

In Pursuit of the Invisible: Measuring $Y \rightarrow \nu \bar{\nu}$ at Super-B

Steve Sekula

Southern Methodist University

Presented at the XI Super-B Workshop

November 30, 2009



SMU



- The lure of the invisible - *predictions*
- The challenge of the invisible – *systematics*
- Pursuing the invisible - *approaches*

The Lure of the Invisible

Predictions for $Y \rightarrow \nu \bar{\nu}$

THE STANDARD MODEL

$$BR(Y(1S) \rightarrow \nu \bar{\nu}) = \frac{N_\nu G_F^2}{48 \pi} \left(1 - \frac{4}{3} \sin^2 \theta_W \right)^2 \frac{f_{Y(1S)}^2 M_{Y(1S)}^3}{\Gamma_{Y(1S)}}$$

$$BR(Y(1S) \rightarrow \nu \bar{\nu}) = (1.03 \pm 0.04) \times 10^{-5}$$

From Yeghiyan (see Reference Backup Slide)

LOW-MASS DARK MATTER

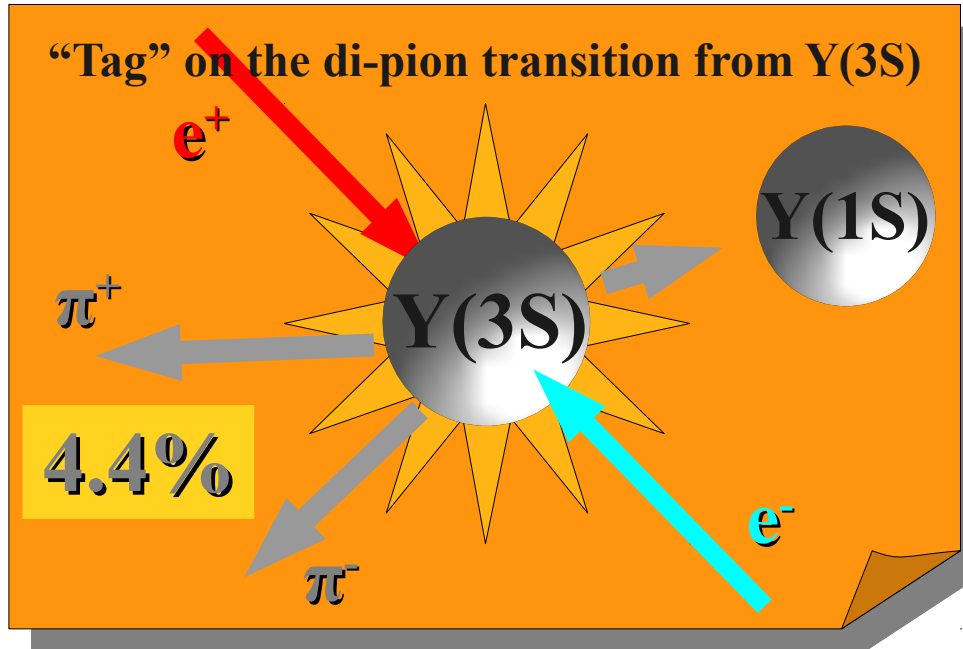
Fayet, McElrath, Yeghiyan, ...

Most recently, Yeghiyan calculated from an effective theory that:

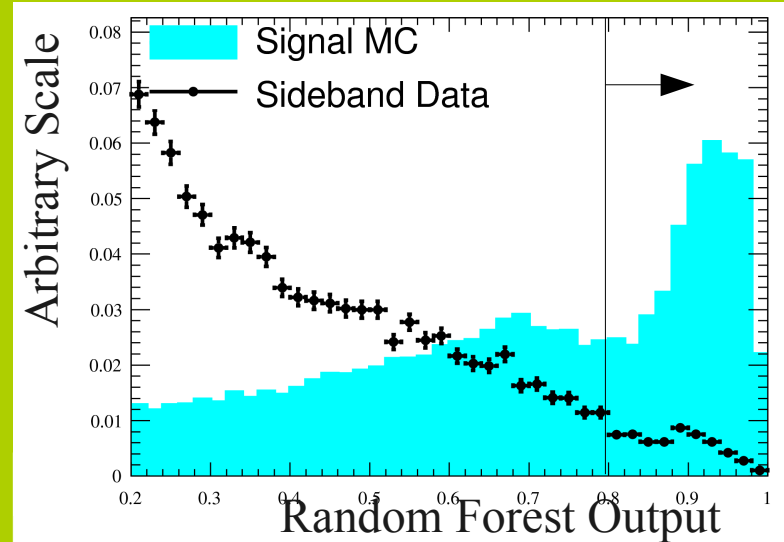
$$BR(Y(1S) \rightarrow \phi \bar{\phi}) = \frac{C_3^2}{\Lambda_H^4} \frac{f_{Y(1S)}^2}{48 \pi \Gamma_{Y(1S)}} \left(M_{Y(1S)}^2 - 4m_\phi^2 \right)^{3/2}$$

where the production of the dark matter is mediated by heavy degrees of freedom whose mass scale is Λ_H and where C_3 is the (real-valued) Wilson coefficient for the term in the effective theory that leads to this final state.

Measurement Technique

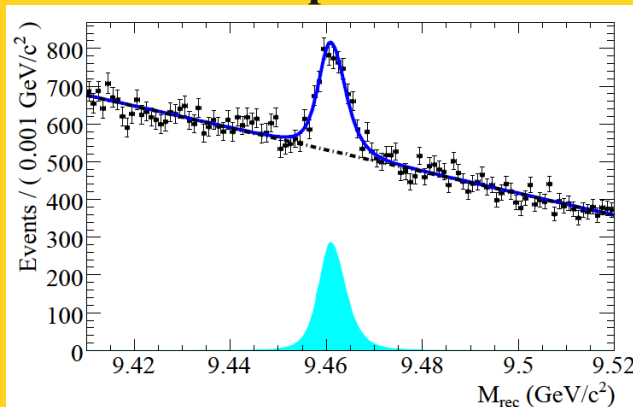


Reject background with a random forest (decision tree algorithm)



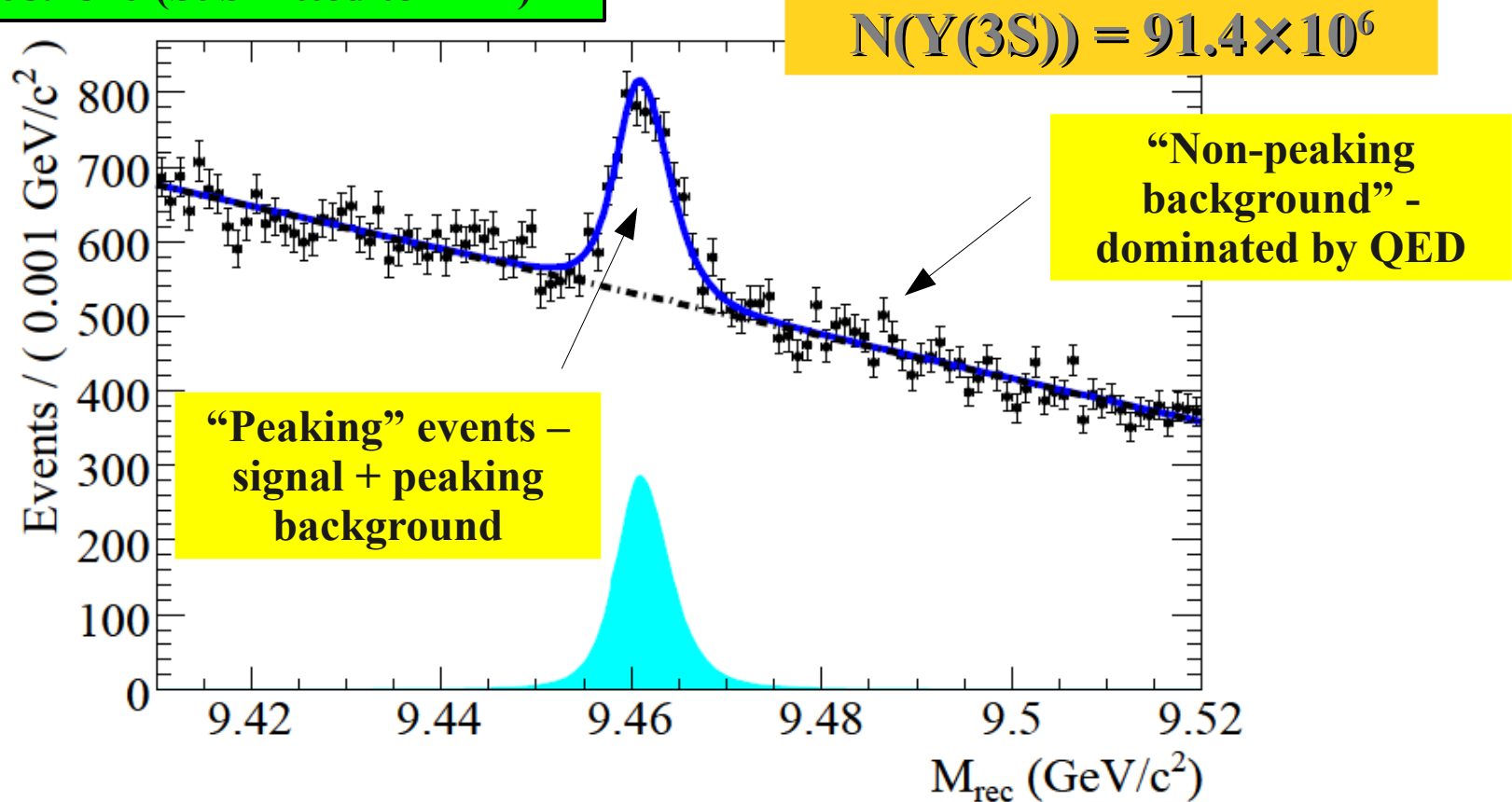
$$\epsilon_{\text{sig}}^{\text{total}} = 17\%$$

Fit for a peak in the mass recoiling against the pions



Present Measurement of $Y \rightarrow \nu\bar{\nu}$

arXiv:0908.2840 (Submitted to PRL)



Peaking Background: 2444 ± 123 events

Signal Yield: $-118 \pm 105 \pm 124$

$\text{BR}(Y \rightarrow \text{invisible}) = (-1.6 \pm 1.4 \pm 1.6) \times 10^{-4}$

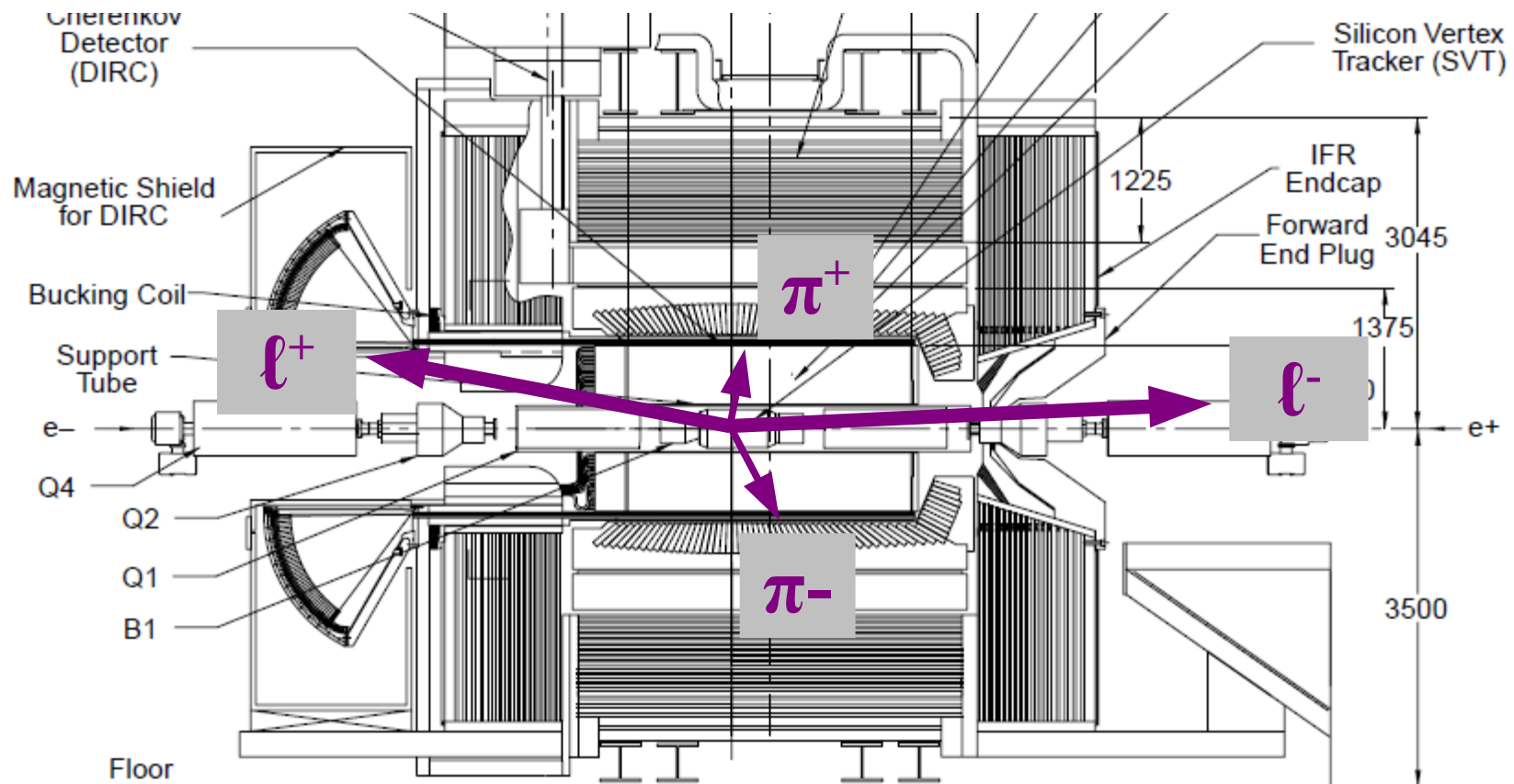
Why Care About $Y \rightarrow \nu\bar{\nu}$?

- Fundamental physics
 - We haven't measured an invisible meson decay
 - This is a straight-forward but rare process in the SM
 - a whole order-of-magnitude of discovery is left!
- Enabling other measurements
 - Measuring this will be challenging
 - Meeting this challenge may make other work easier
 - I will elaborate more on this at the end

The Challenge of the Invisible

A Discussion of Peaking Background

- Peaking background dominated by $Y \rightarrow e^+e^-$ and $\mu^+\mu^-$
 - 4% from $\tau^+\tau^-$ and $<1\%$ from hadrons



A Breakdown of Current Uncertainties

2444.0

Green boxes indicate uncertainties that improve with statistics; yellow indicate those which could improve by other means.

± 28.0

Due to statistics in a control sample (manifests as uncertainty in correction to peaking background)

± 14.0

ibid. (manifests as a fit yield uncertainty on control sample events)

± 22.0

Due to limitation on knowledge of the different trigger efficiencies for control/invisible events

± 15.7

Due to limitation on knowledge of the rate at which hadronic $Y(1S)$ decays mimic the invisible signature

$\pm 0.9\%$

Due to limitation on knowledge of the Level 3 trigger efficiency for signal(-like) events

**The
single
largest
effects**

$\pm 2.1\%$

Due to limitation on knowledge of the Level 1 trigger efficiency for signal(-like) events

$\pm 4.0\%$

Uncertainty on the Random Forest selection efficiency for signal(-like) events

The total systematic uncertainty is about 5%

Pursuing the Invisible

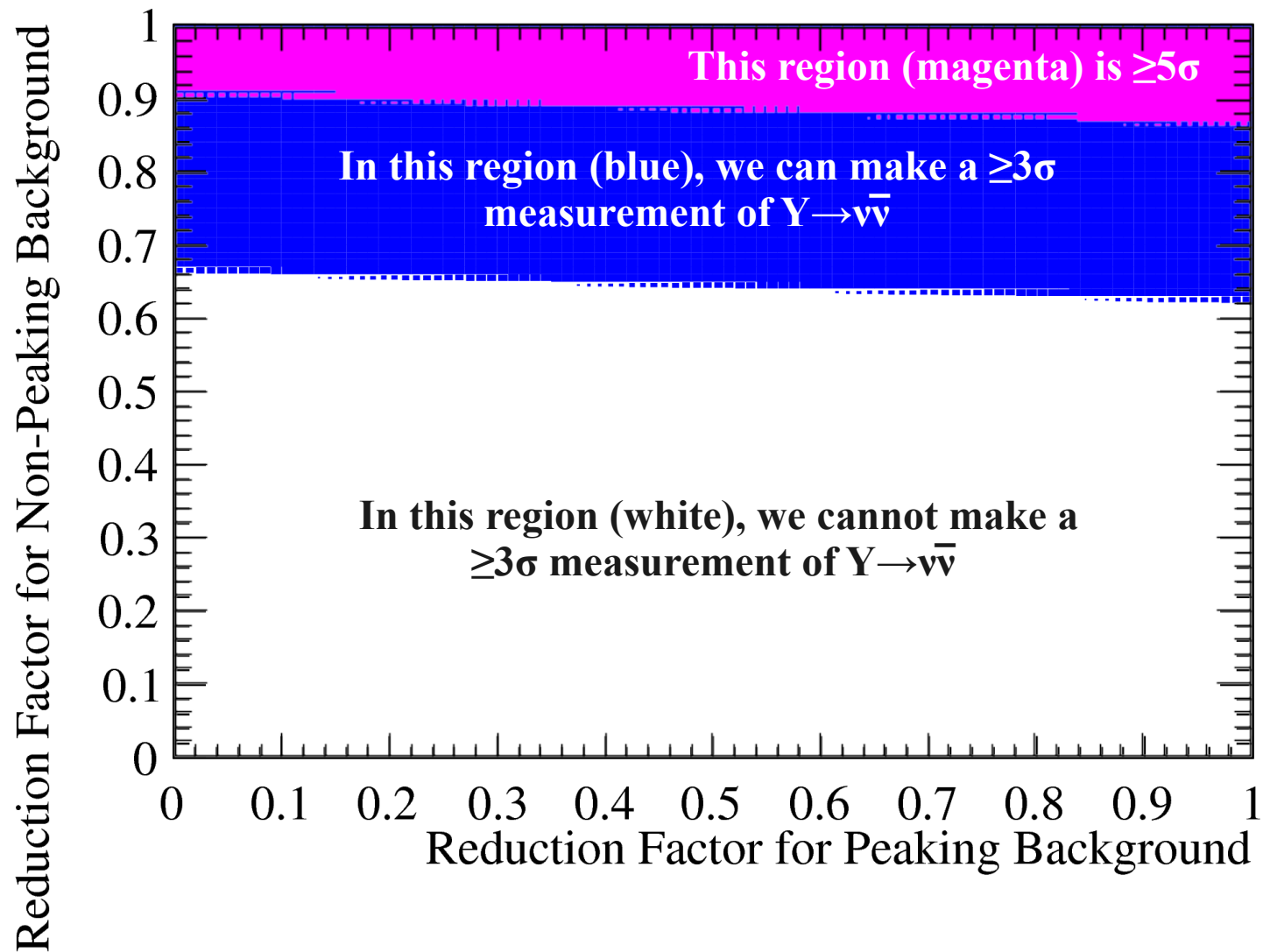
Assumptions for this Study

- 100x more $Y(3S)$ at Super-B ~few months running at full luminosity
- Similar trigger configuration for this running
 - need low- p_T 2-track triggers to catch the pions
- Systematic errors
 - two cases: stay the same as now (unlikely!), or improve by a factor of 2 (likely!)
- Detector design
 - similar to current fiducial coverage, with upgrades/replacements to appropriate systems that yield similar performance

Question: by how much would we need to reduce either the non-peaking or peaking background (or both) in order to achieve at least a 3σ -significant measurement?

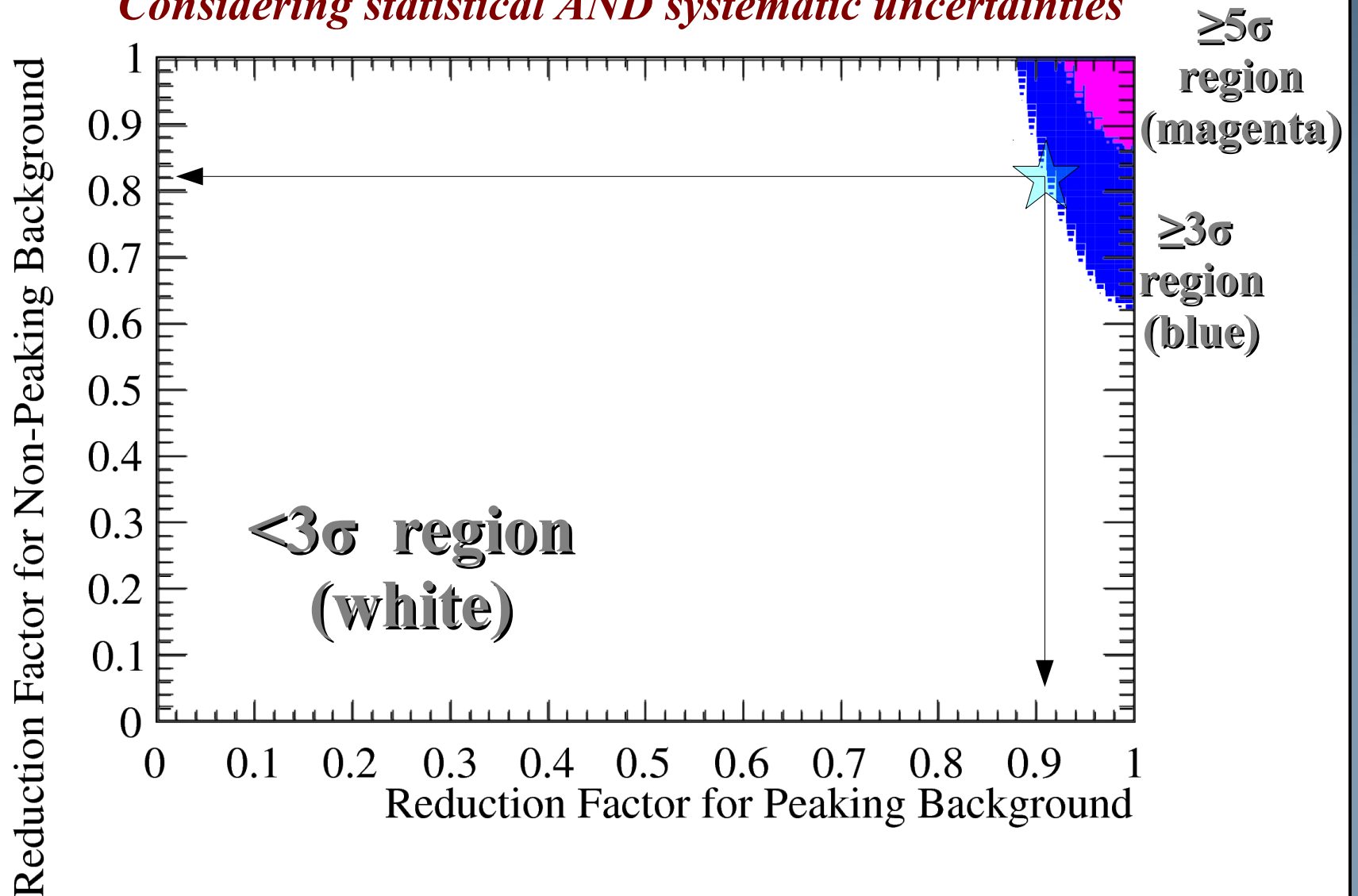
Background Reductions – Part I

Considering ONLY statistical uncertainties



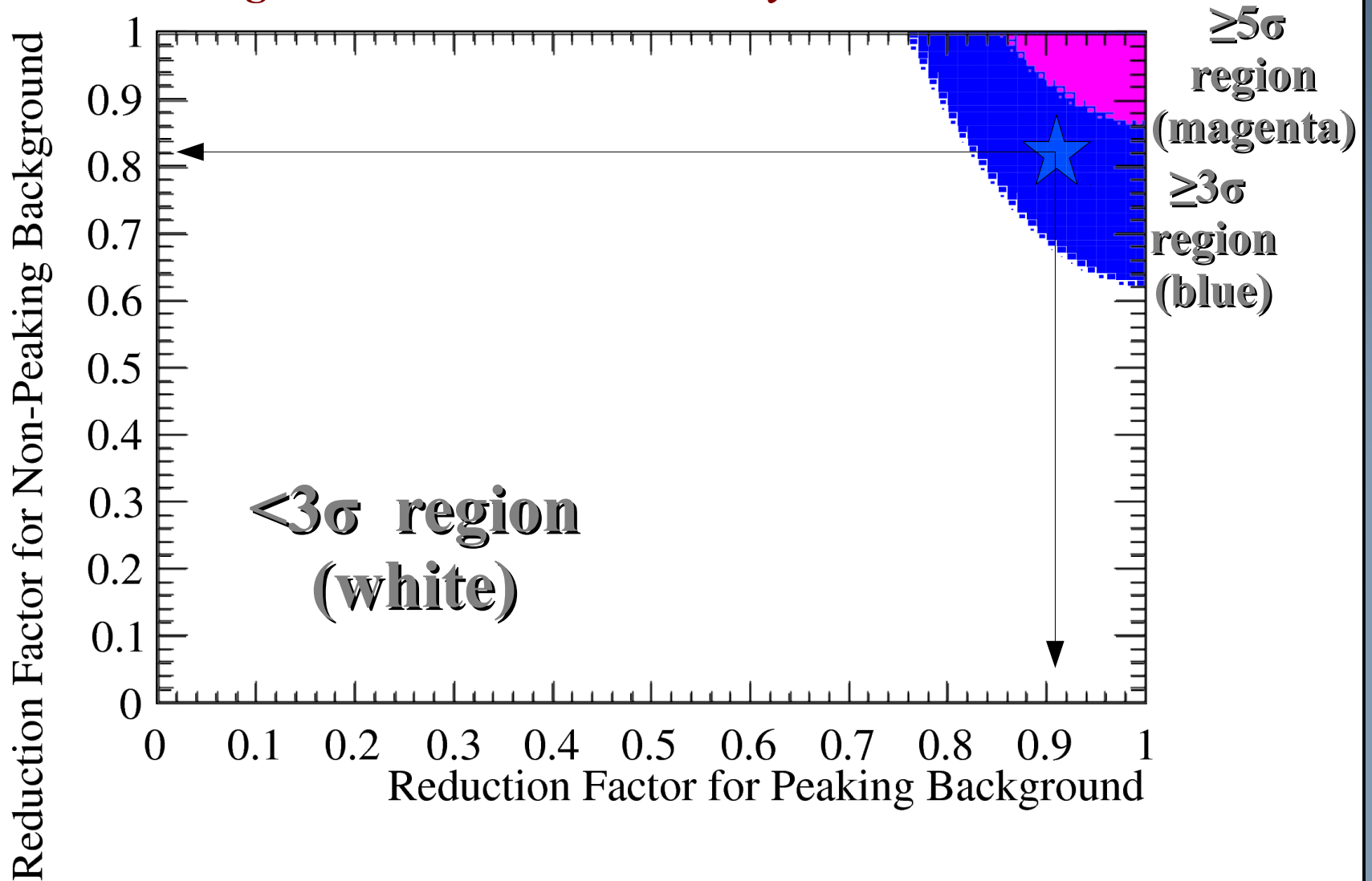
Background Reductions – Part II

Considering statistical AND systematic uncertainties



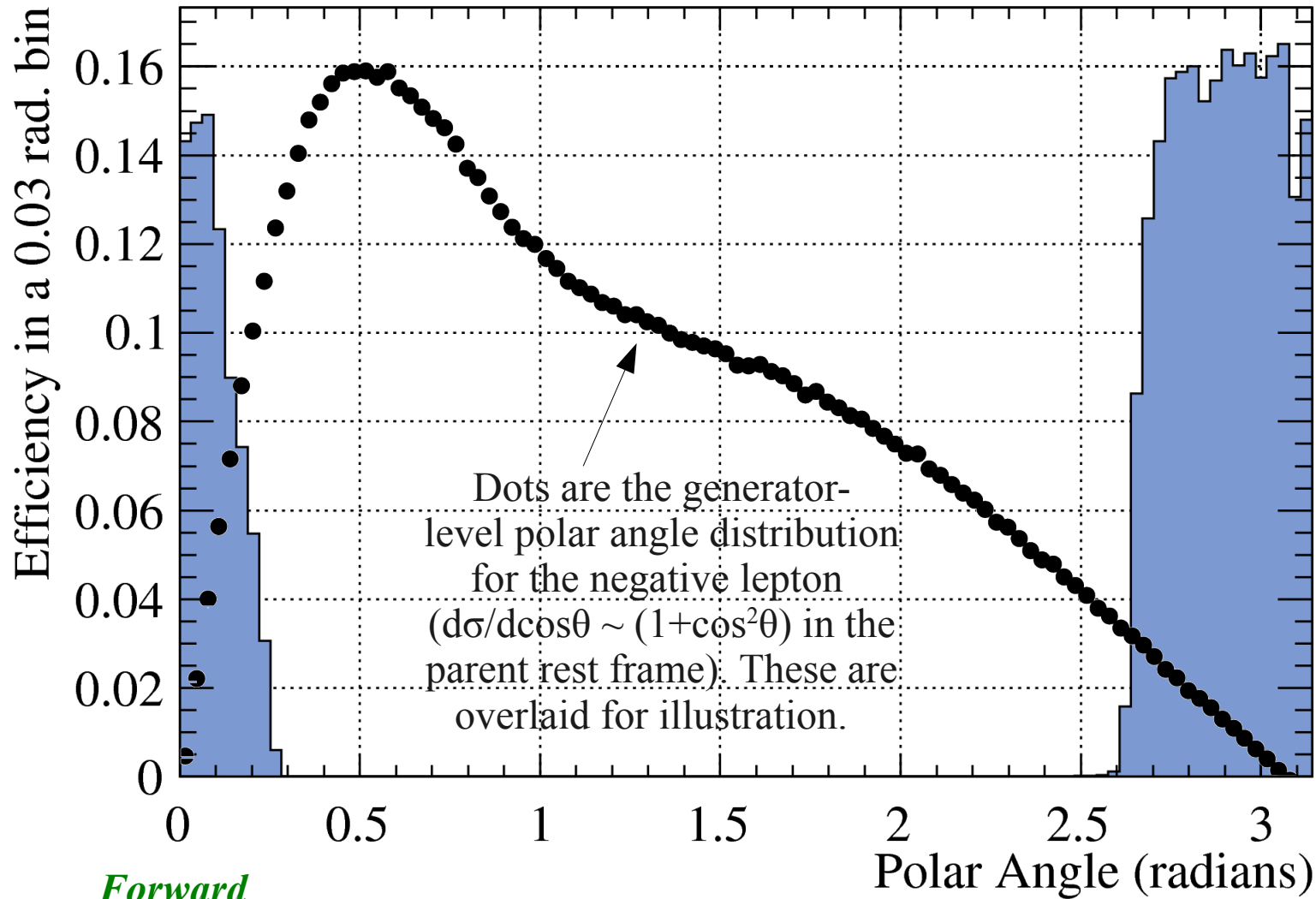
Background Reductions – Part III

Considering statistical AND reduced systematic uncertainties



Peaking Background

Efficiency for selecting an $Y \rightarrow \ell^+ \ell^-$ event as “invisible” vs. polar angle of negatively-charged true lepton

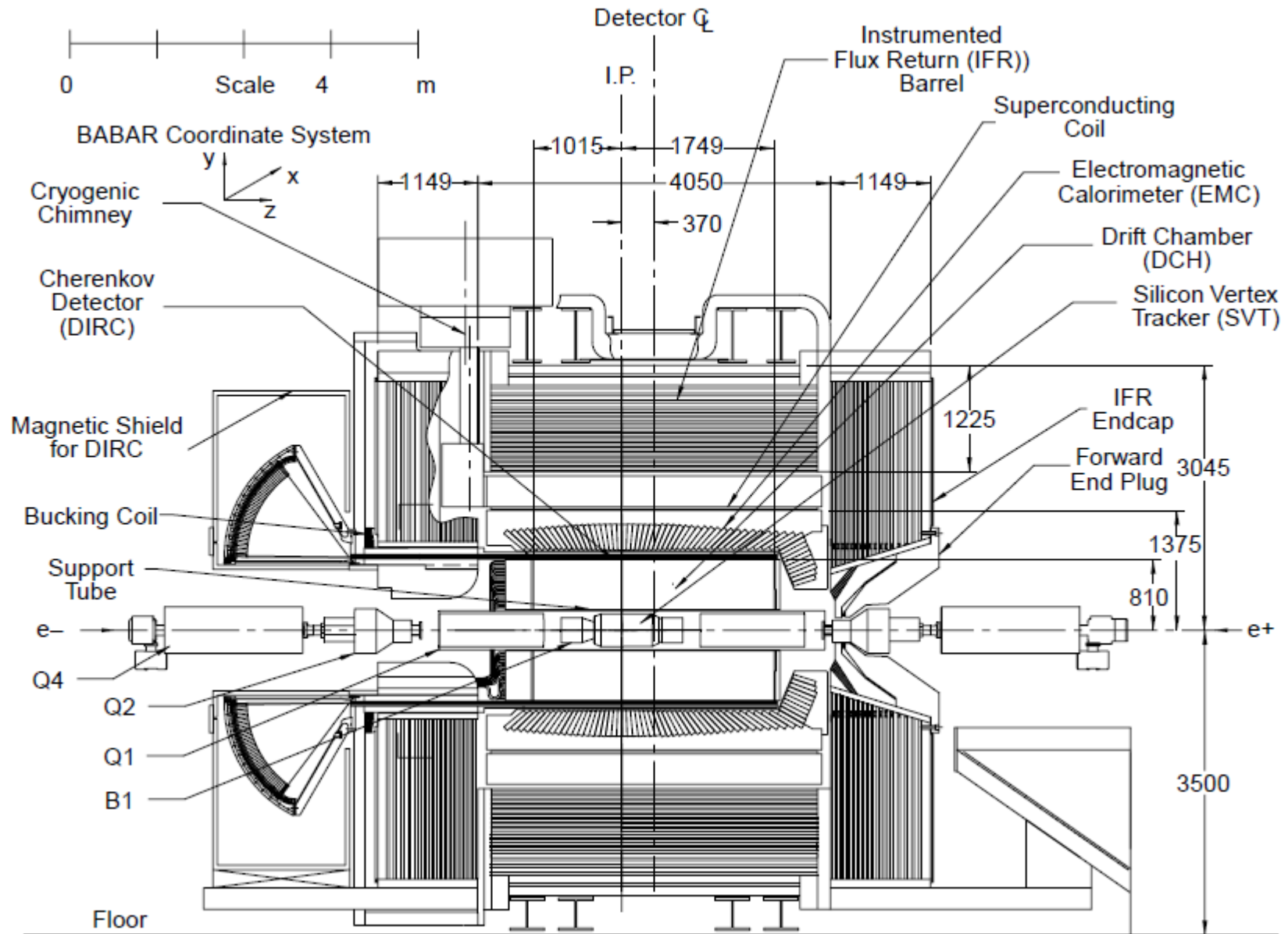


*Forward
Detector Region*

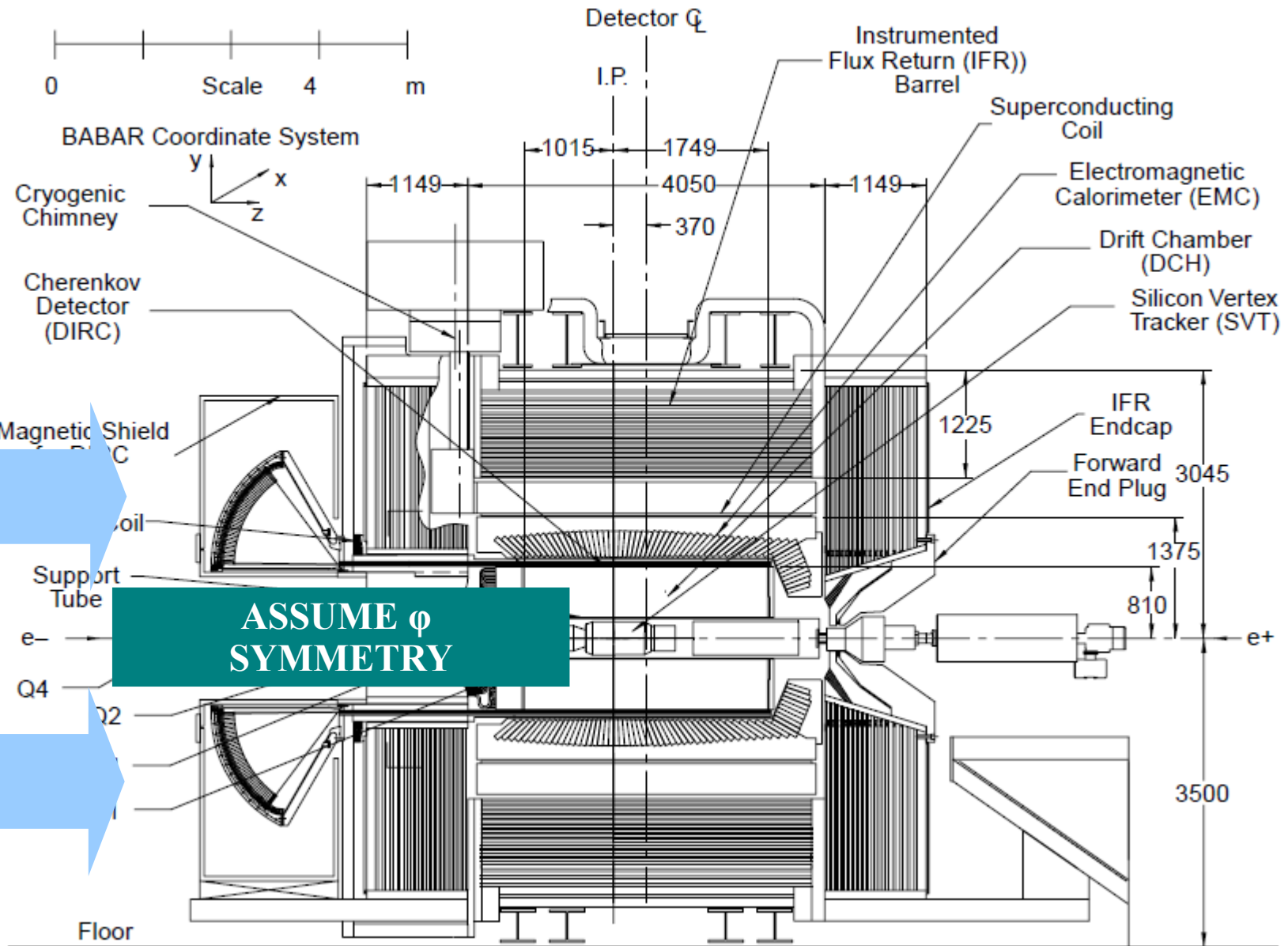
Stephen Sekula - SMU

*Backward
Detector Region* 17

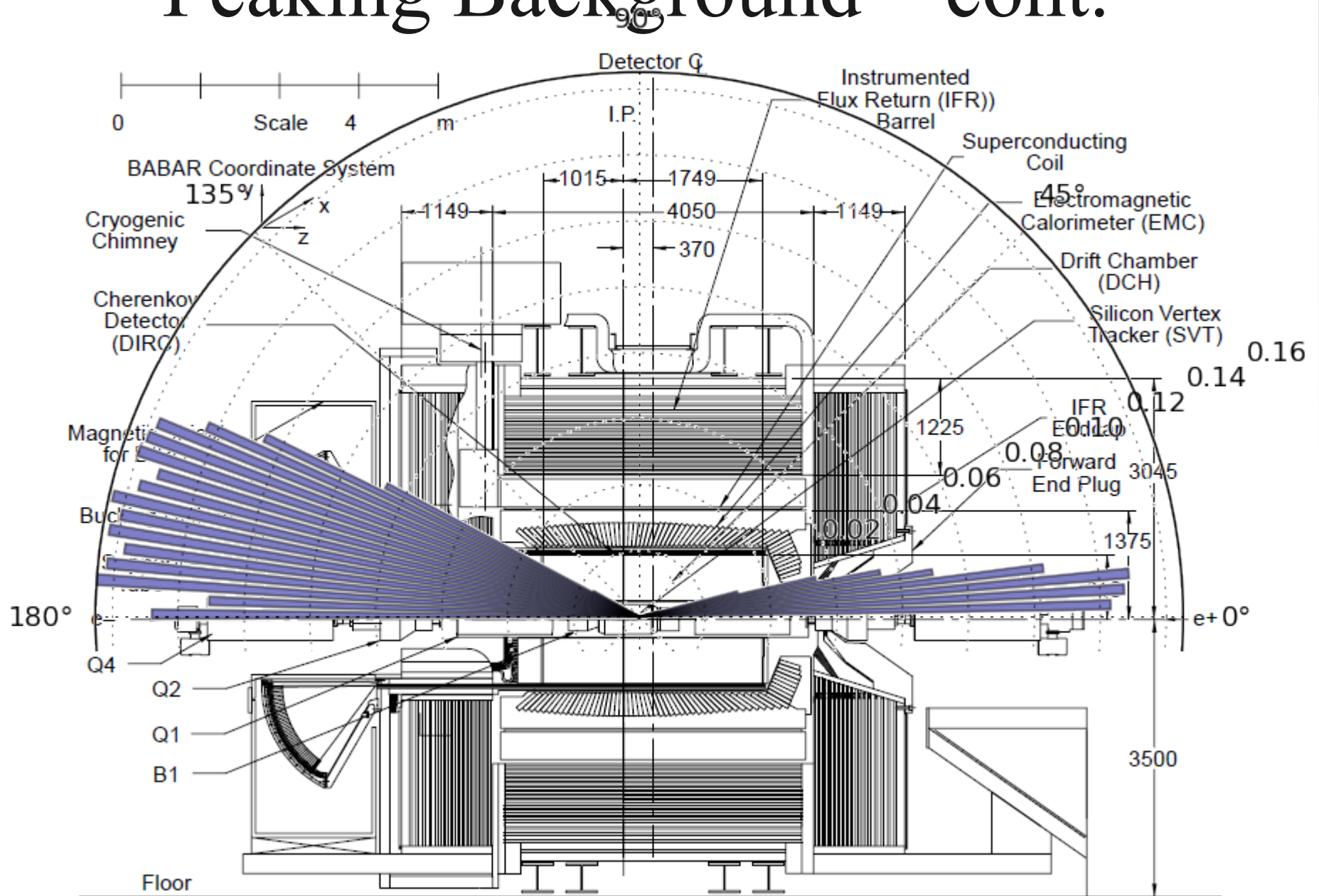
Peaking Background – cont.



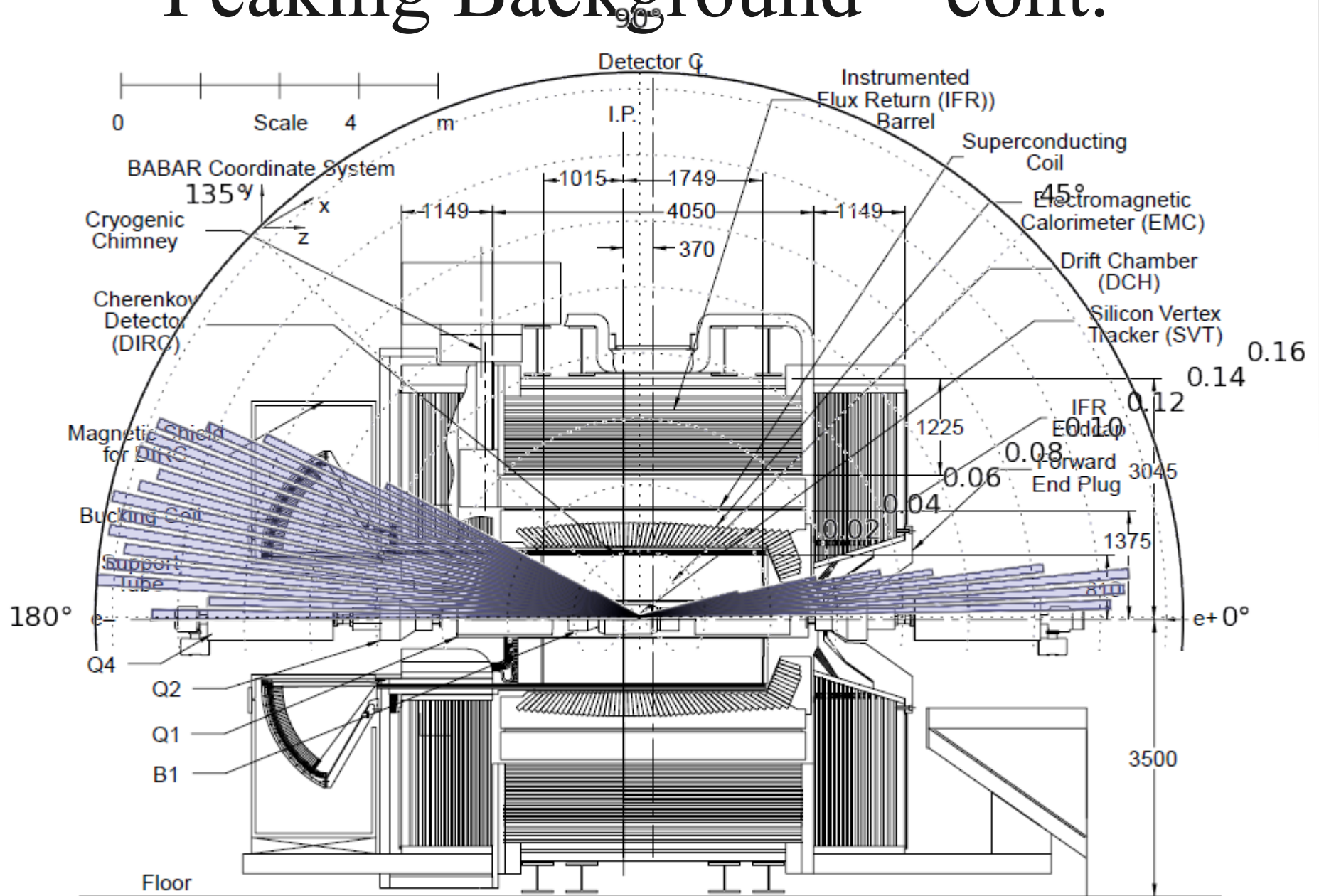
Peaking Background – cont.



Peaking Background – cont.



Peaking Background – cont.

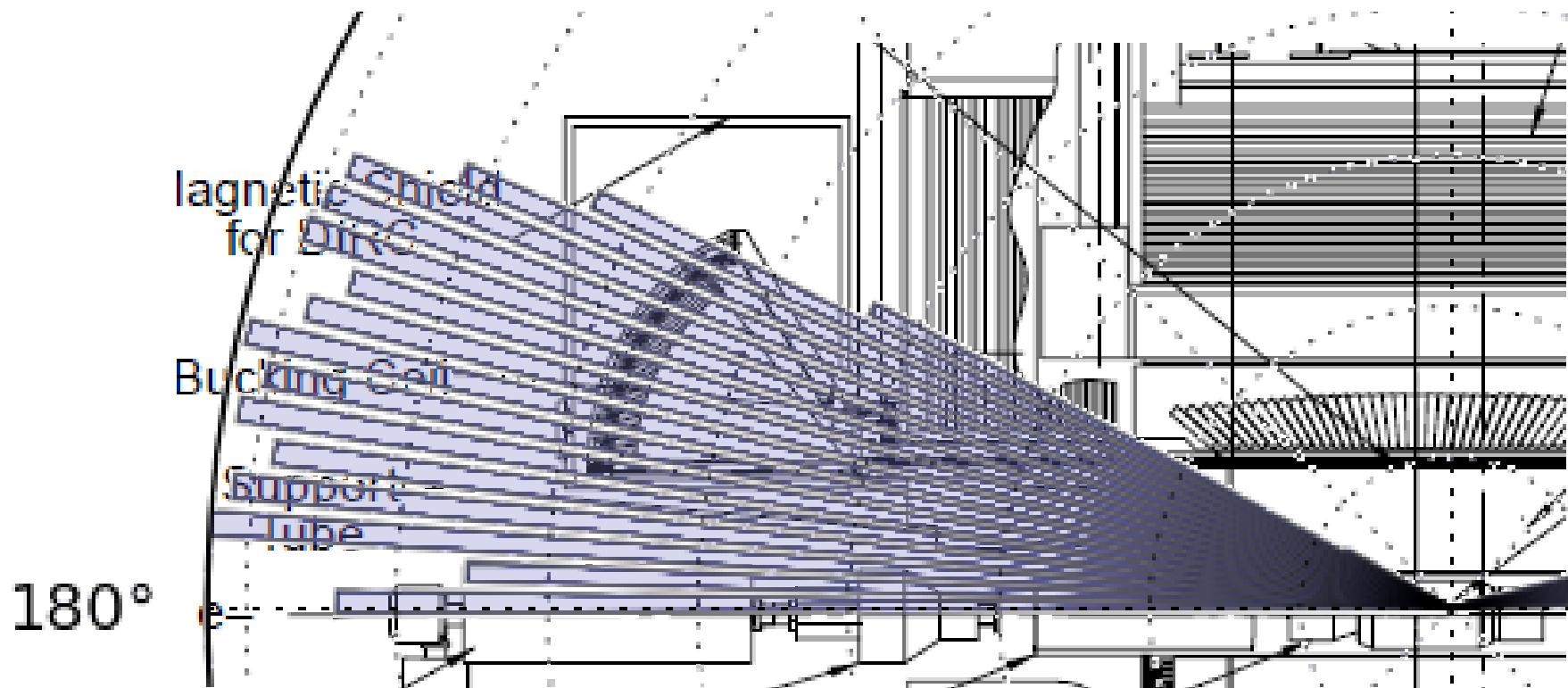


Peaking Background – A Discussion

- How can we further reduce peaking background?
 - we need additional instrumentation or handles
 - question: what can the DIRC SOB do for us?
- Is the benefit isolated to only peaking background?
 - No – QED background happens when the beam e^+ and e^- miss the detector, leaving only the $\gamma^*\gamma^*$ final state.
 - covering more of the solid angle will reject these events as well, though perhaps at a lower rate
 - beam lepton spectra peaked far forward (low scattering angle)

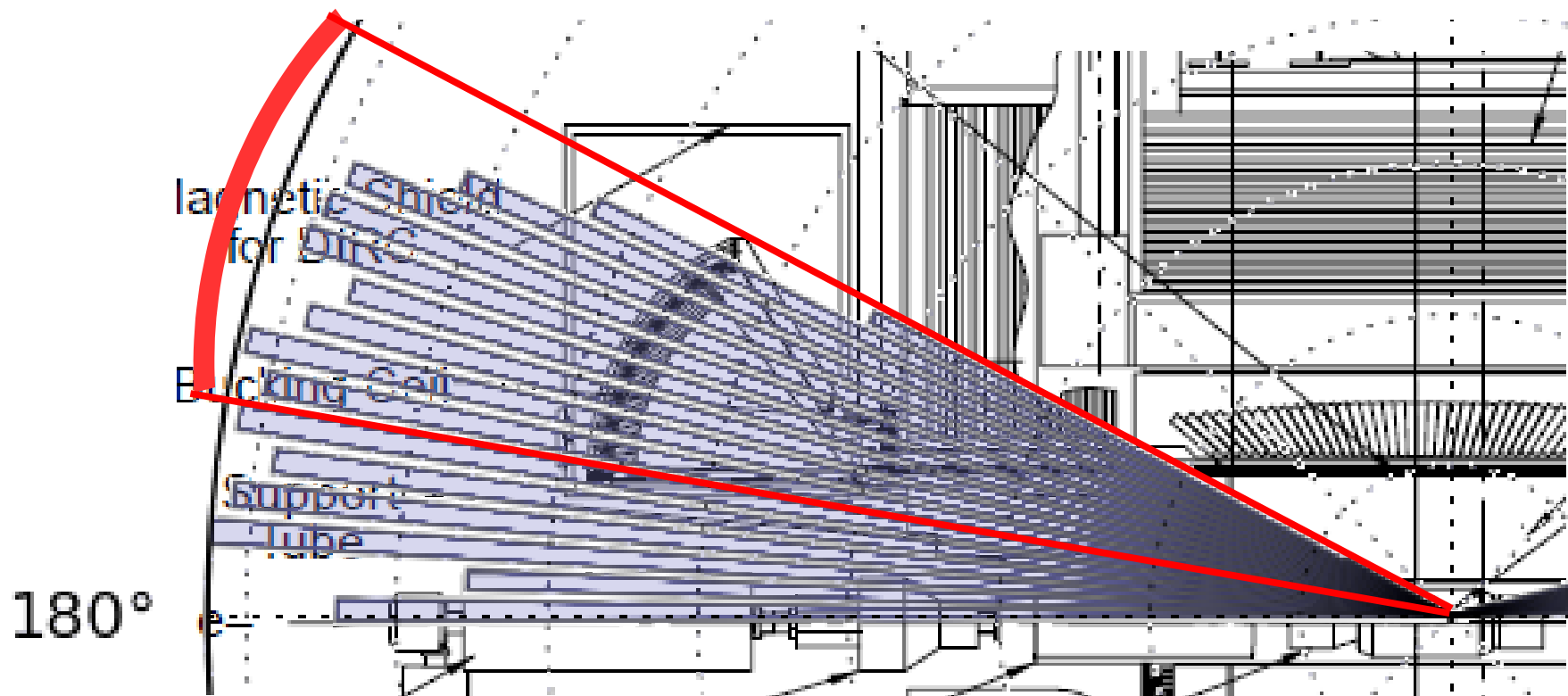
Scenario A: Active Backward Calorimetry

This means finding a way to make the backward end of the detector re-active to the passage of charged particles.



Scenario A: Active Backward Calorimetry

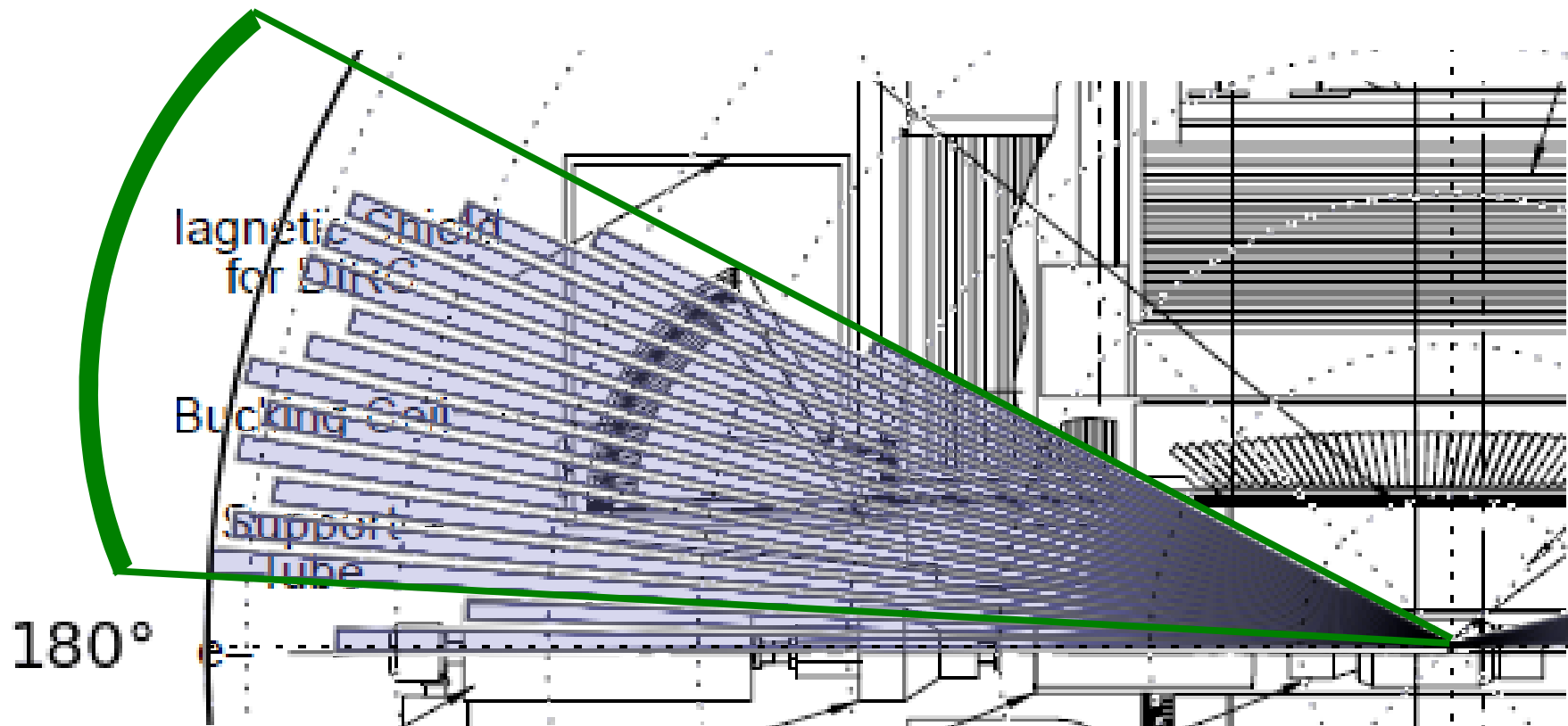
This means finding a way to make the backward end of the detector re-active to the passage of charged particles.



For this scenario, I will estimate the impact of a 90% reduction in events if either true lepton lies in the region enclosed above.

Scenario B: Close the Gap

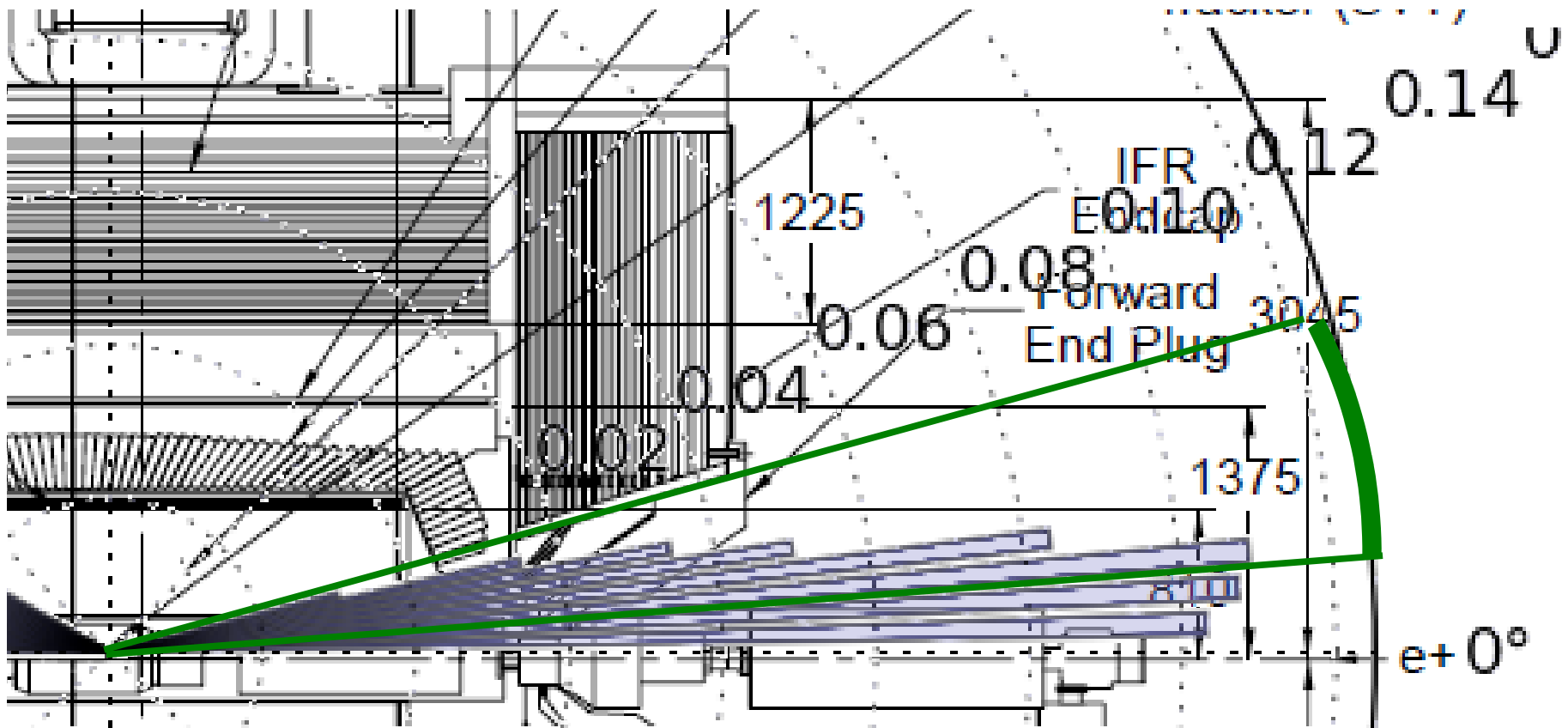
This means instrumenting both extreme ends of the detector with something – hopefully inexpensive – and store information about energy deposition for later vetoing.



For this scenario, I will estimate the impact of a 90% reduction in events if either true lepton lies in the region enclosed above.

Scenario B: Close the Gap

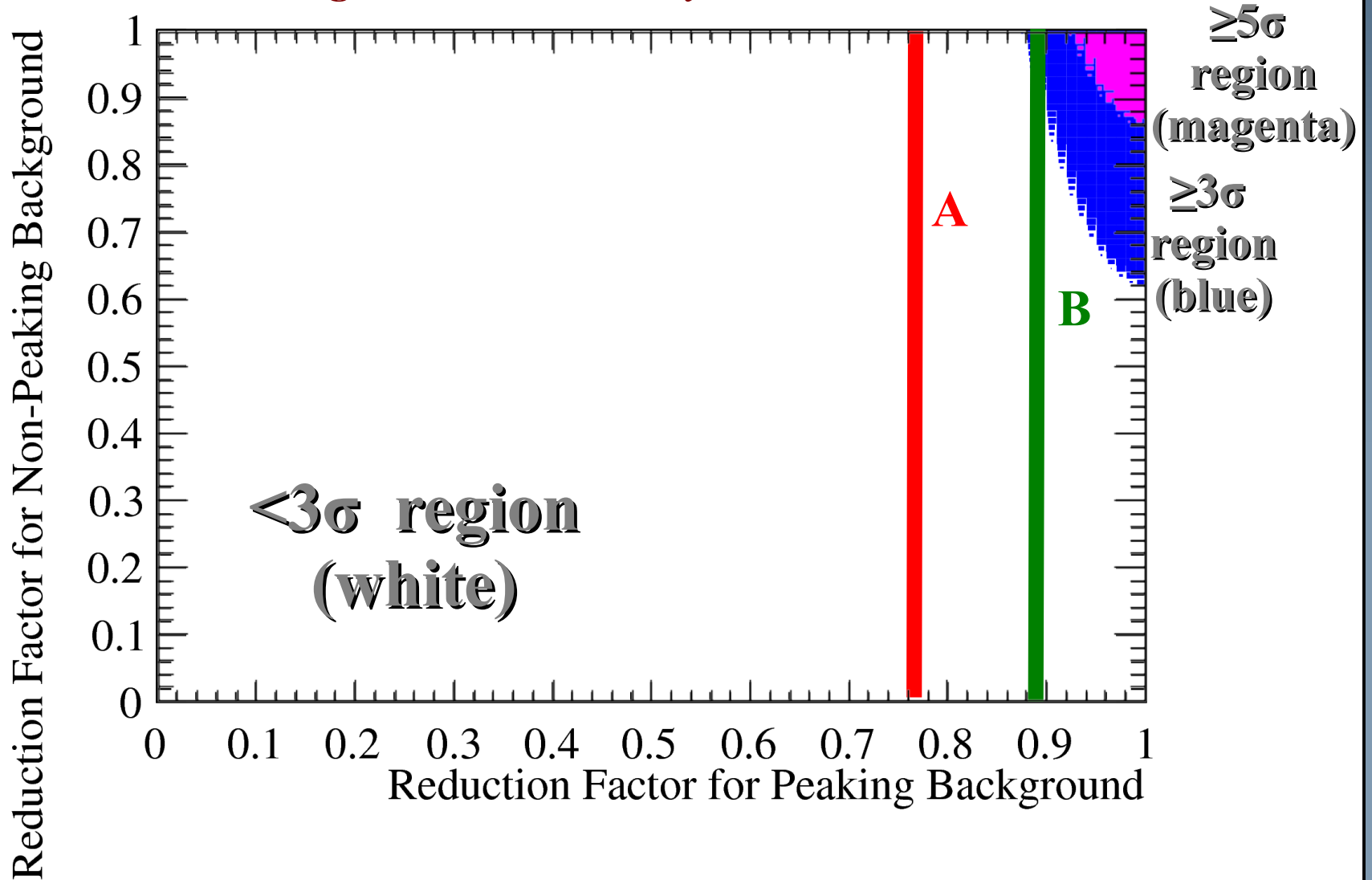
This means instrumenting both extreme ends of the detector with something – hopefully inexpensive – and store information about energy deposition for later vetoing.



For this scenario, I will estimate the impact of a 90% reduction in events if either true lepton lies in the region enclosed above.

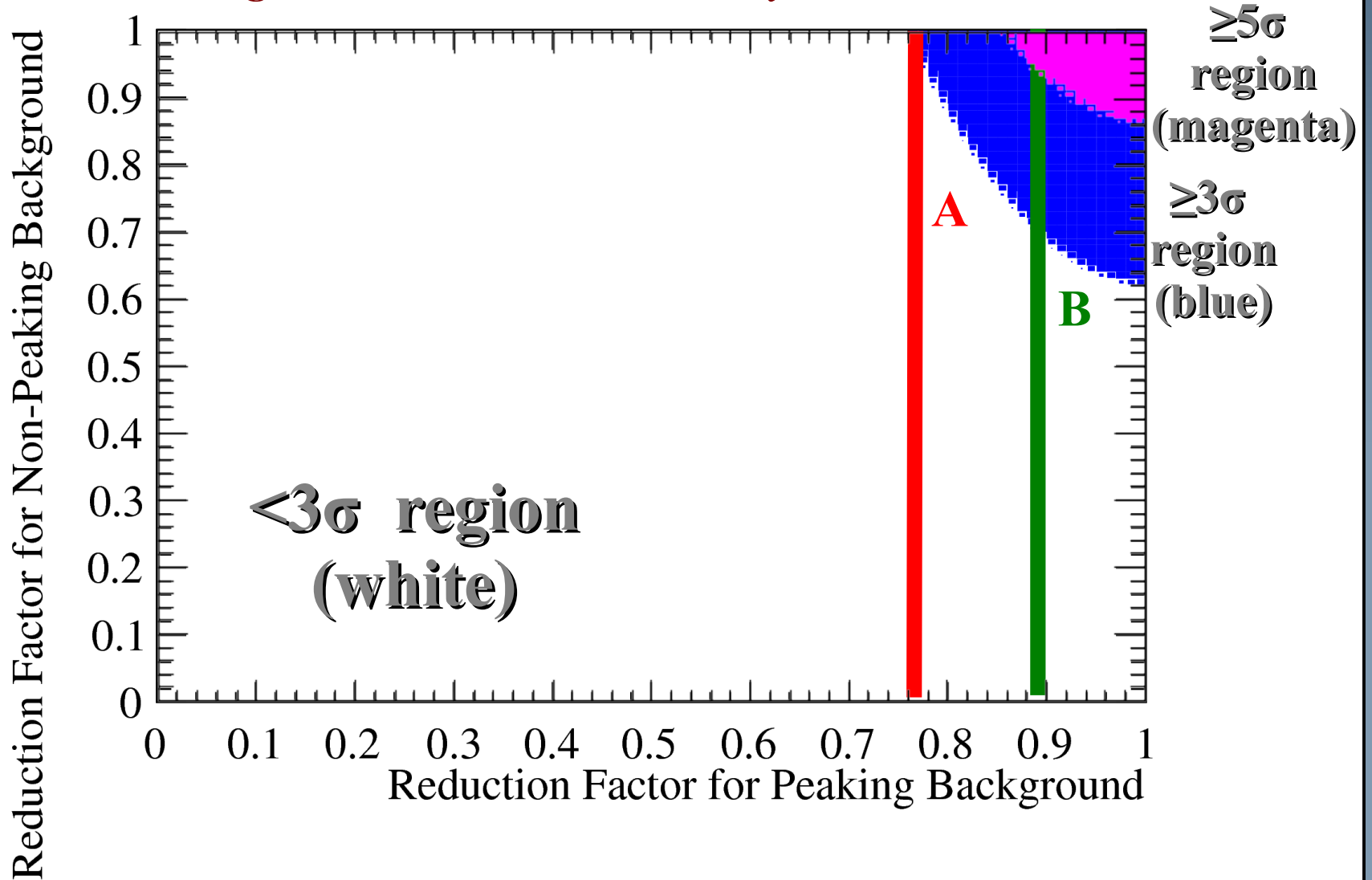
Scenarios: Impacts

Considering statistical AND systematic uncertainties



Scenarios: Impacts

Considering statistical AND reduced systematic uncertainties



Comments on the Scenarios

- Scenario A: Active backward calorimetry
 - It might be just enough to make the measurement
 - this statement assumes no other improvements
- Scenario B: Close the Gap
 - Requires a lot more work
 - Yields some flexibility in how hard to reject non-peaking background
 - might get much of that “for free” from this approach
 - Creates an opportunity for some detector R&D

Conclusions

- $Y \rightarrow \nu \bar{\nu}$ is within the grasp of Super-B
 - a test of a clean SM prediction
 - a “free” order-of-magnitude for discovery
- Evidence for this process requires work
 - reduce systematics, reduce backgrounds
 - need to think seriously about a veto in the “hole”
- Positive impact not limited to this final state
 - To the benefit of other analyses
 - $B \rightarrow \ell \nu$, $B \rightarrow K^{(*)} \nu \bar{\nu}$ – **keystone measurements** for rare decay processes, complimentary to LHC physics

Backup Slides

Comments on $Y(2S)$

- Why haven't I talked about $Y(2S)$?
 - branching fraction to $\pi^+\pi^-$ is bigger by factor of 4
 - production cross-section 2x larger than $Y(3S)$
- Difficult to trigger
 - significantly lower p_T pions are produced here
- Systematics expected to be same scale
 - expect peaking background rate to be larger
 - means bigger contribution to yield uncertainty
- This mode needs careful thought and more work

How do we improve systematics?

- Trigger systematics
 - a dream:
 - a tool that easily lets you
 - remove trigger objects from data which are related to certain particles (e.g. hard leptons)
 - recycle the modified event through the actual hardware trigger, or a close *virtual* analog
 - measure the efficiency of the real trigger on the events
 - I once mentioned this dream to a trigger expert, as relates to our BaBar analysis
 - “You should have thought of that in 1995.”
 - Well, I'm thinking of it now for Super-B, with intent to use it in 2015 or so.
- Others: MC models, etc. need study with BaBar data

References

- **Experimental results**

1. The BABAR Collaboration: B Aubert, “A Search for Invisible Decays of the Upsilon(1S),” 0908.2840 (August 19, 2009), <http://arxiv.org/abs/0908.2840>.

- **Theory**

Pierre Fayet, “Constraints on Light Dark Matter and U bosons, from psi, Upsilon, K+, pi0, eta and eta' decays,” hep-ph/0607318 (July 28, 2006), doi:doi:10.1103/PhysRevD.74.054034, <http://arxiv.org/abs/hep-ph/0607318>.

Bob McElrath, “Light Higgses and Dark Matter at Bottom and Charm Factories,” 0712.0016 (December 3, 2007), <http://arxiv.org/abs/0712.0016>.

Gagik K Yeghiyan, “ Υ Decays into Light Scalar Dark Matter,” 0909.4919 (September 27, 2009), <http://arxiv.org/abs/0909.4919>.