Magnetic Bubble Chambers

Surjeet Rajendran with Phil Bunting, Giorgio Gratta, Michael Nippe, Jeffrey Long, Rupak Mahapatra and Tom Melia





Coherence time of signal too short for phase measurement to work. Energy deposition too small to be been using conventional WIMP calorimeters



Coherence time of signal too short for phase measurement to work. Energy deposition too small to be been using conventional WIMP calorimeters

Need amplification of deposited energy (meV - keV)



Coherence time of signal too short for phase measurement to work. Energy deposition too small to be been using conventional WIMP calorimeters

Need amplification of deposited energy (meV - keV)

Challenge: Need large target mass. Rare dark matter event. Requires amplifier stability > years

Consider magnet with all spins aligned





Consider magnet with all spins aligned

Spins now in metastable excited state with energy

~ g µ B



Consider magnet with all spins aligned

Spins now in metastable excited state with energy \sim g μ B

Dark Matter collides, deposits heat. Causes meta-stable spin to flip



Consider magnet with all spins aligned

Spins now in metastable excited state with energy ~g μ B

Dark Matter collides, deposits heat. Causes meta-stable spin to flip

Spin flip releases stored Zeeman energy (exothermic). Released energy causes other spins to flip, leading to magnetic deflagration (burning) of material.



Consider magnet with all spins aligned

Spins now in metastable excited state with energy ~g μ B

Dark Matter collides, deposits heat. Causes meta-stable spin to flip

Spin flip releases stored Zeeman energy (exothermic). Released energy causes other spins to flip, leading to magnetic deflagration (burning) of material.





Consider magnet with all spins aligned

Spins now in metastable excited state with energy ~g μ B

Dark Matter collides, deposits heat. Causes meta-stable spin to flip

Spin flip releases stored Zeeman energy (exothermic). Released energy causes other spins to flip, leading to magnetic deflagration (burning) of material.





Consider magnet with all spins aligned

Spins now in metastable excited state with energy \sim g μ B

Dark Matter collides, deposits heat. Causes meta-stable spin to flip

Spin flip releases stored Zeeman energy (exothermic). Released energy causes other spins to flip, leading to magnetic deflagration (burning) of material.



Amplifies deposited energy. Like a bubble chamber. Is this possible? Stability?







Will not happen in a ferromagnet - spins are strongly coupled.



Will not happen in a ferromagnet - spins are strongly coupled.

Need weak spin-spin coupling. But need large density - necessary for heat conduction. Can't use gas.



Will not happen in a ferromagnet - spins are strongly coupled.

Need weak spin-spin coupling. But need large density - necessary for heat conduction. Can't use gas.

Organo-Metallic complexes. Central metal complex surrounded by organic material.







Need weak spin-spin coupling. But need large density - necessary for heat conduction. Can't use gas.

Organo-Metallic complexes. Central metal complex surrounded by organic material.

Weak coupling between adjacent metal complexes - but still large density





Will not happen in a ferromagnet - spins are strongly coupled.

Need weak spin-spin coupling. But need large density - necessary for heat conduction. Can't use gas.

Organo-Metallic complexes. Central metal complex surrounded by organic material.

Weak coupling between adjacent metal complexes - but still large density

Each molecule acts as an independent magnet





Will not happen in a ferromagnet - spins are strongly coupled.

Need weak spin-spin coupling. But need large density - necessary for heat conduction. Can't use gas.



Weak coupling between adjacent metal complexes - but still large density

Each molecule acts as an independent magnet

Recently discovered systems. Few 100 known examples. Can make large samples. Magnetic deflagration experimentally observed and well studied in Manganese Acetate complexes



Magnetic Deflagration



System well described by 2 level Hamiltonian. Two states separated by energy barrier.

Turn on magnetic field, metastable state decays to ground state through tunneling

Magnetic Deflagration



System well described by 2 level Hamiltonian. Two states separated by energy barrier.

Turn on magnetic field, metastable state decays to ground state through tunneling

 $\tau \propto \tau_0 \exp\left(U_{\rm eff}/T\right)$

Magnetic Deflagration



System well described by 2 level Hamiltonian. Two states separated by energy barrier.

Turn on magnetic field, metastable state decays to ground state through tunneling

 $\tau \propto \tau_0 \exp\left(U_{\rm eff}/T\right)$

Ultra-long lived state at low temperature - localized heating rapidly decreases life-time, decay results in more energy release

Condition for Deflagration

Initially heat region of size λ to T



Thermal Diffusion, lowers T

Spin flips, releases energy, increases T

Condition for Deflagration

Initially heat region of size λ to T



Thermal Diffusion, lowers T

 $au_{
m D} \propto \lambda^2$

Spin flips, releases energy, increases T

 $\tau \propto \tau_0 \exp\left(U_{\rm eff}/T\right)$

Condition for Deflagration

Initially heat region of size λ to T



Thermal Diffusion, lowers T

 $au_{\rm D} \propto \lambda^2$

Spin flips, releases energy, increases T

 $\tau \propto \tau_0 \exp\left(U_{\rm eff}/T\right)$

Deflagration occurs as long as we heat a sufficiently large region



Initially heat region of size λ to T



Thermal Diffusion, lowers T

 $au_{\rm D} \propto \lambda^2$

Spin flips, releases energy, increases T

 $\tau \propto \tau_0 \exp\left(U_{\rm eff}/T\right)$

Deflagration occurs as long as we heat a sufficiently large region

 U_{eff} and τ_0 sets the detector threshold. Short τ_0 and small U_{eff} means tiny energy deposit will sufficiently heat up material to trigger deflagration. Low threshold



Initially heat region of size λ to T



Thermal Diffusion, lowers T

 $au_{
m D} \propto \lambda^2$

Spin flips, releases energy, increases T

 $\tau \propto \tau_0 \exp\left(U_{\rm eff}/T\right)$

Deflagration occurs as long as we heat a sufficiently large region

U_{eff} and τ₀ sets the detector threshold. Short τ₀ and small U_{eff} means tiny energy deposit will sufficiently heat up material to trigger deflagration. Low threshold

Known examples with $\tau_0 \sim 10^{-13}$ s, $U_{eff} \sim 70$ K, enabling 0.01 eV thresholds

High energy (> MeV) background from radio-active decays.

Detect MeV events using conventional means. Actual background at low energy very low - forward scattering of compton events

Problem: MeV events will constantly set off detector. Reset time vs operation time? Big problem for bubble chambers like COUPP

Expected background ~ 1/(m² s). Initial detector size ~ (10 cm)³ (kg mass), 1 background event ~ 100 s

High energy (> MeV) background from radio-active decays.

Detect MeV events using conventional means. Actual background at low energy very low - forward scattering of compton events

Problem: MeV events will constantly set off detector. Reset time vs operation time? Big problem for bubble chambers like COUPP

Expected background ~ 1/(m² s). Initial detector size ~ (10 cm)³ (kg mass), 1 background event ~ 100 s



With precision magnetometers, don't need entire crystal to flip

High energy (> MeV) background from radio-active decays.

Detect MeV events using conventional means. Actual background at low energy very low - forward scattering of compton events

Problem: MeV events will constantly set off detector. Reset time vs operation time? Big problem for bubble chambers like COUPP

Expected background ~ 1/(m² s). Initial detector size ~ (10 cm)³ (kg mass), 1 background event ~ 100 s



With precision magnetometers, don't need entire crystal to flip

Within ~ 10 µs, flame ~ 10 - 100 µm. Visible with SQUID.

High energy (> MeV) background from radio-active decays.

Detect MeV events using conventional means. Actual background at low energy very low - forward scattering of compton events

Problem: MeV events will constantly set off detector. Reset time vs operation time? Big problem for bubble chambers like COUPP

Expected background ~ 1/(m² s). Initial detector size ~ (10 cm)³ (kg mass), 1 background event ~ 100 s



With precision magnetometers, don't need entire crystal to flip

Within ~ 10 µs, flame ~ 10 - 100 µm. Visible with SQUID.

Shut off B, turn off fuel. Deflagration stops. Lose ~ (10 -100 μm)³ of volume every 100 s.

Potential Reach





Absorption obtained from photoabsorption. Exposure of I kg-year

Trial using Mn-Ac Hall Sensor Reversible B



Two sets of Mn12-Ac and Hall sensors

One with μ Ci Am 241 α source One without source

Trial using Mn-Ac Hall Sensor Reversible B



Two sets of Mn12-Ac and Hall sensors

One with μ Ci Am 241 α source One without source

Metastability? Deflagration?

Results



Avalanche only observed with source

Mn12-Ac has high threshold (~ few MeV) - using new materials now

Conclusions



Poor observational constraints on dark matter

New ideas needed to probe dark matter between 10-4 eV - GeV

Single Molecular Magnets offer unique opportunity for amplifiers in this range with suitable stability

> Promising initial results with Mn12-ac Stay tuned for tests with other compounds