# DAVIDE PINCI - INFN, SEZIONE DI ROMA

# SEARCH FOR LIGHT DARK MATTER WITH GAS DETECTORS: LATEST DEVELOPMENTS

#### **INTERACTION PROBABILITY**

- WIMP detectors look at nuclear recoils
- Probability of creating a nuclear recoil

$$R \propto 0.13 \left( \underbrace{A}_{100} \frac{\sigma_{WN}}{10^{-38} \text{cm}^2} \frac{v}{220 \text{ km/s}} \frac{\rho_0}{0.3 \text{ GeV/cm}^3 \text{ kg}} \right) \left[ \frac{\text{events}}{\text{year}} \right]$$
Target Cross Section

(i.e. High Mass Number)

Target Mass

<b>2</b> 4.003			Gas		$A_{eff}$		kg/m³		A x mass/m	3
He Helium	6 C Carbon	Carbon		r	39.9			1.8	71.0	
<b>8</b> 15.999	[He]2s <sup>2</sup> 2p <sup>2</sup> 9 18.998		He		4.0		(	0.2	0.7	
Oxygen 26 [He]2s <sup>2</sup> 2p <sup>4</sup>	Fluorine 27 (He)2s <sup>2</sup> 2p <sup>5</sup>	СС		$\mathbf{D}_2$	44	4.0	1.8	81.0		
<b>18</b> 39.948 Argon 288 [Ne]3s <sup>2</sup> 3p <sup>6</sup>	16 32.066 S Sulfur 286 [Ne]3s <sup>2</sup> 3p <sup>4</sup>		CF <sub>4</sub>		88.0			3.7	327.4	
			SF	6	14	146.0		6.2	900.8	
4 Xe Xenon 131.294		Liqui	d	А	eff	kg/l		A x mass/l		
	-	Ar		39	9.9	1.4		55.9		
	-	Xe		131.3		3.0		393.9		г
										L

Davide Pinci - INFN Sezione di Roma

### **SIGNAL DETECTION**

- > Primary signal in gaseous detector is always due to ionization of gas molecules;
  - Given the large fraction of energy released in ionising processes, the effective average energy release per e-ion pair is 2-3 times larger than single ionization (i.e. 20-50 eV/pair);





- Gas
   eV/pair

   Ar
   26

   He
   42

   CO2
   33

   CF4
   54

   SF6
   51
- In principle energies releases of few tens of eV are clearly detectable.

### **SIGNAL FORMATION**

- In DM detector, the ionizing particles can be nuclear or electron recoils;
- The maximum energy that can be transferred by a DM particle to a recoil of mass m<sub>N</sub>, is given by:

$$\epsilon = \frac{4\rho}{\left(\rho+1\right)^2}$$
 where  $\rho = \frac{m_N}{m_{WIMP}}$ 

Element	Max E transferred by a 1GeV DM particle	Min DM particle mass with 1 keV threshold
Ar	0.2 keV	5.25 GeV
He	1.2 keV	0.78 GeV
С	0.6 keV	1.76 GeV
F	0.4 keV	2.63 GeV
S	0.2 keV	4.25 GeV
Xe	0.06 keV	16.6 GeV

The use of light atoms would allow larger Energy Transfers

### **ENERGY MEASUREMENT**

Gaseous detectors operated in proportional regime have a good (not exceptional) energy resolution:





Davide Pinci - INFN Sezione di Roma



High energy particles create very long tracks in gas (from centimetres to meters) with a very clear direction.



Low energy particles create sizeable tracks in gas (from millimetres to centimetres) with a less clear direction.



Particle range can be tuned by the choice of the gas mixture

#### **BACKGROUND REJECTION**

#### Low radioactivity:

Gases and the light structures of gas vessels (acrylic, copper) are usually less radioactive than more complex apparati (e.g. cryogenic ones);

#### Directionality:

3D reconstruction of tracks to exploit asymmetries and direction modulation of the Dark Matter signals for background discrimination;

#### Sensitive Volume Fiducialization:

Measurement of the event position within the detector to reject events close the detector walls, very likely due to decays of radioactive materials

#### Particle identification:

dE/dx of electrons and nuclear recoils are quite different giving rise to different signal development in space and time.

Davide Pinci - INFN Sezione di Roma

threshold, excellent energy resolution, single readout channel

Spherical Proportional Counters (SPC): low capacitance, low energy

Energy detection thresholds as low as 10 eV

A 60 cm diameter SPC prototype made of high purity copper has been set-up at the Laboratoire Souterrain de Modane (LSM)



**NEWSG** 





#### NEWSG

Electron diffusion translates into pulses exhibiting longer rise times for energy depositions occurring further away from the sensor



Pulse shape analysis allows the discrimination, if operated at low pressure (100 mbar), between point-like nuclear recoils and short tracks due to electron recoils

on the oss section eV.

1

2

3

WIMP Mass [GeV]

56

4

10<sup>41</sup> 5×10

World-leading constraint on the WIMP elastic scattering cross section for a WIMP mass of 0.5 GeV.

1.4 m diameter detector to be located at SNOLAB in construction

7 8 9 1 0

section [pb

له مorecon cross الم

10

10<sup>-5</sup>

#### WHY TPC?

Time Projection Chambers provide:



total released energy measurement;



3D tracking (position and direction);



dE/dx profile (pid, head-tail);

relatively small amount of readout channels;





- 1.5 m<sup>3</sup> steel
   vacuum vessel
- target gas, typically 30 Torr CS<sub>2</sub> + 10 Torr CF<sub>4</sub>
- MWP readout



#### Located in Boulby mine.

Demonstrates the use of dE/dx measurement, tuned so the detector essentially never triggers on the large background of Compton recoil electrons yet retains high efficiency for nuclear recoils;

A nuclear recoil detection efficiency of ~60% was achieved with a gamma rejection factor, measured using <sup>137</sup>Cs, of 2×10<sup>-7</sup>

### **NEGATIVE ION DRIFT**

With penetrating tracks (e.g. HEP) an external trigger can be exploited to measure the electron drift time to evaluate the event "Z";

In DM applications an alternative technique is the "negative ion drift":

- ionization electrons can be absorbed by a high electronegative component of gas mixture (SF<sub>6</sub>, CS<sub>2</sub>);
- negative ions start to drift in the electric fields;
- at the multiplication stage, electrons are released and avalanches develop;



- The Z of the event is evaluated from the measurement of the difference of the time of arrival of ions with different masses and mobilities

### DRIFT - NEGATIVE ION DRIFT (NID)



Experiment DRIFT observed 185 events in 54.7 shielded day data taking.

All of them were reconstructed to be produced close to the cathode plane and due radioactivity inside the material.

This allowed to set a 90% C.L. upper limits on the WIMP-proton spin-dependent; First directional detector that excluded a fraction of DAMA region.



Davide Pinci - INFN Sezione di Roma





TPC with a Micromegas-based readout;

Charge integrator connected to the mesh coupled to a FADC sampled at 50 MHz

#### Fluorine 6.3 keV (~2 keVee)



Davide Pinci - INFN Sezione di Roma

#### MIMAC - CATHODE SIGNAL TO PLACE 3D-TRACKS

The cathode signal is produced by the primary electrons before the anode signal before the avalanche.



Measurement in a MIMAC chamber of an alpha passing through the active volume parallel to the cathode at 10 cm distance.



#### **NEWAGE**



## µPIC based TPC with CF<sub>4</sub> based gas mixtures





Field Cage with sheet resistor (SR) to overcome distortion of field cage and radioactive background





Better track reconstruction in the whole volume w.r.t wire Field Cage





Davide Pinci - INFN Sezione di Roma

### **D<sup>3</sup>: DIRECTIONAL DARK-MATTER DETECTOR**



Charge diffusion:

- gives a one-to-one relationship between z and charge profile width
- allows for absolute z to be measured within ~ 1 cm





Davide Pinci - INFN Sezione di Roma

### LIGHT: A CHANGE OF PARADIGM

During the multiplication process, photons are produced along with electrons by the gas through atomic and molecular de-excitation;

The CYGNO collaboration proposed to readout the light instead of electric signal.



Optical readout of gas detectors offers several advantages:

- optical sensors are able to provide high granularities along with very low noise level and high sensitivity;
- optical coupling allows to keep sensor out of the sensitive volume (no interference with HV operation and lower gas contamination);
- suitable lens allow to acquire large surfaces with small sensors;

### **CYGNO: SOME IMAGE**

450 MeV electron with a  $\delta$  ray







recoils (a)

#### **PERFORMANCE AT LOW ENERGY**

Light response to 5.9 keV <sup>55</sup>Fe measured with a source 20 cm far from the GEM.



About 1170 photons are detected: i.e. 200 photons per keV released.

Eccentricity: 
$$\rho = \frac{(A/\sigma)_y}{(A/\sigma)_x}$$

 $\rho_{NoDir} = 0.98 \pm 0.01$  and  $\rho_{Dir} = 1.10 \pm 0.03$ 





(pint-like before diffusion) (170 µm before diffusion)

#### FIDUCIALIZATION AND BACKGROUND REJECTION



All information provided by the optical readout, allow to distinguish different particle tracks

10 keV He

MeV electrons due to 4 MeV y

Davide Pinci - INFN Sezione di Roma

He nuclear

recoils (a)

Track shape, brightness, size can be used by suitable algorithms to identify different particles

#### **PMT+CMOS COMBINED READOUT**

Single cluster 3D position reconstruction can be obtained by comparing the light profile along the track (X, Y) and the PMT waveform (t);

A peak finding algorithm was used to highlight the main cluster signals;



Residual distribution to a 3D fit allows to compute a resolution on Z of 100  $\mu$ m.

### **INITIUM: NEGATIVE ION DRIFT WITH OPTICAL READOUT**







After having demonstrated the feasibility of Negative Ion Drift at nearly atmospheric pressure (610 Torr), within the CYGNO collaboration Elisabetta Baracchini (GSSI) won an ERC Consolidator Grant with (INITIUM) for



"the development and operation of the first 1 m<sup>3</sup> Negative Ion TPC (NITPC) with Gas Electron Multipliers (GEMs) amplification [in He/CF<sub>4</sub>/SF<sub>6</sub> mixture] and optical readout with CMOS-based cameras and PMTs"

### **CYGNO DEMOSTRATOR**

1 m<sup>3</sup> of He/CF<sub>4</sub>(/SF<sub>6</sub>) 60/40 (1.6 kg) at atmospheric pressure subdivided in two 50 cm long parts by the cathode with a drift field of about 500 V/cm



Each side equipped by a 3x3 matrix of:

- sCMOS sensor 65 cm away;
- Fast light detector (PMT or SiPM).

One year of operation of CYGNO (expected for 2022) will already explore new low mass regions; A 30-100 m<sup>3</sup> experiment would be able to reach the neutrino floor.



#### CYGNUS

Multi-site Galactic Recoil Directional Observatory for WIMPs and neutrinos; More than 50 signed members UK, Japan, Italy, Spain, China focused on gas TPCs with 2D or 3D direction sensitivity;



THANK YOU!