



# SuperB Physics Workshop: Introduction

30<sup>th</sup> November – 4<sup>th</sup> December 2009  
LNF Frascati

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# The SuperB Physics Programme

Data Taking commences:

$T_0 + 5\text{yr}$  (=2015?)



# The Landscape in 2015

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- Tevatron's era will have come to pass and the age of the LHC will be in full swing.
- LHC will have produced many interesting results.
  - Hopefully there will be some new physics.
  - LHCb will have measured:
    - $\Phi_s$  to
    - $\gamma$  to
    - ... etc.
- MEG will have placed limits down to  $10^{-13}$  on  $\mu \rightarrow e\gamma$   
..... and so on.

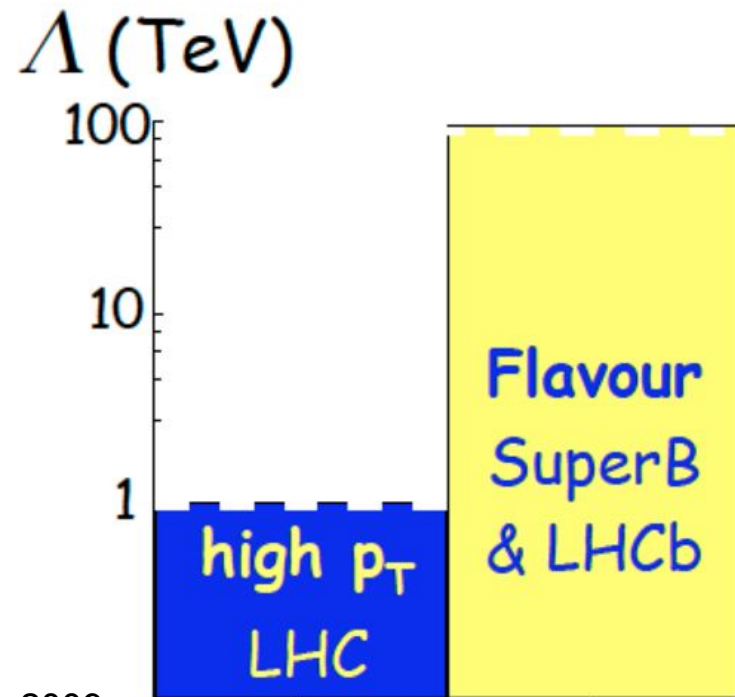
# Piecing together New Physics

$$\mathcal{L}_{\text{eff}}^{\text{NP}} = \mathcal{L}_{\text{SM}} + \sum_k \left( \sum_i C_i^k Q_i^{(k+4)} \right) / \Lambda^k$$

NP flavour effects are governed by two players:

- i) the new physics scale  $\Lambda$
- ii) the effective flavour-violating couplings  $C$ 's

- NP could be flavour blind... but that would be fine tuning.
  - Wouldn't help us with the SUSY flavour problem.
  - unnatural given CKM and MSW.



# SUSY CKM

$$M_{\mathcal{D}}^2 \approx \left( \begin{array}{cccccc}
 m_{dL}^2 & m_d(A_d - \mu \tan \beta) & (\Delta_{12}^d)_{LL} & (\Delta_{12}^d)_{LR} & (\Delta_{13}^d)_{LL} & (\Delta_{13}^d)_{LR} \\
 & m_{dR}^2 & (\Delta_{12}^d)_{RL} & (\Delta_{12}^d)_{RR} & (\Delta_{13}^d)_{RL} & (\Delta_{13}^d)_{RR} \\
 & & m_{sL}^2 & m_s(A_s - \mu \tan \beta) & (\Delta_{23}^d)_{LL} & (\Delta_{23}^d)_{LR} \\
 & & & m_{sR}^2 & (\Delta_{23}^d)_{RL} & (\Delta_{23}^d)_{RR} \\
 & & & & m_{bL}^2 & m_b(A_b - \mu \tan \beta) \\
 & & & & & m_{bR}^2
 \end{array} \right)$$

LHC, ILC - HE frontier
LHCb, SuperB

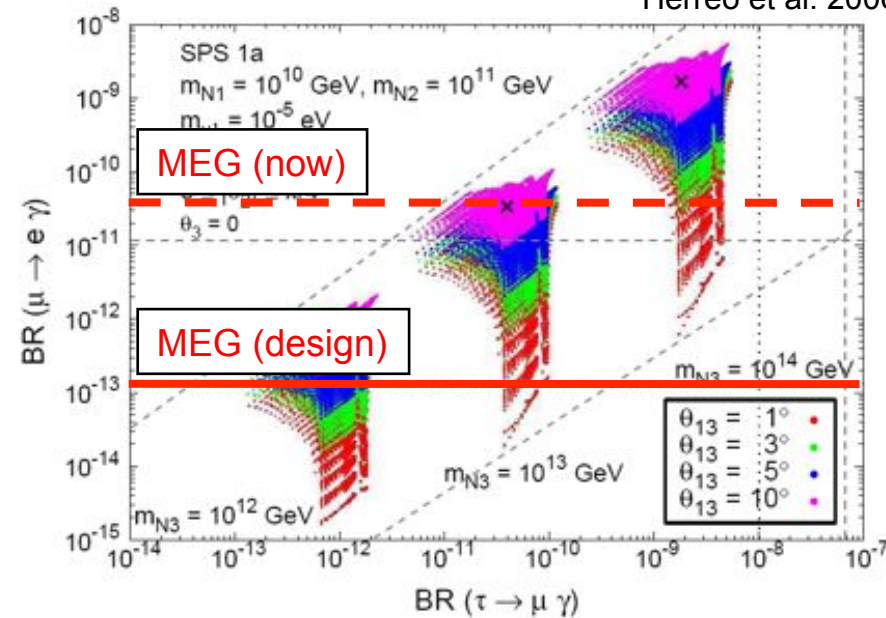
and similarly for  $M_{\mathcal{U}}^2$

# Lepton Flavour Violation ( $\tau$ decay)

- $\tau \rightarrow \mu \gamma$  upper limit can be correlated to  $\theta_{13}$  (neutrino mixing/CPV, T2K etc.) and also to  $\mu \rightarrow e \gamma$ .
- Complementary to flavour mixing in quarks.
- Golden modes:
  - $\tau \rightarrow \mu \gamma$  and  $3\mu$ .
- $e^-$  beam polarization:
  - Lower background
  - Better sensitivity than competition!
- $e^+$  polarization may be used later in programme.
- CPV in  $\tau \rightarrow K_S \pi \nu$  at the level of  $\sim 10^{-5}$ .
- Bonus:
  - Can also measure  $\tau$  g-2 (polarization is crucial).
  - $\sigma(g-2) \sim 2.4 \times 10^{-6}$  (statistically dominated error).

SUSY seasaw = CMSSM +  $3\nu_R + \tilde{\nu}$

Herreo et al. 2006



Process	Expected 90%CL upper limited	4 $\sigma$ Discovery Reach
$B(\tau \rightarrow \mu \gamma)$	$2 \times 10^{-9}$	$5 \times 10^{-9}$
$B(\tau \rightarrow \mu \mu \mu)$	$2 \times 10^{-10}$	$8.8 \times 10^{-10}$

Use  $\mu \gamma/3l$  to distinguish SUSY vs. LHT.

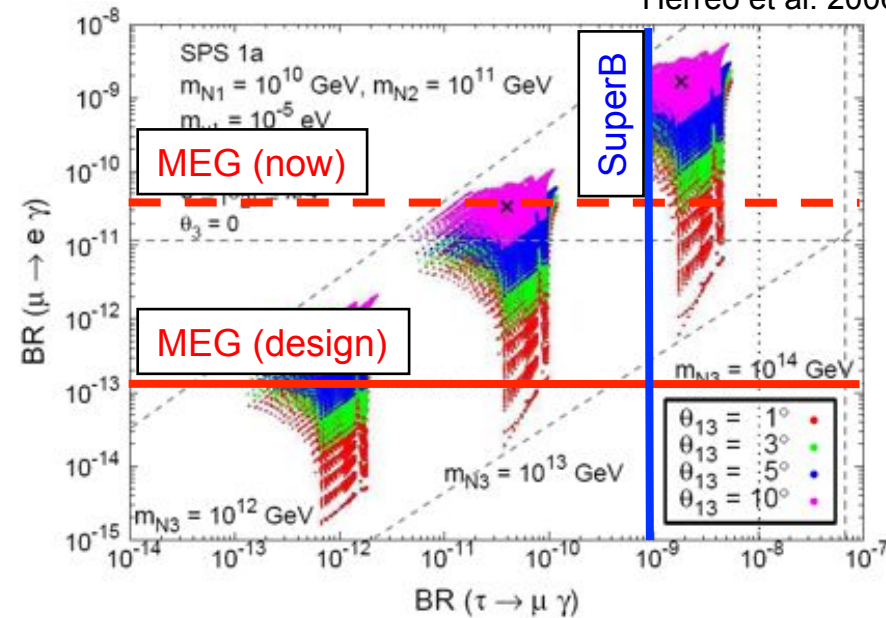


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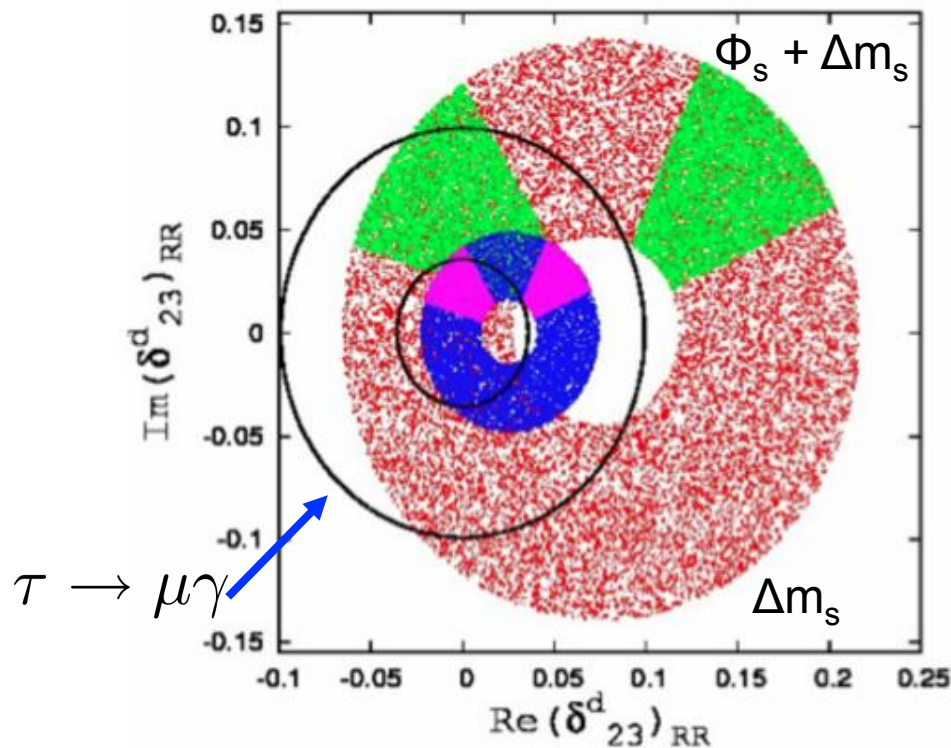
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# Lepton Flavour Violation ( $\tau$ decay)

$m_{\tilde{q}} = 300 \text{ GeV}$     **BLUE**

$m_{\tilde{q}} = 500 \text{ GeV}$     **RED**



- SU(5) SUSY GUT Model (arXiv :0710.5443, Parry and Zhang).

- Model has non-trivial SUSY squark couplings.

- Current  $B_s$  measurements favour  $B(\tau \rightarrow \mu\gamma) > 3 \times 10^{-9}$ .

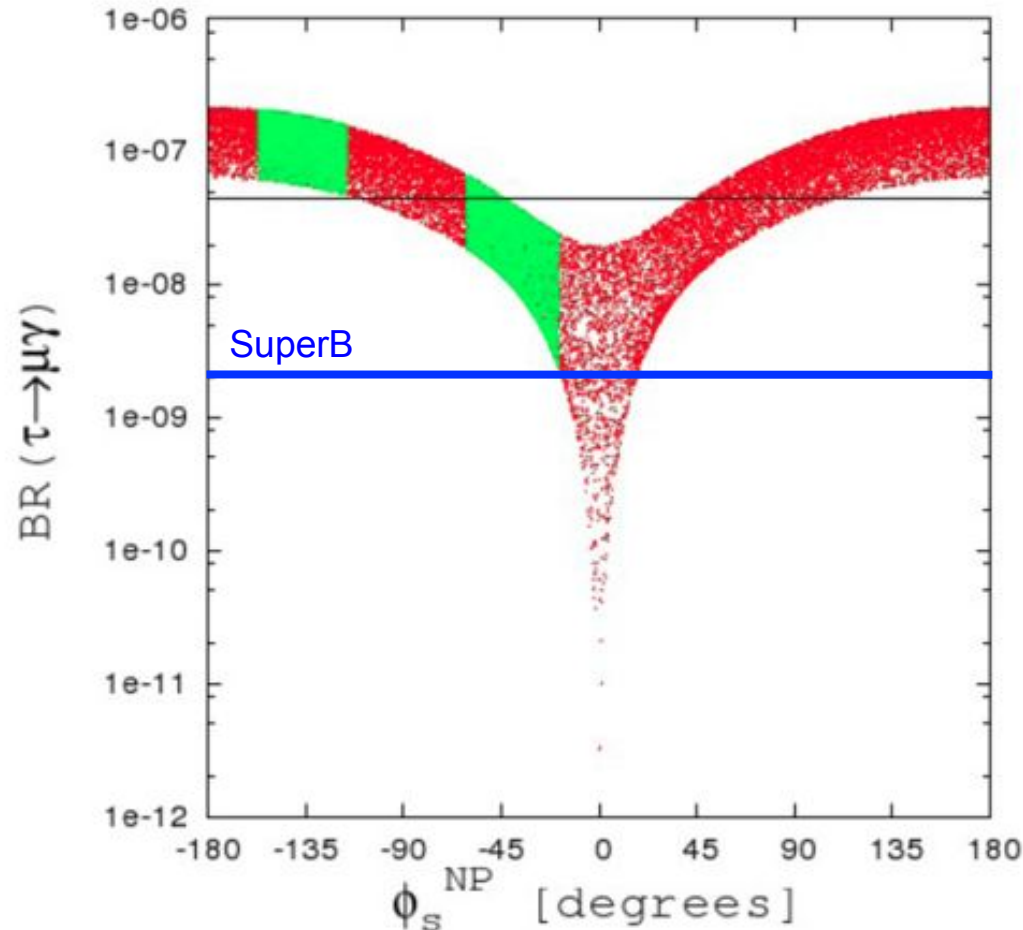
- Need SuperB to probe to this sensitivity.

**N.B. Different New Physics Models have different features, and different hierarchies!**





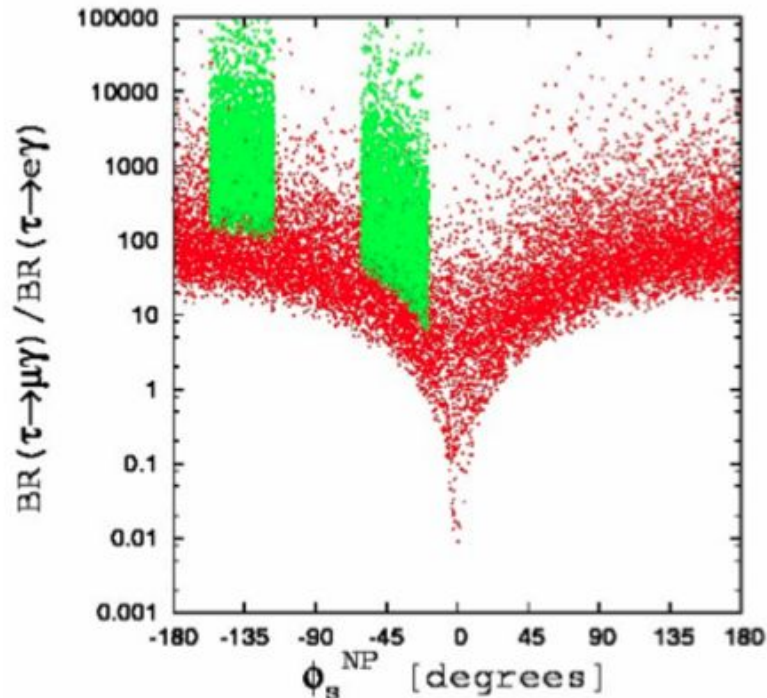
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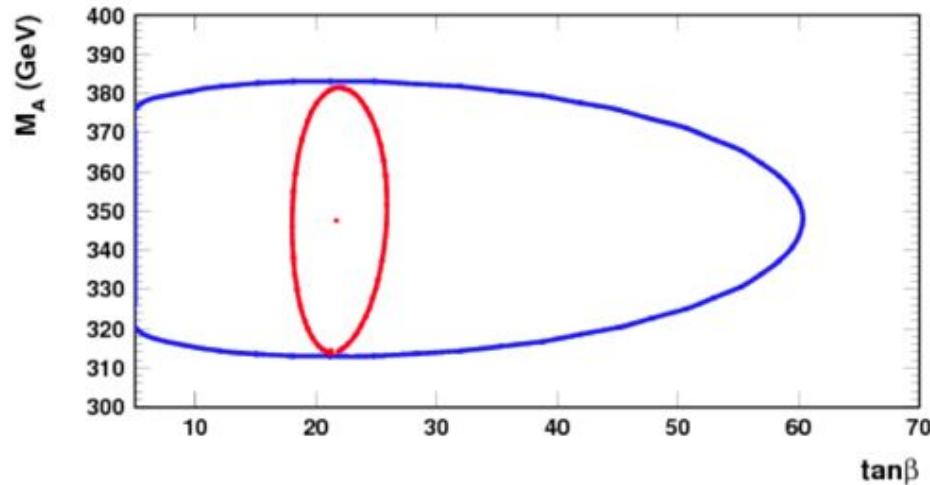


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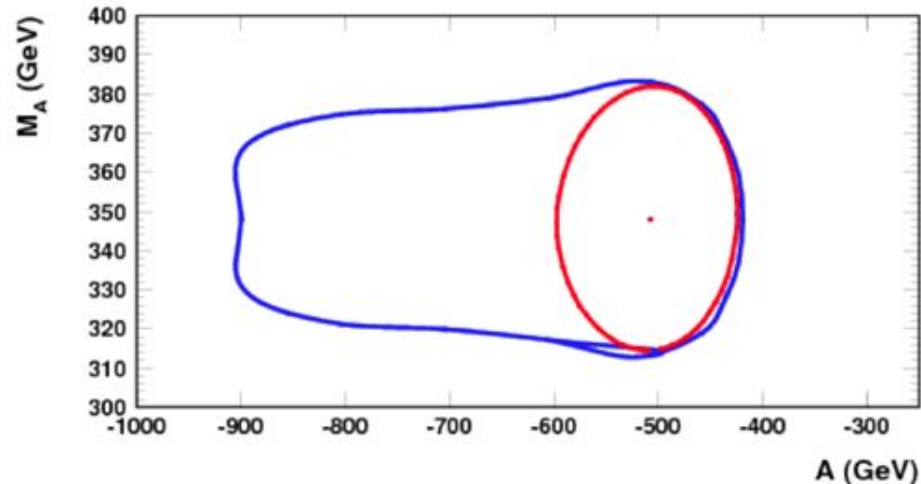


# CMSSM: LHC/SuperB complementarity



**Blue = LHC:**

- Will be able to measure  $m(A)$  [CP odd Higgs mass]
- Poor sensitivity to  $\tan\beta$  [ratio of Higgs vevs]
- Poor sensitivity to  $A$  [coupling]

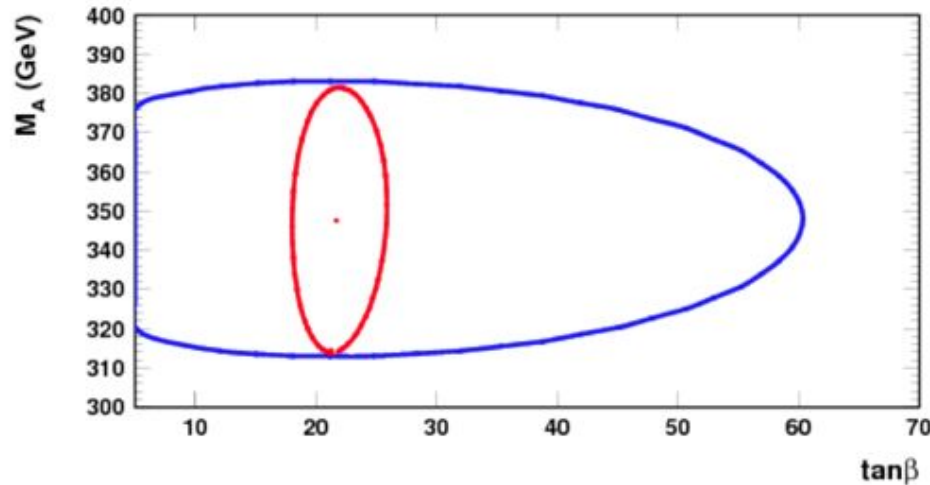


**Red=LHC+EW/Low-energy constraints (includes SuperB):**

Observable	Constraint	theo. error
$R_{BR_{b \rightarrow s\gamma}}$	$1.127 \pm 0.1$	0.1
$R_{\Delta M_s}$	$0.8 \pm 0.2$	0.1
$BR_{b \rightarrow \mu\mu}$	$(3.5 \pm 0.35) \times 10^{-8}$	$2 \times 10^{-9}$
$R_{BR_{b \rightarrow \tau\nu}}$	$0.8 \pm 0.2$	0.1
$\Delta a_\mu$	$(27.6 \pm 8.4) \times 10^{-10}$	$2.0 \times 10^{-10}$
$M_W^{\text{SUSY}}$	$80.392 \pm 0.020 \text{ GeV}$	0.020 GeV
$\sin^2 \theta_W^{\text{SUSY}}$	$0.23153 \pm 0.00016$	0.00016
$M_h^{\text{light}}(\text{SUSY})$	$> 114.4 \text{ GeV}$	3.0 GeV

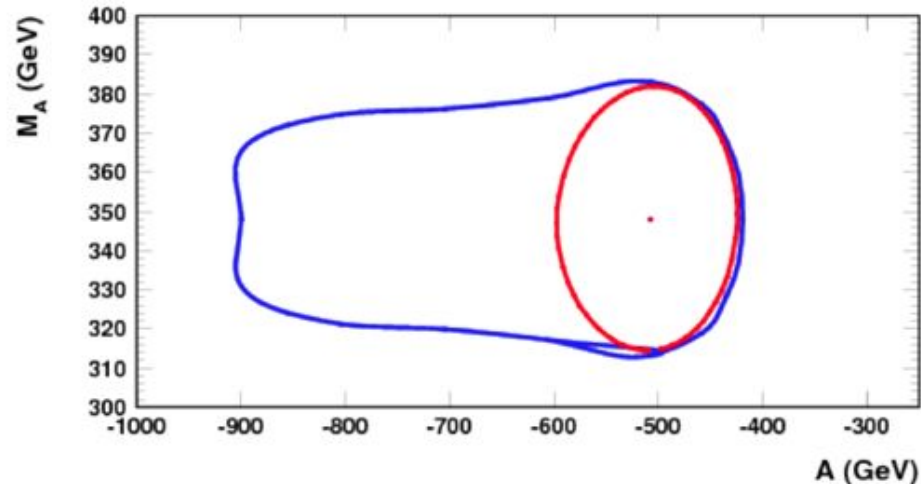
Current analysis of data prefers  $\tan\beta \sim 10$ .

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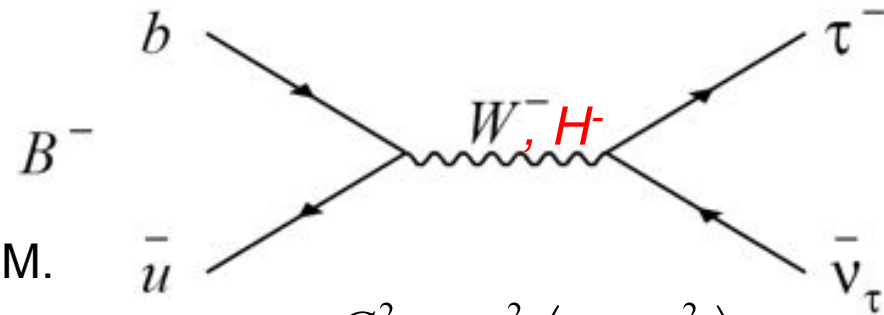
- Can build on the  $m(A)$  measurement to measure  $\tan\beta$ .

Current analysis of data prefers  $\tan\beta \sim 10$ .

Again LHC and SuperB are complementary experiments. Each can contribute significantly to the knowledge of new physics.

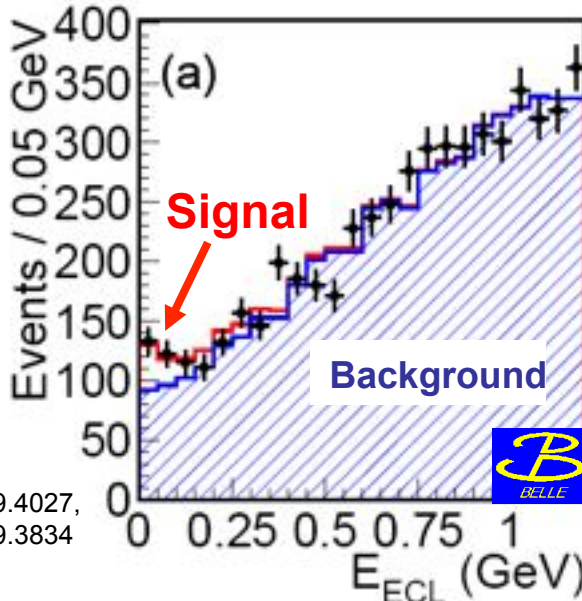
# Charged Higgs: $B^\pm \rightarrow \tau^\pm \nu$

- Within the SM, sensitive to  $f_B$  and  $|V_{ub}|$ :  $\mathcal{B}_{SM} \sim 1.6 \times 10^{-4}$ .
- $B$  affected by new physics.
  - MFV models like 2HDM / MSSM.
  - Unparticles.
- Fully reconstruct the event (modulo  $\nu$ ).

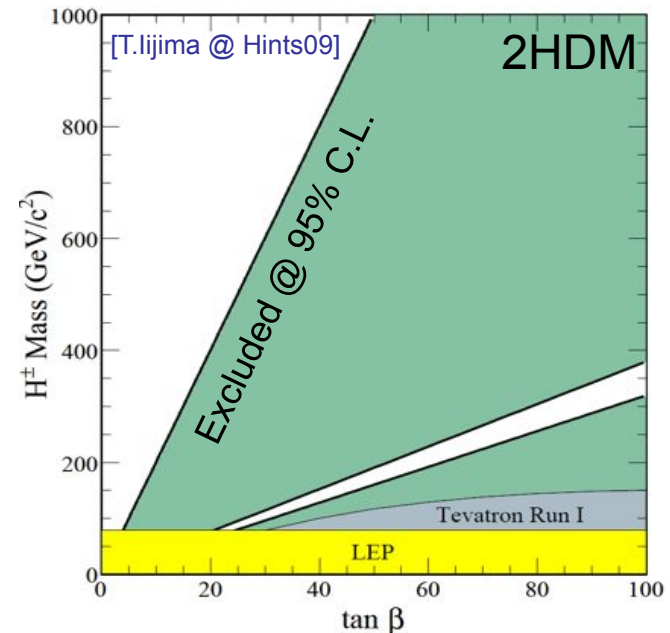


$$\mathcal{B}_{SM}(B^+ \rightarrow l^+ \nu_l) = \frac{G_F^2 m_B m_l^2}{8\pi} \left(1 - \frac{m_l^2}{m_B^2}\right) f_B^2 |V_{ub}|^2 \tau_B$$

$$\mathcal{B}_{WA} = (1.73 \pm 0.35) \times 10^{-4}$$



arXiv:0809.4027,  
arXiv:0809.3834



2HDM: W.-S Hou PRD **48** 2342 (1993)

MSSM: G. Isidori arXiv:0710.5377

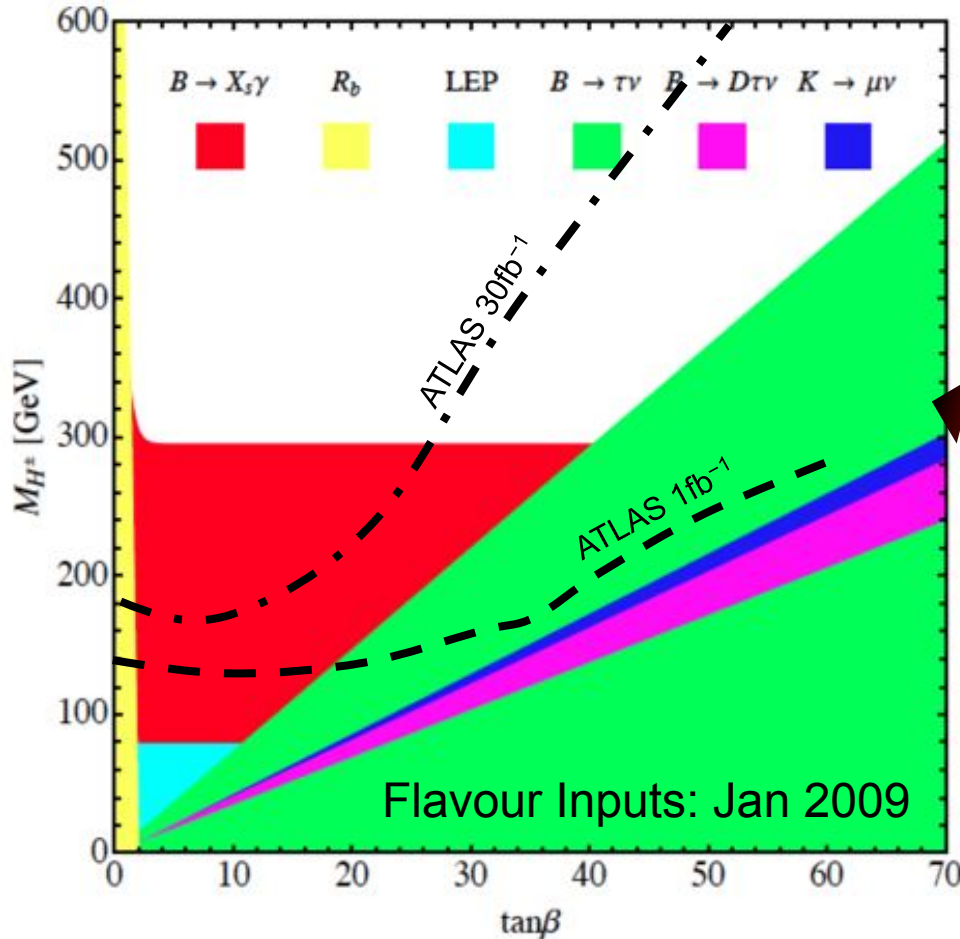
Unparticles: R. Zwicky PRD **77** 036004 (2008)

# Charged Higgs

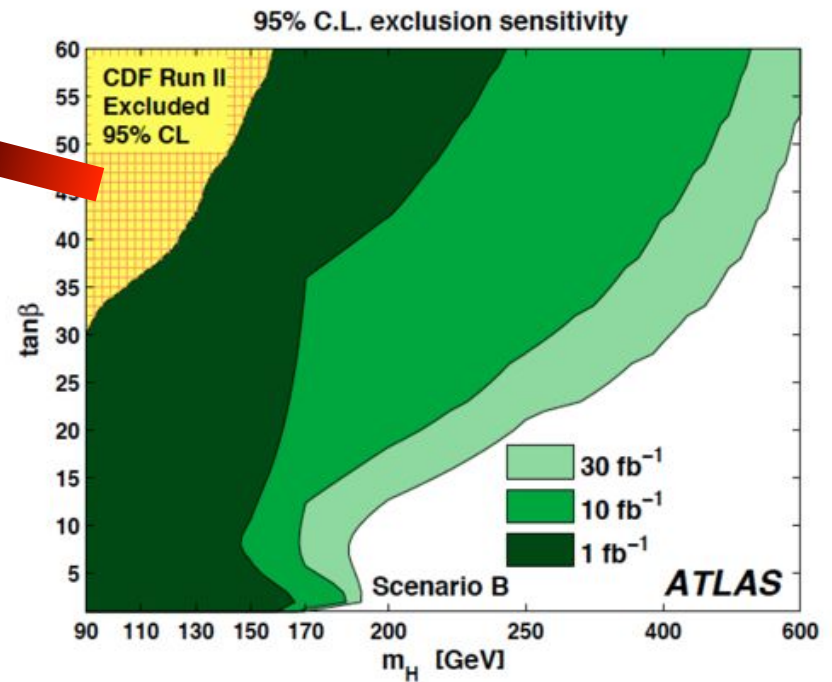
- B-factory searches competitive with LHC era: e.g. 2HDM

Existing Constraints from BaBar and Belle.

Combined Higgs search constraint from ATLAS: arXiv:0901.1502



Converted constraints expected from ATLAS onto the plot by hand.



U. Haisch 0805.2141

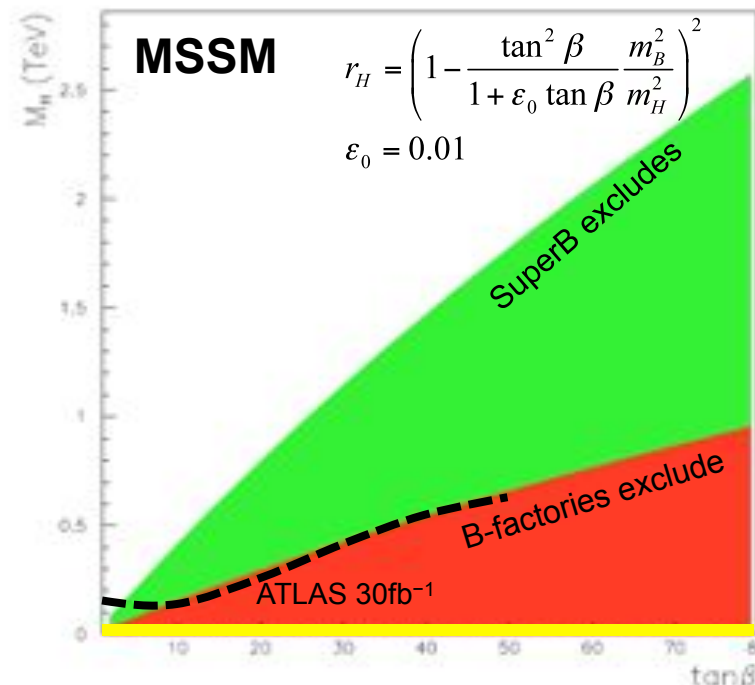
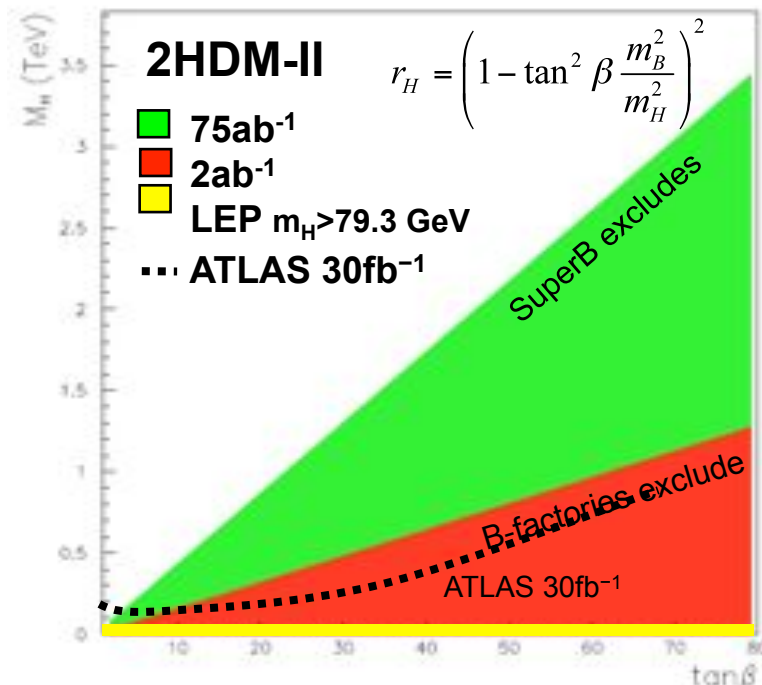
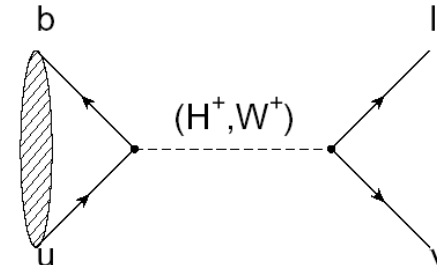
- Higgs mediated MFV:

$$r_H = \frac{\mathcal{B}_{SM+NP}}{\mathcal{B}_{SM}}$$

(Assuming SM branching fraction is measured)

B-factories have  $1.5ab^{-1}$  of data.

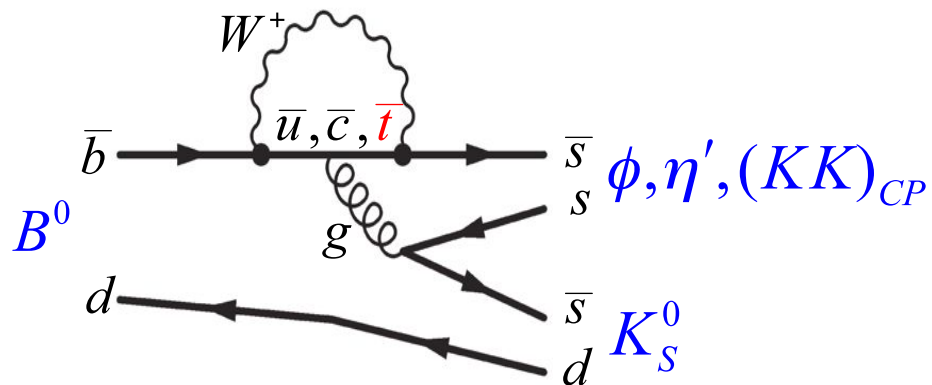
ATLAS sensitivity sketched from combined sensitivity plots in arXiv:0901.0512.



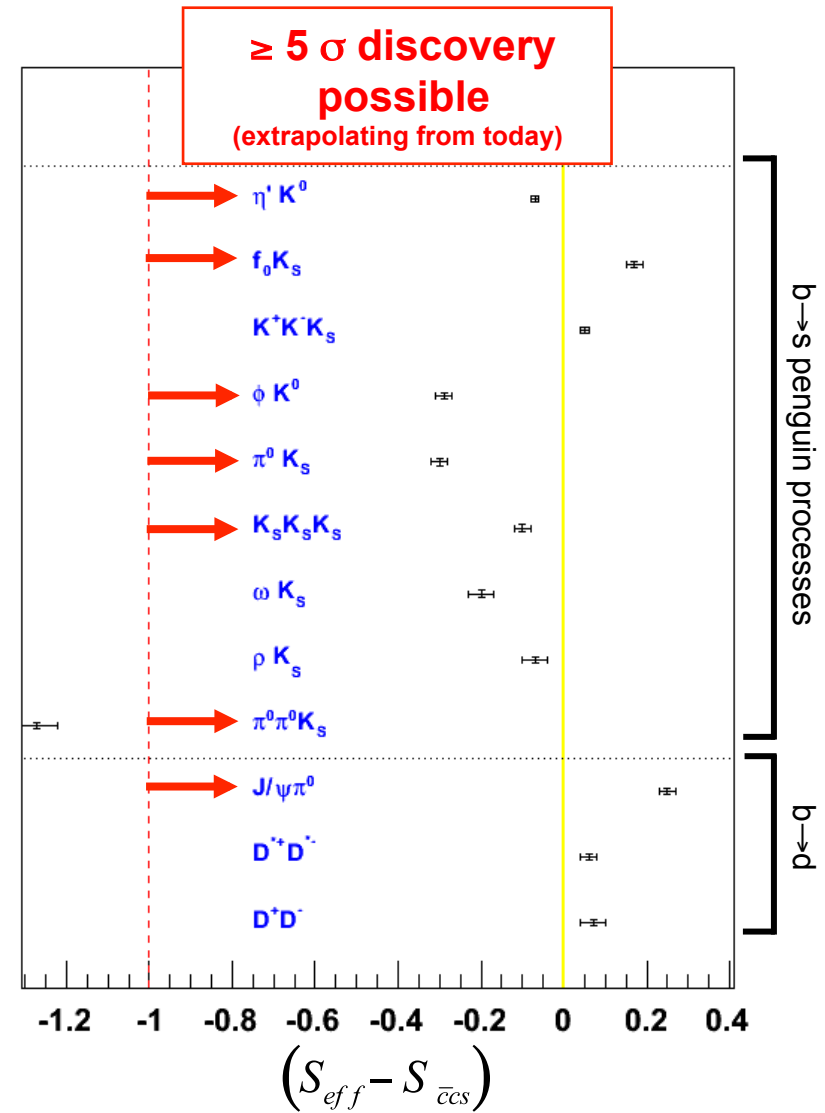
- Multi TeV search capability for large  $\tan\beta$ .
- Includes SM uncertainty  $\sim 20\%$  from  $V_{ub}$  and  $f_B$ .

# $\Delta S$ measurements

- $\beta = (21.1 \pm 0.9)^\circ$  from Charmonium decays.
- Look in many different  $b \rightarrow s$  and  $b \rightarrow d$  decays for  $\sin 2\beta$  deviations from the SM:
- The golden channel is:



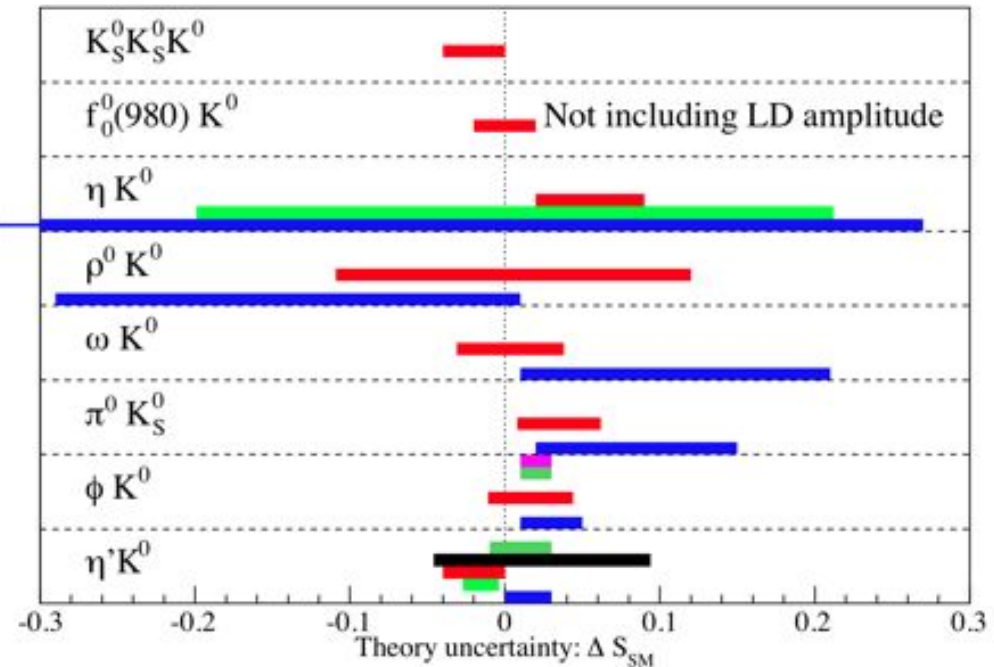
- Deviations would be from high mass particles in loops:  $H, \chi, \dots$





# $\Delta S$ measurements

- The SM uncertainty is strongly mode dependent.
- Golden modes have to be well measured and theoretically clean.
- Prefer to also have robust constraints from more than one theoretical approach.
- Precision measurements of the reference Charmonium decay also have a small SM uncertainty.



- QCDF, Beneke et al., PLB620 143 (2005)
- SCET/QCDF Williamson and Zupan PRD 74 014003 (2006)
- QCDF Cheng, Chua, Soni PRD72, 014006 (2005); PRD 74 094001 (2005)
- SU(3) Gronau, Rosner, Zupan PRD74 093003 (2006)
- QCDF Buchalla, Hiller, Nir, Raz, JHEP 09, 074 (2005)
- Li and Mishima PRD74, 094020 (2006)



# $\Delta S$ measurements

- We were reminded that we should be careful with what we compare:
  - NP could affect  $c\bar{c}s \sin 2\beta$ .

1) Predict  $\sin 2\beta$  from indirect constraints.

$$[\sin(2\beta)]_{\text{no } V_{ub}}^{\text{prediction}} = 0.87 \pm 0.09 \quad \color{green}{\blacksquare}$$

2) Compare to  $c\bar{c}s$  measurement.

$$[\sin 2\beta]_{c\bar{c}s} = 0.672 \pm 0.023 \quad \color{yellow}{\blacksquare}$$

3) Compare to clean penguin measurements.

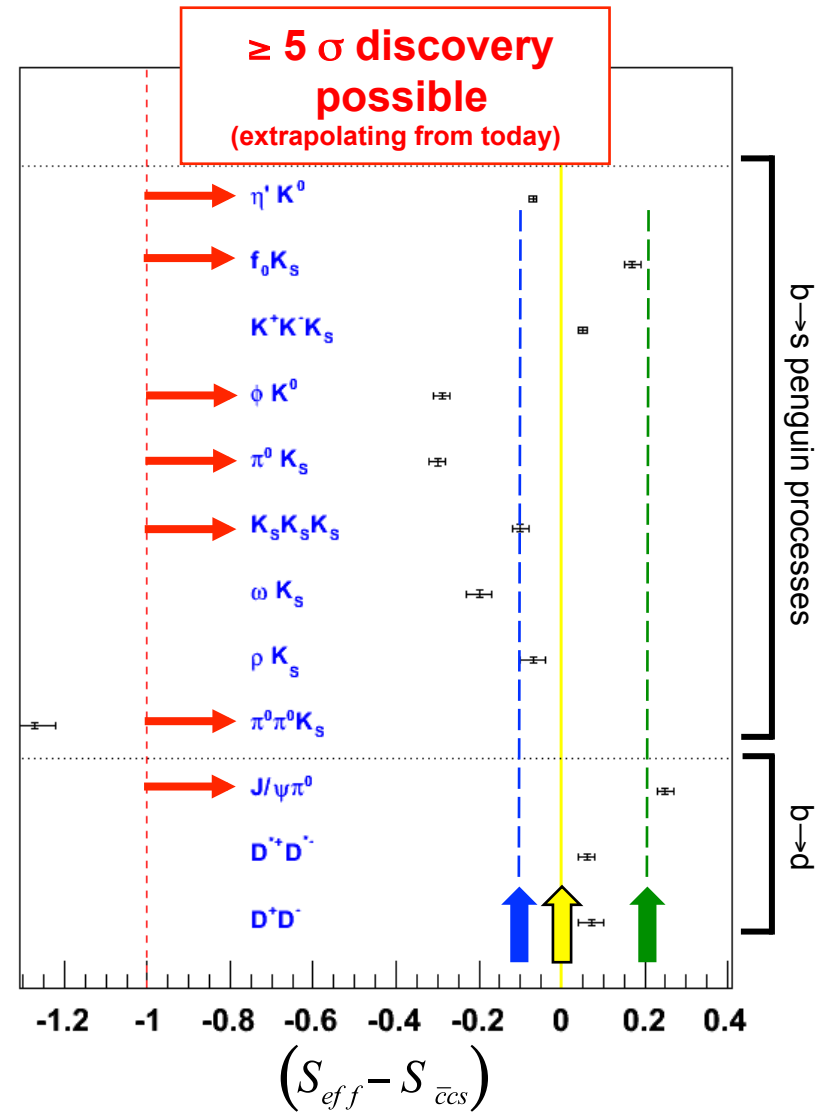
$$[\sin 2\beta]_{b \rightarrow s\text{-penguin}}^{\text{clean}} = 0.58 \pm 0.06 \quad \color{blue}{\blacksquare}$$

(or the average of the two)

**Are these 2.1-2.7 $\sigma$  hints  
for new physics?**

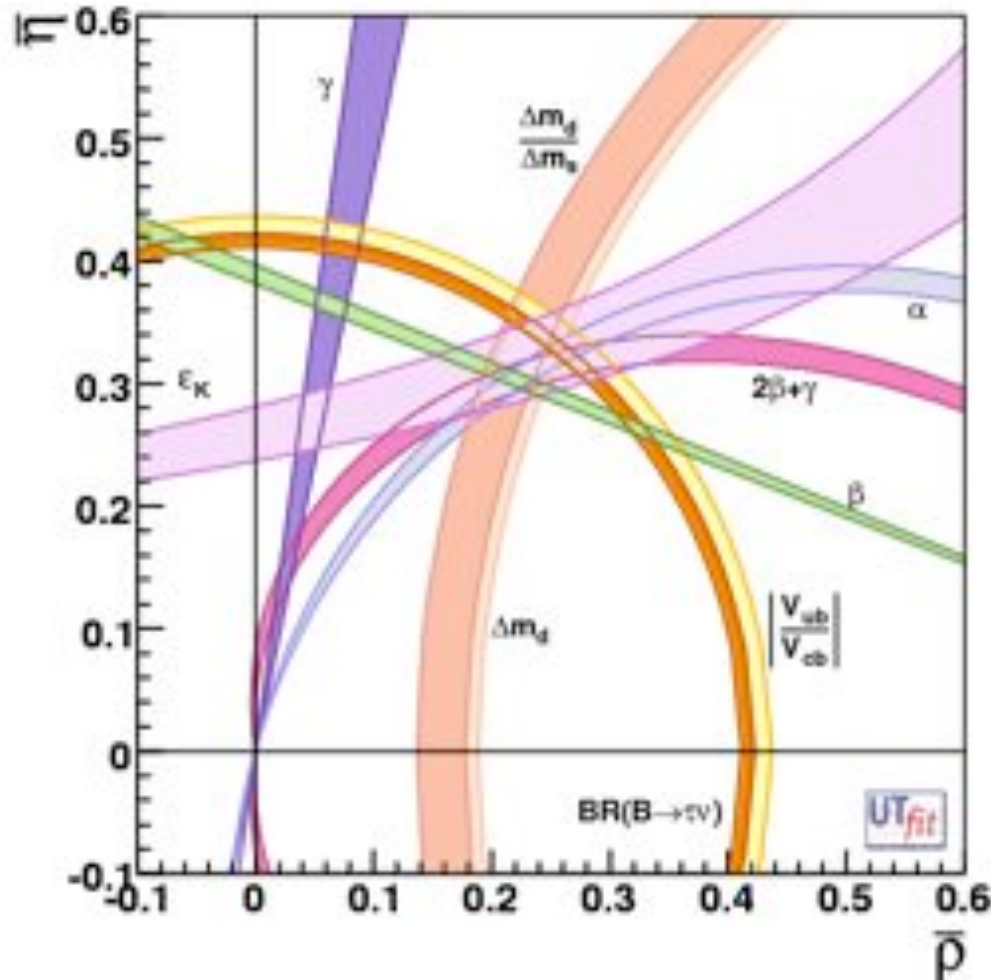
Lunghi and Soni, Phys.Lett.B**666** 162-165 (2008).  
Buras and Guadagnoli Phys Rev D **78** 033005 (2008).

- Can theory error be reduced for other modes?**





# ... and precision CKM



$$\sigma(\beta) < 0.5^\circ$$

$$\sigma(\alpha) \sim 1 - 2^\circ$$

$$\sigma(\gamma) \sim 1 - 2^\circ$$

*etc.*

Can we develop more insight into CKM?

- does it work at the % level?

- what new physics probes become available with precision CKM?

[for SuperB and Other expt.]



# Standard Model measurements.

## B Physics at Y(4S)

Observable	B Factories (2 ab <sup>-1</sup> )	SuperB (75 ab <sup>-1</sup> )
sin(2β) (J/ψ K <sup>0</sup> )	0.018	0.005 (†)
cos(2β) (J/ψ K <sup>0</sup> )	0.30	0.05
sin(2β) (Dh <sup>0</sup> )	0.10	0.02
cos(2β) (Dh <sup>0</sup> )	0.20	0.04
S(J/ψ π <sup>0</sup> )	0.10	0.02
S(D <sup>+</sup> D <sup>-</sup> )	0.20	0.03
S(φ K <sup>0</sup> )	0.13	0.02 (*)
S(φ K <sup>0</sup> )	0.05	0.01 (*)
S(K <sub>s</sub> <sup>0</sup> K <sub>s</sub> <sup>0</sup> )	0.15	0.02 (*)
S(K <sub>s</sub> <sup>0</sup> π <sup>0</sup> )	0.15	0.02 (*)
S(ω K <sub>s</sub> <sup>0</sup> )	0.17	0.03 (*)
S(f <sub>0</sub> K <sub>s</sub> <sup>0</sup> )	0.12	0.02 (*)
γ (B → DK, D → CP eigenstates)	~15°	2.5°
γ (B → DK, D → suppressed states)	~12°	2.0°
γ (B → DK, D → multibody states)	~9°	1.5°
γ (B → DK, combined)	~6°	1-2°
α (B → ππ)	~16°	3°
α (B → ρρ)	~7°	1-2° (*)
α (B → ρπ)	~12°	2°
α (combined)	~6°	1-2° (*)
2β + γ (D <sup>0</sup> ± π <sup>±</sup> , D <sup>±</sup> K <sub>s</sub> <sup>0</sup> π <sup>±</sup> )	20°	5°
V <sub>ub</sub>   (exclusive)	4% (*)	1.0% (*)
V <sub>ub</sub>   (inclusive)	1% (*)	0.5% (*)
V <sub>ub</sub>   (exclusive)	8% (*)	3.0% (*)
V <sub>ub</sub>   (inclusive)	8% (*)	2.0% (*)
B(B → τν)	20%	4% (†)
B(B → μν)	visible	5%
B(B → Dτν)	10%	2%
B(B → ργ)	15%	3% (†)
B(B → ωγ)	30%	5%
A <sub>CP</sub> (B → K <sup>+</sup> γ)	0.007 (†)	0.004 († *)
A <sub>CP</sub> (B → ργ)	~0.20	0.05
A <sub>CP</sub> (b → sγ)	0.012 (†)	0.004 (†)
A <sub>CP</sub> (b → (s + d)γ)	0.03	0.006 (†)
S(K <sub>s</sub> <sup>0</sup> π <sup>0</sup> γ)	0.15	0.02 (*)
S(μ <sup>0</sup> γ)	possible	0.10
A <sub>CP</sub> (B → K <sup>+</sup> ℓℓ)	7%	1%
A <sup>FB</sup> (B → K <sup>+</sup> ℓℓ)s <sub>0</sub>	25%	9%
A <sup>FB</sup> (B → X <sub>s</sub> ℓℓ)s <sub>0</sub>	35%	5%
B(B → Kνℓ)	visible	20%
B(B → πνℓ)	-	possible

## Rare Charm Decays: 1 month at ψ(3770)

Channel	Sensitivity
D <sup>0</sup> → e <sup>+</sup> e <sup>-</sup> , D <sup>0</sup> → μ <sup>+</sup> μ <sup>-</sup>	1 × 10 <sup>-8</sup>
D <sup>0</sup> → π <sup>0</sup> e <sup>+</sup> e <sup>-</sup> , D <sup>0</sup> → π <sup>0</sup> μ <sup>+</sup> μ <sup>-</sup>	2 × 10 <sup>-8</sup>
D <sup>0</sup> → ηe <sup>+</sup> e <sup>-</sup> , D <sup>0</sup> → ημ <sup>+</sup> μ <sup>-</sup>	3 × 10 <sup>-8</sup>
D <sup>0</sup> → K <sub>s</sub> <sup>0</sup> e <sup>+</sup> e <sup>-</sup> , D <sup>0</sup> → K <sub>s</sub> <sup>0</sup> μ <sup>+</sup> μ <sup>-</sup>	3 × 10 <sup>-8</sup>
D <sup>+</sup> → π <sup>+</sup> e <sup>+</sup> e <sup>-</sup> , D <sup>+</sup> → π <sup>+</sup> μ <sup>+</sup> μ <sup>-</sup>	1 × 10 <sup>-8</sup>
D <sup>0</sup> → e <sup>±</sup> μ <sup>∓</sup>	1 × 10 <sup>-8</sup>
D <sup>+</sup> → π <sup>+</sup> e <sup>±</sup> μ <sup>∓</sup>	1 × 10 <sup>-8</sup>
D <sup>0</sup> → π <sup>0</sup> e <sup>±</sup> μ <sup>∓</sup>	2 × 10 <sup>-8</sup>
D <sup>0</sup> → ηe <sup>±</sup> μ <sup>∓</sup>	3 × 10 <sup>-8</sup>
D <sup>0</sup> → K <sub>s</sub> <sup>0</sup> e <sup>±</sup> μ <sup>∓</sup>	3 × 10 <sup>-8</sup>
D <sup>+</sup> → π <sup>-</sup> e <sup>+</sup> e <sup>+</sup> , D <sup>+</sup> → K <sup>-</sup> e <sup>+</sup> e <sup>+</sup>	1 × 10 <sup>-8</sup>
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## τ: LFV / CPV / ...

Process	Sensitivity
B(τ → μ γ)	2 × 10 <sup>-9</sup>
B(τ → e γ)	2 × 10 <sup>-9</sup>
B(τ → μ μ μ)	2 × 10 <sup>-10</sup>
B(τ → e e e)	2 × 10 <sup>-10</sup>
B(τ → μ η)	4 × 10 <sup>-10</sup>
B(τ → e η)	6 × 10 <sup>-10</sup>
B(τ → ℓ K <sub>s</sub> <sup>0</sup> )	2 × 10 <sup>-10</sup>

## Charm Mixing

Mode	Observable	Υ(4S) (75 ab <sup>-1</sup> )	ψ(3770) (300 fb <sup>-1</sup> )
D <sup>0</sup> → K <sup>+</sup> π <sup>-</sup>	x' <sup>2</sup>	3 × 10 <sup>-5</sup>	
	y'	7 × 10 <sup>-4</sup>	
D <sup>0</sup> → K <sup>+</sup> K <sup>-</sup>	y <sub>CP</sub>	5 × 10 <sup>-4</sup>	
D <sup>0</sup> → K <sub>s</sub> <sup>0</sup> π <sup>+</sup> π <sup>-</sup>	x	4.9 × 10 <sup>-4</sup>	
	y	3.5 × 10 <sup>-4</sup>	
	q/p	3 × 10 <sup>-2</sup>	
	φ	2°	
ψ(3770) → D <sup>0</sup> D <sup>0</sup>	x <sup>2</sup>		(1-2) × 10 <sup>-5</sup>
	y		(1-2) × 10 <sup>-3</sup>
	cos δ		(0.01-0.02)

See CDR and Valencia report for details of the SM measurements and other possible NP searches.

## B Physics at Y(5S)

Observable	Error with 1 ab <sup>-1</sup>	Error with 30 ab <sup>-1</sup>
ΔΓ	0.16 ps <sup>-1</sup>	0.03 ps <sup>-1</sup>
Γ	0.07 ps <sup>-1</sup>	0.01 ps <sup>-1</sup>
β <sub>s</sub> from angular analysis	20°	8°
A <sub>SL</sub> <sup>s</sup>	0.006	0.004
A <sub>CH</sub>	0.004	0.004
B(B <sub>s</sub> → μ <sup>+</sup> μ <sup>-</sup> )	-	< 8 × 10 <sup>-9</sup>
V <sub>td</sub> /V <sub>ts</sub>	0.08	0.017
B(B <sub>s</sub> → γγ)	38%	7%
β <sub>s</sub> from J/ψφ	10°	3°
β <sub>s</sub> from B <sub>s</sub> → K <sup>0</sup> K <sup>0</sup>	24°	11°



# Other measurements of interest

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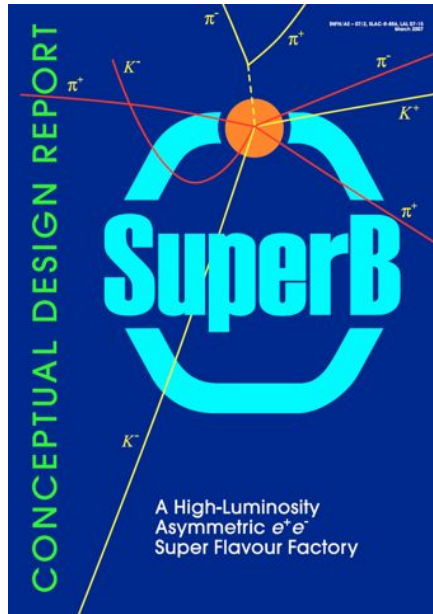
- Let us not forget the other opportunities for searching for new physics at SuperB:
  - Charm decays: CPV and parallels of rare B decays.
  - LFV in light mesons
  - Light (few GeV) Higgs Searches
  - CPT Violation test / EPR correlation tests
  - Dark matter ... and Dark Forces



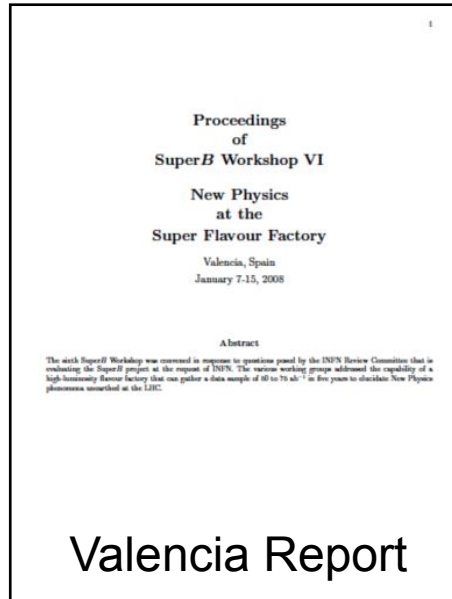
# Recent (and imminent) milestones



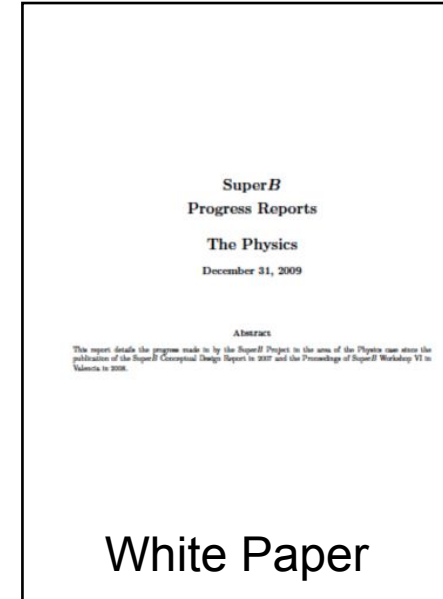
# Milestones in the programme...



2007



2008



2009/2010

- The white paper is a stepping stone toward the TDR (end of 2010).
- During this workshop we aim to review the state of the art in terms of flavour physics capabilities of SuperB.



# Recent Progress

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- A lot has happened since Warwick:
  - Dark Forces has gathered a lot of interest in the community (See F. Bossi).
  - First results from MEG (See F. Renga).
  - LHC data taking has re-started (See T. Hurth).
  - Full and Fast Sim. and a SuperB distributed computing model has come online (See the General Meeting).
    - We are now starting to optimize SuperB based on a few key benchmark channels.
- All of these factors are additional inputs to refine our goals.
- This work will continue through 2010 as we re-focus for the TDR.





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# Physics White Paper



# Purpose

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- Provide a coherent description of the physics case of the experiment:
  - New Physics Search Capability:
    - This is the reason why SuperB is relevant in 2015 (when we plan to go online).
    - What other inputs are necessary to decode the behaviour of new physics more comprehensively?
    - How the SuperB measurements push back our understanding of the viable models?
  - Standard Model Physics:
    - Precision CKM: as a test of the CKM ansatz.
    - As a control of SM uncertainties for other (non-)SuperB benchmark channels to search for new physics.



# Scope

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- Aim: 40-50 pages:
  - Have made a good start with the CDR & Valencia document.
  
- Several key features have changed:
  - D mixing has been established, and CPV searches in charm are more mature.
  - NP probes in B and  $\tau$  decays have advanced since last time.
  - Lattice QCD has matured: we can compare against CDR predictions.
  - Tevatron has measured  $\Phi_s$  more precisely.
  
- This document should be stand alone.



# Content

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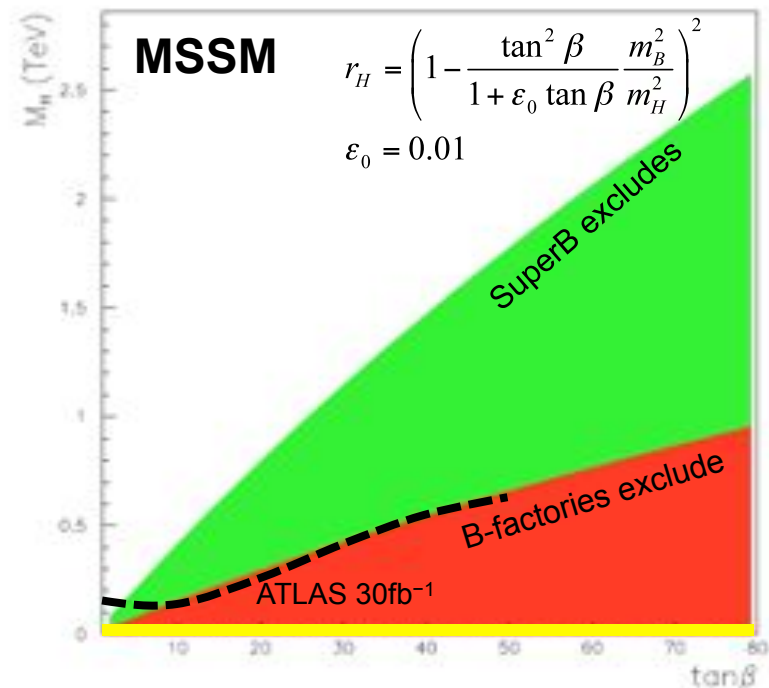
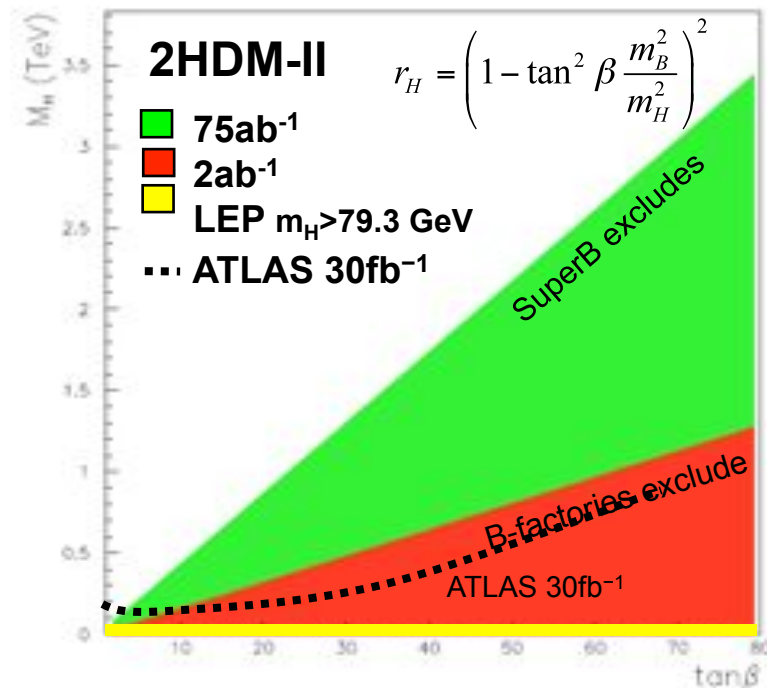
- Introduction
- B<sub>u,d,s</sub> physics
- Charm physics
- $\tau$  physics
- Spectroscopy
- Role of Lattice QCD
- Interplay with other experiments
- Conclusion

The content is based on previous documents. If you think that a new section is needed – then propose it to the conveners in charge of that section.



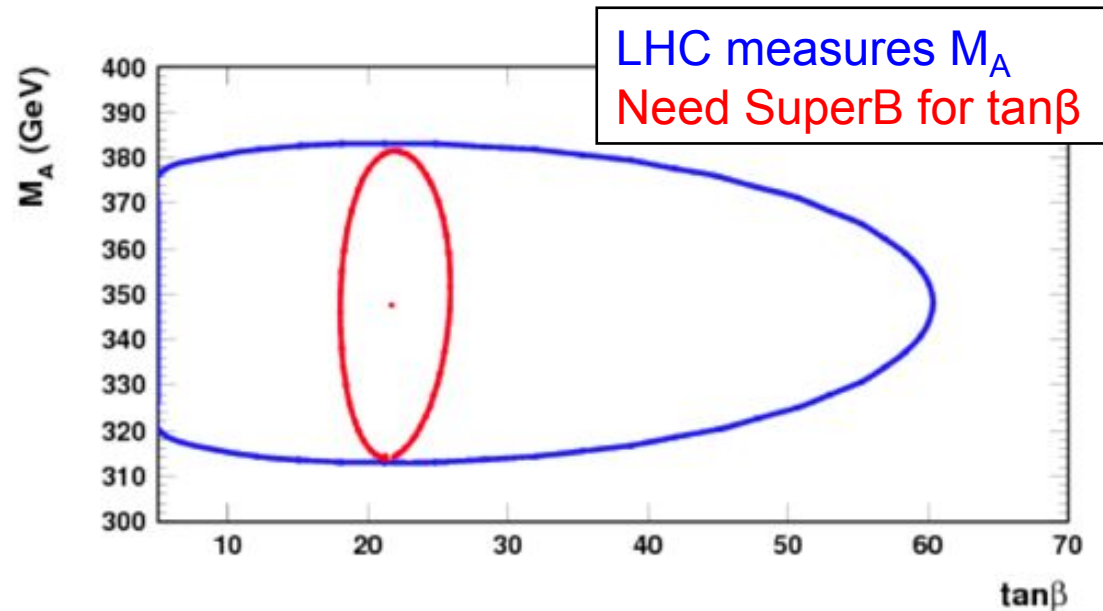
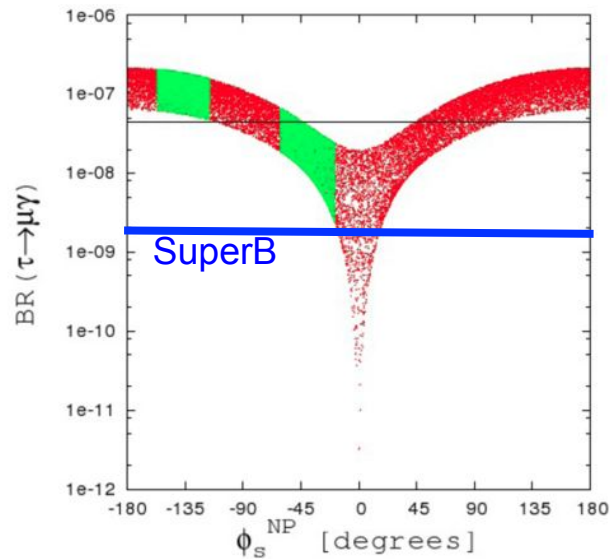
# Why have an interplay section?

- This is how we can show what SuperB will teach us about nature; in the context of existing and planned experimental programmes.



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- This is how we can show what SuperB will teach us about nature; in the context of existing and planned experimental programmes.

	$H^+$ high $\tan \beta$	MFV	Non-MFV	NP Z-penguins	Right-handed currents	LTH	SUSY
$\mathcal{B}(B \rightarrow X_s \gamma)$		L	M		M		
$\mathcal{A}_{CP}(B \rightarrow X_s \gamma)$			L		M		
$\mathcal{B}(B \rightarrow \tau \nu)$	L-CKM						
$\mathcal{B}(B \rightarrow X_s \ell \ell)$			M	M	M		
$\mathcal{B}(B \rightarrow K \nu \bar{\nu})$			M	L			
$S_{K_S \pi^0 \gamma}$					L		
The angle $\beta$ ( $\Delta S$ )			L-CKM		L		
$\tau \rightarrow \mu \gamma$							L
$\tau \rightarrow \mu \mu \mu$						L	

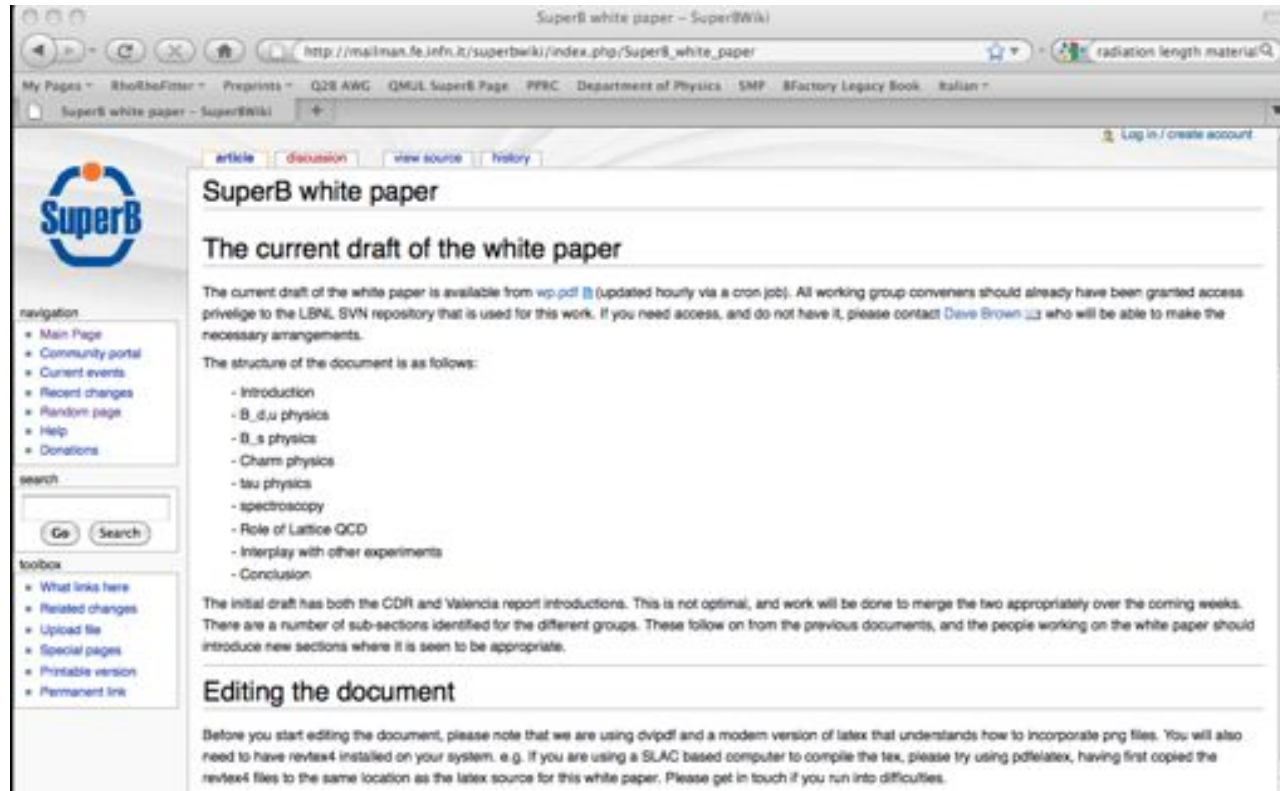
Need to elaborate on the golden matrix of observables that can be used to decode new physics.

This is more than just rare B and  $\tau$  decays... See the talks later this morning!



# Where can I find the Tex source?

- The source is in an SVN repository at LBNL.
- Instructions for accessing this are on the Physics Wiki Page.
  - Physics Portal -> Documentation -> White Paper







# How do I get to see the current version?

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- A trscron job updates the pdf version of the white paper every hour.
  - Can download a current copy that corresponds to whatever has been committed.
  - Will always be up-to-date within 60min (*or broken*)

<http://www.slac.stanford.edu/~bevan/superbID/wp.pdf>

**SuperB**  
**Progress Reports**

**The Physics**

**December 31, 2009**

**Abstract**

This report details the progress made in by the SuperB Project in the area of the Physics case since the publication of the SuperB Conceptual Design Report in 2007 and the Proceedings of SuperB Workshop VI in Valencia in 2008.



# How do I get an account to contribute?

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- This work uses the original LBNL SuperB repository.
- Working group conveners and theorists have been given accounts to access this repository.
- If you have not already been informed of this and would like to access the Tex to contribute to the white paper, please see Dave Brown: [Dave\\_Brown@lbl.gov](mailto:Dave_Brown@lbl.gov).



# Timescale

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- We want to have the white paper finalized as soon as possible.
  - Get as much work done as possible this week.
  - Follow up with this work through the Christmas vacation.
  - Finish the document early next year at the latest: end Dec/Start Jan.
- *We can come back to the white paper at the end of the morning and take discussions through lunch!*



# Summary

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- Plenary talks this morning:
  - Dark Forces (F. Bossi)
  - MEG (F. Renga)
  - Interplay (T. Hurth)
  
- Parallel sessions this afternoon.
  
- Lattice plenary talk on Tuesday 11.14 (V. Lubicz).
  
- Parallel sessions on the remaining afternoons.