Angular Distribution

Gamma Spectra





Dept. of Physics and Astronomy University of Padova

AGATA-MUGAST-VAMOS Comissioning Analysis Report

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Angular Distribution

Gamma Spectra

Conclusion









- $\rightarrow\,$ Angular distribution for transfer to g.s. and 1/2+ of $^{17}{\rm O}$
- \rightarrow Proton-gamma coincidence for the 870 keV gamma decay
- $\rightarrow\,$ AGATA efficiency estimate.

Angular Distribution

Setup







- ¹⁶O beam on CD2 Target
 - \rightarrow 6 MeV/u 10⁴pps
 - \rightarrow 1 mg/cm²
- AGATA backward
 → 37 crystal
- MUGAST backward
 - \rightarrow 5 trapezoidal DSSD
 - \rightarrow 1 anular DSSD
 - ightarrow 2 square DSSD at 90°
- MUST2 forward
 - \rightarrow 4 Telescope
- VAMOS

Angular Distribution









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Angular Distribution ●○○○ Gamma Spectra

Conclusion







11 h. acquisition long at $\sim 4~10^4~\text{pps}$



Angular Distribution

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Kinematic Lines





Angular Distribution

Gamma Spectra



Kinematic Lines





Angular Distribution

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Conclusion



Angular Distribution





Angular Distribution

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Differential Cross Section



Normalized with simulated angular efficiency ↓ Fit over the theoretical distribution

Efficiency normalization to be checked in the anular acceptance

DWBA calculation from Jesus Casal Università degli Studi di Padova INFN Sezione di Padova

Optical potential from An et al.,PRC 73.5 (2006) Watson et al., PR 182.4 (1969)



Angular Distribution

Gamma Spectra ●○○











Angular Distribution

Gamma Spectra ○●○ Conclusion



Doppler Correction





- Beta computation from proton direction
- First interaction from tracking (51 mm AGATA shift)

The peak is 1.8 keV shifted from 1/2+ state adopted energy.

Angular Distribution

Gamma Spectra ○●○



Doppler Correction



Minimization of resolution with scan over different shifts



Angular Distribution

Gamma Spectra ○●○ Conclusion



Doppler Correction



Best Resolution and Energy match with 33 mm shift



00	Angular Distribution	Gamma Spectra 00●	Conclusion 00
INFN	Deduced	Efficiency	
30 25- 20- 15- 10- 10- 780 800 820 840	860 880 900 920 940 960 Gamma Energy [keV]	 870 keV excitation peak is (background remove) 2562 ± 50 protons 870 keV gamma peak in (background remove) Add Back → 177 ± Tracking → 160 ± Efficiency Estimate: Add Back → 7.0 ± 0.5 Tracking → 6.3 ± 0.5 	integral ed) itegral ed) 13 γ 13 γ % %

Angular Distribution

Gamma Spectra



Summary



- Kinematic of the outgoing proton and $^{17}{\rm O}$ (5/2+ and 1/2+) final levels selection from excitation energy.
- Angular distribution normalized and fitted to the theoretical differential cross section.
- Coincidence with AGATA and 870 keV peak doppler correction.
- AGATA efficiency estimate

Perspectives

- Investigation on mugast simulated angular efficiency.
- Check VAMOS performance.
- Investigation on AGATA resolution.

Angular Distribution

Gamma Spectra

Conclusion ○●



Acknowledgments



Thanks to AGATA+GRIT+VAMOS collaborations and thank you for attention!