

CUORE

O.Cremonesi
LNF - 10/12/2018

Outline

- Neutrinoless Double Beta Decay
- CUORE
 - Construction
 - Operation
 - Results
- Future perspectives

NEUTRINO-LESS DOUBLE BETA DECAY

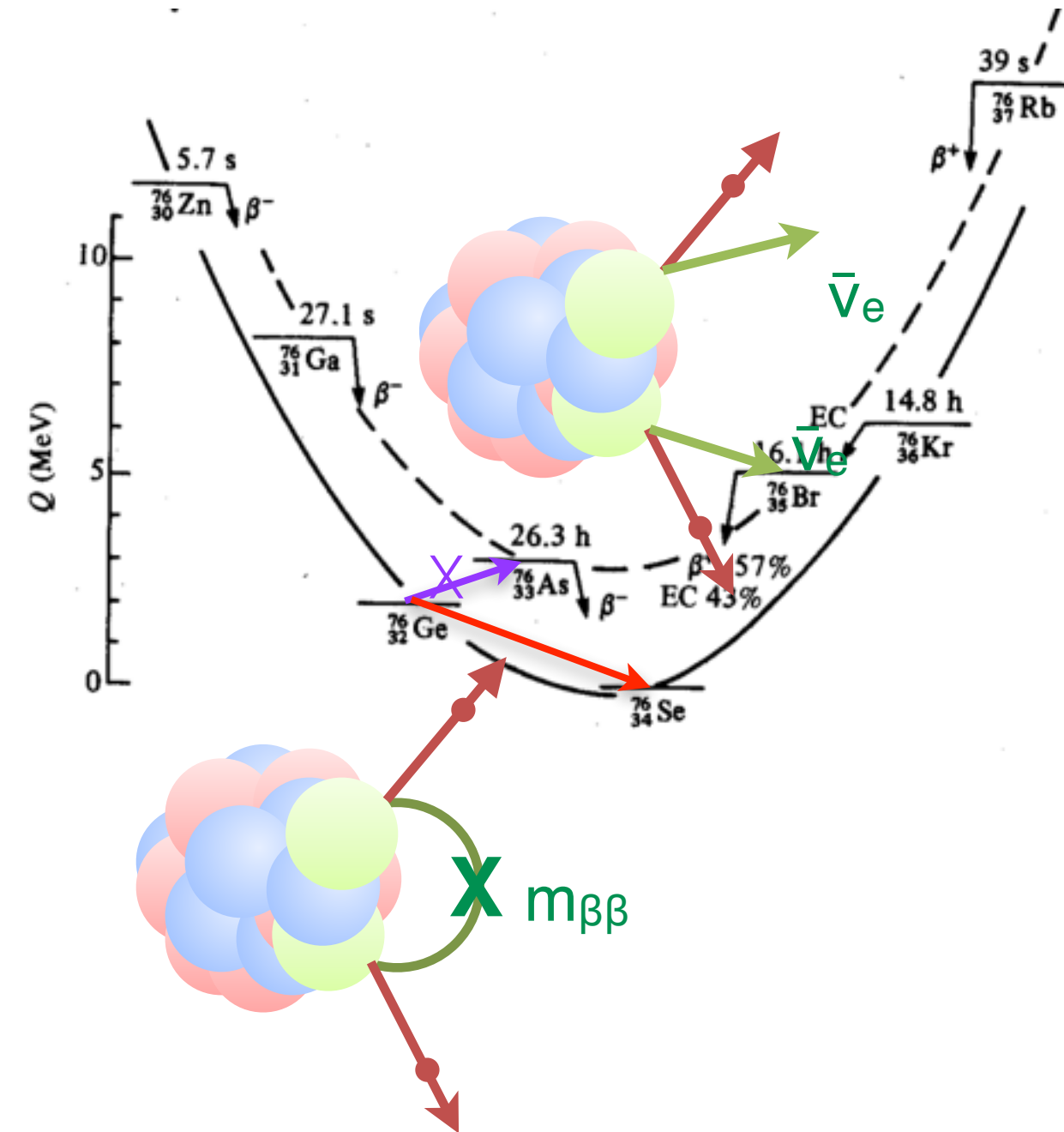
$\beta\beta$ decay

Rare nuclear decay between isobar nuclei with $|\Delta Q|=2$

Even-even nuclei: **ONLY** direct $0^+ \rightarrow 0^+$ transition

Different modes:

- $(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\bar{\nu}$ ($2\nu\beta\beta$)
 - Standard 2nd order weak nuclear transition
 - Long lifetime (10^{18-24} yr)
- **$(A,Z) \rightarrow (A,Z+2) + 2e^-$ ($0\nu\beta\beta$)**
 - **Lepton number violation (LNV): physics beyond the SM**
 - Only possible if neutrinos are Majorana fermions
 - LNV ($\Delta L=2$)
 - absolute neutrino mass scale
 - Majorana phases
- SM extensions
- $(A,Z) \rightarrow (A,Z+2) + n\chi$ (exotic modes)



$0\nu\beta\beta$ decay width

Under non trivial approximations it is possible to separate **atomic**, **nuclear** and **particle** contributions and factoring the transition width as

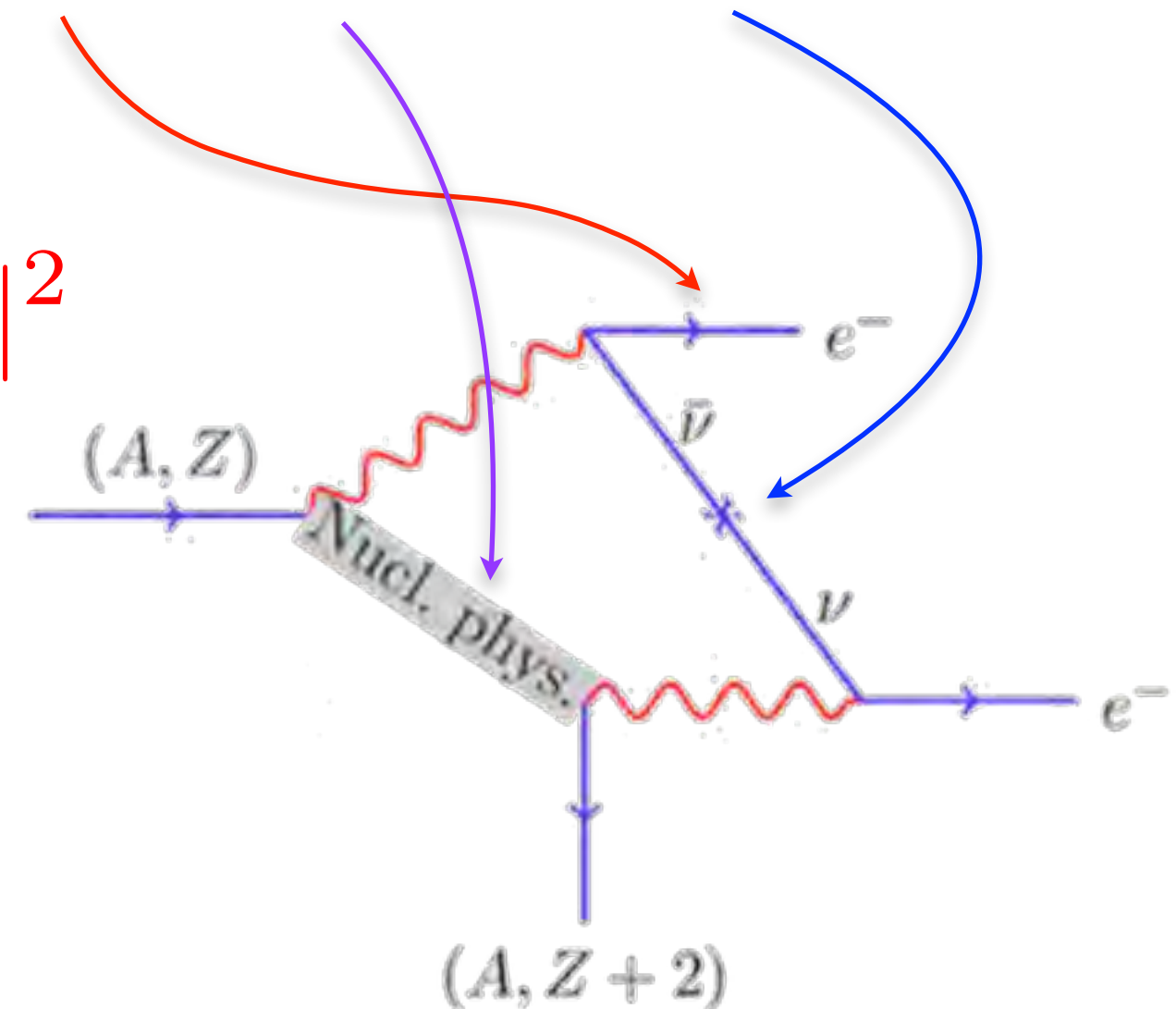
$$\Gamma^{0\nu} = G_x(Q, Z) |M_x(A, Z)|^2 |\eta_x|^2$$

$G_x(Q, Z)$ = phase space factor \rightarrow precisely calculable

$M_x(A, Z)$ = nuclear matrix element \rightarrow problematic

η_x = particle physics parameter \rightarrow model dependent

- massive Majorana neutrinos
- GUT's
- SUSY
- ...

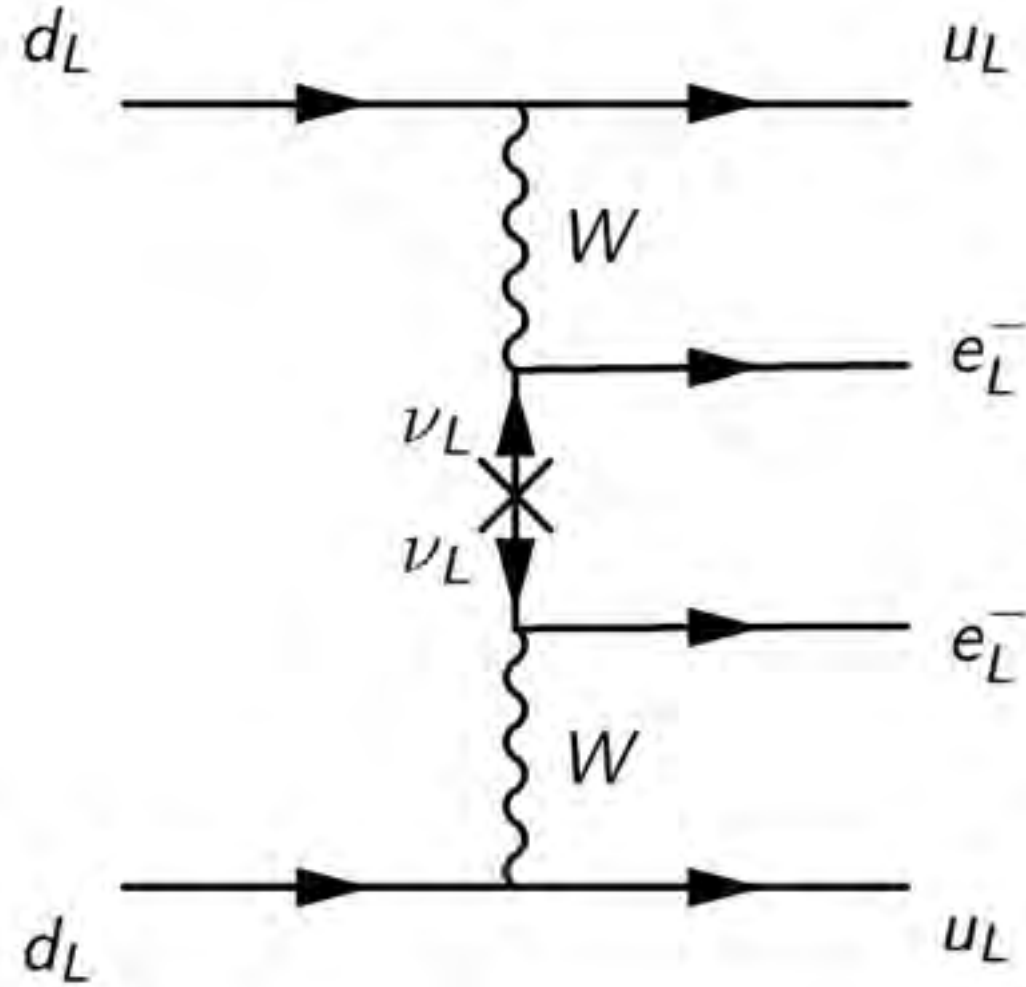


Lepton number violation (LNV) and $0\nu\beta\beta$

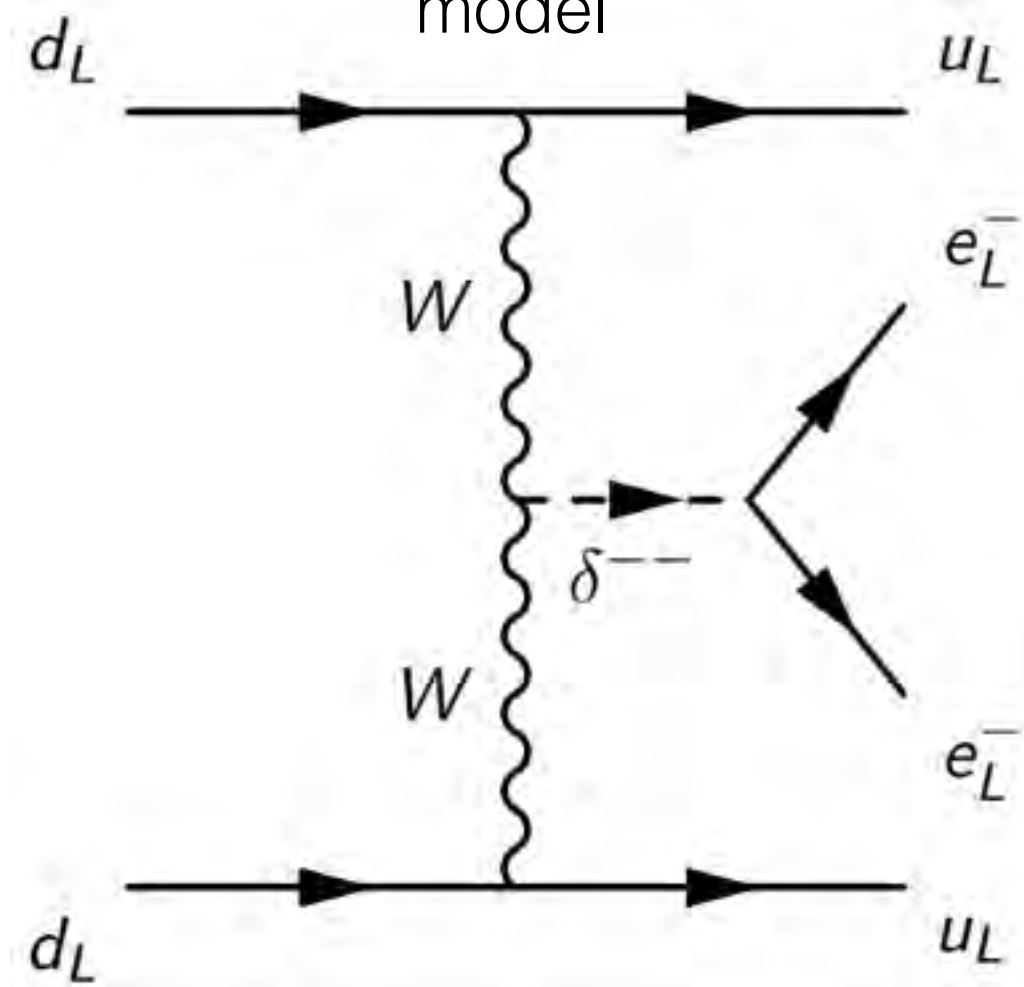
$$A_{\beta\beta} \sim m_{ee}$$

$$A_{\beta\beta} \sim \text{LNV parameters}$$

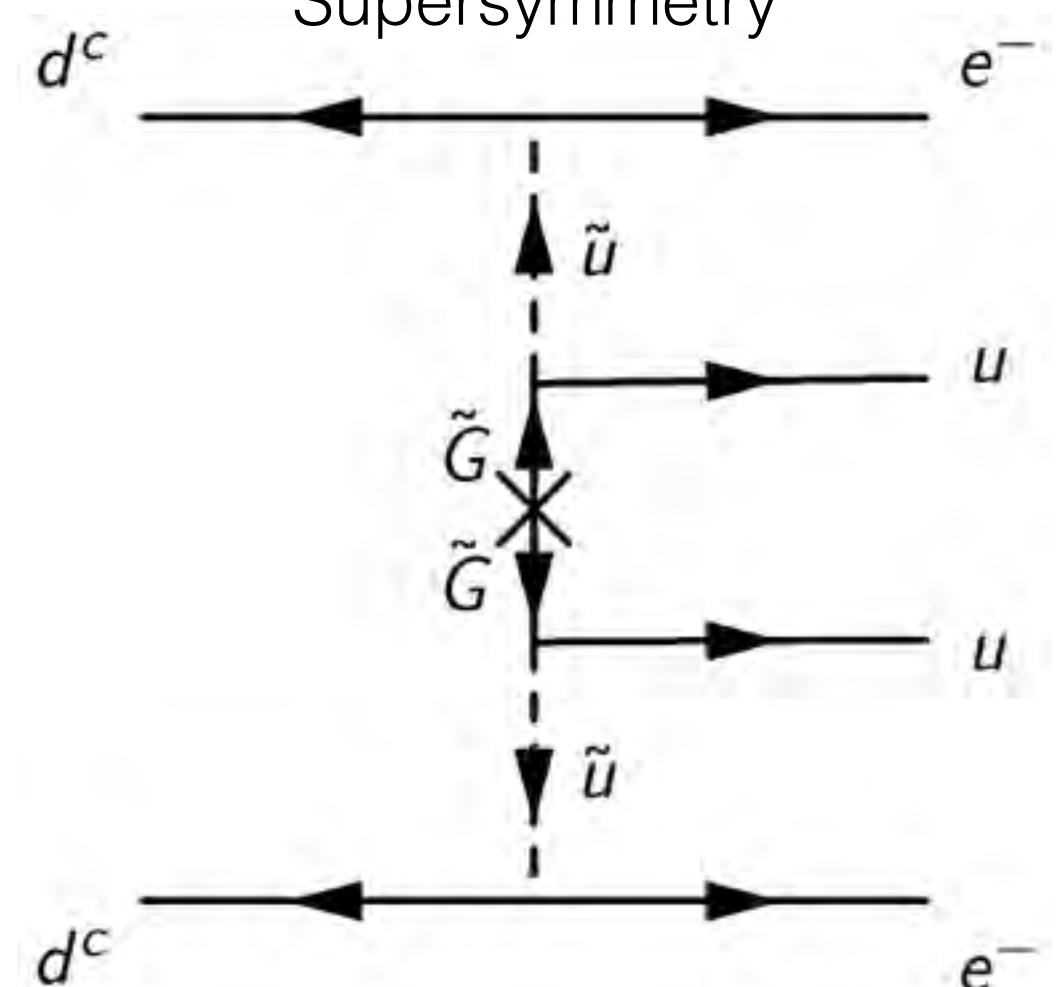
Mass mechanism



Higgs triplet model



Supersymmetry

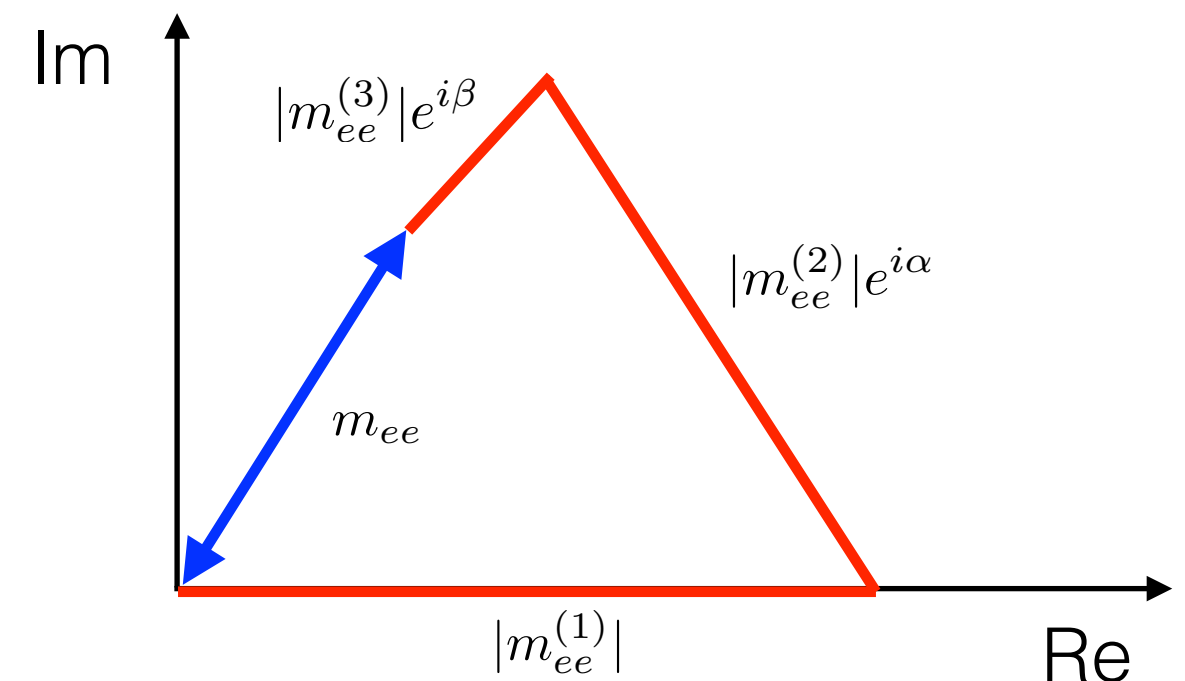


$\beta\beta 0\nu$ standard interpretation: light Majorana neutrino

- Neutrinoless Double Beta Decay is mediated by light massive Majorana neutrinos and all other potential mechanisms give negligible or no contribution

$$\begin{aligned}\eta_x = \langle m_{ee} \rangle &= \sum_k U_{ek}^2 m_k \\ &= c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 e^{i\alpha} m_2 + s_{13}^2 e^{i\beta} m_3\end{aligned}$$

- The transition amplitude is proportional to coherent sum of neutrino masses
- Majorana phases play a crucial role: possible cancellations

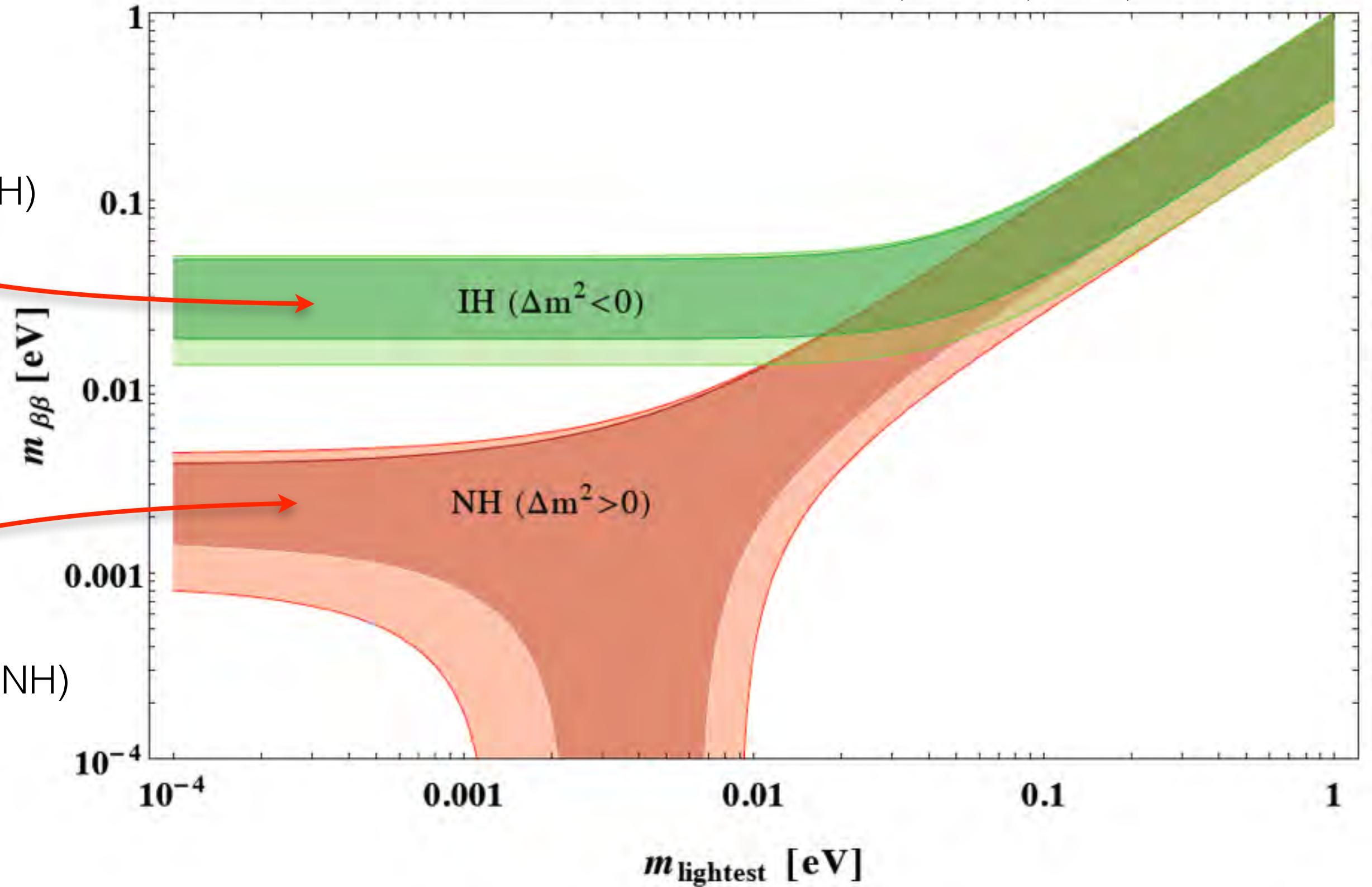


Light Majorana neutrino

Dell'Oro, Marcocci, Vissani, PRD 90.033005

Inverse Ordering (IO or IH)

Normal Ordering (NO or NH)



Distinguishing $0\nu\beta\beta$ mechanisms

A single measurement of total rate cannot pin down underlying $0\nu\beta\beta$ mechanisms.

- Half life ratios of different isotopes

Deppisch, Päs 06 , Gehman, Elliot 07 , Fogli, Lisi, Rotunno 09

- Electron angular correlations

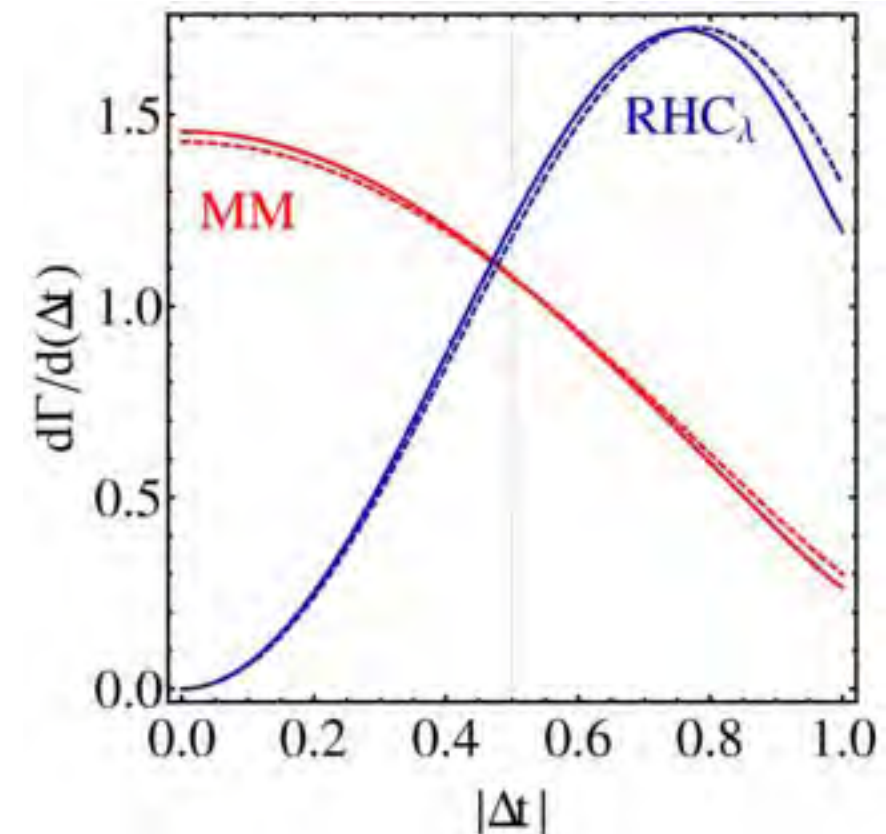
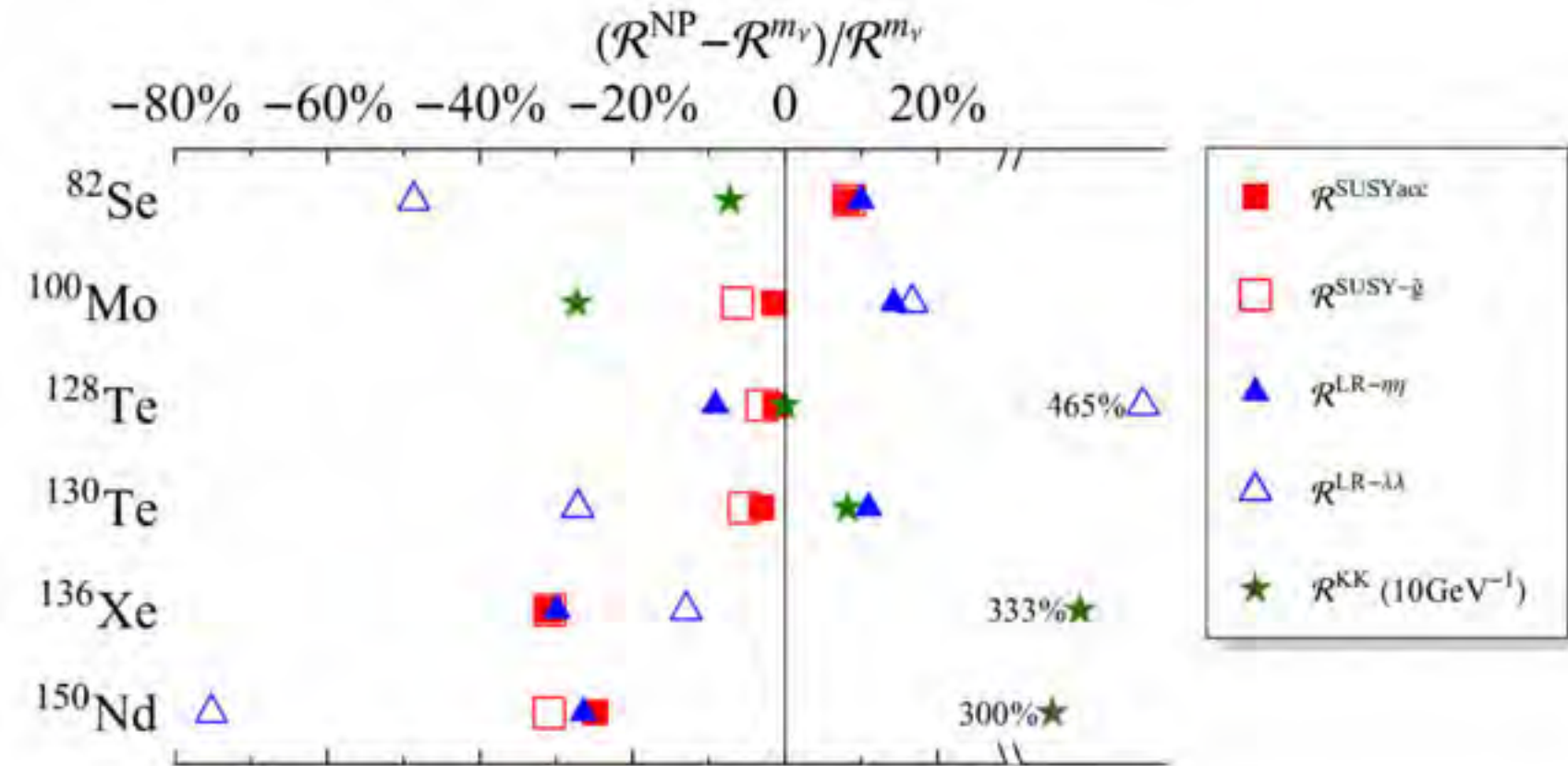
Ali, Borisov, Zhuridov 07

- Decay to excited states and other rare decays

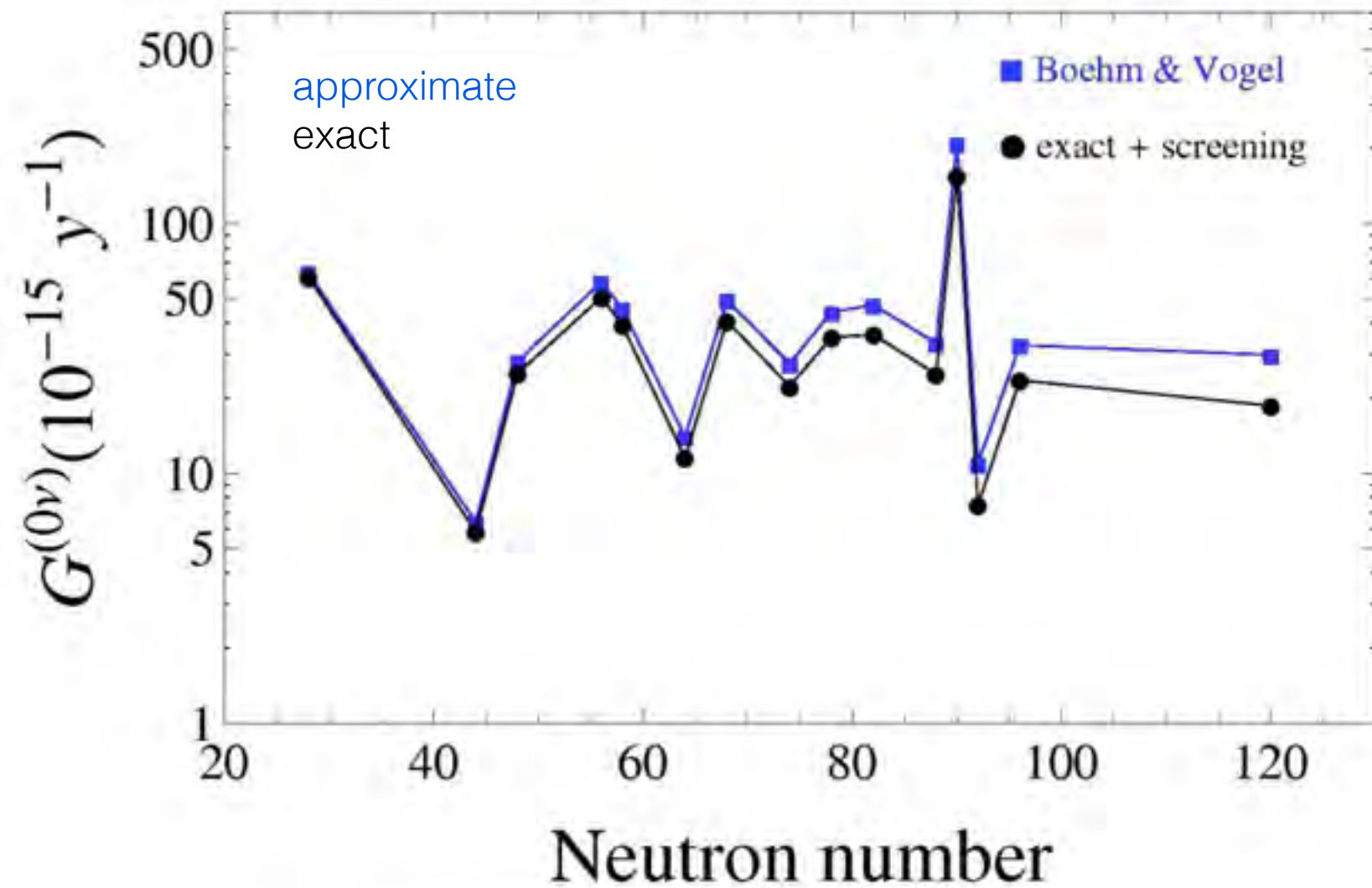
Faessler et.al. 94

- LNV processes at the LHC

Allanach, CHK, Päs 09



Extract m_{ee} from data: a hard job

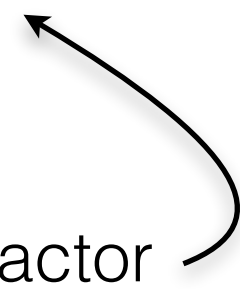


F. Böhm and P. Vogel, *loc. cit.*

J. Kotila and F. Iachello, Phys. Rev. C 85, 034316 (2012)

$$\Gamma^{0\nu} = G_{0\nu}(Q, Z) |M_{0\nu}(A, Z)|^2 |m_{ee}|^2$$

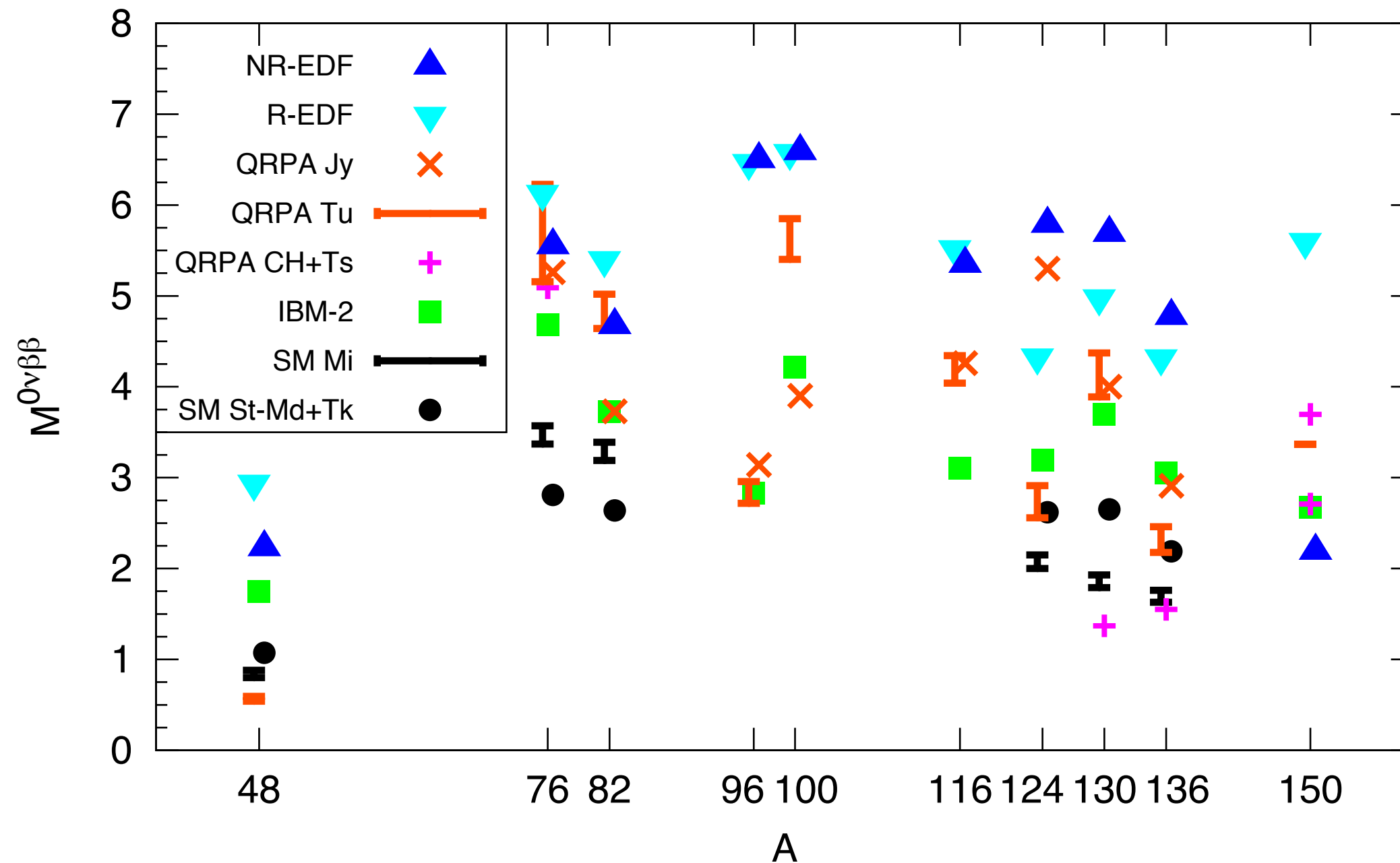
The easy part:
phase space factor
→ precisely calculable



Extract m_{ee} from data: a hard job

Full estimated range of $M^{0\nu}$ within different calculation methods

$$\Gamma^{0\nu} = G_{0\nu}(Q, Z) |M_{0\nu}(A, Z)|^2 |m_{ee}|^2$$



Much more problematic:
nuclear matrix elements

- Significant spread
- Absolute values inaccuracy

quenching of g_A ?

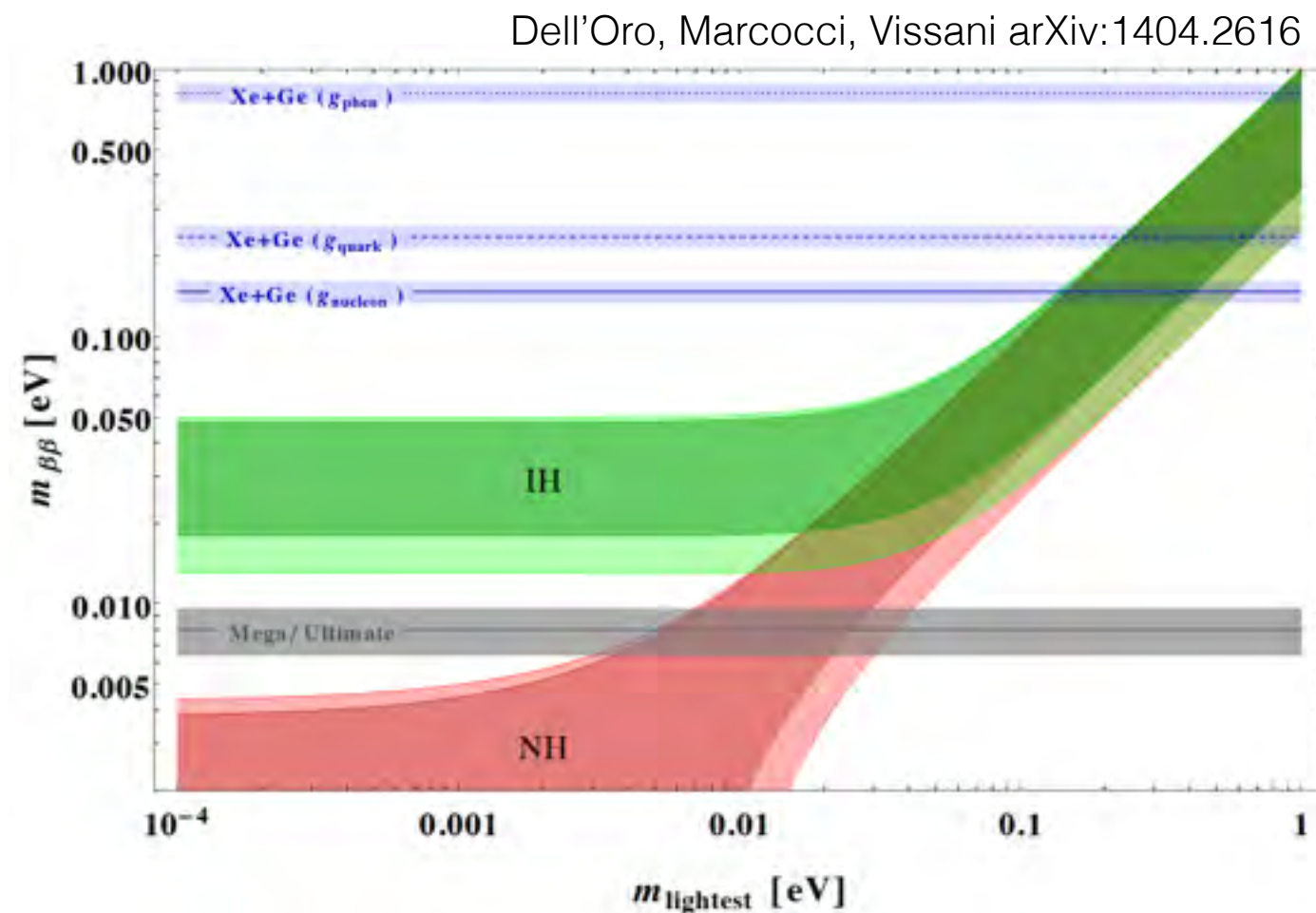
$$\Gamma_{\beta\beta} \sim g_A^4$$

g_A quenching

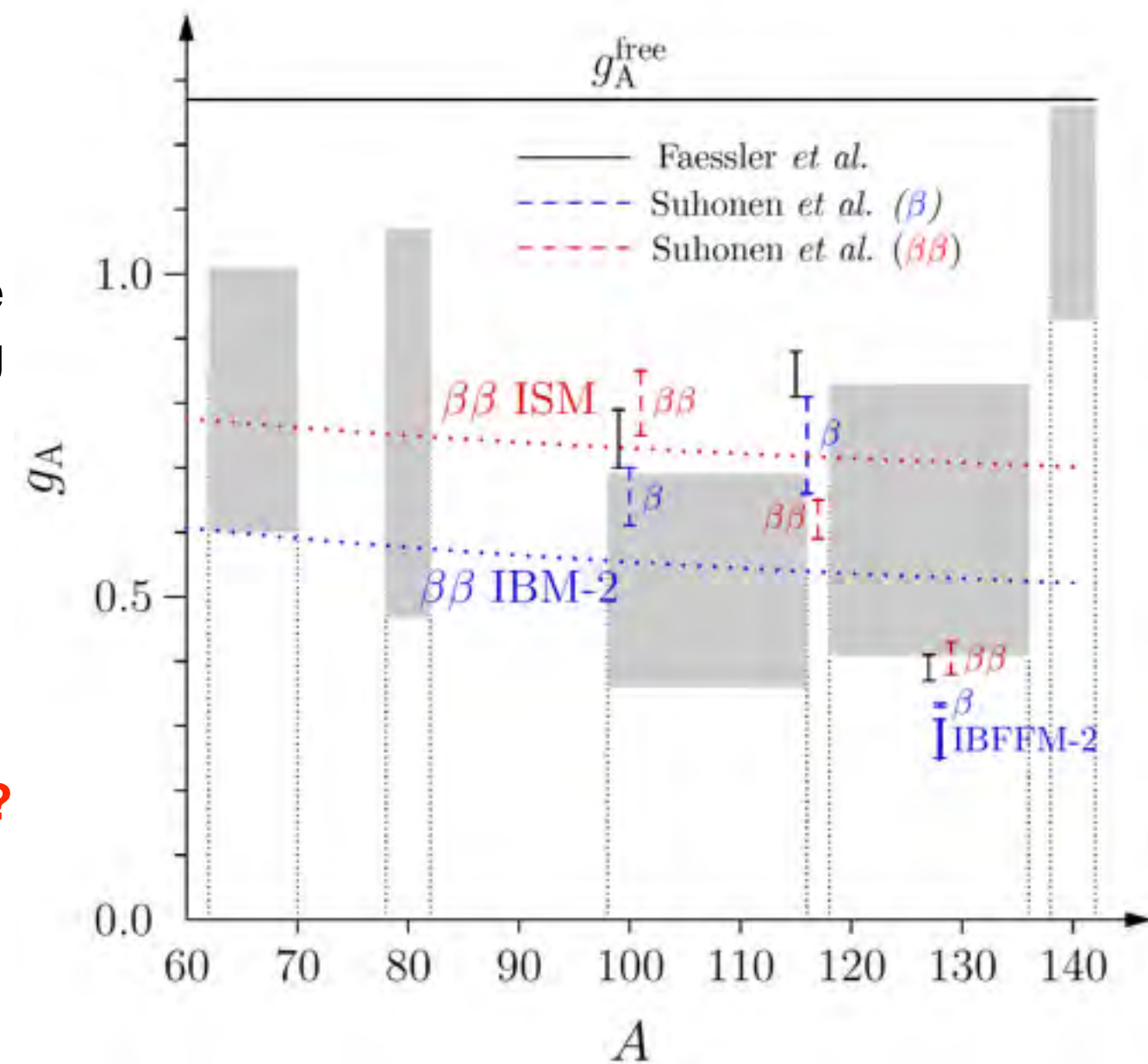
Allowed “beta” decays

- Recent analysis of pn-QRPA, ISM and IBM based calculations of single β and $2\nu\beta\beta$ decays shows an evident dependence of the g_A quenching on A (J.Suhonen, Neutrino 2018):

Mass range	76-82	100-116	122-136
g_A	0.7-0.9	0.5	0.5-0.7



What about $0\nu\beta\beta$?



Forbidden “beta” decays

- Possible extrapolation to $0\nu\beta\beta$ (large momentum transfer)
- Crucial role of high forbiddance non unique transitions

^{94}Nb , ^{98}Tc , ^{99}Tc , ^{113}Cd , ^{115}In , ^{138}Cs

Caveat: $0\nu\beta\beta$ decay is a high-momentum transfer process ($q \sim 100$ MeV) \rightarrow less quenching

Experimental signature



- A new (ionised) isotope
- Two electrons

Minimal information:

- two e^- energy sum spectrum (total deposited energy)

$0\nu\beta\beta$ exhibits a peak at Q over $2\nu\beta\beta$ tail (and background contributions)

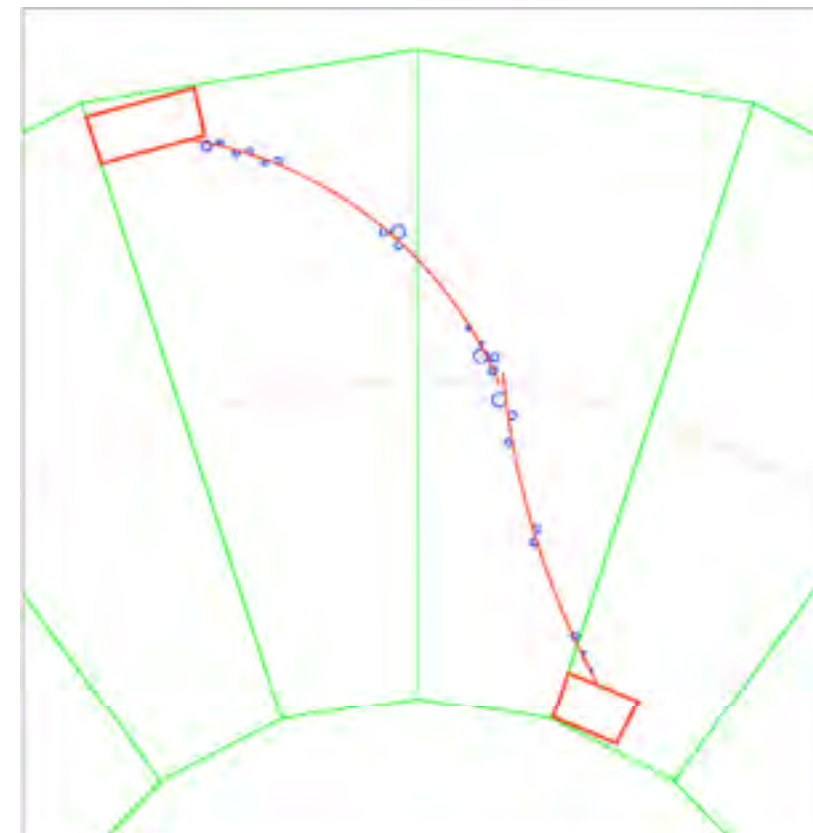
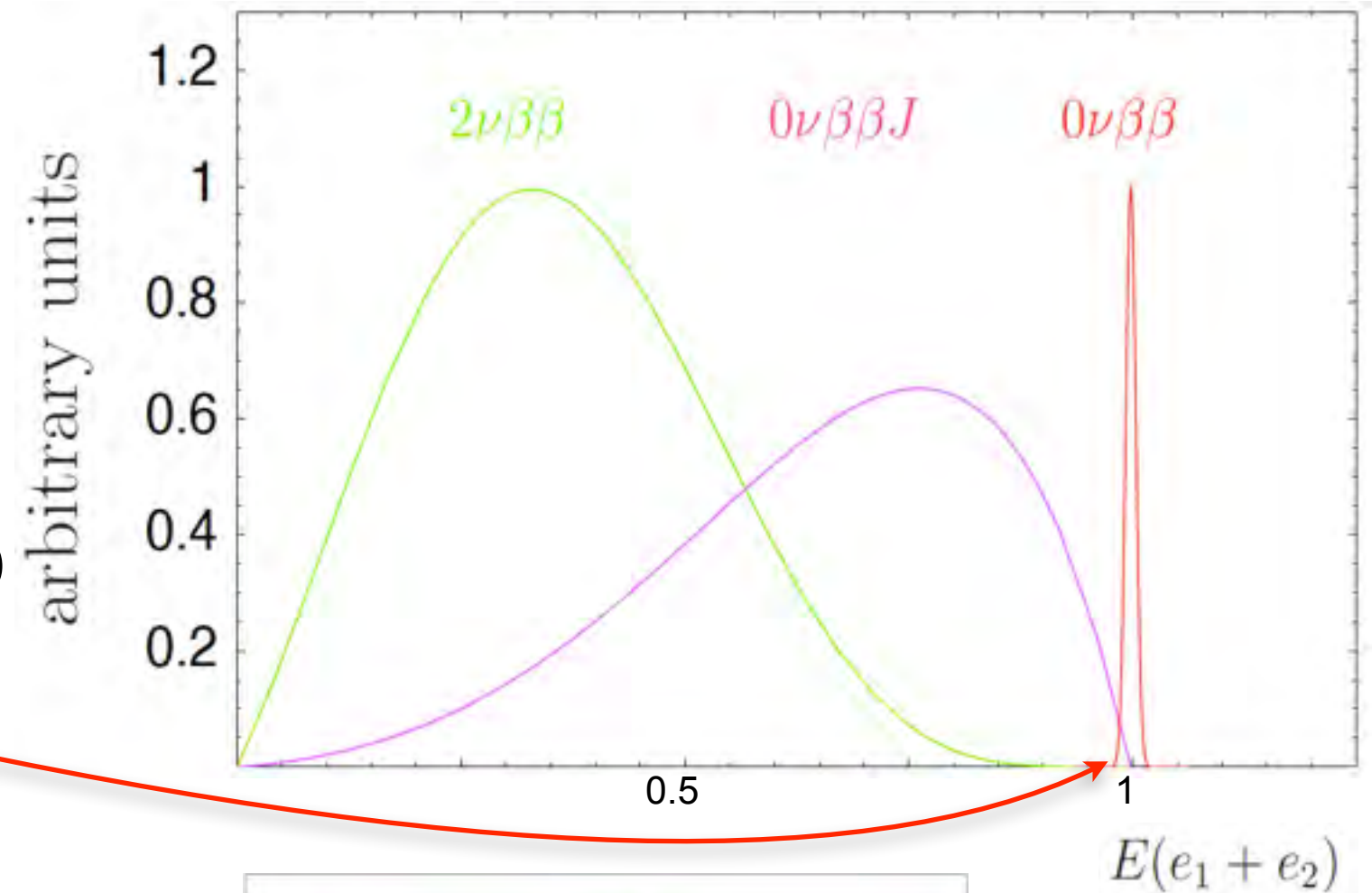
Additional signatures:

- Daughter nuclear species
- Single electron energy spectrum
- Angular correlation between the two electrons
 - Track and event topology
 - Time Of Flight

Two experimental approaches:

- homogeneous (source \equiv detector) → calorimeters
- inhomogeneous (source \neq detector) → trackers

... and mixed solutions



Experimental sensitivity

$$\tau_{1/2}^{0\nu} = \ln 2 \frac{\epsilon N_{nuclei} t_{meas}}{N_{\beta\beta}}$$

Lifetime corresponding to the minimum detectable number of events over background at a given confidence level

$$N_{\beta\beta} \leq \sqrt{bkg \cdot \Delta E \cdot M \cdot t_{meas}}$$

N_{nuclei}	number of active nuclei in the experiment
t_{meas}	measuring time [y]
M	detector mass [kg]
ϵ	detector efficiency
$i.a.$	isotopic abundance
A	atomic number
ΔE	energy resolution [keV]
bkg	background [c/keV/y/kg]

$N_B = bkg \cdot \Delta E \cdot T \cdot M$ number of background events expected along the experiment lifetime

$N_B \gg 1$

$$S_{1/2}^{0\nu} \propto \epsilon \frac{i.a.}{A} \sqrt{\frac{M \cdot t_{meas}}{bkg \cdot \Delta E}}$$

$N_B \leq O(1) \rightarrow$ "zero background"

$$S_{1/2}^{0\nu} \propto \epsilon \frac{i.a.}{A} M \cdot t_{meas}$$

- **Isotopical abundance**
- **Mass**
- **Energy resolution**
- **Background level**

Performance

Scale

$$N_B = S \cdot P \equiv 1$$

$$S_{FB} \sim \sqrt{\frac{S}{P}}$$

$$S_{ZB} \sim S$$

$$\frac{1}{S_{1/2}^{0\nu}(m_{ee})} \propto \sqrt{S_{1/2}^{0\nu} \cdot G^{0\nu} |M^{0\nu}|}$$

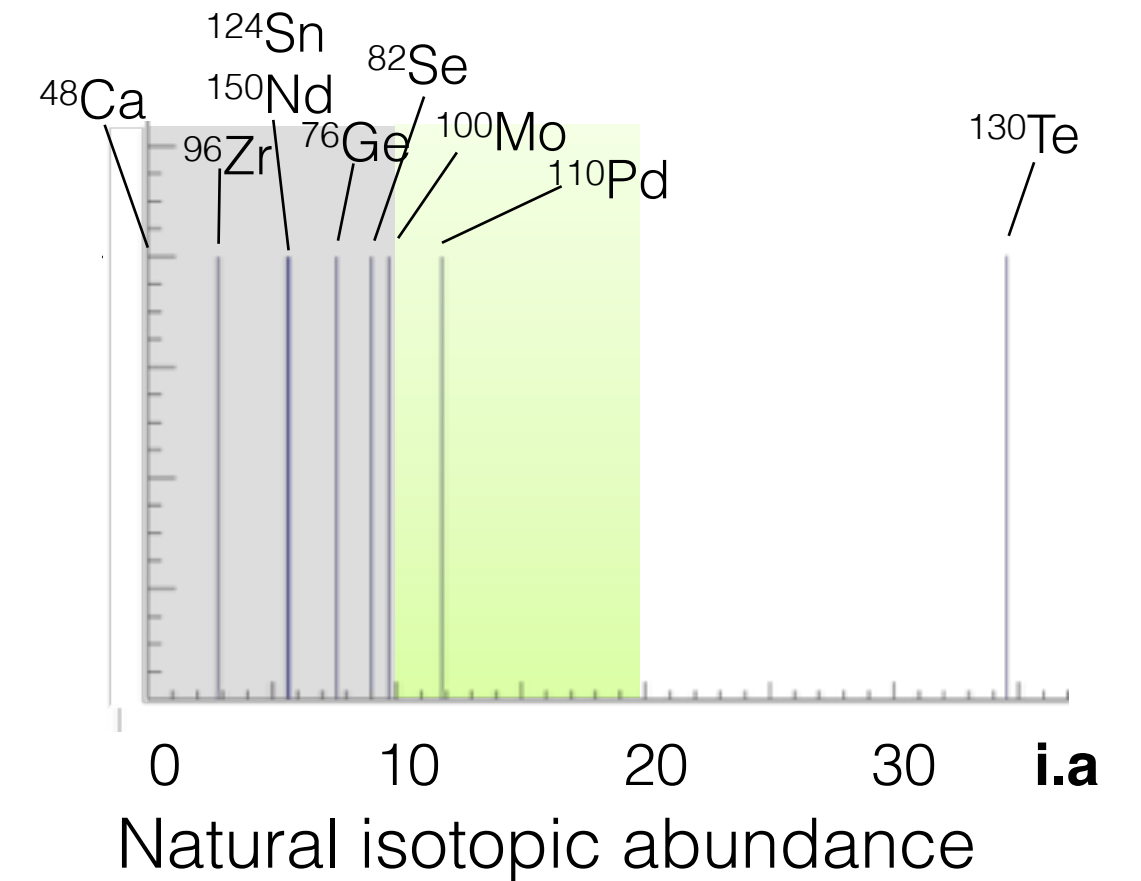
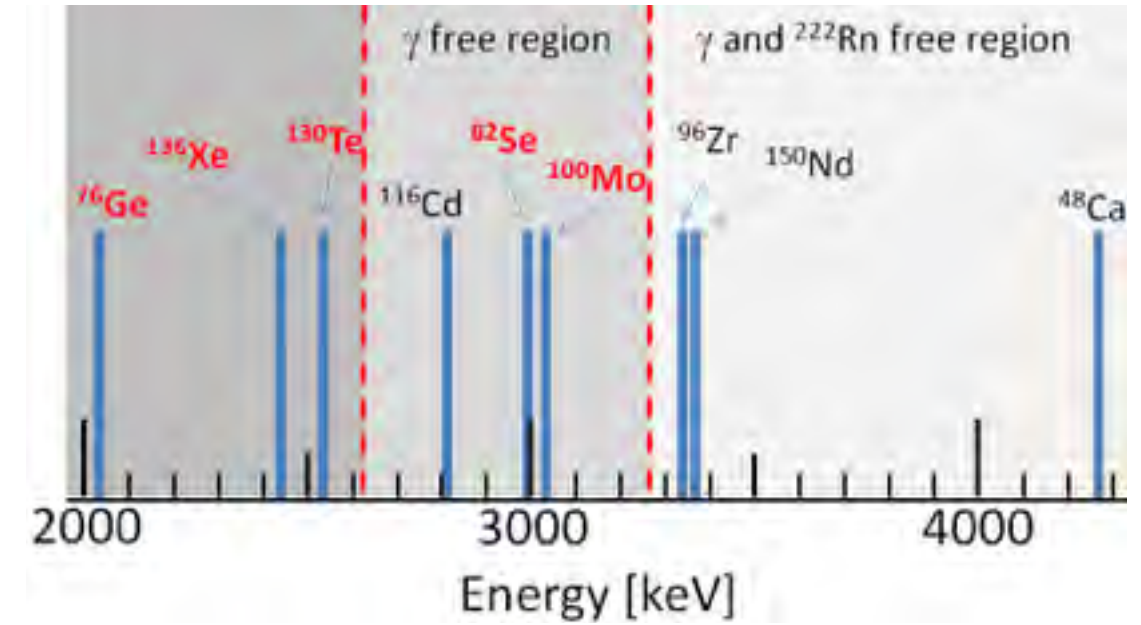
◦ **Isotope choice**

Choice of the isotope

Transition	T0 (keV)	Abundance (%)	first 2+ (keV)
46Ca→46Ti	985	0.0035	889
48Ca→48Tit	4272	0.187	984
70Zn→70Ge	1001	0.62	-
76Ge→76Se	2045	7.8	559
80Se→80Kr	136	49.8	-
82Se→82Kr	3005	9.2	776
86Kr→86Sr	1249	17.3	1077
94Zr→94Mo	1148	17.4	871
96Zr→96Mo	3350	2.8	778
98Mr→98Ru	111	24.1	-
100Mo→100Ru	3033	9.6	540
104Ru→104pd	1301	18.7	556
110Pd→110Cd	2014	11.8	658
114Cd→114Sn	540	28.7	-
116Cd→116Sn	2808	7.5	1294
122Sn→122Te	358	4.56	-
124Sn→124Te	2278	5.64	603
128Te→128Xe	869	31.7	443
130Te→130Xe	2533	34.5	536

Transition	T0 (keV)	Abundance (%)	first 2+ (keV)
134Xe→134Ba	843	10.4.	605
136Xe→136Ba	2481	8.9	819
142Ce→142Nd	1414	11.1	-
146Nd→146Sm	61	17.2	-
148Nd→148Sm	1928	5.7	550
150Nd→150Sm	3367	5.6	334
154Sm→154Gd	1250	22.6	123
160Gd→160Dy	1731	21.8	87
170Er→170Yb	655	14.9	84
176Yb→176Hf	1077	12.6	88
186W→186Os	489	28.6	137
192Os→192Pt	408	41.0	317
198Pt→198Hg	1043	7.2	412
204Hg→204Pb	414	6.9	-
232Th→232U	850	100	48
232U→232Pu	1146	99.275	44

Transition energy



DBD experiments summary

Experiment	Isotope	$m_{\text{fid}}(\beta\beta)$ [kg]	Technique	Laboratory	Status
CANDLES	48Ca	305	CaF2 crystals - liq. scintillator	Kamioka	Construction
CARVEL	48Ca		48CaWO4 crystal scint.		R&D
GERDA I	76Ge	14	Ge diodes in LAr	LNGS	Complete
GERDA II	76Ge	31	Point contact Ge in LAr	LNGS	Operating
Majorana D	76Ge	26	Point contact Ge	SURF	Operating
LEGEND-200	76Ge	172	Point contact Ge in LAr	LNGS	CDR→TDR
NEMO3	100Mo/82Se	6.9/0.9	Foils with tracking	LSM	Complete
SuperNEMO D	82Se	6.3	Foils with tracking	LSM	Construction
SuperNEMO	82Se	126	Foils with tracking		R&D
CUPID-0	82Se	5	ZnSe scint. bolometer	LNGS	Operating
AMoRE	100Mo	50	CaMoO4 scint. bolometer	Y2L	R&D
MOON	100Mo	200	Mo sheets		R&D
CUPID	100Mo	200	Li2MoO4	LNGS	R&D→CDR
COBRA	116Cd	10/183	CdZnTe detectors	LNGS	R&D
CUORICINO	130Te	10	TeO2 Bolometer	LNGS	Complete
CUORE-0	130Te	11	TeO2 Bolometer	LNGS	Complete
CUORE	130Te	206	TeO2 Bolometer	LNGS	Operating
SNO+	130Te	55	0.1% natNd suspended in Scint	SNOlab	Commissioning
KamLAND-ZEN	136Xe	380	2.7% in liquid scint.	Kamioka	Operating
NEXT-100	136Xe	90	High pressure Xe TPC	LSC	Construction
EXO-200	136Xe	60	Xe liquid TPC	WIPP	Operating
nEXO	136Xe	450/3330	Xe liquid TPC	SNOlab	R&D
PandaX-III	136Xe	200	High pressure Xe TPC	CJPL-II	R&D
DCBA	150Nd		Nd foils & tracking chambers		R&D

CUORE

CUORE



Cryogenic Underground Observatory for Rare Events

Primary goal: search for $0\nu\beta\beta$ decay in ^{130}Te

Detector design:

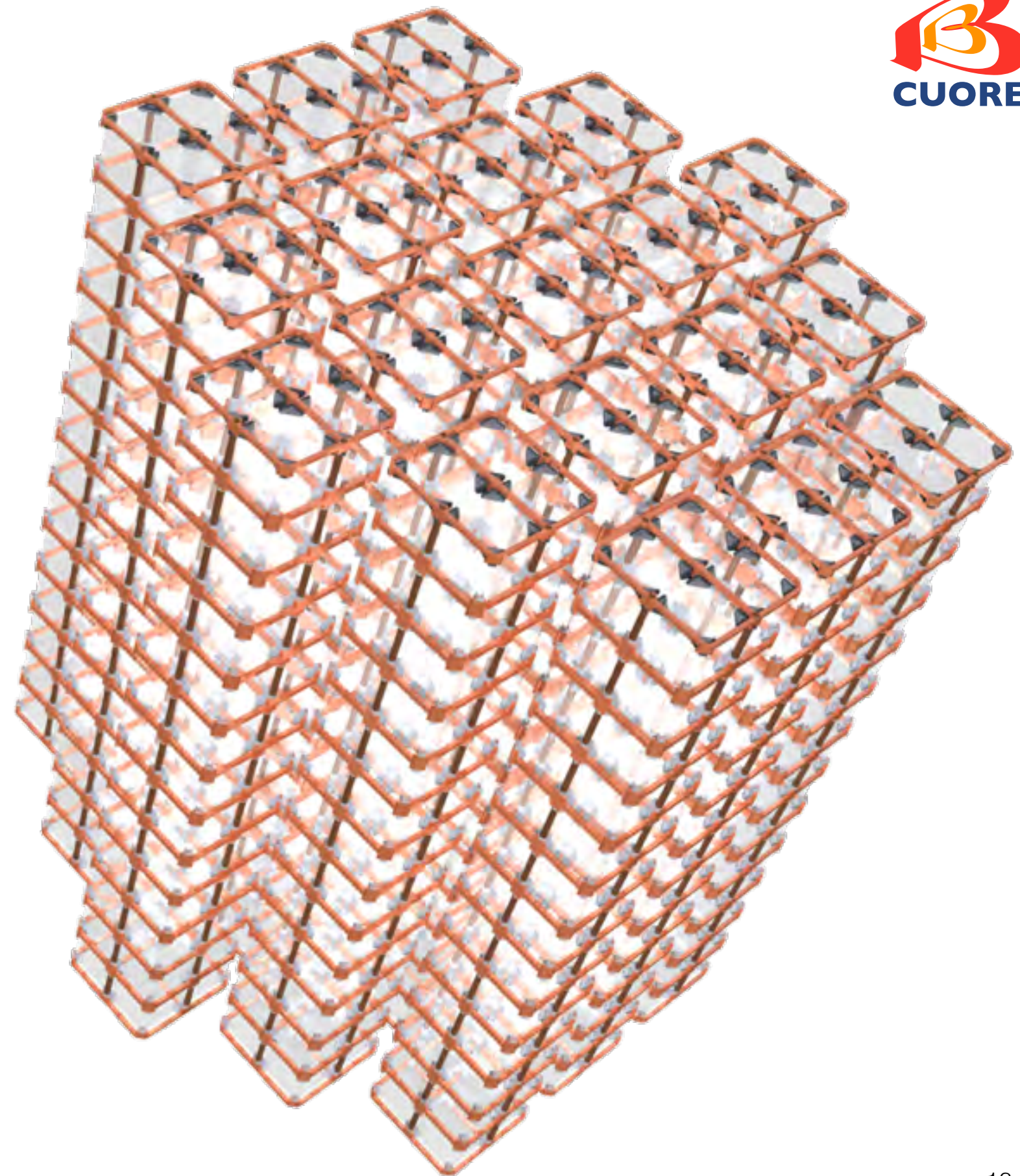
closely packed array of 988 TeO_2 crystals arranged in 19 towers

Design parameters:

- mass of TeO_2 : **742 kg** (206 kg of ^{130}Te)
- low background aim: **10^{-2} c/(keV·kg·yr)**
- target energy resolution: **5 keV** FWHM in the Region Of Interest (ROI)
- high granularity
- deep underground location
- strict radio-purity controls on materials and assembly

CUORE projected sensitivity (5 years, 90% C.L.):

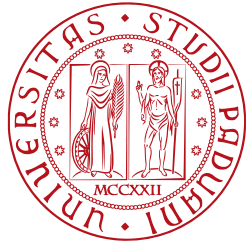
$$T_{1/2} > 9 \times 10^{25} \text{ yr}$$



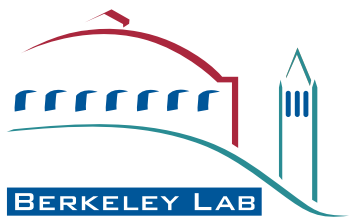
The CUORE Collaboration



UCLA



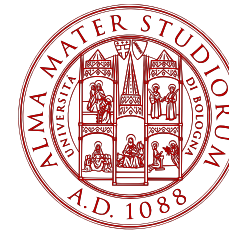
Yale



CAL POLY
SAN LUIS OBISPO



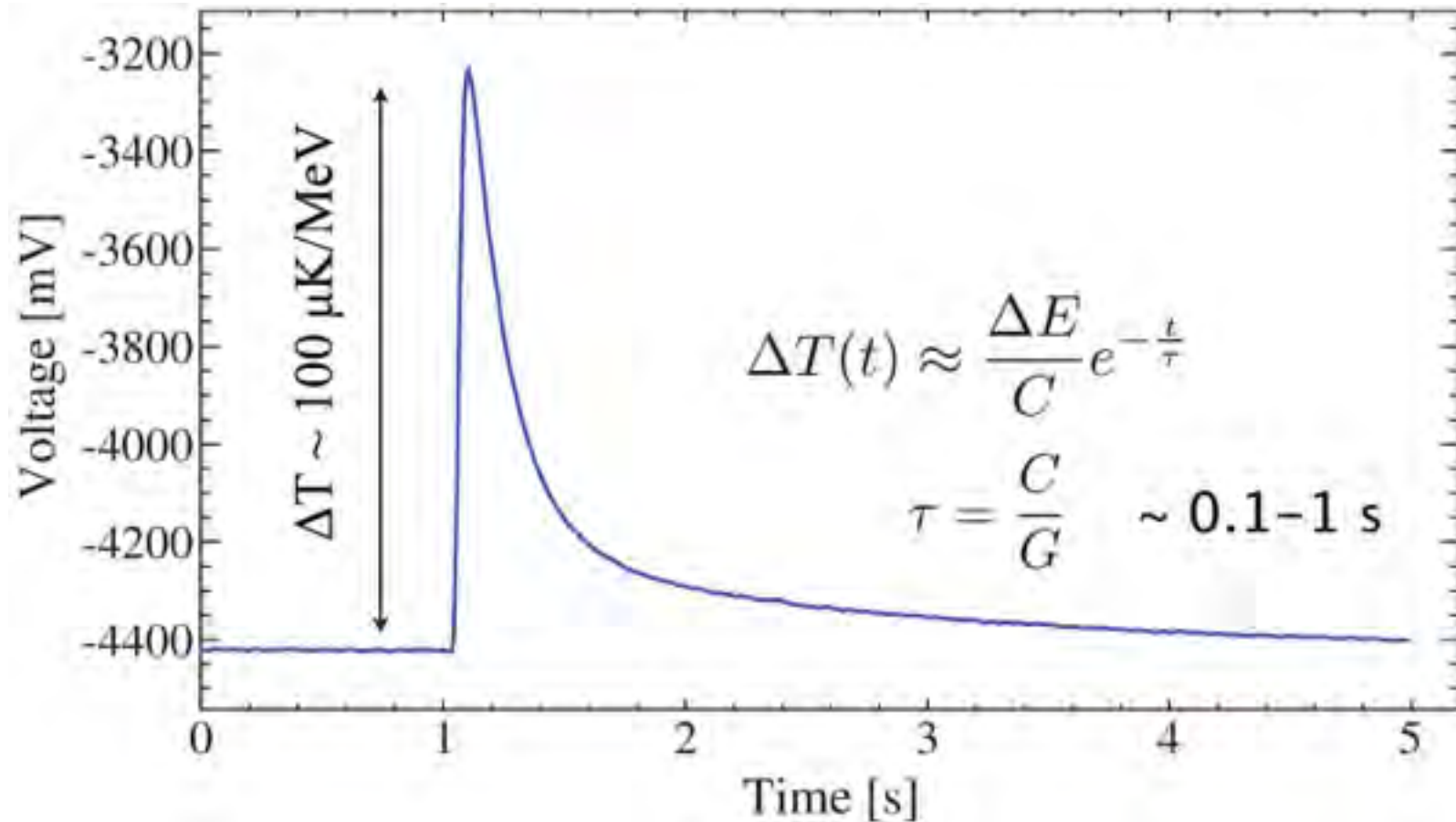
SAPIENZA
UNIVERSITÀ DI ROMA



The CUORE detectors

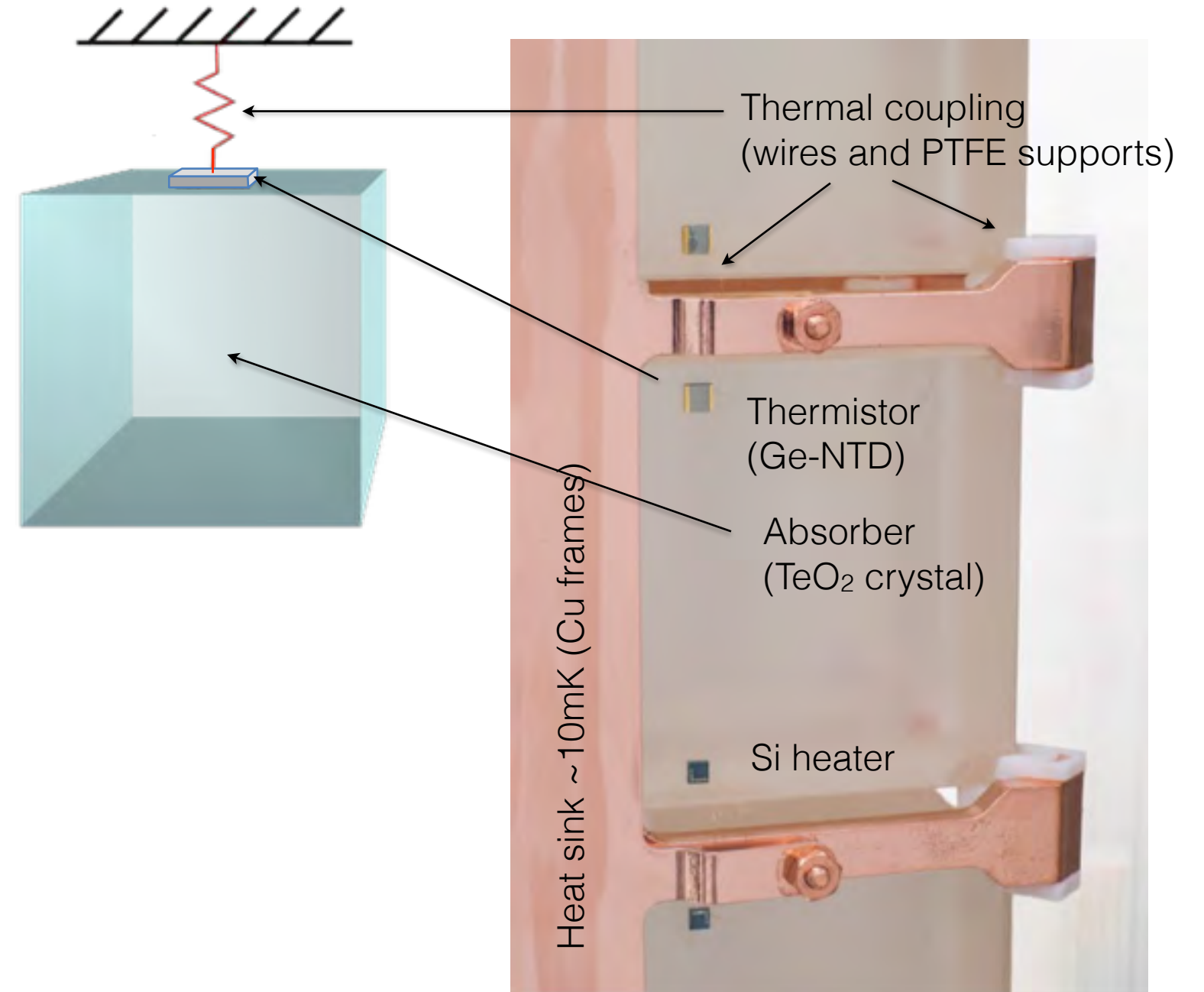
CUORE is a large array of thermal detectors

- low heat capacity @ T_{work} ($C \sim T^3$)
- excellent energy resolution ($\sim 0.2\%$ FWHM)
- same detector response for different particles
- slowness (suitable for rare event searches)

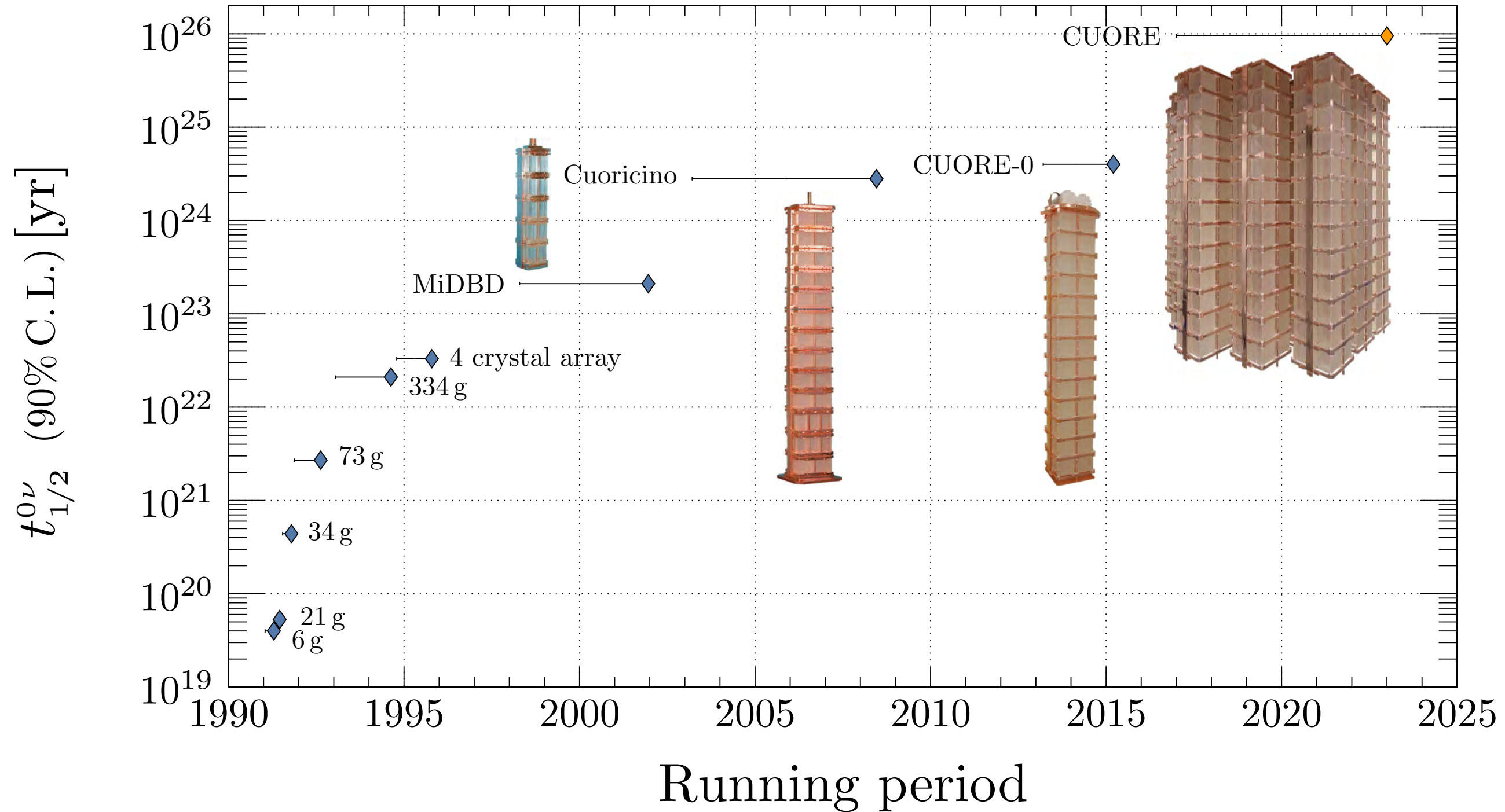


The concept: a pure calorimeter

The absorbed energy is converted into a variation of the crystal temperature, measured by the thermistor



TeO₂ arrays

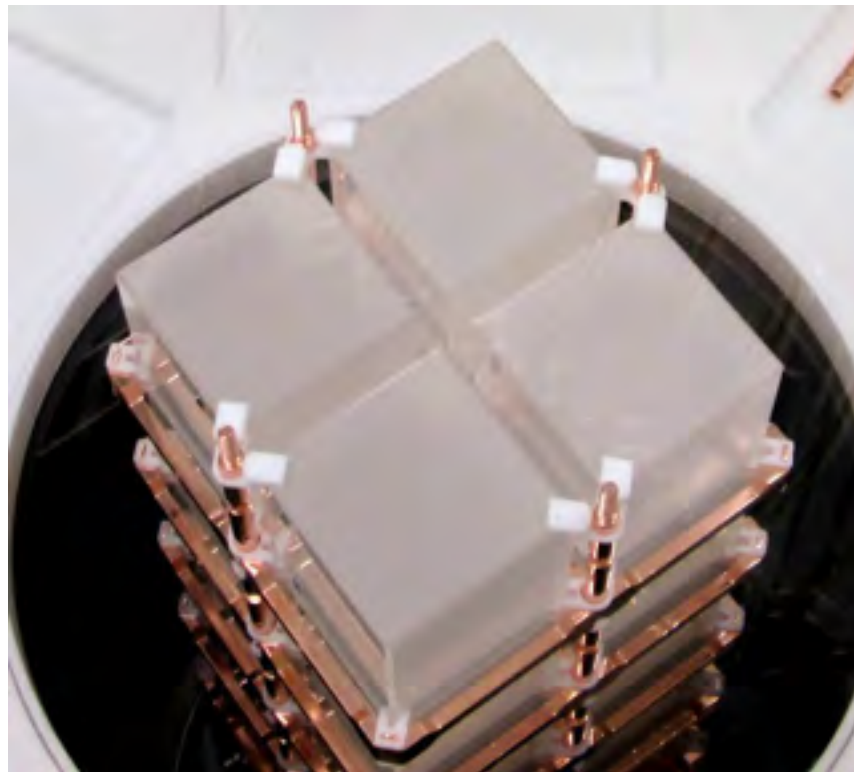


CUORE is the latest step in a long series of TeO₂ detectors which included two large demonstrators:

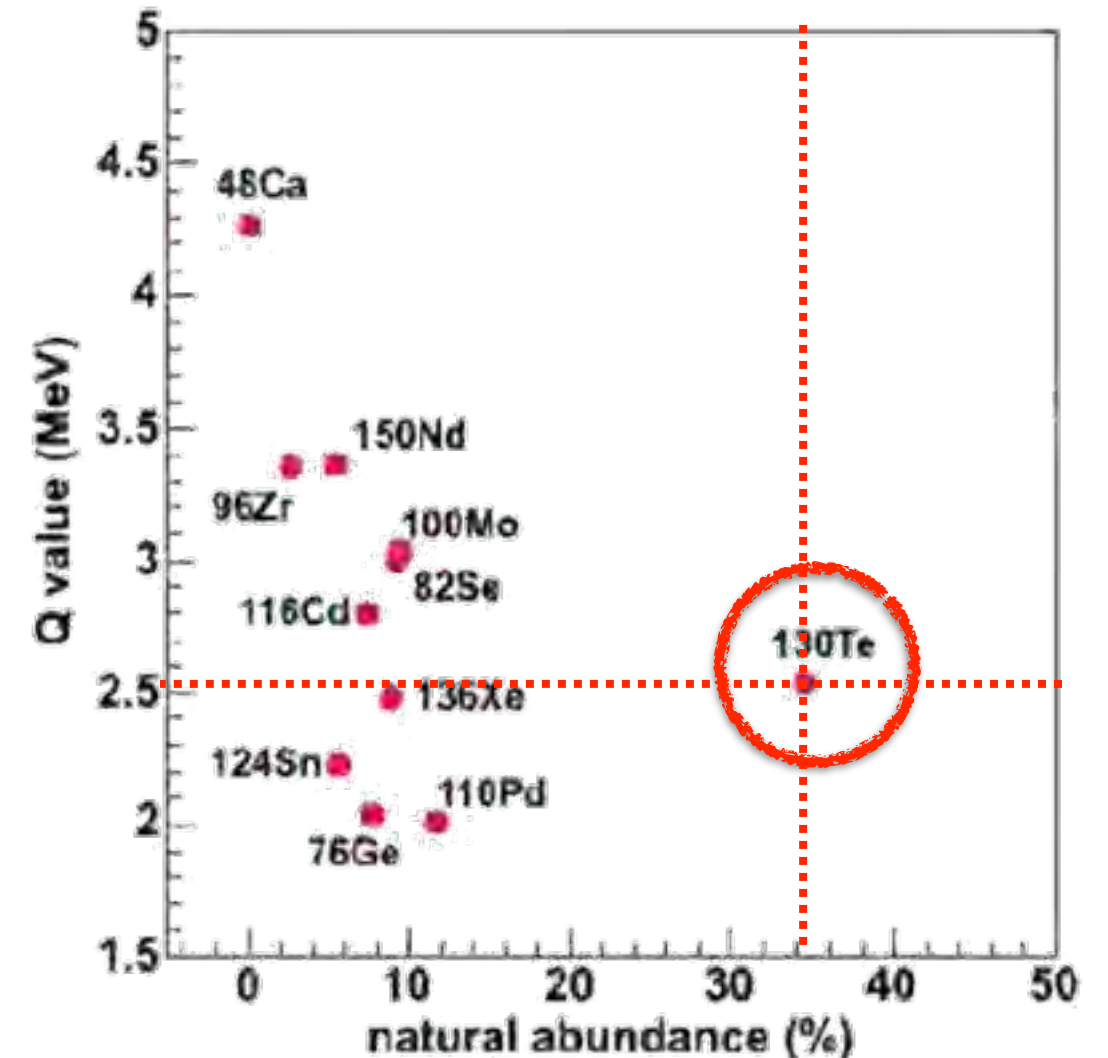
- Cuoricino
- CUORE-0

Why Tellurium?

- High natural isotopic abundance of source isotope (^{130}Te)
- ^{130}Te included in the detector: high efficiency
- $Q_{\beta\beta} = 2527.5 \text{ keV}$, in a region with relatively low β/γ background
- Excellent energy resolution (5 keV FWHM @ $Q_{\beta\beta}$)
- Reproducible growth of high quality crystals



Isotope	i.a.(%)	Q [MeV]
^{48}Ca	0.187	4.263
^{76}Ge	7.8	2.039
^{82}Se	9.2	2.998
^{96}Zr	2.8	3.348
^{100}Mo	9.6	3.035
^{116}Cd	7.6	2.813
^{130}Te	34.1	2.527
^{136}Xe	8.9	2.459
^{150}Nd	5.6	3.371



CUORE @ LNGS



1400 m of rock (~ 3600 m.w.e.) deep

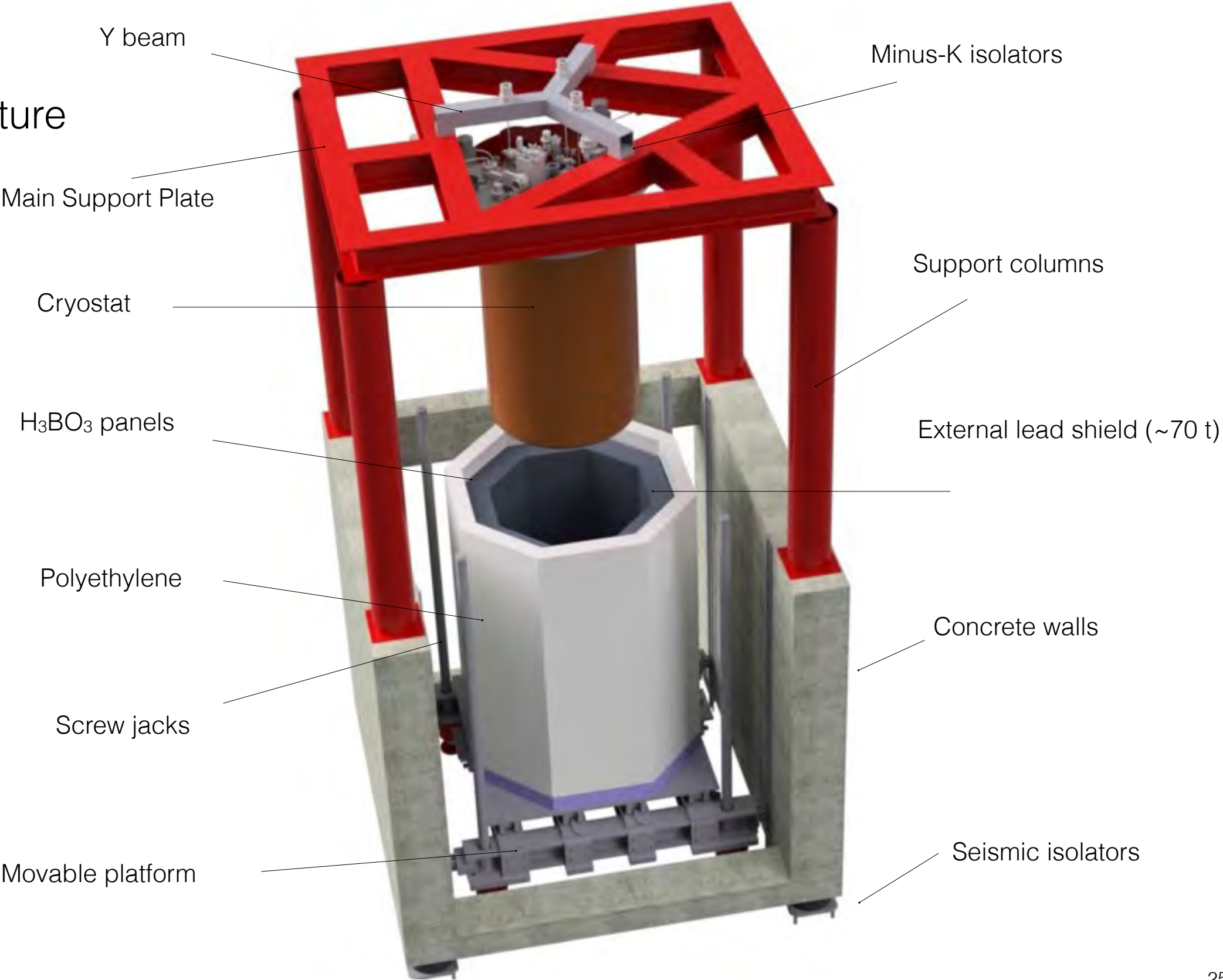
- μ 's: $\sim 3 \times 10^{-8} / (\text{s} \cdot \text{cm}^2)$
- γ 's: $\sim 0.73 / (\text{s} \cdot \text{cm}^2)$
- neutrons: $4 \times 10^{-6} \text{ n}/(\text{s} \cdot \text{cm}^2)$ below 10 MeV

CUORE @ LNGS



Underground Laboratory

- Three-story building
- Hosting the cryostat supporting structure

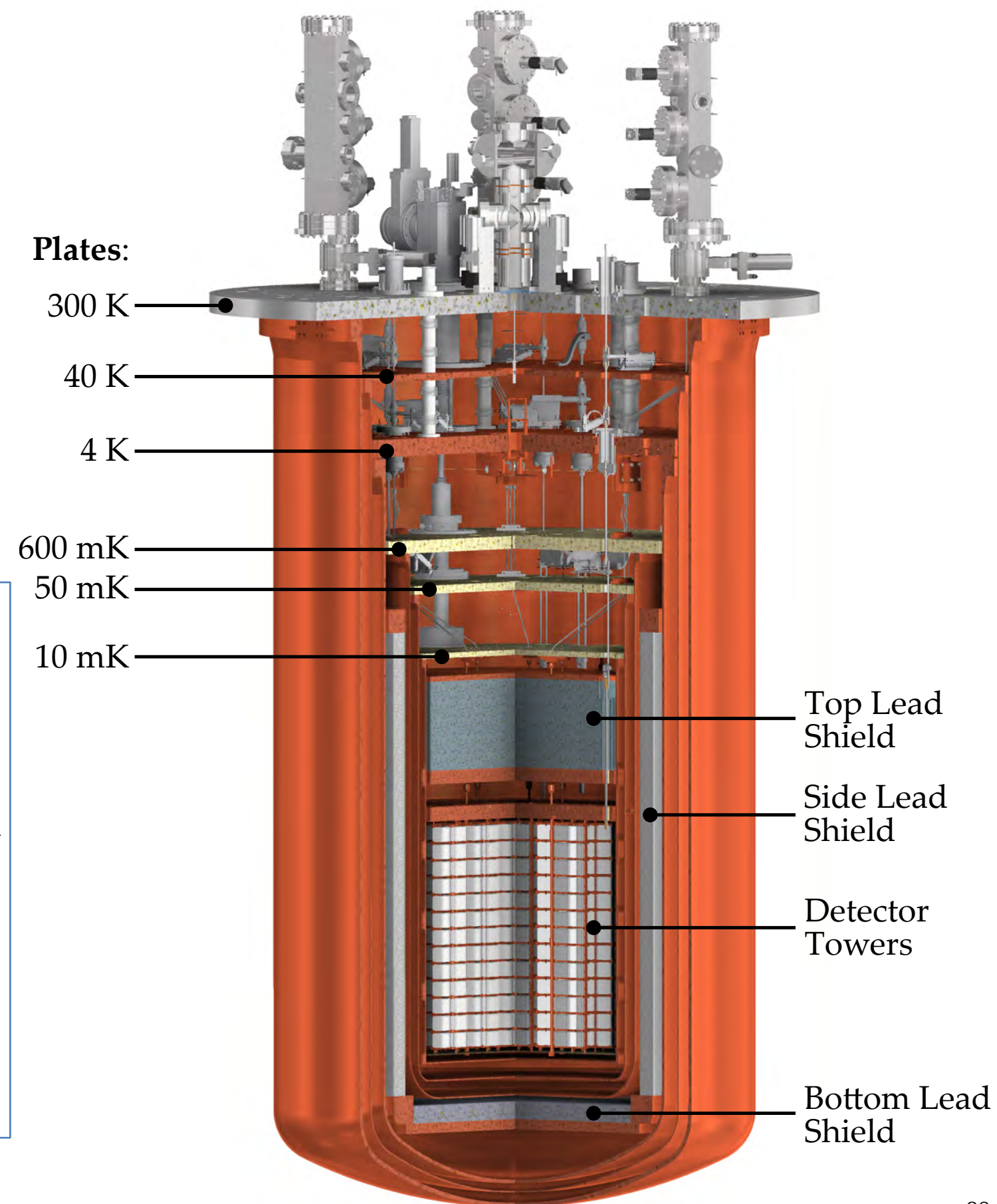


The CUORE cryostat

Challenges:

- Cool down ~1 ton detector to ~10 mK
- Mechanically decoupled for extremely low vibrations
- Low background environment

- Cryogen-free cryostat
- Fast Cooling System (^4He gas) down to ~50K
- 5 pulse tubes cryocooler down to ~4K
- Dilution refrigerator down to operating temperature ~10 mK
- Nominal cooling power: $3 \mu\text{W}$ @ 10mK
- Cryostat total mass ~30 tons
- Mass to be cooled < 4K: ~15 tons
- Mass to be cooled < 50 mK: ~3 tons (Pb, Cu and TeO_2)

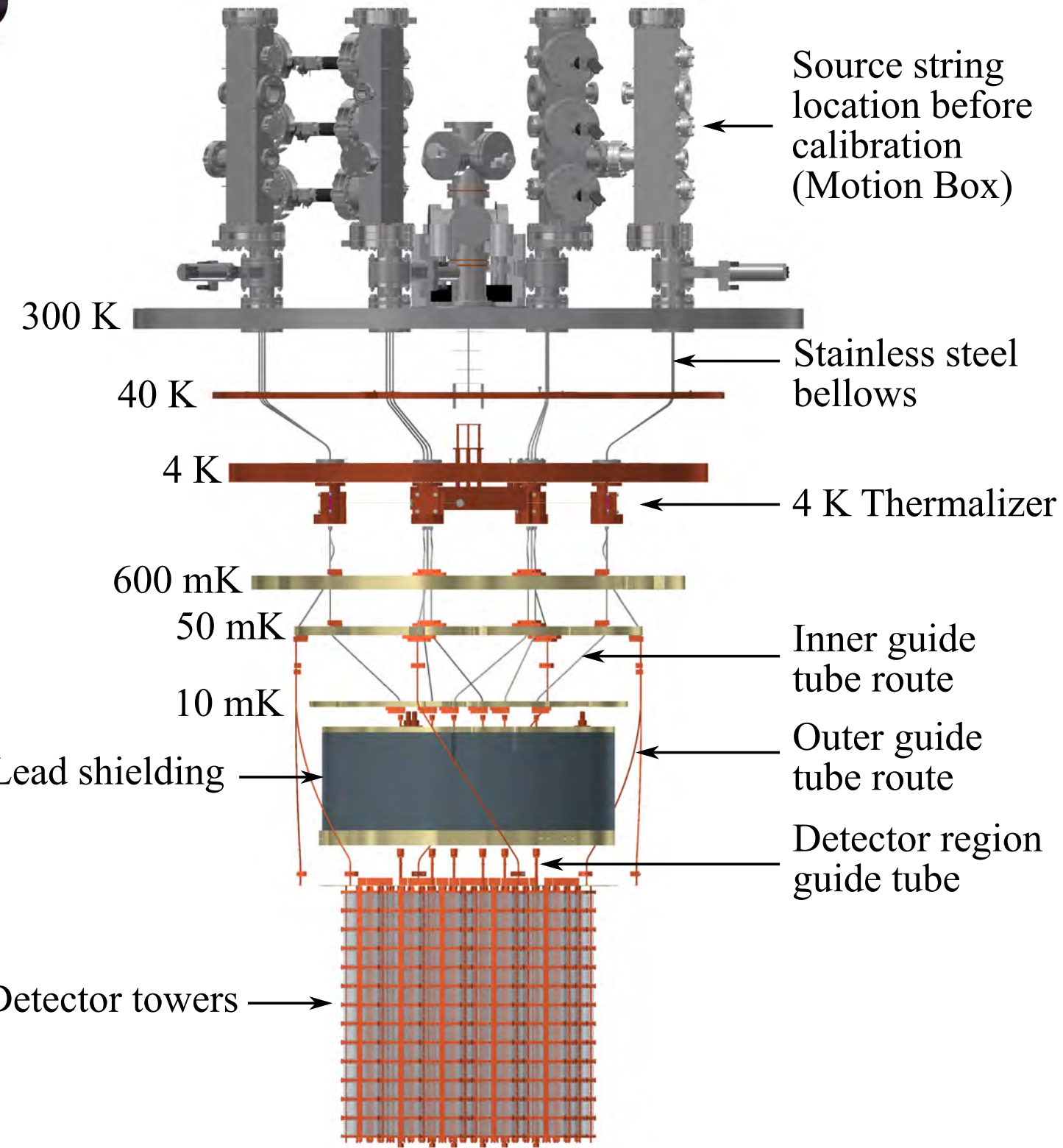
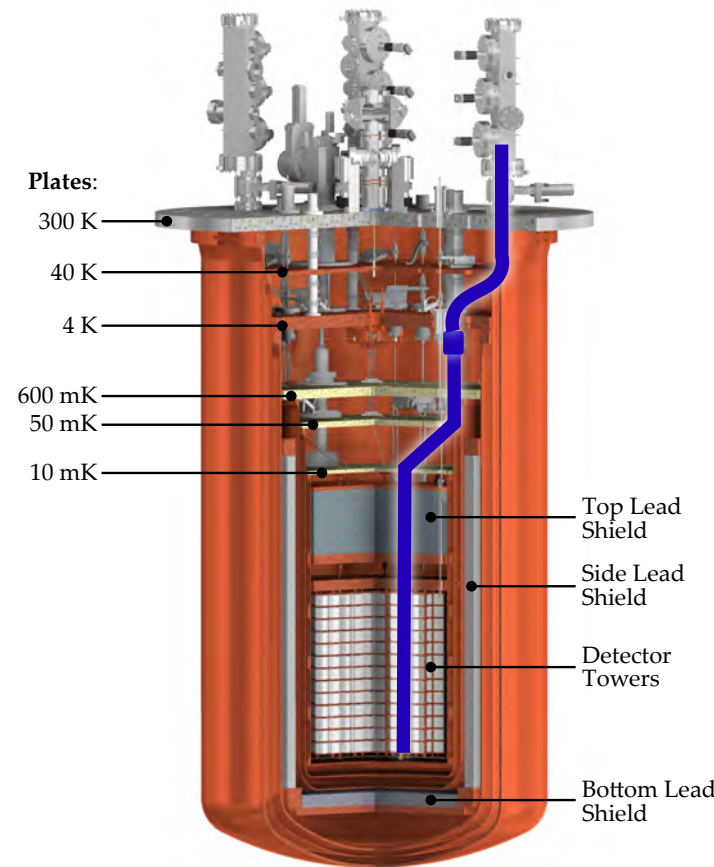
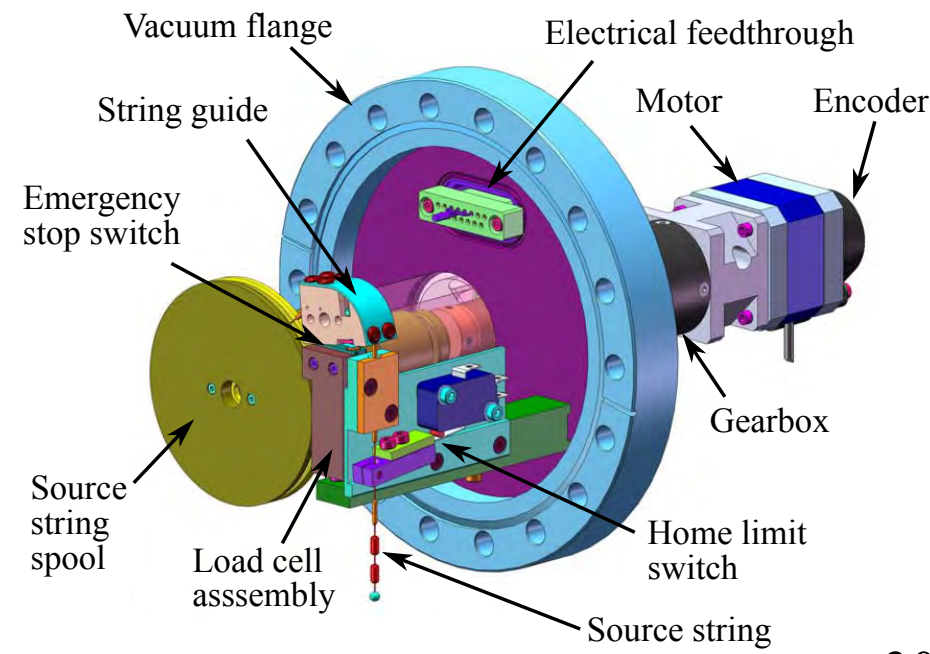
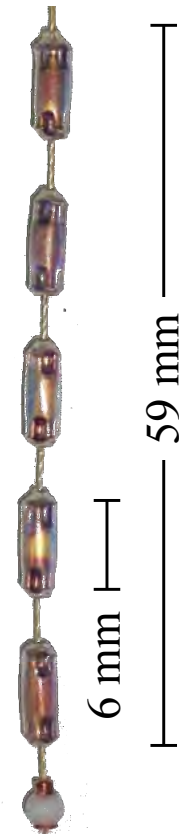
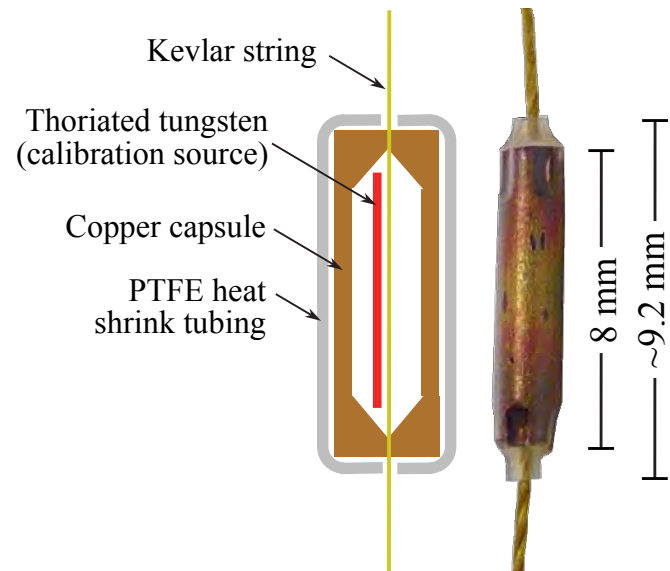
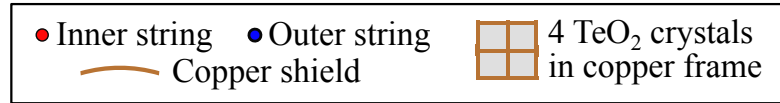
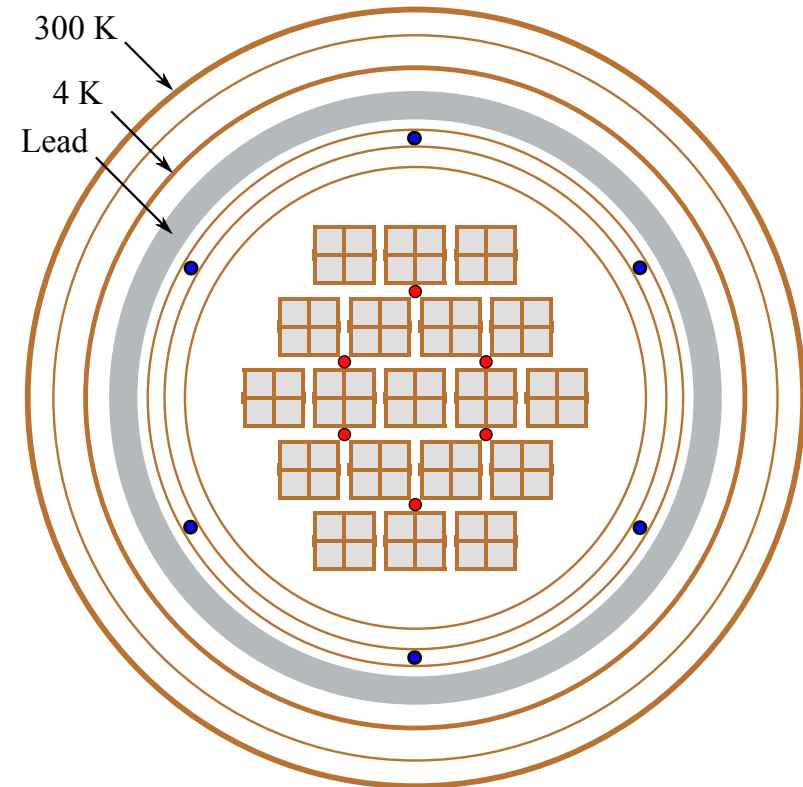


Passive shielding

- Protect the detectors with a heavy shield against gamma and neutron activity from external sources (~70 tonnes lead + H_3BO_3)
- Select materials that don't contribute themselves to the background level (ancient roman lead and selected NOSV copper)
- Cool down inner layers of the shielding to the correct temperature (2.5 tonnes @ 50mK + 5.5 tonnes @ 4K)



DCS



Detector installation

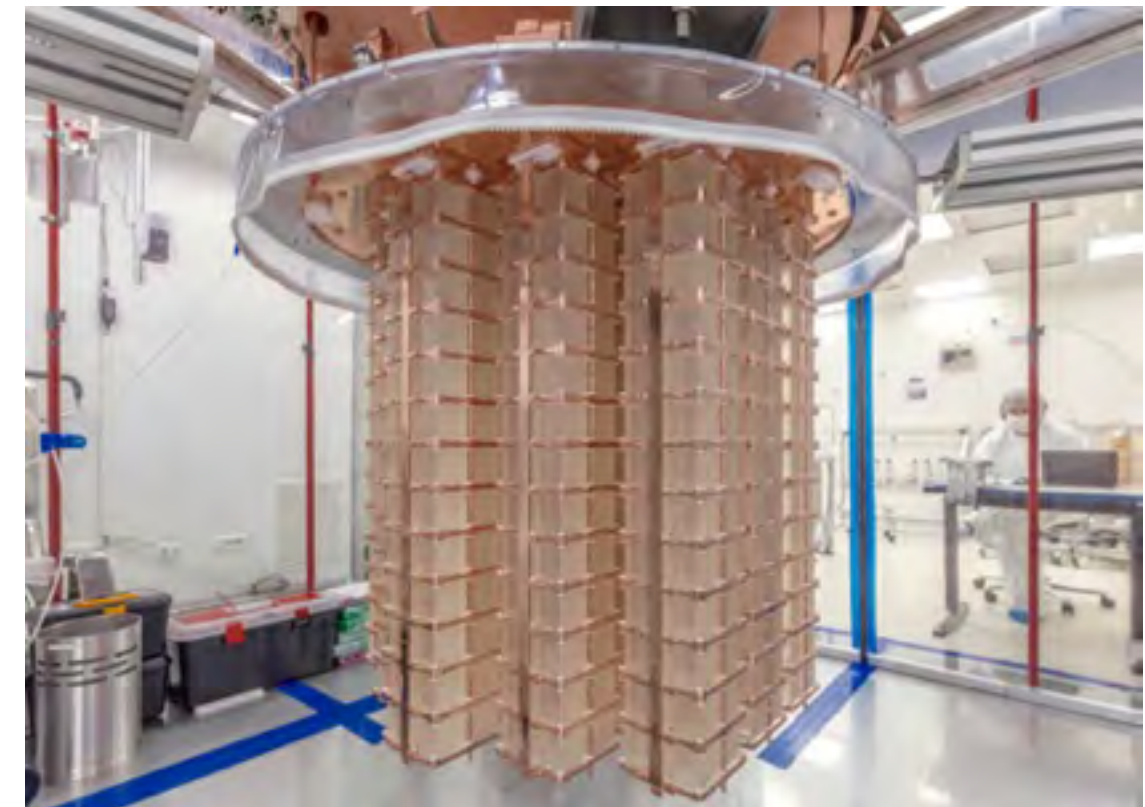
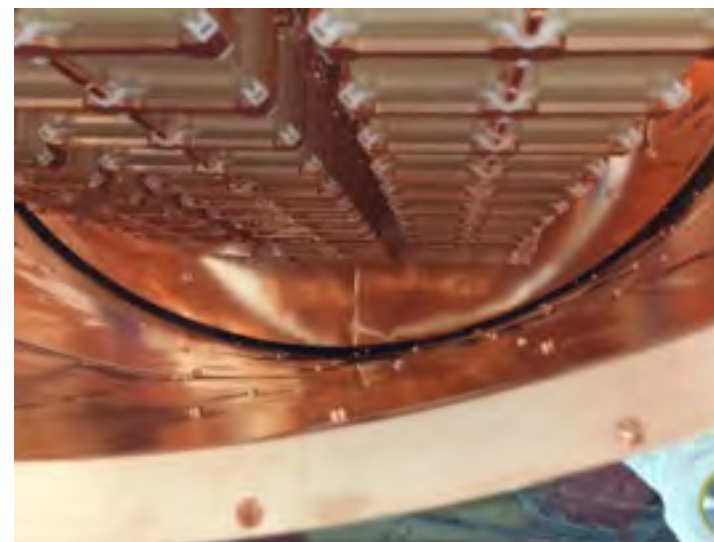
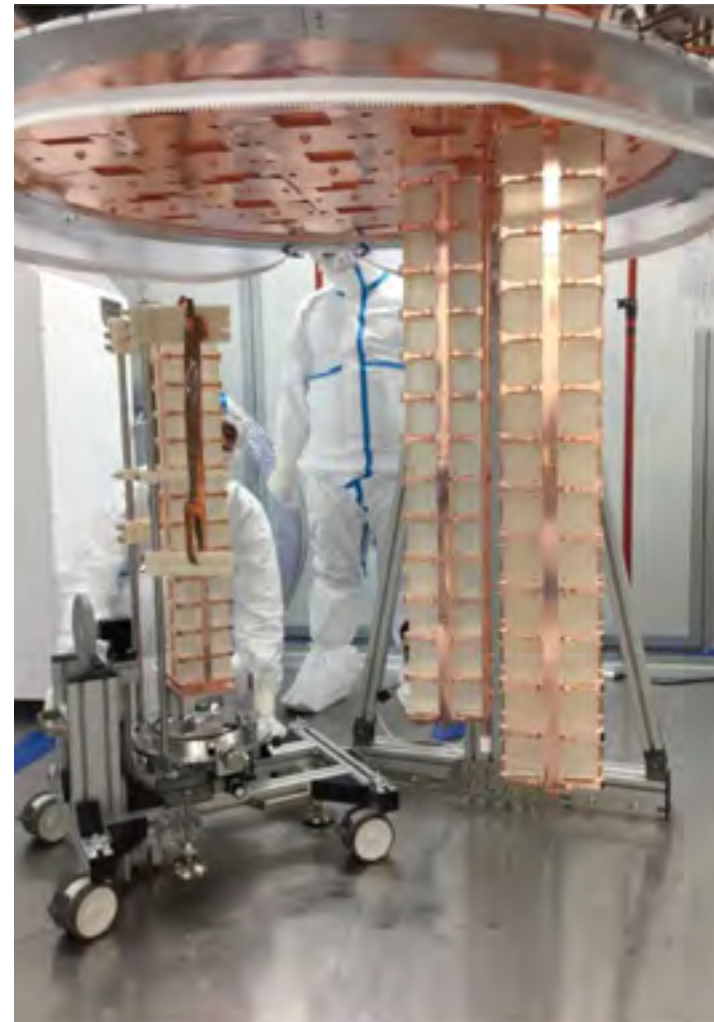
Performed in a radon-free environment:

- protected area inside the CUORE clean room flushed with radon-free air (Rn concentration $< 0.1 \text{ Bq/m}^3$) for operators life support
- protective bags flushed with nitrogen for overnight and emergency storage
- teams composed of 3 operators spending the minimum amount of time in the cleanroom, following strict protocols developed during months of training and test with mockup components

Towers installation completed on August 26, 2016

September-October 2016:

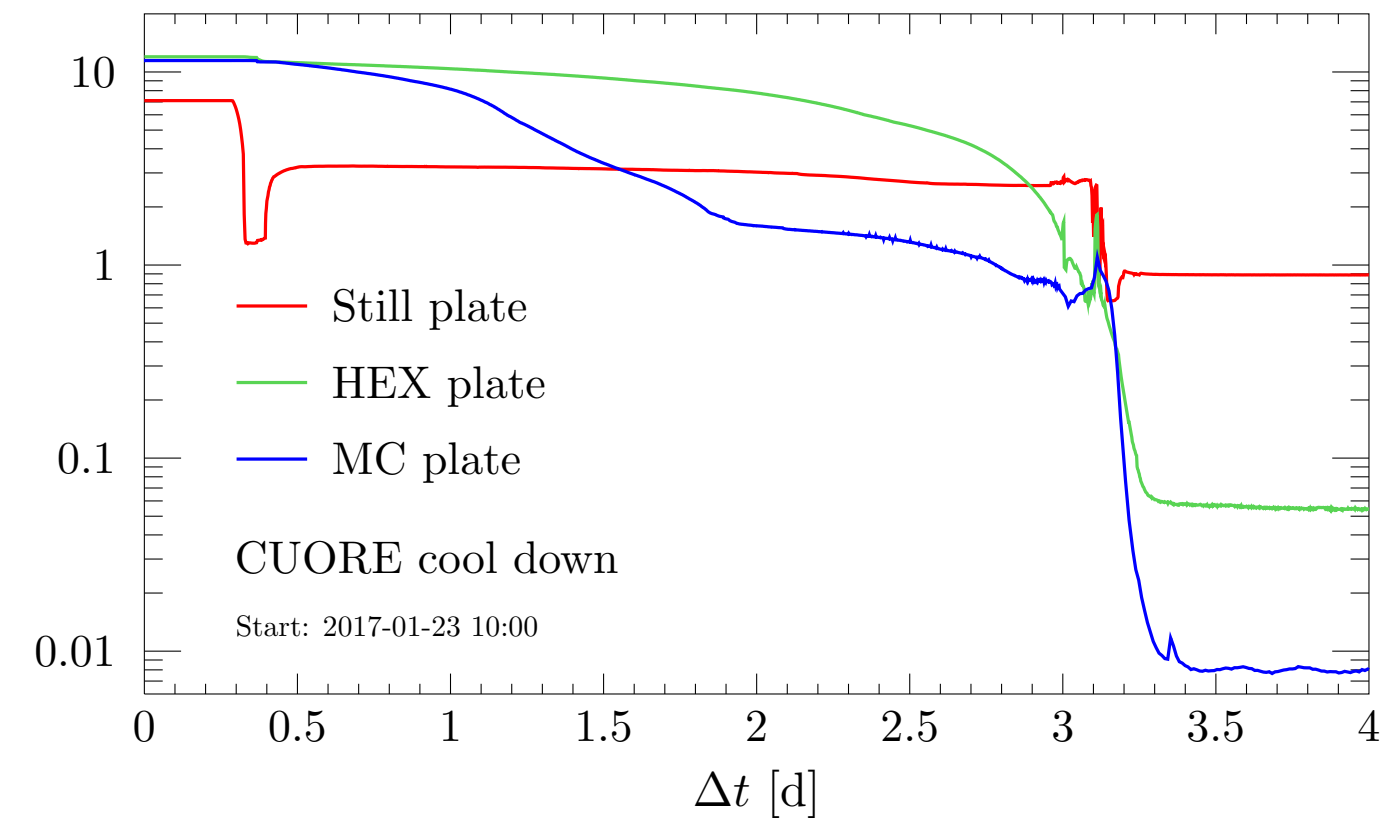
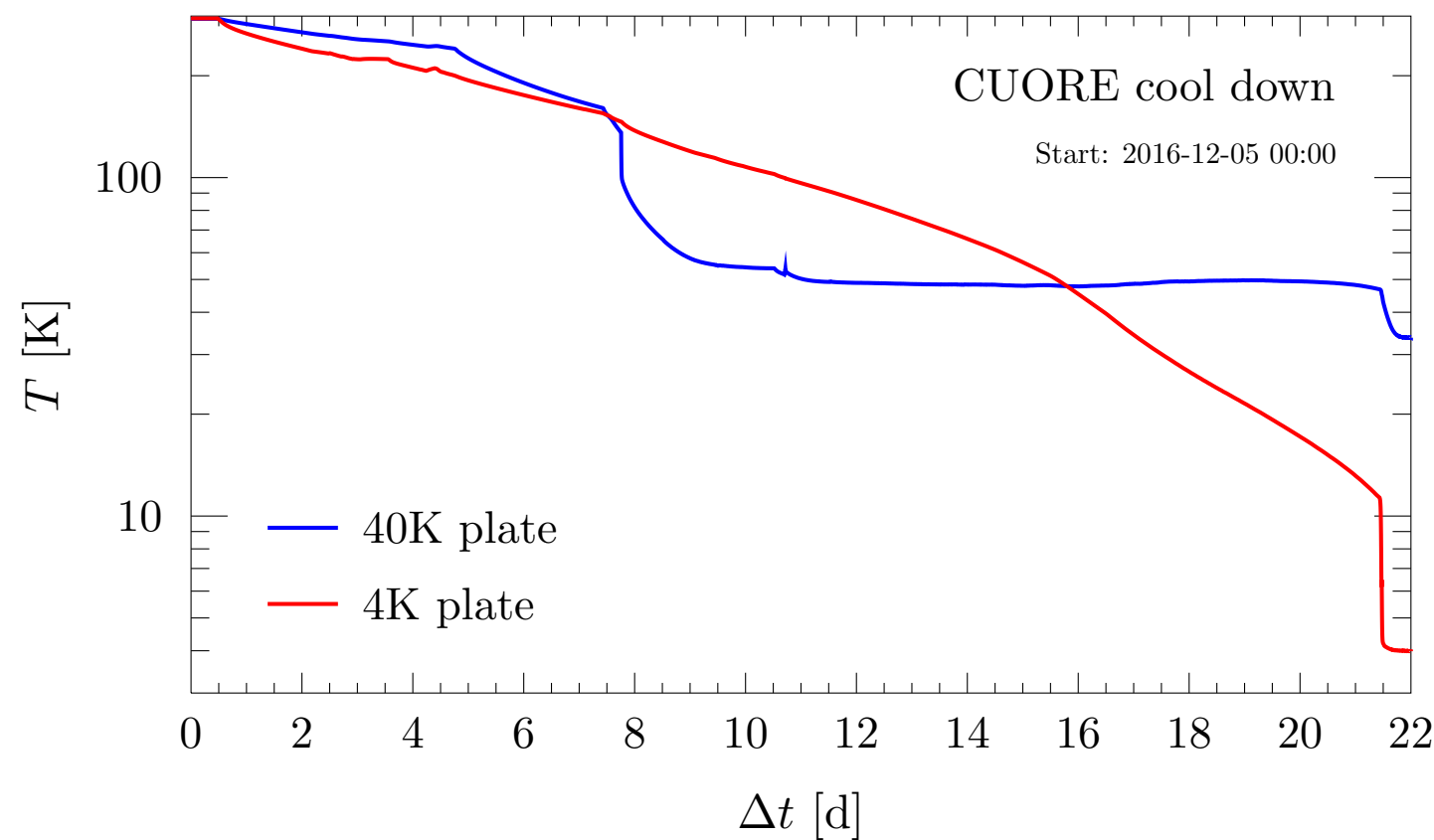
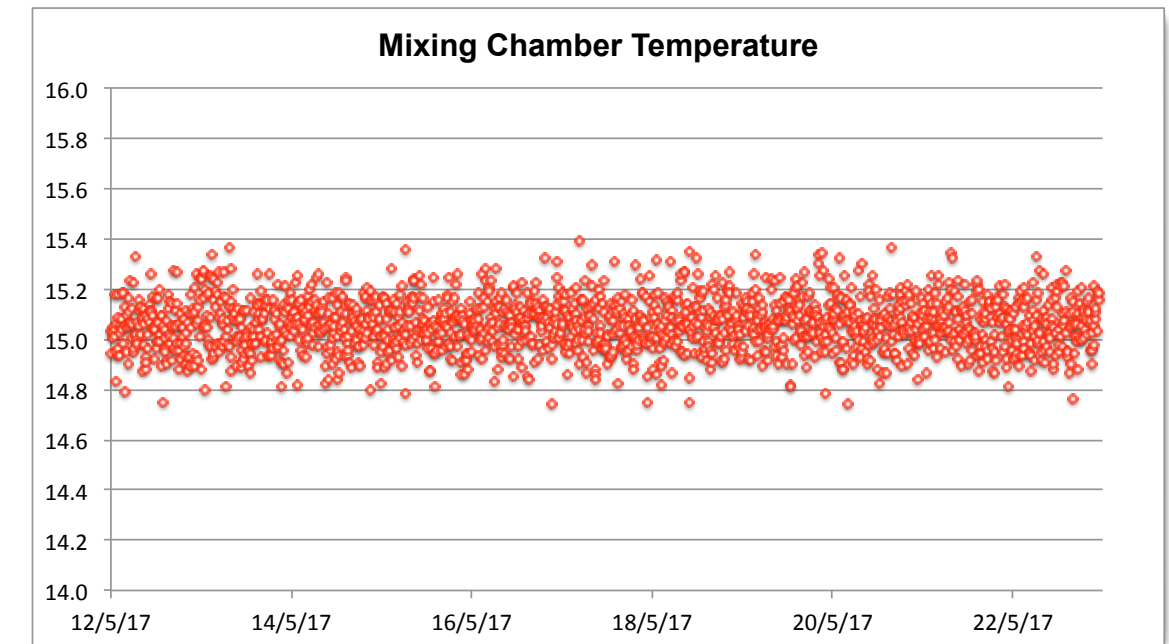
- installation of the cryostat interfaces (protective tiles) and radiation shields
- read-out tests



Detector cool down

Started at the beginning of December 2016:

- 300 K \rightarrow 4 K in about 22 days
- 4 K \rightarrow 7 mK in 3.5 days
- reached a stable base temperature of \sim 7 mK on Jan 27, 2017
- lowest observed temperature: 6.7 mK
- observed first detector pulses just after the cool down without any optimisation!



OPERATION & PERFORMANCE

CUORE (pre-)operation

02/12/2016

Construction is complete

05/12/2016 - 23/01/2017

Cooldown

09/01/2017 - 12/04/2017

Detector commissioning (optimisation phase I)

14/04/2017 - 14/06/2017

Science run n. 1

11/06/2017 - 26/07/2017

Optimisation phase II (warm-up to ~100K)

27/07/2017 - 17/09/2017

Science run n. 2

18/09/2017 - 13/05/2018

Optimisation phase III (new warm-up)

19/05/2018 - 03/09/2018

Science run

03/09/2018 - 04/12/2018

System maintenance

04/12/2018 - 01/02/2019

Science run

01/02/2019 - 17/04/2019

System maintenance

17/04/2019 -

Science run

2 scientific runs:

Phys. Rev. Lett. 120 (2018) 132501

Detector pre-operation

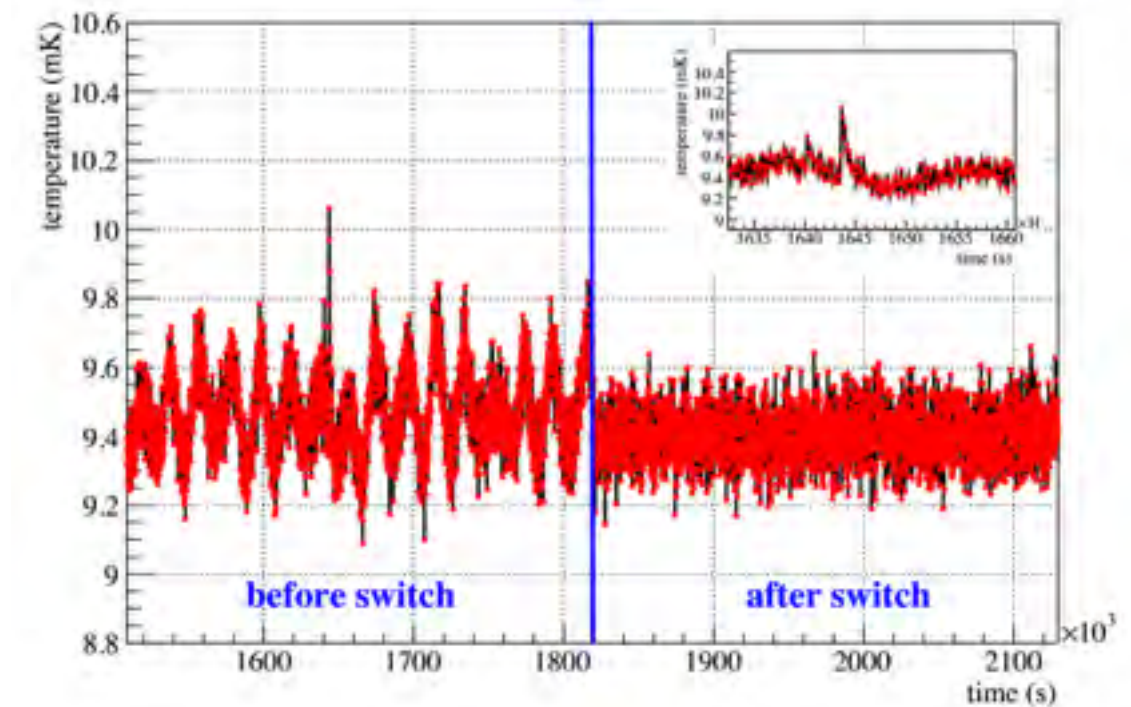
After the successful cool-down we faced the challenge to operate a thousand bolometers in a completely new system.

A long list of tests and activities

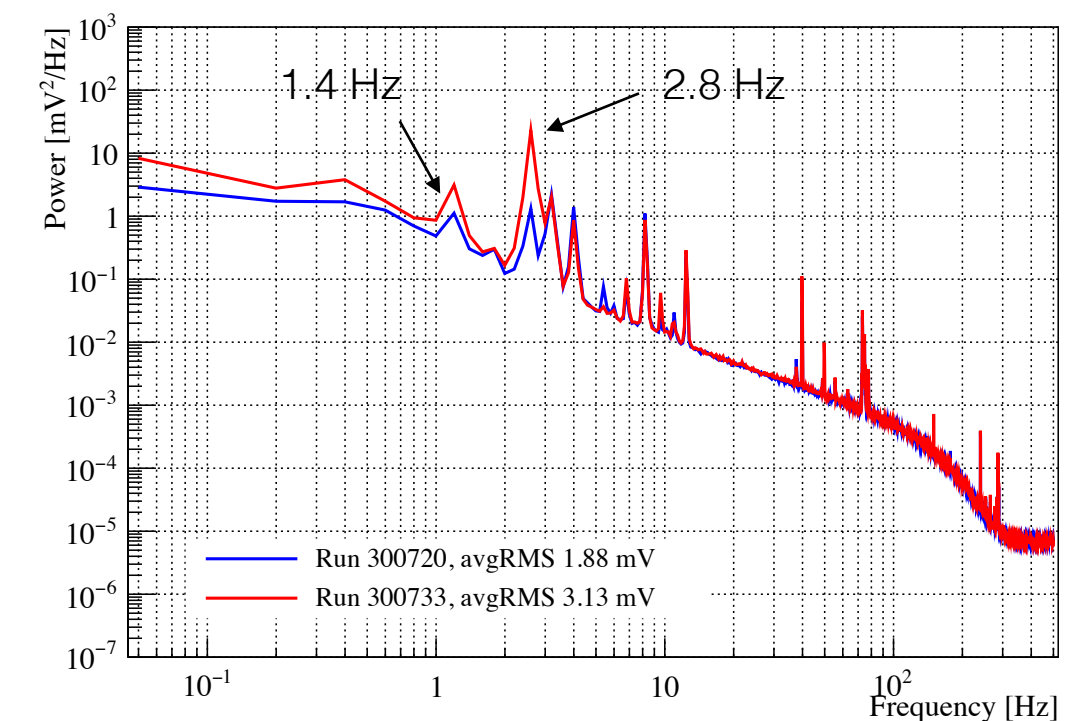
- DAQ and front-end electronics optimization
- Detector working points
 - Select representative subset
 - Load curves (to select optimal working points)
 - Temperature scan for the best operating conditions
- Noise reduction
 - Linear drives to control the pulse tube (PT) motor-heads
 - Monitor and control the relative phase shifts between different PT's using pressure sensors installed on the PT lines
 - Impressive results both in terms of temperature stabilisation and noise abatement

End of March 2017:

- Closed first optimisation phase
- **Ready to start calibrations and science runs**
- **Selected working temperature: 15 mK**



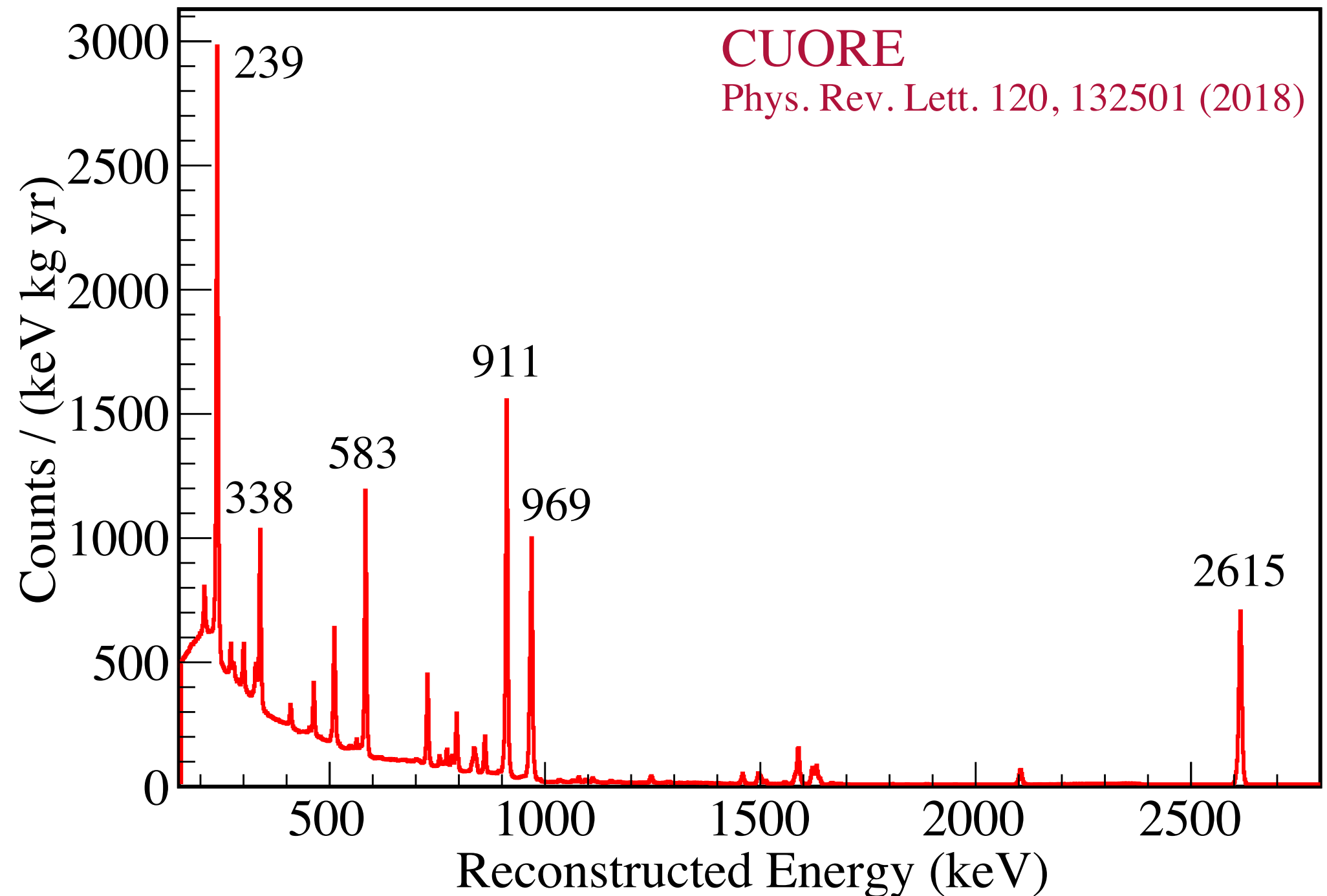
Average Noise Power Spectrum: ch. 142 runs 300720, 300733



Calibration spectrum

239 keV - ^{212}Pb
338, 911, 969 keV - ^{228}Ac
583, 2615 keV - ^{208}Tl

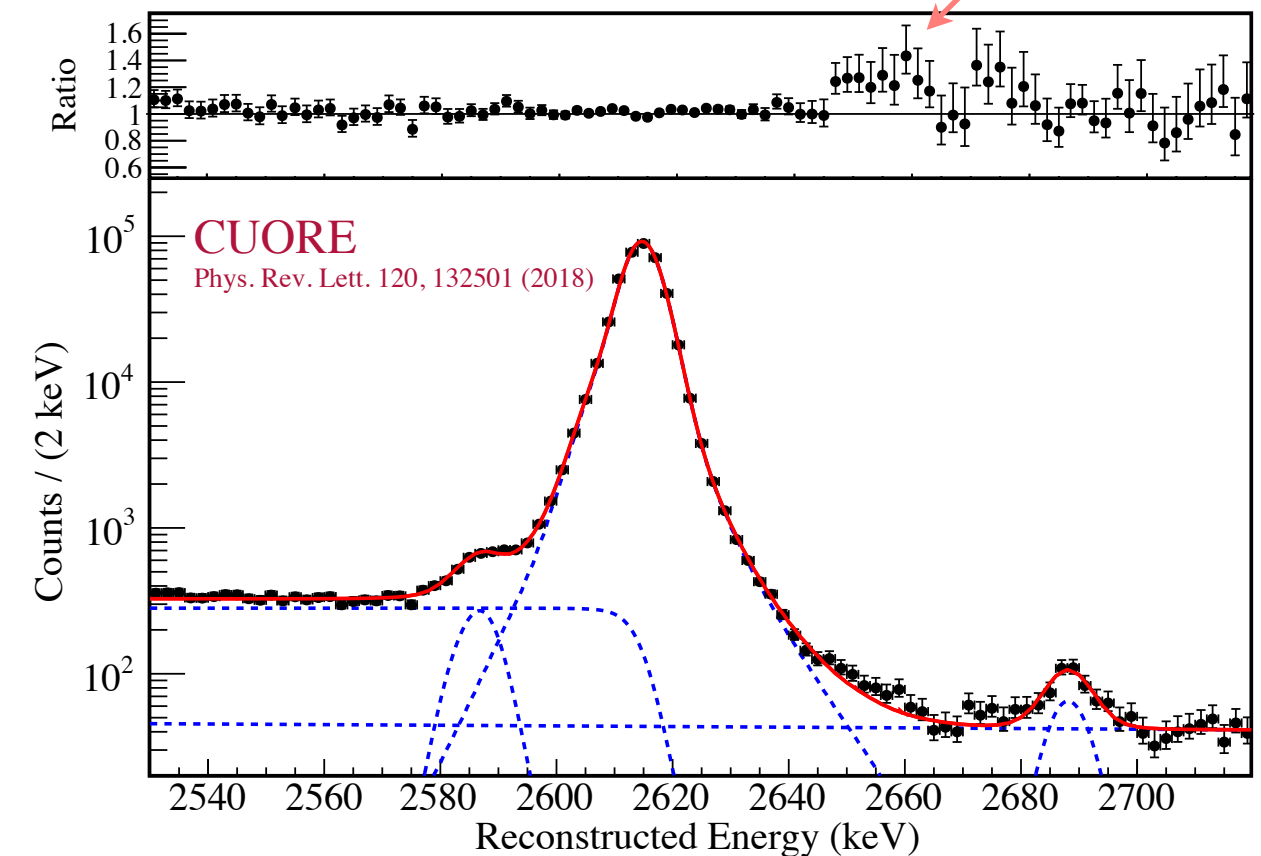
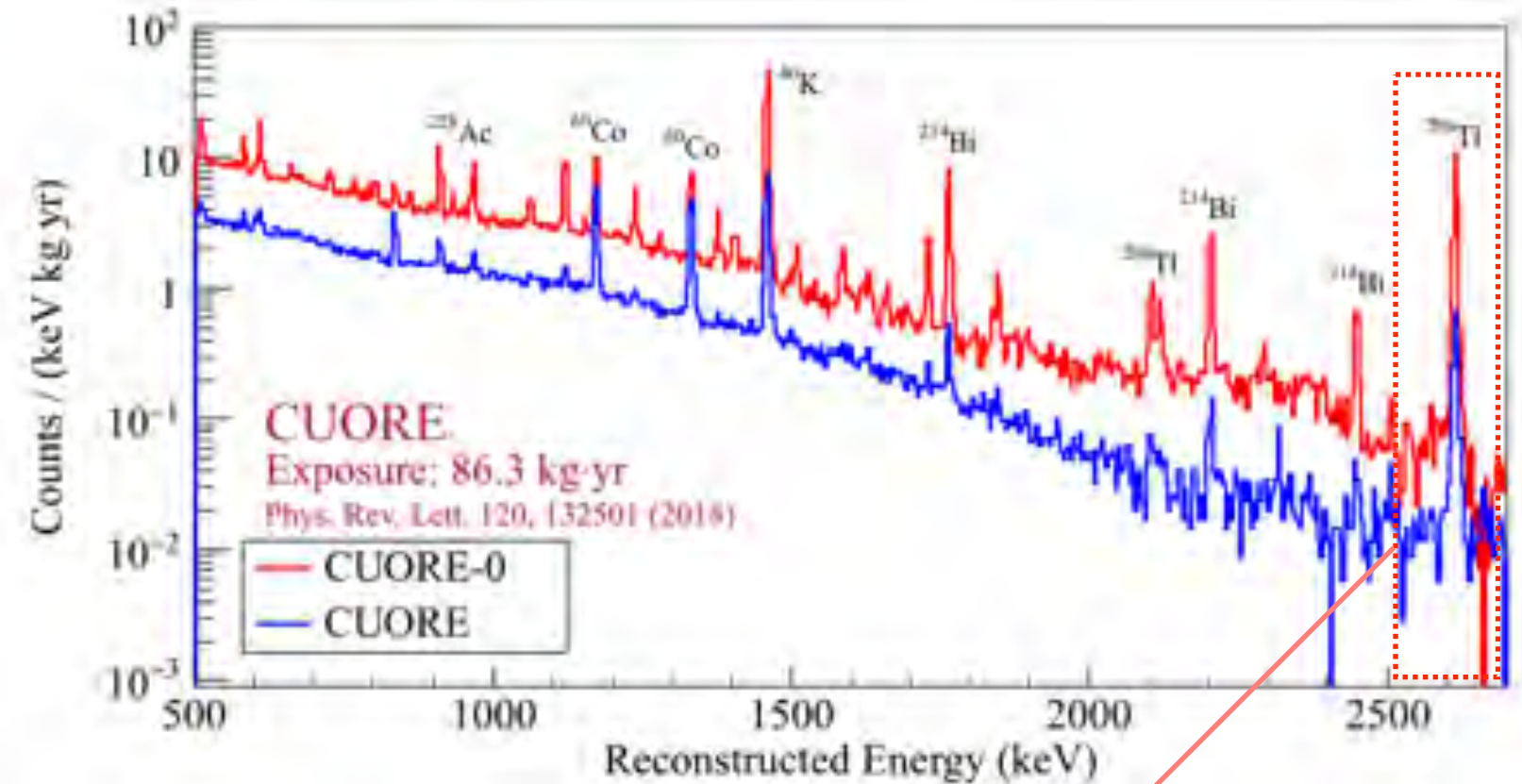
- Calibration strings deployed inside the CUORE detector
- Summed energy spectrum of all the CUORE detectors-datasets
- Calibration data used for:
 - ▶ energy scale calibration
 - ▶ thermal gain stabilisation
 - ▶ detector response (line shape) study



Detector performance: line shape



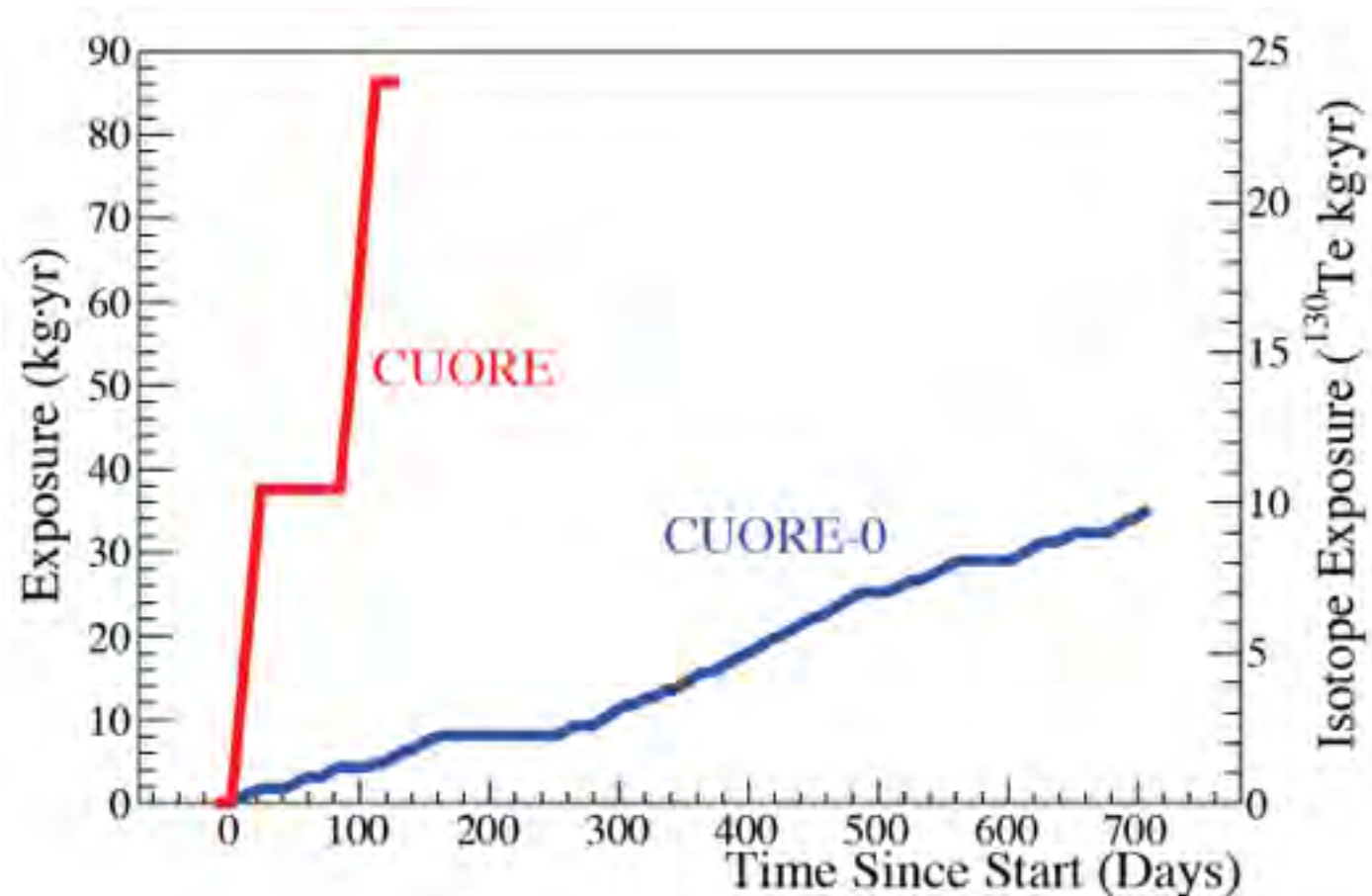
- The prominent ^{208}Tl line in the calibration spectrum are used to model the detector response to a monochromatic electron-like energy deposition
- 19 simultaneous fit on the data from all the channels of each tower, with some parameters (backgrounds) common to the whole tower and some defined channel-by-channel (resolutions, normalisations)
- Eventually only channel-dependent parameters (**signal**) are used to build the PDF for the ROI fit
- Fit components:
 - triple gaussian for the photopeak**
 - step-wise smeared multi-compton background
 - combination of gaussian X-rays escape lines
 - linear background
 - single gaussian line for the coincident absorption of 2615-keV and 583-keV followed by a single escape process



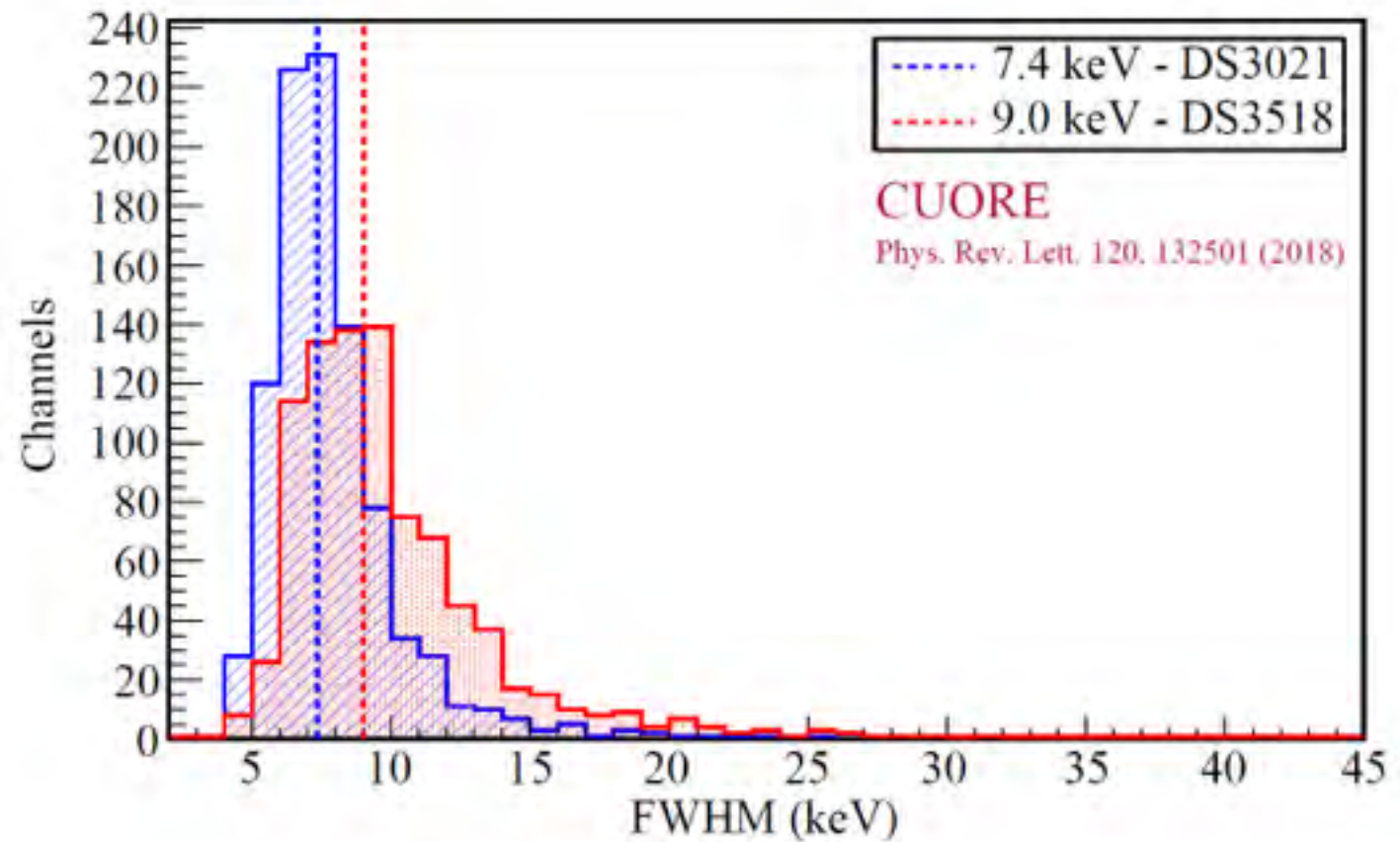
CUORE: 2017 science runs

Collected 86.3 kg·yr of TeO_2 over 7 weeks in summer 2017

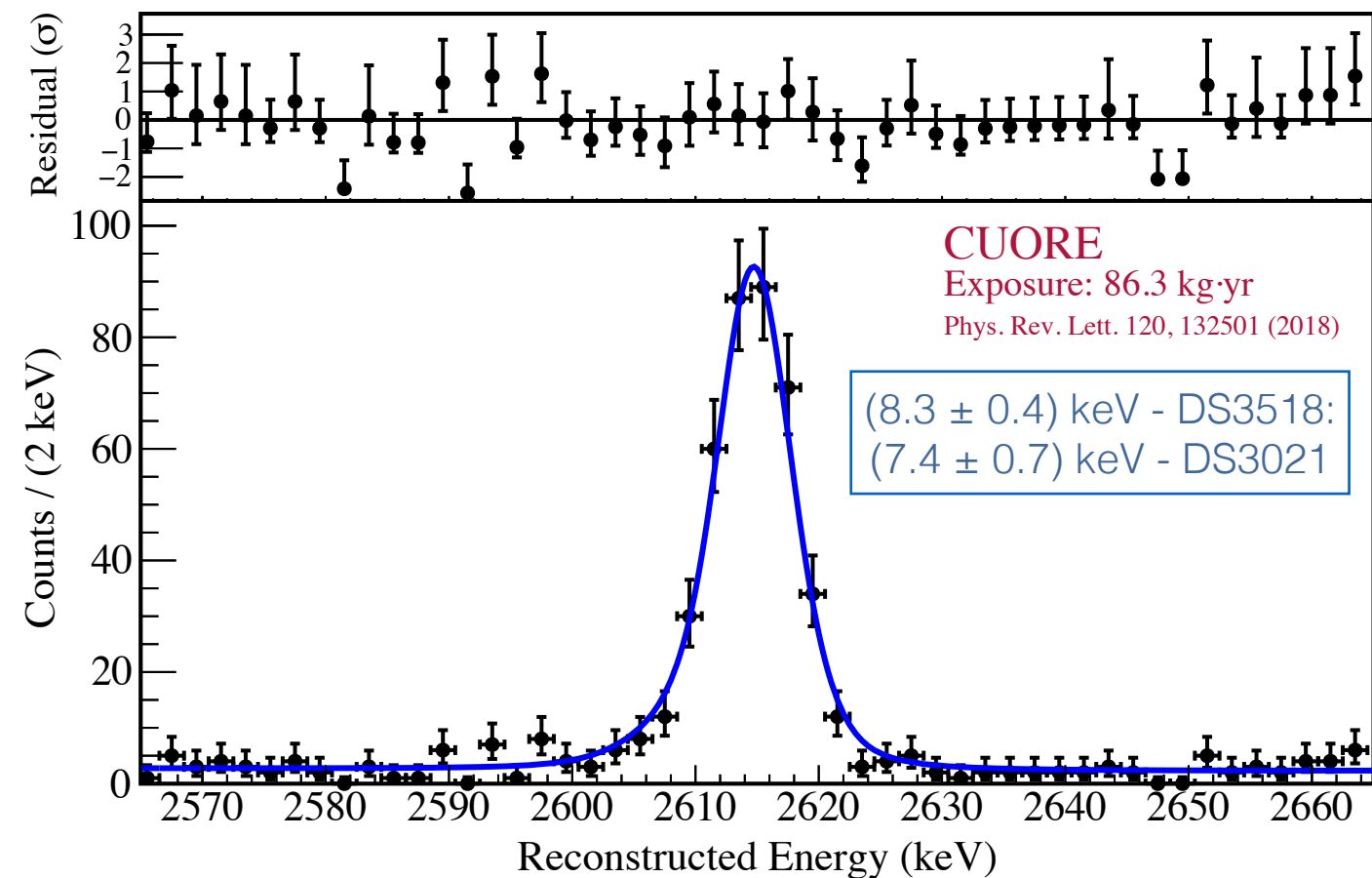
- 99.6% of channels active
- 92% of channels passing analysis cuts
- average (exposure weighed) energy resolution: (7.7 ± 0.5) keV FWHM
- signal efficiency: $\sim 80\%$



Calibrations



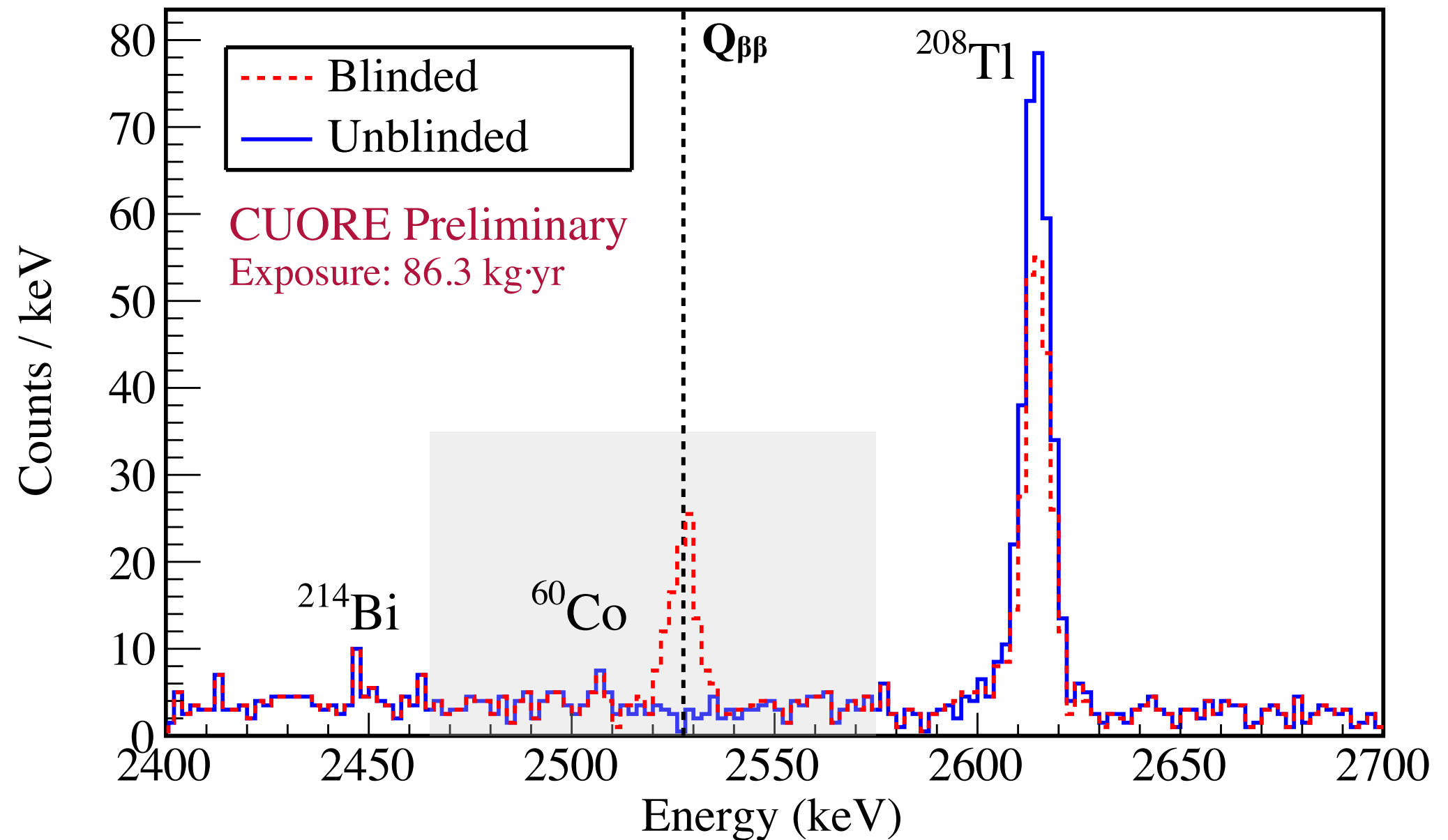
Science runs



RESULTS

Blinding procedure

- To blind our data we randomly move a fraction of events from +/- 20 keV of 2615 keV to the Q-value and vice versa
- The blinding algorithm produces an artificial peak around the $0\nu\text{DBD}$ Q-value hiding the real $0\nu\text{DBD}$ rate of ^{130}Te
- This method of blinding the data preserves the integrity of the possible $0\nu\text{DBD}$ events while maintaining the spectral characteristics with measured energy resolution and introducing no discontinuities in the spectrum
- When all data analysis procedures are fixed the data are eventually unblinded



Fit in the ROI

- Simultaneous **UEML** (Unbinned Extended Maximum Likelihood) fit in the energy region **2465-2575 keV**
- The fit has 3 components:
 1. **posited peak at the Q-value** of ^{130}Te :
 - energy scale defined relative to the ^{208}Tl line in calibration data to account for residual mis-calibration between channels
 - signal normalisation common to all detectors-datasets (1 free parameter)
 2. **floating peak to account for the ^{60}Co sum gamma line** (2505 keV):
 - energy scale defined relative to the ^{208}Tl line in calibration data to account for residual mis-calibration between channels
 - rate common to all detectors-dataset, with a correction accounting for the time elapsed between the two datasets (1 free parameter)
 3. **flat background**, attributed to multi scatter Compton events from ^{208}Tl and surface alpha events:
 - common to all detectors in a single dataset, two independent parameters for the two datasets to account for differences in the background rejection efficiency (2 free parameters)
- The peaks in each channel-dataset are fitted with its own line shape (fixed from calibration data)

Fit in the ROI: results

Region of interest: **2465 to 2575 keV**

Overall signal efficiency: **$(75.7 \pm 3.0)\%$ - ds3018**

$(83.0 \pm 2.6)\%$ - ds3021

Events in the region of interest: **155**

ROI background index: **$(1.49_{-0.17}^{+0.18}) \times 10^{-2} \text{ c}/(\text{keV} \cdot \text{kg} \cdot \text{yr})$**

$(1.35_{-0.18}^{+0.20}) \times 10^{-2} \text{ c}/(\text{keV} \cdot \text{kg} \cdot \text{yr})$

Best fit for ^{60}Co mean: **$(2506.4 \pm 1.2) \text{ keV}$**

Best fit decay rate: **$(-1.0_{-0.3}^{+0.4} \text{ (stat.)} \pm 0.1 \text{ (syst.)}) \times 10^{-25} / \text{yr}$**

No evidence of signal

Limit calculation

Profile likelihood integrated on the physical region ($\Gamma^{0\nu} > 0$)

Decay rate limit (90% CL, including systematics):

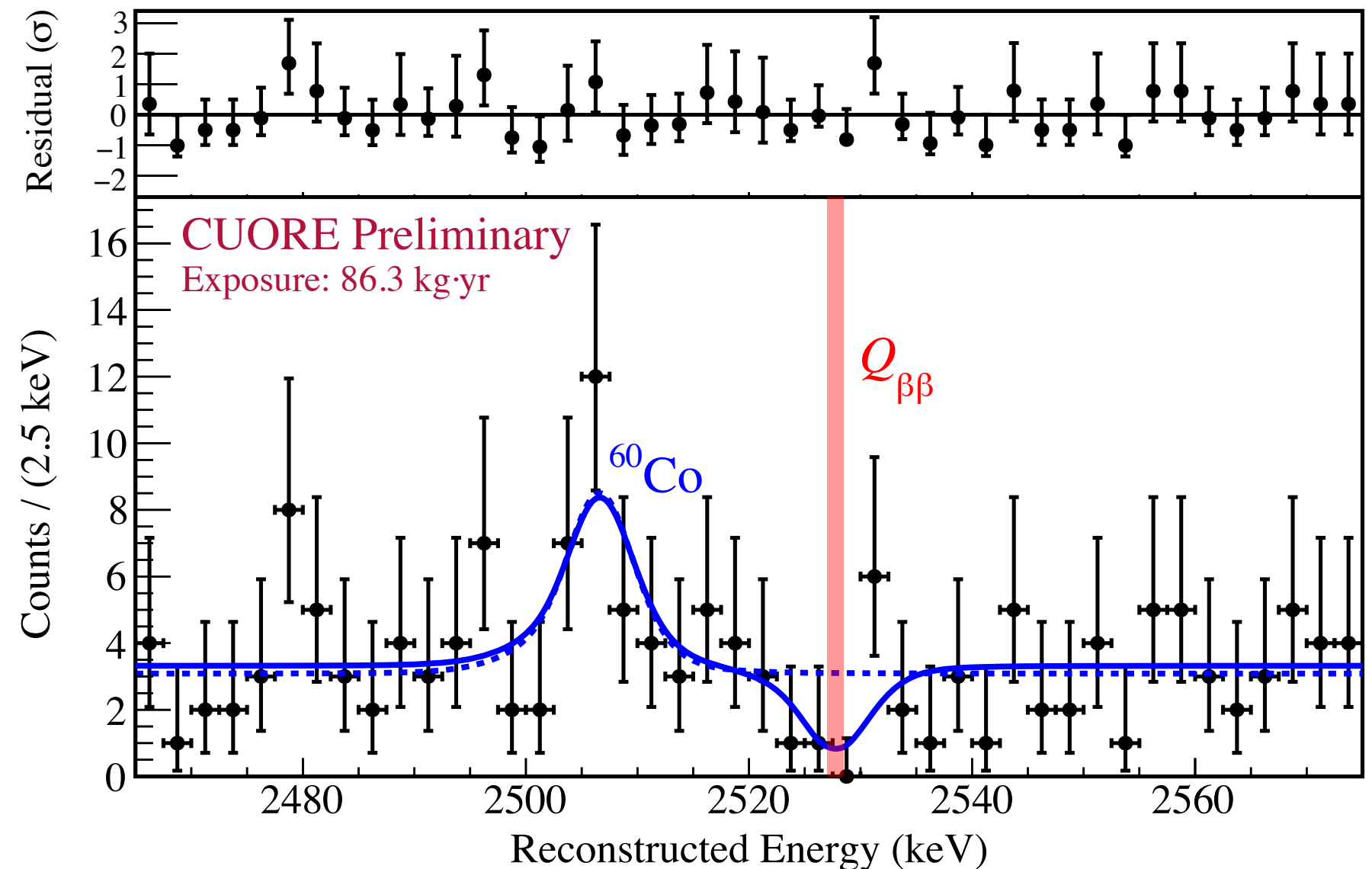
$0.51 \times 10^{-25} / \text{yr}$

Half-life limit (90% CL, including systematics):

$1.3 \times 10^{25} \text{ yr}$

Median expected sensitivity:

$7.0 \times 10^{24} \text{ yr}$



Combination with previous results

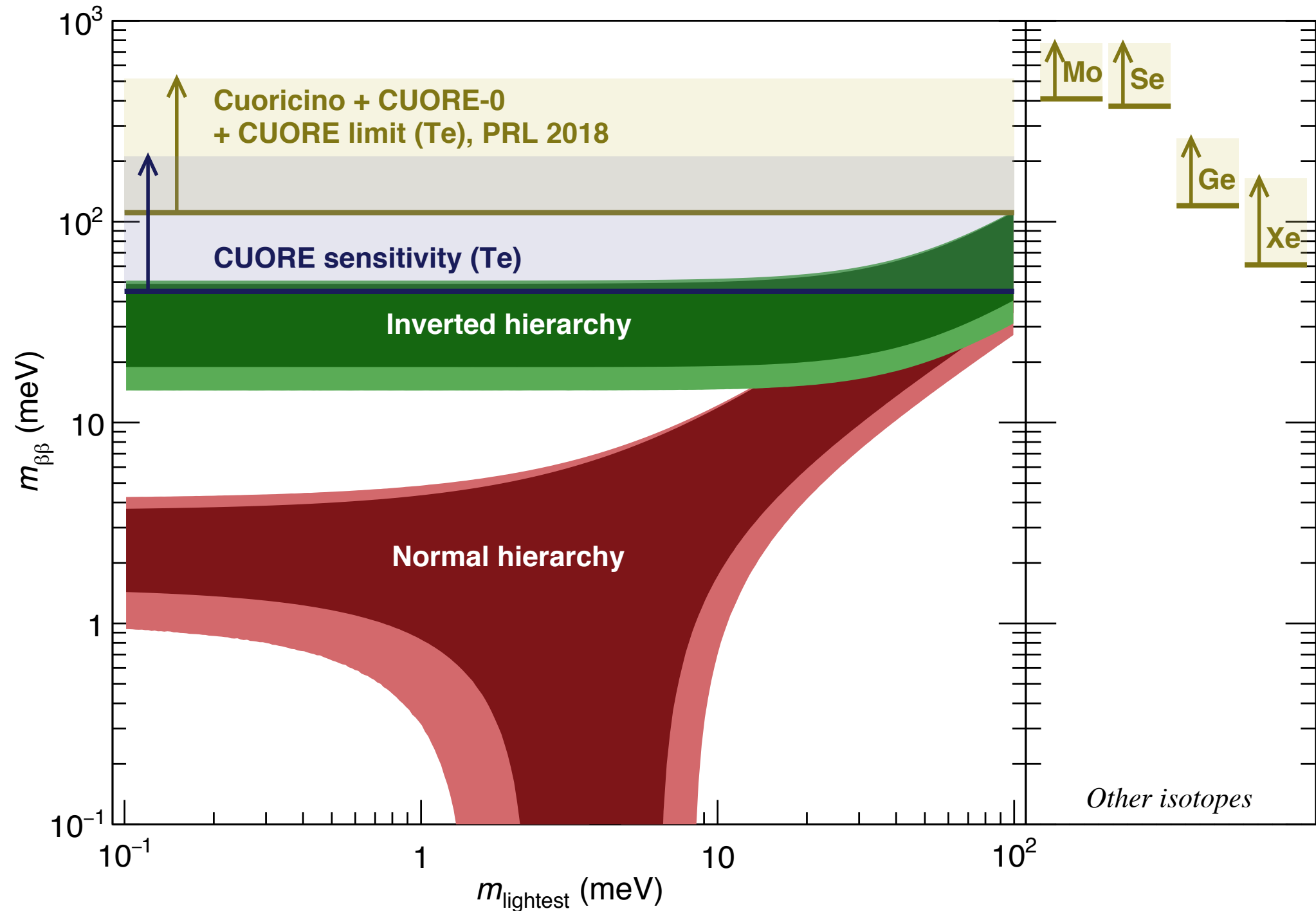
- We combined the CUORE result with the existing ^{130}Te data:
 - 19.75 kg·yr of Cuoricino
 - 9.8 kg·yr of CUORE-0
- The combined 90% C.L. limit is
 - $T_{0\nu} > 1.5 \times 10^{25}$ yr**
 - $m_{\beta\beta} < 110\text{--}520$ meV**

NME:

- JHEP02 (2013) 025
- Nucl. Phys. A 818, 139 (2009)
- Phys. Rev. C 87, 045501 (2013)
- Phys. Rev. C 87, 064302 (2014)
- Phys. Rev. C 91, 034304 (2015)
- Phys. Rev. C 91, 024613 (2015)
- Phys. Rev. C 91, 024309 (2015)
- Phys. Rev. C 91, 024316 (2015)
- Phys. Rev. Lett. 105, 252503 (2010)
- Phys. Rev. Lett. 111, 142501 (2013)

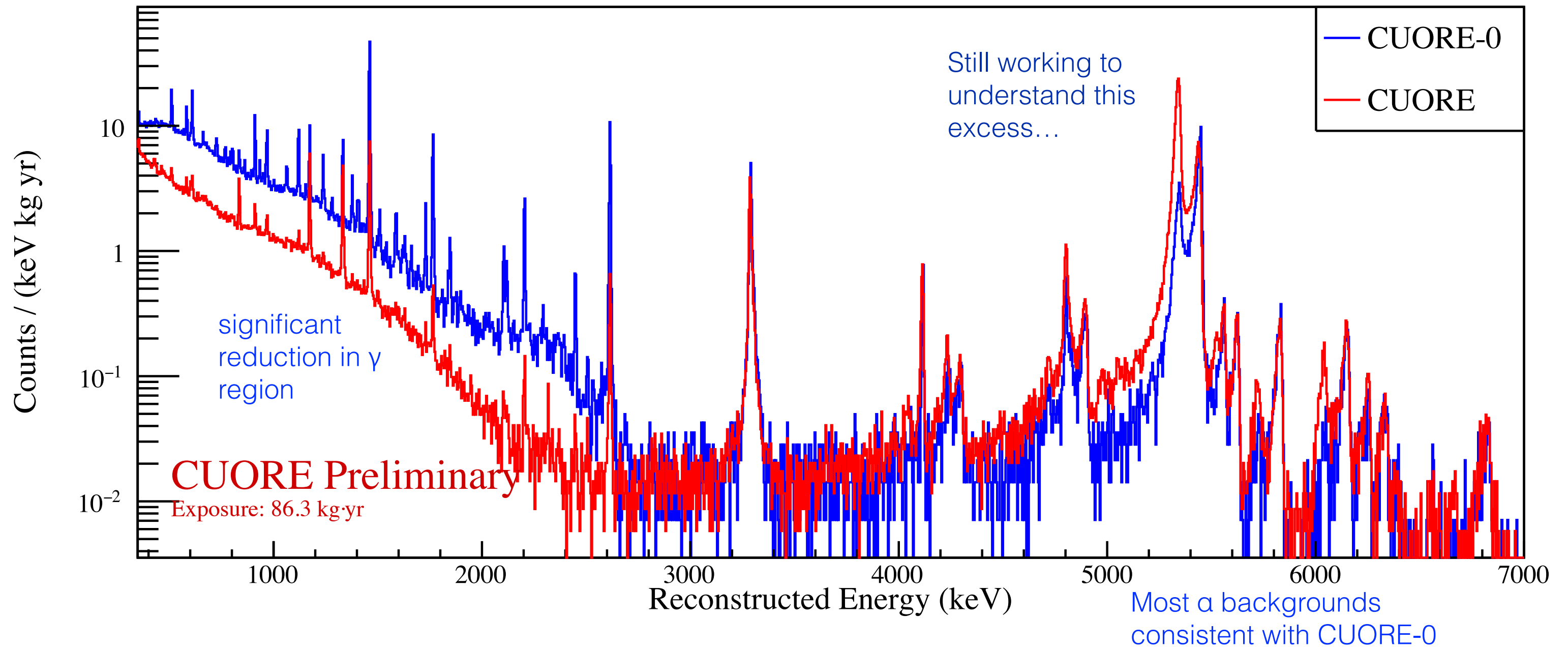
Experiments:

- ^{130}Te : 1.5×10^{25} yr from PRL 120, 132501 (2018)
- ^{76}Ge : 8.0×10^{25} yr from PRL 120, 132503 (2018)
- ^{136}Xe : 1.1×10^{26} yr from Phys. Rev. Lett. 117, 082503 (2016)
- ^{100}Mo : 1.1×10^{24} yr from Phys. Rev. D 89, 111101 (2014)
- ^{82}Se : 2.4×10^{24} yr from Phys. Rev. Lett. 120, 232502 (2018)
- CUORE sensitivity: 9.0×10^{25} yr



Background spectra

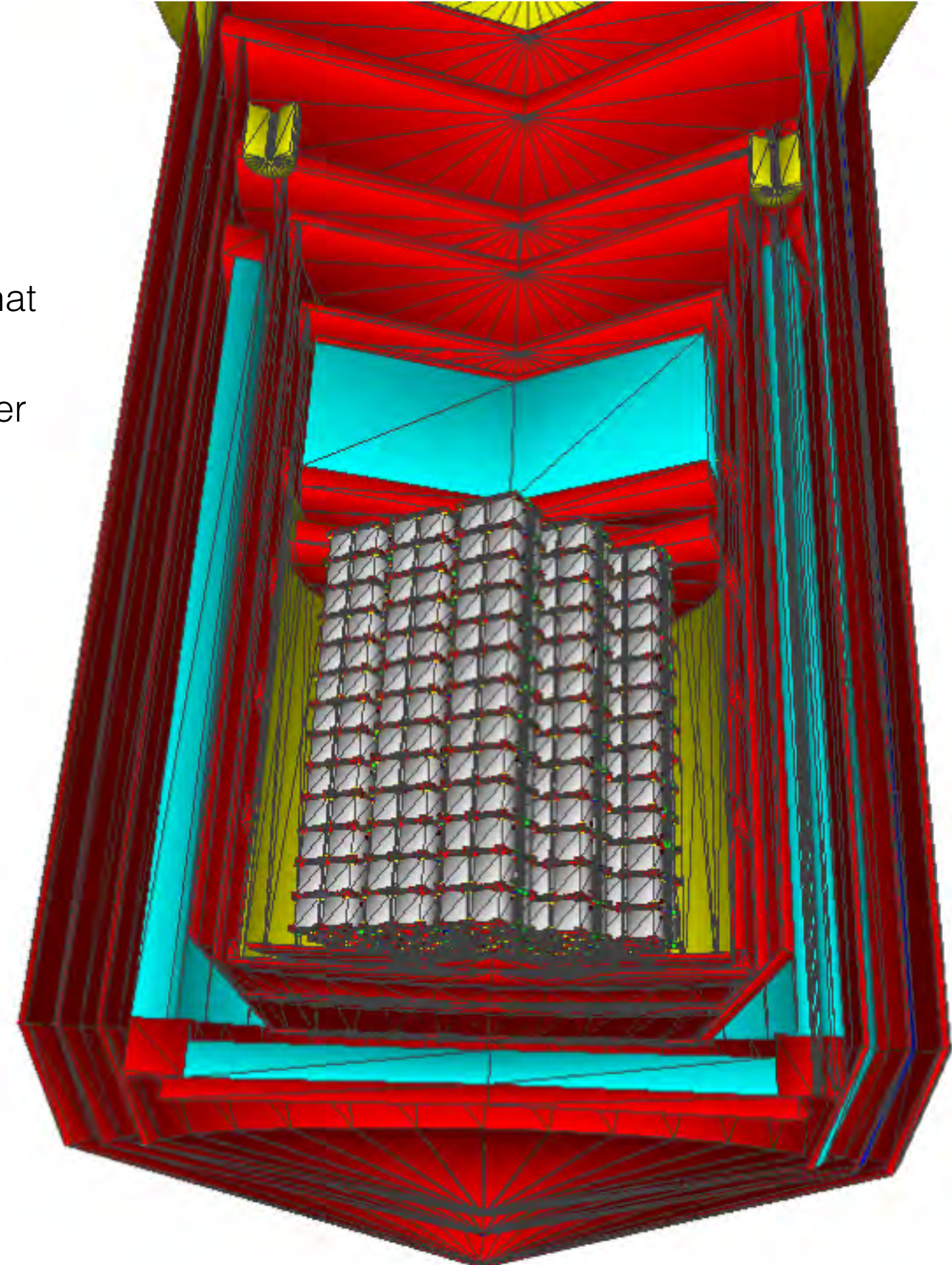
- Backgrounds generally consistent with expectations
- ^{210}Po excess appears to be from shallow contamination in copper around the detectors (estimated contribution to ROI at the level of $\sim 10^{-4}$ cnts/(keV kg yr))



Background model

- Simulate the contaminations coming from different cryostat components using a detailed Geant4 MC simulation
- ~60 independent parameters representing various contaminations that could contribute to the CUORE background model
- Perform a large Bayesian fit to the data using a MCMC Gibbs sampler
- Flat priors on all parameters

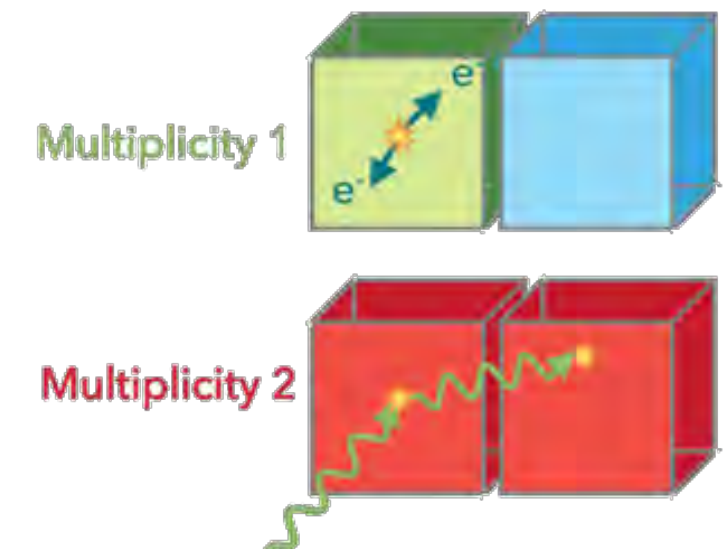
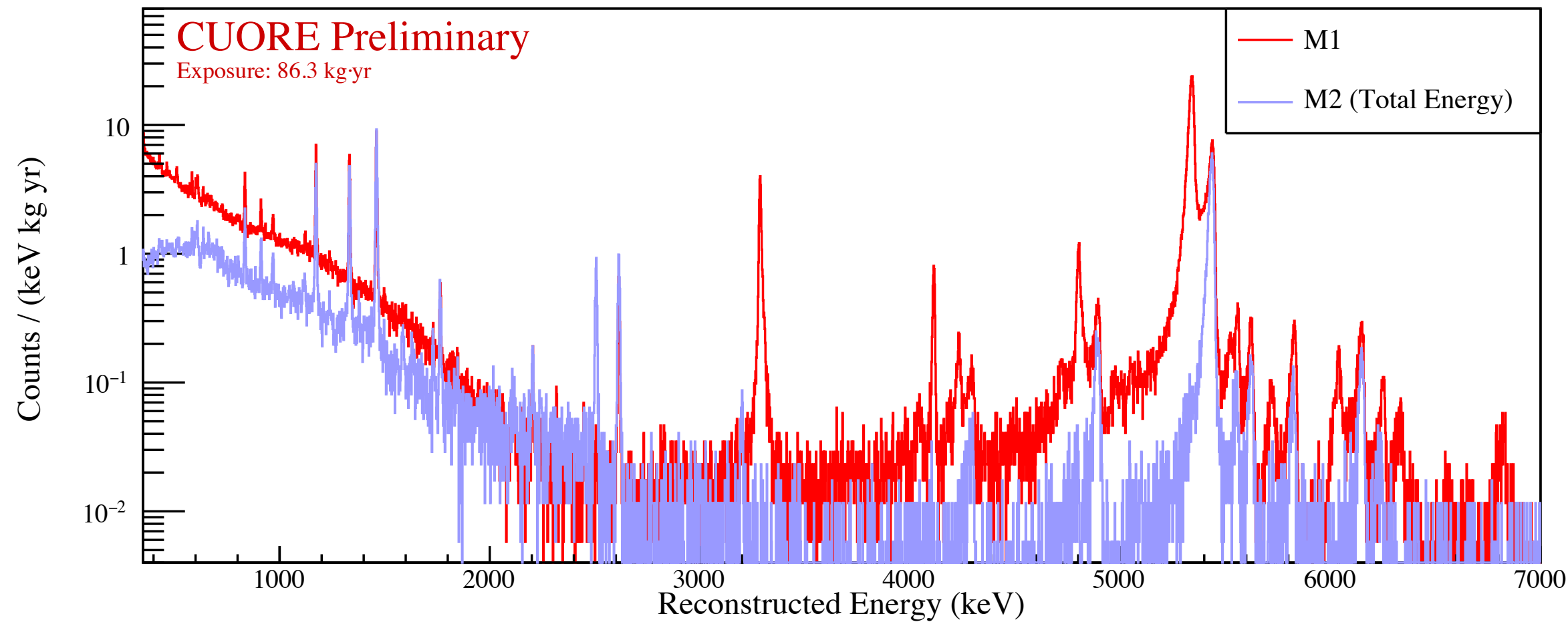
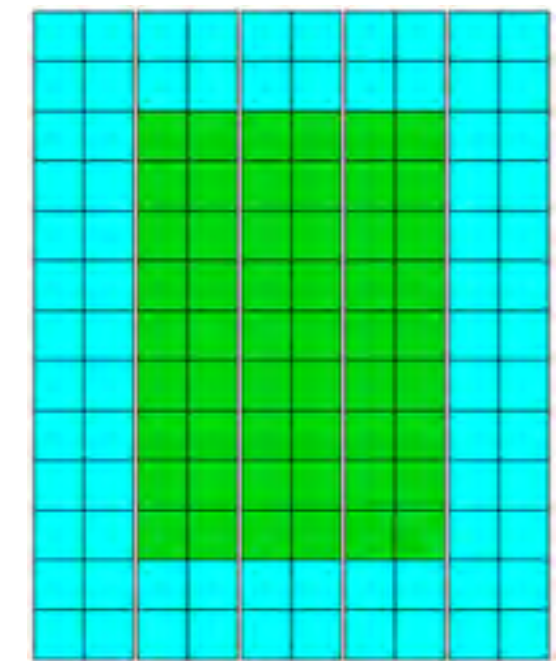
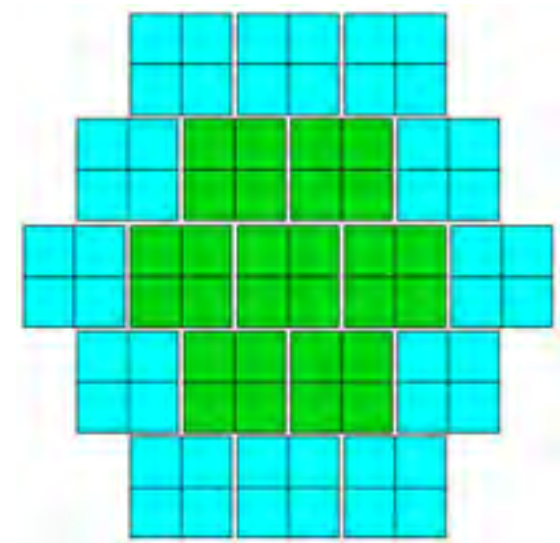
Volume	Type	Components
TeO ₂	Bulk	$2\nu\beta\beta$, ^{210}Pb , ^{232}Th , ^{228}Ra - ^{208}Pb , ^{238}U - ^{230}Th , ^{230}Th , ^{226}Ra - ^{210}Pb , ^{40}K , ^{60}Co , ^{125}Sb , ^{190}Pt
TeO ₂	Surface (0.01 μm)	^{232}Th , ^{228}Ra - ^{208}Pb , ^{238}U - ^{230}Th , ^{226}Ra - ^{210}Pb , ^{210}Pb
TeO ₂	Surface (1 μm)	^{210}Pb
TeO ₂	Surface (10 μm)	^{210}Pb , ^{232}Th , ^{238}U
CuNOSV	Bulk	^{232}Th , ^{238}U , ^{40}K , ^{60}Co , ^{54}Mn
CuNOSV	Surface (0.01 μm)	^{210}Pb , ^{232}Th , ^{238}U
CuNOSV	Surface (1 μm)	^{210}Pb , ^{232}Th , ^{238}U
CuNOSV	Surface (10 μm)	^{210}Pb , ^{232}Th , ^{238}U
Roman lead	Bulk	^{232}Th , ^{238}U , $^{108\text{m}}\text{Ag}$
Top lead	Bulk	^{232}Th , ^{238}U , ^{210}Bi
Ext. lead	Bulk	^{210}Bi
CuOFE	Bulk	^{232}Th , ^{238}U , ^{60}Co
External	-	Cosmic muons



Background model

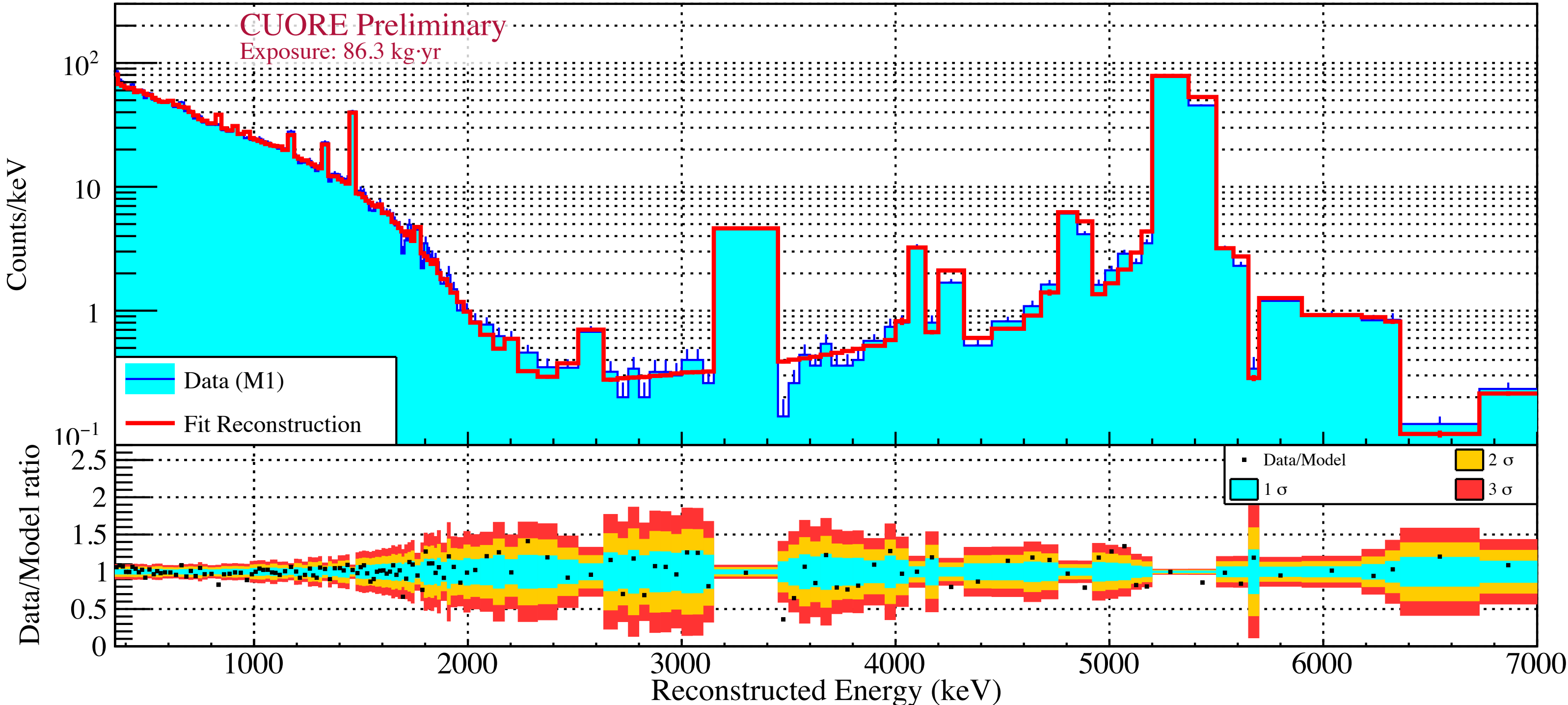
86.3 kg·yr of TeO₂ from summer 2017

- split data into inner and outer layers
- split data into Multiplicity 1 (M1), Multiplicity 2 (M2), and Multiplicity 2 Sum ($\Sigma 2$) spectra
 - higher multiplicity spectra sensitive to different backgrounds



Background model

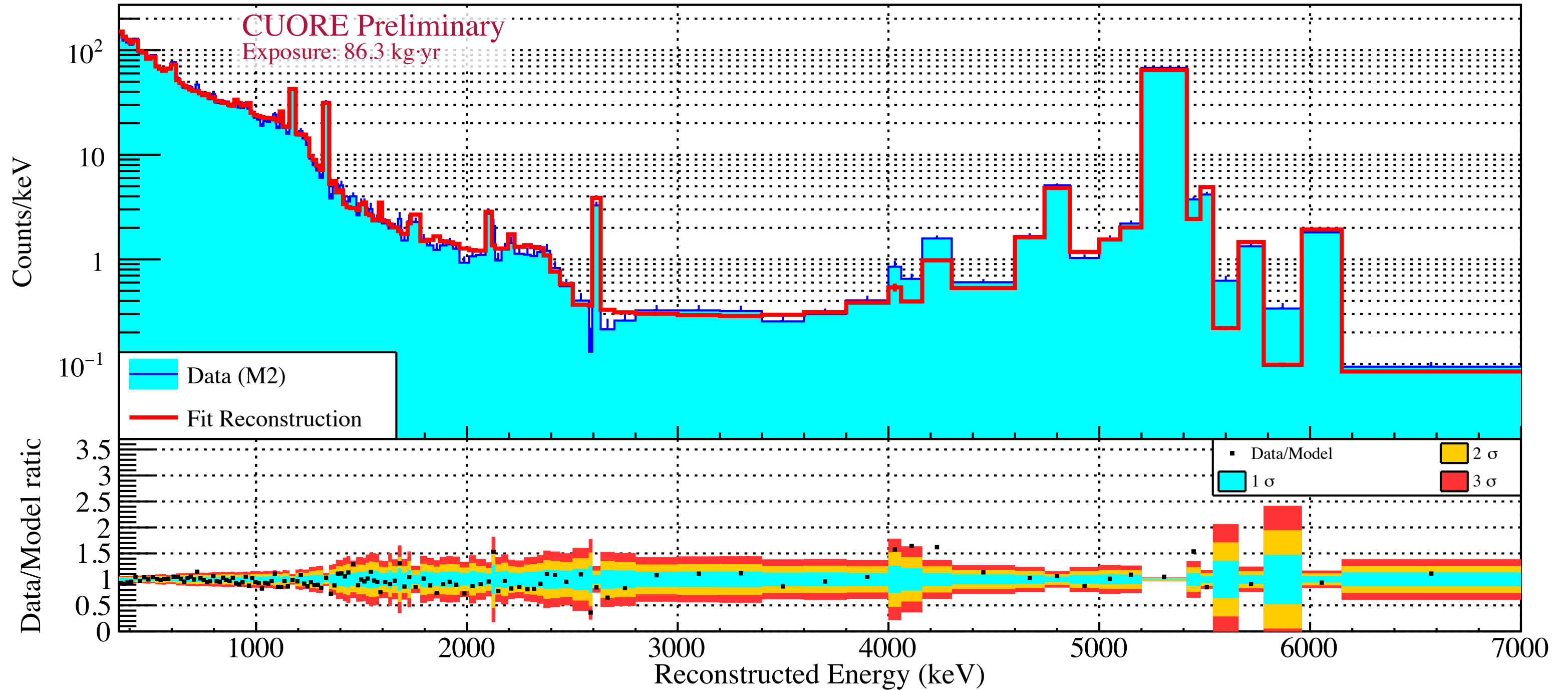
Multiplicity 1 - Inner Layer



Able to reconstruct the major features of the observed spectrum in CUORE

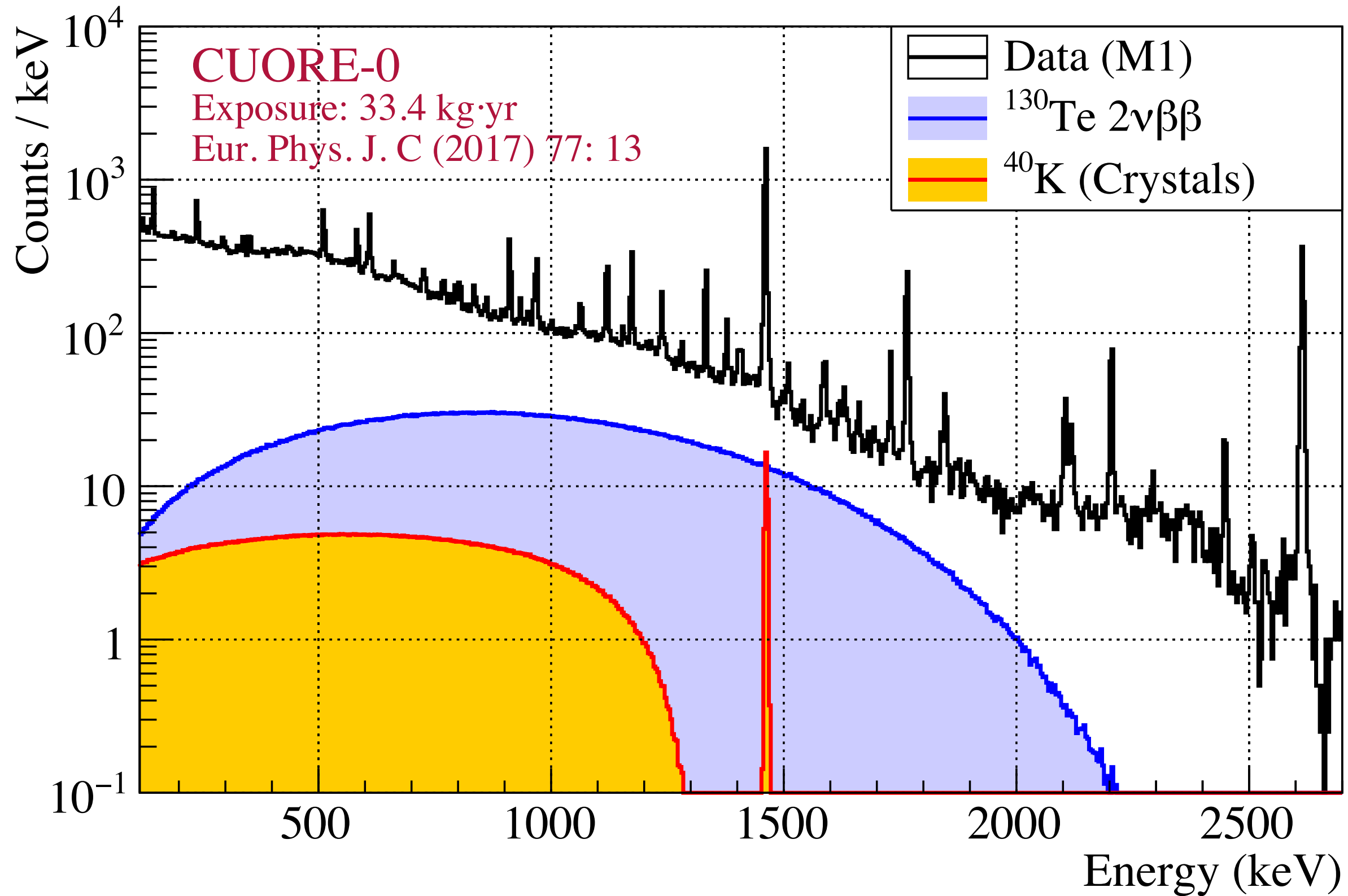
Background model

Multiplicity 2



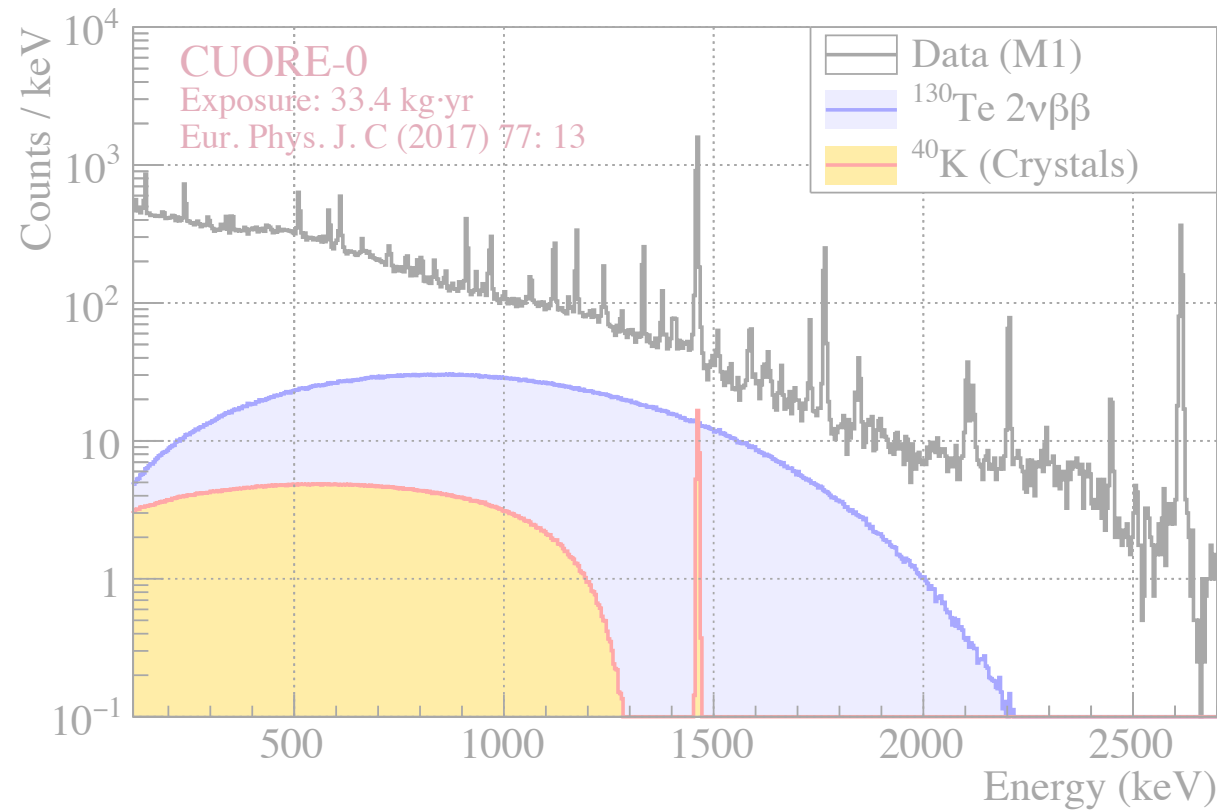
Many contaminations constrained by the higher multiplicity spectra

$\beta\beta 2\nu$ half-life measurement: CUORE-0



In CUORE-0, $2\nu\beta\beta$ decay spectrum accounts for ~20% of the signal in the range 1 - 2 MeV

$\beta\beta 2\nu$ half-life measurement: CUORE



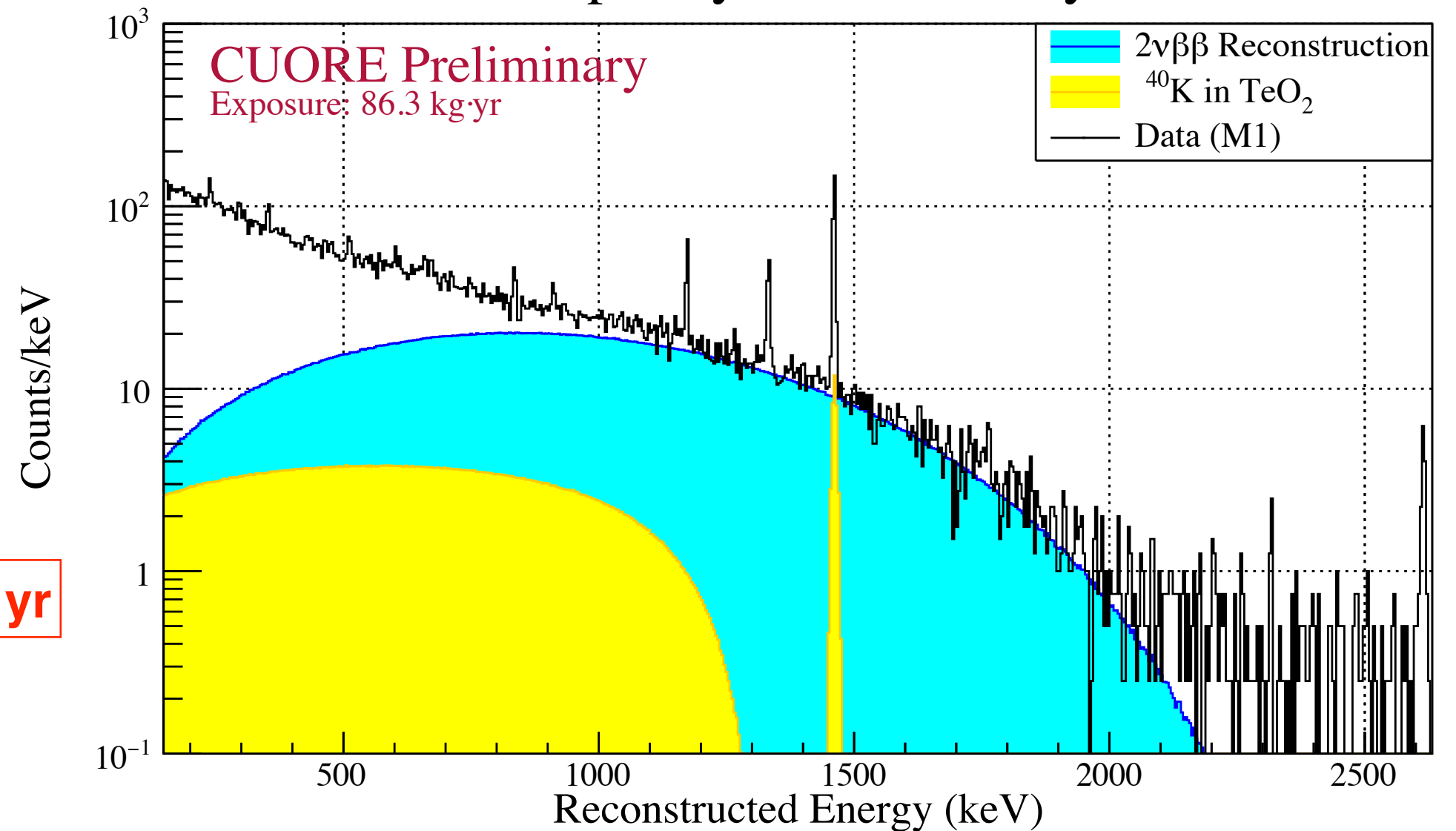
In CUORE-0, $2\nu\beta\beta$ decay spectrum accounts for ~20% of the signal in the range 1 - 2 MeV

In CUORE, $2\nu\beta\beta$ decay spectrum accounts for nearly all of the signal in the range 1 - 2 MeV

$$T_{1/2} = [7.9 \pm 0.1 \text{ (stat.)} \pm 0.2 \text{ (syst.)}] \rightarrow 10^{20} \text{ yr}$$

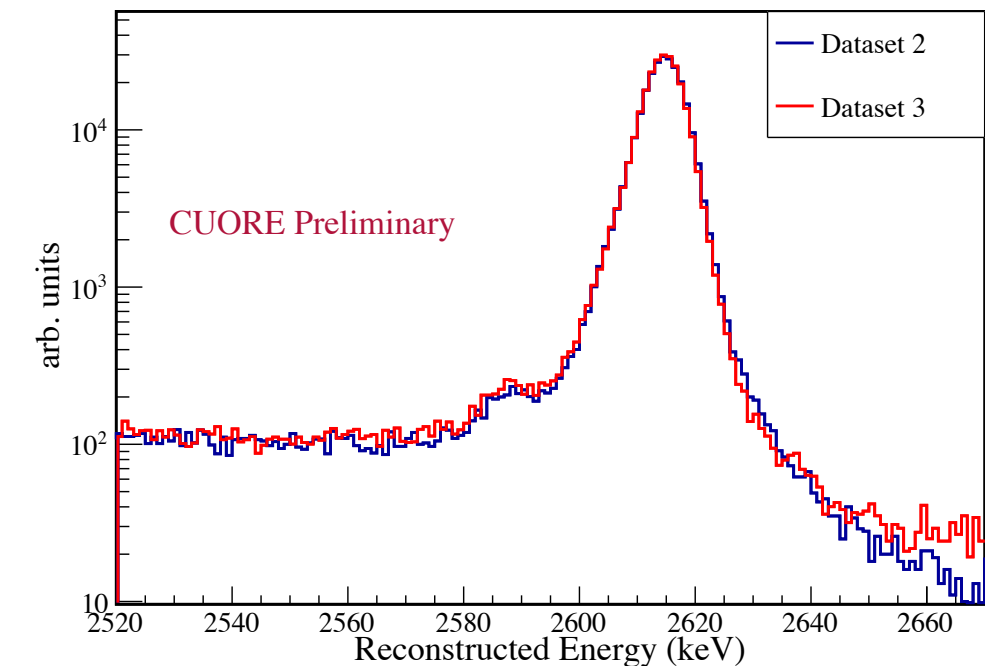
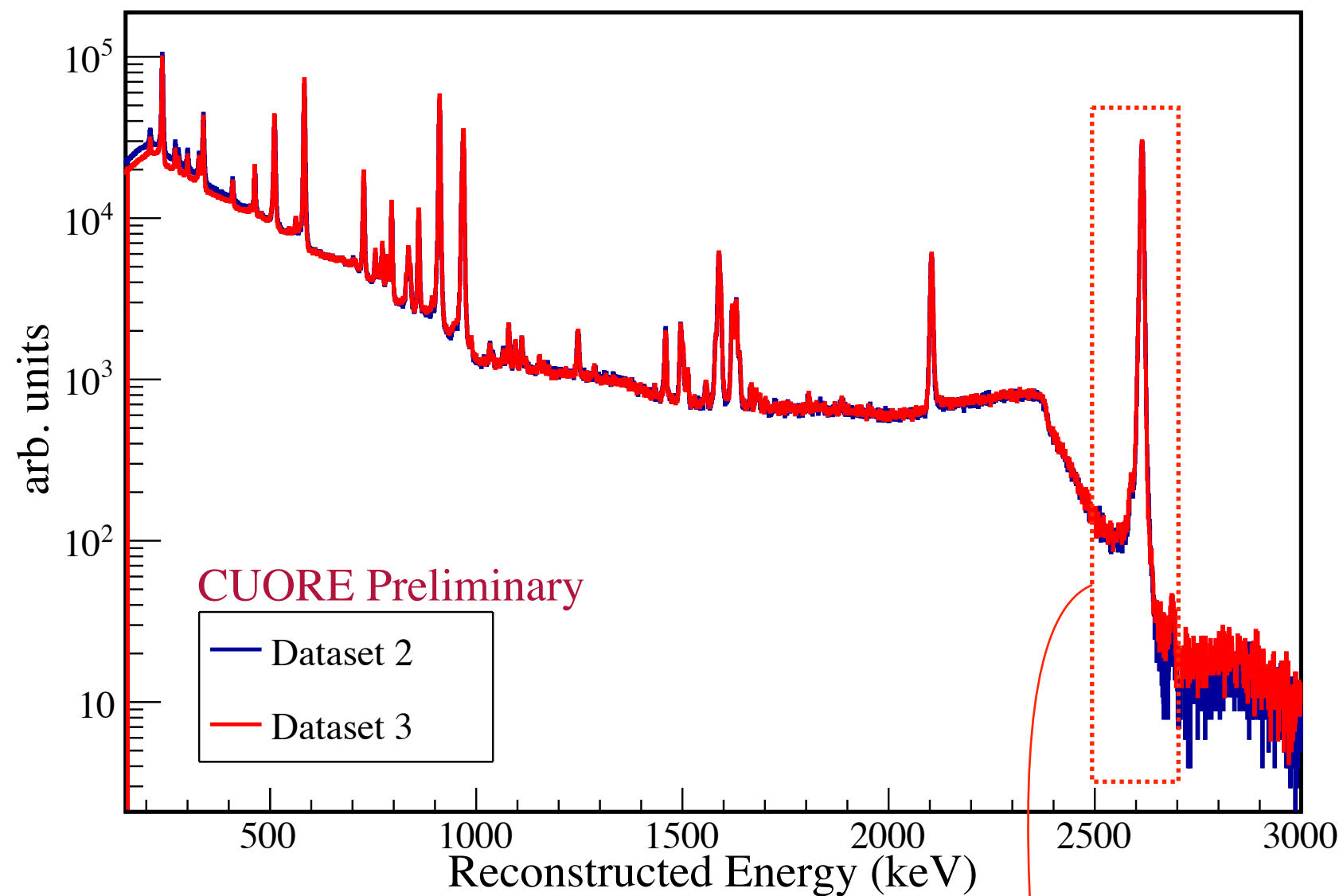
CUORE-0 : $[8.2 \pm 0.2 \text{ (stat.)} \pm 0.6 \text{ (syst.)}] \rightarrow 10^{20} \text{ yr}$
 NEMO-3 : $[7.0 \pm 0.9 \text{ (stat.)} \pm 1.1 \text{ (syst.)}] \rightarrow 10^{20} \text{ yr}$

Multiplicity 1 -- Inner Layer



2018 science runs

- Operating temperature: 11 mK
- April calibration data characterized by energy resolution of 7.6 keV FWHM with 93% of channels passing cuts (using same processing procedures)
- Still working to achieve the energy resolution goal of 5 keV FWHM
- Back to stable physics data taking in May 2018



Conclusions

- CUORE is the first ton-scale cryogenic detector array in operation, more than an order of magnitude larger than its predecessors
- The successful commissioning and operation of this large-mass, low-background cryogenic bolometer array represents a major advancement in the application of this technique to $\beta\beta 0\nu$ decay searches and demonstrates the feasibility of future large-mass bolometer arrays for rare-event searches.
- Thanks to the increased mass of the detector a number of physics processes can be studied with high precision
- With the first data of CUORE:
 - Set the most stringent limit on $0\nu\beta\beta$ half-life of ^{130}Te to date
 - Made the most precise measurement of $2\nu\beta\beta$ half-life of ^{130}Te
- After a period of detector optimization, data taking has restarted in May 2018
- CUORE will continue taking data in the coming years, with an ultimate sensitivity to $0\nu\beta\beta$ half-life in ^{130}Te of $T_{1/2}^{0\nu} > 9 \times 10^{25}$ yr

FUTURE PERSPECTIVES: CUPID

CUORE

CUORE is collecting data successfully

- 5 y projected half-life sensitivity: $\sim 10^{26}$ y ($m_{\beta\beta} < 50 - 190$ meV)
- Background according to expectations: $(1.4 \pm 0.2) \times 10^{-2}$ c/(keV · kg · yr)
- Energy resolution close to expectations: 7.6 keV FWHM (still margins for improvement)
- Analysis of ~ 1000 individual bolometers is handable

Two important messages from CUORE

1. A tonne-scale bolometric detector is feasible
2. An infrastructure to host a bolometric next-generation $0\nu\beta\beta$ experiment is already available

CUPID is the natural evolution of CUORE



CUPID

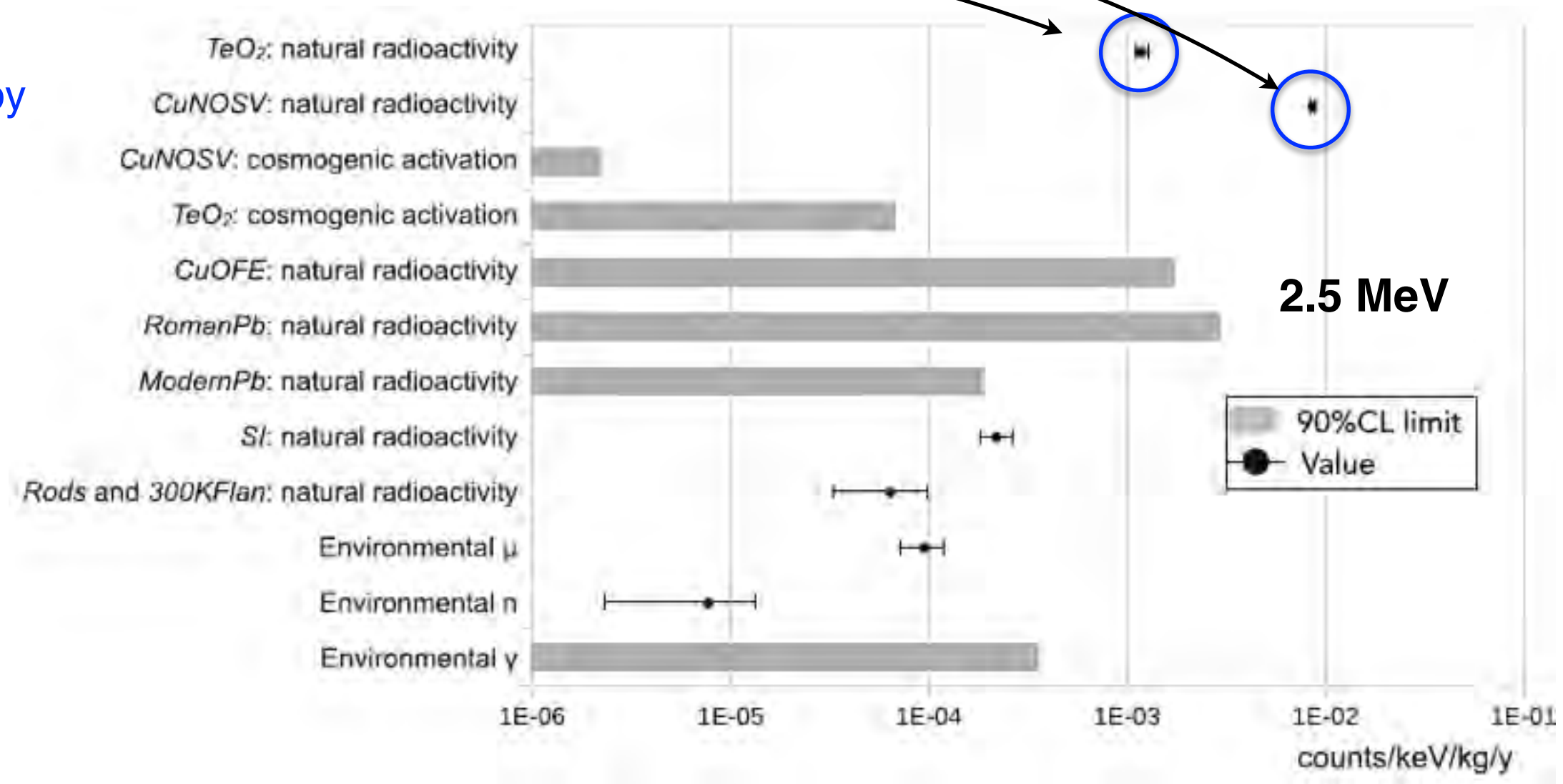
CUORE background presently dominated by alpha particles from surface contamination

Moving to CUPID requires:

1. Rejection of alphas
2. Control of the residual background

New detector technology:

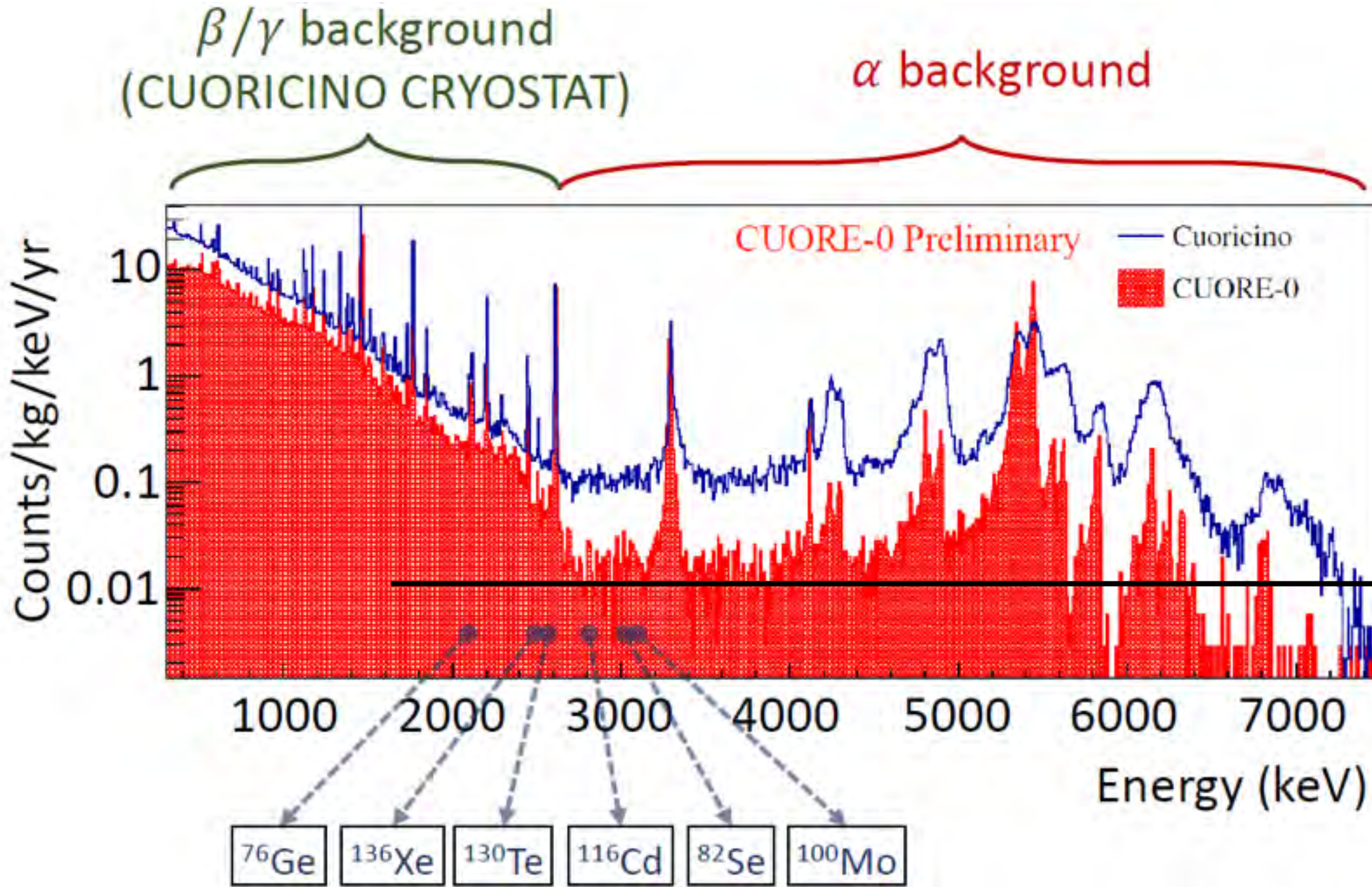
- luminescent bolometers
- Isotopic enrichment



R&D and demonstrators

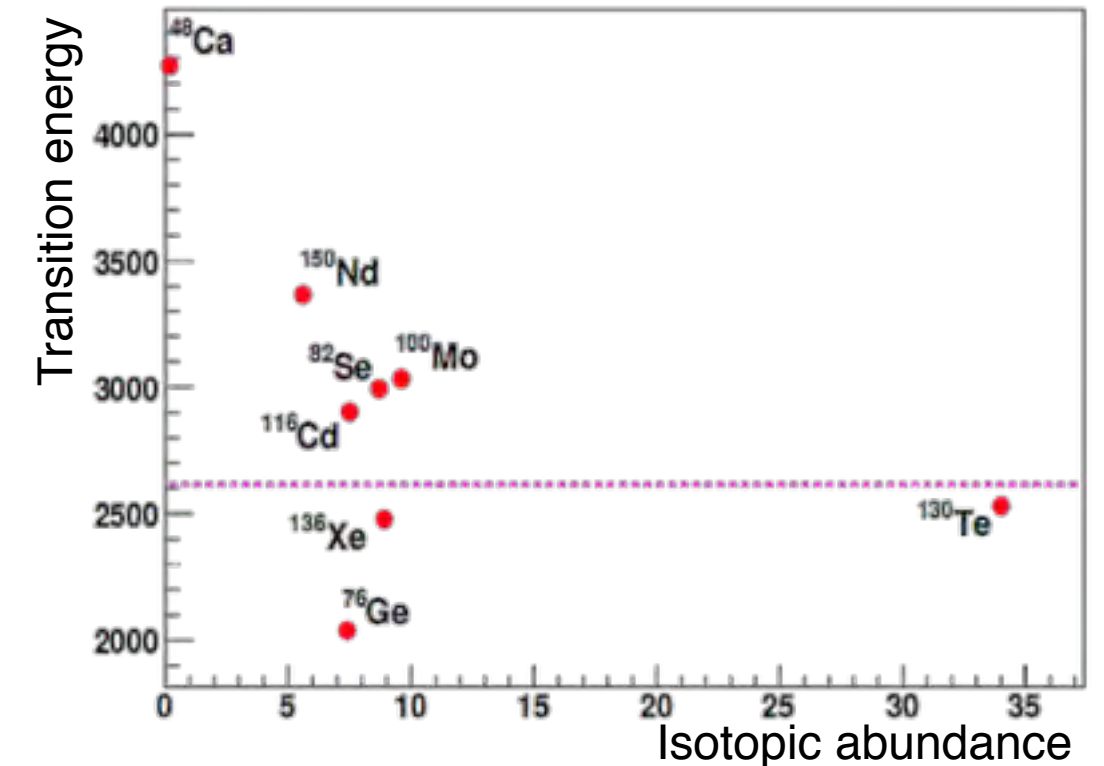
- Full CUORE background model + information from demonstrators
 - ¹³⁰Te in TeO₂ crystals and detection of Cherenkov light → prototypes
 - ⁸²Se in scintillating ZnSe crystals → demonstrator CUPID-0
 - ¹⁰⁰Mo in scintillating Li₂MoO₄ crystals → demonstrator CUPID-Mo

CUPID



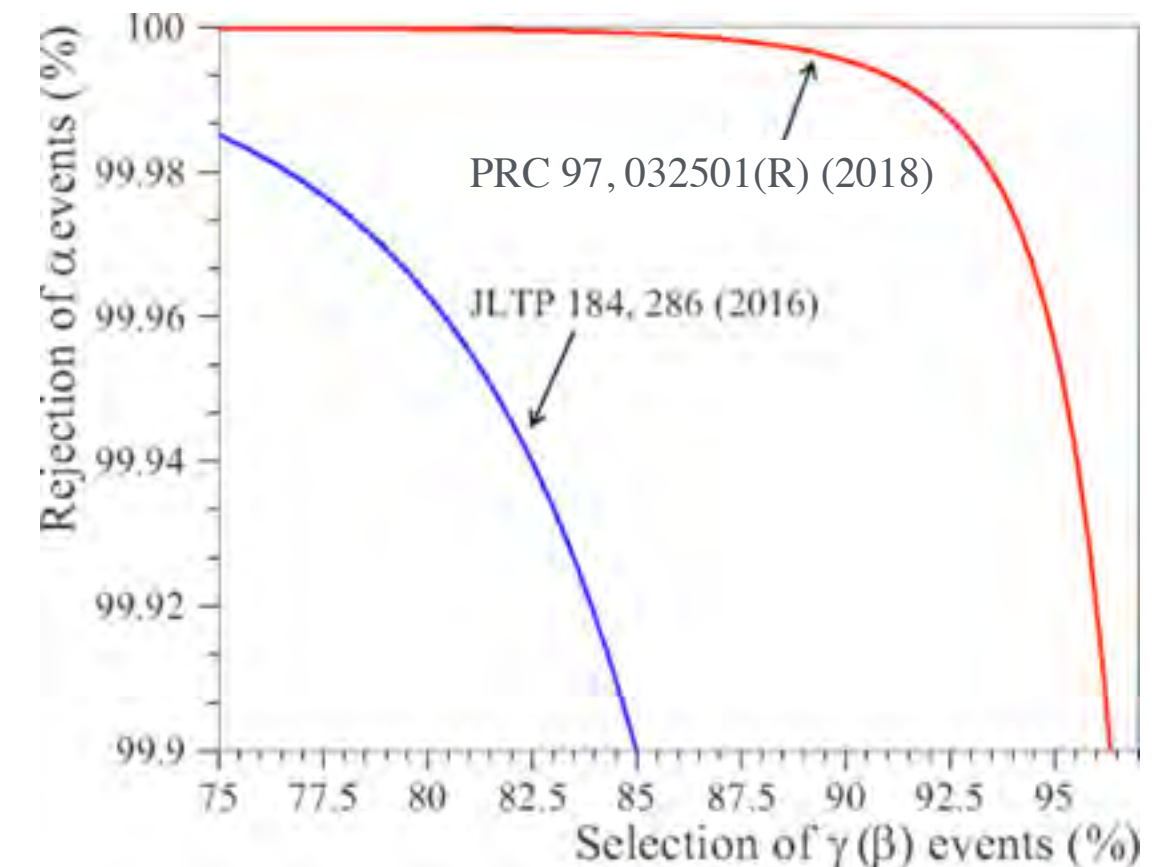
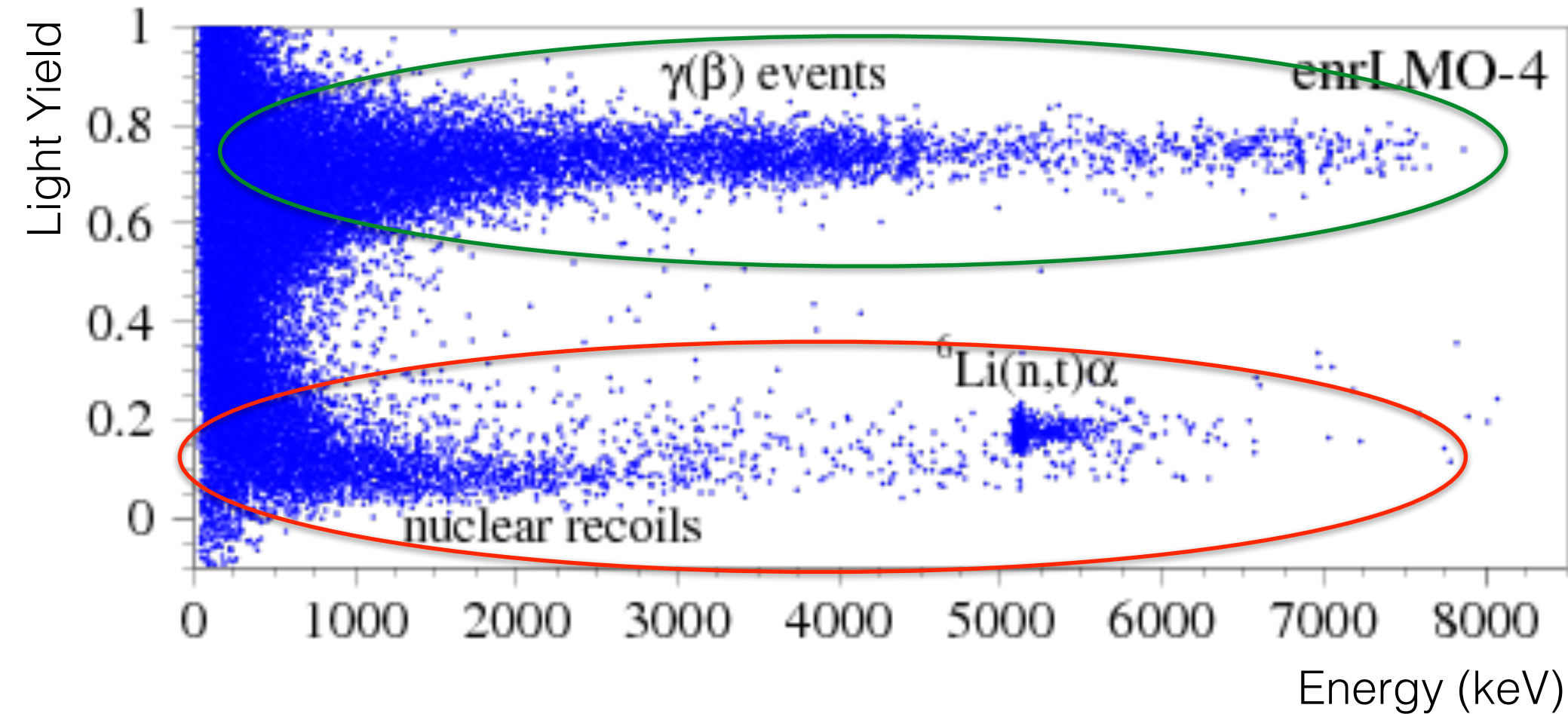
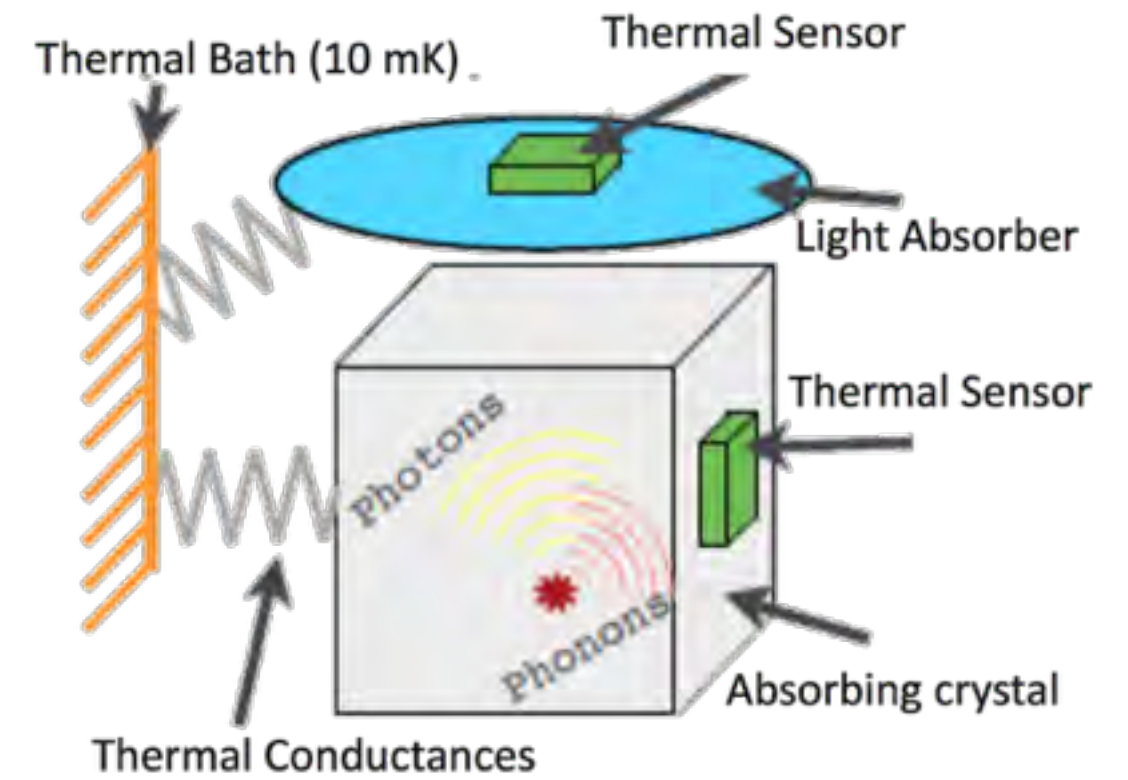
Energy-degraded alpha background
→ α/β discrimination is needed

A number of viable $\beta\beta$ isotopes
+ Background reduction
- Isotopic enrichment needed



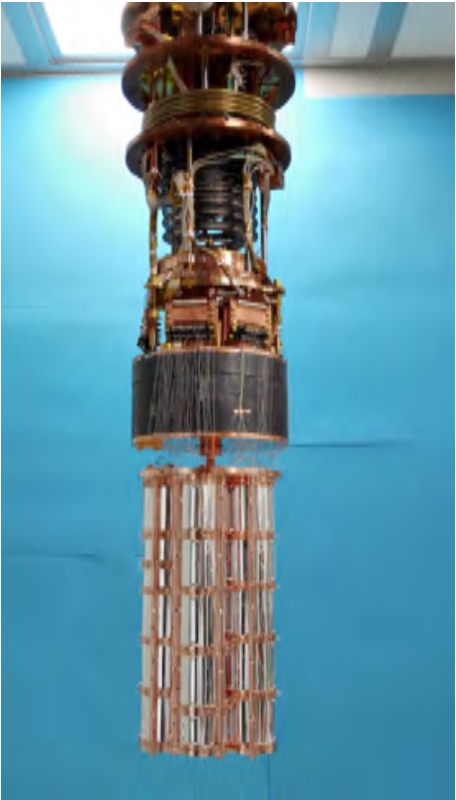
CUPID CUORE Upgrade with Particle ID

- Goal of reducing the background in the ROI by rejecting all α events with particle ID
- Add the ability to read out the light emitted in a particle interaction (scintillation/Cherenkov)
- Combine the energy resolution of bolometers with the background discrimination of a dual channel detector

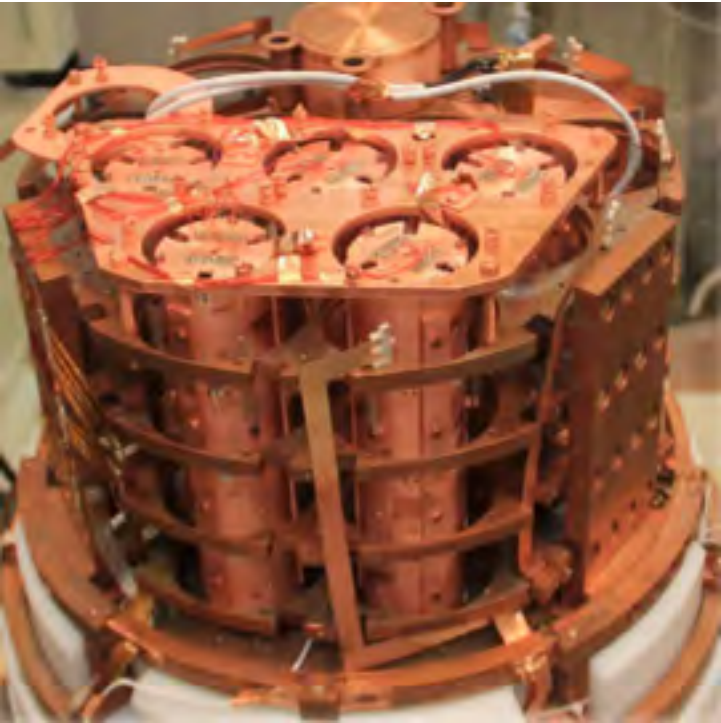


CUPID CUORE Upgrade with Particle ID

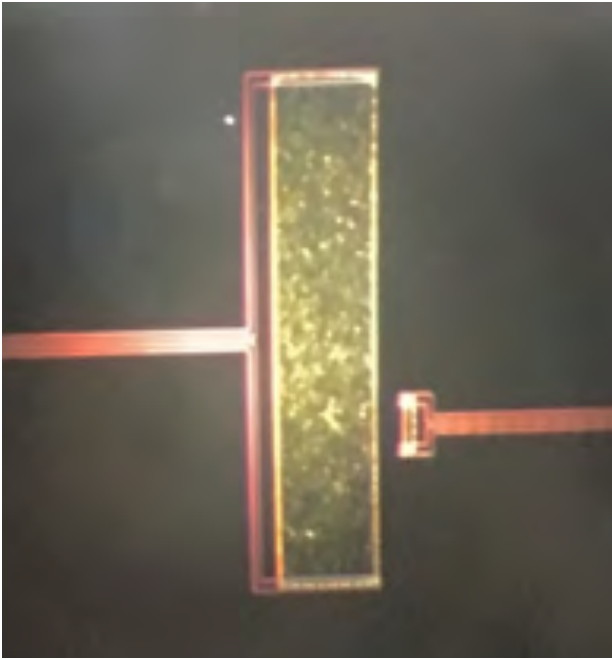
- Next generation of $0\nu\beta\beta$ decay experiments seek to be sensitive to the full IH region ($m\beta\beta \sim 6 - 20$ meV, $T_{1/2} \sim 10^{27}$ yr)
- ~ 1000 enriched light emitting bolometers mounted in the CUORE cryostat
- Nearly zero background goal of ~ 0.1 cnts/(ROI·yr)
- Worldwide effort focused on demonstrating readiness to construct a tonne-scale bolometric experiment



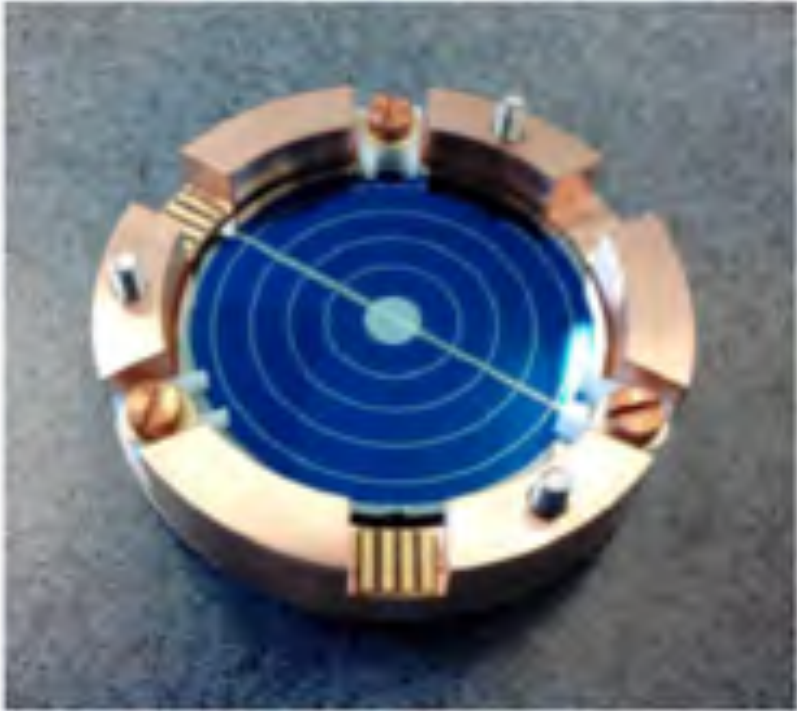
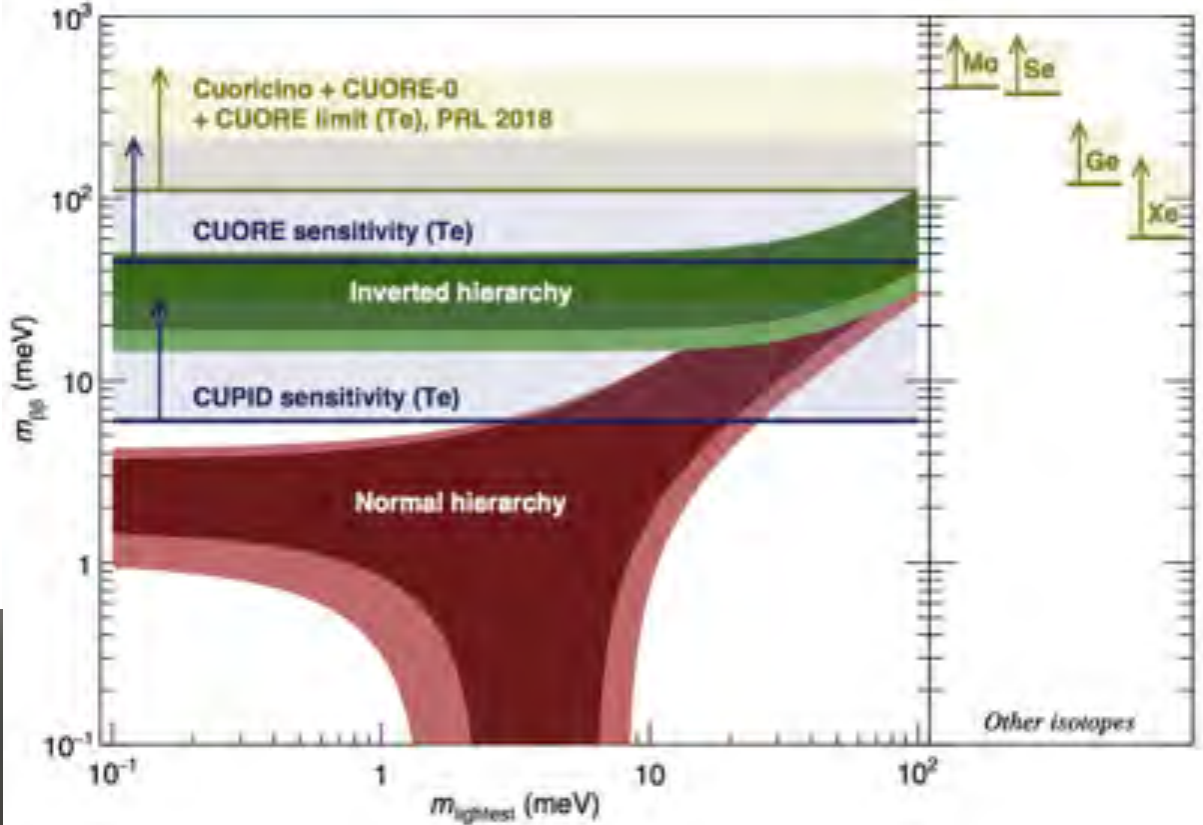
CUPID-0/Se at LNGS: $Zn^{82}Se$



CUPID-Mo at Modane: $Li_2^{100}MoO_4$



TES-based and Neganov-Luke light detectors for TeO_2



CUPID prospects

Results of the ongoing R&D and demonstrators + CUORE background model

- **$\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers** → **promising baseline option for CUPID**
- **$^{130}\text{TeO}_2$ Cherenkov bolometers** → **mature viable alternative**
- Fast and high-sensitivity light detectors are a common feature
 - Detection of Cherenkov light in TeO_2
 - Rejection of $2\nu\beta\beta$ random coincidences in $\text{Li}_2^{100}\text{MoO}_4$

The purpose of CUPID is to fully explore the IO region:

- Mission: half-life sensitivity higher than 10^{27} y
- With background < 0.1 counts/(ton·y) in the ROI:
 - ^{100}Mo sensitivity is $2.1 \cdot 10^{27}$ y
 - $m_{\beta\beta} < 6 - 17$ meV

CUPID ... moving forward

- CUORE will execute its scientific program to completion.
- CUORE's success motivates a next-generation bolometric experiment
- Based on the results of the ongoing R&D and demonstrator experiments, **Li₂MoO₄ scintillating bolometers have been identified as the most promising baseline** for a next-generation high-sensitivity bolometric experiment.
- Enriched **TeO₂ is a mature viable alternative**, with R&D towards demonstrating its sensitivity proceeding.
- Light detector technologies being developed for both TeO₂ and Li₂MoO₄ readout are common and R&D is beneficial to both.
- CUPID plan is under development
- We aim to form the CUPID collaboration in the near future:
 - CDR under preparation
 - kickoff meeting is being planned

