

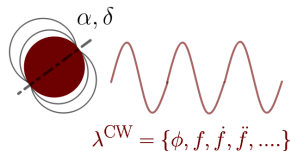
Detecting long-duration gravitational wave signals

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Continuous gravitational waves (CWs)

- quasi-sinusoidal emission
- sources: **isolated** spinning neutron stars (pulsars);
- slowly varying frequency
- almost constant amplitude
- emission mechanisms:
 - **non-axisymmetric deformations** (mass-quadrupole)
 - **currents** (current-quadrupole)
- CWs are weak but **always emitted** as long as the neutron star is deformed / the currents exist.
- (almost) **infinite duration**



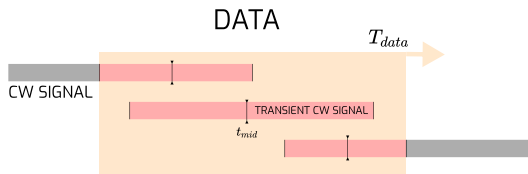
CWs are **long-duration** (from months to years) gravitational waves from pulsars with almost **unchangeable amplitude** and slowly varying frequency.

Transient CW-signal: physical sources

- **finite duration**: from **days** to **months**.
- a phase-coherent CW is **broken** due to a *sudden spin-up* (**glitch**) in the pulsar's timing.
- glitching sources:
 - observed: Vela, Crab, the fastest young pulsar **PSR J0537-6910** (Antonopoulou et al., 2018);
 - expected: the **newly born pulsars** according to the *r-mode emission* theory (Andersson, 1998)
- actual **time parameters** of a signal are **unknown** from electromagnetic observations.
- Example: no glitch information for the regularly glitching pulsar **PSR J0537-6910** in **O1 and O2 LIGO** runs.

In the **absence of the glitch information**, we need to search for a **transient signal** with **unknown** start time and duration in the available data.

Transient CW-signal in the data



Transient CW-signal in the data is when the signal overlaps the data from detectors.

Matched- Itering transient search

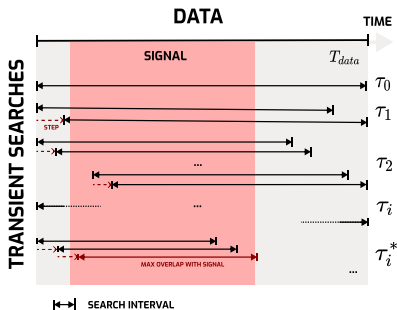


Figure: Scheme of the matched-filtering transient searches. Figure credit: A. Vishnevskaya

Method (Prix et al., 2011)

- Searches over various **time-intervals** with **steps** (shifts) in time and over the frequency range with an appropriate **template grid**.
- From every search the maximum statistical value ($2F$) is recorded.

Detection

- Decision about the signal detection is based on the maximum value over all the searches.
- The corresponding search duration is taken as an **estimate** of the signal duration.

The **matched- Itering** is the **most sensitive method** to search for long-duration transients with unknown signal frequency and time parameters

New approach: Estimation of the middle time of a transient signal

Estimation of the middle time of a signal

- 1 There is a transient signal of unknown time parameters in the data.
- 2 Perform a **template search** in frequency over all the available data.
- 3 The SNR-reduction profile from the search has a **slope** depending on the distance between the reference time of the search and the mid-time of the signal.

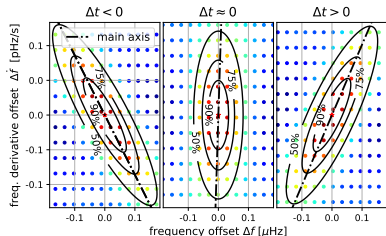


Figure: The SNR-reduction profile for three values of the distance Δt between the reference time of the search and the mid-time of the signal.

Based on the information recovered in an all data search, one can **estimate the mid-time** of a transient signal.

Post-following transient searches

Figure: Scheme of the post-following transient searches around an estimated middle time of a signal *Figure credit: A. Vishnevskaya*

Method (Fesik and Papa, 2022)

- Searches over various time-intervals with the **common reference time** – the estimated middle time of a signal, and over the frequency range with an appropriate template grid.
- From every search is recorded the maximal statistical value ($2F$).

Detection

- Decision about the signal detection is based on the maximum value from the searches.
- The corresponding search duration is taken as an **estimate** of the signal duration.

With the estimated middle time of a signal, we are able to **localise the signal in duration** and recover its signal-to-noise ratio.

The recovered SNR and signal duration

Estimation of the signal duration

- an **error in the estimation** of the middle time produces **over- or underestimate** of the signal duration.
- the larger is an error, the greater is the deviation of the recovered duration.

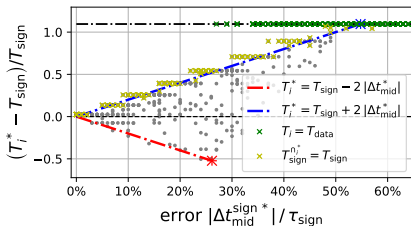


Figure: The recovered signal duration T_i^* as a function of the errors $|\Delta t_{\text{mid}}^{\text{sign}*}|$ in the estimate of the mid-time of a transient signal.

Recovered transient-signal SNR

- The **minimum** recovered signal SNR from the transient searches is limited by the value recovered in the initial search.
- The deviation of the **recovered signal duration** is also **limited**.

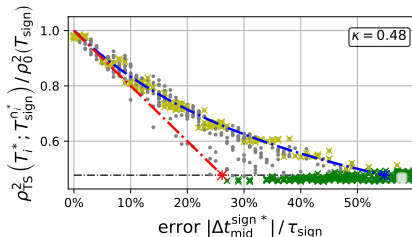


Figure: The value $\rho_{\text{TS}}^2(T_i^*; T_{\text{sign}}^{\text{ri}}) / \rho_0^2(T_{\text{sign}})$ as a function of the errors $|\Delta t_{\text{mid}}^{\text{sign}*}|$.

Application of the method

The method can be applied in the **absence of the glitch information** during the observation period.

Possible objects to search

- Glitching pulsar **PSR J0537-6910**: there were **no timing information** about the glitches during the **O1 and O2 LIGO** observing runs.
- Other objects with accident or periodic **glitch activity**, which have not been explored yet.

Summary

- 1 Initial search over all available data in the frequency range.
- 2 Estimation of the middle time of a signal.
- 3 Search with various time-intervals around the estimated middle time.
- 4 The best statistical result is recorded as signal's candidate and the corresponding interval is an estimate of the signal's duration.

We propose the method, which allows to localise a transient CW-signal in time and recover the signal's SNR in the absence of the glitch information.

References

- N. Andersson. A New class of unstable modes of rotating relativistic stars, 1998.
- D. Antonopoulou, C. M. Espinoza, L. Kuiper, and N. Andersson. Pulsar spin-down: the glitch-dominated rotation of PSR J0537-6910, 2018.
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- R. Prix, S. Giampanis, and C. Messenger. Search method for long-duration gravitational-wave transients from neutron stars. *Phys. Rev.*, D84:023007, 2011. doi: 10.1103/PhysRevD.84.023007.

Thank you for your attention!

