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INFN – Sezione di Firenze

*for the* **FAZIA collaboration**

Charged particle  
identification  
using **time of flight**  
with FAZIA

**FATA 2019**

Acireale, September 3<sup>rd</sup> – 5<sup>th</sup>, 2019



# FURBO project

## FAZIA Upgrade for Radioactive Beam Operation (INFN grant for new staff researchers)

# FURBO project

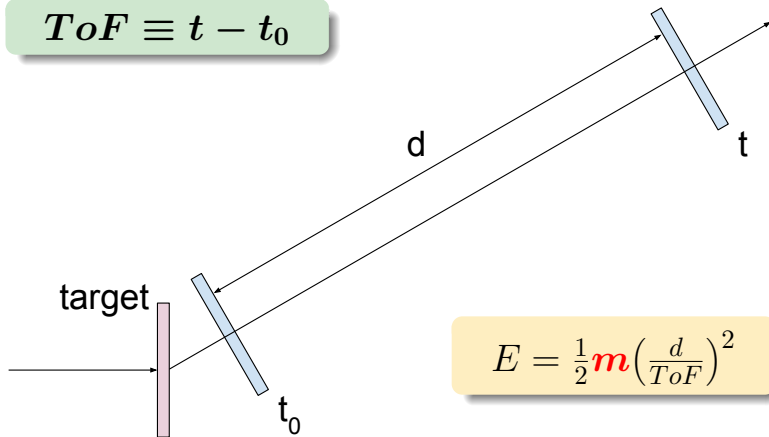
## FAZIA Upgrade for Radioactive Beam Operation (INFN grant for new staff researchers)

### Reduction of identification thresholds

- Fundamental task to measure in future ISOL facilities (SPES, Spiral2, . . .)
- Physics of quasi-target
- Different possible solutions:
  - **time of flight** implementation (discussed here)
  - use of **thin** Si detectors as first stage
  - use of **alternative** detectors

# Time of flight identification

$$ToF \equiv t - t_0$$

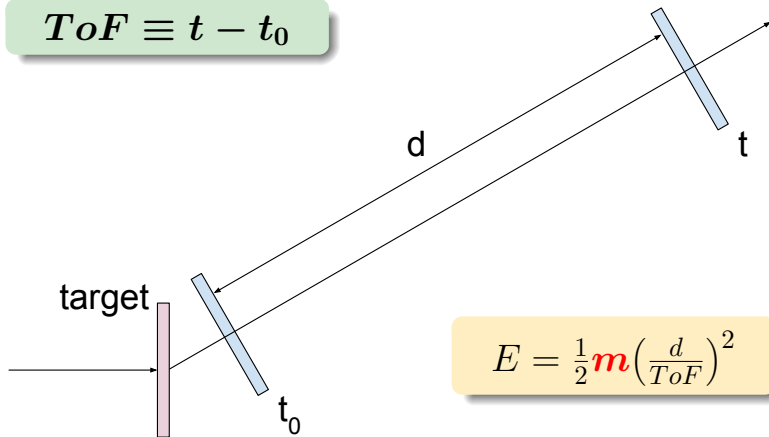


$$E = \frac{1}{2} \textcolor{red}{m} \left( \frac{d}{ToF} \right)^2$$

Recover **mass** of particles from ToF and energy

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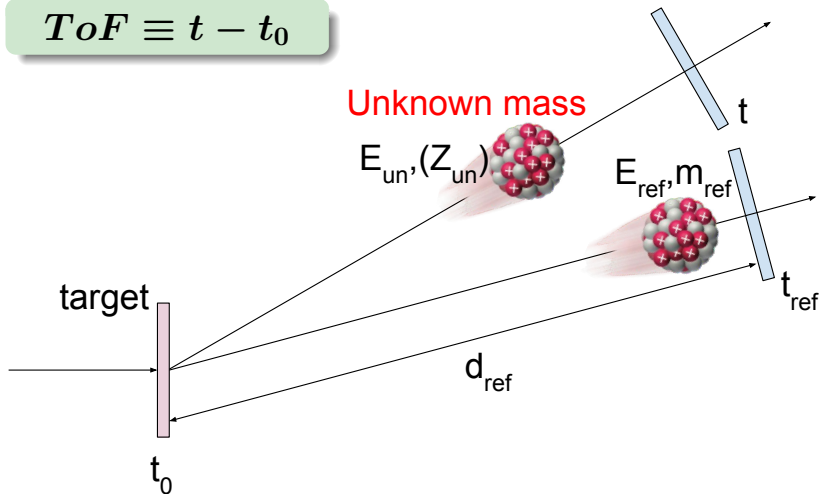


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**Start time mark needed!**

# Time of flight identification (only $M > 1$ events)

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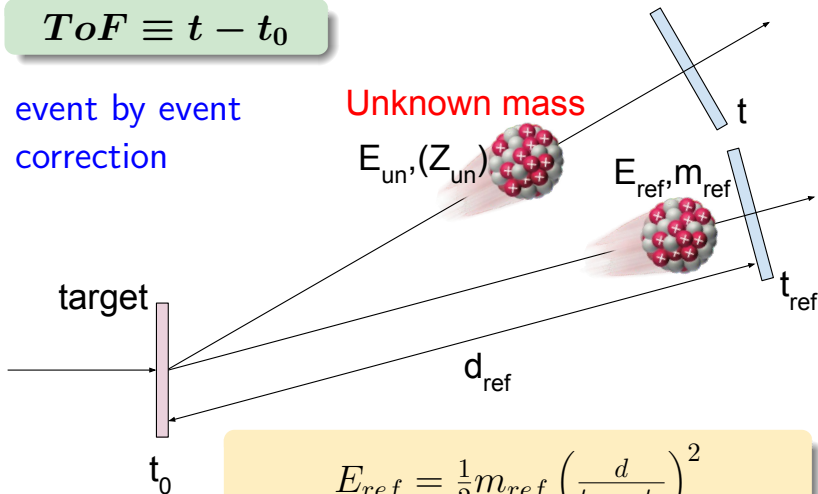


Proposed solution **without a start detector**

# Time of flight identification (only $M > 1$ events)

$$ToF \equiv t - t_0$$

event by event  
correction

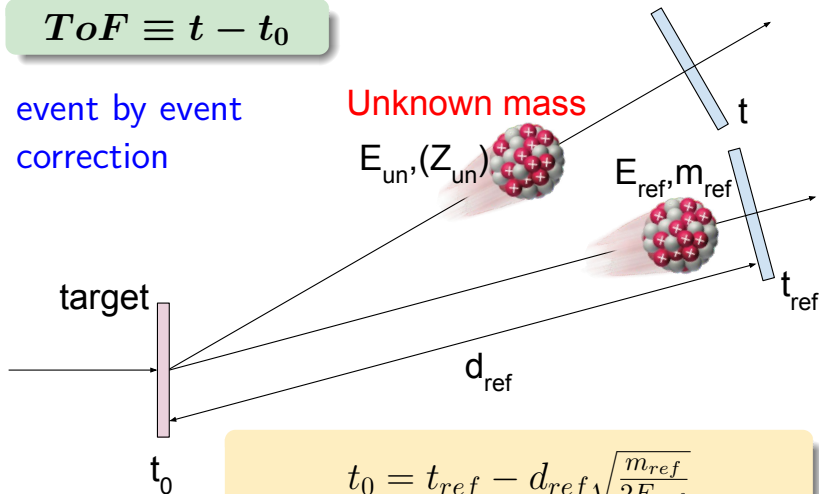


$$E_{ref} = \frac{1}{2} m_{ref} \left( \frac{d}{t_{ref} - t_0} \right)^2$$

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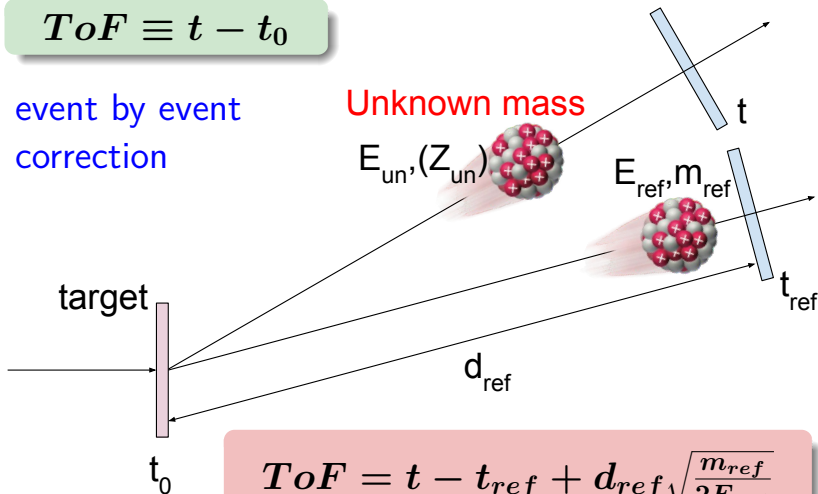
$$t_0 = t_{ref} - d_{ref} \sqrt{\frac{m_{ref}}{2E_{ref}}}$$



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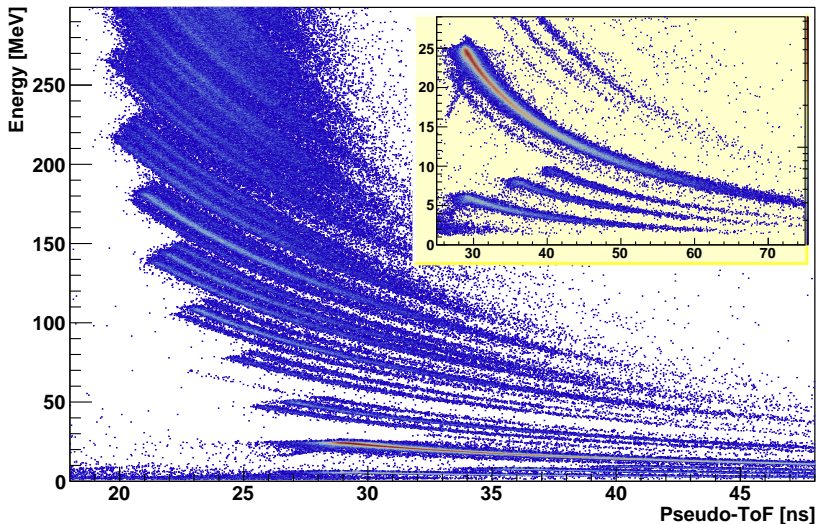
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$$ToF = t - t_{ref} + d_{ref} \sqrt{\frac{m_{ref}}{2E_{ref}}}$$

# Time of flight identification



“Spoiler” of FAZIA discrimination capabilities

# The FAZIA telescope

## The telescope stages

- ❶ 300  $\mu\text{m}$  reverse-mounted Si detector;
- ❷ 500  $\mu\text{m}$  reverse-mounted Si detector;
- ❸ 10 cm CsI(Tl) cristal read by a photodiode.

*To achieve the best possible energy resolution and  $A$  and  $Z$  identification Si detectors come from a  $n\text{TD}$  ingot cut at random angle to avoid channeling effects.*

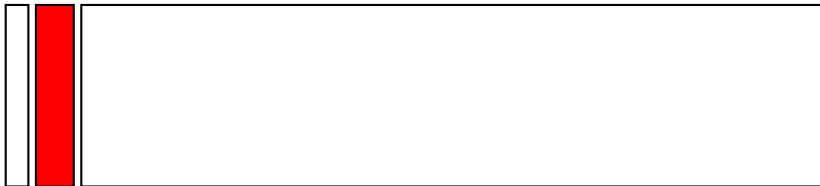


# The FAZIA telescope

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- 1 300  $\mu\text{m}$  reverse-mounted Si detector;
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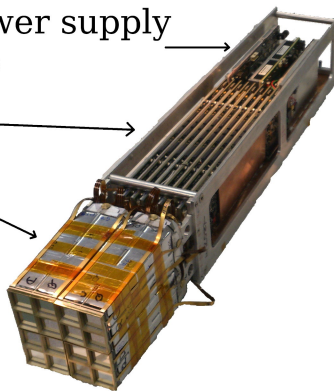


# The FAZIA block

Block card, power supply  
and half bridge

FEE cards

Detectors



***16 telescopes**, together with **front-end electronics**,  
form a **block** operating in **vacuum**.*

# FAZIA electronics

## Front-end

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- Compactness and modularity
- Very good isotopic discrimination capabilities
- Thresholds suited for Fermi energies (2–10 MeV/u)

# Identification methods

## $\Delta E - E$ correlation

- exploits the Bethe-Bloch energy loss relation
- identification threshold due to first layer thickness

## Pulse Shape Analysis<sup>a</sup>

- charge collection depending on the impinging nuclei
- identification threshold corresponding to  $\sim 50 \mu\text{m}$  penetration

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## $E - \text{ToF}$ correlation

- FAZIA implementation proposed here
- lowest identification threshold

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# Time of flight with FAZIA

Not the first heavy-ion experiment to implement ToF<sup>a</sup>

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# Time of flight with FAZIA

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## Our challenges:

- **large area** ( $2 \times 2 \text{ cm}^2$ ), **reverse-mounted** Si detectors;
- signal slowed down by anti-aliasing filter;
- time mark extraction from 250 MS/s sampled signals;
- not using beam radiofrequency;
- 1 m flight base

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# FAZIA time mark

For time mark extraction, after some tests we decided to adopt a **digital ARC-CFD<sup>a</sup>** with  $t_D = 20 \text{ ns}$  and  $f = 20 \%$

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FAZIA time mark is digitally extracted from the acquired signal:

- we used the first layer **low range** signal ( $\sim 300$  MeV range, 14-bit @ 250 MS/s);
- all signals are referred to the same validation time, which must be subtracted to obtain the true time mark:

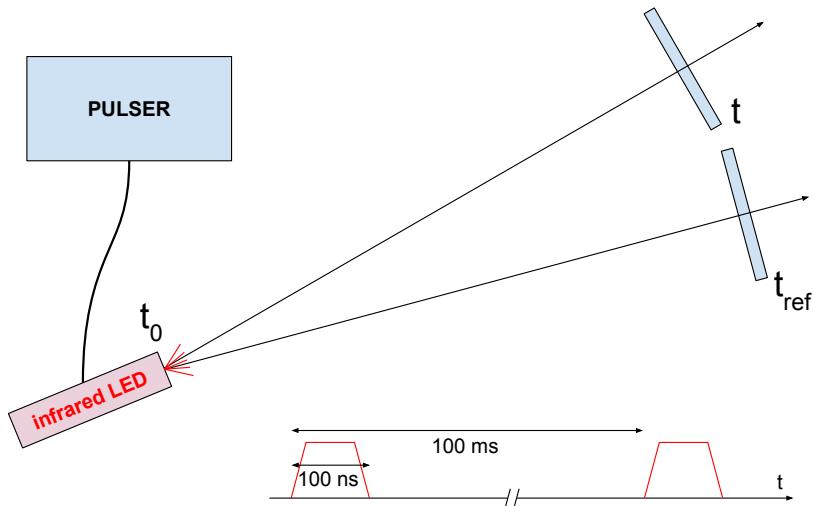
$$t^{(\text{ev,det})} = t_{\text{CFD}}^{(\text{ev,det})} - t_{\text{val}}^{(\text{ev})}$$

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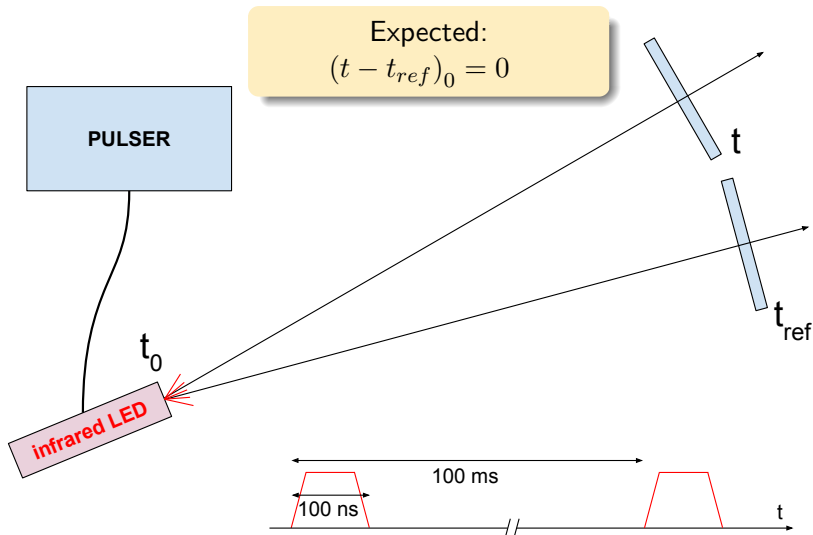
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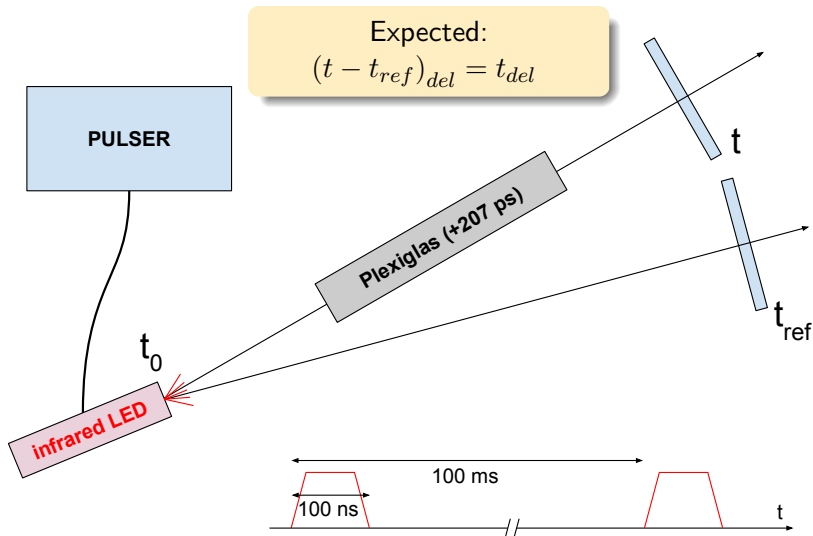
# Timing accuracy test in Florence



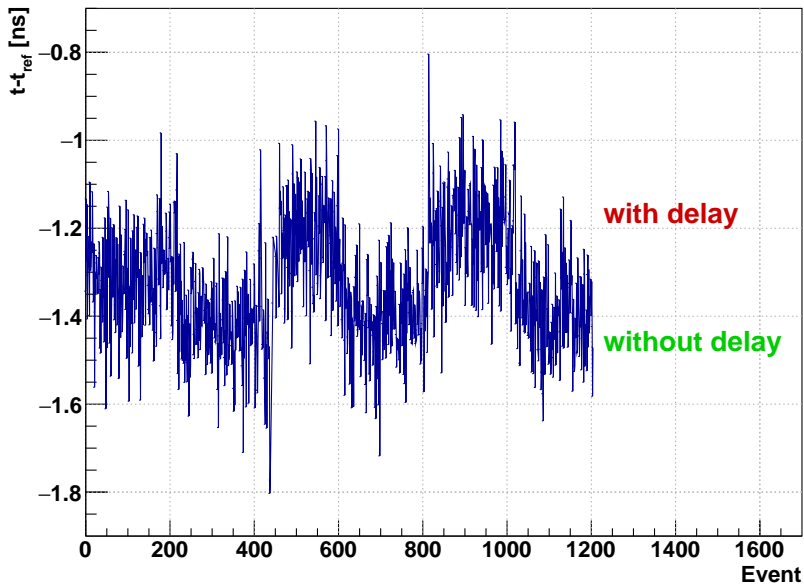
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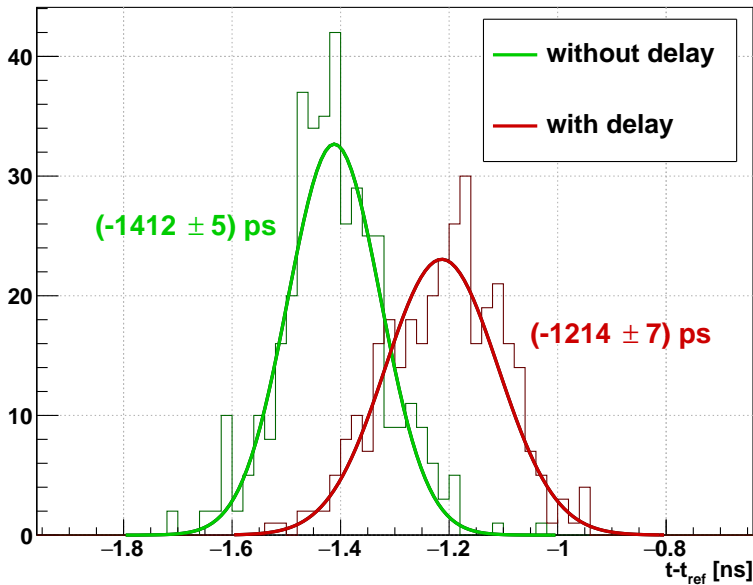
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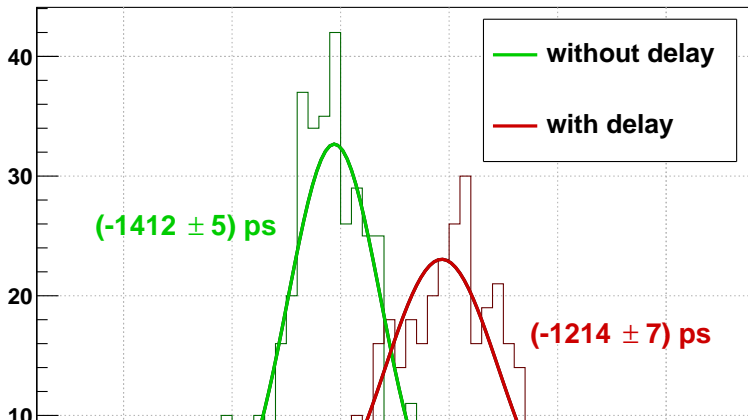
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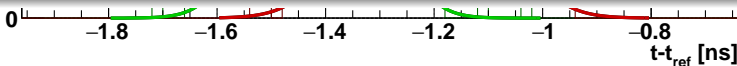


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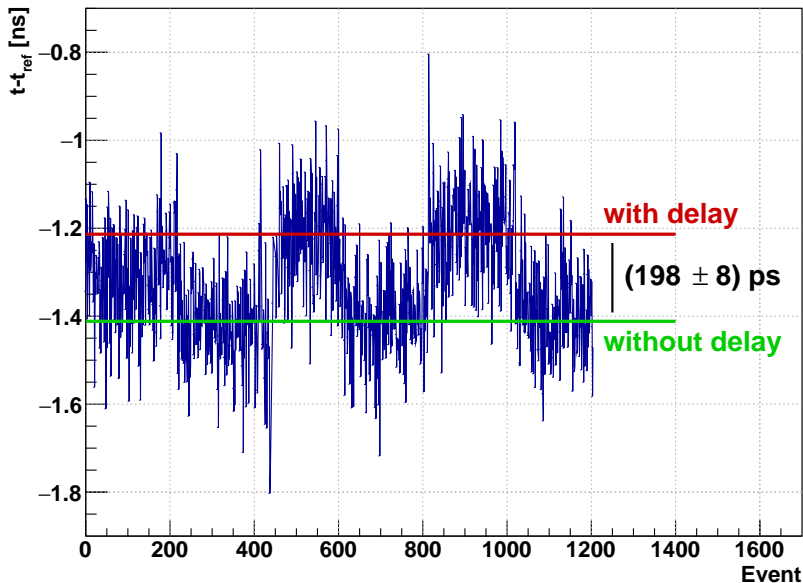


$$(t - t_{ref})_0 = -1412 \text{ ps} \neq 0$$

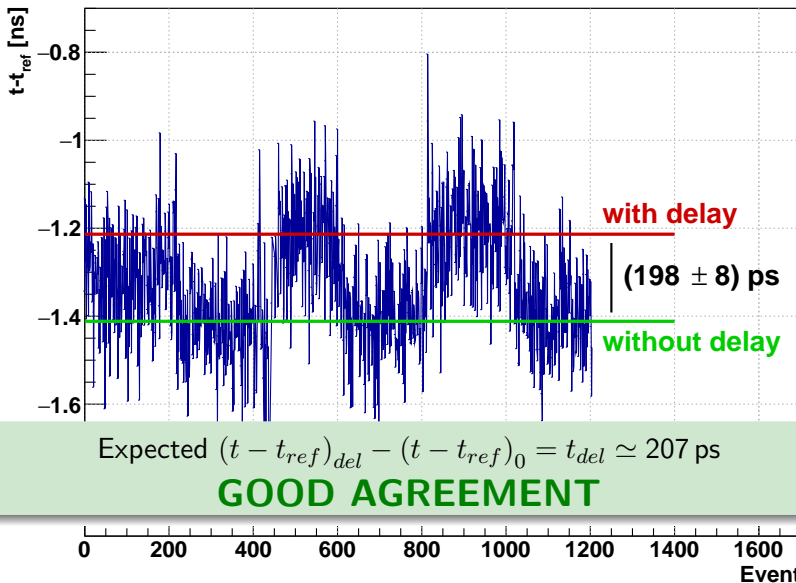
**FIXED DELAY BETWEEN TWO CHANNELS**



# Timing accuracy test in Florence



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# FAZIAPRE experiment at LNS

$^{40,48}\text{Ca} + ^{12}\text{C}$  @ 25, 40 AMeV (6 FAZIA blocks)

## Timing test

*The same timing test performed on the test bench was repeated during the mounting of FAZIAPRE experiment at LNS giving a measured delay of  $(203 \pm 13)$  ps (added delay was nominally 207 ps)*

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## Permanent infrared LED

*During the FAZIAPRE experiment, the infrared LED was mounted inside the scattering chamber and was kept on during all the shift (at a 0.1 Hz rate) to trace channel delays*

# FAZIAPRE experiment at LNS

## Results of LED analysis:

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- fixed delays up to 4 ns between channels  
**(NOT EVENT DEPENDENT):**

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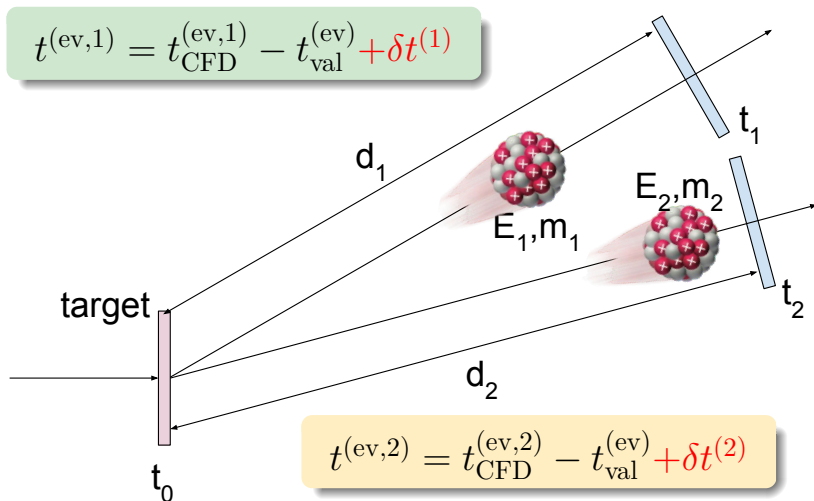
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I also studied an alternative synchronization method  
which can be used **without LED pulses**

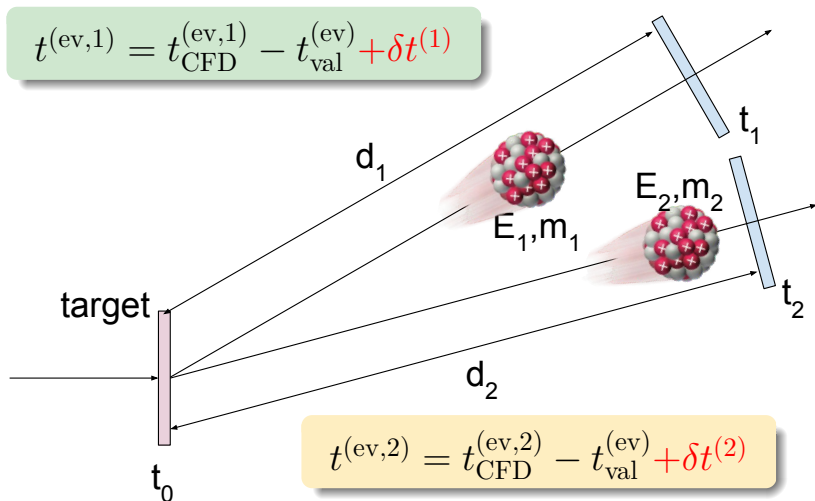
# Auto-synchronization



Take events with (at least) 2 well known particles



# Auto-synchronization



$$\Delta^{(1,2)} \equiv \delta t^{(2)} - \delta t^{(1)} = t^{(ev,2)} - t_{CFD}^{(ev,2)} - t^{(ev,1)} + t_{CFD}^{(ev,1)}$$

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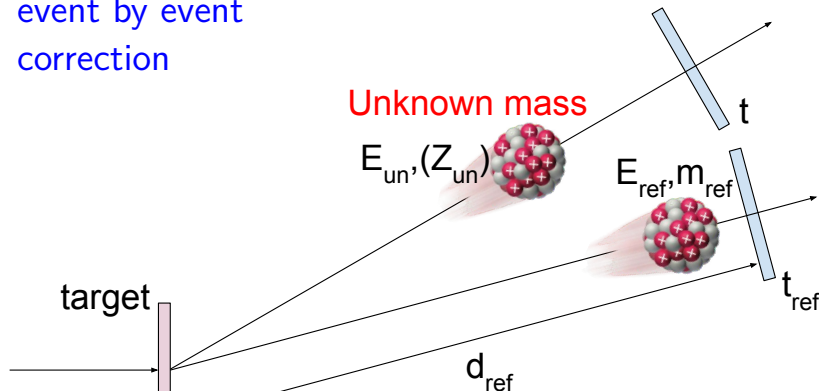
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# Time of flight identification (only $M > 1$ events)

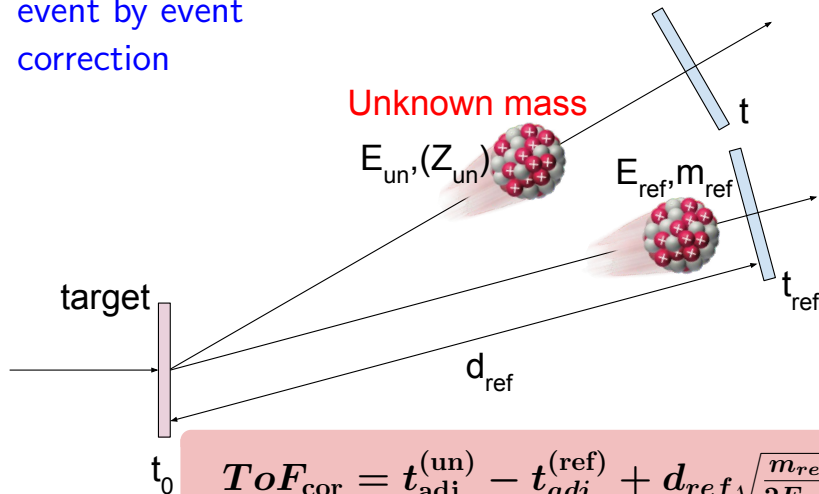
event by event  
correction



$$ToF_{RAW} = t_{CFD}^{(un)} - t_{CFD}^{(ref)} + d_{ref} \sqrt{\frac{m_{ref}}{2E_{ref}}}$$

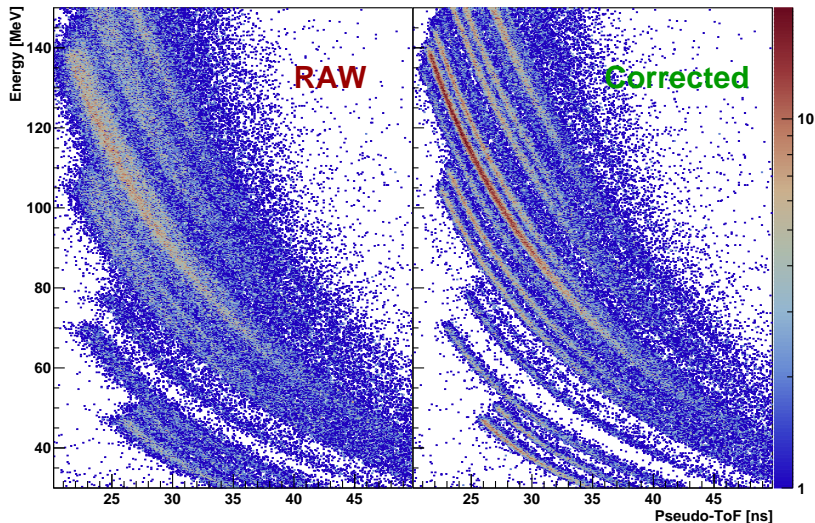
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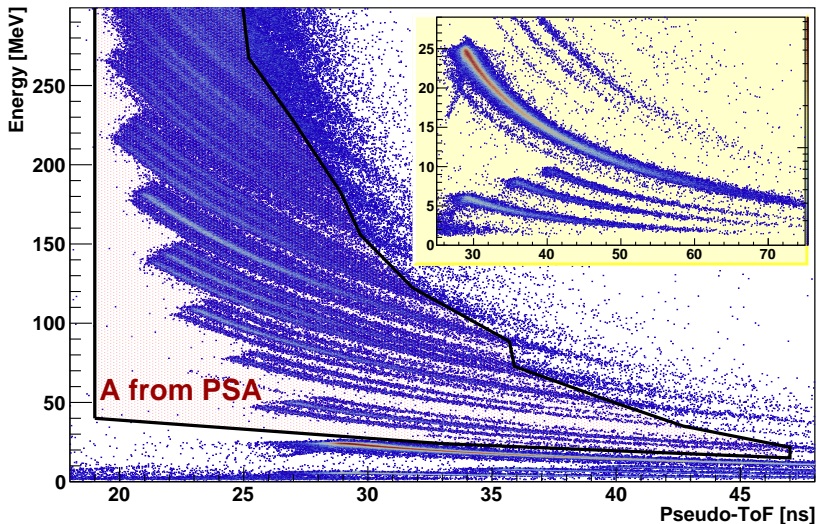




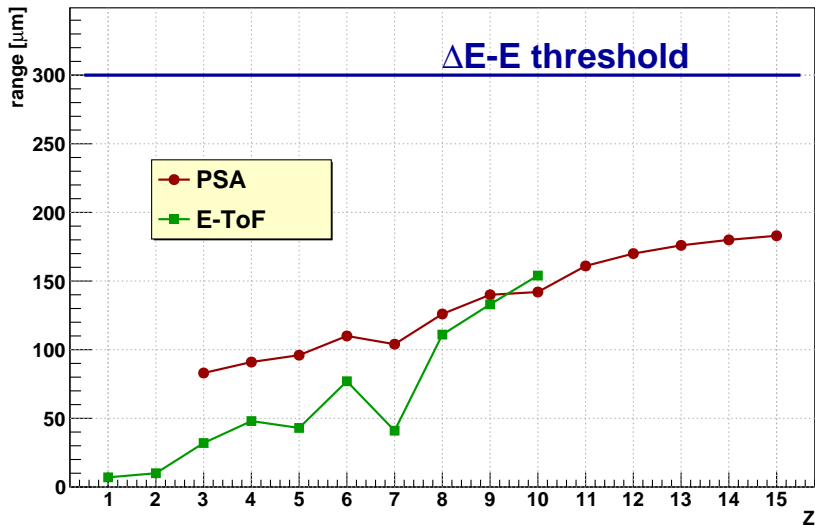
# Delay correction (LED or auto-sync)



# Final $E - ToF$ correlation



# Improvement of isotope discrimination



# Summary and conclusions

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- Opening the door to QT physics with FAZIA.

# FAZIA collaboration

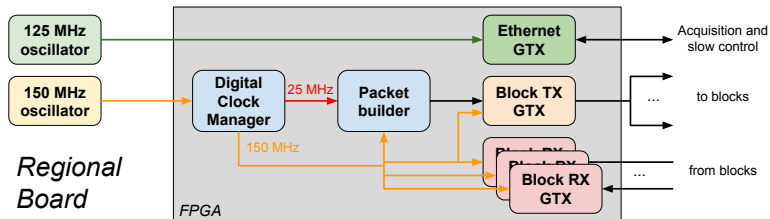


Thanks for your attention

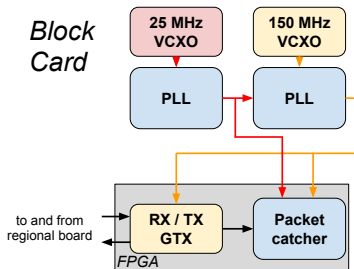
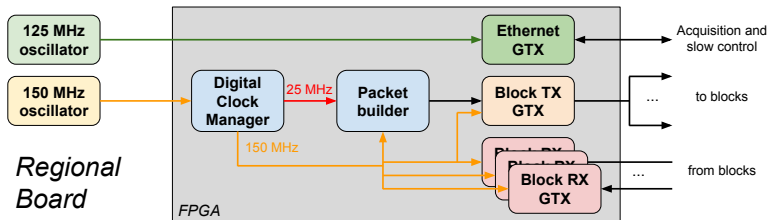


**Backup slides**

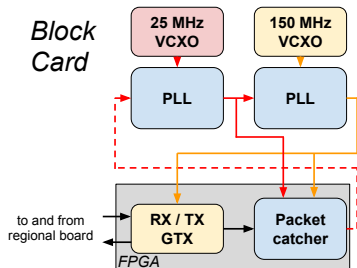
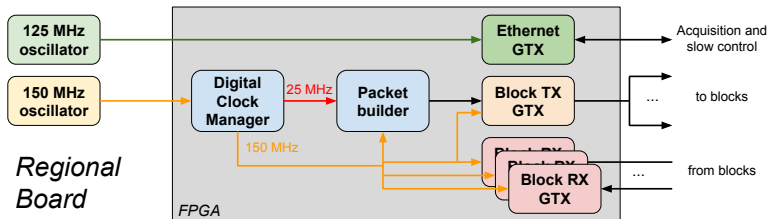
# Clock tree



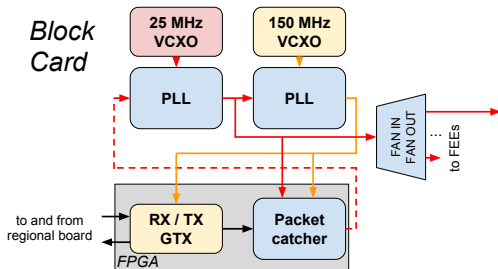
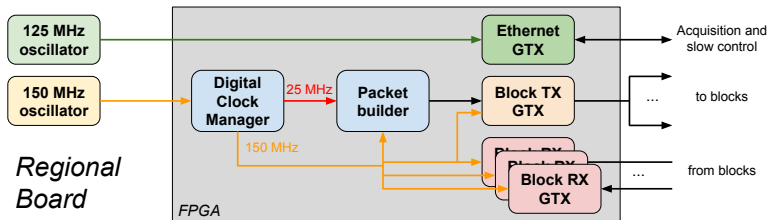
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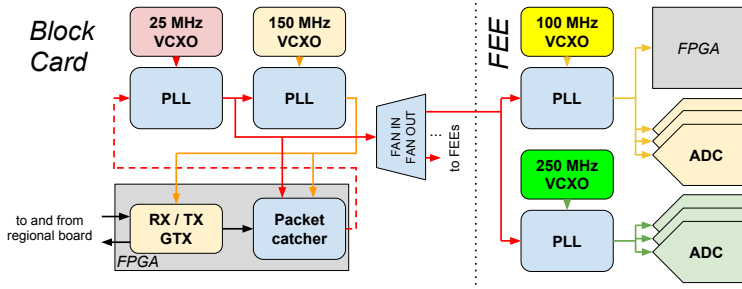
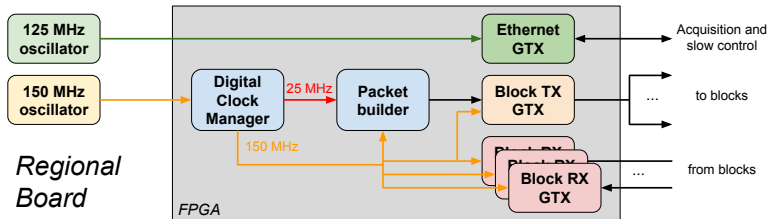
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# FAZIA collaboration

## Publications

- S. Barlini *et al*, Nucl. Instr. and Meth. A 600 (644–650), 2009
- L. Bardelli *et al*, Nucl. Instr. and Meth. A 654 (272), 2011
- S. Carboni *et al*, Nucl. Instr. and Meth. A 664 (251), 2012
- N. Le Neindre *et al*, Nucl. Instr. and Meth. A 701 (145), 2013
- S. Barlini *et al*, Nucl. Instr. and Meth. A 707 (89), 2013
- S. Barlini *et al*, Phys. Rev. C 87 (054607), 2013
- S. Piantelli *et al*, Phys. Rev. C 88 (064607), 2013
- R. Bougault *et al*, Eur. Phys. Jour. A 50 (47), 2014
- G. Pasquali *et al*, Eur. Phys. Jour. A 50 (86), 2014
- A. J. Kordyasz *et al*, Eur. Phys. Jour. A 51 (15), 2015
- F. Salomon *et al*, J. Instrum. 11 (C01064), 2016
- D. Gruyer *et al*, Nucl. Instr. and Meth. A 847 (142), 2017
- G. Pastore *et al*, Nucl. Instr. and Meth. A 860 (42), 2017