



# Commissioning of the ATLAS Pixel Detector

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#### On behalf of the ATLAS Collaboration

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## The ATLAS Experiment



#### The ATLAS Pixel Detector

✓ The Pixel Detector is the innermost sub-system of the ATLAS experiment.

✓ It operates within a magnetic field of 2 Tesla.
✓ It helps to determine the momenta of charged

particles and to reconstruct tracks and vertexes. It permits to obtain high precision on the determination of the impact parameter.

$$\Box \Rightarrow \sigma(d_0) = 10 \mu m \oplus 140 \mu m / p_T [GeV/c]$$

✓ It's made by Silicon pixel sensors.

✓ 80 M channels.

✓ The first sensor is only 5 cm away from the interaction point.

✓ The system covers up to  $|\eta|$  < 2.5 with three measured points.

✓ Installed in the pit in June 2007.
 ✓ Fully connected and evaporative cooling commissioned in August 2008.
 ✓ First example as an but he Direct

✓ First cosmic track seen by the Pixel detector on the 14th September 2008.



#### Barrel

✓ 3 layers

#### Disks

- 1456 modules
- 3 disks
- 288 modules

#### Module

- ✓ sensor of (60.8mm X 16.4mm X 250µm)
- ✓ segmentation of pixels: 400µm X 50µm
- ✓ 16 front-end chips
- ✓ bump-bonding between sensor and FE

## The cosmic rays data taking – 2008/2009



#### Performance

After 2009 data taking – 98% of the Pixel detector is operational.



Intrinsic detector efficiency: 99.8%.

Number of pixel masked (online and offline) due to high noise. The fraction of problematic pixels is well below 0.01%.



### Pixel detector calibration



✓ The Pixel detector has been fully calibrated.

✓ The thresholds have been tuned for the different types of pixels.
✓ 4000 e threshold with 40 e dispersion.



✓ Pixel cluster charge distributions in data is in good agreement with expectations.

### The Lorentz angle

Due to the presence of the magnetic field, the charge drift along the Lorentz angle direction.  $\bigotimes^{\vec{B}}$ 

> tan  $\alpha_L = \mu_H B$ The point displacement due to the Lorentz angle is  $\approx 30 \ \mu m$ , bigger than the intrinsic detector resolution.

Width of the cluster (i.e. group of adiacent pixels) as a function of the track incidence angle. The minimum of the distribution determines the Lorentz angle.



## Alignment & Resolution

✓ Using cosmic ray data it was possible to align the detector for the first time.

✓ Alignment is an iterative process and is obtained in many steps. Bigger structures (shells and staves) are aligned first and then all the substructures, down to the module level.





500

-0.5 -0.4 -0.3 -0.2 -0.1 -0

direction.

0.2 0.3 0.4 0.

y residual [mm]

0.1

### Alignment & Resolution

✓ Cosmic tracks crossing the entire Inner Detector leave hits in both the upper and lower halves of the ID. These tracks can be split near the interaction point and fit separately, resulting in two collision-like tracks that can then be compared.



### **Pixel Detector Commissioning**

✓ The ATLAS Pixel Detector has been fully commissioned thanks to the high statistic of cosmic tracks collected in 2008 and 2009.

✓ The 98% of the detector is operational and the intrinsic efficiency of the detector is 99.8%.

✓ The number of noisy pixels is less than 0.01% and the noise occupancy is of the order of  $10^{-10}$  hits/pixel/event/BC.

 $\checkmark$  The detector has been fully calibrated.

 $\checkmark$  The whole Inner Detector has been aligned.



✓ Using splitted tracks it has been possible to determine the impact parameter resolution. It amounts to 22  $\mu$ m for high momentum tracks and it s already sufficient for initial physics measurements.

✓ The ATLAS Pixel Detector is ready for LHC beams this year.