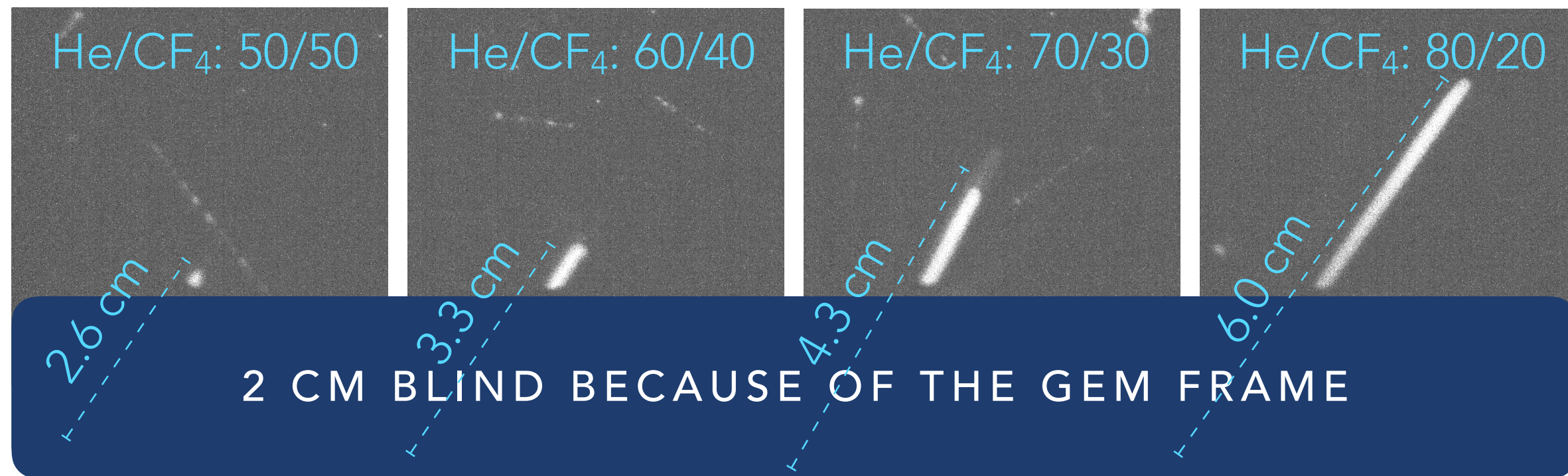
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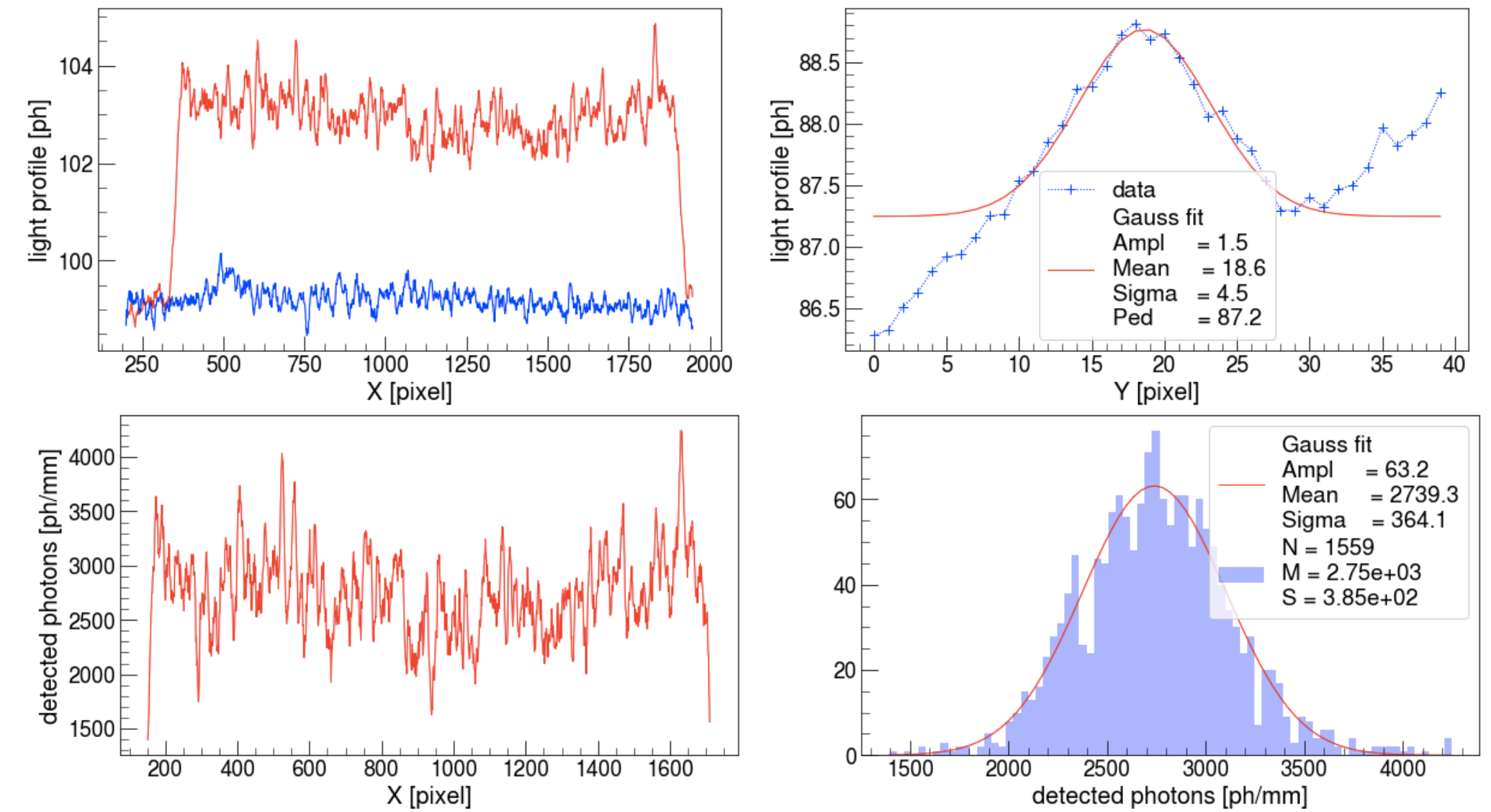
10-12 July @ ROMA

Range and recoil energy released

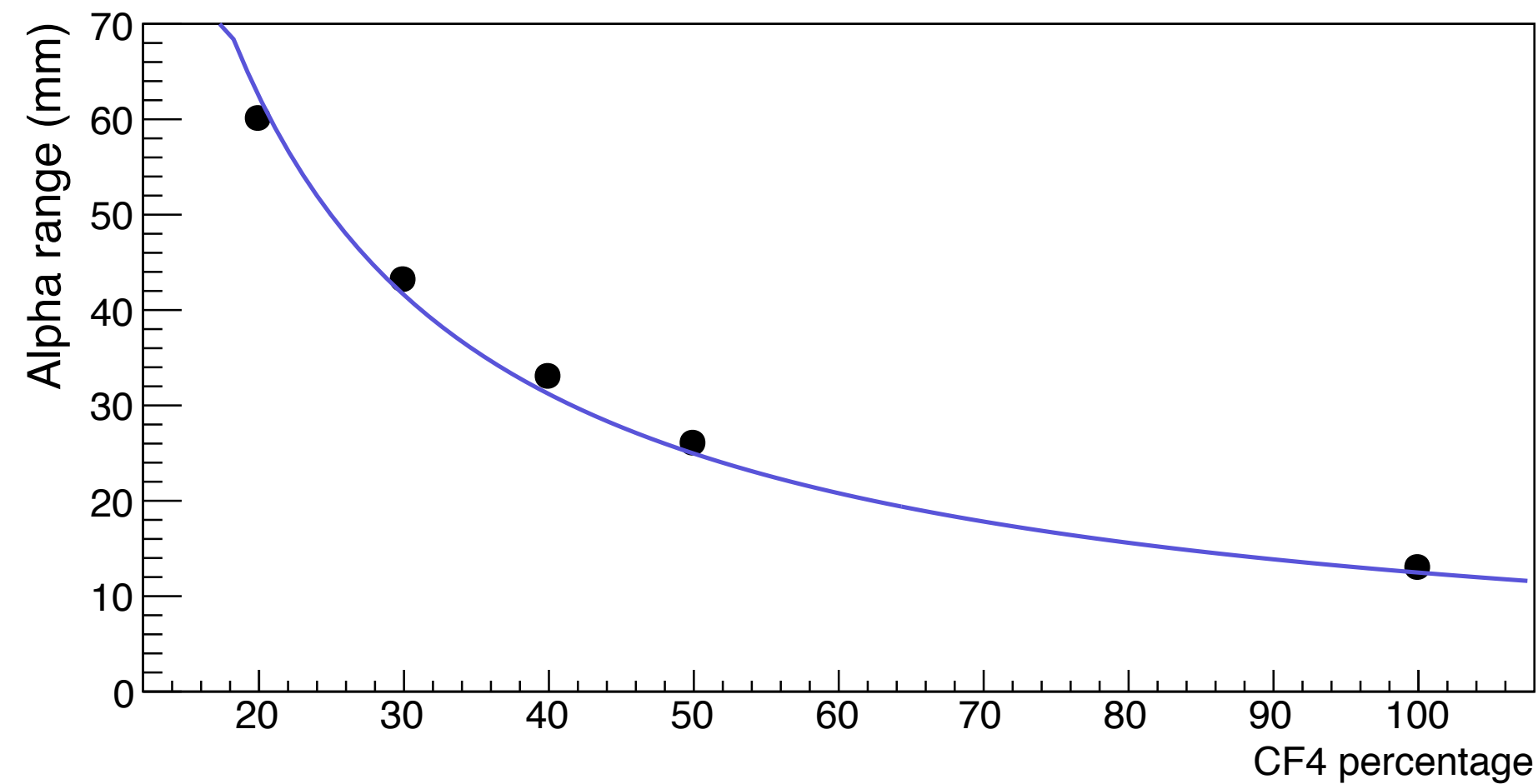
Am alpha source (5.48 MeV);
alpha range in HeCF₄



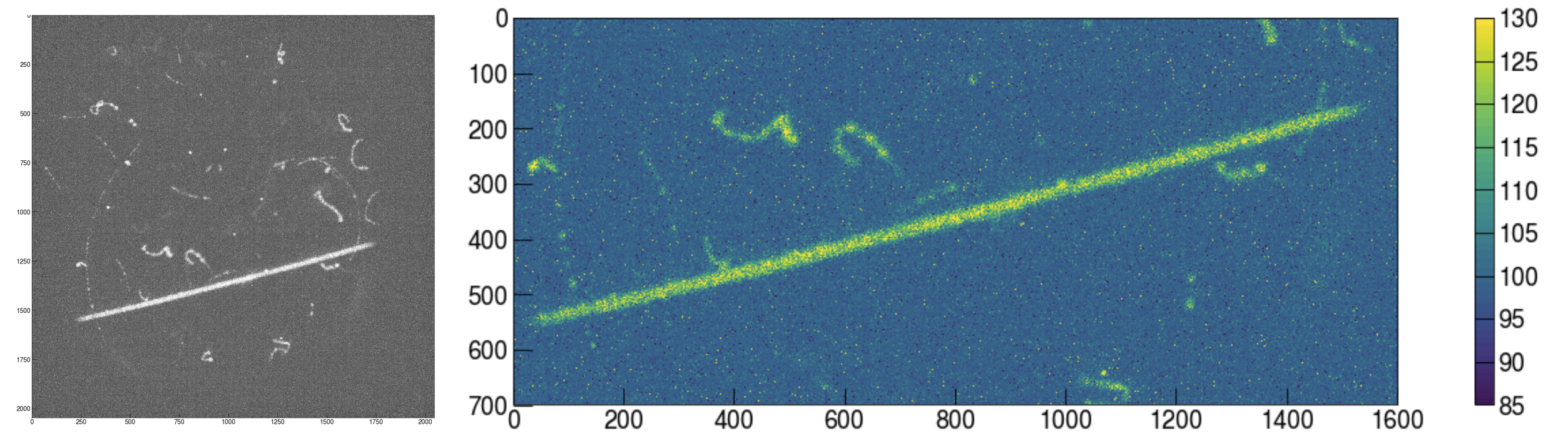
AmBe (5-11 MeV) neutron source;
nuclear recoil in HeCF₄ / 60:40



GEM 440/2kVcm - drift 1kV/cm - 2000 ms

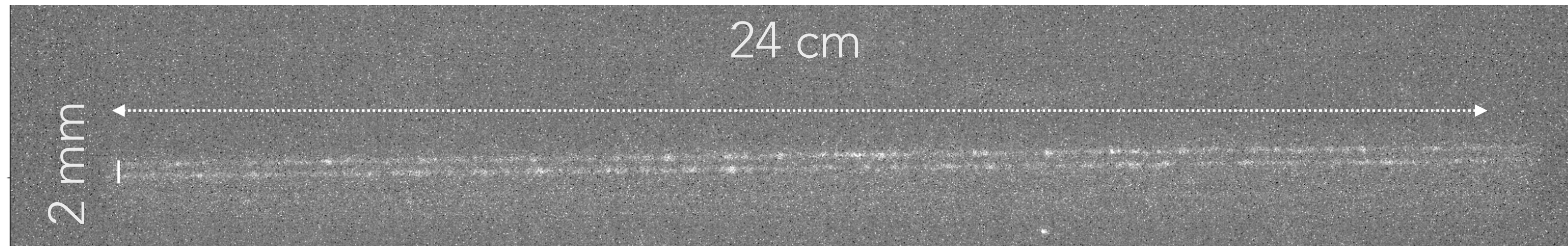


alpha range seems to be “determined” only by
CF₄ and to decrease linearly with its amount

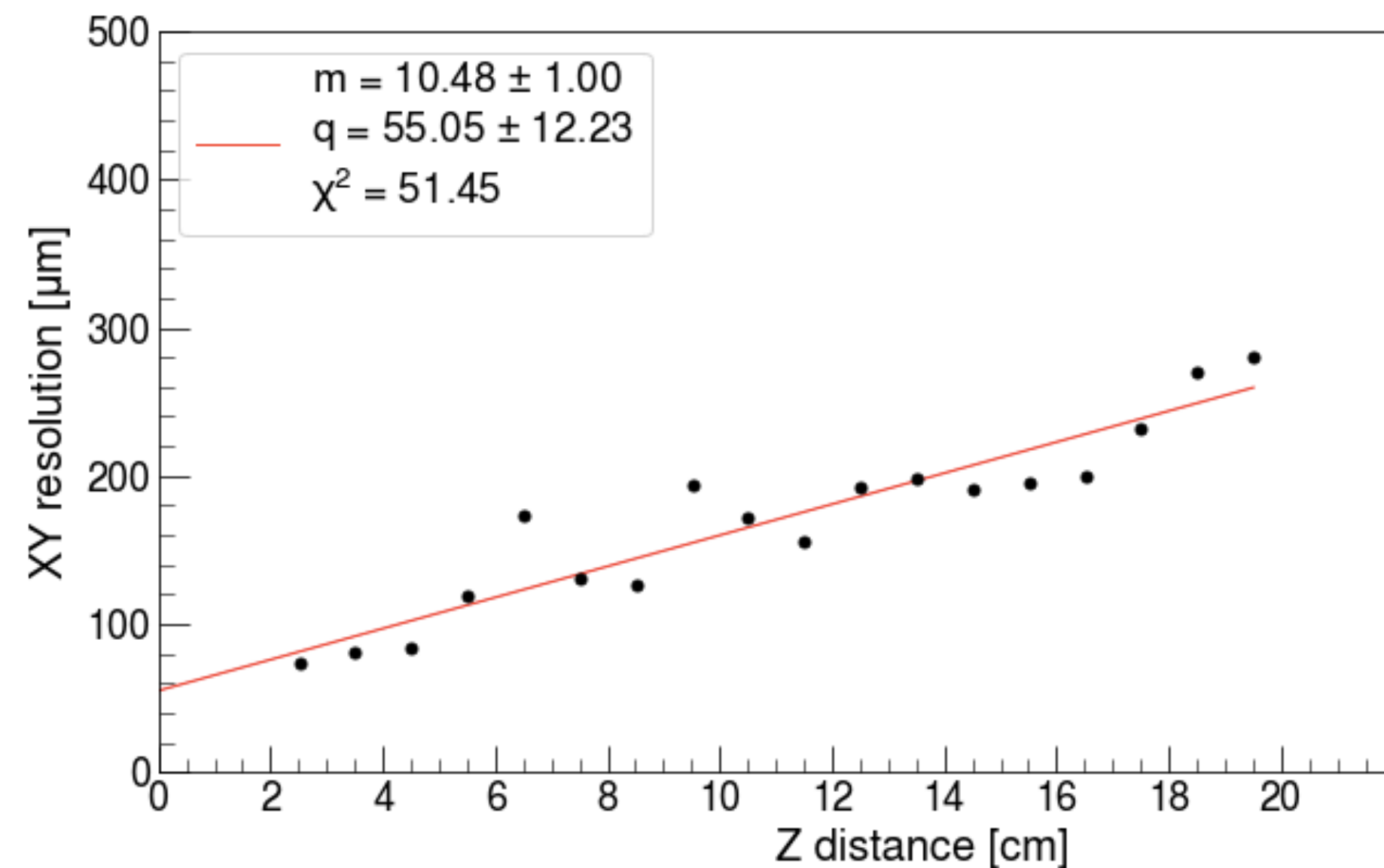


2740 [ph/mm] / 1.5 [ph/eV] → 1.7 keV/mm (55 μm resolution)

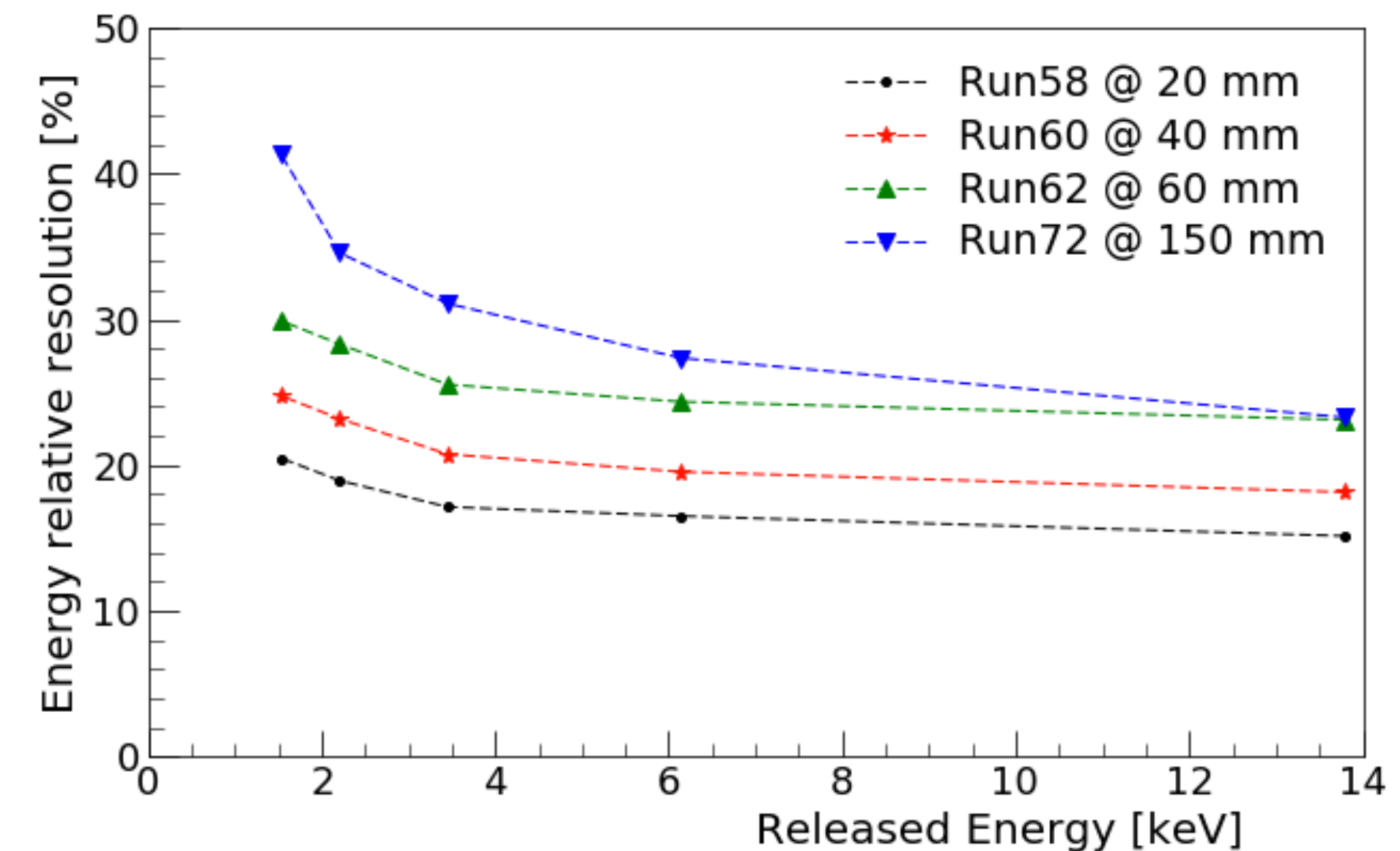
XY and energy resolution



XY resolution vs depths (Z)



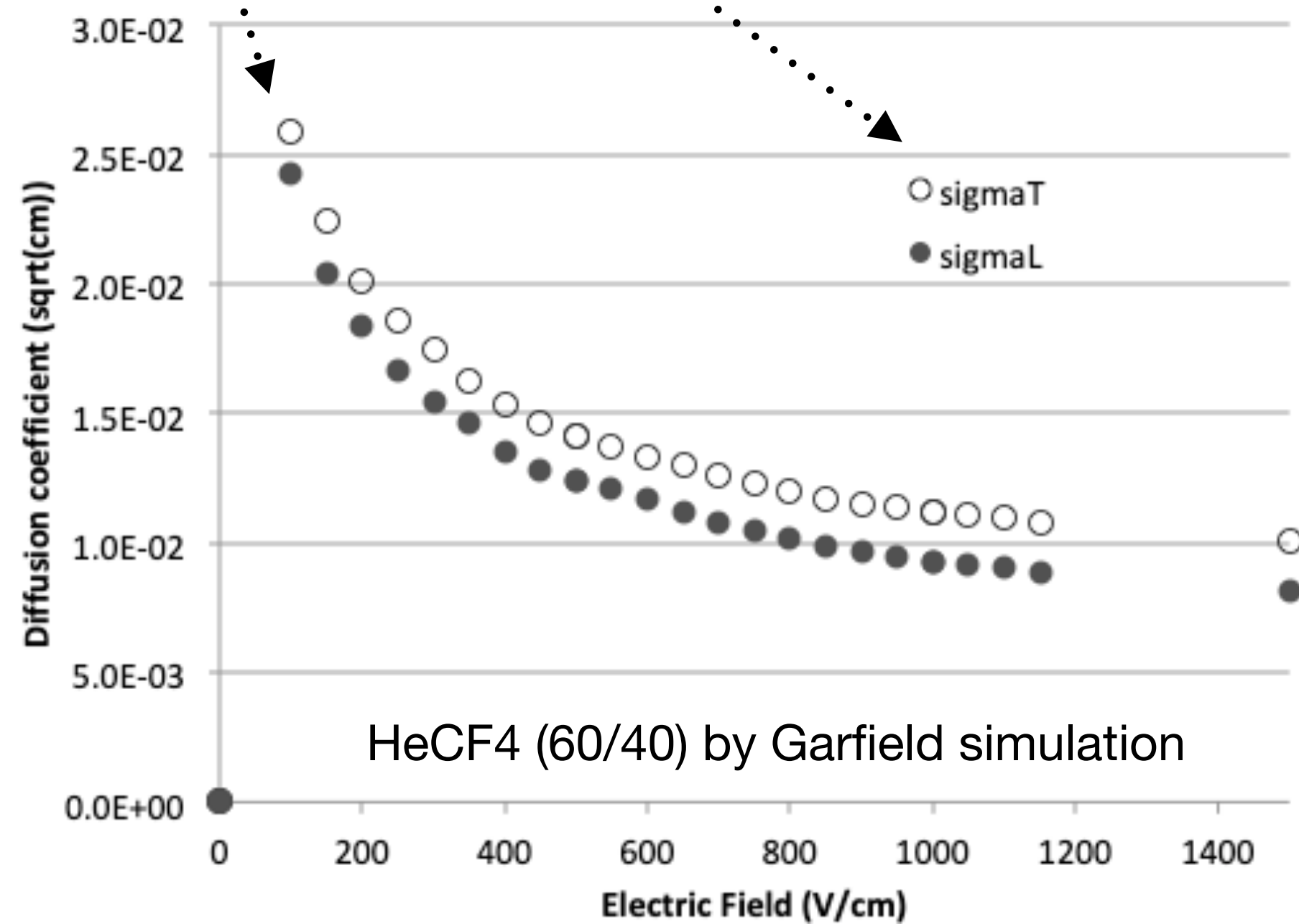
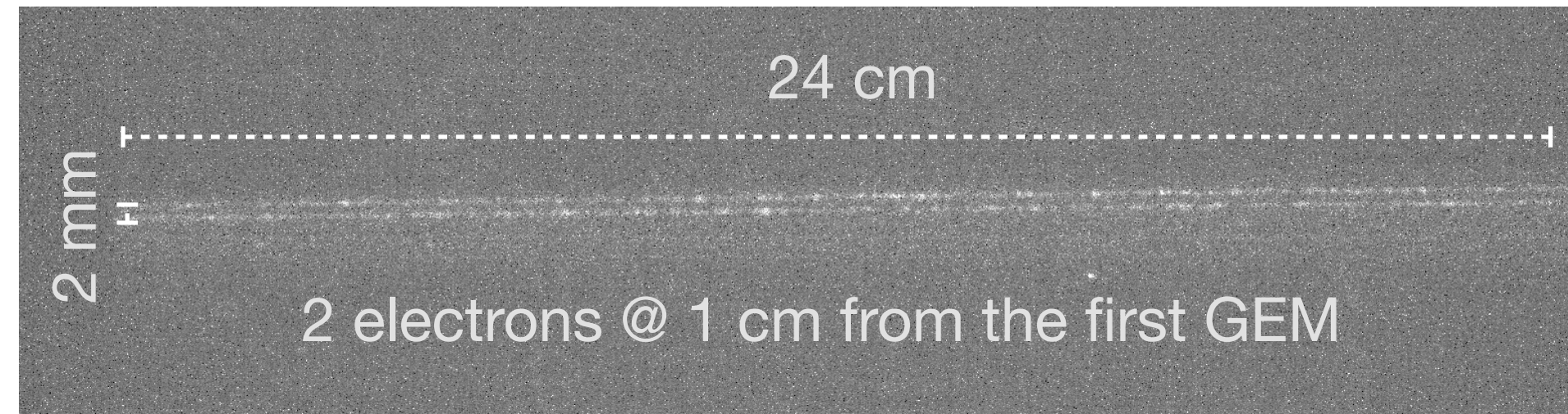
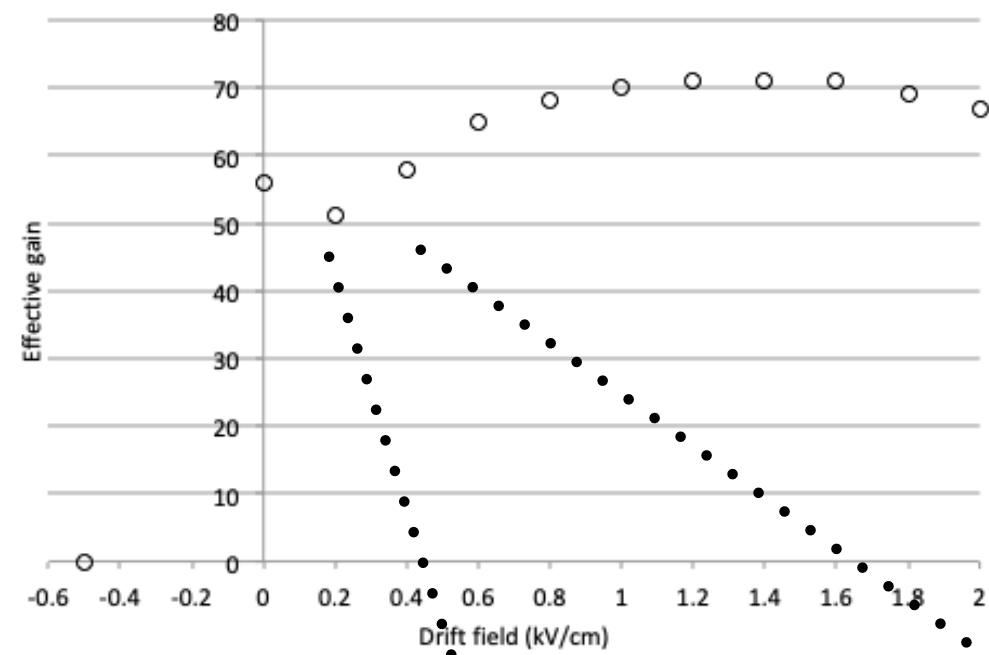
energy resolution @ depths (Z)



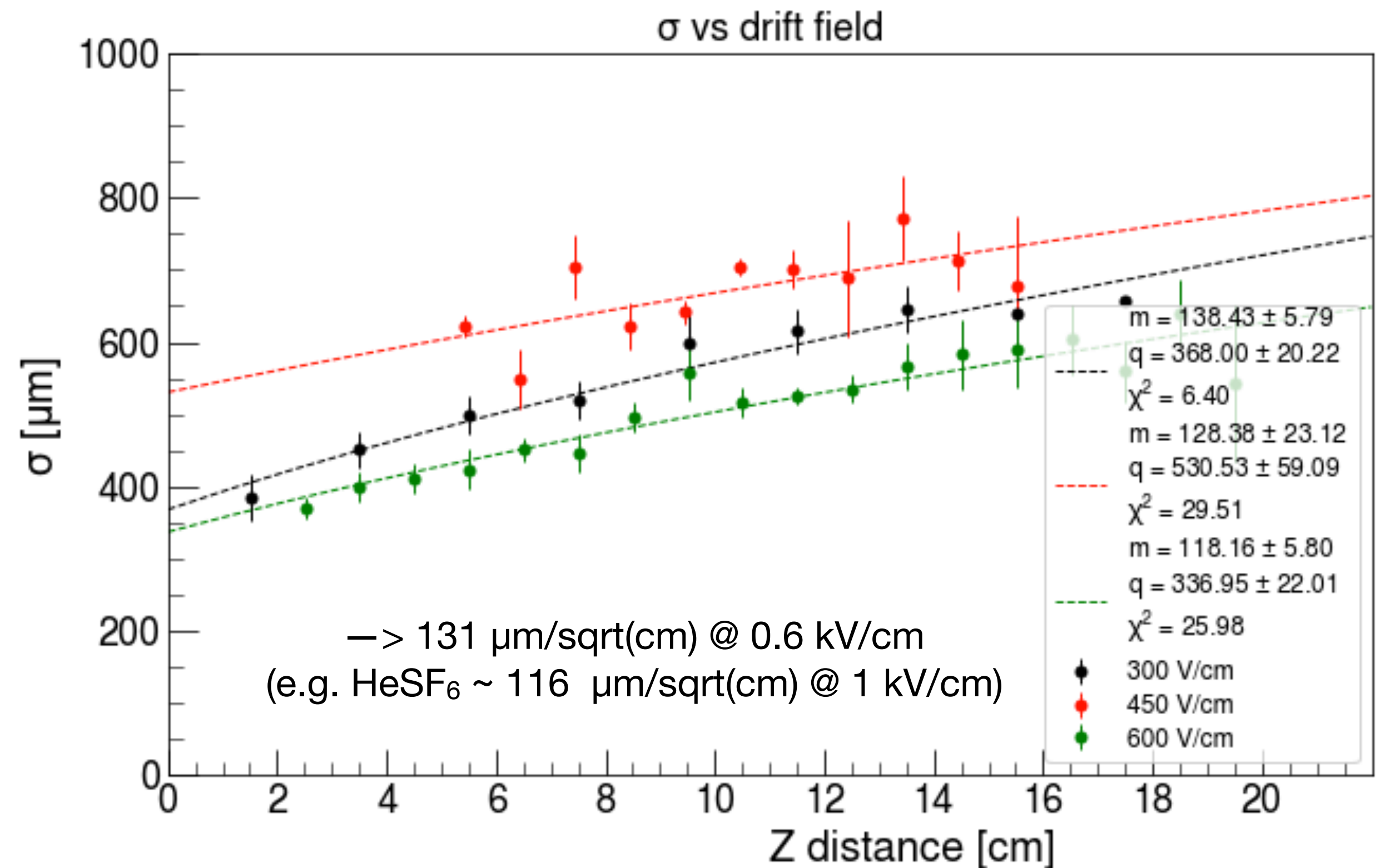
resolution obtained with a drift field to 0.6kV/cm

in the few keV region a relative resolution of 20%-30% is achieved

diffusion in HeCF₄ for electrons



$V_{\text{drift}} = 4.36 \text{ cm}/\mu\text{s} @ 0.6 \text{ kV/cm}$



→ 131 $\mu\text{m}/\sqrt{\text{cm}}$ @ 0.6 kV/cm
(e.g. HeSF₆ ~ 116 $\mu\text{m}/\sqrt{\text{cm}}$ @ 1 kV/cm)

LEMON data @ Frascati BTF

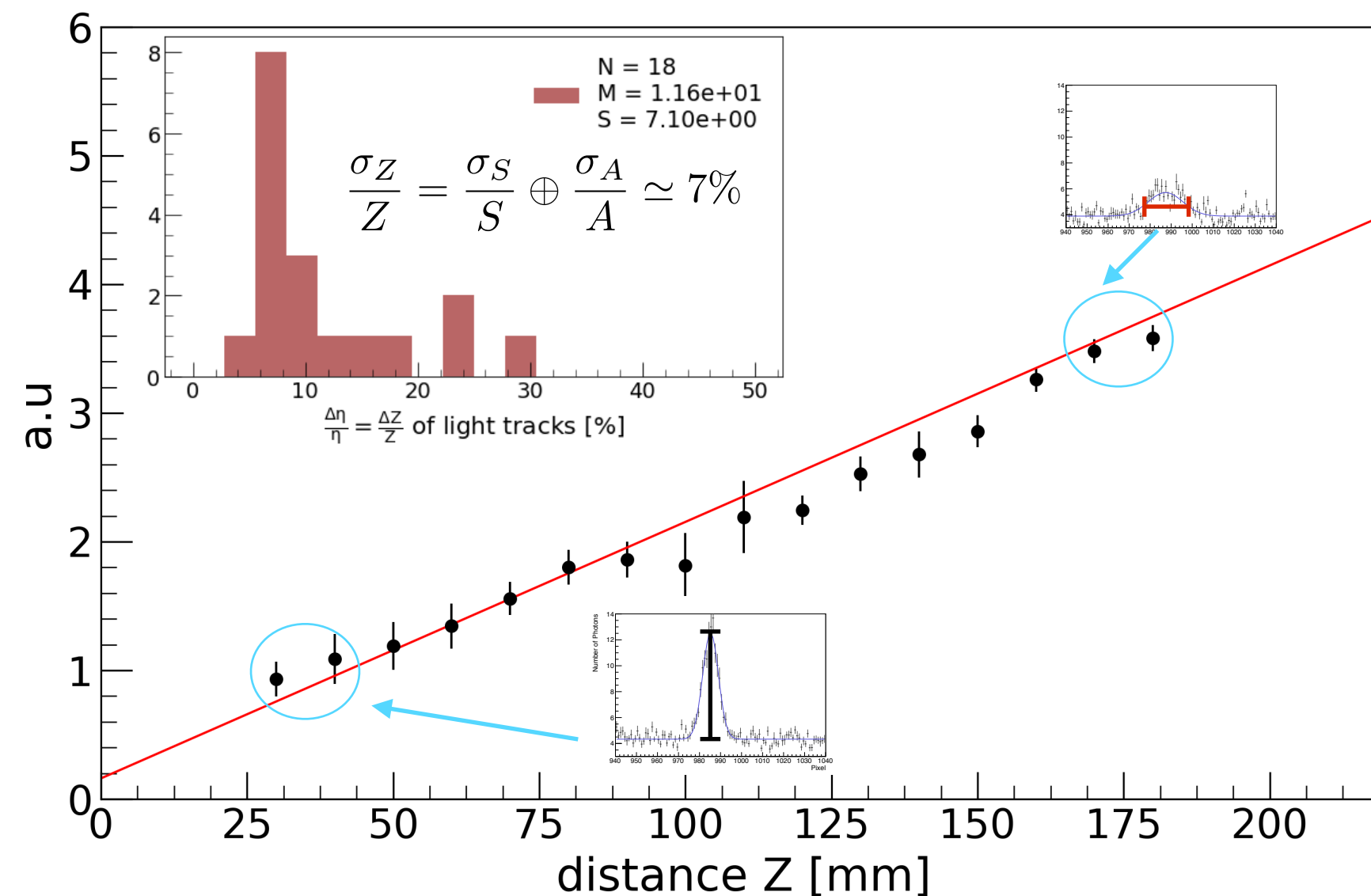
Z resolution in LEMOn



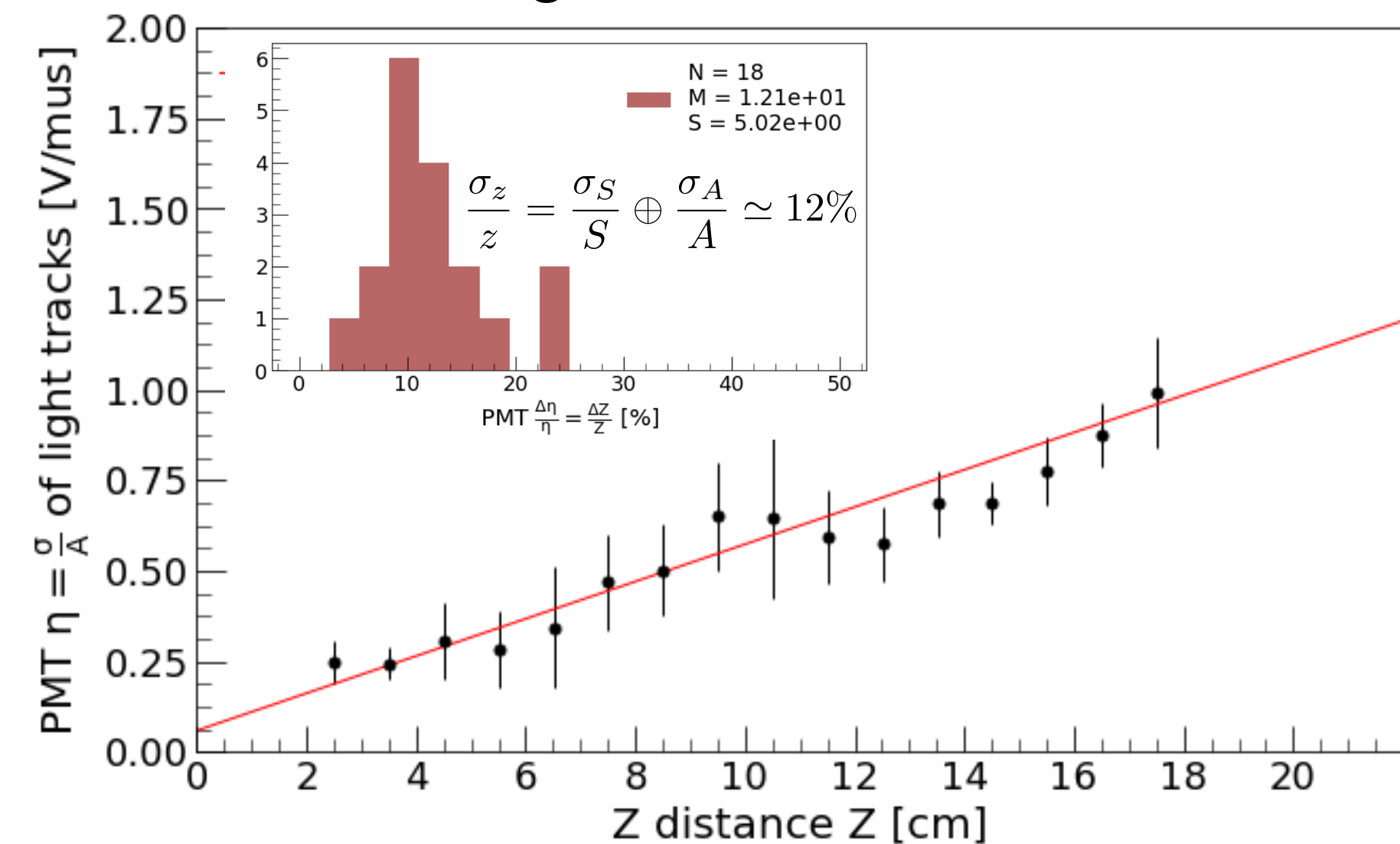
Electron diffusion in the drift gap can be exploited to evaluate the Z of the event. The transverse light profile and the PMT signal waveform are expected to become lower and larger as long as the event is far from the GEM;

Since the amplitude (A) decreases and the width (S) increases with Z, their ratio $\eta = S/A$ increases (independently from the amount of produced light);

transverse diffusion



longitudinal diffusion



both methods gives 10% precision: $\sigma_z \sim 2 \text{ cm @ } 20 \text{ cm}$