

SuperB Physics Workshop: Introduction

30th November – 4th December 2009 LNF Frascati

Adrian Bevan, Marco Ciuchini, Dave Brown, Achille Stocchi



The SuperB Physics Programme

Data Taking commences:

$$T_0$$
+5yr (=2015?)

The Landscape in 2015

- Tevatron's era will have come to pass and the age of the LHC will be in full swing.
- LHC will have produced many interesting results.
 - Hopefully there will be some new physics.
 - LHCb will have measured:
 - Φ_s to
 - γ to
 - ... etc.
- MEG will have placed limits down to 10⁻¹³ on $\mu \to e \gamma$ and so on.

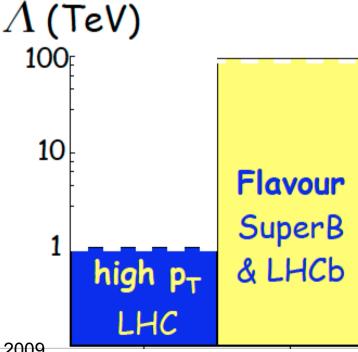


Piecing together New Physics

$$\mathcal{L}_{eff}^{NP} = \mathcal{L}_{SM} + \sum_{k} (\sum_{i} C_{i}^{k} Q_{i}^{(k+4)}) / \Lambda^{k}$$

NP flavour effects are governed by two players:

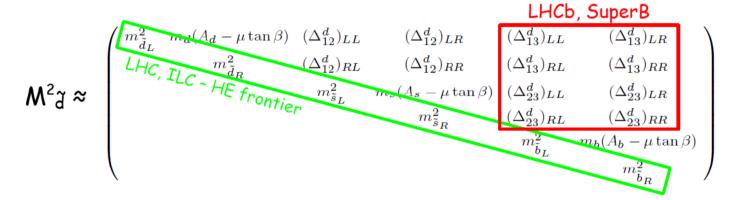
- i) the new physics scale Λ
- ii) the effective flavour-violating couplings C's
- NP could be flavour blind... but that would be fine tuning.
 - Wouldn't help us with the SUSY flavour problem.
 - unnatural given CKM and MSW.



SuperB Frascati Dec 2009



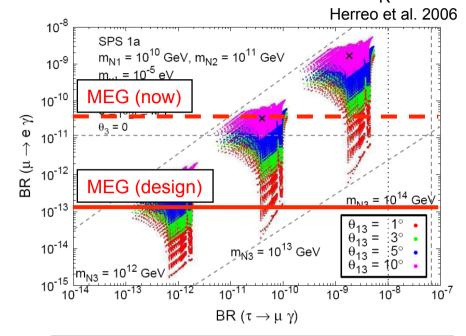
SUSY CKM



and similarly for M22



- τ→μγ upper limit can be correlated to θ₁₃ (neutrino mixing/CPV, T2K etc.) and also to μ→eγ.
 SUSY seasaw = CMSSM + 3v_R + γ
- Complementary to flavour mixing in quarks.
- Golden modes:
 - $\tau \rightarrow \mu \gamma$ and 3μ .
- e⁻ beam polarization:
 - Lower background
 - Better sensitivity than competition!
- e⁺ polarization may be used later in programme.
- CPV in $\tau \rightarrow K_S \pi v$ at the level of ~10⁻⁵.
- Bonus:
 - Can also measure τ g-2 (polarization is crucial).
 - $\sigma(g-2) \sim 2.4 \times 10^{-6}$ (statistically dominated error).

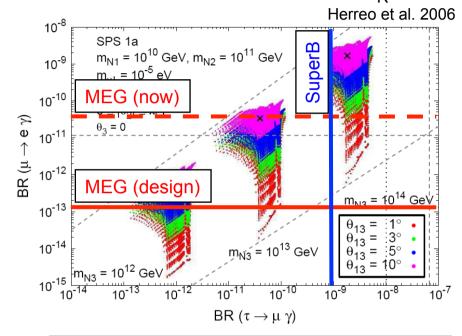


Process	Expected 90%CL	4σ Discovery
	upper limited	Reach
$\mathcal{B}(au o \mu \gamma)$	2×10^{-9}	5×10^{-9}
$\mathcal{B}(au o \mu \mu \mu)$	2×10^{-10}	8.8×10^{-10}

Use $\mu \gamma/3I$ to distinguish SUSY vs. LHT.



- $\tau \rightarrow \mu \gamma$ upper limit can be correlated to θ_{13} (neutrino mixing/CPV, T2K etc.) and also to $\mu \rightarrow e \gamma$. SUSY seasaw = CMSSM + $3v_R$ + $\overline{\gamma}$
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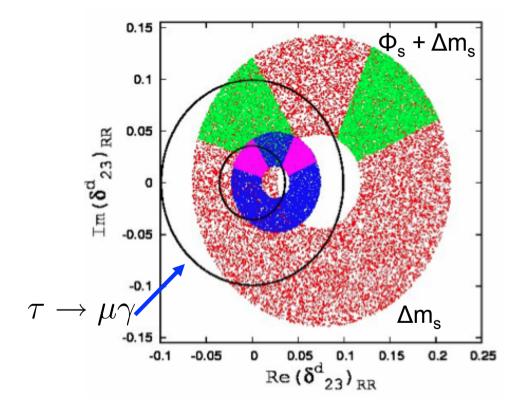


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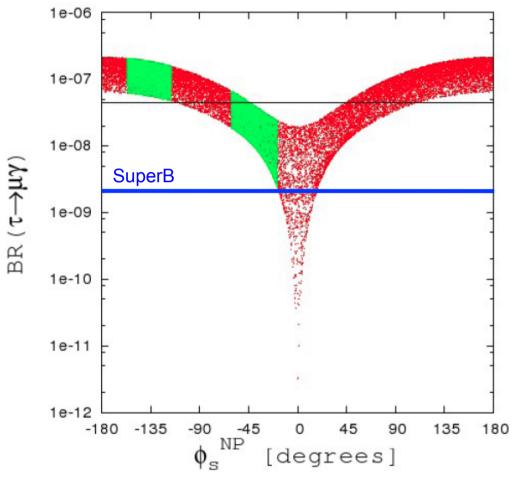
$$m_{\tilde{q}} = 300\,GeV$$
 BLUE $m_{\tilde{q}} = 500\,GeV$ RED



- SU(5) SUSY GUT Model (arXiv :0710.5443, Parry and Zhang).
- Model has non-trivial SUSY squark couplings.
- Current B_S measurements favour B($\tau \rightarrow \mu \gamma$)>3×10⁻⁹.
- Need SuperB to probe to this sensitivity.

N.B. Different New Physics Models have different features, and different hierarchies!

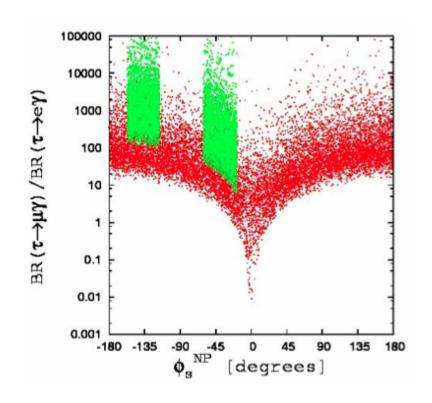




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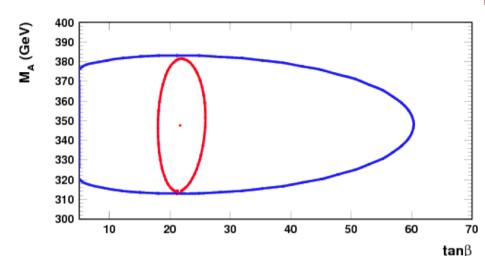


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N.B. Different New Physics Models have different features, and different hierarchies!



CMSSM: LHC/SuperB complementarity



400 M_A (GeV) 390 380 370 360 350 340 330 320 310 300 E -1000 -900 -800 -700 -600 -500 -400 -300 A (GeV)

Current analysis of data prefers $\tan \beta \sim 10$.

SuperB Frascat

Blue = LHC:

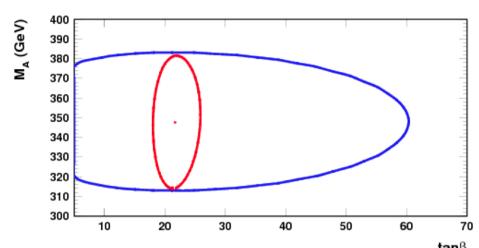
- Will be able to measure m(A) [CP odd Higgs mass]
- Poor sensitivity to tanβ [ratio of Higgs vevs]
- Poor sensitivity to A [coupling]

Red=LHC+EW/Low-energy constraints (includes SuperB):

Observable	Constraint	theo. error	
$R_{ extbf{BR}_{b ightarrow s\gamma}}$	1.127 ± 0.1	0.1	
$R_{\Delta M_s}$	0.8 ± 0.2	0.1	
$\mathrm{BR}_{b o\mu\mu}$	$(3.5\pm0.35)\times10^{-8}$	2×10^{-9}	
$R_{{ m BR}_{b ightarrow au u}}$	0.8 ± 0.2	0.1	
Δa_{μ}	$(27.6 \pm 8.4) \times 10^{-10}$	2.0×10^{-10}	
$M_W^{ m SUSY}$	$80.392 \pm 0.020{\rm GeV}$	0.020 GeV	
$\sin^2 heta_W^{ m SUSY}$	0.23153 ± 0.00016	0.00016	
$M_h^{ m light}({ m SUSY})$	> 114.4 GeV	$3.0\mathrm{GeV}$	



CMSSM: LHC/SuperB complementarity



400 M_A (GeV) 390 380 370 360 350 340 330 320 310 -300 -1000 -900 -800 -700 -600 -500 -400 A (GeV)

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Blue = LHC:

- Will be able to measure m(A) [CP odd Higgs mass]
- Poor sensitivity to tanβ [ratio of Higgs vevs]
- Poor sensitivity to A [coupling]

Red=LHC+EW/Low-energy constraints (includes SuperB):

• Can build on the m(A) measurement to measure tanβ.

Again LHC and SuperB are complementary experiments. Each can contribute significantly to the knowledge of new physics.



Charged Higgs: $B^{\pm} \rightarrow \tau^{\pm} \nu$

 B^{-}



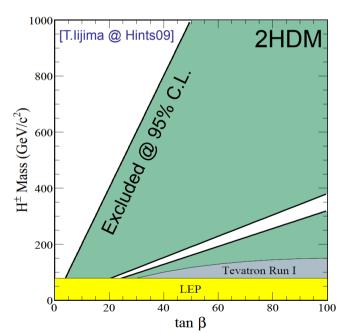
- Within the SM, sensitive to f_B and $|V_{ub}|$: $\mathcal{B}_{SM} \sim 1.6 \times 10^{-4}$.
- lacksquare ${\mathcal B}$ affected by new physics.
 - MFV models like 2HDM / MSSM.
 - Unparticles.

MSSM.
$$u \nearrow v_{\tau}$$

$$\mathcal{B}_{SM}(B^+ \to l^+ v_l) = \frac{G_F^2 m_B m_l^2}{8\pi} \left(1 - \frac{m_l^2}{m_B^2}\right) f_B^2 |V_{ub}|^2 \tau_B$$

• Fully reconstruct the event (modulo v).

$$\mathcal{B}_{WA} = (1.73 \pm 0.35) \times 10^{-4}$$
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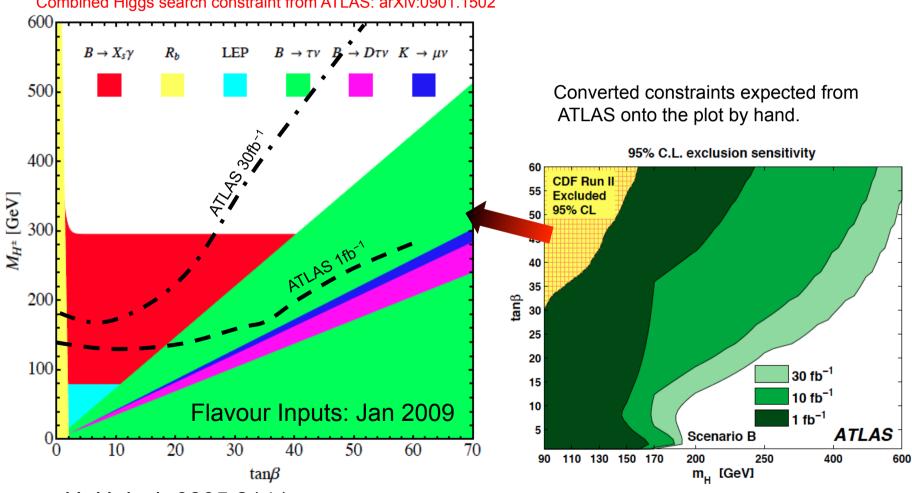
2HDM: W.-S Hou PRD **48** 2342 (1993) MSSM: G. Isidori arXiv:0710.5377 Unparticles: R. Zwicky PRD**77** 036004 (2008)



Charged Higgs

B-factory searches competitive with LHC era: e.g. 2HDM

Existing Constraints from BaBar and Belle.
Combined Higgs search constraint from ATLAS: arXiv:0901.1502



U. Haisch 0805.2141

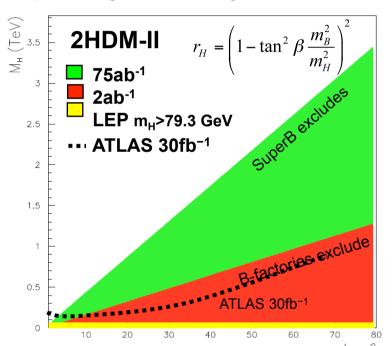


Charged Higgs

Higgs mediated MFV:

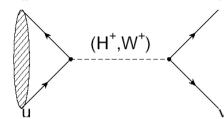
$$r_H = \frac{\mathcal{B}_{SM+NP}}{\mathcal{B}_{SM}}$$

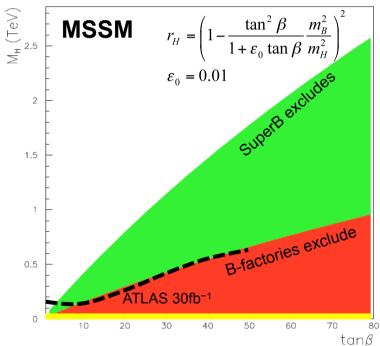
(Assuming SM branching fraction is measured)





ATLAS sensitivity sketched from combined sensitivity plots in arXiv:0901.0512. b



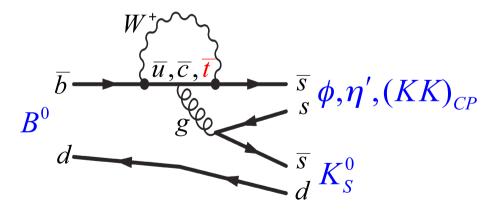


- Multi TeV search capability for large tanβ.
- Includes SM uncertainty ~20% from V_{ub} and f_B.

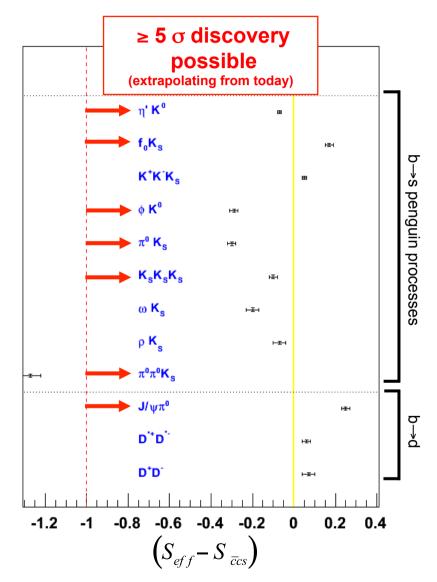


∆S measurements

- β=(21.1±0.9)° from Charmonium decays.
- Look in many different b→s and b→d decays for sin2β deviations from the SM:
- The golden channel is:



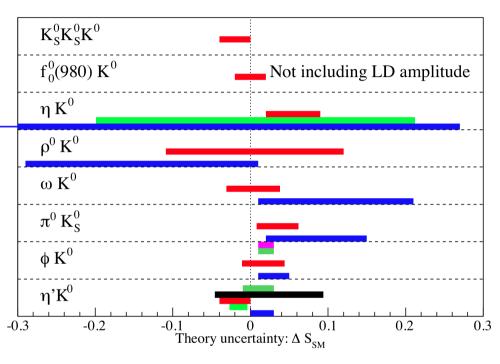
 Deviations would be from high mass particles in loops: H, χ, ...

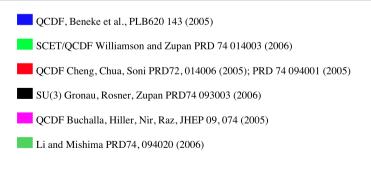




∆S measurements

- The SM uncertainty is strongly mode dependent.
- Golden modes have to be well measured and theoretically clean.
- Prefer to also have robust constraints from more than one theoretical approach.
- Precision measurements of the reference Charmonium decay also have a small SM uncertainty.







∆S measurements

- We were reminded that we should be careful with what we compare:
 - NP could affect cc̄s sin2β.
- 1) Predict $\sin 2\beta$ from indirect constraints.

$$\left[\sin(2\beta)\right]_{\text{no }V_{ub}}^{\text{prediction}} = 0.87 \pm 0.09$$
 .

2) Compare to ccs measurement.

$$[\sin 2\beta]_{c\bar{c}s} = 0.672 \pm 0.023$$



3) Compare to clean penguin measurements.

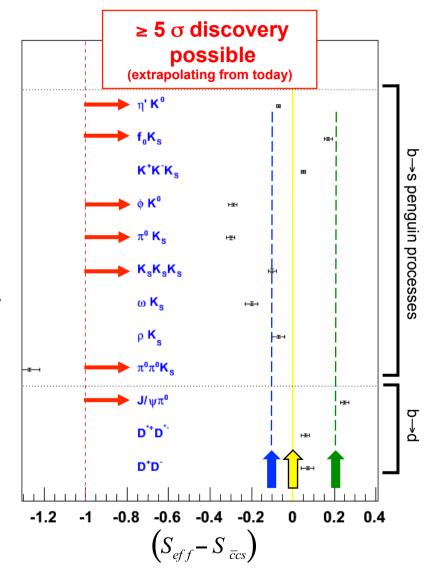
$$[\sin 2\beta]_{b \to s-p \, enguin}^{clean} = 0.58 \pm 0.06$$

(or the average of the two)

Are these 2.1-2.7 σ hints for new physics?

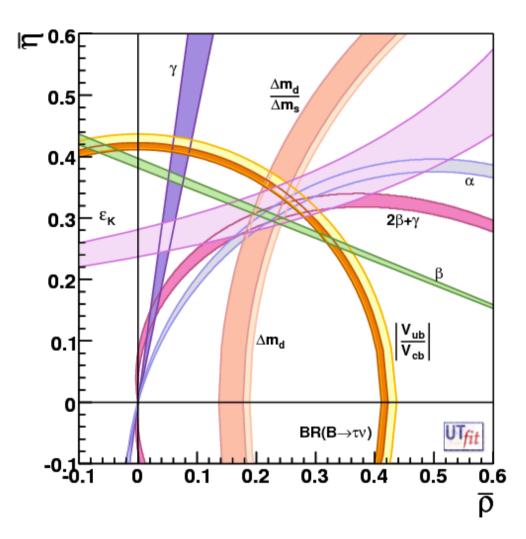
Lunghi and Soni, Phys.Lett.B**666** 162-165 (2008). Buras and Guadagnoli Phys Rev D **78** 033005 (2008).

Can theory error be reduced for other modes?





... and precision CKM



$$\sigma(\beta) < 0.5^{\circ}$$

$$\sigma(\alpha) \sim 1 - 2^{\circ}$$

$$\sigma(\gamma) \sim 1 - 2^{\circ}$$

$$etc.$$

Can we develop more insight into CKM?

- does it work at the % level?
- what new physics probes become available with precision CKM?

[for SuperB and Other expt.]



Standard Model measurements.

Observable	B Factories (2 ab^{-1})	Super B (75 ab ⁻¹)		
$\sin(2\beta) (J/\psi K^0)$	0.018	0.005 (†)		
$cos(2\beta) (J/\psi K^{*0})$	0.30	0.05		
$\sin(2\beta) (Dh^0)$	0.10	0.02		
$cos(2\beta)$ (Dh^0)	0.20	0.04		
$S(J/\psi \pi^0)$	0.10	0.02		
$S(D^{+}D^{-})$	0.20	0.03		
$S(\phi K^0)$	0.13	0.02 (*)		
$S(\eta'K^0)$	0.05	0.01 (*)		
$S(K_s^0 K_s^0 K_s^0)$	0.15	0.02 (*)		
$S(K_s^0\pi^0)$	0.15	0.02 (*)		
$S(\omega K_s^0)$	0.17 0.12	0.03 (*)		
$S(f_0K_s^0)$	0.12	0.02 (*)		
γ (B \rightarrow DK, D \rightarrow CP eigenstate	∞s) ~ 15°	2.5°		
γ (B \rightarrow DK, D \rightarrow suppressed st		2.0°		
γ (B \rightarrow DK, D \rightarrow multibody sta		1.5°		
γ (B \rightarrow DK, combined)	~ 6°	1–2°		
$\alpha (B \rightarrow \pi \pi)$	$\sim 16^{\circ}$	3°		
$\alpha (B \rightarrow \rho \rho)$	$\sim 7^{\circ}$	1-2° (*)		
$\alpha \ (B \rightarrow \rho \pi)$	$\sim 12^{\circ}$	2°		
α (combined)	$\sim 6^{\circ}$	1-2° (*)		
$2\beta + \gamma \ (D^{(*)\pm}\pi^{\mp}, \ D^{\pm}K_{S}^{0}\pi^{\mp})$	20°	5°		
$ V_{cb} $ (exclusive)	4% (*)	1.0% (*)		
$ V_{cb} $ (inclusive)	1% (*)	0.5% (*)		
$ V_{ub} $ (exclusive)	8% (*)	3.0% (*)		
$ V_{ub} $ (inclusive)	8% (*)	2.0% (*)		
$B(B \rightarrow \tau \nu)$	20%	4% (†)		
$\mathcal{B}(B \to \mu \nu)$	visible	5%		
$\mathcal{B}(B \to D\tau\nu)$	10%	2%		
S(B · B(v)	1070	270		
$B(B \rightarrow \rho \gamma)$	15%	3% (†)		
		5%		
$\mathcal{B}(B \to \omega \gamma)$	30%			
$A_{CP}(B \rightarrow K^* \gamma)$	0.007 (†)	0.004 († *)		
$A_{CP}(B \rightarrow \rho \gamma)$	~ 0.20	0.05		
$A_{CP}(b \rightarrow s\gamma)$	0.012 (†)	0.004 (†)		
$A_{CP}(b \rightarrow (s + d)\gamma)$	0.03	0.006 (†)		
$S(K_S^0\pi^0\gamma)$	0.15	0.02 (*)		
$S(\rho^0\gamma)$	possible	0.10		
. "	-			
$A_{CP}(B \rightarrow K^*\ell\ell)$	7%	1%		
$A^{FB}(B \rightarrow K^*\ell\ell)s_0$	25%	9%		
$A^{FB}(B \rightarrow X_s \ell \ell) s_0$ $A^{FB}(B \rightarrow X_s \ell \ell) s_0$	35%	9% 5%		
$\mathcal{B}(B \to K \nu \overline{\nu})$	visible	20%		
$B(B \rightarrow \pi \nu \bar{\nu})$	_	possible		

B Physics at Y(4S) Rare Charm Decays: 1 month at $\psi(3770)$

Channel	Sensitivity
$D^0 \to e^+ e^-, D^0 \to \mu^+ \mu^-$	1×10^{-8}
$D^0 \to \pi^0 e^+ e^-, D^0 \to \pi^0 \mu^+ \mu^-$	2×10^{-8}
$D^0 \to \eta e^+ e^-, D^0 \to \eta \mu^+ \mu^-$	3×10^{-8}
$D^0 \to K_s^0 e^+ e^-, D^0 \to K_s^0 \mu^+ \mu^-$	3×10^{-8}
$D^+ \to \pi^+ e^+ e^-, D^+ \to \pi^+ \mu^+ \mu^-$	1×10^{-8}
$D^{0} \rightarrow e^{\pm} \mu^{\mp}$	1×10^{-8}
$D^+ \to \pi^+ e^{\pm} \mu^{\mp}$	1×10^{-8}
$D^0 \to \pi^0 e^{\pm} \mu^{\mp}$	2×10^{-8}
$D^0 \to \eta e^{\pm} \mu^{\mp}$	3×10^{-8}
$D^0 \to K_s^0 e^{\pm} \mu^{\mp}$	3×10^{-8}
$D^+ \to \pi^- e^+ e^+, D^+ \to K^- e^+ e^+$	1×10^{-8}
$D^+ \to \pi^- \mu^+ \mu^+, D^+ \to K^- \mu^+ \mu^+$	1×10^{-8}
$D^+ \to \pi^- e^{\pm} \mu^{\mp}, D^+ \to K^- e^{\pm} \mu^{\mp}$	1×10^{-8}

Mode Observable $\Upsilon(4S)$ $\psi(3770)$ Charm Mixing (75 ab^{-1}) (300 fb^{-1}) $D^0 \rightarrow K^+\pi^ 3 \times 10^{-5}$ 7×10^{-4} $D^0 \rightarrow K^+K^ 5 \times 10^{-4}$ $D^0 \rightarrow K_S^0 \pi^+ \pi^ 4.9 \times 10^{-4}$ 3.5×10^{-4} 3×10^{-2} $\psi(3770) \rightarrow D^0 \overline{D}^0$ $(1-2) \times 10^{-5}$ $(1-2) \times 10^{-3}$ $\cos \delta$ (0.01 - 0.02)

See CDR and Valencia report for details of the SM measurements and other possible NP searches.

τ: LFV / CPV / ...

Process	Sensitivity
$\mathcal{B}(\tau \to \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \to e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \to \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \to eee)$	2×10^{-10}
$\mathcal{B}(\tau \to \mu \eta)$	4×10^{-10}
$\mathcal{B}(\tau \to e \eta)$	6×10^{-10}
$\mathcal{B}(\tau \to \ell K_s^0)$	2×10^{-10}

B Physics at Y(5S)

Observable	Error with 1 ab^{-1}	Error with 30 ab^{-1}
ΔΓ	$0.16 \ \mathrm{ps^{-1}}$	$0.03~{\rm ps^{-1}}$
Γ	$0.07 \ \mathrm{ps^{-1}}$	$0.01 \ \mathrm{ps^{-1}}$
β_s from angular analysis	20°	8°
$A^s_{ m SL}$	0.006	0.004
A_{CH}	0.004	0.004
$\mathcal{B}(B_s \to \mu^+ \mu^-)$	-	$<8\times10^{-9}$
$ V_{td}/V_{ts} $	0.08	0.017
$\mathcal{B}(B_s \to \gamma \gamma)$	38%	7%
β_s from $J/\psi\phi$	10°	3°
$\beta_s \text{ from } B_s \to K^0 \bar{K}^0$	24°	11°



Other measurements of interest

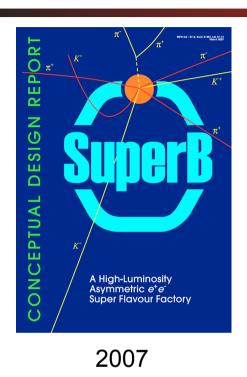
- Let us not forget the other opportunities for searching for new physics at SuperB:
 - Charm decays: CPV and parallels of rare B decays.
 - LFV in light mesons
 - Light (few GeV) Higgs Searches
 - CPT Violation test / EPR correlation tests
 - Dark matter ... and Dark Forces

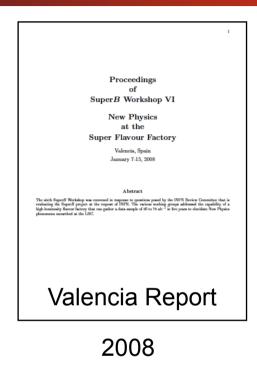


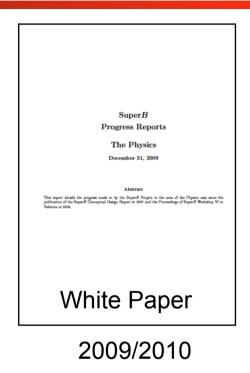
Recent (and imminent) milestones



Milestones in the programme...







- The white paper is a stepping stone toward the TDR (end of 2010).
- During this workshop we aim to review the state of the art in terms of flavour physics capabilities of SuperB.



Recent Progress

- A lot has happened since Warwick:
 - Dark Forces has gathered a lot of interest in the community (See F. Bossi).
 - First results from MEG (See F. Renga).
 - LHC data taking has re-started (See T. Hurth).
 - Full and Fast Sim. and a SuperB distributed computing model has come online (See the General Meeting).
 - We are now starting to optimize SuperB based on a few key benchmark channels.
- All of these factors are additional inputs to refine our goals.
- This work will continue through 2010 as we re-focus for the TDR.



Physics White Paper



Purpose

- Provide a coherent description of the physics case of the experiment:
 - New Physics Search Capability:
 - This is the reason why SuperB is relevant in 2015 (when we plan to go online).
 - What other inputs are necessary to decode the behaviour of new physics more comprehensively?
 - How the SuperB measurements push back our understanding of the viable models?
 - Standard Model Physics:
 - Precision CKM: as a test of the CKM anzatz.
 - As a control of SM uncertainties for other (non-)SuperB benchmark channels to search for new physics.



Scope

- Aim: 40-50 pages:
 - Have made a good start with the CDR & Valencia document.
 - Several key features have changed:
 - D mixing has been established, and CPV searches in charm are more mature.
 - NP probes in B and τ decays have advanced since last time.
 - Lattice QCD has matured: we can compare against CDR predictions.
 - Tevatron has measured Φ_s more precisely.
 - This document should be stand alone.



Content

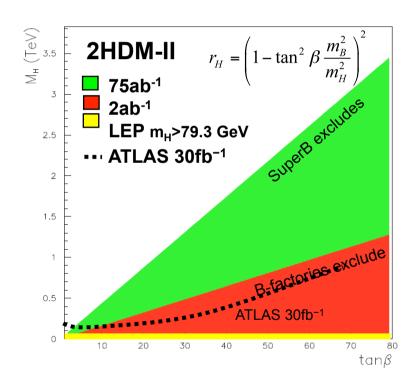
- Introduction
- B_u,d,s physics
- Charm physics
- т physics
- Spectroscopy
- Role of Lattice QCD
- Interplay with other experiments
- Conclusion

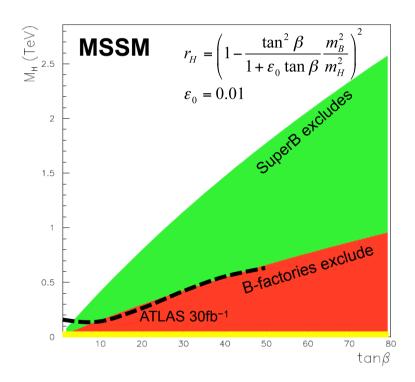
The content is based on previous documents. If you think that a new section is needed – then propose it to the conveners in charge of that section.



Why have an interplay section?

 This is how we can show what SuperB will teach us about nature; in the context of existing and planned experimental programmes.

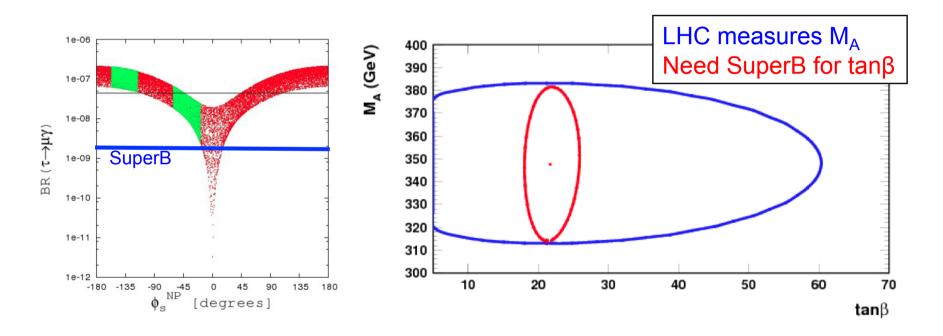






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Why have an interplay section?

 This is how we can show what SuperB will teach us about nature; in the context of existing and planned experimental programmes.

	H^+	MFV	Non-MFV	NP	Right-handed	LTH S	SUSY
	high $\tan \beta$			Z -penguins	currents		
$\mathcal{B}(B \to X_s \gamma)$		\mathbf{L}	M		M		
$\mathcal{A}_{CP}(B o X_s \gamma)$			\mathbf{L}		\mathbf{M}		
$\mathcal{B}(B \to \tau \nu)$	L-CKM						
$\mathcal{B}(B o X_s \ell \ell)$			\mathbf{M}	M	\mathbf{M}		
$\mathcal{B}(B \to K \nu \overline{\nu})$			\mathbf{M}	\mathbf{L}			
$S_{K_S\pi^0\gamma}$					L		
The angle β (ΔS)			L-CKM		L		
$ au o \mu \gamma$							L
$ au ightarrow \mu \mu \mu$						\mathbf{L}	

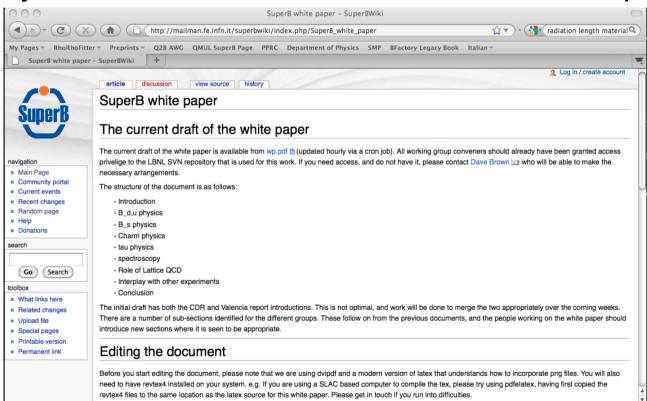
Need to elaborate on the golden matrix of observables that can be used to decode new physics.

This is more than just rare B and τ decays... See the talks later this morning!



Where can I find the Tex source?

- The source is in an SVN repository at LBNL.
- Instructions for accessing this are on the Physics Wiki Page.
 - Physics Portal -> Documentation -> White Paper





How do I get to see the current version?

- A trscron job updates the pdf version of the white paper every hour.
 - Can download a current copy that corresponds to whatever has been committed.
 - Will always be up-to-date within 60min (or broken)

http://www.slac.stanford.edu/~bevan/superbID/wp.pdf

 ${f Super} B$ Progress Reports

The Physics

December 31, 2009

Abstract

This report details the progress made in by the Super Project in the area of the Physics case since the publication of the Super Propertial Design Report in 2007 and the Proceedings of Super Workshop VI in Valencia in 2008.



How do I get an account to contribute?

- This work uses the original LBNL SuperB repository.
- Working group conveners and theorists have been given accounts to access this repository.
- If you have not already been informed of this and would like to access the Tex to contribute to the white paper, please see Dave Brown: Dave_Brown@lbl.gov.



Timescale

- We want to have the white paper finalized as soon as possible.
 - Get as much work done as possible this week.
 - Follow up with this work through the Christmas vacation.
 - Finish the document early next year at the latest: end Dec/Start Jan.
- We can come back to the white paper at the end of the morning and take discussions through lunch!



Summary

- Plenary talks this morning:
 - Dark Forces (F. Bossi)
 - MEG (F. Renga)
 - Interplay (T. Hurth)
- Parallel sessions this afternoon.
- Lattice plenary talk on Tuesday 11.14 (V. Lubicz).
- Parallel sessions on the remaining afternoons.