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## Radio morphology-accretion mode link in FR II low-excitation radio galaxies

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Radio Galaxies (RG) are Radio-Loud Active Galactic Nuclei (AGN) characterized by powerful relativistic jets, oriented at relatively large inclination angles with respect to the observer's line of sight. They are extraordinarily relevant to address important issues such as the interaction between Super Massive Black Holes, the radio jets and their environment.

RGs are commonly classified on the basis of their radio morphology into low radio power FRI (core-dominated) and high radio power FR II (lobe-dominated). Almost all FR IIs hide an efficient accretion disk, while the central engine of FR Is is probably an ADAF, suggesting a strong link between ejection of relativistic plasma and accretion. However, there is a group of FR II sources that does not fit into this picture. They exhibit powerful extended radio structures but inefficient accretion, as attested by their low excitation optical spectra.

In order to investigate their nature, an X-ray systematic analysis of all FR IIs belonging to the 3CR sample with  $z < 0.3$  and available Chandra and XMM-Newton observations was performed. We investigate different scenarios and conclude that the evolutionary one is the most probable. FR II-LERGs have indeed intermediate properties (X-ray luminosity, column density and accretion rate in terms of Eddington luminosity) with respect to classical FR IIs and FR Is.

The nuclear power seems to suffer of a depletion of the cold gas reservoir. It is then plausible to think that this information has not yet reached the large-scales radio structures at kpc distances from the central engine. The powerful lobes are the heritage of a past AGN activity at higher efficiency.

**Primary author:** MACCONI, Duccio (OAS-Bologna, DIFA Bologna)

**Co-authors:** TORRESI, Eleonora (INAF-OASBO); GRANDI, Paola (INAF-OASBO); BOCCARDI, Bia (Max Planck Institute for Radio Astronomy, Bonn); VIGNALI, Cristian (DIFA Bologna)

**Presenter:** MACCONI, Duccio (OAS-Bologna, DIFA Bologna)

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