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Motivazioni teoriche per assioni leggeri, $(0.1 \div 1) \mu\text{eV}$

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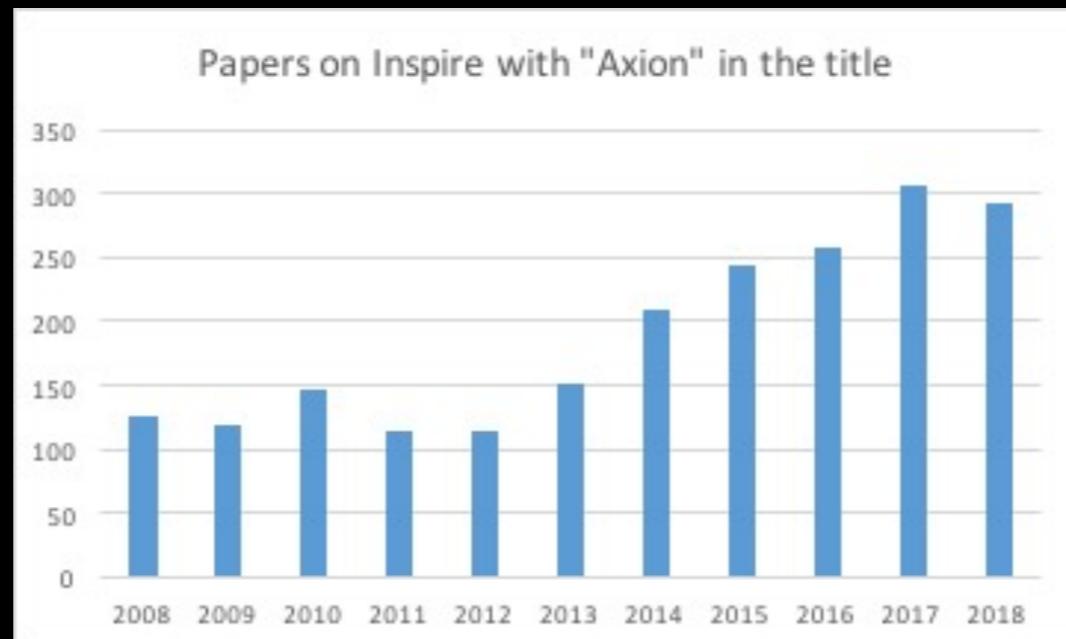
Based on:

LV, Redondo [1808.01879]

LV, Baum, Redondo, Freese, Wilczek, PLB **777**, 64 (2018) [1710.08910]

LV, Gondolo, PRL **113** 011802 (2014) [1403.4594]

Steady growth in the interest on the axion



PATRAS @ DESY 2009



PATRAS @ DESY 2018



Motivations

The value of $\bar{\theta}$ controls the matter-antimatter asymmetry in QCD

$$\mathcal{L}_{\text{strong,CP}} = \bar{\theta} \frac{\alpha_s}{2\pi} \text{Tr} (E^\mu B_\mu)$$

A similar term arises from EW, $\theta = \bar{\theta} + \theta_{\text{weak}} \sim \mathcal{O}(1)$

No observation of C and CP violation in Nature, $|\theta| \lesssim 10^{-10}$

Motivations

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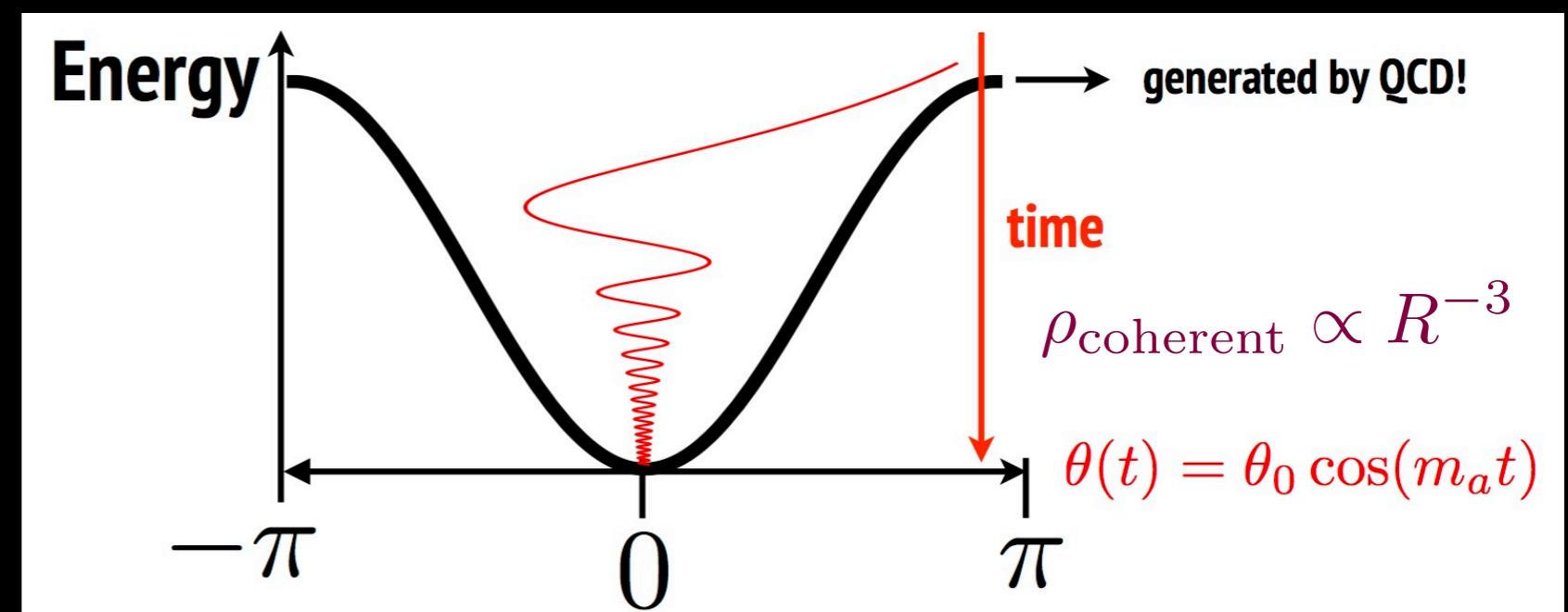
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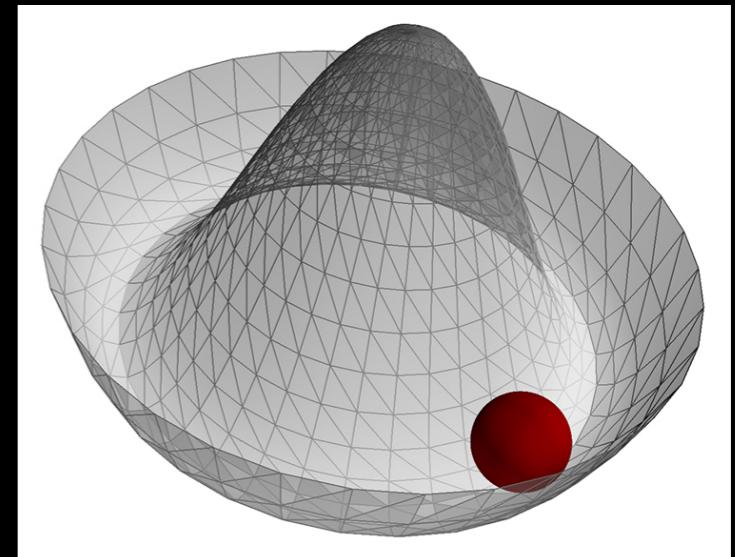
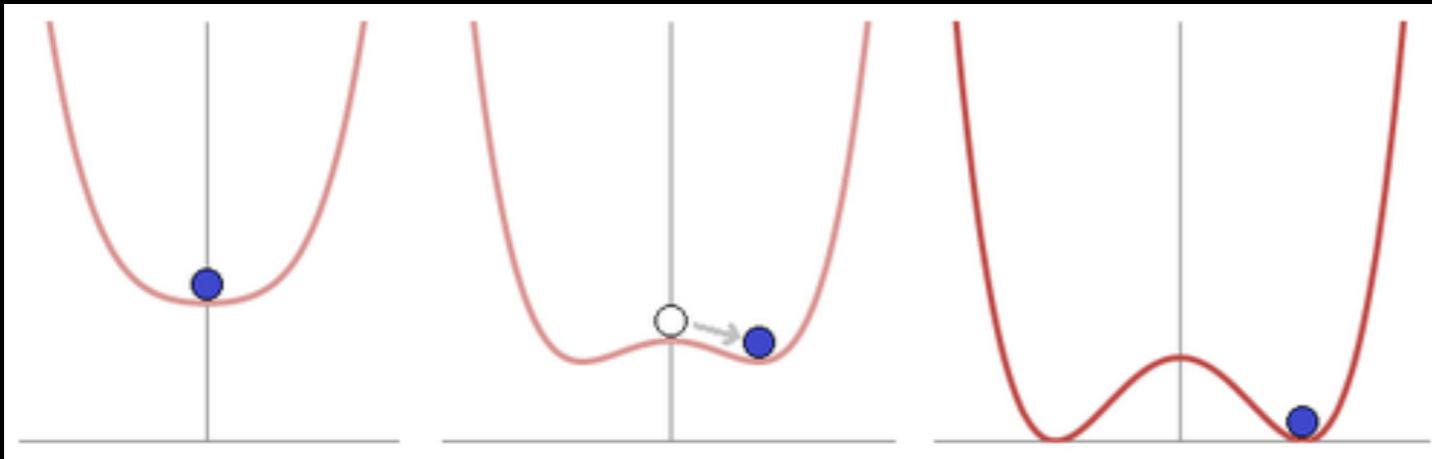
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Maybe it is a dynamical field?

$$\theta = \theta(t, \mathbf{x})$$



Early-Universe dynamics of the axion

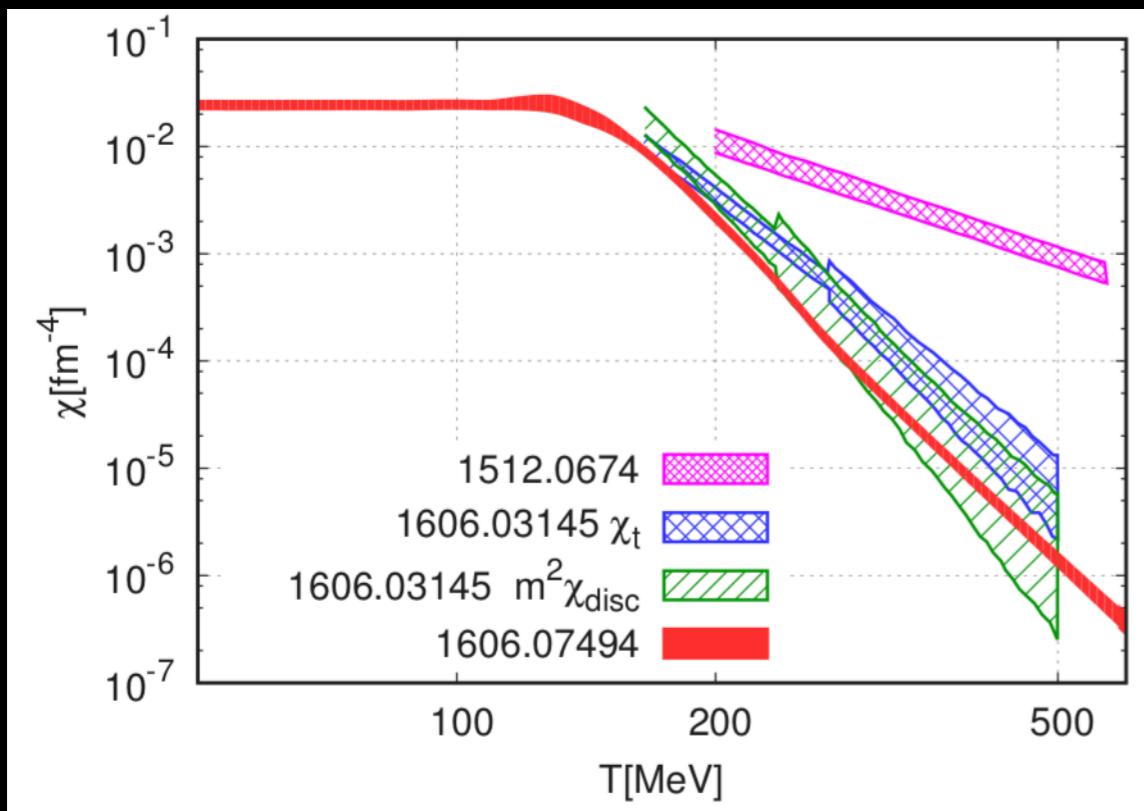


$$T \gtrsim f_a$$

$$T \sim f_a$$

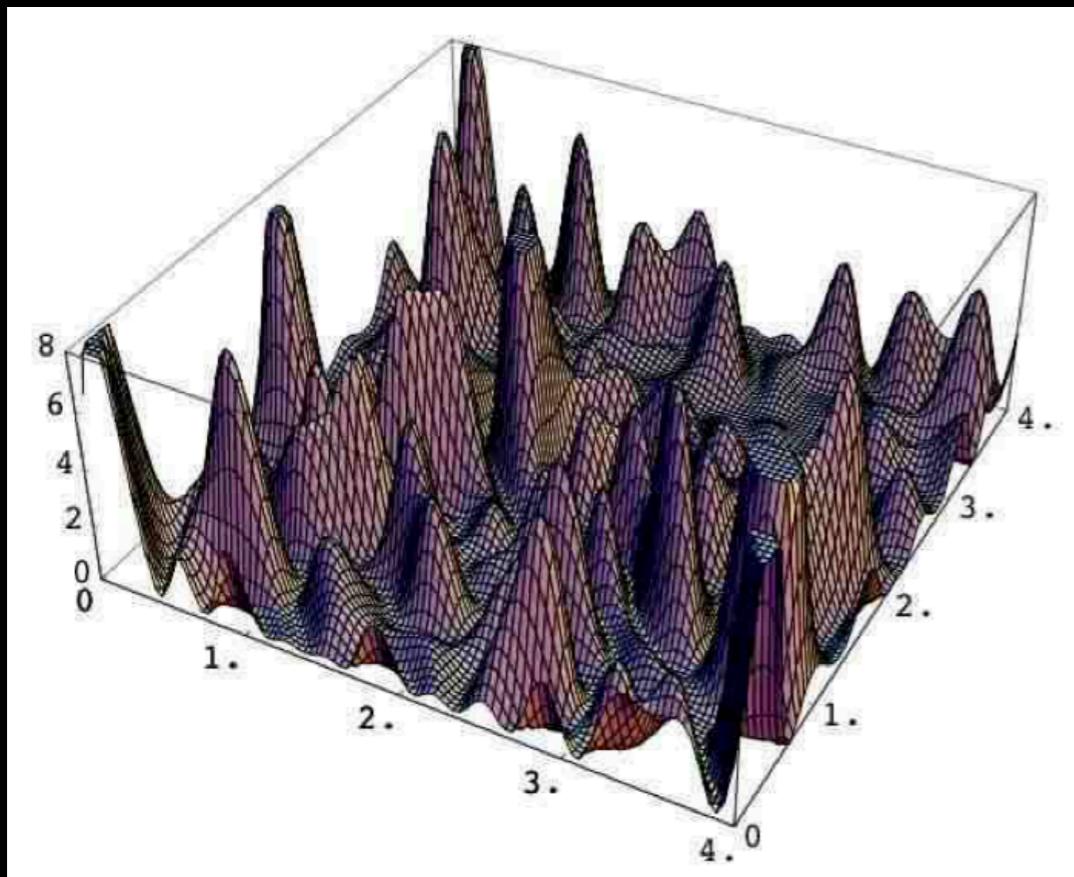
$$T \lesssim f_a$$

$$T \sim \Lambda_{\text{QCD}} \ll f_a$$



Details on the temperature dependence still debated
(semi-analytical,
lattice simulations)

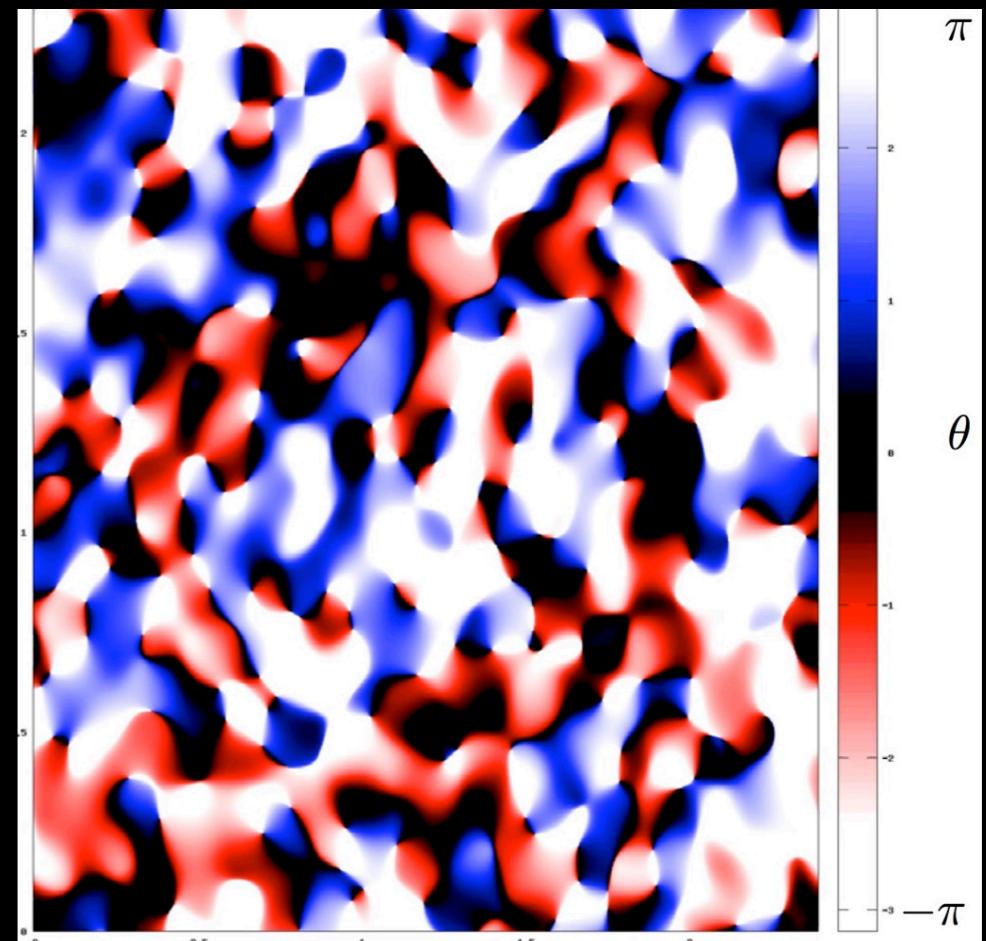
Post-inflation scenario



Random initial condition
→ we need a UV theory
Axion: angle of a complex scalar field

$$\mathcal{L} = \frac{1}{2} \left| \frac{d\phi}{dt} \right|^2 - \frac{1}{2a^2} |\nabla \phi|^2 - V(\phi)$$

$$\phi = |\phi| e^{i\theta}$$



$$T \sim f_a$$

The Kibble mechanism (Kibble76)
leads to string network formation

$$T \sim T_{\text{QCD}}$$

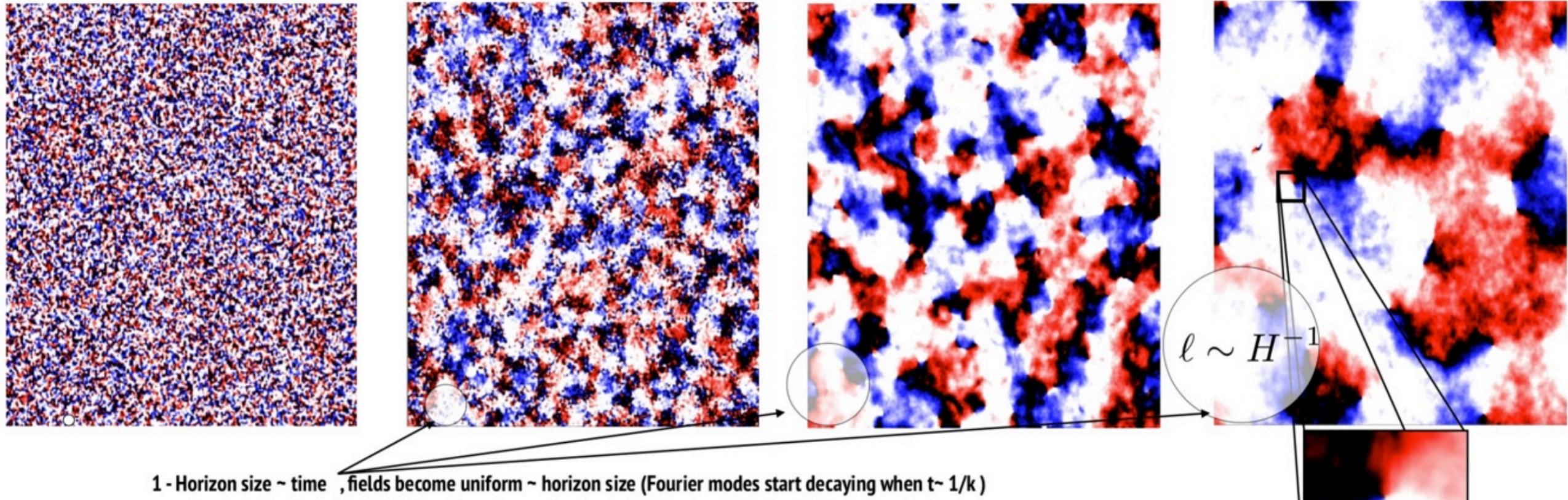
Domain walls form and dissipate
(Sikivie82; Georgi+82)

Destruction of the string-wall network

- Axion strings form by Kibble mechanism

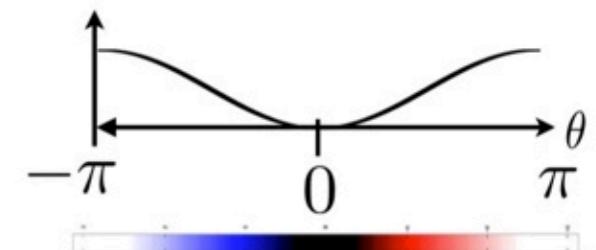
- Energy logarithmically distributed around, tension $\mu \simeq \pi f_A^2 \log\left(\frac{f_A}{H}\right)$

Courtesy of J. Redondo

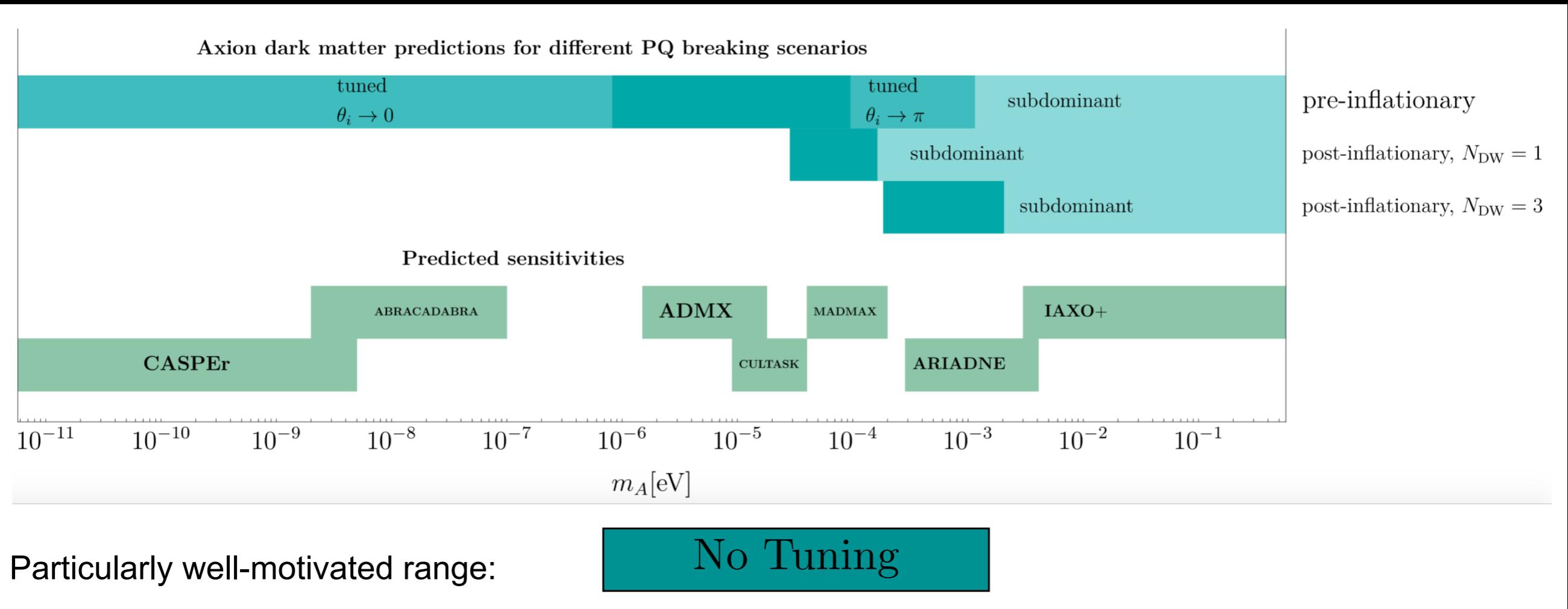


Computationally challenging:
all scales from the size of the box H_{QCD}^{-1}
to the string core f_a^{-1} have to be resolved!

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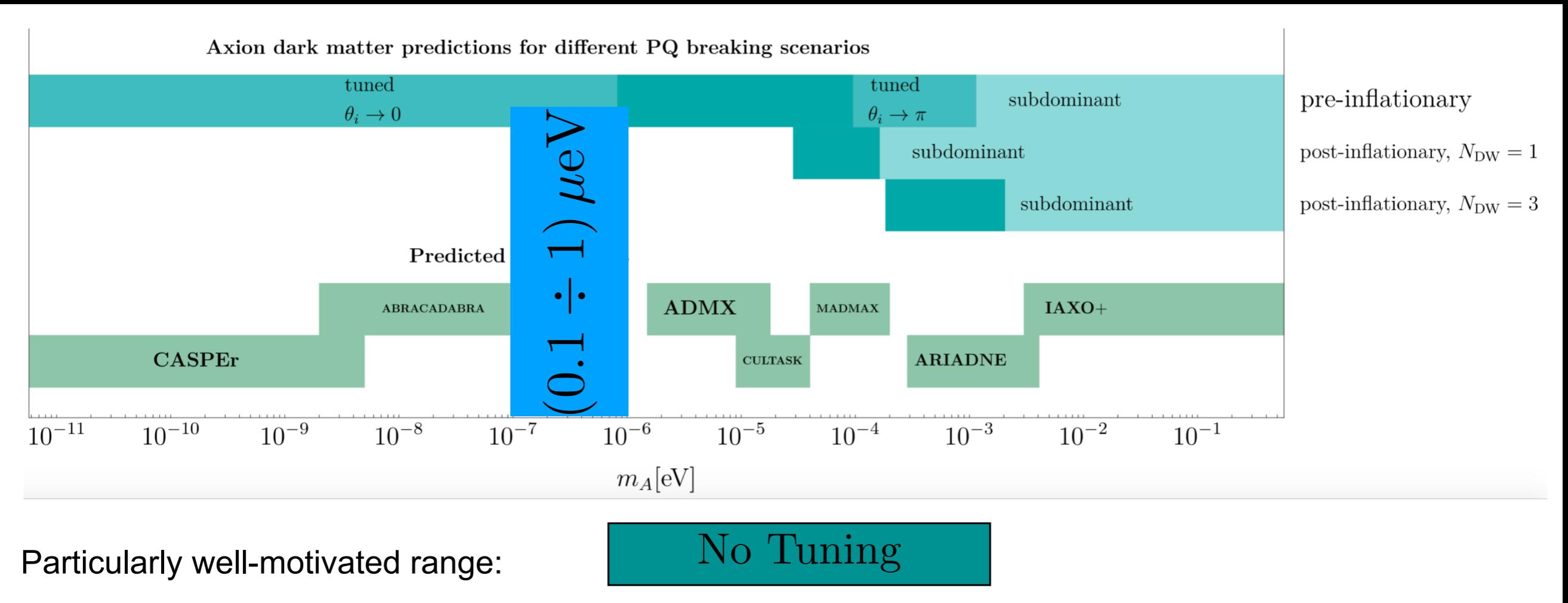


Less crowded than other mass ranges for the QCD axion



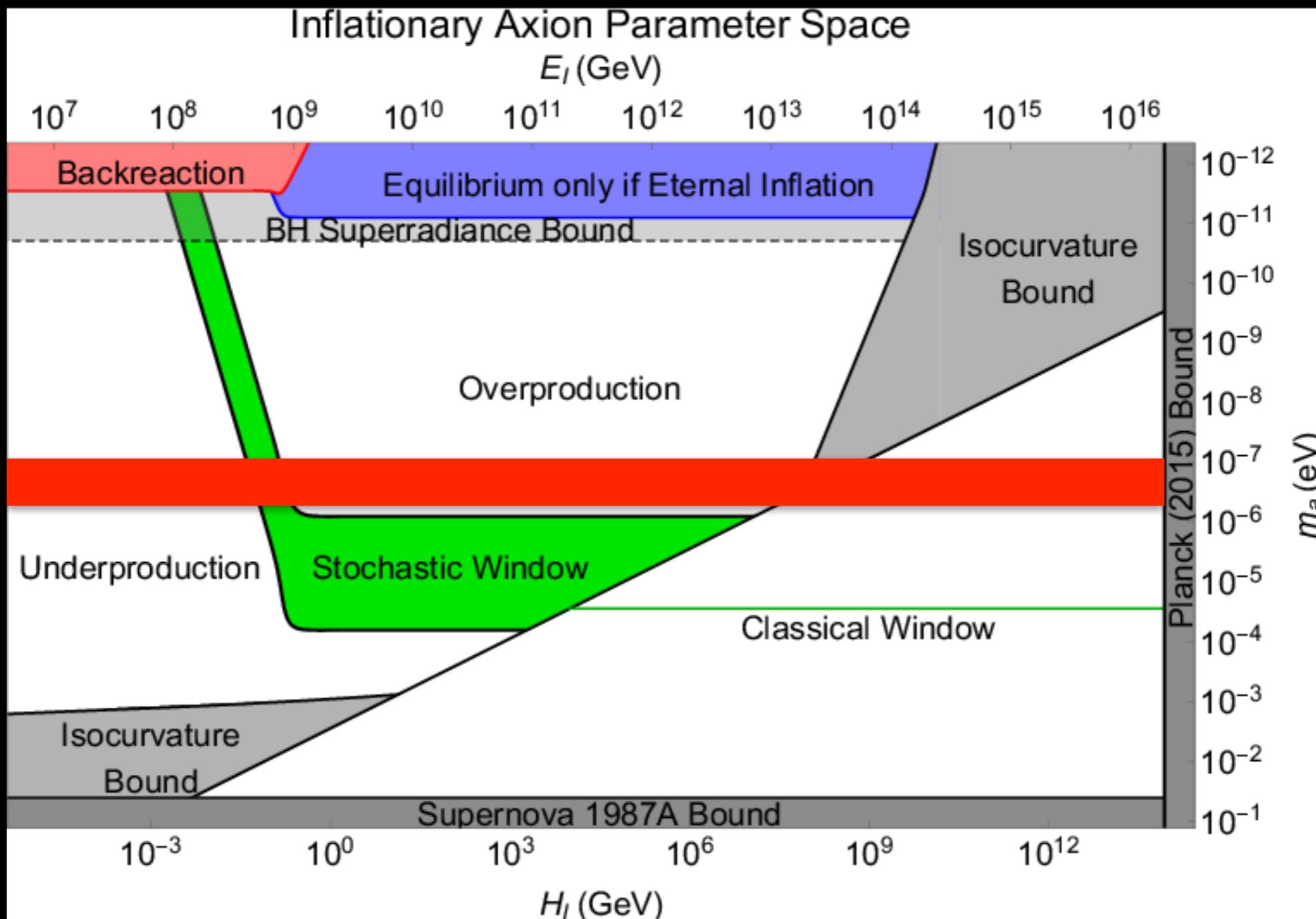
Andreas Ringwald 2018

Less crowded than other mass ranges for the QCD axion



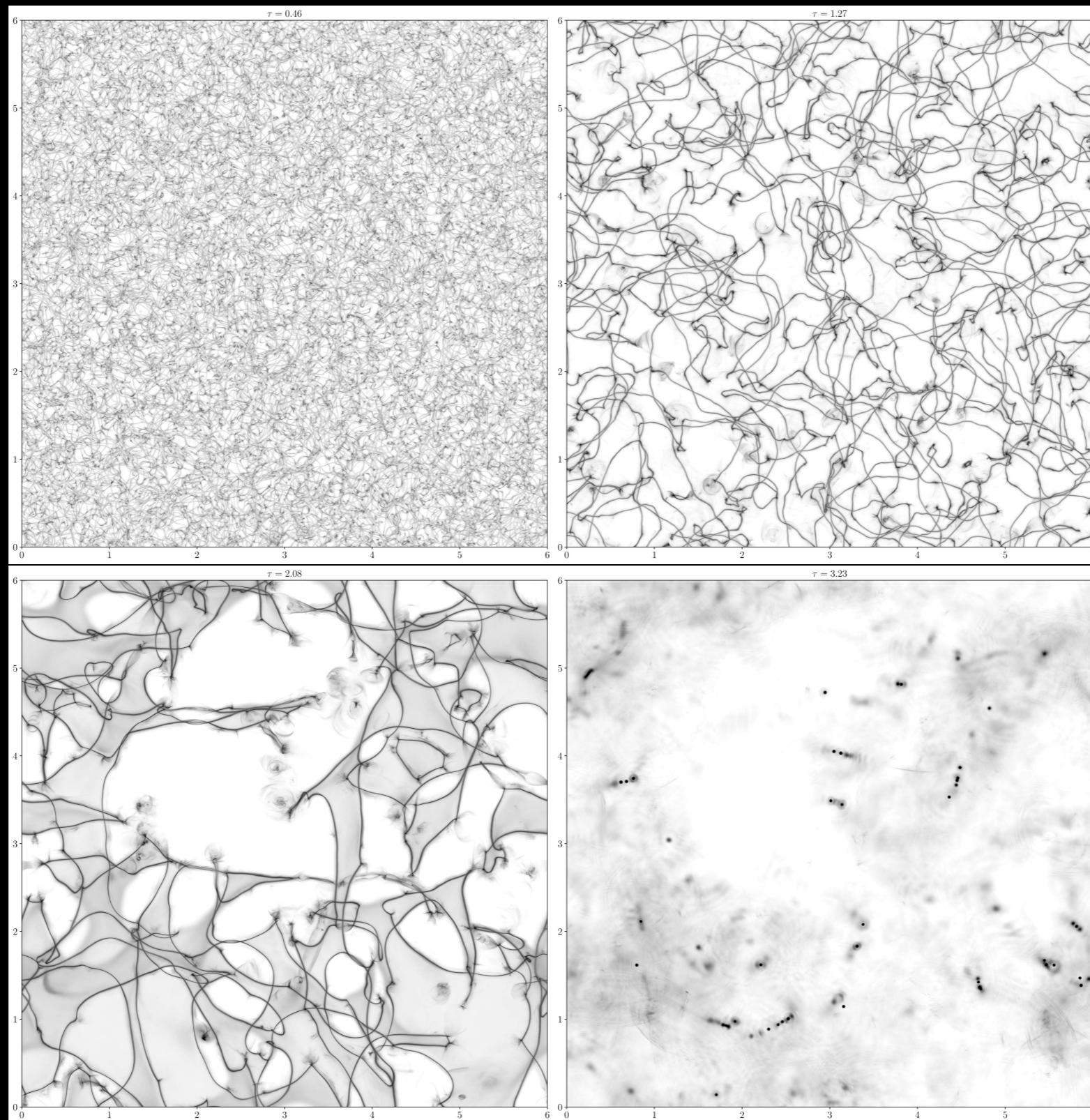
Andreas Ringwald 2018

Axion as the dark matter?



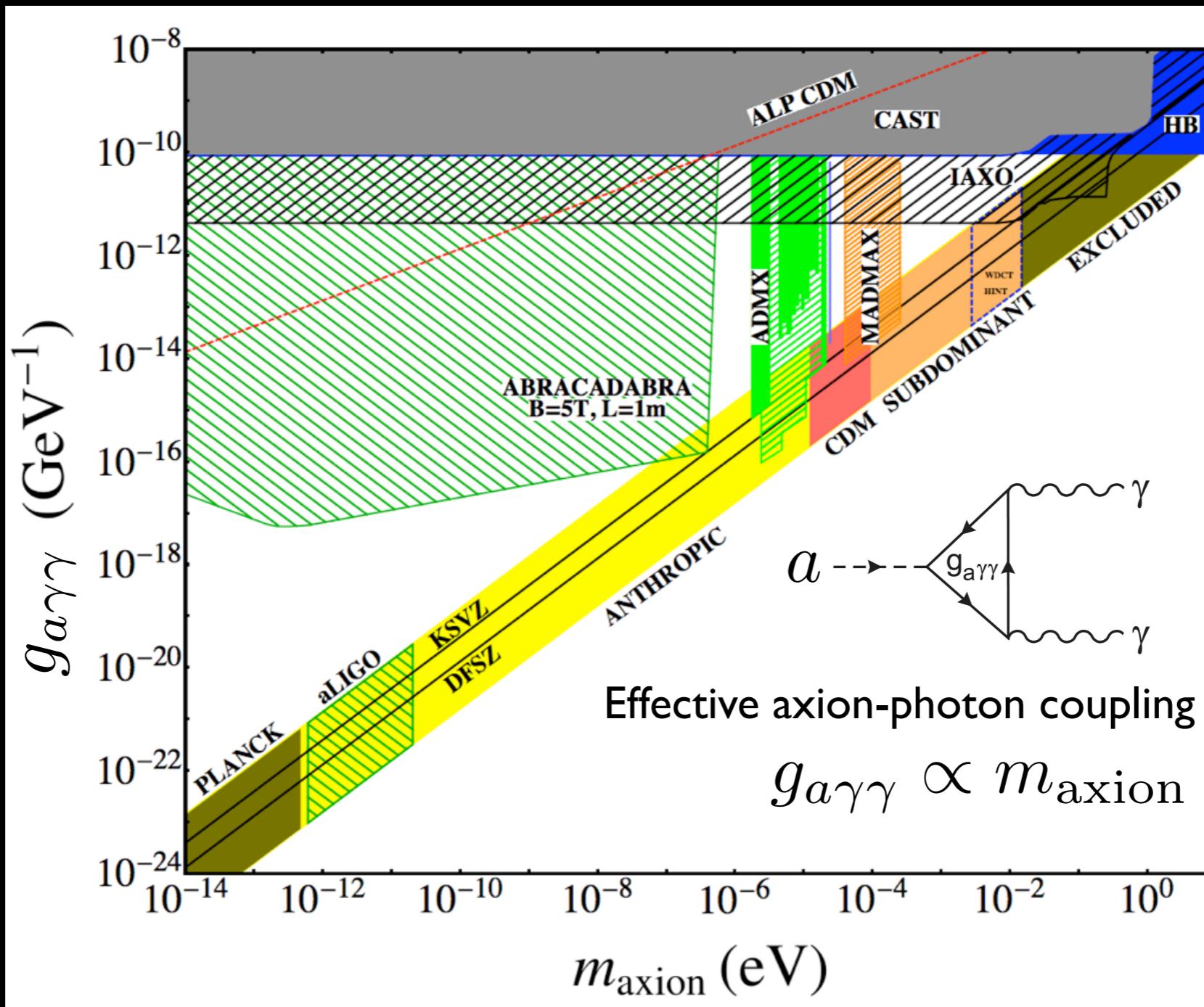
Graham et al 2018
Takahashi et al 2018

Destruction of the string-wall network



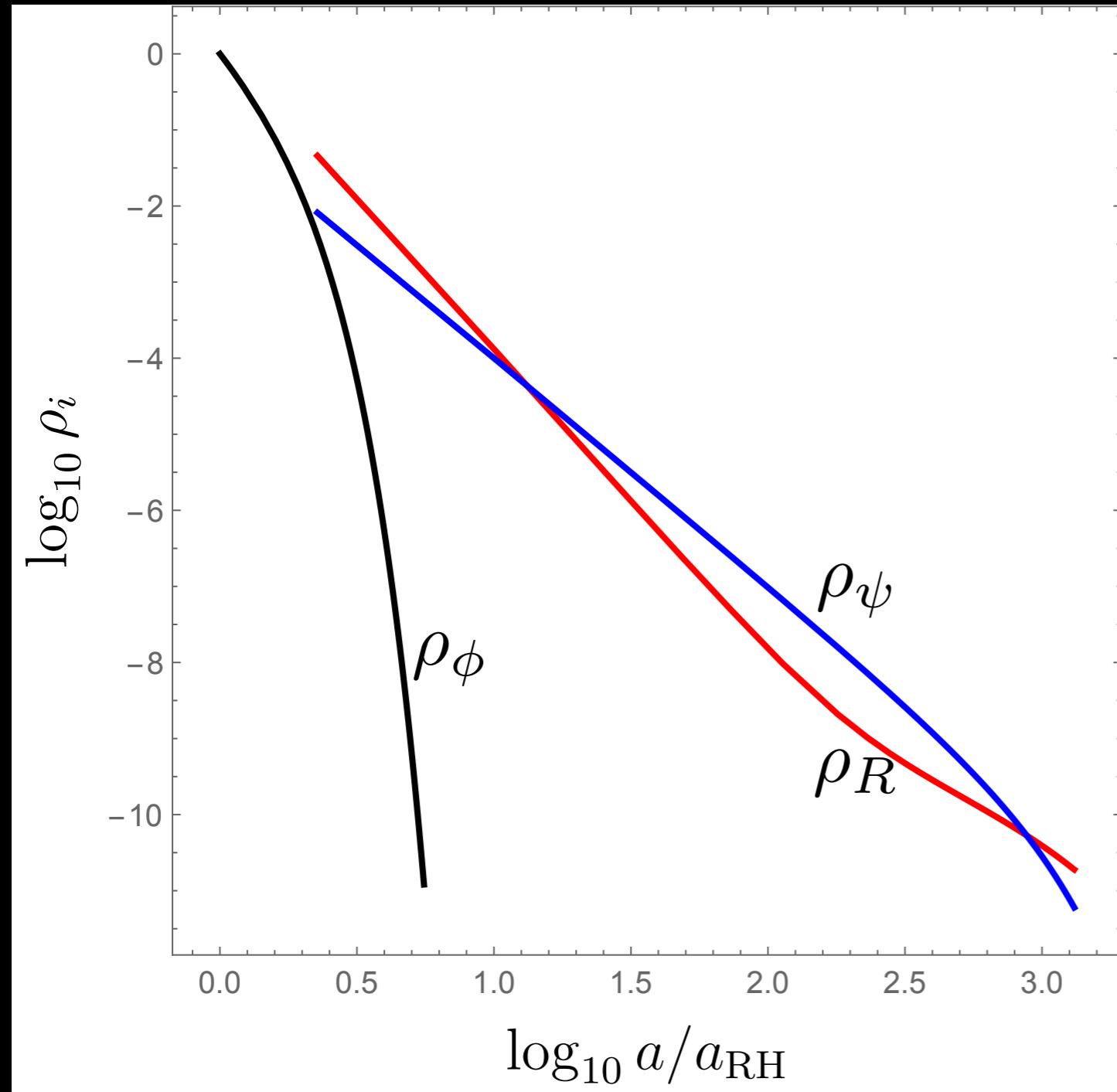
Vaquero+2018

One-parameter theory, falsifiable



Alternative DM candidate: the QCD Axion
One-parameter theory, falsifiable

Modified cosmologies



Early matter-dominated cosmology

$$\rho_\phi + 3H\rho_\phi = -\Gamma\rho_\phi$$

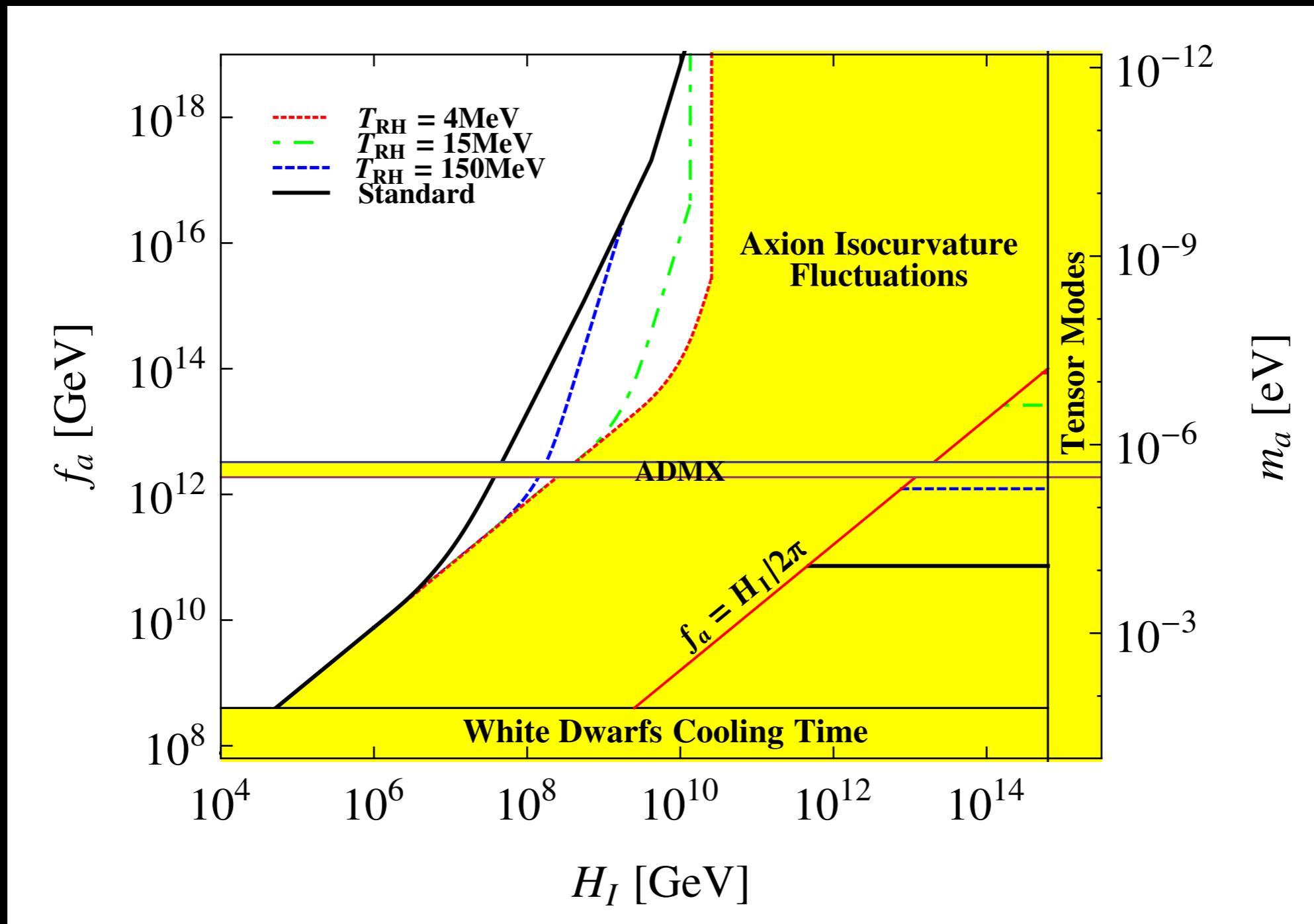
$$\rho_R + 4H\rho_R = b\Gamma\rho_\phi$$

$$\rho_\psi + 3H\rho_\psi = (1 - b)\Gamma\rho_\phi$$

Structures grow linearly....

$m_a \sim 0.1 \div 10 \mu\text{eV} \ll$ standard scenario

$f_a \sim 10^{15} \text{ GeV}$



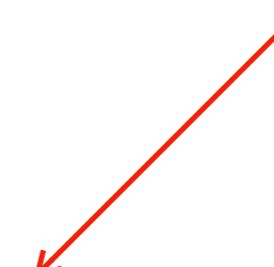
Hidden photons

$$\mathcal{L} = -\frac{1}{4}F^{\mu\nu}F_{\mu\nu} - \frac{1}{4}B^{\mu\nu}B_{\mu\nu} - \frac{1}{2}\chi F^{\mu\nu}B_{\mu\nu} + \frac{1}{2}m_{\gamma'}^2 B^\mu B_\mu$$

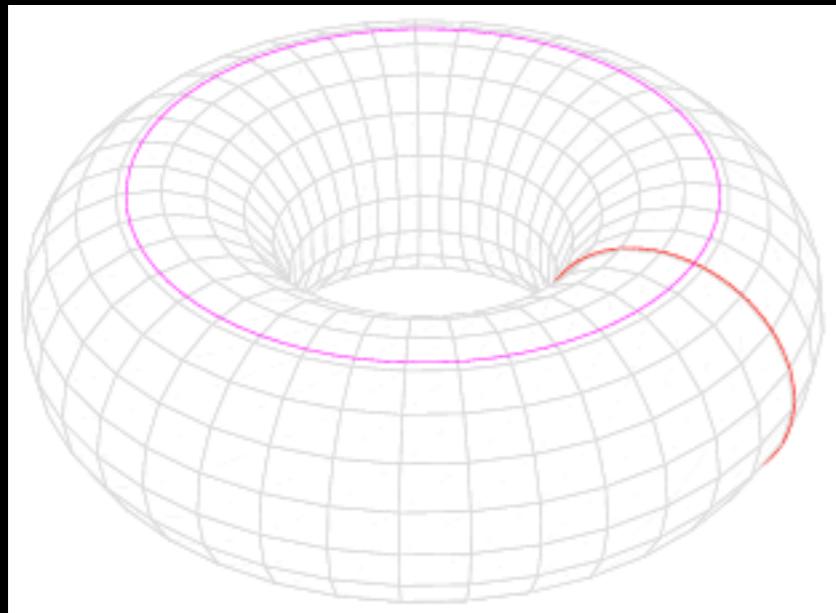
Cambio di base $B^\mu \rightarrow \tilde{B}^\mu - \chi A^\mu$

$$\mathcal{L} = -\frac{1}{4}F^{\mu\nu}F_{\mu\nu} - \frac{1}{4}\tilde{B}^{\mu\nu}\tilde{B}_{\mu\nu} + \frac{1}{2}m_{\gamma'}^2(\tilde{B}^\mu\tilde{B}_\mu - 2\chi\tilde{B}^\mu A_\mu + \chi^2 A^\mu A_\mu)$$

photon - hidden photon mixing



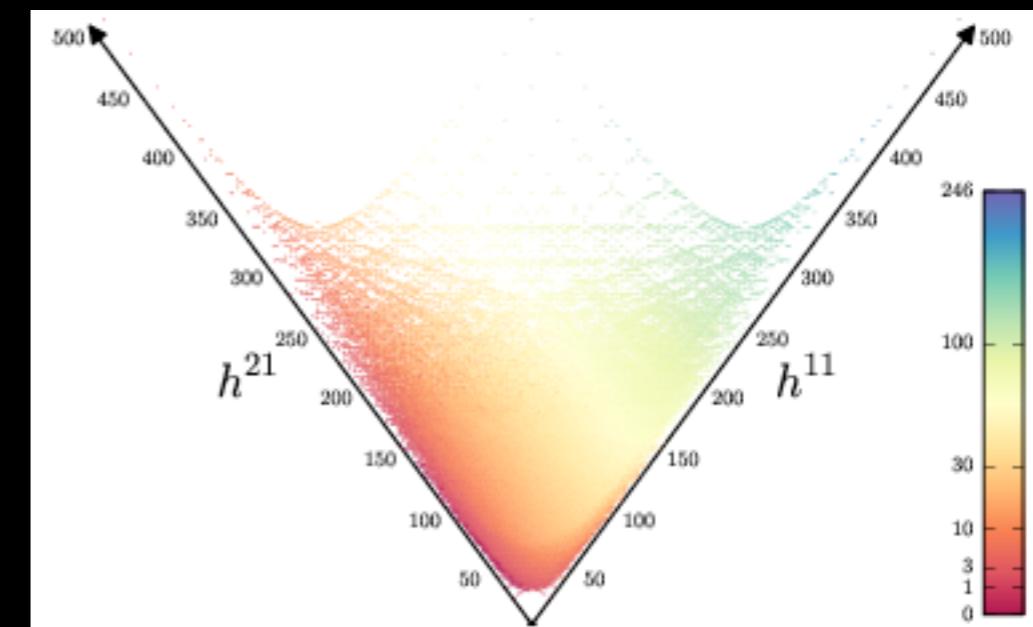
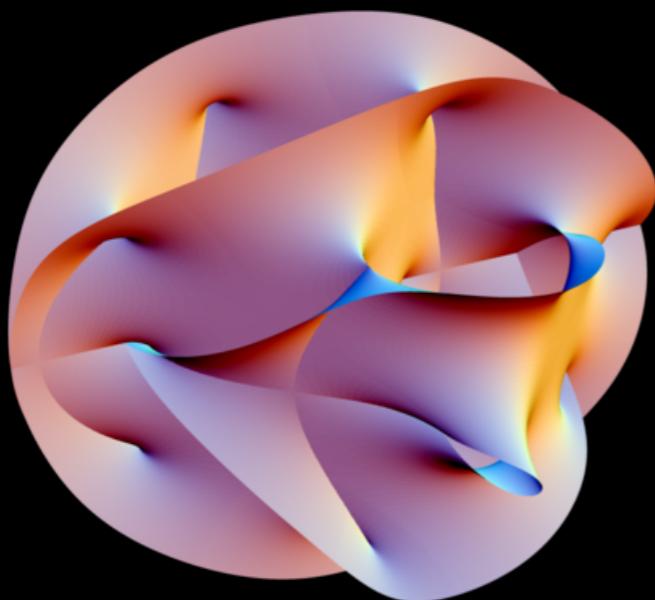
Axions from String Theory



A torus has two “moduli” fields
In general, # fields given by the
topology of the manifold

Moduli are paired with axions

A 6D manifold is more complicated than a torus...



Density of Hodge numbers for all
known Calabi-Yau three-folds

Axion-like particles for multiple purposes



Strong-CP

Peccei, Quinn; Weinberg; Wilczek

Dark Matter

Preskill+; Abbot+; Dine+; Baldeschi+

Baryogenesis

Ballestreros+

Dark Energy

Frieman+; Coble+; Arvanitaki+

Inflation

Freese+