Collective excitations in nuclei: The isoscalar and isovector electric giant resonances and spin-isospin chargeexchange modes

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## **Vibrations of a liquid drop in weightlessness**





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 university of groningen In the following: **IS** = **Iso-Scalar IV** = **Iso-Vector** S = SpinG = GiantM = Monopole  $\mathbf{D} = \mathbf{Dipole}$ Q = Quadrupole **O** = **Octupole** 

e.g., ISGMR = Isoscalar giant monopole resonance ISGDR = Isoscalar giant dipole resonance IVGDR = Isovector giant dipole resonance IVSGMR= Isovector spin giant monopole resonance IVSGDR = Isovector spin giant dipole resonance







### **The Collective Response of the Nucleus: Giant Resonances**





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# **Operators and Microscopic Structure**



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✓ university of groningen Microscopic picture: GRs are coherent (1p-1h) excitations induced by single-particle operators.

- Excitation energy depends on

   multipole L (Lħω, since radial operator ∝ r<sup>L</sup>; except for ISGMR and ISGDR, 2ħω & 3ħω, respectively),
   strength of effective interaction and
   collectivity.
- Exhaust appreciable % of EWSR
- Acquire a width due to coupling to continuum and to underlying 2p-2h configurations.





## **Microscopic structure of ISGMR & ISGDR**

**Transition operators:** 



 $3\hbar\omega$  excitation (overtone of c.o.m. motion)



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Nucleus		Many-bo	dy system	with a finit	te size
Vibratior	is	Multipole expansion with <i>r</i> , $Y_{lm}$ , $\tau$ , $\sigma$			
	∆S=0, ∆T=0	$\Delta S=0, \Delta T=1$	∆S=0, ∆T=1	$\Delta S=1, \Delta T=1$	$\Delta S=1, \Delta T=1$
<i>L=0</i> : Monopole	<b>ISGMR</b> $r^2Y_0$	$\mathbf{IAS}\\\boldsymbol{\tau Y}_{\boldsymbol{\theta}}$	<b>IVGMR</b> $\tau r^2 Y_0$	$\frac{\text{GTR}}{\tau\sigma Y_0}$	IVSGMR $\tau \sigma r^2 Y_0$
<i>L=1</i> : Dipole	<b>ISGDR</b> $(r^3 - 5/3 \langle r^2 \rangle r)$	Y <sub>1</sub>	IVGDR $ au r Y_1$		IVSGDR $\tau \sigma r Y_1$
<i>L=2</i> : Quadrupo	ble ISGQR $r^2Y_2$		<b>IVGQR</b> $\tau r^2 Y_2$		IVSGQR $\tau\sigma r^2 Y_2$
<i>L=3</i> : Octupole LEOR, HEOR $r^3Y_3$			<b>Dropped</b> $\Delta S=1$ , $\Delta T=0$ operators because excitations are very weak		



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## **Decay of giant resonances**

- Width of resonance
  - $\Gamma, \Gamma^{\uparrow}, \Gamma^{\downarrow} (\Gamma^{\downarrow\uparrow}, \Gamma^{\downarrow\downarrow})$
  - Γ<sup>↑</sup>: direct or escape width
  - Γ<sup>1</sup>: spreading width
    - $\Gamma^{\downarrow\uparrow}$ : pre-equilibrium,  $\Gamma^{\downarrow\downarrow}$ : compound
- Decay measurements
  - $\Rightarrow$  Direct reflection of damping processes

Allows detailed comparison with theoretical calculations









# **Energy-Weighted Sum Rules**



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#### **Intermezzo: Sum rules**

# **Consider the most general electric multipole operator (Bohr & Mottelson 69), neglecting the current term:**

$$\mathcal{M}(E\lambda,\mu) = \frac{(2\lambda+1)!!}{q^{\lambda}(\lambda+1)} \int \rho(\vec{r}) \frac{\partial}{\partial r} (rj_{\lambda}(qr)) Y_{\lambda\mu}(\hat{r}) d\tau$$
$$j_{\lambda}(qr) = \frac{(qr)^{\lambda}}{(2\lambda+1)!!} \left(1 - \frac{1}{2} \frac{(qr)^2}{(2\lambda+3)} + \cdots\right) \qquad \text{Bessel function}$$

This leads in  $1^{st}$  order (long-wave length limit, i.e.  $qr \ll 1$ ):

$$\mathcal{M}(E\lambda,\mu) = \frac{(2\lambda+1)!!}{q^{\lambda}(\lambda+1)} \int \rho(\vec{r}) \frac{\partial}{\partial r} \left(\frac{r(qr)^{\lambda}}{(2\lambda+1)!!}\right) Y_{\lambda\mu}(\hat{r}) d\tau$$

$$\mathcal{M}(E\lambda,\mu) = \int \rho(\vec{r})r^{\lambda} Y_{\lambda\mu}(\hat{r})d\tau$$

Using 
$$\rho(\vec{r}) = \sum_{k} e\left(\frac{1}{2} - t_{zk}\right)\delta(\vec{r} - \vec{r}_{k})$$







we get:

$$\mathcal{M}(E\lambda,\mu) = \sum_{k} e\left(\frac{1}{2} - t_{zk}\right) r_{k}^{\lambda} Y_{\lambda\mu}(\Omega_{k})$$
$$\mathcal{M}(E\lambda,\mu) = \frac{1}{2} e \sum_{k} r_{k}^{\lambda} Y_{\lambda\mu}(\Omega_{k}) - e \sum_{k} t_{zk} r_{k}^{\lambda} Y_{\lambda\mu}(\Omega_{k})$$

For the isoscalar E0 and E1, 1<sup>st</sup> order leads to a constant and c.o.m. coordinate, respectively. Expanding to 2<sup>nd</sup> order (taking only dependence on r) we get:

$$\mathcal{M}(E0) = \frac{1}{4}e \sum_{k} r_{k}^{2} - \frac{1}{2}e \sum_{k} t_{zk}r_{k}^{2}$$
  
**Isoscalar**  

$$\mathcal{M}(E1,\mu) = \frac{1}{4}e \sum_{k} r_{k}^{3}Y_{1\mu}(\Omega_{k})$$
**Isovector term**  
**neglected**



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Thomas Reiche Kuhn (TRK) sum rule is originally obtained for an atomic system assuming an electric field directed along z-axis:  $S_e(E1) = \sum_f (E_f - E_i) |\langle f | \sum_k z_k | i \rangle|^2$ 

The total absorption cross section is in the long-wave length limit:

$$\int_{0}^{\infty} \sigma(E_{\gamma}) dE_{\gamma} = \frac{4\pi^2 e^2}{\hbar c} \sum_{f} \left( E_f - E_i \right) |\langle f| \sum_k z_k |i\rangle|^2$$

For a Hermitian operator and using closure relation

$$(\sum_{f} |f\rangle < f| = 1)$$
, we obtain:  
$$\int_{0}^{\infty} \sigma(E_{\gamma}) dE_{\gamma} = \frac{4\pi^{2}e^{2}}{\hbar c} \frac{1}{2} \langle i | [\sum_{k} z_{k}, [H, \sum_{k} z_{k}]] | i \rangle$$



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Consider only kinetic term of Hamiltonian:

$$\int_{0}^{\infty} \sigma(E_{\gamma}) dE_{\gamma} = \frac{4\pi^{2}e^{2}}{\hbar c} \frac{1}{2} \left\langle i \right| \left[ \sum_{k} z_{k} , \left[ \frac{p_{z}^{2}}{2m_{e}}, \sum_{k} z_{k} \right] \right] \left| i \right\rangle$$
$$\int_{0}^{\infty} \sigma(E_{\gamma}) dE_{\gamma} = \frac{4\pi^{2}e^{2}}{\hbar c} \frac{\hbar^{2}I}{2m_{e}}$$

I is number of electrons. For a nucleus (see later):

$$e_{eff}^2 I = Z e_{peff}^2 + N e_{neff}^2 = \frac{NZ}{A} e^2$$

Therefore:

$$\int_{0}^{\infty} \sigma(E_{\gamma}) dE_{\gamma} = \frac{2\pi^{2}e^{2}\hbar}{mc} \frac{NZ}{A} = 60 \frac{NZ}{A} \quad MeV \ mb$$

This is the TRK sum rule for a nucleus.



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$$B(E\lambda, J_i \to J_f) = \sum_{\mu M_f} |\langle \Psi_f | \mathcal{M}(E\lambda, \mu) | \Psi_i \rangle|^2$$

$$B(E\lambda, J_i \to J_f) = \sum_{\mu M_f} \langle J_i M_i \lambda \mu | J_f M_f \rangle^2 | \langle \Psi_f \| \mathcal{M}(E\lambda) \| \Psi_i \rangle |^2$$

$$B(E\lambda, J_i \to J_f) = \frac{2J_f + 1}{2J_i + 1} |\langle \Psi_f || \mathcal{M}(E\lambda) || \Psi_i \rangle|^2$$

$$S_{\lambda}(E\lambda) = \sum_{f} (E_{f} - E_{i}) |\langle f | \mathcal{M}(E\lambda, \mu) | i \rangle|^{2}$$
  
$$\Rightarrow S_{\lambda}(E\lambda) = \frac{1}{2} |\langle i | [\mathcal{M}(E\lambda, \mu), [H, \mathcal{M}(E\lambda, \mu)]] | i \rangle$$

Introducing for  $\mathcal{M}(E\lambda,\mu)$  the isoscalar E0, E1 and  $E\lambda$  operators, and using a similar procedure as for TRK sum rule (using Hermitian property and closure relation), we obtain the isoscalar E0, E1 and  $E\lambda$  energy-weighted sum rules (EWSR).



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$$\begin{split} P_{0\mu} &= \frac{1}{2} \sum_{i} r_{i}^{2} \\ \sum_{n} (E_{n} - E_{0}) B(E0, 0 \to n) = S_{0} = \frac{\hbar^{2}}{2m} A < r^{2} > \\ P_{1\mu} &= \frac{1}{2} \sum_{i} r_{i}^{3} Y_{1\mu}(\hat{r}_{i}) \\ \sum_{n} (E_{n} - E_{0}) B(E1, 0 \to n) = S_{1} = \frac{\hbar^{2}}{8\pi m} \frac{3}{4} A[11 < r^{4} > -\frac{25}{3} < r^{2} >^{2} -10\varepsilon < r^{2} >] \\ \varepsilon &= (\frac{4}{E_{2}} + \frac{5}{E_{0}}) \frac{\hbar^{2}}{3mA} \\ Q_{\lambda\mu} &= \sum_{i} r_{i}^{\lambda} Y_{\lambda\mu}(\hat{r}_{i}) \\ \sum_{n} (E_{n} - E_{0}) B(E\lambda, 0 \to n) = S_{\lambda} = \frac{\hbar^{2}}{8\pi m} \lambda (2\lambda + 1)^{2} A < r^{2\lambda - 2} > \end{split}$$



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**Isovector E1**  
operator
$$\mathcal{M}(E1) = \sum_{k=1}^{A} e\left(\frac{1}{2} - t_{zk}\right) \vec{r}_{k}^{int}$$

$$\vec{r}_{k} = \vec{R} + \vec{r}_{k}^{int} \quad where \ \vec{R} = \sum_{k} \vec{r}_{k}/A$$

$$\mathcal{M}(E1) = e\sum_{k=1}^{A} \left(\frac{1}{2} - t_{zk}\right) (\vec{r}_{k} - \vec{R})$$

$$\mathcal{M}(E1) = -e\sum_{k=1}^{A} t_{zk} (\vec{r}_{k} - \vec{R})$$

$$\mathcal{M}(E1) = e\sum_{k=1}^{A} \left(\frac{N-Z}{2A} - t_{zk}\right) \vec{r}_{k}$$

 $\Rightarrow$  Effective charges for neutrons and protons

$$e_{D} = e\left(\frac{N-Z}{2A} - t_{zk}\right) = \begin{cases} \frac{N}{A}e & \text{for proton} \\ -\frac{Z}{A}e & \text{for neutron} \end{cases}$$





$$\sum_{n} (E_n - E_0) B(E\lambda, 0 \to n) = S_{\lambda} = \frac{\hbar^2}{8\pi m} \lambda (2\lambda + 1)^2 A \langle r^{2\lambda - 2} \rangle$$

For isovector *E1*,  $\lambda=1$  and *A* becomes  $Ze_{peff}^2 + Ne_{neff}^2$ , which leads to:

$$\sum_{n} (E_n - E_0) B(E1, 0 \to n) = \frac{\hbar^2}{8\pi m} 9 \left[ Z \left( \frac{N}{A} \right)^2 e^2 \right]$$



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## **Consider isovector electric dipole excitations.**

How does a nucleus respond to an external perturbation, e.g., real photons?

#### $\Rightarrow$ Photo-absorption cross section

 $\gamma$ -rays from bremsstrahlung or positron capture in flight





## Nuclear Collective response Giant Resonances

Isovector Electric Giant Resonances

Monopole (IVGMR)

Isovector

Photo-neutron cross sections



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Dipole (IVGDR)

Quadrupole (IVGQR)



#### **Isovector Giant Dipole Resonances: Photo-neutron cross section**



B. L. Berman and S. C. Fultz, Rev. Mod. Phys. 47 (1975) 713

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Measurement of the giant dipole resonance with mono-energetic photons B.L. Berman and S.C. Fultz Rev. Mod. Phys. 47 (1975) 713

Nucleus	Centroid	Width	
	(MeV)	(MeV)	
<sup>116</sup> Sn	15.68	4.19	
<sup>117</sup> Sn	15.66	5.02	
<sup>118</sup> Sn	15.59	4.77	
<sup>119</sup> Sn	15.53	<b>4.81</b>	
<sup>120</sup> Sn	15.40	<b>4.89</b>	
<sup>124</sup> Sn	15.19	<b>4.81</b>	

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# Photo-neutron cross section in deformed nuclei:

### **Deformed Nucleus**

$$R(\theta,\phi) = R_0(1 + \beta_2 Y_{20}(\theta,\phi))$$

 $\beta_2 (^{150}\text{Nd}) = 0.285(3)$ 



Excitation energies:  $E_2/E_1 = 0.911\eta + 0.089$ Where  $\eta = b/a$  $S_1/S_2 = 1/2$ 

B. L. Berman and S. C. Fultz, Rev. Mod. Phys. 47, 713 (1975)



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#### **Experimental Tool: Electromagnetic excitation at high energies**



Determination of 'photon energy' (excitation energy) via a kinematically complete measurement of the momenta of all outgoing particles (invariant mass)

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#### **Experimental Scheme: The LAND reaction setup @GSI**



## **Dipole Strength Distribution of n-Rich Nuclei**





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## **Dipole strength distributions in neutron-rich Sn isotopes**

Electromagnetic-excitation cross section

**Photo-neutron cross section** 





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## **Dipole strength distributions in <sup>68</sup>Ni**

Simultaneous fit of spectra with 8 individual energy bins as free fit parameters: "deconvolution"





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**Distribution of isovector** dipole strength for the three closed-(sub)shell nickel isotopes <sup>56</sup>Ni, <sup>68</sup>Ni, and <sup>78</sup>Ni calculated in **HF-plus-RPA** using the **FSUGold interaction** parameter set.



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## **Experiments at RCNP, Osaka University**

- ➤ (p,p') reaction at 295 MeV
  - High-resolution spectrometer "Grand Raiden"





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A. Tamii et al., Phys. Rev. Lett. 107 (2011) 062502



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0°-1.8° inelastic proton scattering spectrum shows in addition to IVGDR low-lying E1 and M1 structures.

Peaks with (\*) have been selected for multipoledecomposition analysis.

3.75°-5.25° inelastic proton scattering spectrum is almost structure-less.



C. Iwamoto et al., Phys. Rev. Lett. 108 (2012) 262501

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E1 (dash-dotted line)

M1 (solid line)

E2 (dashed line)



C. Iwamoto et al., Phys. Rev. Lett. 108 (2012) 262501



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# Histograms MDA of 100 keV bins and red circles with errors are of selected peaks.

C. Iwamoto et al., Phys. Rev. Lett. 108 (2012) 262501





#### **Decay of IVGDR built on excited states**





Left: Statistical decay of IVGDR in <sup>156</sup>Dy selected on different angular momentum bins. Curves fits CASCADE calculations with dashed curve increased by 5%. Right: Same as left linearized by multiplying with  $e^{E\gamma/Teff}$ 

A. Stolk et al., Phys. Rev. C40 (1989) R2454



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Role of isospin in the statistical decay of the IVGDR built on excited states

Clebsch-Gordon coefficient for isospin coupling <0010|00>=0

Dotted: pure ISGQR Dashed: pure isospin Dash-dotted: complete isospin mixing Solid: isospin mixing (~ 5%)

M.N. Harakeh et al., Phys. Lett. B176 (1986) 297



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# **Compression Modes ISGMR & ISGDR**



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In fluid mechanics, **compressibility** is a measure of the relative volume change of a fluid as a response to a pressure change.

 $\beta = -\frac{1}{V} \frac{\partial V}{\partial P}$ 

where *P* is pressure, *V* is volume.

**Incompressibility or bulk modulus (***K***) is a measure of a substance's resistance to uniform compression and can be formally defined:** 

 $K = -V \frac{\partial P}{\partial V}$ 





# For the equation of state of symmetric nuclear matter at saturation nuclear density:

$$\left[\frac{d(E/A)}{d\rho}\right]_{\rho=\rho_0} = 0$$

and one can derive the incompressibility<sup>0</sup> of nuclear matter: -20

0

$$K_{nm} = \left[9\rho^2 \frac{d^2(E/A)}{d\rho^2}\right]_{\rho = \rho}$$

**E/A:** binding energy per nucleon

**ρ** : nuclear density

J.P. Blaizot, Phys. Rep. 64 (1980) 171

 $\rho_0$  : nuclear density at saturation









120 HARD 100 SOFT 80 E/A (MeV/nucleon) 60 40 20 -20 0.1 0.3 0.4 0.5 0.2 0.6  $(1/fm^{3})$ Q

Equation of state (EOS) of nuclear matter

More complex than for infinite neutral liquids Neutrons and protons with different interactions Coulomb interaction of protons

- 1. Governs the collapse and explosion of giant stars (supernovae)
- 2. Governs formation of neutron stars (mass, radius, crust)
- **3.** Governs collisions of heavy ions.
- 4. Important ingredient in the study of nuclear properties.







## **Isoscalar Excitation Modes of Nuclei**

**Hydrodynamic models/Giant Resonances** Coherent vibrations of nucleonic fluids in a nucleus.

**Compression modes: ISGMR, ISGDR** 

**In Constrained and Scaling Models:** 

$$E_{ISGMR} = \hbar \sqrt{\frac{K_A}{m \langle r^2 \rangle}}$$

$$E_{ISGDR} = \hbar \sqrt{\frac{7}{3} \frac{K_A + \frac{27}{25} \varepsilon_F}{m \langle r^2 \rangle}}$$

 $\varepsilon_F$  is the Fermi energy and the nucleus incompressibility:

$$K_A = [r^2 (d^2 (E/A)/dr^2)]_{r=R_0}$$

J.P. Blaizot, Phys. Rep. 64 (1980) 171



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# **Giant resonances**

- Macroscopic properties:  $E_x$ ,  $\Gamma$ , %EWSR
- Isoscalar giant resonances; compression modes

ISGMR, ISGDR ⇒ Incompressibility, symmetry energy

$$K_{A} = K_{vol} + K_{surf}A^{-1/3} + K_{sym}((N-Z)/A)^{2} + K_{Coul}Z^{2}A^{-4/3}$$



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#### Nucleus, e.g., <sup>208</sup>Pb

# **Inelastic** *α* **scattering**



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# ISGQR, ISGMR



$$\Leftarrow$$
<sup>208</sup>Pb(α,α') at E<sub>α</sub>=120 MeV

# Large instrumental background and nuclear continuum!

M. N. Harakeh et al., Phys. Rev. Lett. 38 (1977) 676

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**BBS@KVI** 

(p,p') at  $E_p \sim 300$  $(\alpha,\alpha')$  at  $E_{\alpha} \sim 400$ & 200 MeV at RCNP & KVI, respectively

**RCNP** 



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Si-ball 16 Si-detectors at 10 cm from the target total solid angle: 1 sr



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total solid angle: 0.37 sr

## **KVI Big-Bite Spectrometer (BBS)**



#### **ISGQR at 10.9 MeV**

#### **ISGMR at 13.9 MeV**







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## Multipole decomposition analysis (MDA)

$$\left(\frac{d^2\sigma}{d\Omega dE}(\vartheta_{c.m.}, E)\right)^{\exp.} = \sum_{L} a_L(E) \left(\frac{d^2\sigma}{d\Omega dE}(\vartheta_{c.m.}, E)\right)_{L}^{calc.}$$

$$\left(\frac{d^2\sigma}{d\Omega dE}(\theta_{c.m.},E)\right)^{\exp}$$

: Experimental cross section



 $\left(\frac{d^2\sigma}{d\Omega dE}(\vartheta_{c.m.}, E)\right)_{I}^{calc.}$ : DWBA cross section (unit cross section)

 $a_{I}(E)$ : EWSR fraction

- **ISGR** (L<15)+ **IVGDR** (through Coulomb excitation) **a**.
- **b.** DWBA formalism; single folding  $\Rightarrow$  transition potential

$$\delta U_{L}(r,E) = \int d\vec{r}' \,\delta \rho_{L}(\vec{r}',E) [V(|\vec{r}-\vec{r}'|,\rho_{0}(r')) + \rho_{0}(r')\frac{\partial V(|\vec{r}-\vec{r}'|,\rho(r'))}{\partial \rho_{0}(r')}]$$

$$U(r) = \int \vec{dr'} V(|\vec{r} - \vec{r'}|, \rho_0(r'))\rho_0(r')$$





# **Transition density**

**ISGMR Satchler, Nucl. Phys. A472 (1987) 215** 

$$\delta \rho_0(r, E) = -\alpha_0 [3 + r\frac{d}{dr}]\rho_0(r)$$
$$\alpha_0^2 = \frac{2\pi\hbar^2}{mA < r^2 > E}$$

ISGDR Harakeh & Dieperink, Phys. Rev. C23 (1981) 2329

$$\begin{split} &\delta\!\rho_1(r,E) = -\frac{\beta_1}{R\sqrt{3}} [3r^2 \frac{d}{dr} + 10r - \frac{5}{3} < r^2 > \frac{d}{dr} + \varepsilon(r\frac{d^2}{dr^2} + 4\frac{d}{dr})]\rho_0(r) \\ &\beta_1^2 = \frac{6\pi\hbar^2}{mAE} \frac{R^2}{(11 < r^4 > -(25/3) < r^2 >^2 - 10\varepsilon < r^2 >)} \end{split}$$

Other modes Bohr-Mottelson (BM) model

$$\delta \rho_L(r, E) = -\delta_L \frac{d}{dr} \rho_0(r)$$
  
$$\delta_L^2 = (\beta_L c)^2 = \frac{L(2L+1)^2}{(L+2)^2} \frac{2\pi\hbar^2}{mAE} \frac{\langle r^{2L-2} \rangle}{\langle r^{L-1} \rangle^2}$$



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Uchida et al., Phys. Lett. B557 (2003) 12 Phys. Rev. C69 (2004) 051301

 $(\alpha, \alpha')$  spectra at 386 MeV



10 <sup>3</sup> (a)

10

10

10

10 (b)  $E_{\rm v} = 14.5 \, {\rm MeV}$ 

10

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E<sub>x</sub>= 24.5 MeV

θ<sub>c.m.</sub> (deg.)

<sup>116</sup>Sn

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d<sup>2</sup>d/dΩdE (mb/sr MeV)





#### In HF+RPA calculations,

$$K_{nm} = \left[9\rho^2 \frac{d^2(E/A)}{d\rho^2}\right]_{\rho = \rho_0}$$

**Nuclear matter** 

 $K_A$ : incompressibility

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**E**/A: binding energy per nucleon

- **ρ** : nuclear density
- $\rho_0$  : nuclear density at saturation





## From GMR data on <sup>208</sup>Pb and <sup>90</sup>Zr,

## $K_{\infty} = 240 \pm 10 \text{ MeV}$

[See, e.g., G. Colò et al., Phys. Rev. C 70 (2004) 024307]

## This number is consistent with both ISGMR and ISGDR Data and with non-relativistic and relativistic calculations



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Isoscalar GMR strength distribution in Sn-isotopes obtained by Multipole Decomposition Analysis of singles spectra obtained in <sup>A</sup>Sn(α,α') measurements at incident energy 400 MeV and angles from 0° to 9°





$$K_A \sim K_{vol} (1 + cA^{-1/3}) + K_{\tau} ((N - Z)/A)^2 + K_{Coul} Z^2 A^{-4/3}$$
$$K_A - K_{Coul} Z^2 A^{-4/3} \sim K_{vol} (1 + cA^{-1/3}) + K_{\tau} ((N - Z)/A)^2$$

~ Constant +  $K_{\tau}((N - Z)/A)^2$ 

We use  $K_{Coul} = -5.2$  MeV (from Sagawa) (N - Z)/A $^{112}Sn - ^{124}Sn: 0.107 - 0.194$ 



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## Sn isotopes $\Rightarrow K_{\tau} = -550 \pm 100 \text{ MeV}$



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# **Monopole strength Distribution**






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#### Data from H. Sagawa et al., Phys. Rev. C 76 (2007) 034327



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Colò *et al.*: Non-relativistic RPA (without pairing) reproduces ISGMR in <sup>208</sup>Pb and <sup>90</sup>Zr.

Piekarewicz: Relativistic RPA (FSUGold model) reproduces g.s. observables and ISGMR in <sup>208</sup>Pb, <sup>144</sup>Sm and <sup>90</sup>Zr [ $K_{\infty} = 230$  MeV]

Vretenar: Relativistic mean field (DD-ME2: densitydependent mean-field effective interaction).

[ $K_{\infty} = 240$  MeV]. Possibly agreement is fortuitous since strength distributions are not much different from those by Colò *et al.* and Piekarewicz.

Tselyaev *et al.*: Quasi-particle time-blocking approximation (QTBA) (T5 Skyrme interaction)  $[K_{\infty} = 202 \text{ MeV}?!]$ 

### Softness of Sn-nuclei is still unresolved



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#### Splitting of the ISGMR under deformation

#### *K* projection of *J* on symmetry axis is good quantum number in deformed nuclei Coupling of ISGMR with *K*=0 component of ISGQR



#### **Isoscalar Giant Resonances in Nd isotopes: QRPA calculations**



K. Yoshida and T. Nakatsukasa, Phys. Rev. C 88 (2013) 034309



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#### **Effect of deformation on Isoscalar Giant Resonances: Sm isotopes**





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## **Decay of Giant Resonances**



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## **Decay of giant resonances**

- Width of resonance
  - $\Gamma, \Gamma^{\uparrow}, \ \Gamma^{\downarrow} \ (\Gamma^{\downarrow\uparrow}, \Gamma^{\downarrow\downarrow})$
  - Γ<sup>†</sup>: direct or escape width
  - Γ<sup>1</sup>: spreading width
    - $\Gamma^{\downarrow\uparrow}$ : pre-equilibrium,  $\Gamma^{\downarrow\downarrow}$ : compound
- Decay measurements
  - $\Rightarrow$  Direct reflection of damping processes

Allows detailed comparison with theoretical calculations







# Excitation of ISGDR in <sup>208</sup>Pb

- In <sup>208</sup>Pb located around 22 MeV and width of 4 MeV
- L=1 angular distribution peaks close to a scattering angle of 3°
- Difficult to identify in nuclear continuum and rides on instrumental background

Singles <sup>208</sup>Pb( $\alpha,\alpha'$ ) At  $E_{\alpha}$  = 200 MeV



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### **Microscopic structure of ISGDR**

### **Transition operator**



Spurious centerOvertoneof mass motion

 $3\hbar\omega$  excitation (overtone of c.o.m. motion)



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Si-ball 16 Si-detectors at 10 cm from the target total solid angle: 1 sr



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total solid angle: 0.37 sr

### **KVI Big-Bite Spectrometer (BBS)**



# **Proton-decay detection**



# α-*p* separation using rise time of signal Si(Li)



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## **Neutron-decay detection**





<sup>208</sup>Pb( $\alpha,\alpha' p \text{ or } n$ )





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# <sup>208</sup>Pb( $\alpha,\alpha'$ ) followed by *p* decay

30  $^{208}$ Pb( $\alpha, \alpha p$ )<sup>2</sup> 25 E\_=200 MeV **Decay to hole** .5°<Θ, <6.0° states in <sup>207</sup>Tl; 20 branching ratios ө W) 15 Ш <sup>207</sup>TI (E,=0, 0.35 MeV predicted by <sup>207</sup>TI (E. =1, **3**5, 1,68 M Gorelik et al. 10 =0.MeV5 M. Hunyadi et al., Phys. Lett. B576 (2003) 253 ISCOR M. Hunyadi et al., Phys. Rev. C75 (2007) 014606 0 30 10 15 20 25 35 5 E, in <sup>208</sup>Pb (MeV)



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# **Branching ratios for decay**





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# ISGDR in <sup>208</sup>Pb in *p* decay





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# **Branching ratios for decay**



This work M.L. Gorelik *et al.*, PRC 62 (2000) 047301; Continuum RPA; Landau-Migdal Parameters:  $f^{ex}, f'$ ; Smearing parameter  $\Delta$  energy-dependent



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# **Overtone of the ISGQR?** $[r^4Y_2]$



# <sup>208</sup>Pb( $\alpha,\alpha'$ ) followed by *n* decay





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### Branching ratios for the direct neutron-decay channel

From the simplified MDA of angular distributions



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Final-state spectra in <sup>207</sup>Pb obtained from neutron decay of

(a) continuum underlying ISGMR in <sup>208</sup>Pb and
(b and c) ISGMR proper.
(b) Fit with 100% statistical
(c) Fit with 60% statistical

$I_{j}$	$E_x$ (MeV)	Γ¦ (keV), expt.	$\Gamma_i^{\dagger}$ (keV), theory
<i>p</i> 1/2	0	$140 \pm 35$	5
$l_{13/2}$	1.630	140 ± 55	6
$f_{5/2}$	0.570	$70 \pm 15$	92
$p_{3/2}$	0.890	$50 \pm 10$	8
$f_{1/2}$	2.340	$165 \pm 40$	174



S. Brandenburg et al., Nucl. Phys. A466 (1987) 29



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### **Data analysis: Proton decay**







# **Experimental results**

M. Hunyadi et al., Phys. Rev. C 80 (2009) 044317



# **Experimental results**



## Strength distribution of ISGDR in <sup>58</sup>Ni



Spectra of L = 1 strengths obtained with DOS method in percentage of isoscalar EWSR; a) coincidence data gated on g.s. decay and b) singles data.



Differential cross section of resonance structure fitted with L = 1and L = 3 DWBA calculations.





#### Proton-decay branching ratios Normalized to 100%

		Exp. (%)	Cal. (%)
		(24-38 MeV)	(15-40 MeV)
	7/2-	61.3 (with 5/2 <sup>-</sup> )	47
	3/2-	7.9	3.1
	3/2 <sup>-</sup> , 1/2 <sup>-</sup>	9.9	2.2 (only for 1/2-)
	5/2-	3.2±3.4	-
	$1/2^{+}$	2.0±4.2	13.4
	3/2+, 5/2+	15.9	34.3
Σ		100 %	100%

#### Calculations: M.L. Goerlik, I.V. Safonov, and M.H. Urin, Phys. Rev. C69 (2004) 054322





## Conclusions!

- There has been much progress in understanding ISGMR & ISGDR macroscopic properties
  - Systematics:  $E_x$ ,  $\Gamma$ , %EWSR
  - $\Rightarrow K_{\rm nm} \approx 240 {
    m MeV}$
  - $\Rightarrow K_{\tau} \approx -500 \text{ MeV}$
- Sn nuclei are softer than <sup>208</sup>Pb and <sup>90</sup>Zr.
- Recently, Microscopic Structure for a few nuclei CRPA has some success in <sup>208</sup>Pb & <sup>58</sup>Ni but fails badly in <sup>116</sup>Sn & <sup>90</sup>Zr.
- Possible observation quadrupole compression mode, i.e. overtone of ISGQR





# Gamma-Decay Neutron-Skin Thickness Pygmy Dipole Resonance







## **Setup at KVI**



D. Savran et al., Nucl. Inst. and Meth. Phys. Res. A 564 (2006) 267



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### ISGQR at 10.8 MeV ISGMR at 13.8 MeV

Hatched area ⇒ IVGDR contribution (Coulomb + nuclear)





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Isoscalar transition density in the Goldhaber-Teller model for excitation of IVGDR in inelastic  $\alpha$ -scattering.

$$g_1^{10}(r) = g_1^{n}(r) - g_1^{p}(r) = \alpha_1 \gamma \left(\frac{N-Z}{A}\right) \left(\frac{d\rho(r)}{dr} + \frac{1}{3}c\frac{d^2\rho(r)}{dr^2}\right).$$

$$\frac{\Delta R_{\rm PN}}{R_0} = \frac{R_{\rm n} - R_{\rm p}}{\frac{1}{2}(R_{\rm n} + R_{\rm p})} = \gamma \frac{2(N - Z)}{3A}.$$

Here,  $\gamma$  is related to the proton and neutron central density distributions and thus to  $\Delta R_{pn}$ .  $\alpha_1$  is deformation length obtained from TRK sum rule. Therefore, DWBA cross sections can be calculated as function of  $\Delta R_{pn}/R_0$  for the Goldhaber-Teller model and similarly for the Steinwedel-Jensen model.

# γ-decay branching ratios are known from photo-absorption experiments.



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Full line is the calculated  $\alpha \gamma_0$ coincidence cross section, averaged over the solid angle of the  $\alpha$ -particle and integrated over the full  $\gamma$ -ray solid angle  $(4\pi)$  and over the  $\Delta E$  energy range as function of  $\Delta R_{pn}/R$ . The experimental  $\alpha \gamma_0$  cross sections for the IVGDR are shown as full circles with vertical error bars. The deduced values for  $\Delta R_{pn}/R$  with the associated uncertainty (full circles with horizontal error bars) are also indicated.

Isotope	Present work $\Delta R_{\rm PN}/R_0$ (%)	Present work $\Delta R_{\rm PN}$ (fm)	Batty et al. [2] $\Delta R_{\rm PN}$ (fm)	Angeli et al. [6] ⊿R <sub>PN</sub> (fm)
<sup>116</sup> Sn	0.5 ± 2.7	$0.02 \pm 0.12$	0.15 ± 0.05	0.13
<sup>124</sup> Sn	4.4 ± 2.4	$0.21 \pm 0.11$	$0.25 \pm 0.05$	0.22
<sup>208</sup> Pb	$3.5^{+1.5}_{-1.6}$	0.19 ± 0.09	0.14 ± 0.04	0.22



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#### <sup>140</sup>Ce( $\alpha, \alpha' \gamma$ ) - coincidence matrix



<sup>140</sup>Ce( $\alpha,\alpha'\gamma$ ) vs. <sup>140</sup>Ce( $\gamma,\gamma'$ )



D. Savran et al., Phys. Rev. Lett. 97 (2006) 172502



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#### Multipole assignment with $\alpha$ - $\gamma$ angular correlation



#### Multipole assignment with $\alpha$ - $\gamma$ angular correlation



Comparison of  $(\alpha, \alpha' \gamma)$  with  $(\gamma, \gamma')$  on <sup>138</sup>Ba



E1 strength distribution in <sup>140</sup>Ce, <sup>138</sup>Ba, <sup>124</sup>Sn, and <sup>94</sup>Mo





Nantes; 19-20 January 2015









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#### Identification of PDR structure in $(\alpha, \alpha' \gamma)$

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E. Lanza et al., PRC 89 (2014) 041601(R)

- Good reproduction of experimental results using RQTBA transition densities + semi-classical reaction model
- ⇒ Different response to complementary probes allows identification of PDR structure







- The grey histogram corresponds to the total unresolved strength.
- Top panel: Discrete level in α scattering
- Centre panel: Discrete levels in <sup>17</sup>O scattering
- **Bottom panel: photon scattering**
- L. Pellegri et al., Phys. Lett. B738 (2014) 519



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# **Future Prospects**



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# Outlook

Radioactive ion beams will be available at energies where it will be possible to study excitation of ISGMR and ISGDR RIKEN, FAIR, SPIRAL2, NSCL, EURISOL

Determine ISGMR and ISGDR in unstable Sn nuclei. A = 106 to 134 possible

 $\Rightarrow$  A more precise determination of  $K_{\tau}$ 



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Nuclear structure studies with reactions in inverse kinematics - Possible at GSI/FAIR, RIKEN, GANIL (beam energies of 50-100 MeV/u are needed!)





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#### Nuclear structure studies with reactions in inverse kinematics Challenges with exotic beams

• Inverse kinematics





- Intensity of exotic beams is very low ( $\sim 10^4 10^5$  pps)
- To get reasonable yields thick target is needed
- Very low energy (~ sub MeV) recoil particle will not come out of the thick target





## **Active target**

A gas detector where the target gas also acts as a detector

- Good angular coverage
- Effective target thickness can be increased without much loss of resolution
- > Detection of very low energy recoil particle is possible

#### MAYA active-target detector at GANIL



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### Schematic view of MAYA active target detector





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#### **GSI Storage Ring**

**Experimental Storage Ring** 

Luminosity: 10<sup>26</sup> – 10<sup>27</sup> cm<sup>-2</sup>s<sup>-1</sup>

EPJ Web Conf. 66, 03093 (2014)



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# Advantages and disadvantages of storage-ring experiments

Advantages: Large intensities in the ring Little energy loss in the target No target window (no background) High resolution of the beam (cooling) Forward focusing for high-energy particles Low-energy threshold

Disadvantages: Ultra high vacuum Very small recoil energies for low q Thin targets







## **Detection system @ FAIR**



Figure 1: Schematic view of the EXL detection systems. Left: Set-up built into the NESR storage ring. Right: Target-recoil detector surrounding the gas-jet target.

#### Use of EXL recoil detector has been under evaluation



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# Thank you for your attention



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