

GR or not GR?

Testing General Relativity through gravitational wave observations

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Outlook

Introduction to gravitational waves

- GWs in general relativity
- Detection and data analysis

Tests of General Relativity

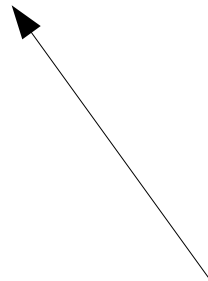
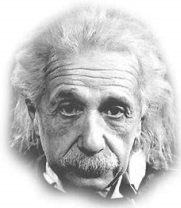
- **Consistency tests**
 - Compatibility of residuals with noise
 - IMR consistency test
 - Ringdown test
- **Production effects**
 - Waveform expansion
 - Polarisation tests
- **Propagation effects**
 - Modified dispersion relation
 - Higher dimensions

Conclusions and perspectives

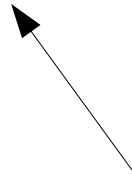
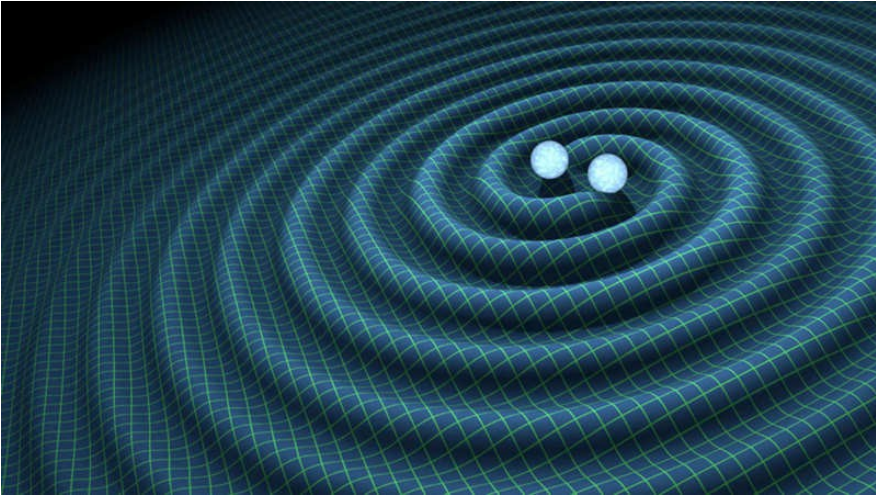
Introduction to gravitational waves



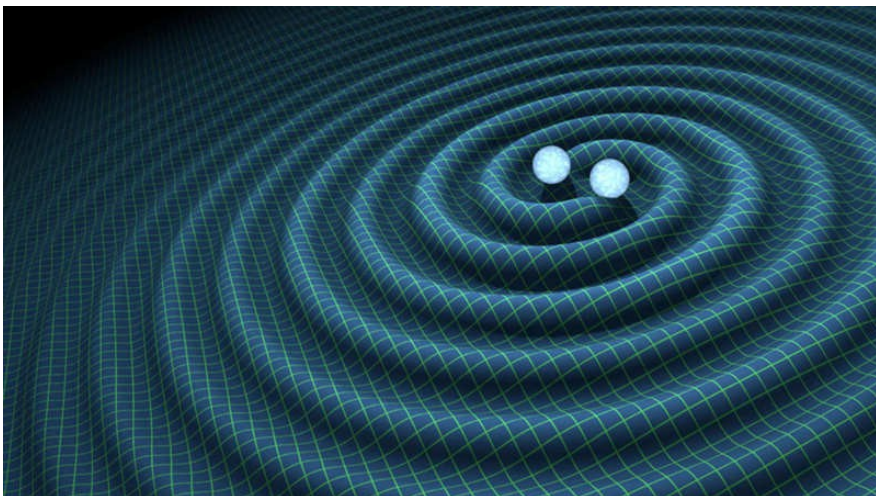
$$\square \bar{h}_{\mu\nu} = -\frac{16\pi G}{c^4} T_{\mu\nu}$$



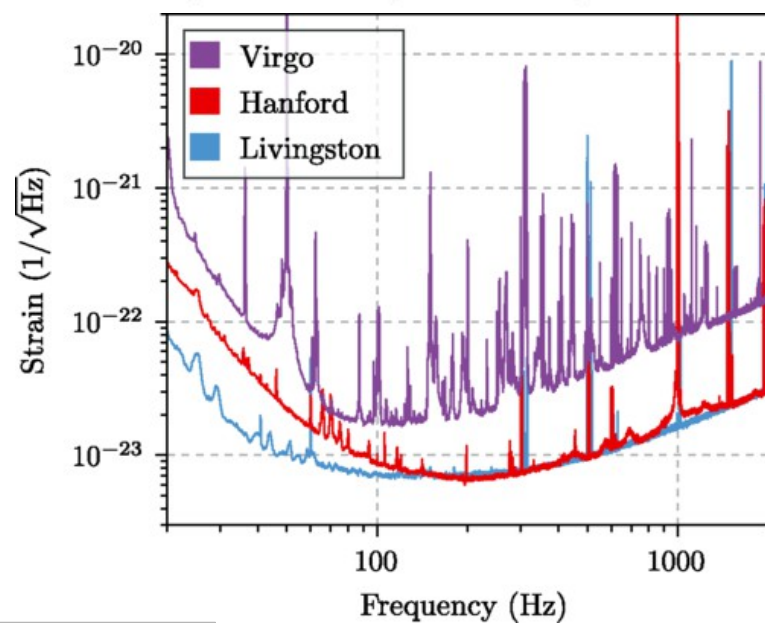
$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}$$



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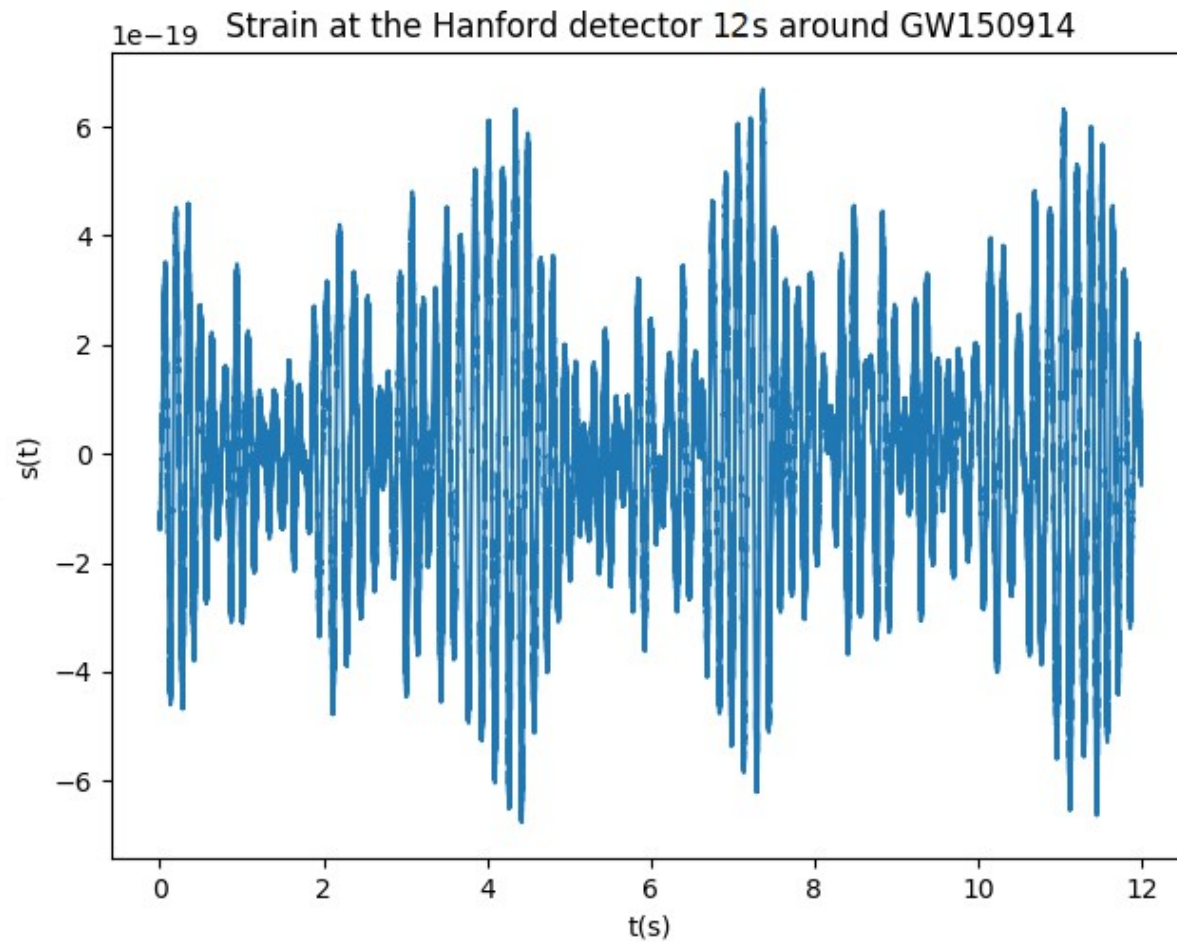


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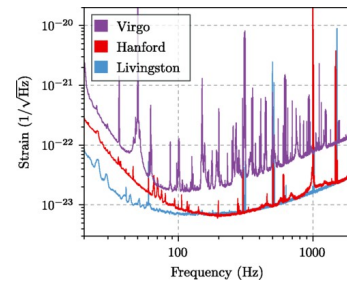
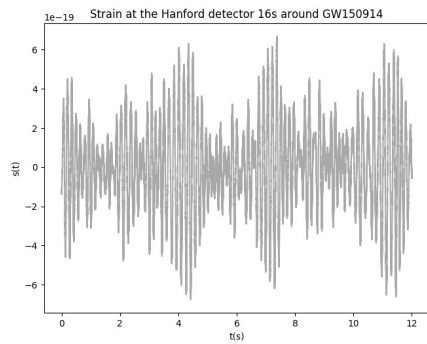
Phys. Rev. Lett., 119, 141101
(2017)

Buried in noise



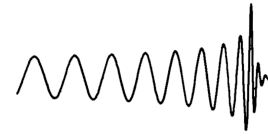
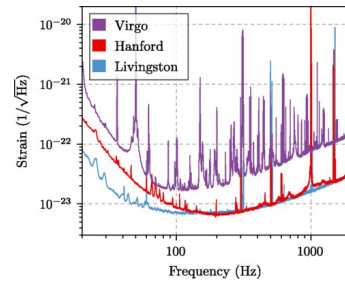
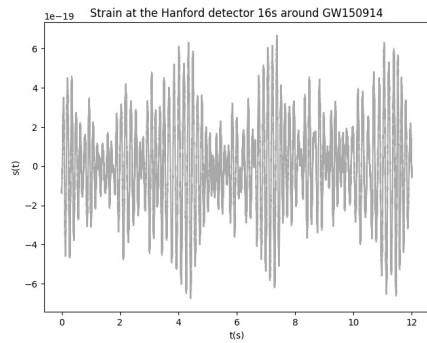
Buried in noise

$$d(t) = n(t) + h(t, \vec{\theta})$$



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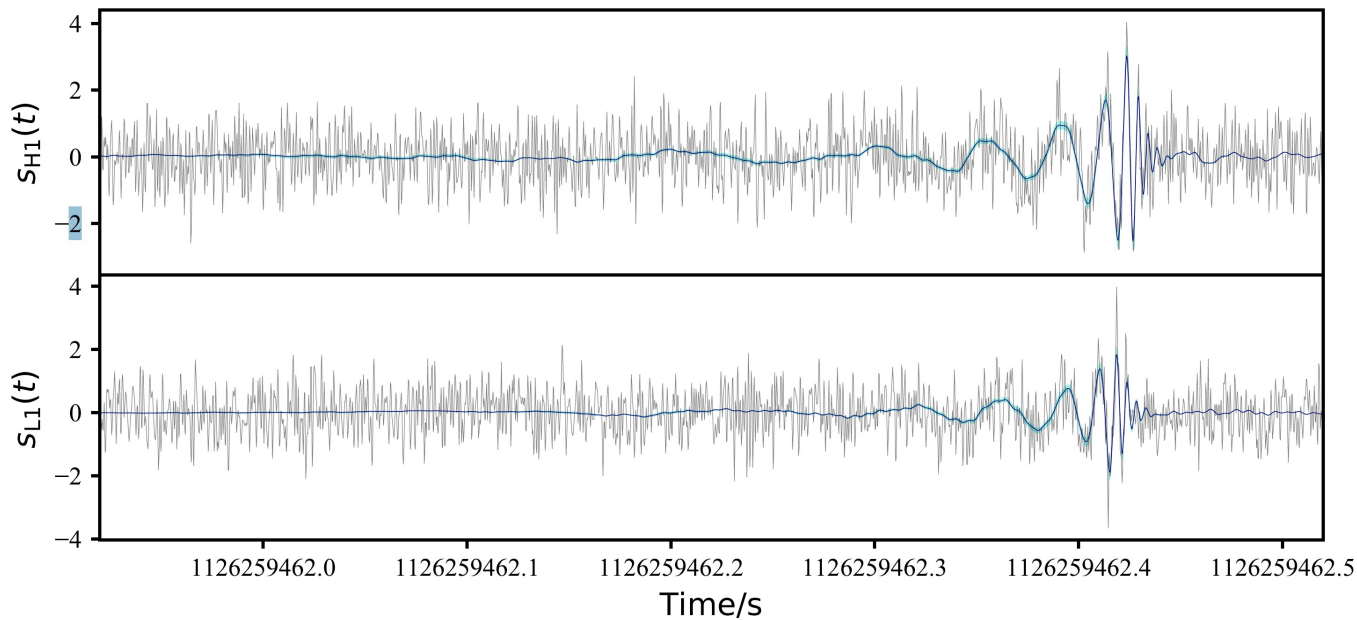
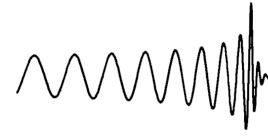
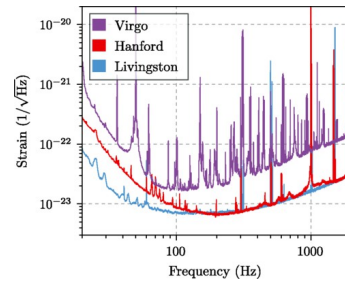
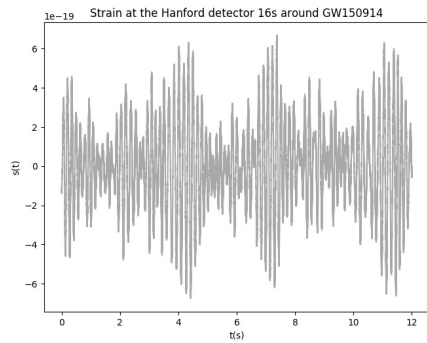
GR



Prediction

Buried in noise

$$d(t) = n(t) + h(t, \vec{\theta})$$



GR



Prediction



Match with data

GR tests: basics

- How to test for the presence of a **GR** waveform in the data?
- How to test for **violations**?

$$d(t) - h(t, \vec{\theta}) = n(t)$$

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 : non **GR**

 $h(t, \vec{\theta})$

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Test residuals
against the
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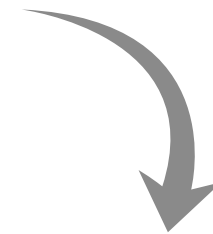
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$h(t, \vec{\theta})$



Test residuals
against the
**distribution of
noise**



• Constrain/measure **parameters**



• Model **selection**

GR tests: Bayesian inference

- **Probability distributions** are recovered for each parameter
- Uses **Bayes' theorem**

Bayes' theorem

$$p(\vec{\theta} | dHI) = p(\vec{\theta} | HI) \cdot \frac{p(d | \vec{\theta} HI)}{p(d | HI)}$$

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The data

$d_H(t), d_L(t), d_V(t)$

The model



GR



non **GR**

Prior information

Any **information** available
before analysing the **data**

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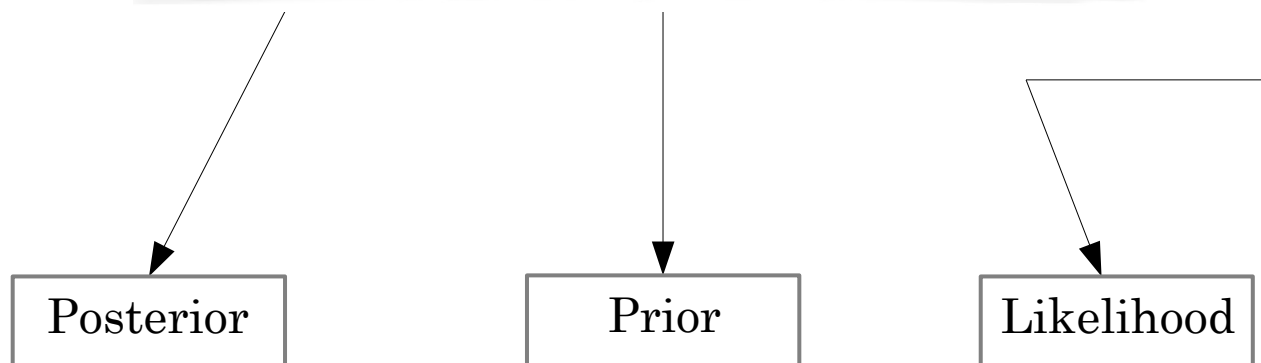
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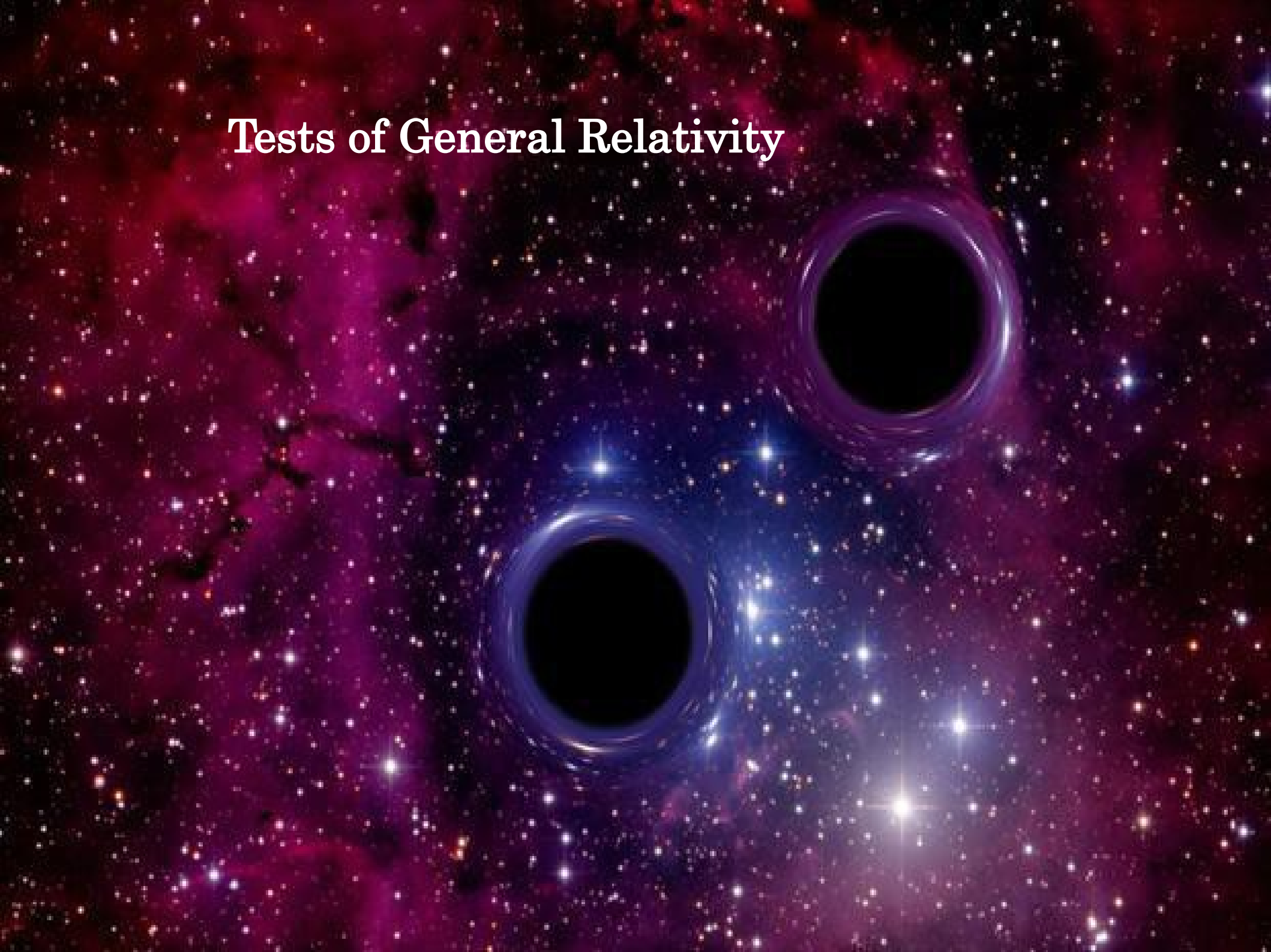
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Tests of General Relativity



Consistency tests

Distribution of the residuals

- Subtract the **most probable GR waveform** from the data
- Analyse the residuals **only** looking for **coherence** (BayesWave)

Bayes' theorem

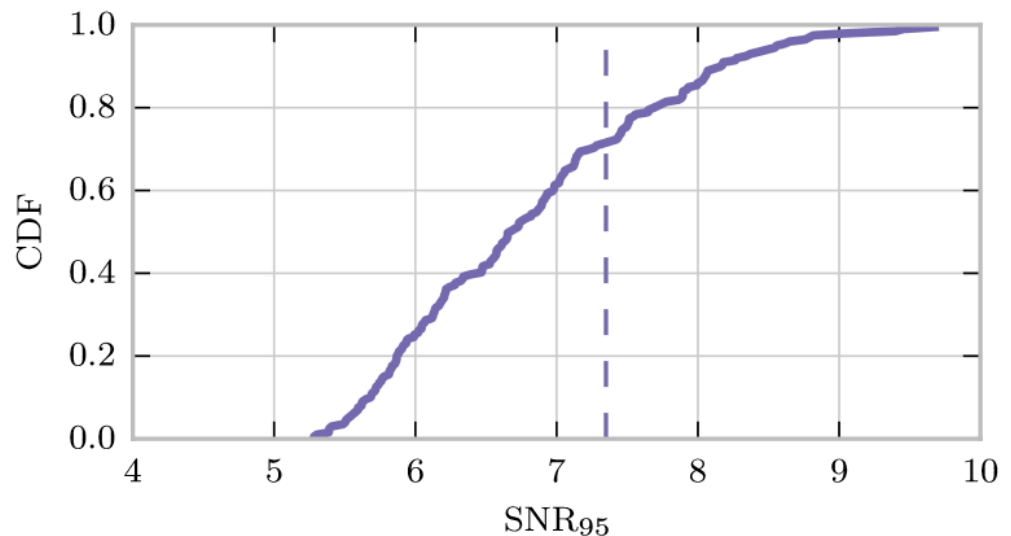
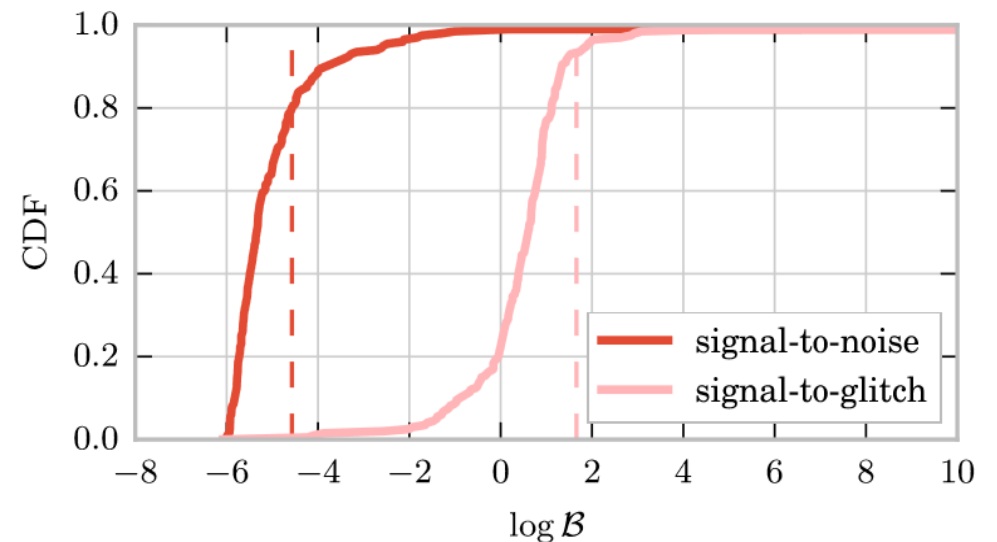
$$p(H_1|DI) = p(H_1|I) \frac{p(D|H_1I)}{p(D|I)}$$

$$p(H_2|DI) = p(H_2|I) \frac{p(D|H_2I)}{p(D|I)}$$

Odds ratio and Bayes factor

$$O_{1,2} \equiv \frac{\cancel{p(H_1|I)} p(D|H_1I)}{\cancel{p(H_2|I)} p(D|H_2I)}$$

GW150914



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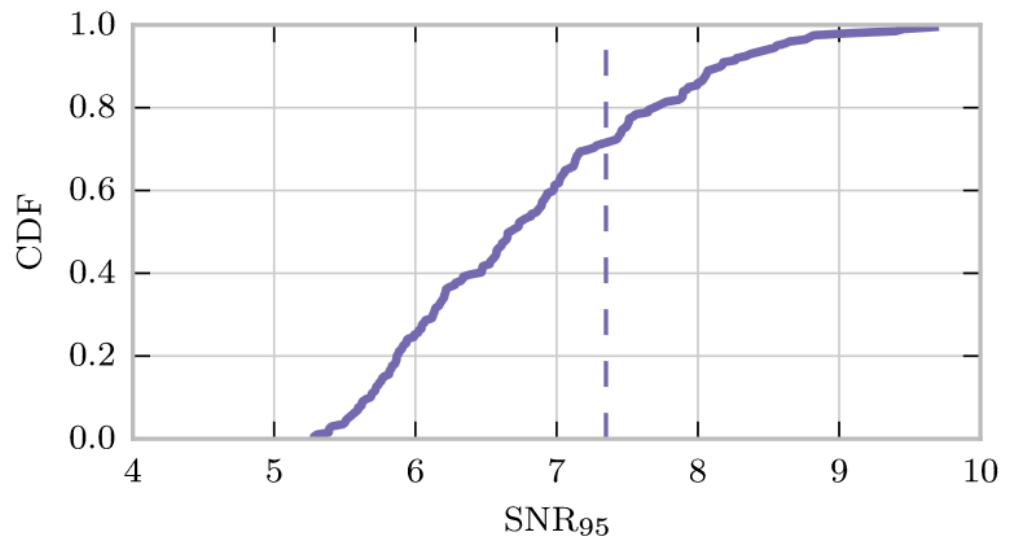
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GW150914

$$\text{SNR}_{\text{res}}^2 = (1 - \text{FF}^2) \text{FF}^{-2} \text{SNR}_{\text{det}}^2$$



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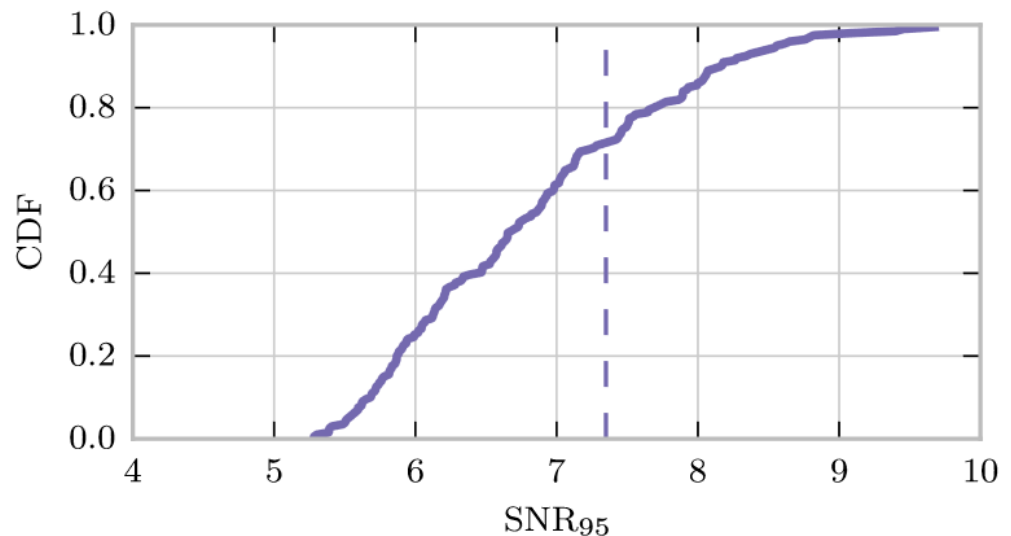
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GW150914

$$\text{SNR}_{\text{res}}^2 = (1 - \text{FF}^2) \text{FF}^{-2} \text{SNR}_{\text{det}}^2$$

$$\text{SNR}_{\text{res}} \leq 7.3 \quad \Rightarrow \quad \text{FF} \geq 0.96$$

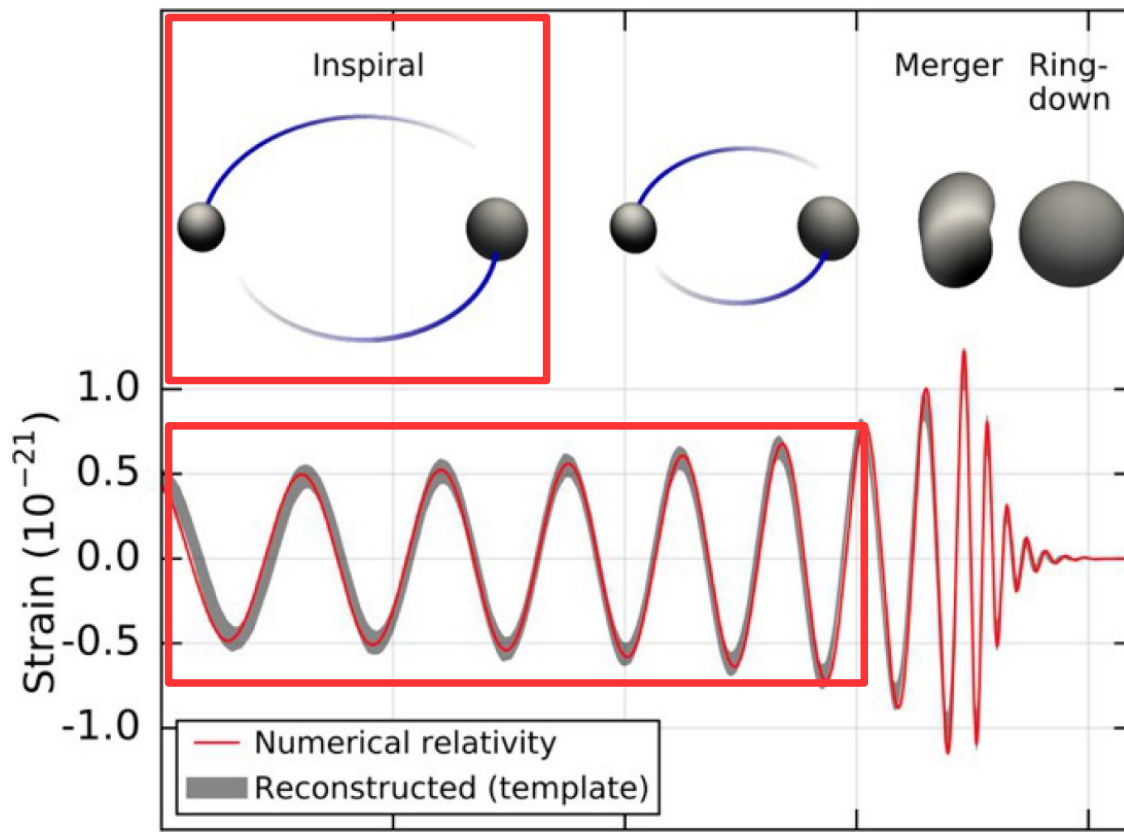
GR deviations less than 4%



Consistency tests

Inspiral-Merger-Ringdown (IMR)

- Test that the **entire coalescence** does not deviate from GR predictions



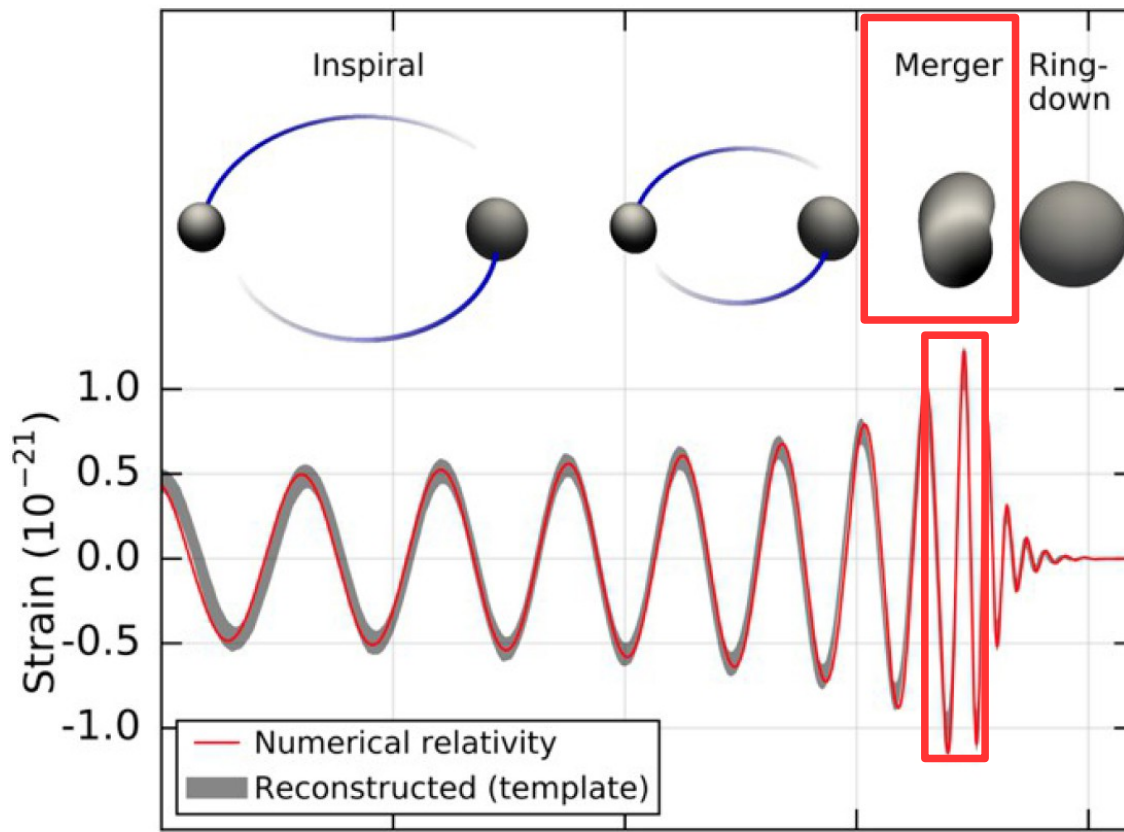
- **Post Newtonian** expansion (PN) of the phase in $v/c \sim f^{1/3}$ up to $(v/c)^7$

Phys. Rev. Lett. 116, 221101 (2016)

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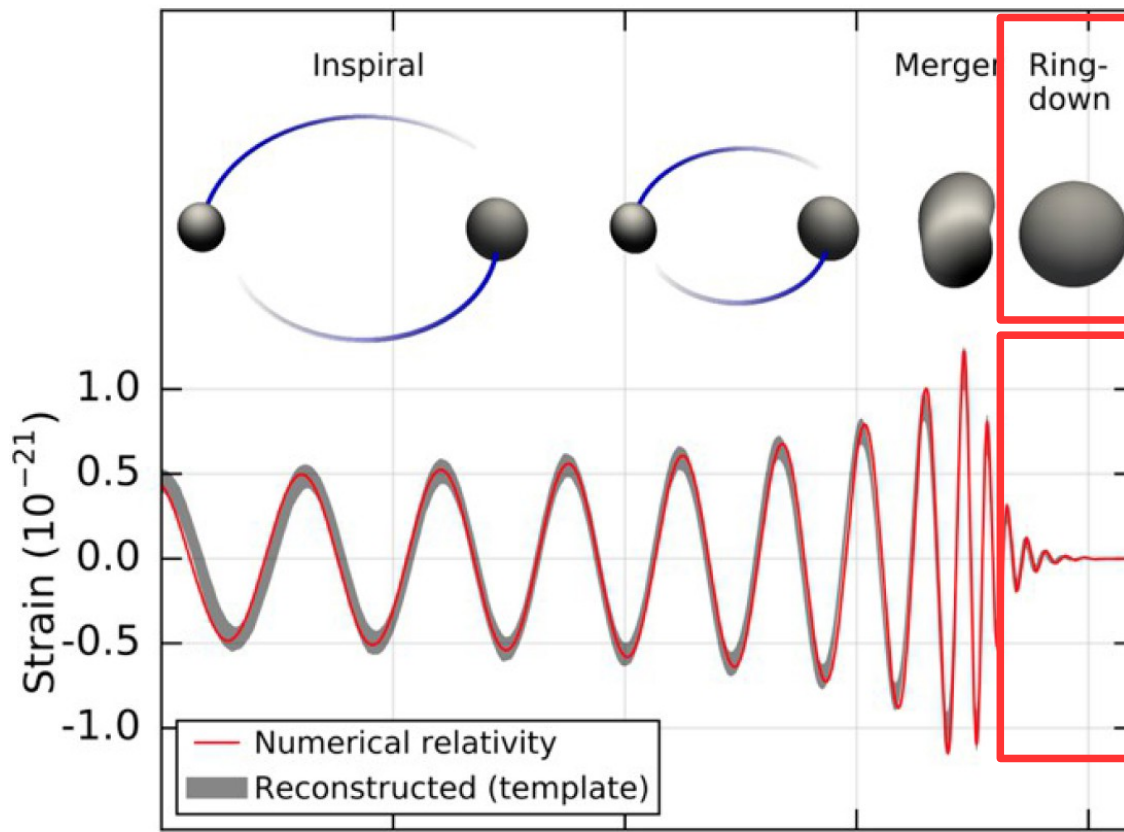
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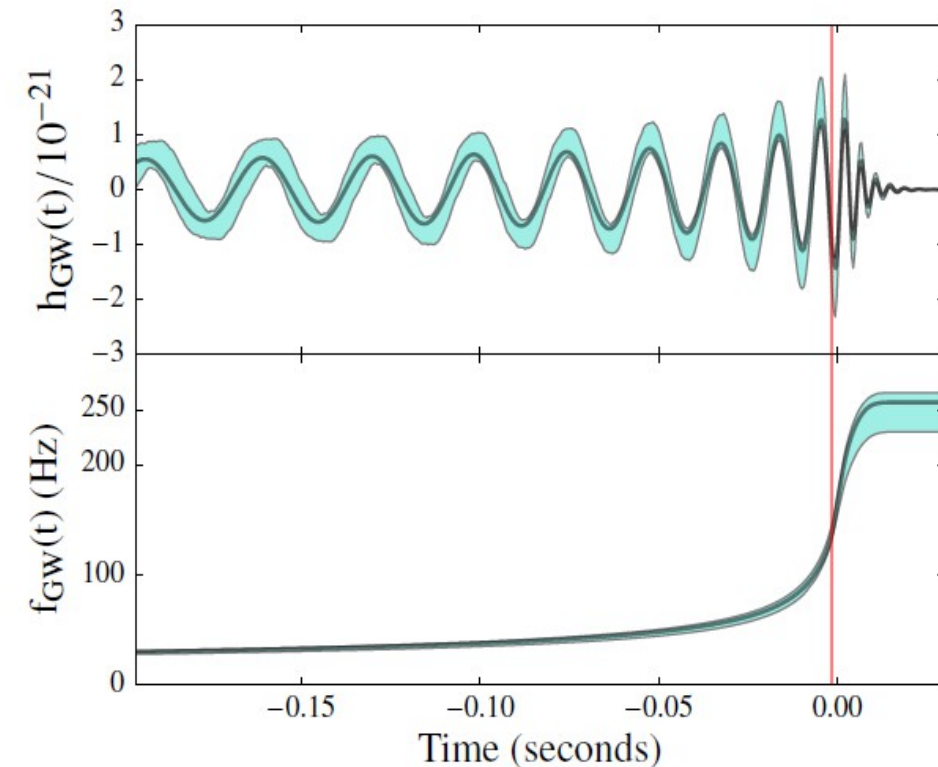
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- **Post Newtonian** expansion (PN) of the phase in $v/c \sim f^{1/3}$ up to $(v/c)^7$
- **Numerical relativity (NR)** regime
- **Perturbation theory**

Consistency tests

Inspiral-Merger-Ringdown (IMR)

- Test that the **entire coalescence** does not deviate from **GR predictions**



Phys. Rev. Lett. 116, 221101 (2016)

- Infer m_1, m_2 and a_1, a_2 from the **low frequency** and the **high frequency part** of the waveform (**Kerr ISCO** cutoff)
- Obtain two **independent estimates** for M_f, a_f through **NR fits**

BUT

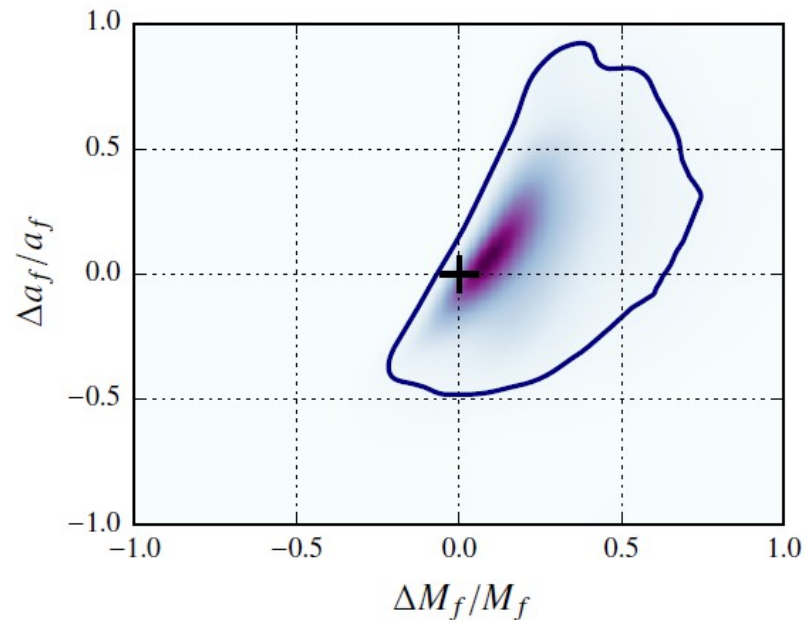
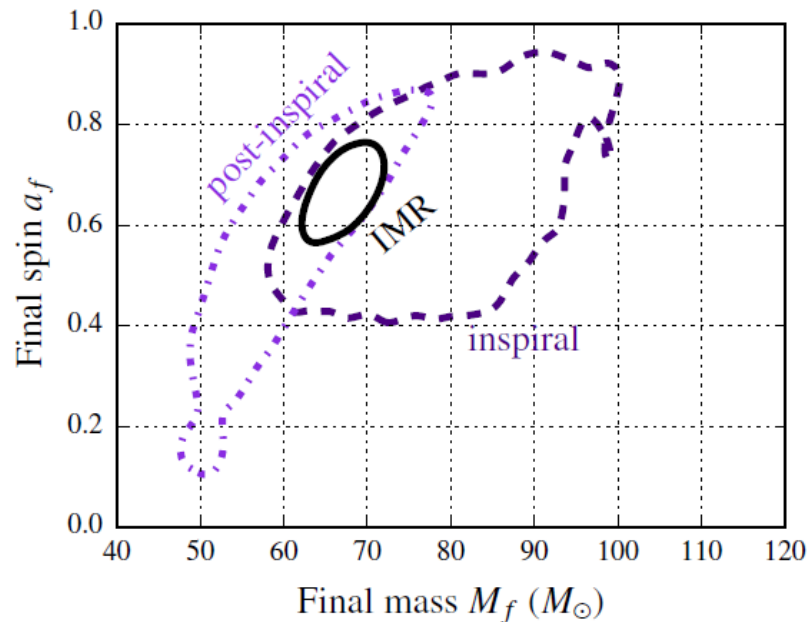
- **GR** predicts a **unique solution**
- **Estimates** must be **compatible**

Consistency tests

Inspiral-Merger-Ringdown (IMR)

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GW150914

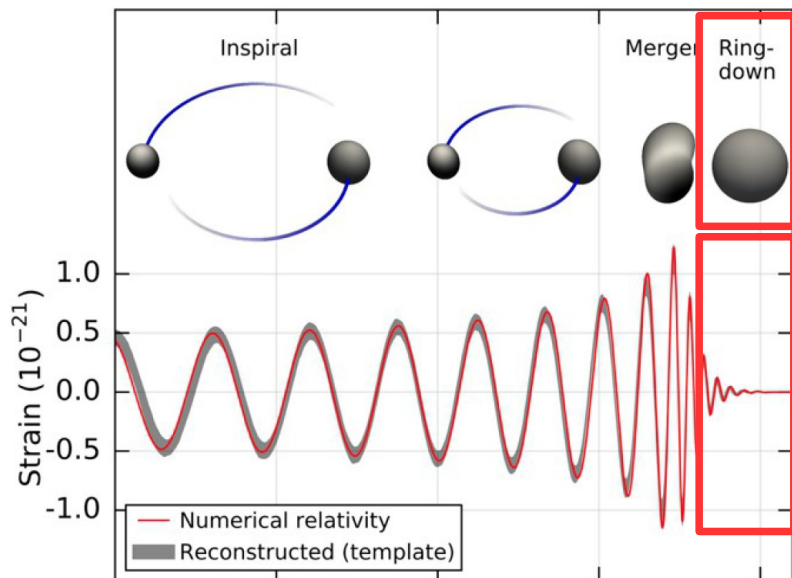


Phys. Rev. Lett. 116, 221101 (2016)

Consistency tests

Ringdown test

- Superposition of **damped sinusoids** when the **linear regime** takes over after the merger



Phys. Rev. Lett. 116, 221101 (2016)

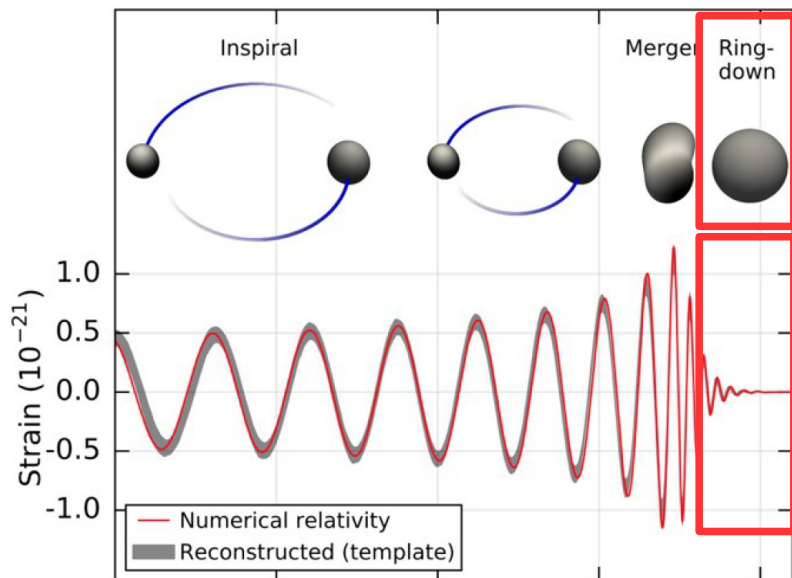
- $l = m = 2$ (spherical harmonics indices) $n = 0$ (overtone) mode is the **least damped**

$$h(t \geq t_0) = A e^{-(t-t_0)/\tau} \cos [2\pi f_0 (t - t_0) + \phi_0]$$

Consistency tests

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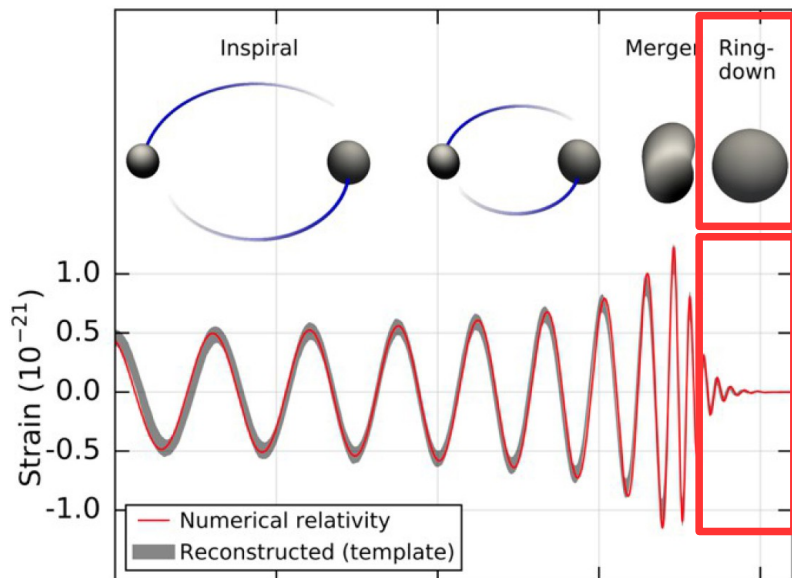
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**predicted by the remnant's
mass and spin**

Consistency tests

Ringdown test

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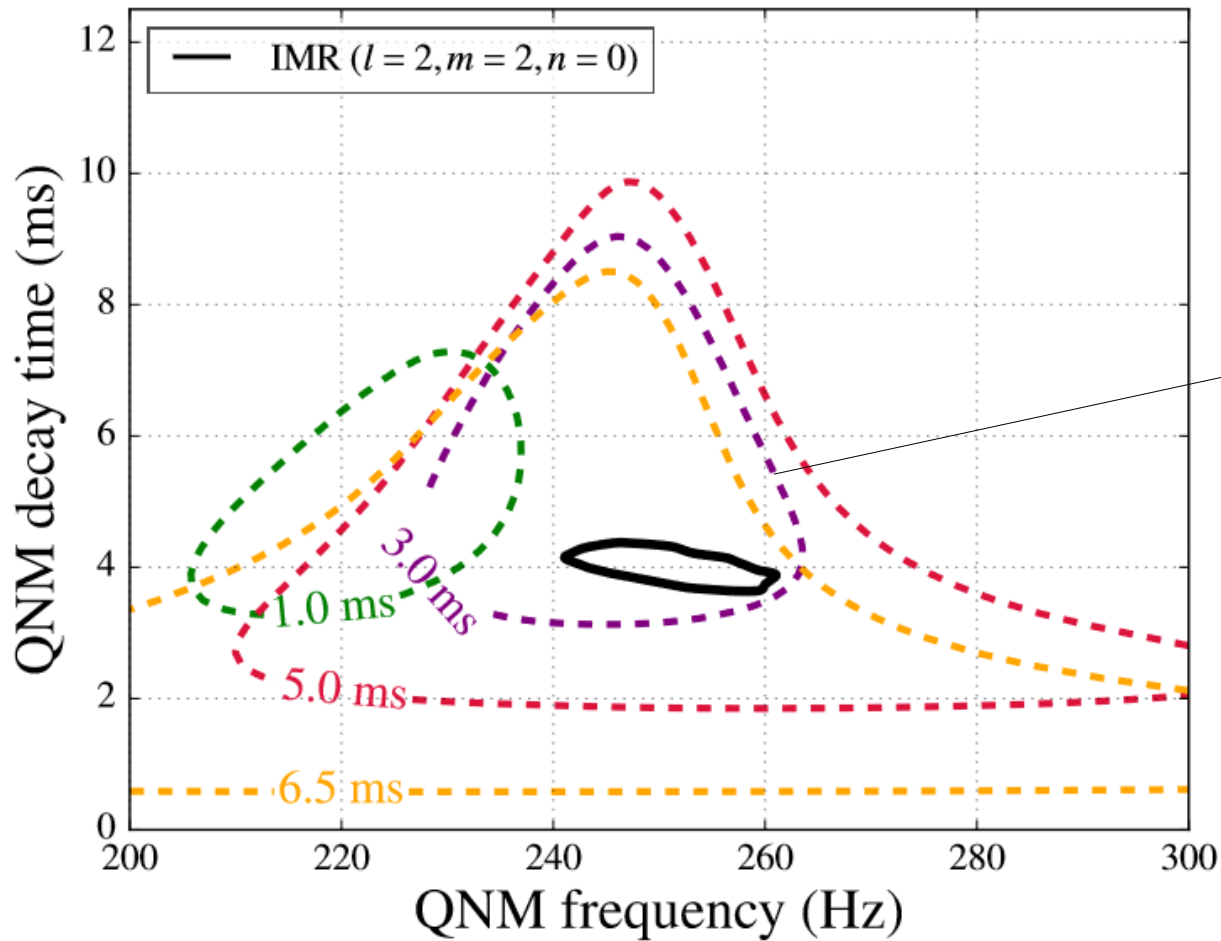
$$h(t \geq t_0) = A e^{-(t-t_0)/\tau} \cos [2\pi f_0 (t - t_0) + \phi_0]$$

- Measure its **frequency** and **damping time**
- **Compare** with the **GR prediction** given by M_f, a_f inferred from **IMR**

Consistency tests

Ringdown test

GW150914



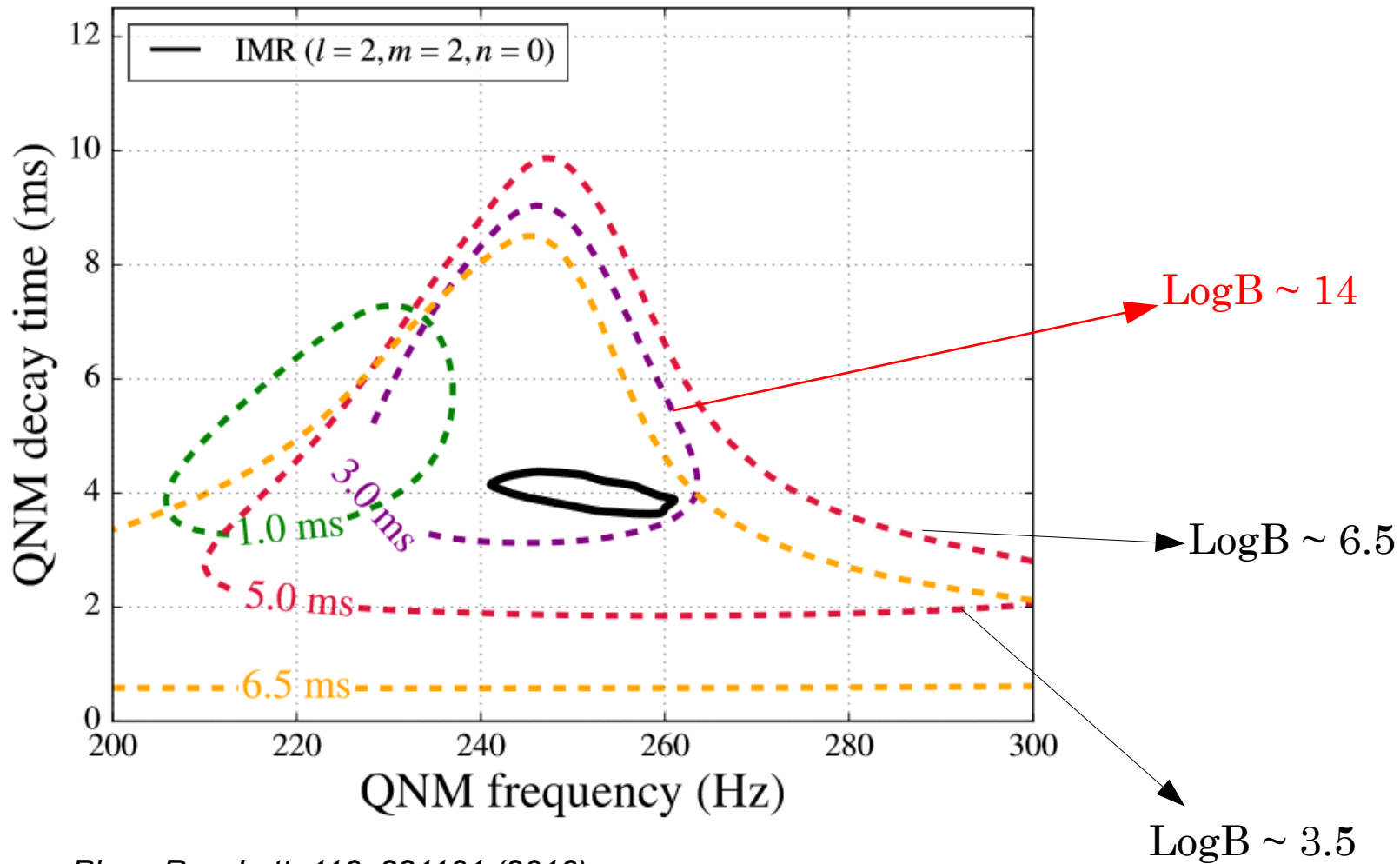
Starts to be consistent after
 $3\text{ms} \sim 10M$

Phys. Rev. Lett. 116, 221101 (2016)

Consistency tests

Ringdown test

GW150914



Phys. Rev. Lett. 116, 221101 (2016)

Production effects

Test **alternative models** which **change the waveform at its generation** (source)

Waveform expansion

- Consider an analytical, **parametric waveform**
- Treat the **GR predicted coefficients as free parameters**
- **Measure and compare** with GR predictions

Production effects

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Post Newtonian (PN) expansion of the inspiral phase + effective terms

- Expansion in $v/c \sim f^{1/3}$, **analytically known** up to $(v/c)^7$ for the **inspiral phase**
- **Effective waveform** for the **other stages**, calibrated on NR/EOB

$$\psi_{3.5}(f) = \sum_{i=0}^7 \varphi_i f^{(i-5)/3} + \varphi_{5l} \ln(f) + \varphi_{7l} f^{1/3} \ln(f) + e f f.$$

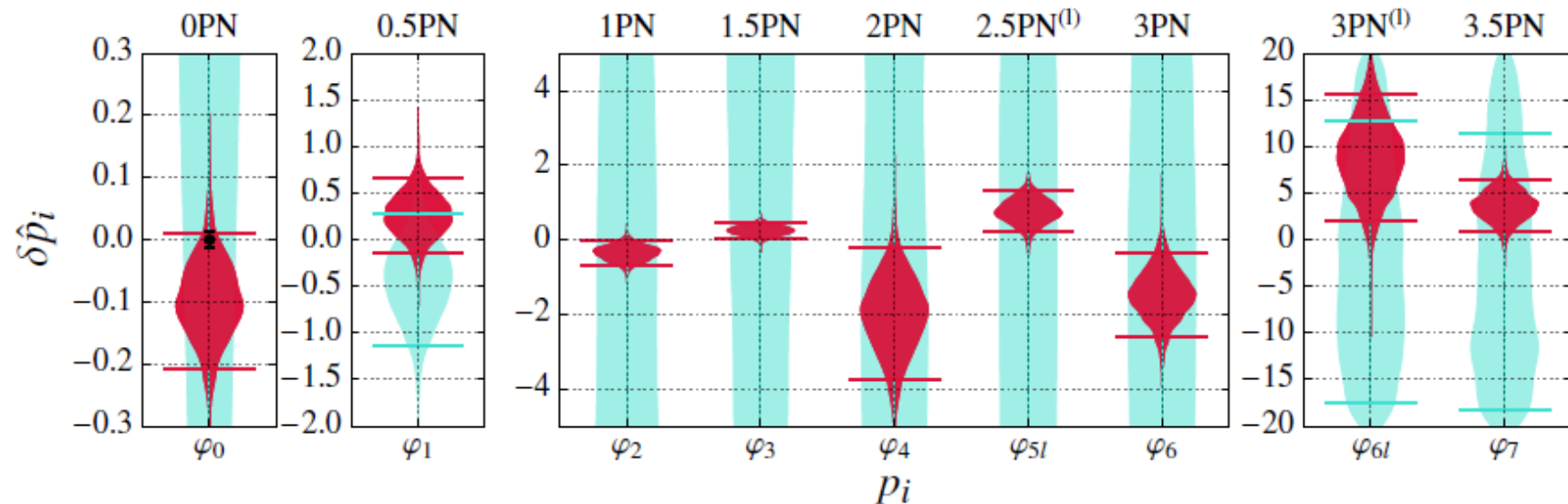
$$\varphi_i(m_1, m_2, a_1, a_2)$$

Production effects

Waveform expansion

- Add a **fractional deviation** and infer its value from **data**
- $\varphi_i \longrightarrow \varphi_i(1 + \delta\hat{\varphi}_i)$

GW150914

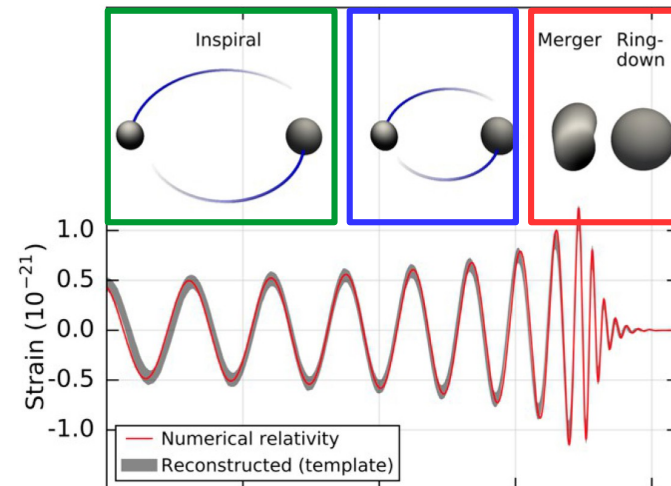
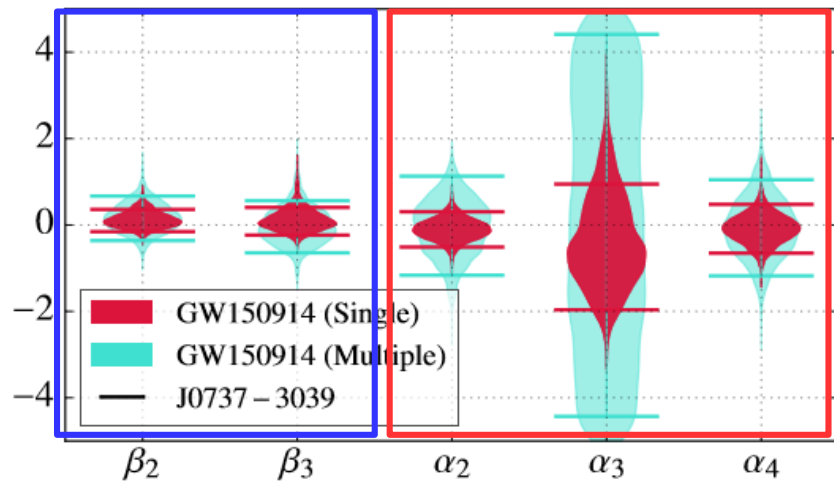
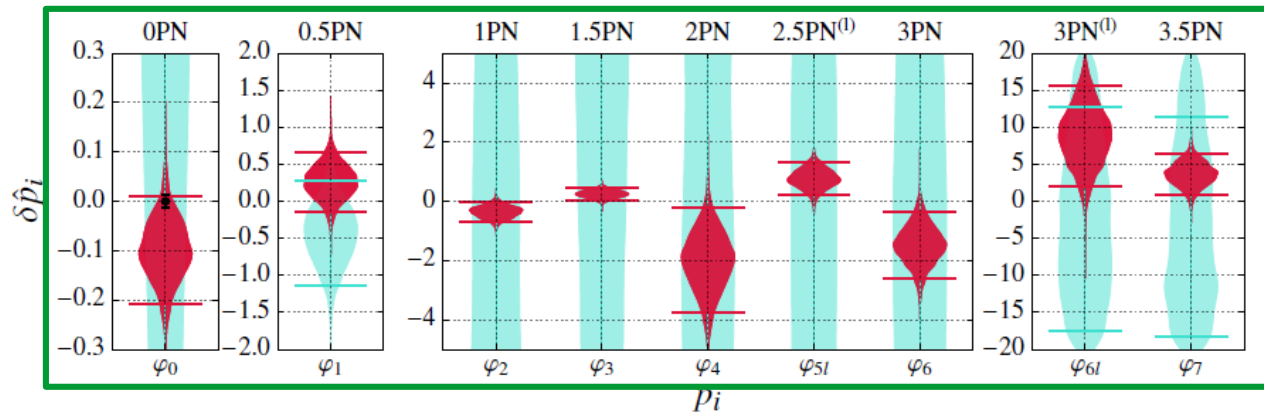


Phys. Rev. Lett. 116, 221101 (2016)

Production effects

Waveform expansion

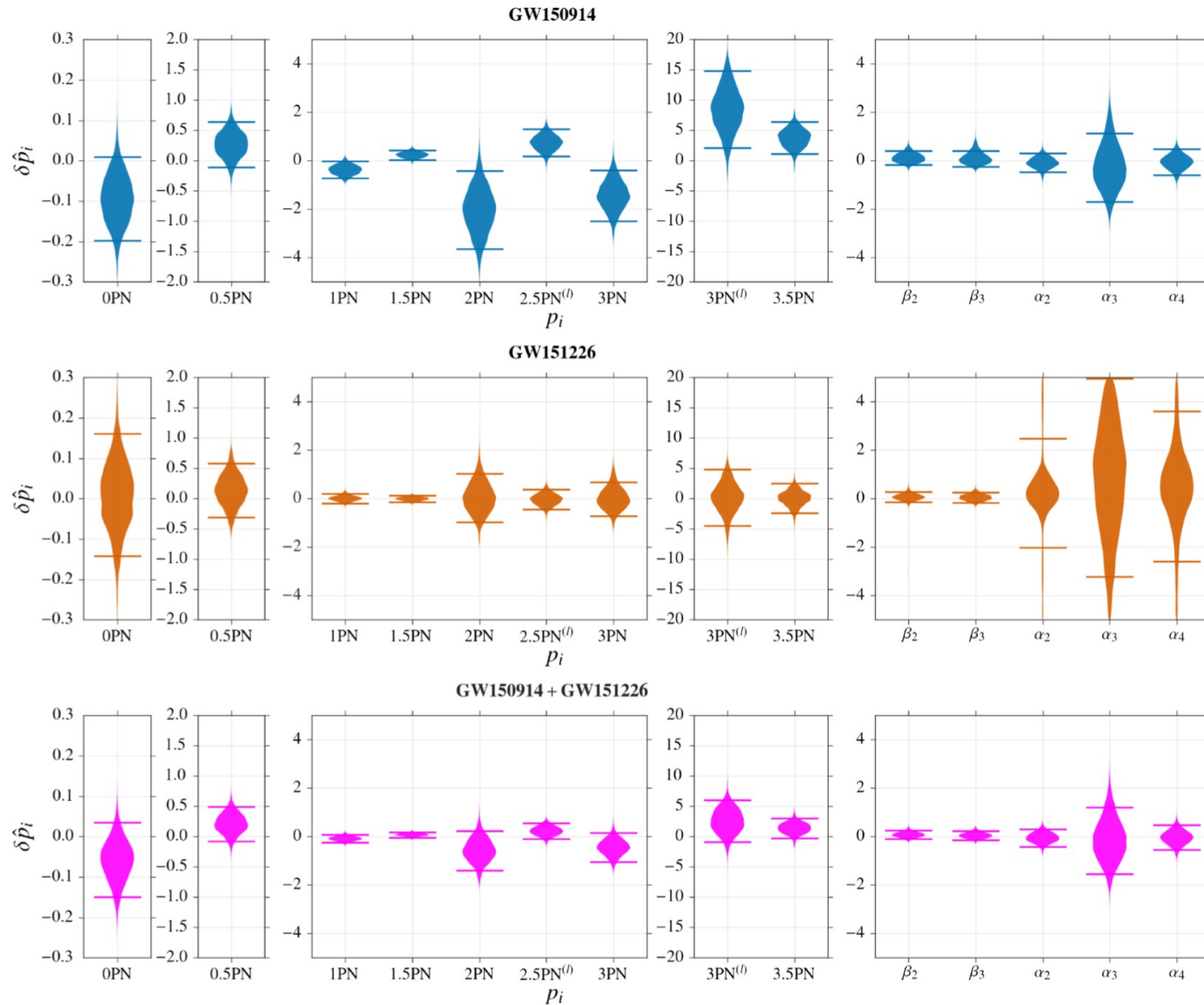
GW150914



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Production effects

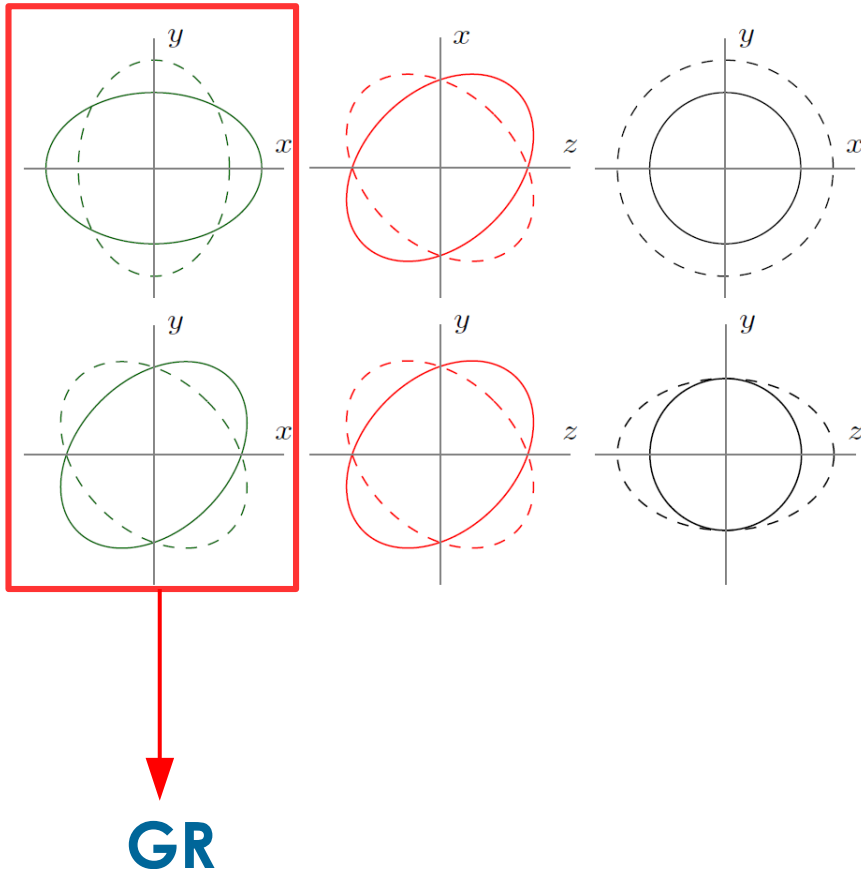
Waveform expansion



Production effects

Polarisation tests

Polarisations of a **generic theory** of gravity



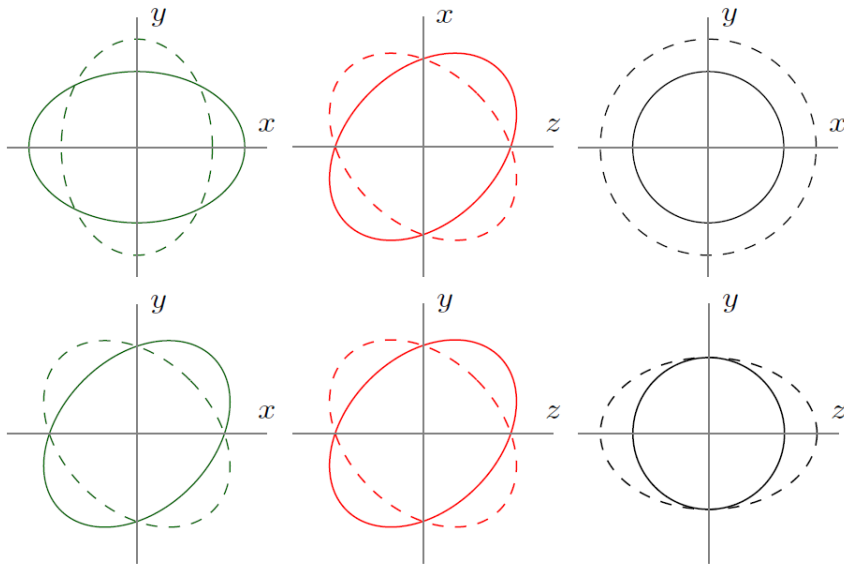
$$h(t, \vec{\theta}) = F^A(\hat{\Omega}) \cdot h_A(t, \vec{\theta})$$

- The polarisation content enters the **projection into the detector**
- Need **more than 3 detectors** to test all the combinations

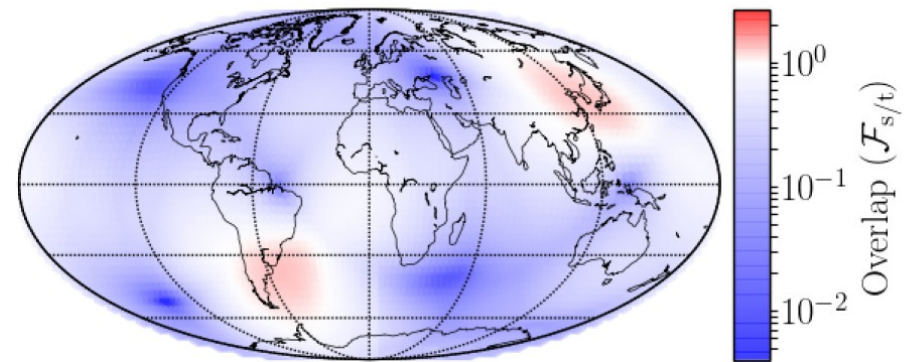
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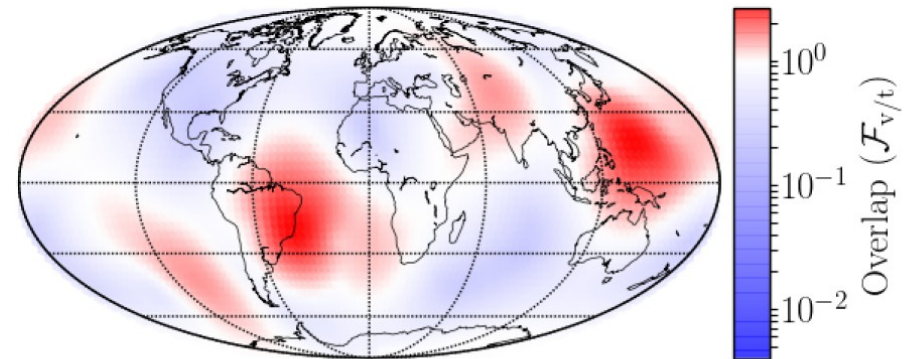
Polarisations of a **generic theory** of gravity



Sensitivity of the **LIGO/Virgo** network



(a) Scalar



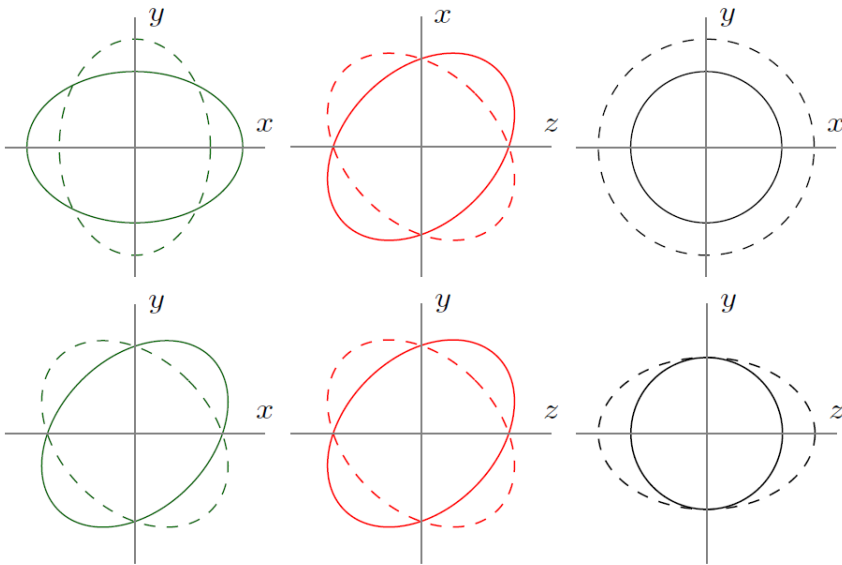
(b) Vector

Isi&Weinstein, arXiv:1710.03794v1

Production effects

Polarisation tests

Polarisations of a **generic theory** of gravity



	T/V	T/S
Bayes factor	200	1000

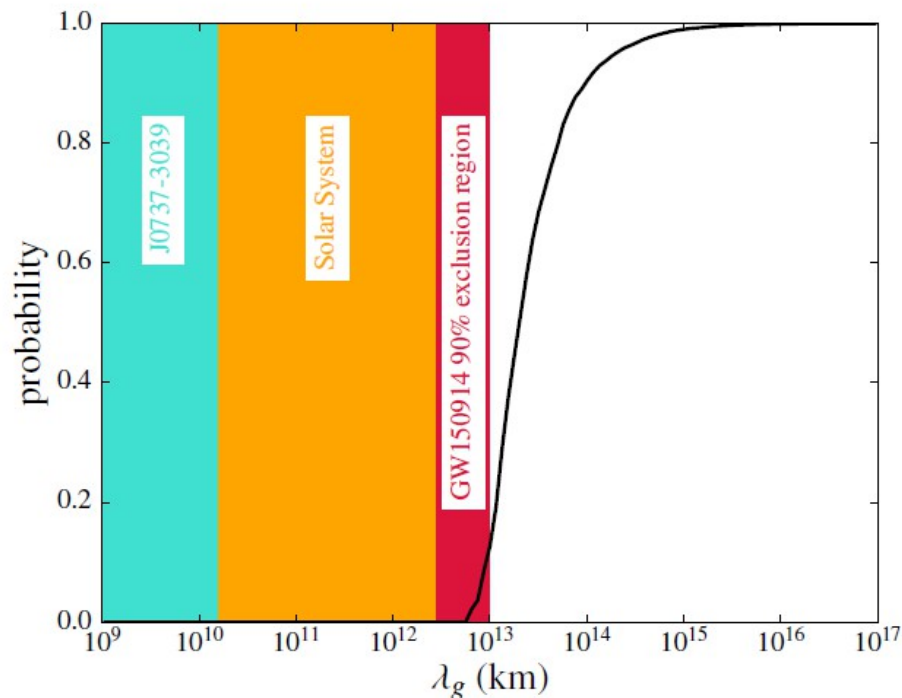
**NO vector/scalar
polarisations in
GW170814**

Propagation effects

Modified dispersion relation

- Consider a **massive graviton**: $E^2 = p^2 c^2 + m_g^2 c^4$ (Will, 2014)
- The **group velocity** gets modified by $v_g^2/c^2 = 1 - h^2 c^2 / (\lambda_g^2 E^2)$, $\lambda_g = h/(m_g c)$
- Which **adds a 1PN term** to the **phase** of the waveform: $\Phi_m = -(\pi D c) / [\lambda_g^2 (1+z) f]$

GW150914



Phys. Rev. Lett. 116, 221101 (2016)

GW150914

$$\lambda_g > 10^{13} \text{ km}$$

$$m_g < 1.2 \cdot 10^{-22} \text{ eV}/c^2$$

GW170817

$$-3 \times 10^{-15} \leq \frac{\Delta v}{v_{\text{EM}}} \leq +7 \times 10^{-16}$$

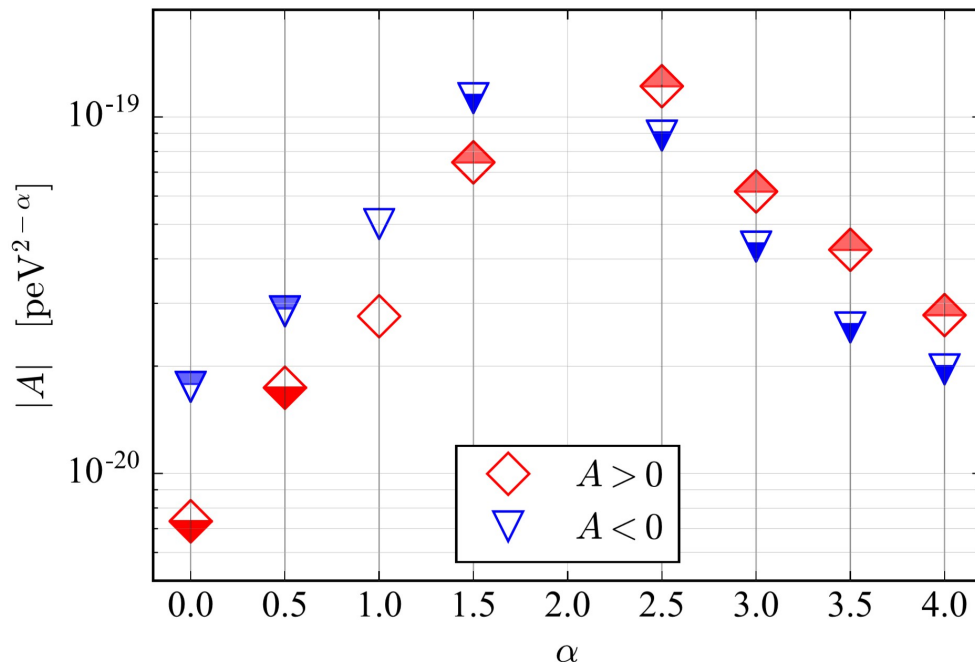
Propagation effects

More generally

- $E^2 = p^2 c^2 + A p^\alpha c^\alpha, \alpha \geq 0$
- $v_g/c \simeq 1 + (\alpha - 1) A E^{\alpha-2} / 2$

Massive graviton	Multifractal spacetime	Doubly special relativity	Extra dimensions
$\alpha = 0, A > 0$	$\alpha = 2.5$	$\alpha = 3$	$\alpha = 4$

GW150914 + GW151226 + GW170104



- **Weaker constraints** than **photon/neutrino** observations
- Some theories predict violations in just **one sector**
- **First bound** in the **gravitational sector**

Phys. Rev. Lett. 118, 221101 (2017)

Propagation effects

Higher dimensions

- A **leakage** of GWs in **higher dimensions** cause them to travel for a **greater distance** w.r.t. electromagnetic waves

$$h \propto \frac{1}{d_L^{\text{GW}}} = \frac{1}{d_L^{\text{EM}}} \left[1 + \left(\frac{d_L^{\text{EM}}}{R_c} \right)^n \right]^{-(D-4)/(2n)}$$

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Higher dimensions

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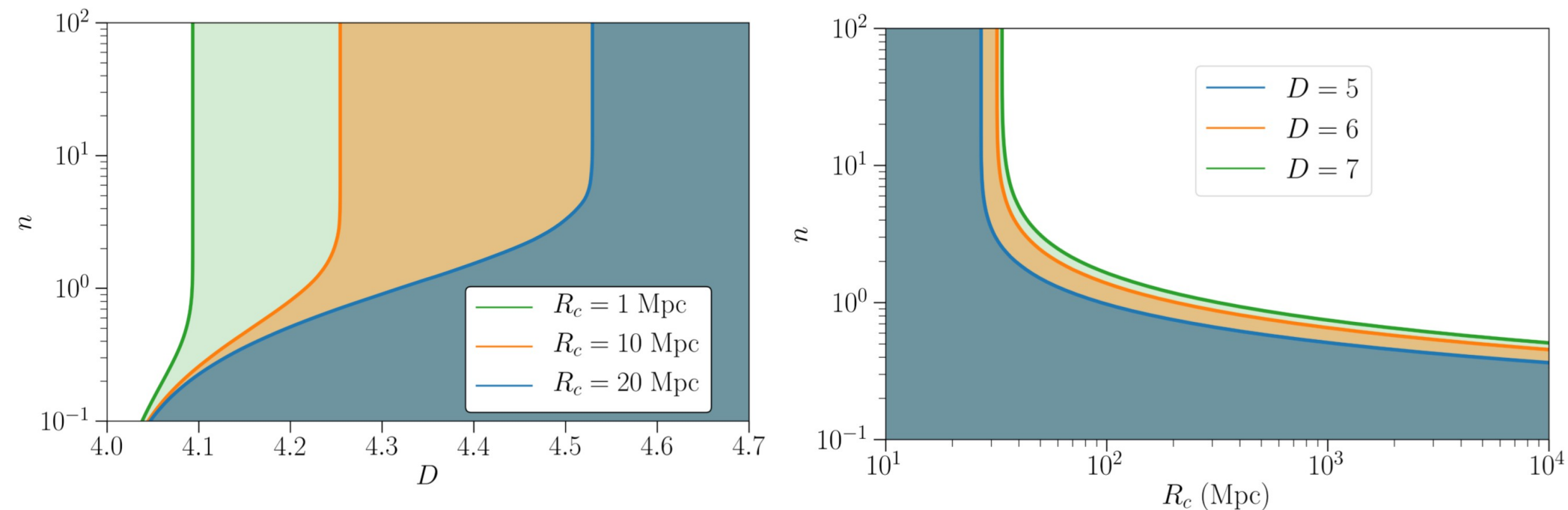
$$h \propto \frac{1}{d_L^{\text{GW}}} = \frac{1}{d_L^{\text{EM}}} \left[1 + \left(\frac{d_L^{\text{EM}}}{R_c} \right)^n \right]^{-(D-4)/(2n)}$$

Propagation effects

Higher dimensions

- A **leakage** of GWs in **higher dimensions** cause them to travel for a **greater distance** w.r.t. electromagnetic waves

GW170817: a binary neutron star signal



Abbott et al., arXiv 1811.00364v2

Conclusions and perspectives



Conclusions and perspectives

- General Relativity tested in the **strong field, non linear regime** thanks to **BBHs** and **BNSs**
- **Generation** and **consistency** of the signal **as expected** from Einstein's theory
 - The **power** accumulates as predicted (**residuals** compatible with **noise**)
 - **Consistency** of the whole **IMR** waveform
 - **Ringdown** regime **consistent** with GR for BBHs
 - Only **tensor degrees of freedom**
- **Propagation** of gravitational waves as expected. **For now** compatible with:
 - **Massless** graviton
 - **Light speed** propagation
 - **4 dimensions**

Conclusions and perspectives

....as many more detections are expected in the upcoming months

- As **statistics** accumulates, **posteriors** will be **combined** and **higher precision** will be reached in the tests presented
- **Higher SNR** signals will allow for more **refined tests** (multimodal ringdown tests, spin interactions, polarisations etc.)
- ?

Far future

- Possible detection of a **stochastic background**
 - The **cosmological background** is a window on **Planck scale energies**
and on the **early universe**, where an **extended theory of gravity** could come into play