

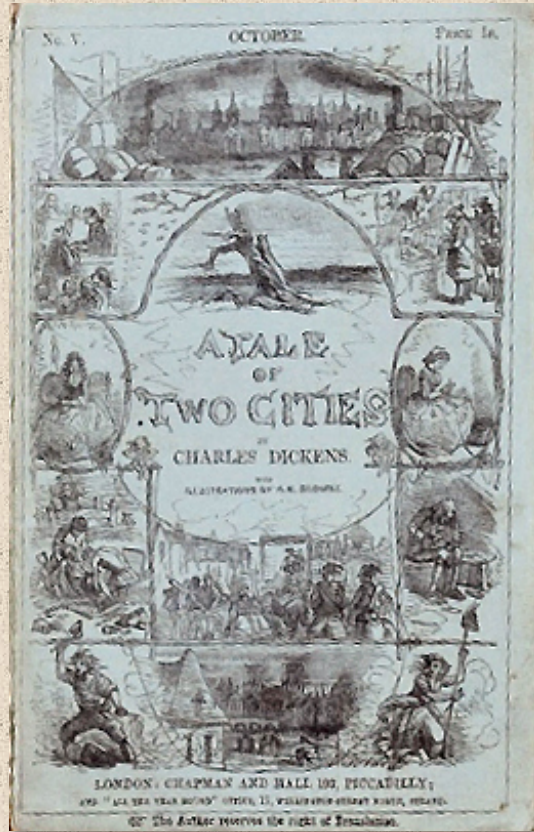
Cosmology in the multimessenger era: an overview

PUMA22
Sestri Levante
25/09/2022

Enzo Branchini



Charles Dickens
1812-1870



A Tale of Two Cities.

1859

Over 200 million copies

One of the best selling
novels of all times.

It was the best of times,
it was the worst of times,
it was the age of wisdom,
it was the age of foolishness,
it was the epoch of belief,
it was the epoch of incredulity,
it was the season of Light,
it was the season of Darkness,
it was the spring of hope,
it was the winter of despair,

Cosmology is the scientific study of the Universe as a unified whole, from its earliest moments through its evolution to its ultimate fate. (Britannica)

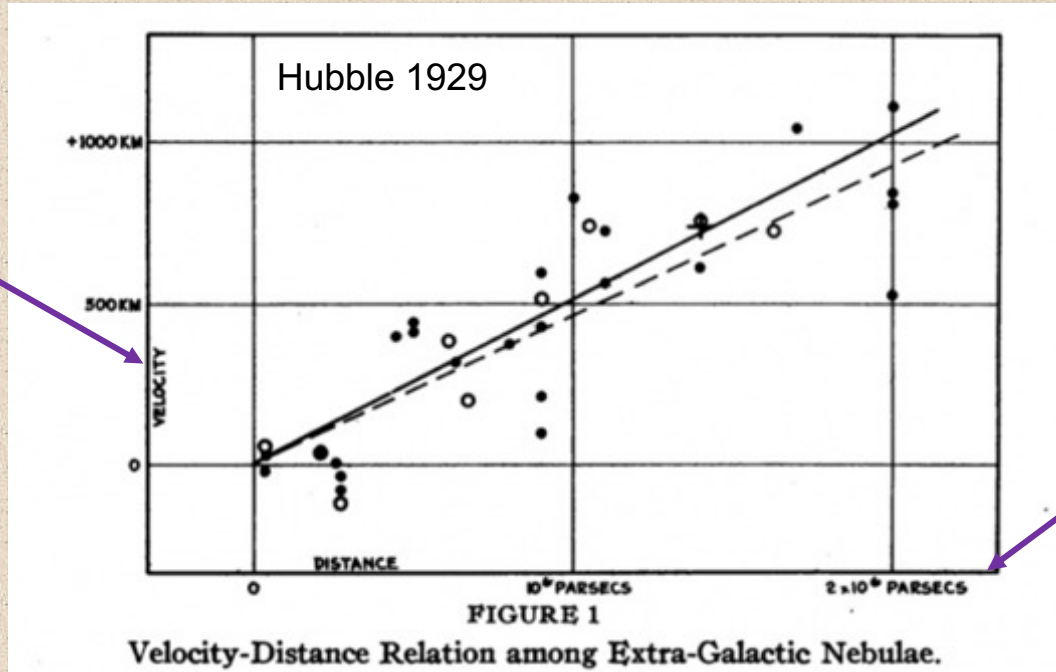
Physical Cosmology deals with the observable Universe. Which is assumed to be a realization of a stationary random process. A fair sample. *One only*. To compare our models with. Hence the importance for Cosmology to go **multi-wavelength**, **multi-probe** and now **multi-messenger**.

Λ CDM “standard” model.

Observational pillars. Hubble law.

Redshifts from
optical spectra

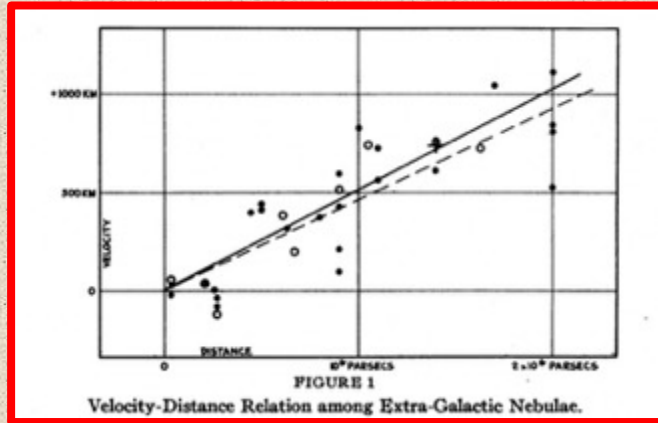
Distance
indicator:
Cepheids



Distances from
Optical
photometry

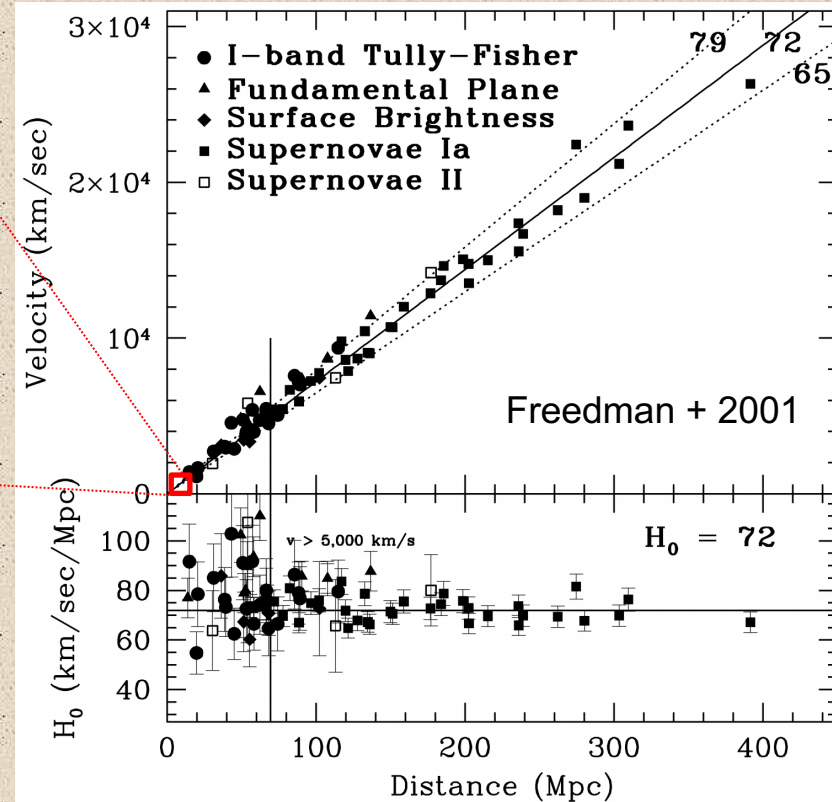
Λ CDM “standard” model.

Observational pillars. Hubble Law.

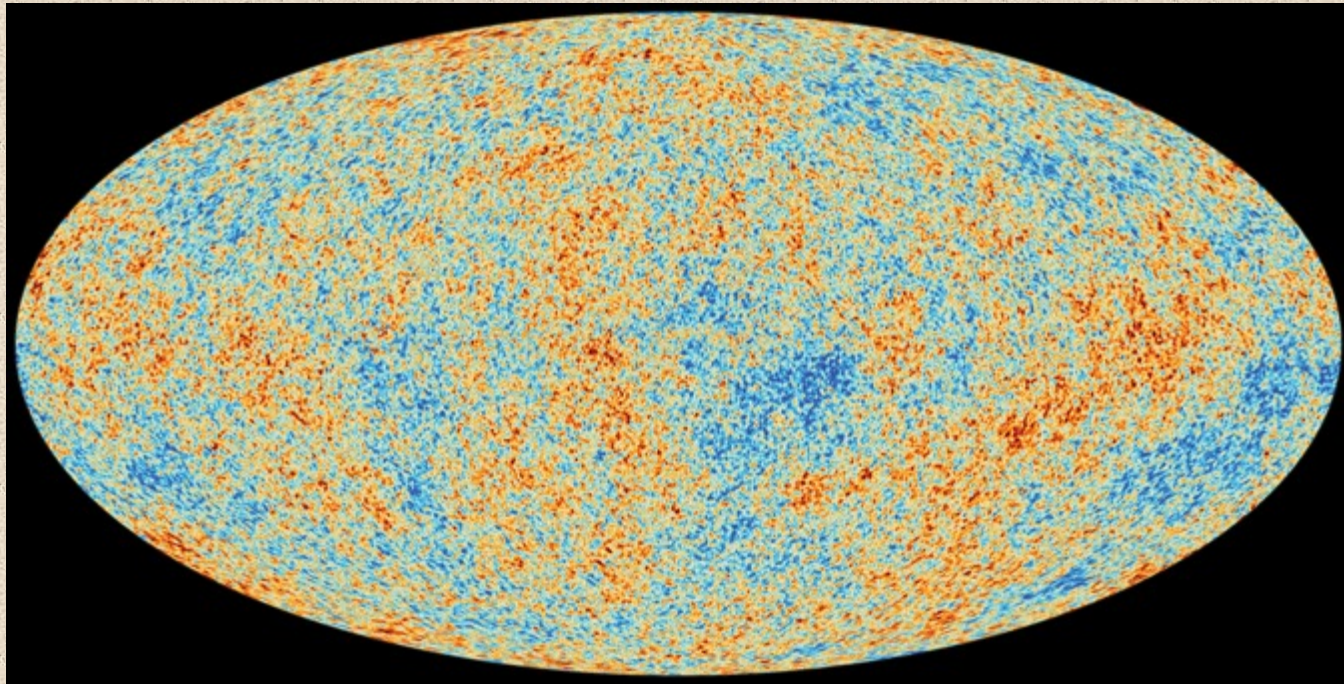


Optical + Near Infrared

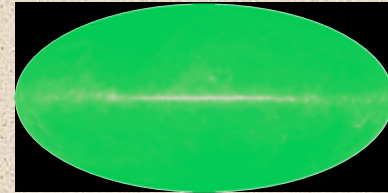
Several distance indicators



Λ CDM “standard” model. Observational pillars. CMB.

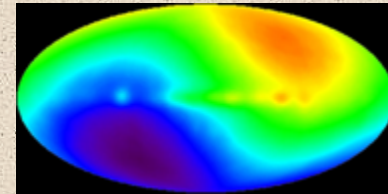


-



Monopole

-



Dipole

Microwave band

Q1422+231 - 18.8 hours
with HIRES (Keck I)

Flux

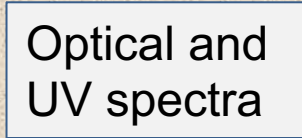
Ly α Forest

QSO's
Ly α emission

Wavelength (\AA)

Ellison+ 2000

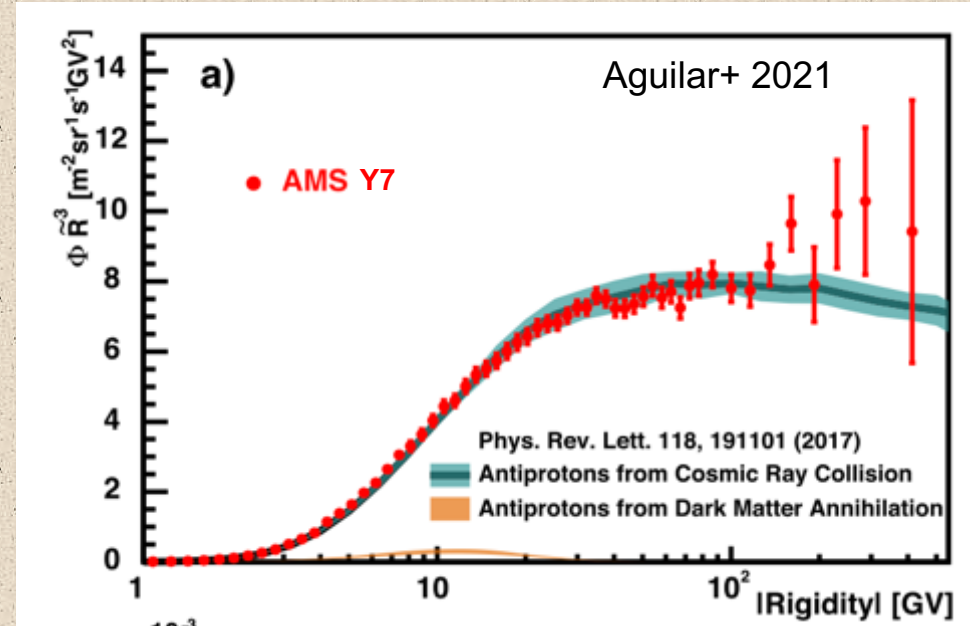
This figure is a spectral plot showing the flux of a quasar (Q1422+231) as a function of wavelength. The x-axis represents Wavelength in Angstroms (\AA), ranging from approximately 4200 to 6600. The y-axis represents Flux, ranging from 0 to 6000. The plot shows a broad Ly α emission line centered around 5600 \AA , which is labeled 'QSO's Ly α emission'. A region of the spectrum between approximately 4300 and 5500 \AA is labeled 'Ly α Forest', indicating the presence of numerous absorption lines from intervening galaxies. The text 'Q1422+231 - 18.8 hours with HIRES (Keck I)' is in the upper left, and 'Ellison+ 2000' is in the lower left.



Pettini+ 2008

Λ CDM “standard” model.

Key observations. Antimatter abundance.



Charged particles

Λ CDM

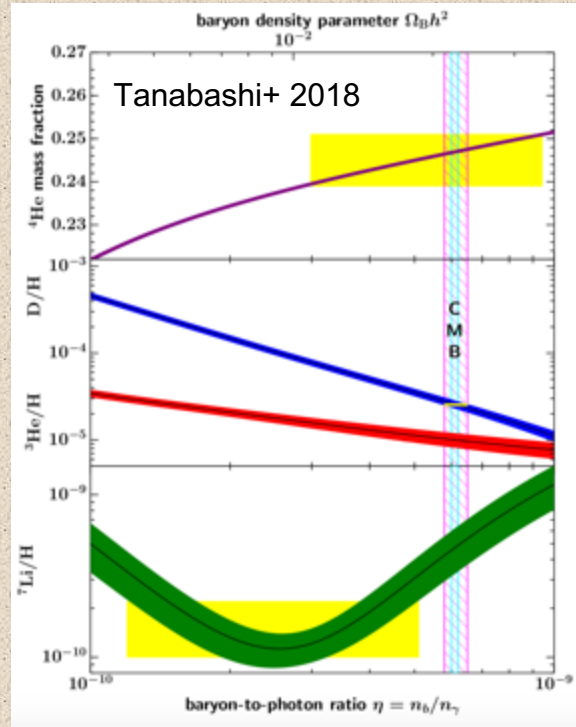
$\Omega_b, \Omega_M, \Omega_\gamma, \Omega_\Lambda, A_s, n_s, H_0, \tau \mid \Omega_v \sim 0, \text{Inflation}$

Λ CDM is a simple and yet very successful model.
It is described by a small number of free parameters.
Is it correct ? Is it satisfactory ?

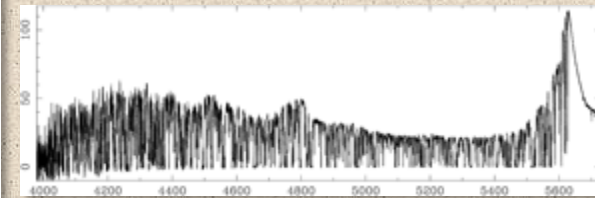
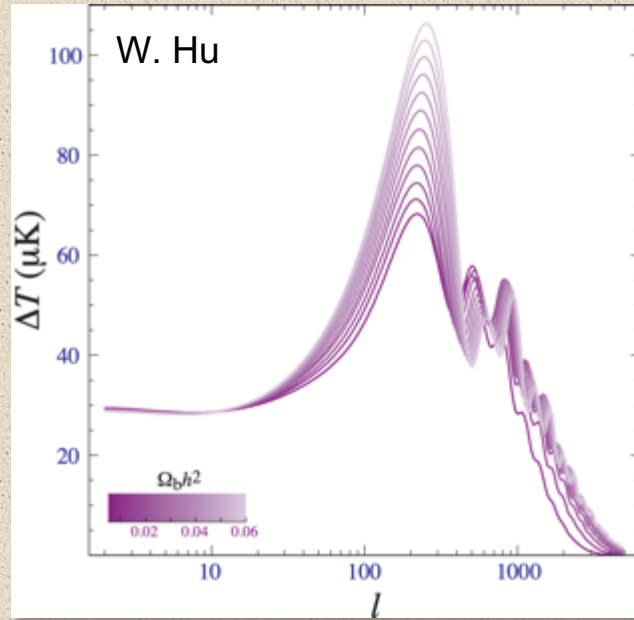
«.....the empirical case that the Λ CDM theory is a good approximation to reality remains compelling. But....we have empirical evidence that there is a still better theory to be found.»

J. Peebles 2022

Λ CDM predictions consistently agrees with a wide range diverse observations. Ex 1. baryon density.



$$\Omega_b h^2 = 0.02237 \pm 0.00015$$

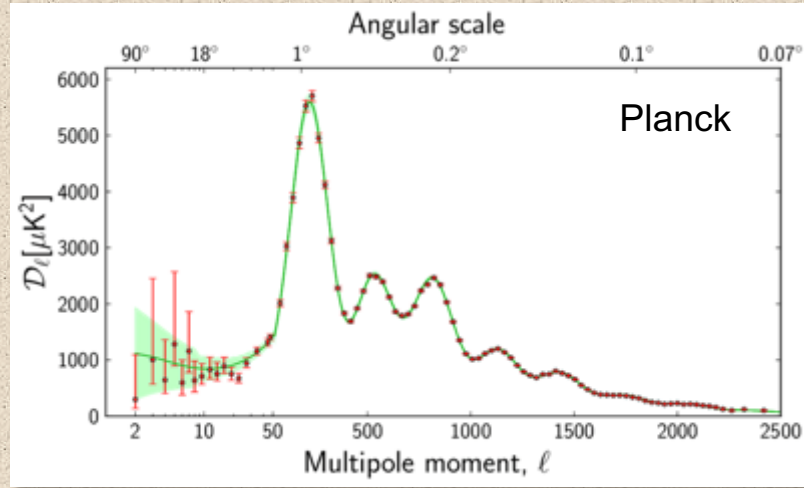


Lyman- α forest $z \sim 1$

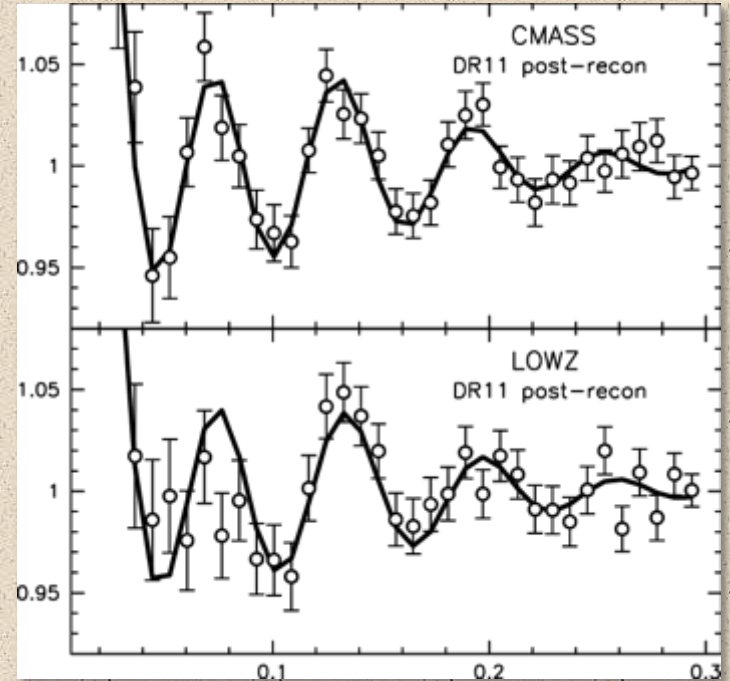
Light elements abundance $z \sim 10^9$

CMB angular spectrum $z \sim 10^3$

Λ CDM predictions consistently agrees with a wide range diverse observations. Ex 2. BAOs



In the plasma @ $z \sim 1000$



In the spatial distribution
of galaxies @ $z \sim 1$

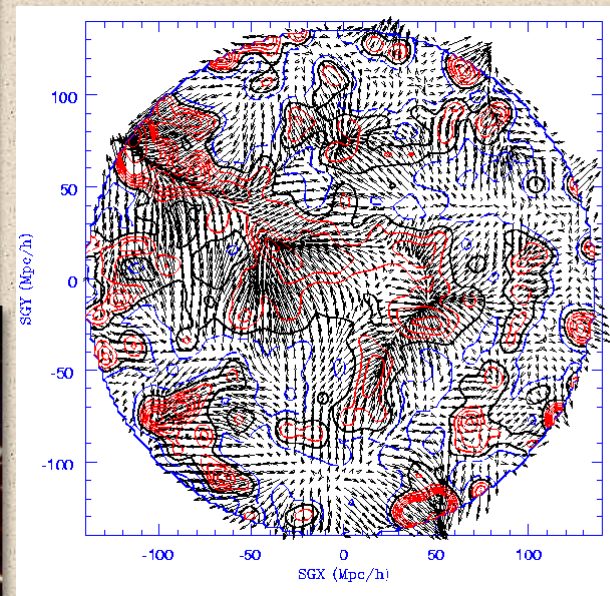
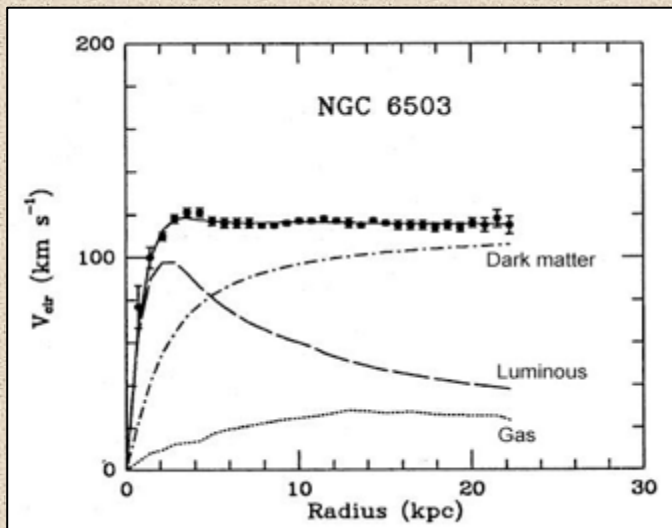
Λ CDM

...provides a good description to a large number of independent observations over a broad range of spatial and temporal scales.

$\Lambda + \text{CDM}$

....but it relies on two fundamental ingredients of which we have convincing phenomenological evidence but little physical understanding.

Λ CDM



+ clustering evolution...

Λ CDM

Compelling evidence from dynamical anomalies on a broad range of scales.

Some phenomenological constraints (\sim collisionless, massive, \sim stable, cold).

No detections from either direct or indirect searches so far.

No unique theoretical prediction.

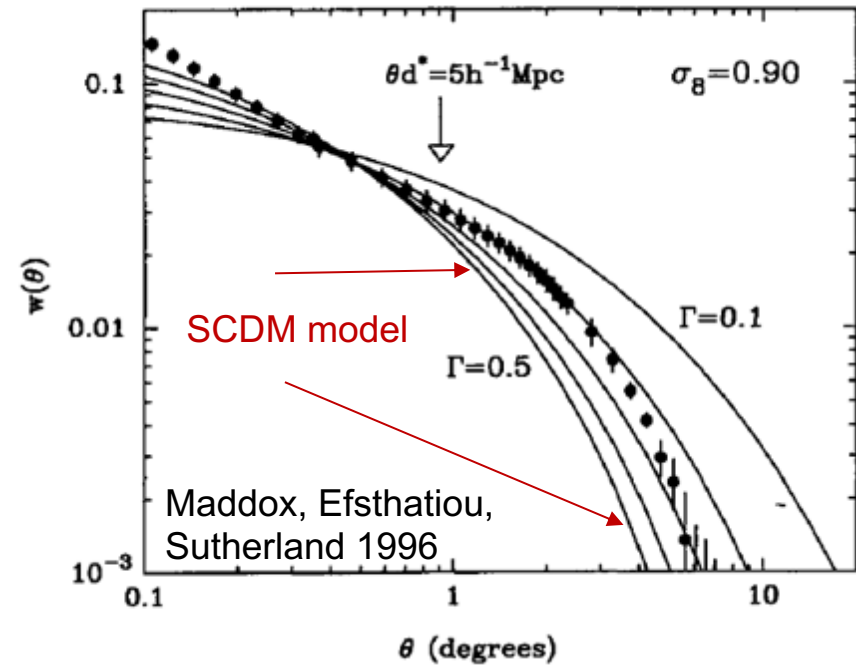
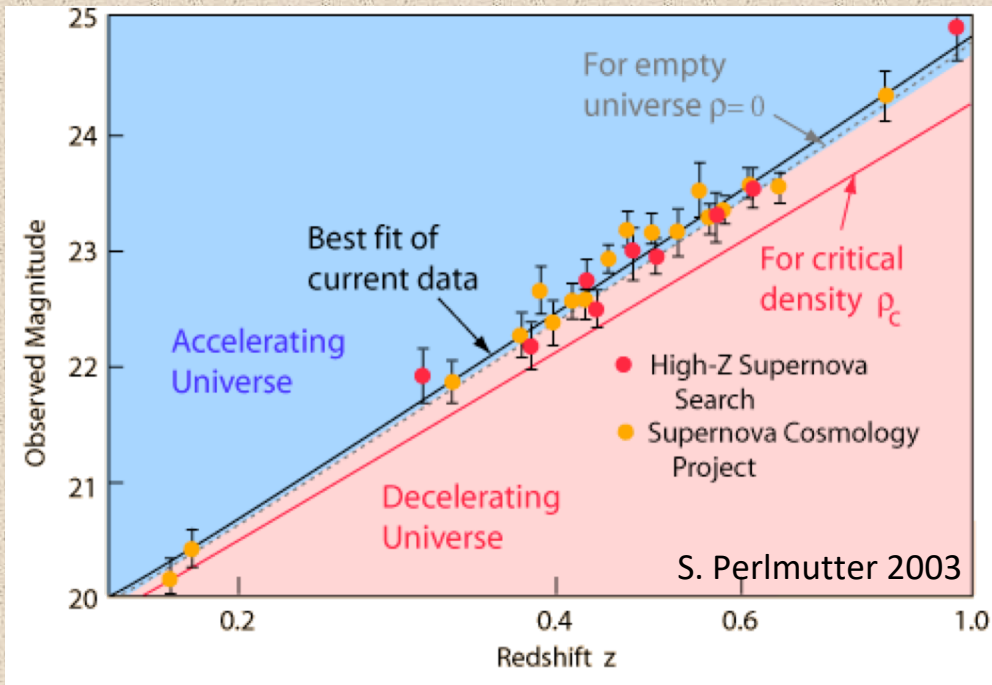
Are we capturing the complexity of the dark sector ?

Open issue of fundamental importance

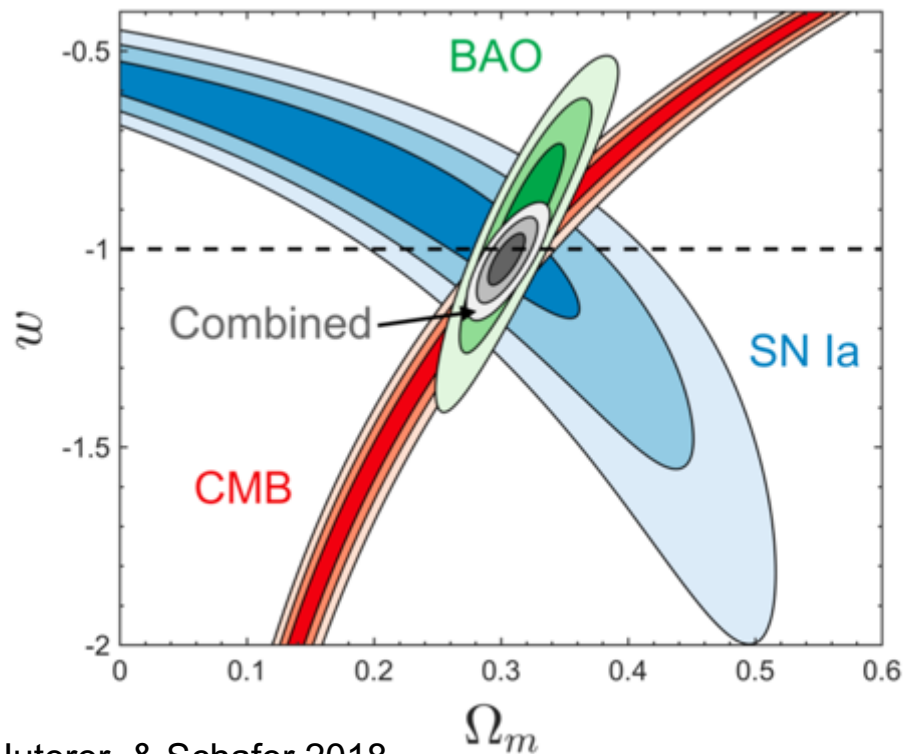
Λ CDM

SN1a - distances

Galaxies - clustering



Λ CDM



Huterer & Schafer 2018

Λ CDM

Compelling evidence from different probes.

Some observational constraints.

Broad range of theoretical predictions, from Dark Energy to Modified Gravity models.

2 separate issues: the nature of the accelerated expansion and cosmological relevance of the zero-point vacuum fluctuations.

Open issue of fundamental importance

Λ CDM

Λ CDM theory is probably a good but unsatisfactory approximation to reality.

How good this approximation is ?

Anomalies, tensions, inconsistencies, if genuine, can show the way to a better model.

The many Λ CDM crises, tensions and anomalies.

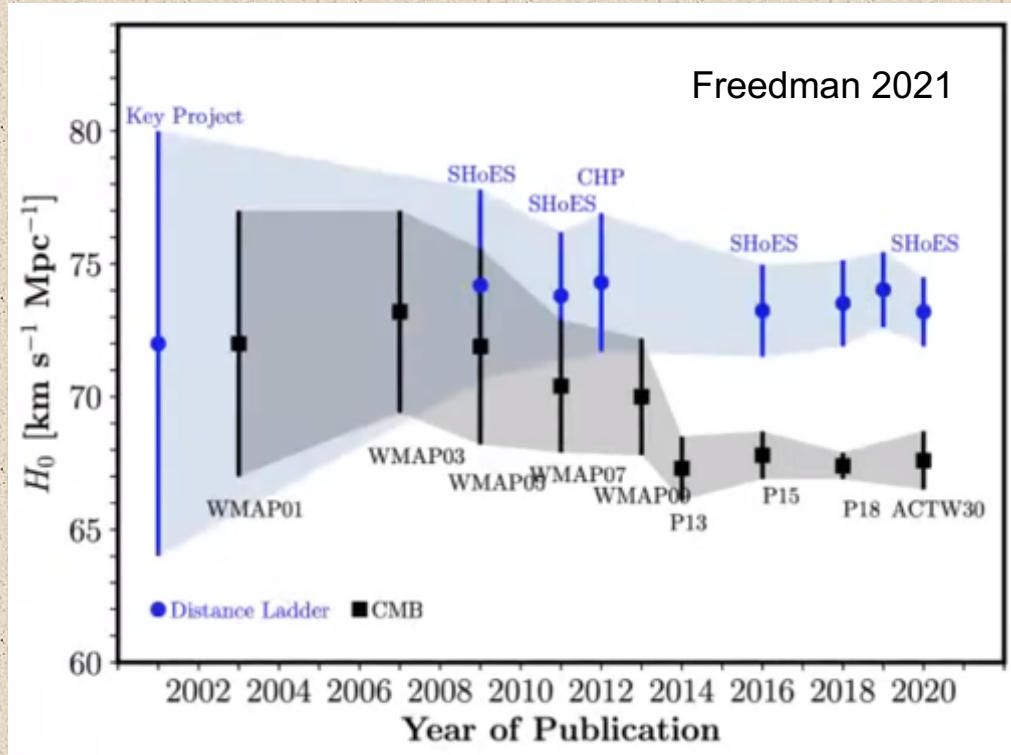
- H_0 tension.
- σ_8 tension.
- CMB anomalies.
- Cosmological dipoles and bulk flows.
- Missing baryons.
- Too big to fail / substructures abundance.
- Core-cusp problem.
- Supermassive black holes.
- Lithium abundance.
-

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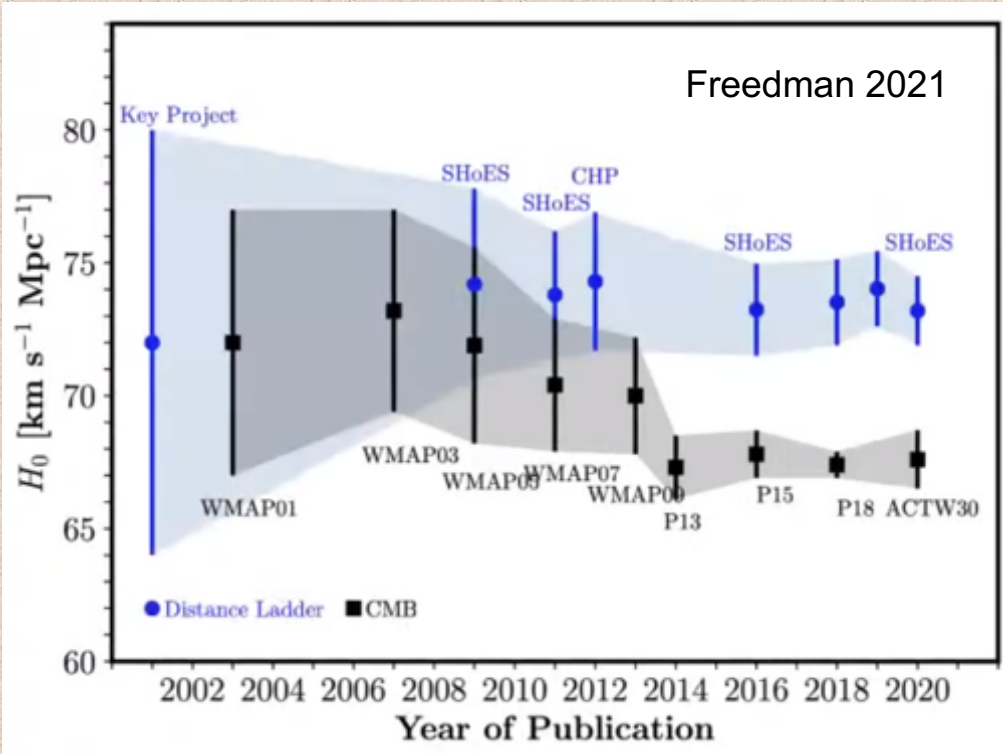
Λ CDM crises, tensions and anomalies.

H_0 tension.



Λ CDM crises, tensions and anomalies.

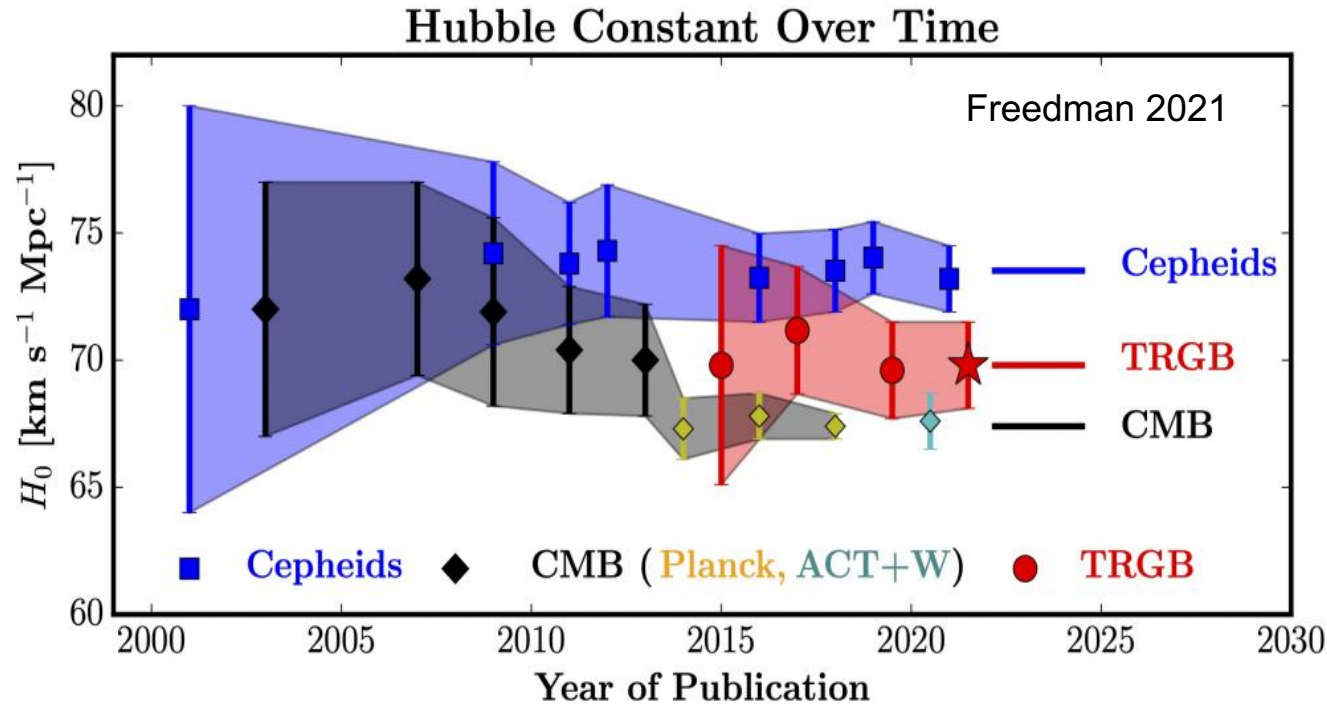
H_0 tension.



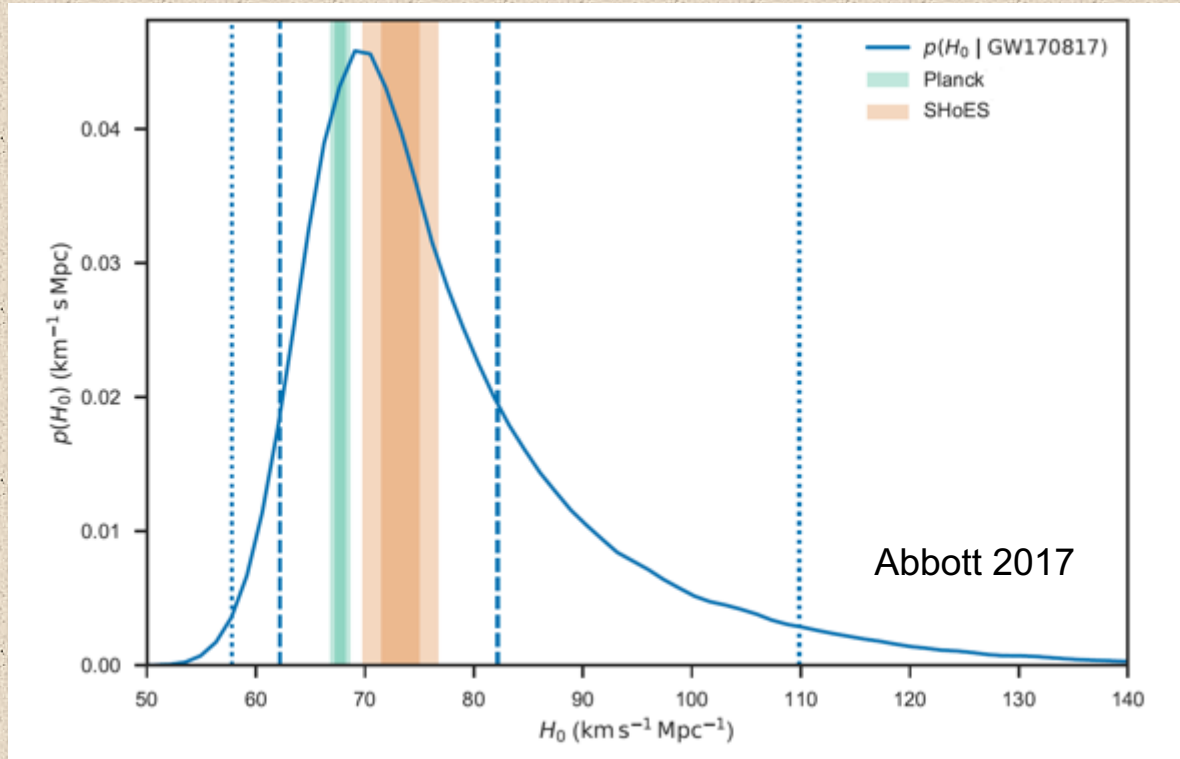
**This is a 10%
discrepancy over a
factor 1000 in the
cosmic expansion !!!!!**

Λ CDM crises, tensions and anomalies.

H_0 tension.



Λ CDM crises, tensions and anomalies.

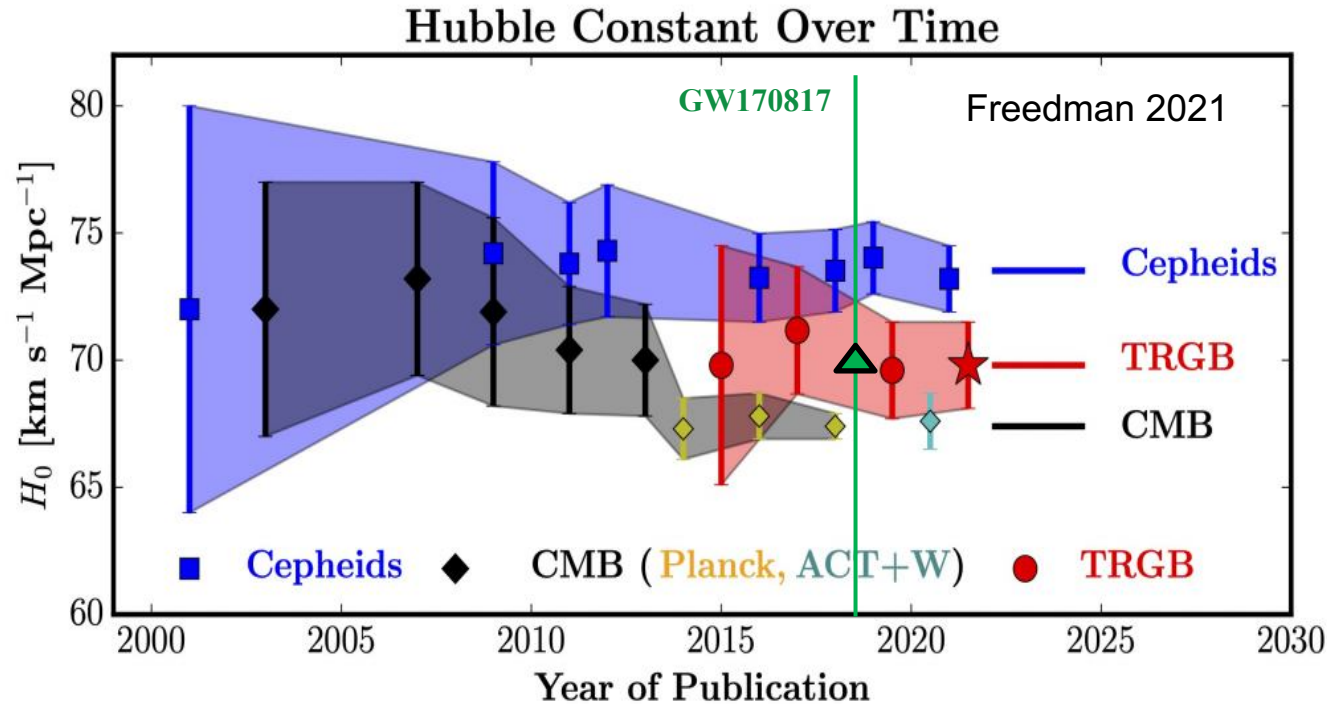


H_0 tension.

GW from bright
and dark sirens

Λ CDM crises, tensions and anomalies.

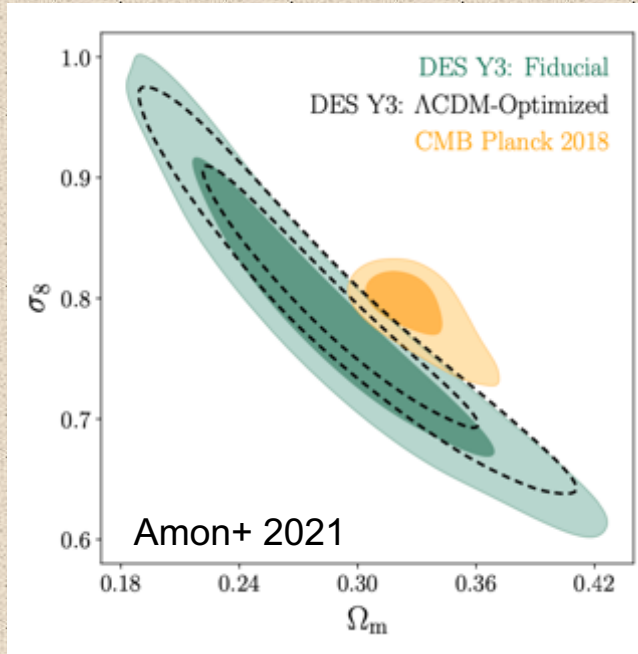
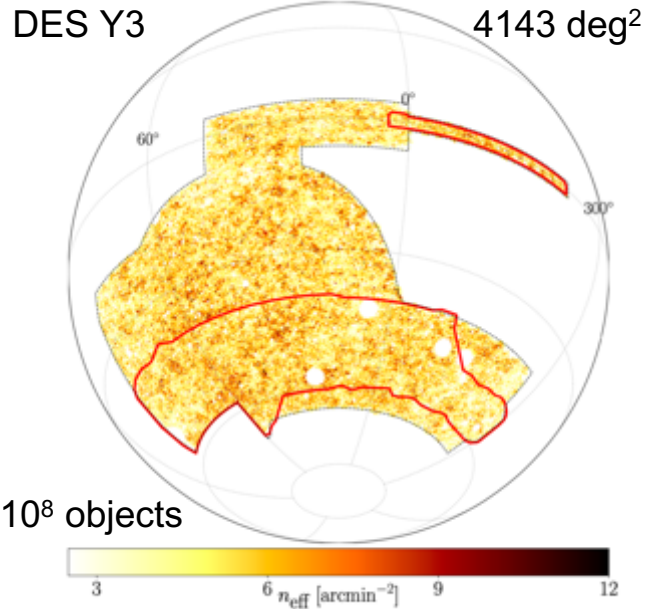
H_0 tension.



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Λ CDM crises, tensions and anomalies.



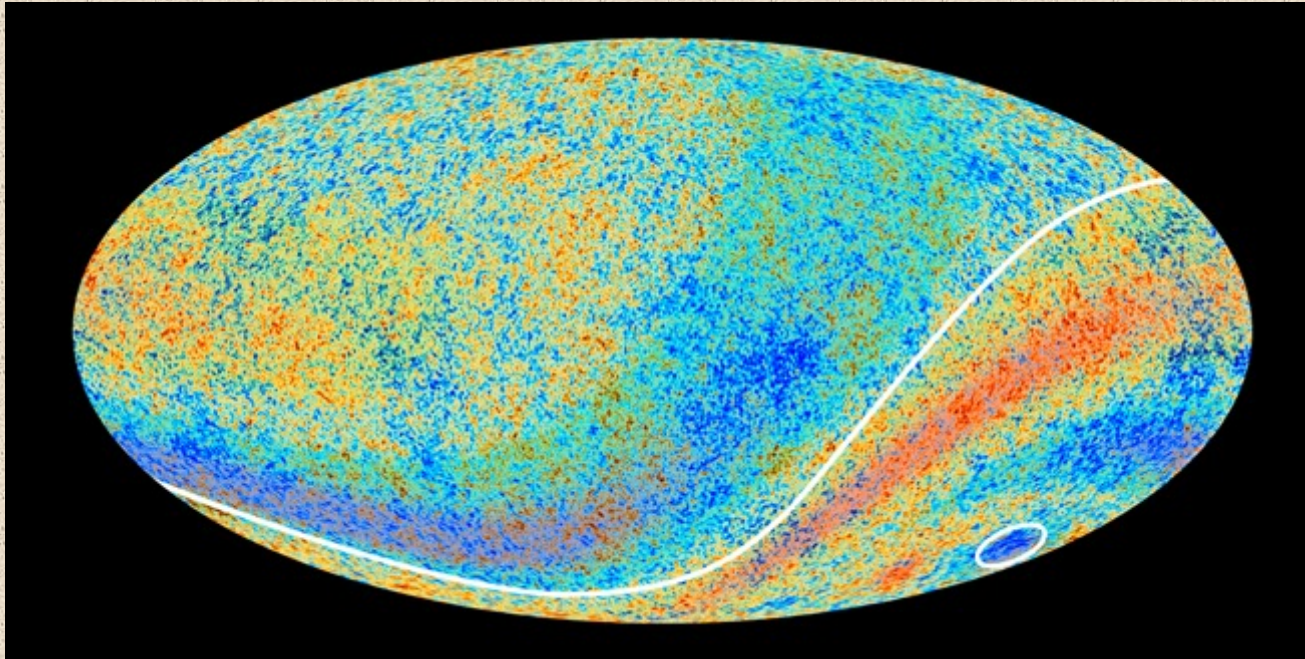
σ_8 tension.

Same tension found from weak lensing analyses of independent datasets i.e. KiDS.

The many Λ CDM crises, tensions and anomalies.

- H_0 tension.
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Λ CDM crises, tensions and anomalies.



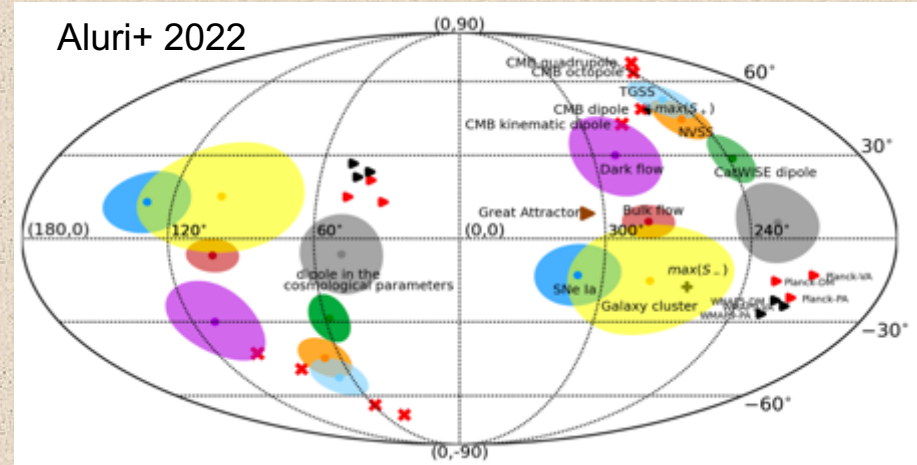
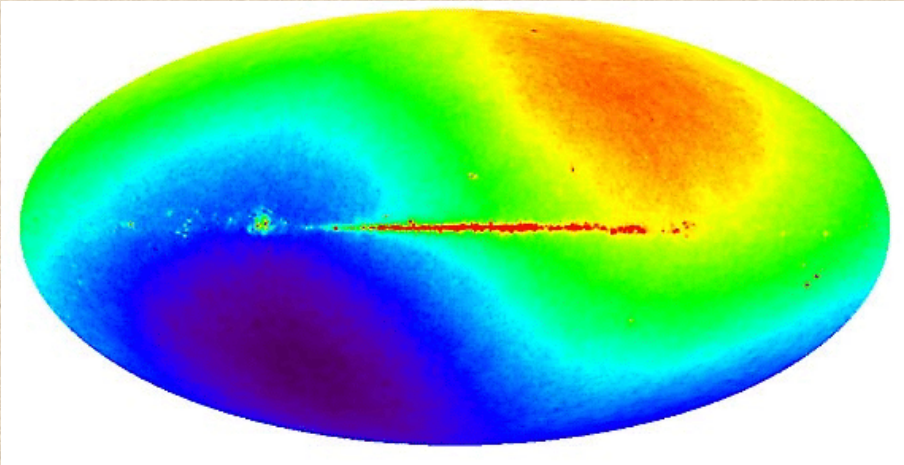
CDM anomalies
Cold spot.
Hemispheric
asymmetry.
Low quadrupole.

ESA Archive

The many Λ CDM crises, tensions and anomalies.

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Λ CDM crises, tensions and anomalies.



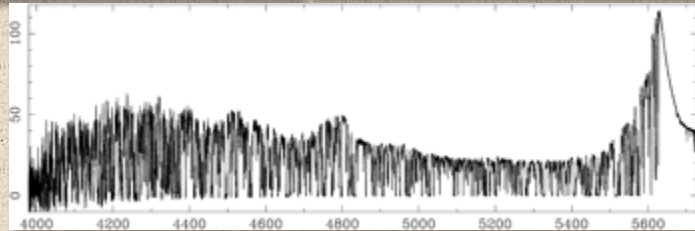
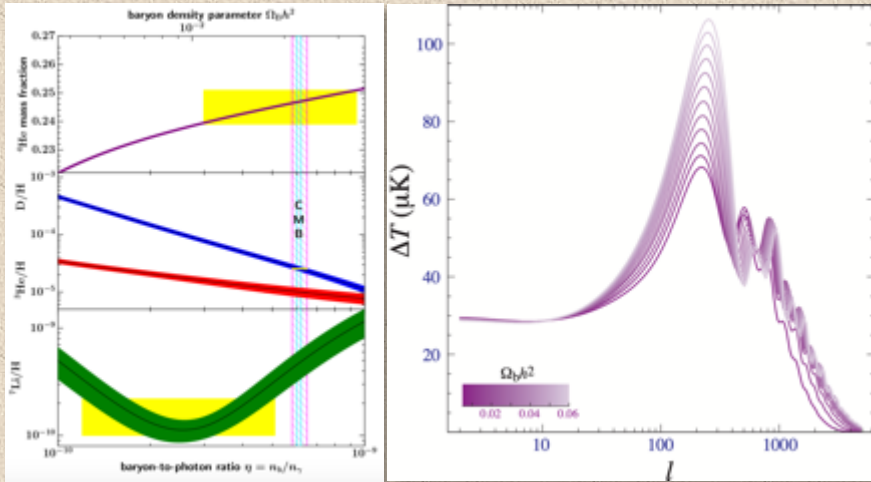
Anomalies: Kinematic Dipole, Galaxy Dipole, Bulk Flow

The many Λ CDM crises, tensions and anomalies.

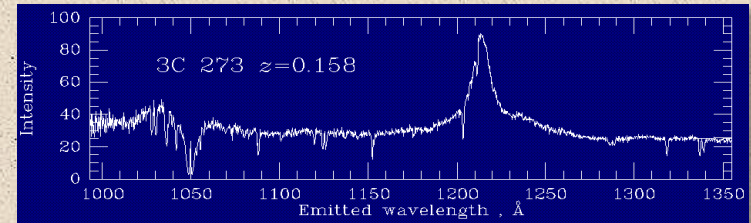
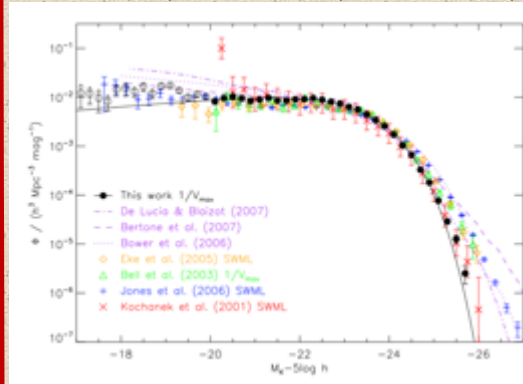
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Λ CDM crises, tensions and anomalies.

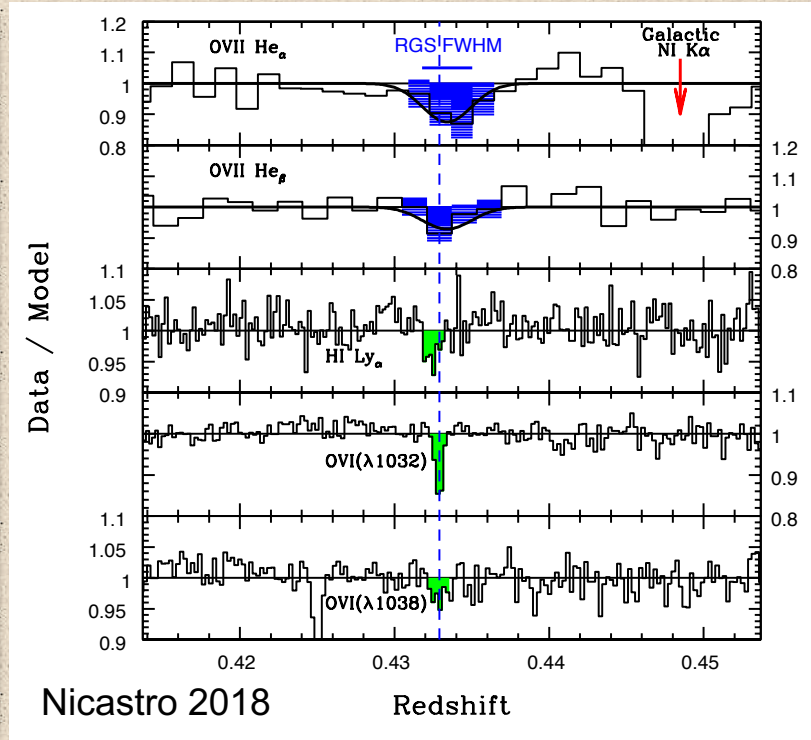
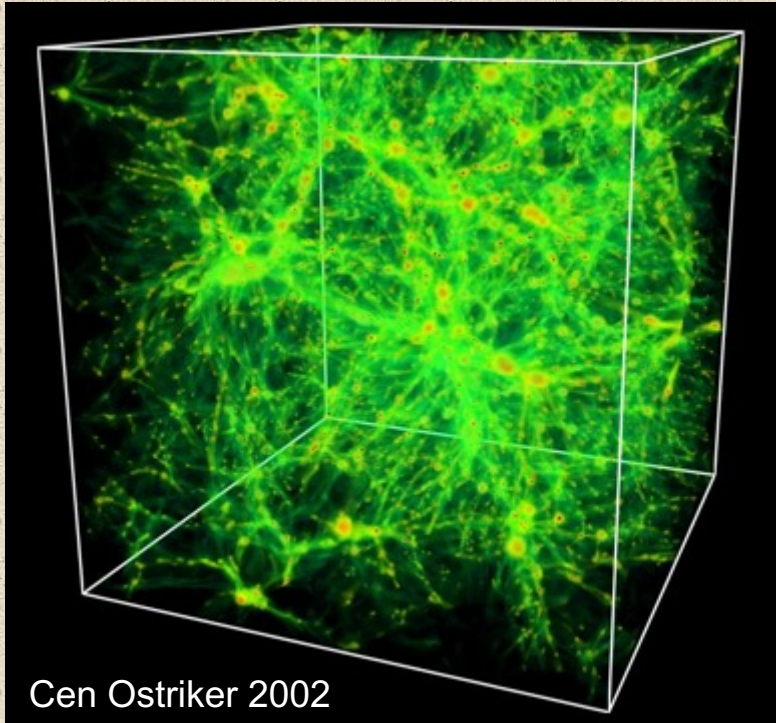
$z > 1$ Consistent baryon density



$z < 1$ Missing baryons



Λ CDM crises, tensions and anomalies.



Missing
baryons



From precision cosmology to accurate cosmology

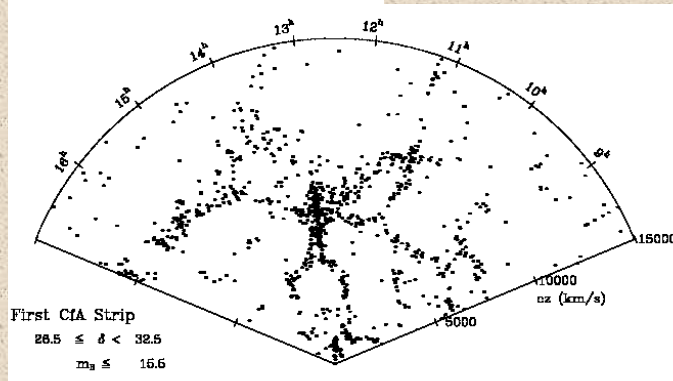
To effectively investigate the reality of these tensions an exquisite control of systematic errors is required.

% accuracy and precision is the goal

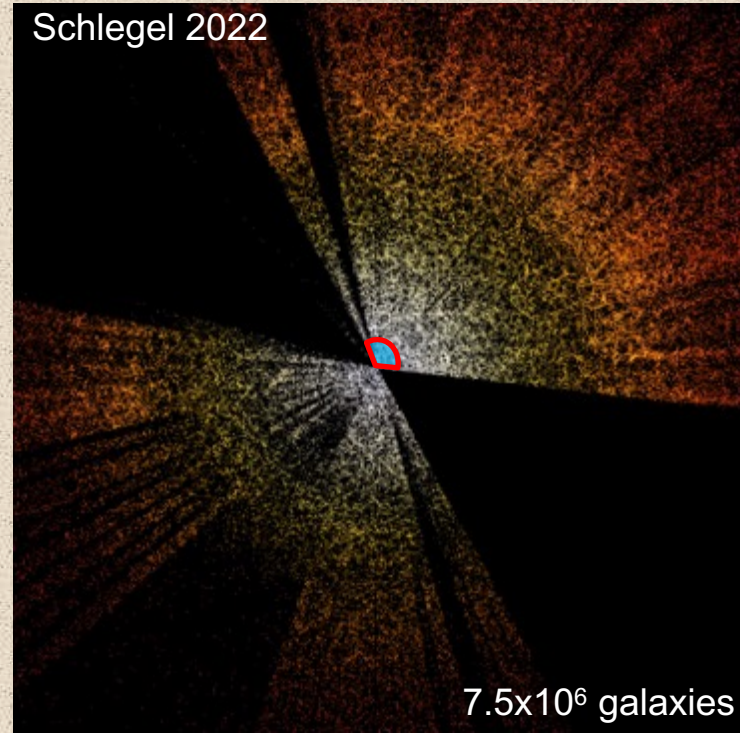
- **Optimized observational strategies.**
- **Combination of different probes (and possibly messengers).**
- Efficient and unbiased data compression methods.
- Theoretical predictions of matching accuracy.
- Effective data vs. model comparisons.

Optimizing observational strategies: DESI

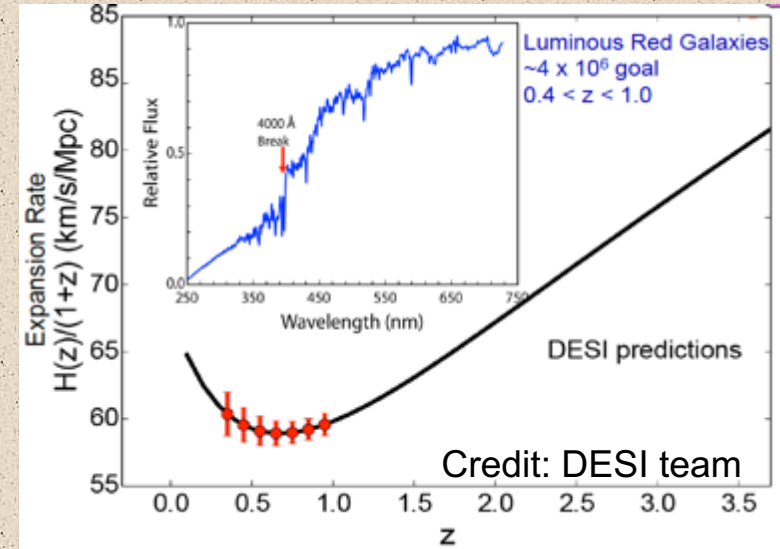
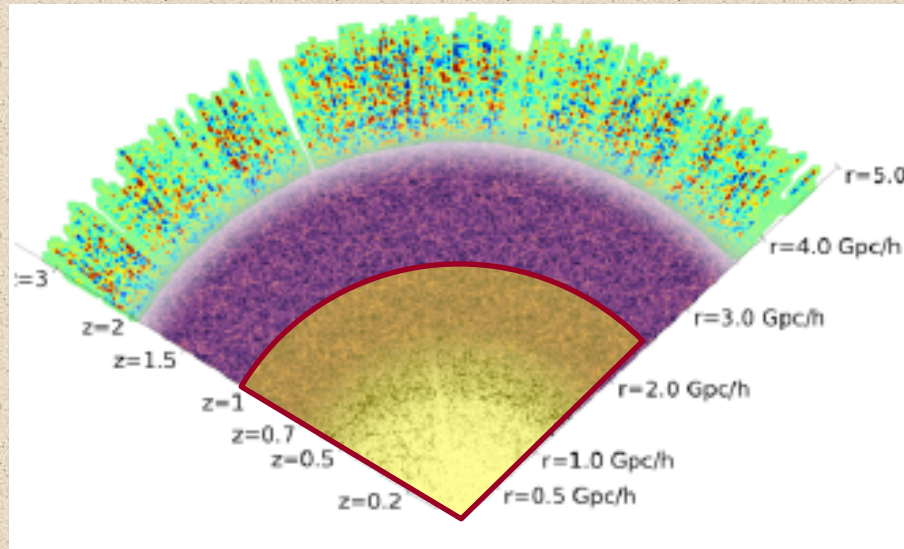
de Lapparent+ 1985



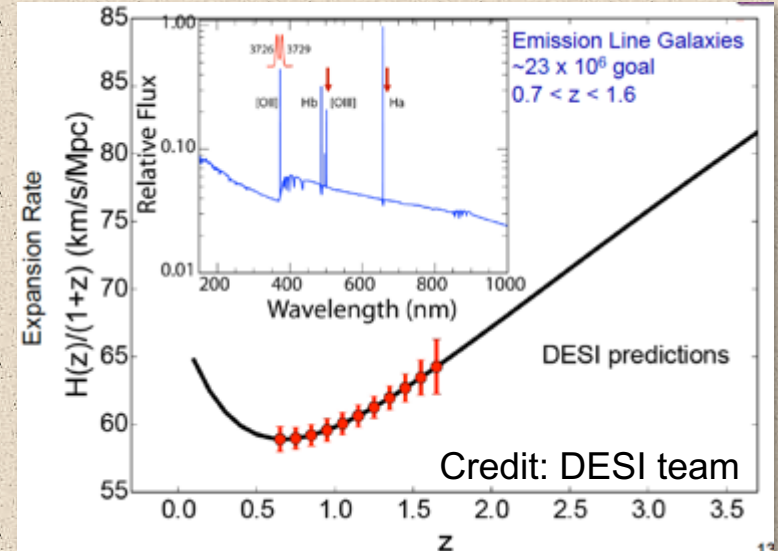
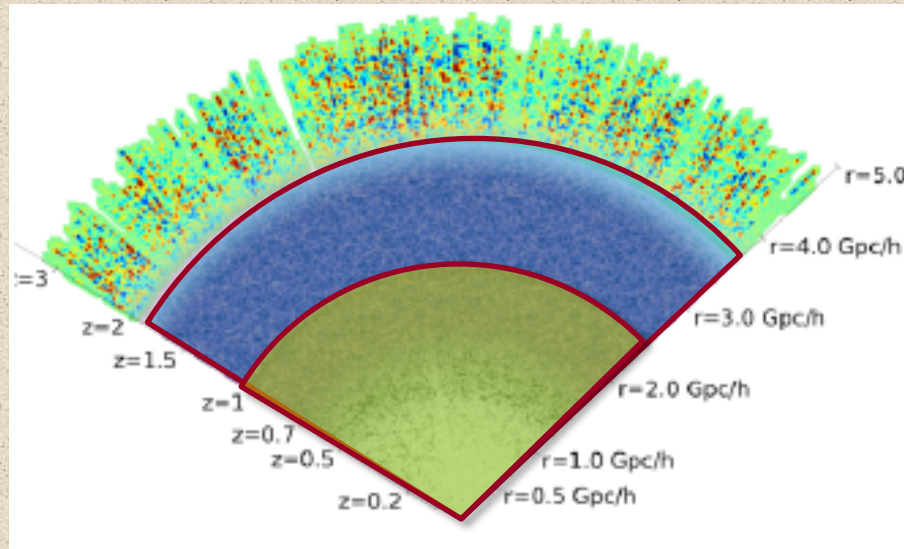
Schlegel 2022



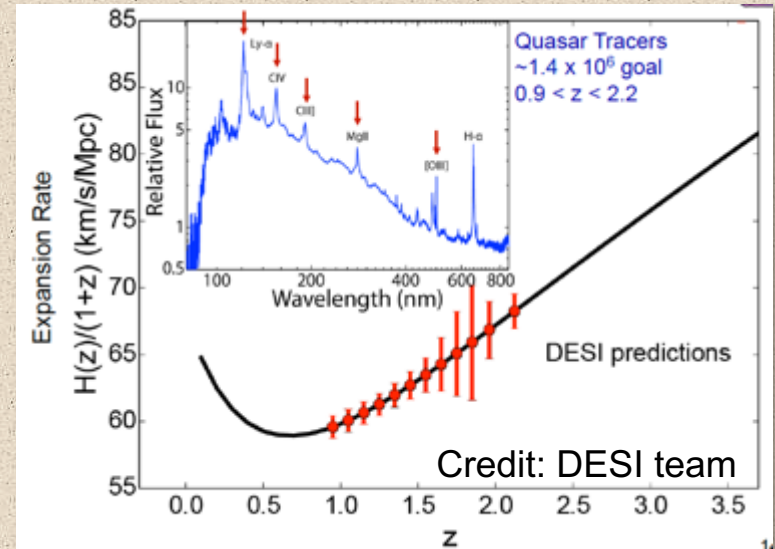
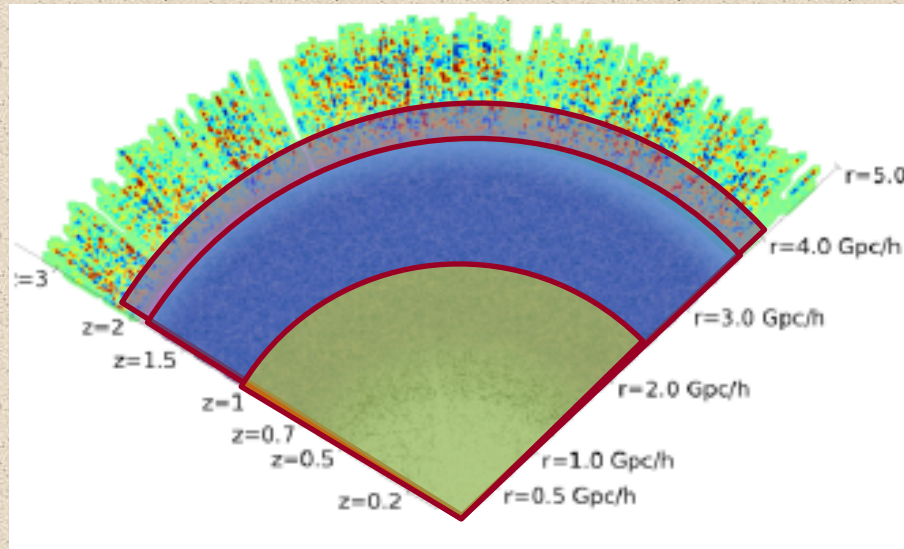
Optimizing observational strategies: DESI



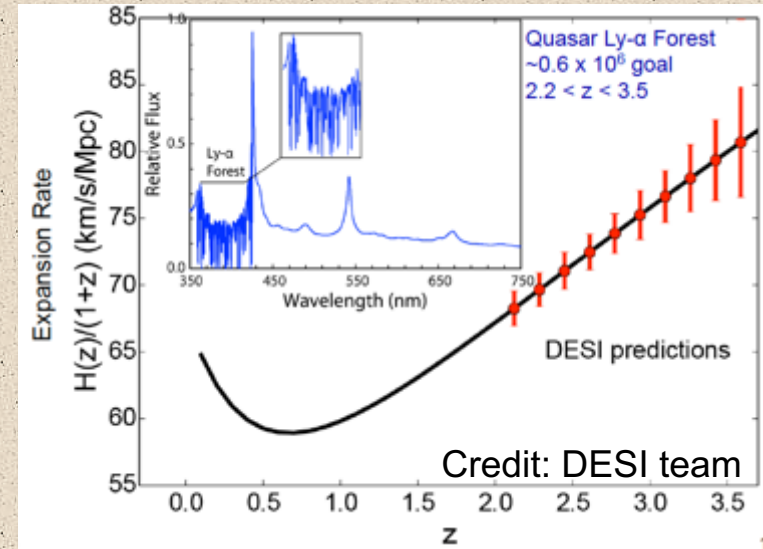
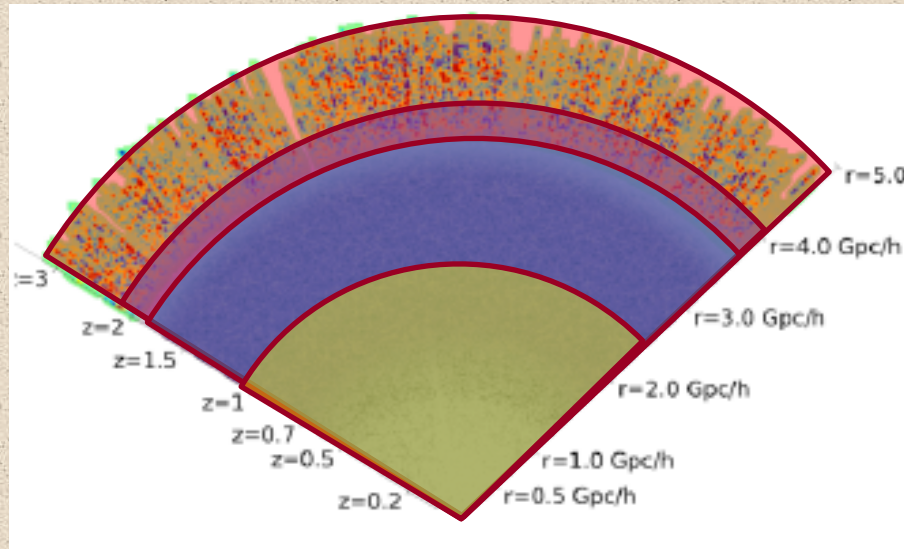
Optimizing observational strategies: DESI



Optimizing observational strategies: DESI



Optimizing observational strategies: DESI



Combining probes: clustering and (weak) lensing



Combining probes: clustering and (weak) lensing (beyond Euclid)

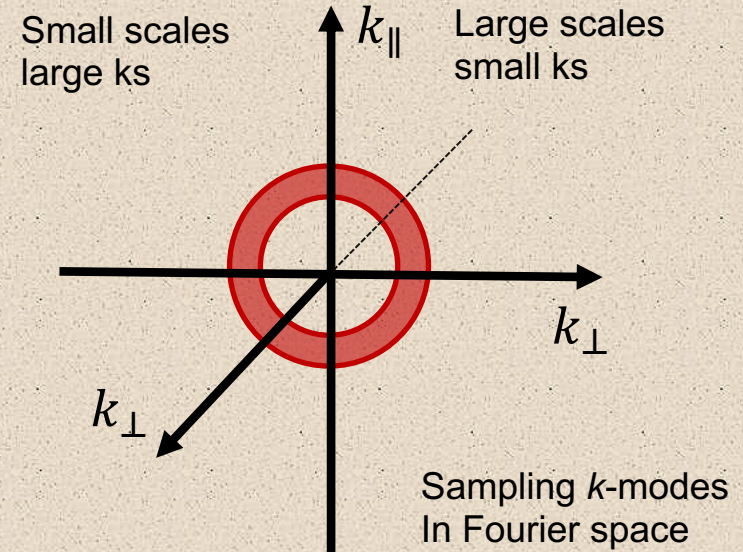
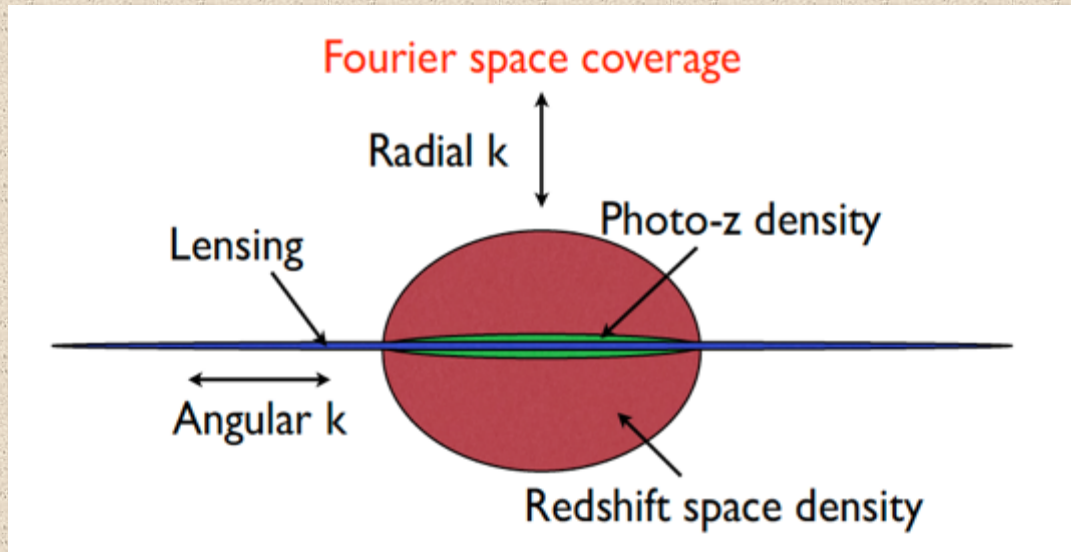


Vera C. Rubin Observatory (2024)
20B galaxies



Nancy Grace Roman Space telescope (2027)
1B photometry 15M spectroscopy

Combining probes: clustering and (weak) lensing



To wrap up.

- Λ CDM is a very effective model.
- A better theory is still to be found.
- Λ CDM is probably a good starting point.
- Investigating tensions is the best way to proceed.
- High accuracy is required.
- Combining probes and messengers is the key.
- Ongoing and future observational campaigns are being designed accordingly.

To wrap up.

Is it the best of times ?

Maybe not.

Is it the worst of times ?

Certainly not.

Is it exciting times ?

Definitely.