



The neutrino-blazar and the new era of multi-messenger Astronomy

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And on behalf of the MAGIC collaboration

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Overview

- Blazars as key objects of the multi-wavelenght Astrophysics
- The **neutrino-blazar** and the birth of multi-messenger Astrophysics
- Extreme blazars as new candidates of extreme Physics?

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3

Multi-wavelenght Astrophysics of blazars

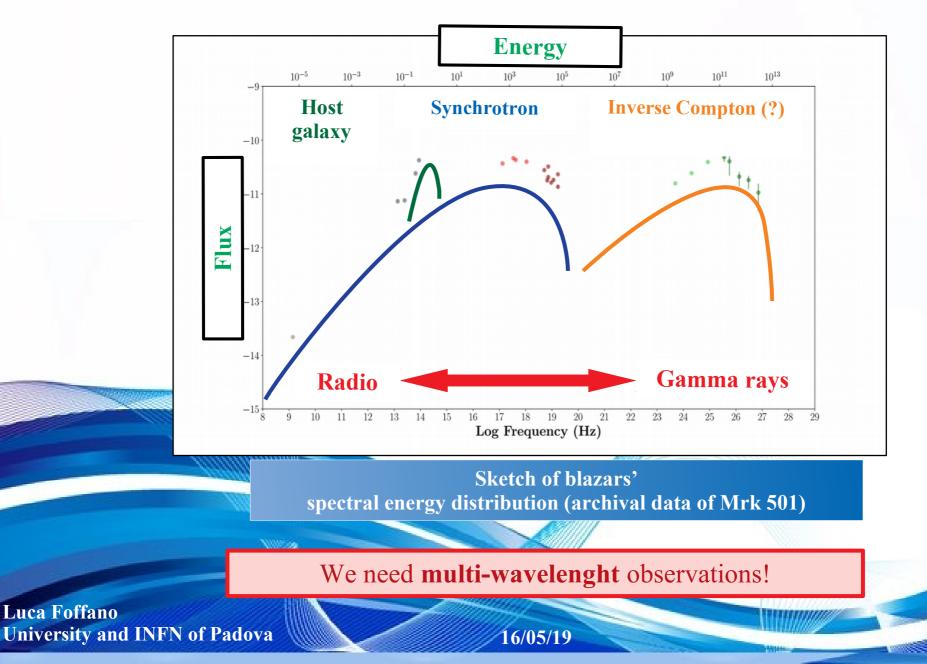
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A simplified view







Multi-wavelenght and multi-messenger Astrophysics of blazars

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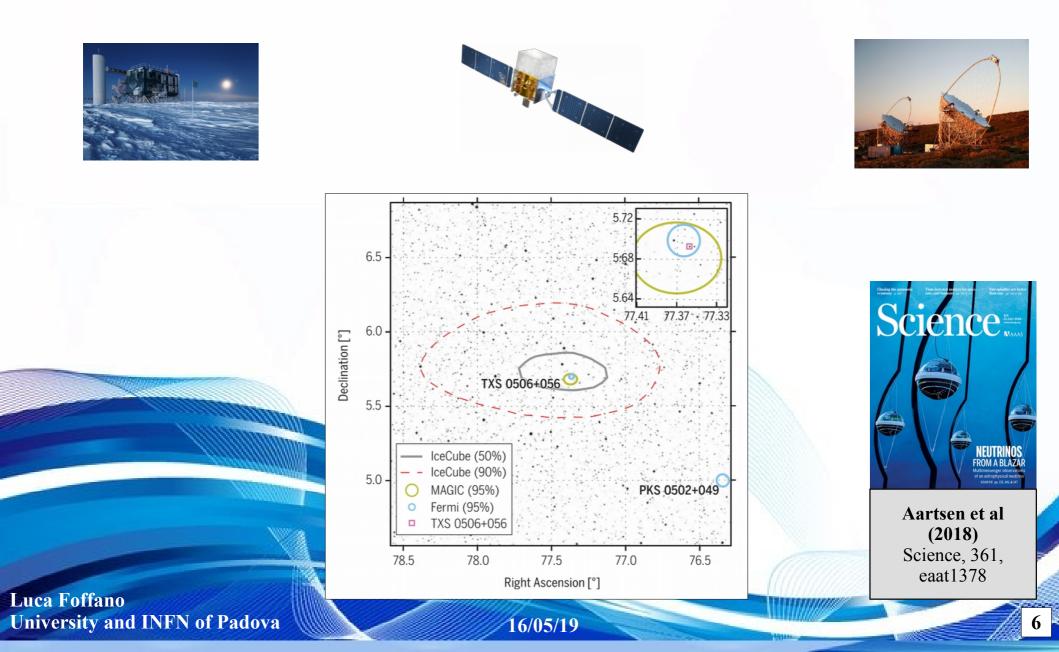
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5



The neutrino-blazar detection







Neutrino follow-up observations with MAGIC





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- IACTs agree on efforts in a transient follow-up program since 2012
- ▶ During the event EHE-170922A, all of them pointed to the blazar TXS 0506+056:
 - → **H.E.S.S.:** fastest follow-up after only $4h \rightarrow no$ detection
 - ▶ **VERITAS:** first obs. after $\sim 12h \rightarrow$ no detection
 - > MAGIC:
 - No observations during the event due to bad weather conditions
 - First observations after ~12h on Sep 24th, non-optimal weather, only 1h used for flux upper-limit estimation
 - ➢ Regular observations from Sep 28th to Oct 4th
 - \rightarrow detection reported on ATel#10817

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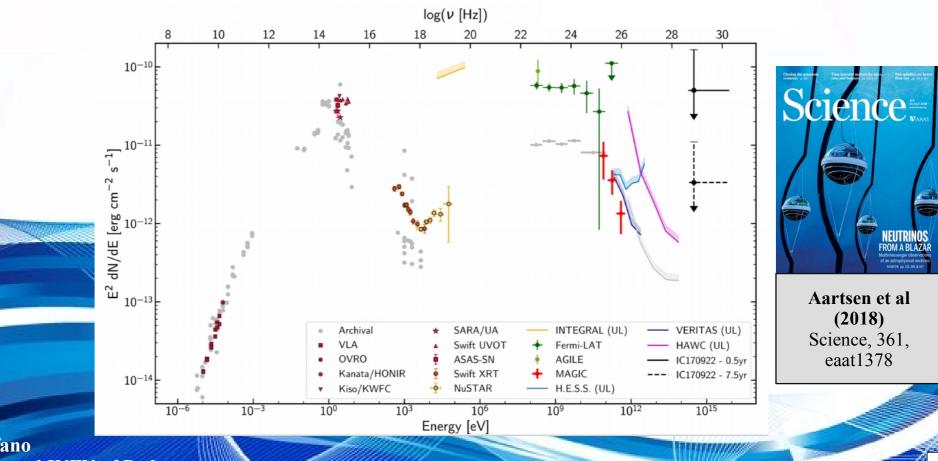
First multi-wavelenght and multi-messenger SED!











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Further MAGIC observations of the blazar TXS 0506+056

Further observations with MAGIC of the blazar TXS 0506+056 resulted in:

- More than **41h** of good quality data collected between Sep 24th Nov 2nd
- ➤ 2 flares detected on Oct 3rd- 4th and on Oct 31st
- Lower state detected above 90 GeV

Light-curve:

- day-scale variability
- constant flux hypothesis below 0.3% probability

> Spectrum:

- no spectral index variability measured
- simple PL, index softer than *Fermi*-LAT
- additional internal absorption at play (apart from EBL effect)

The MAGIC Collaboration, ApJL 863, 1, arXiv:1807.04300

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10

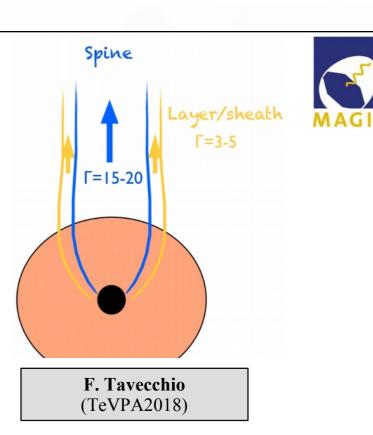
The MAGIC collaboration provided an interpretation of the neutrino-blazar coincidence in terms of the spine-layer model (Ghisellini+2005)

- Components:
 - leptonic (synchrotron, SSC, EC)
 - and hadronic (photo-meson cascade, BH cascade, synch. rad. from pions and muons)

The MAGIC Collaboration, ApJL 863, 1, arXiv:1807.04300

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Interpretation

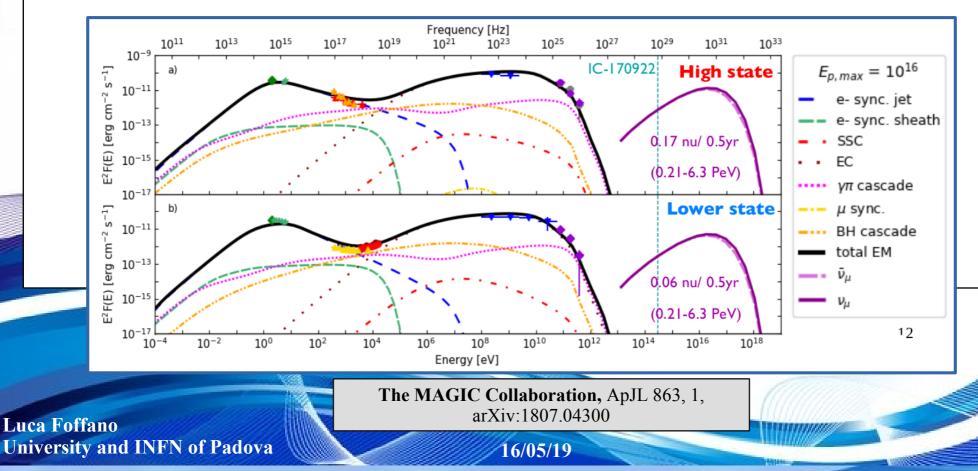


- ▷ **Day-scale variability** \rightarrow Size of emitting region ~10¹⁶ cm
- > Spectral break at VHE compatible with internal absorption: $\tau_{\gamma\gamma} \sim 1$ at $E_{\gamma} \sim 100$ GeV
- **Constraints on max proton energy** $E_{p,max}$ set by X-ray and VHE gamma-ray data:
- > Scan of $E_{p,max}$:10¹⁴ -10¹⁸ eV in the co-moving frame



11

\rightarrow TXS 0506+056 able to accelerate CR to UHE!









13

Blazars and neutrinos

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14

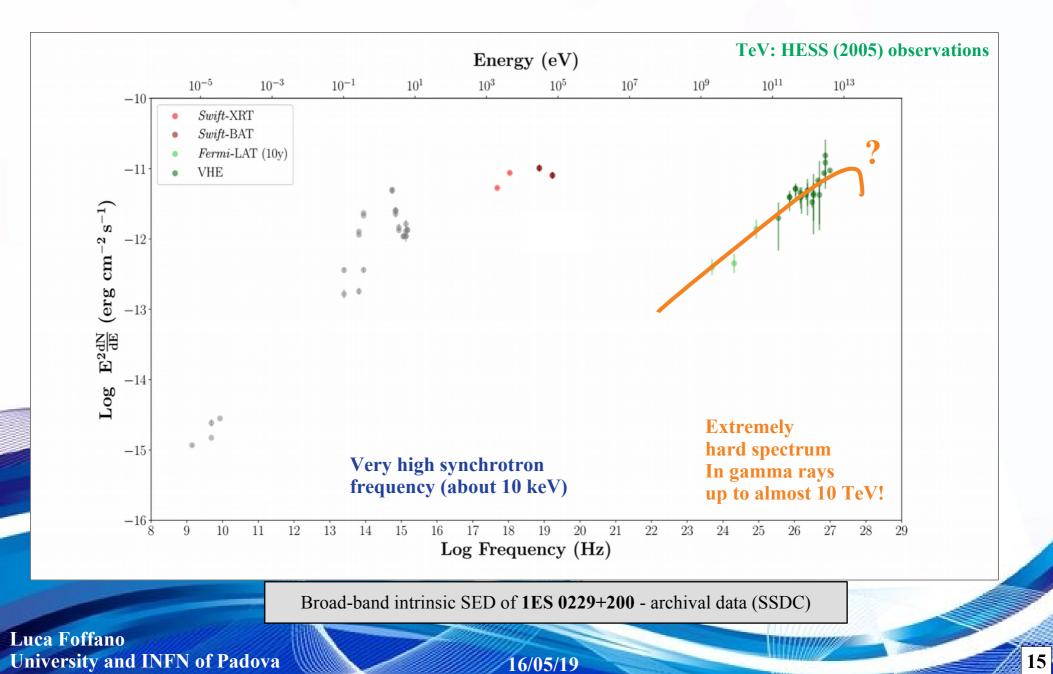
Extreme blazars and neutrinos

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Extreme blazars







Extreme blazars



- They produce the **highest photon energies** ever observed in blazars
- Only few objects currently classified as extreme blazars in gamma rays
- They need large integration time to be detected in gamma rays
- We need **MWL studies** in order to select them
- Their emission mechanism could be contributed by hadronic processes \rightarrow **neutrinos**
- Their spectral energy distribution is challenging for the standard modeling of blazars

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Looking for new extreme blazars



within the MAGIC collaboration

Mrk 501 during some flaring activities showed EHBL behaviour with synchrotron peak above 10¹⁷ Hz

 \rightarrow Being EHBL might be a temporary state

Astronomy & Astrophysics manuscript no. Mrk501_MW2012 August 16, 2018 ©ESO 2018

Ahnen M. L., et al., 2018, A&A

The extreme HBL behaviour of Markarian 501 during 2012

M. L. Ahnen¹, S. Ansoldi^{2,19}, L. A. Antonelli³, C. Arcaro⁴, A. Babić⁵, B. Banerjee⁶, P. Bangale⁷, U. Barres de Almeida^{7,22}, J. A. Barrio⁸, J. Becerra González⁹, W. Bednarek¹⁰, E. Bernardini^{11,23}, A. Berti^{2,24}, W. Bhattacharyya¹¹, O. Blanch¹², G. Bonnoli¹³, R. Carosi¹³, A. Carosi³, A. Carosi³, A. Carosi⁴, A. Berti^{2,24}, W. Bhattacharyya¹¹, O. Blanch¹², G. Bonnoli¹³, R. Carosi¹³, A. Carosi³, A. Carosi⁴, A. Berti^{2,24}, W. Bhattacharyya¹¹, O. Blanch¹², B. Burli¹³, B. Carosi¹³, A. Carosi³, A. Carosi⁴, A. Barti⁴, A. Barti⁴, B. Carosi⁴, A. Barti⁴, B. Carosi⁴, A. Barti⁴, B. Carosi⁴, B. Carosi⁴, B. Carosi⁴, A. Carosi⁴, A. Carosi⁴, A. Barti⁴, B. Carosi⁴, B. Carosi⁴, A. Barti⁴, B. Carosi⁴, B. Carosi⁴,

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COMING SOON

18

Looking for new extreme blazars



within the MAGIC collaboration

The MAGIC collaboration is preparing a mini catalog paper with the results on the observations of new EHBLs

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COMING SOON

19

Looking for new extreme blazars



First detection of very-high-energy gamma-ray emission from the extreme blazar PGC 2402248 with the MAGIC telescopes

ATel #11548; Razmik Mirzoyan (Max-Planck-Institute for Physics, Munich), on behalf of the MAGIC collaboration

Paper on PGC 2402248 and modeling within the extreme blazar context

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20

Looking for new extreme blazars

Presented in L. Foffano, E. Prandini, A. Franceschini, S. Paiano (2019)

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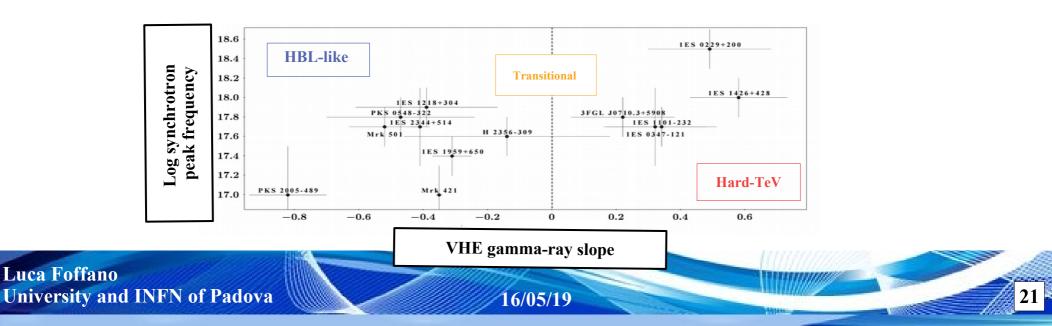


Results



Some results of this analysis:

- EHBLs might present **sub-classifications** depending on their **TeV gamma-ray properties**
- Such differences might be correlated to **differences in the particle acceleration mechanism**!
- MWL studies are needed
- We need good candidates to be observed in the TeV gamma-ray band
- In our work, we propose a sample of new EHBL candidates with promising properties in TeV gamma rays



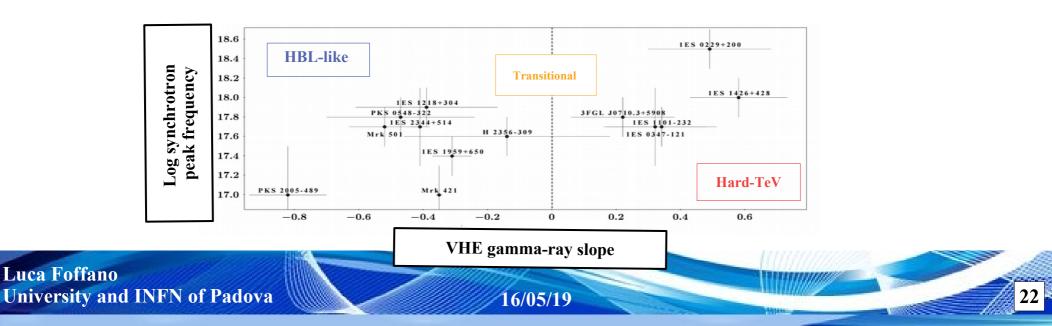


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Looking for new TeV EHBL candidates



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Energy (eV) 10^{10} 10^{2} 10^{6} 10^{-6} 10^{-2} 10^{-9} 1ES 0229+200 1RXS J225146.9-320614 NVSS 10^{-10} WISE Swift-UVOT Swift-XRT ${\rm E}^{2{\rm dN}\over{
m dE}}~({
m erg}~{
m cm}^{-2}~{
m s}^{-1})$ 10^{-11} Swift-BAT Fermi-LAT 10^{-12} 10^{-13} 10^{-14} 10^{-15} $10^{-16} \downarrow 10^{6}$ 10^{14} 10^{16} 1018 1020 1010 10^{12} 10^{22} 10^{24} 10^{26} 10^{28} 10^{8} Frequency (Hz)

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HIMM

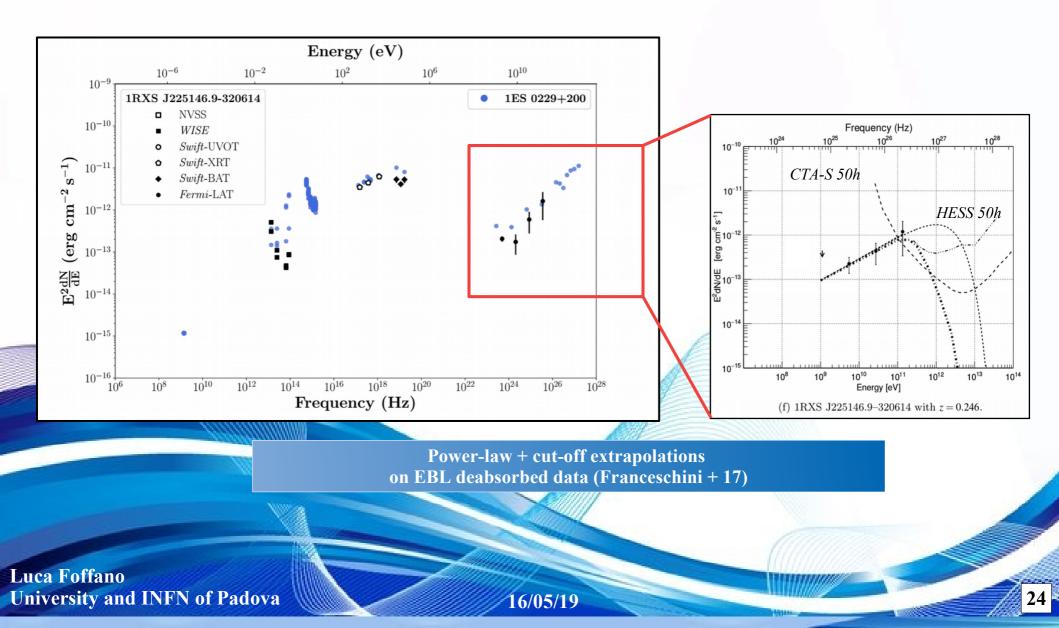
23



Looking for new TeV EHBL candidates



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Conclusions



- We are facing with the opening era of **multi-messenger** and **multi-wavelenght** Astrophysics
- Science will gain a lot from this
- We need to work on the interface between the different messengers and waveleghts, and their instruments

Our contribute in this exciting field:

- We studied a new sample of extreme blazars, finding hints of **sub-classification**
- We propose new targets to be observed with multi-wavelenght observations
- These objects might reveal different physical **interpretation** in term of jet emission models

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Thank you!