

# Detecting long-duration gravitational wave signals

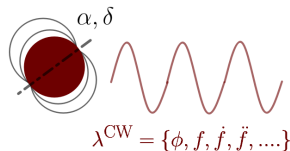
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# Continuous gravitational waves (CWs)

- quasi-sinusoidal emission
- sources: **isolated** spinning neutron stars (pulsars);
- slowly varying frequency
- almost constant amplitude
- emission mechanisms:
  - **non-axisymmetric deformations** (mass-quadrupole)
  - **currents** (current-quadrupole)
- CWs are weak but **always emitted** as long as the neutron star is deformed / the currents exist.
- (almost) **infinite duration**



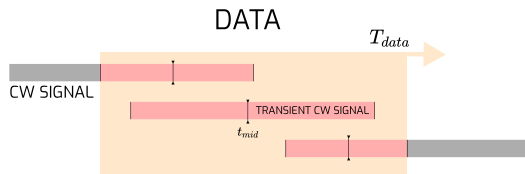
**CWs** are **long-duration** (from months to years) gravitational waves from pulsars with almost **unchangeable amplitude** and slowly varying frequency.

## Transient CW-signal: physical sources

- **finite duration**: from **days** to **months**.
- a phase-coherent CW is **broken** due to a *sudden spin-up* (**glitch**) in the pulsar's timing.
- glitching sources:
  - observed: Vela, Crab, the fastest young pulsar **PSR J0537-6910** (Antonopoulou et al., 2018);
  - expected: the **newly born pulsars** according to the *r-mode emission* theory (Andersson, 1998)
- actual **time parameters** of a signal are **unknown** from electromagnetic observations.
- Example: no glitch information for the regularly glitching pulsar **PSR J0537-6910** in **O1 and O2 LIGO** runs.

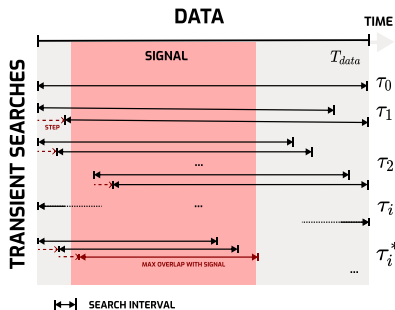
In the **absence of the glitch information**, we need to search for a **transient signal** with **unknown** start time and duration in the available data.

# Transient CW-signal in the data



Transient CW-signal in the data is when the signal overlaps the data from detectors.

# Matched-filtering transient search



**Figure:** Scheme of the matched-filtering transient searches. *Figure credit: A. Vishnevskaya*

## Method (Prix et al., 2011)

- Searches over various **time-intervals** with **steps** (shifts) in time and over the frequency range with an appropriate **template grid**.
- From every search the maximum statistical value ( $2\mathcal{F}$ ) is recorded.

## Detection

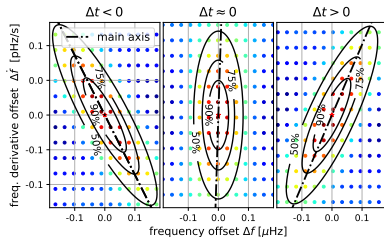
- Decision about the signal detection is based on the maximum value over all the searches.
- The corresponding search duration is taken as an **estimate** of the signal duration.

The **matched-filtering** is the **most sensitive method** to search for long-duration transients with unknown signal frequency and time parameters

# New approach: Estimation of the middle time of a transient signal

## Estimation of the middle time of a signal

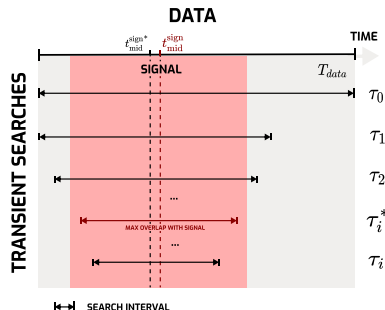
- 1 There is a transient signal of unknown time parameters in the data.
- 2 Perform a **template search** in frequency over all the available data.
- 3 The SNR-reduction profile from the search has a **slope** depending on the distance between the reference time of the search and the mid-time of the signal.



**Figure:** The SNR-reduction profile for three values of the distance  $\Delta t$  between the reference time of the search and the mid-time of the signal.

Based on the information recovered in an all data search, one can **estimate the mid-time** of a transient signal.

# Post-following transient searches



**Figure:** Scheme of the post-following transient searches around an estimated middle time of a signal *Figure credit: A. Vishnevskaya*

## Method (Fesik and Papa, 2022)

- Searches over various time-intervals with the **common reference time** – the estimated middle time of a signal, and over the frequency range with an appropriate template grid.
- From every search is recorded the maximal statistical value ( $2\mathcal{F}$ ).

## Detection

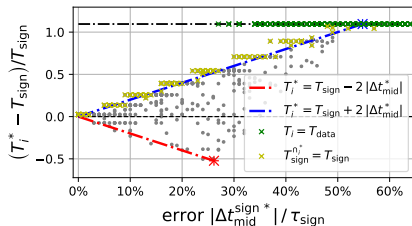
- Decision about the signal detection is based on the maximum value from the searches.
- The corresponding search duration is taken as an **estimate** of the signal duration.

With the estimated middle time of a signal, we are able to **localise the signal in duration** and recover its signal-to-noise ratio.

# The recovered SNR and signal duration

## Estimation of the signal duration

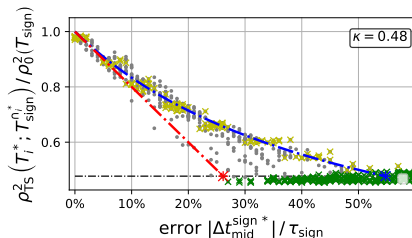
- an **error in the estimation** of the middle time produces **over- or underestimate** of the signal duration.
- the larger is an error, the greater is the deviation of the recovered duration.



**Figure:** The recovered signal duration  $T_i^*$  as a function of the errors  $|\Delta t_{\text{mid}}^*|$  in the estimate of the mid-time of a transient signal.

## Recovered transient-signal SNR

- The **minimum** recovered signal SNR from the transient searches is limited by the value recovered in the initial search.
- The deviation of the **recovered signal duration** is also **limited**.



**Figure:** The value  $\rho_{\text{TS}}^2(T_i^*; T_{\text{sign}}^0)/\rho_0^2(T_{\text{sign}})$  as a function of the errors  $|\Delta t_{\text{mid}}^*|$ .



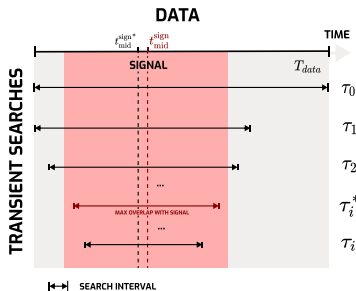
# Application of the method

The method can be applied in the **absence of the glitch information** during the observation period.

## Possible objects to search

- Glitching pulsar **PSR J0537-6910**: there were **no timing information** about the glitches during the **O1 and O2 LIGO** observing runs.
- Other objects with accident or periodic **glitch activity**, which have not been explored yet.

# Summary



- ① **Initial search** over all available data in the frequency range.
- ② **Estimation of the middle time** of a signal.
- ③ **Search with various time-intervals** around the estimated middle time.
- ④ **The best statistical result** is recorded as signal's candidate and the corresponding interval is an estimate of the signal's duration.

We propose the method, which allows to **localise a transient CW-signal in time** and recover the signal's SNR in the **absence of the glitch information**.



# References

- N. Andersson. A New class of unstable modes of rotating relativistic stars, 1998.
- D. Antonopoulou, C. M. Espinoza, L. Kuiper, and N. Andersson. Pulsar spin-down: the glitch-dominated rotation of PSR J0537-6910, 2018.
- L. Fesik and M. A. Papa. Search for a transient signal. Part I. Tools. *Phys. Rev.*, 2022. to be published.
- R. Prix, S. Giampanis, and C. Messenger. Search method for long-duration gravitational-wave transients from neutron stars. *Phys. Rev.*, D84:023007, 2011. doi: 10.1103/PhysRevD.84.023007.



Thank you for your attention!

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