

Light dark matter at colliders

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Der Wissenschaftsfonds.

ÖAW 🕷

ÖSTERREICHISCHE AKADEMIE DER WISSENSCHAFTEN 10-11/10/2018 Outline

- 1) Kinetic mixing and the dark photon
- 2) A light & dark Z' boson
- **3)** Decays of Y(1S) to dark matter



Long lived dark photon search

PART 2

PART 1



The general idea: kinetic mixing γ -A'

Dark photon first proposed in

P. Fayet, Phys. Lett. B **95**, 285 (1980),P. Fayet Nucl. Phys. B **187**, 184 (1981).

 (Holdom, 1986) A boson belonging to an additional U(1)' symmetry would mix kinetically with the photon:



- → The kinetic mixing is a term in the Lagrangian expressed by $\frac{1}{2} \epsilon F_{\mu\nu}^{Y} F'^{\mu\nu}$.
- For the dark photon to acquire mass an extended Higgs sector can be required to break the new U(1)' symmetry.

Note: ϵ is the strength of the kinetic mixing and it is supposed to be small, 10^{-5} - 10^{-2} , **the smaller the value of \epsilon the longer A' lifetime** (i.e. **long lived**). The Mass of the new boson should be in the range few MeV to few Gev (Nima Arkani-Hamed et al. Phys. Rev. D **79**, 015014, 2009).

Consideration about dark sectors

A'= dark photon, h_D = dark Higgs boson, χ = dark matter

Most dark sector models require an additional U(1) symmetry responsible for the "interactions" between dark sector particles and SM particles through its gauge boson A'.



P. Fayet, Phys. Lett. B **95**, 285 (1980),
P. Fayet Nucl. Phys. B **187**, 184 (1981).
B. Holdom, Phys. Lett. B **166**, 196 (1986)

Kinetic mixing strength

- $M(A') \sim GeV$ scale \rightarrow mixing with the photon, decays to SM final states possible
- $M(A') \sim EW$ scale \rightarrow mixing with Z⁰, effects in rare decays (Y, B, ..) through loops¹, decays to SM final states possible
- M(A') ~ TeV scale \rightarrow effects in rare decays (Y, B, ..) through loops¹
- $M(h_D)$ ~ GeV scale \rightarrow dark higgs-strahlung, rare decays
- $M(\chi) \sim GeV \text{ scale: } B \rightarrow \chi\chi, B \rightarrow \nu\chi; Y(1S) \rightarrow \chi\chi; Y(3S) \rightarrow \chi\chi\gamma, A' \rightarrow \chi\chi$

Invisible B/Y decays not accessible at hadron colliders \rightarrow Belle II

¹Remember the lesson from the past, new heavy particles are first seen indirectly or in loops: Z⁰, charm, top

Dark photon searches in e+e- collisions: BABAR, KLOE, BESIII, Belle/Belle II



Dark photon searches in e+e- collisions KLOE, BES III, BABAR & Belle + Belle II



A'= dark photon.

A' decays to SM final states through kinetic mixing (if allowed by kinematics). Low multiplicity final states. 2 charged tracks and 1 photon, prompt.

"A' " decays depend on
$$M_{A'}$$
:

-Decays to leptons require $M_{A'}$ >1.02 MeV/c² -Decays to hadrons require $M_{A'}$ >0.36 GeV/c²



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Dark photon searches in e+e- collisions KLOE, BES III, BABAR & Belle + Belle II

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Dark photon searches in e+e- collisions BES III







Dark photon searches in e+e- collisions BES III



Dark photon searches in e+e- collisions KLOE-2



$$\varepsilon^2 = \frac{N_{CL_{\rm S}}/\epsilon_{\rm eff}}{H \cdot I \cdot L}$$

 $N_{CL_S} = UL$ on number of U-boson candidates at 90% CL (CL_S technique)

$$H = \frac{d\sigma_{\pi\pi\gamma}/dM_{\pi\pi}}{\sigma(\pi^+\pi^- \to \pi^+\pi^-, M)}$$
$$I = \int \sigma_{\pi\pi}^{U} dM_{\pi\pi}$$
$$\epsilon_{eff} = 2 - 40\%$$
Systematic error $\leq 1\%$
$$L = 1.93 \text{ fb}^{-1}$$

Phys. Lett. B 757 (2016) 356



Dark photon searches in e+e- collisions KLOE-2



Systematic error $\leq 1\%$ L=1.93 fb⁻¹

Phys. Lett. B784 (2018) 336-341, arXiv:1807.02691 [hep-ex] 10-4 10^{-5} ϵ^2 10^{-7} BaBar — KLOE_{μμ + ππ} ----- KLOE_{μμ} LHCb ······ KLOE Na48 10 200 800 1000 400 600 $M_{\rm II}$ (MeV)

"The combined upper limit, obtained after averaging the statistical fluctuations by a smoothing procedure, excludes values of ε^2 greater than (13–2)×10⁻⁷ in the U-mass range 519–987MeV".

Dark photon searches in e+e- collisions BABAR



A'= dark photon, L= long lived light gauge boson.

A' decays to SM final states through kinetic mixing (if allowed by kinematics). Low multiplicity final states. 2 charged tracks and 1 photon, prompt or displaced vertex decay. Could require dedicated trigger to increase efficiencies, especially for the displaced vertex case.

"A' " decays depend on $M_{A'}$:

-Decays to leptons require $M_{A'}$ >1.02 MeV/c² -Decays to hadrons require $M_{A'}$ >0.36 GeV/c²



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Dark photon searches in e+e- collisions BABAR See R. Essig et al. JHEP11 (2013) 167.

 $e^{+}(e^{-})$ γ $e^{-}(e^{+})$ A' $\overline{\chi}$

A'= dark photon, χ = dark matter particle (neutral under SU(3)xSU(2)xU(1)) A' decays to dark matter. One on-shell (mono-energetic) or one off-shell (broad spectrum) photon with different gamma spectrum .

> radiative production in e+e- collisions only one photon in the final state with $E_{\gamma}^* = (s - M_{A'}^2)/2\sqrt{s}$ (*on-shell*) No existing limits

Requires high rate single photon trigger, not available in Belle, however the analysis is ongoing using photon conversion. The BaBar Collaboration implemented a single photon trigger (arXiv:0808.0017 [hep-ex]). Single photon trigger available at Belle II.

Dark photon searches in e+e- collisions BABAR



*The level 1 trigger required a single energy cluster > 800 MeV in the laboratory frame. Rate: 350 Hz (8% of the total) at a typical luminosity of 8 × 10^{33} running on the Y(2S) or Y(3S) resonances

Dark photon searches in e+e- collisions BABAR vs. Belle II



→ See Belle II Physics book for the Belle II details: arXiv:1808.10567

*The level 1 trigger required a single energy cluster > 800 MeV in the laboratory frame. Rate: 350 Hz (8% of the total) at a typical luminosity of 8 × 10^{33} running on the Y(2S) or Y(3S) resonances.

BABAR: invisible channels



Y(*nS*): bound state of a *b* quark and a *b* antiquark

 $\frac{BR(Y(1S) \rightarrow v \bar{v})}{BR(Y(1S) \rightarrow e^+ e^-)} = \frac{27G^2 M_{Y(1S)}^4}{64\pi^2 \alpha^2} (-1 + \frac{4}{3} \sin^2 \theta_W)^2 = 4.14 \times 10^{-4}$ $BR(Y(1S) \rightarrow v \bar{v}) \sim 9.9 \times 10^{-6}$

- → Low mass dark matter particles however might might play a role in the decays of Y(1S), having Y(1S) → χχ if kinematic allowed.
 [Phys. Rev. D 80, 115019, 2009]
- → Also, new mediators (Z', A⁰, h⁰) or SUSY particles might enhance Y(1S) $\rightarrow \nu\nu(\gamma)$. [Phys. Rev. D **81**, 054025, 2010]
- → In absence of new physics enhancement, Belle2 should be able to observe the SM $Y(1S) \rightarrow vv$



 $M_{Y(3S)} = 10.355 \, GeV/c^2$, $M_{Y(2S)} = 10.023 \, GeV/c^2$, $M_{Y(1S)} = 9.460 \, GeV/c^2$



Belle2 Simulation $Y(3S) \rightarrow \pi^+\pi^-Y(1S),$ $Y(1S) \rightarrow \nu\nu$

Charge=1, PDG=211 (pi+)
pT=0.420365, pZ=0.000692372
Y=(-0.00, -0.00, -0.03)
Mother: MCParticles[0] (Upsilon(3S))



Charge=-1, PDG=-211 (pi-) pT=0.344016, pZ=0.118851 Y=(-0.00, -0.00, -0.03) Mother: MCParticles[0] (Upsilon(3S))

~ 540 MeV available for $P_{\pi\pi}$

$$\frac{BR(Y(1S) \rightarrow v \bar{v})}{BR(Y(1S) \rightarrow e^+ e^-)} = \frac{27G^2 M_{Y(1S)}^4}{64\pi^2 \alpha^2} (-1 + \frac{4}{3} \sin^2 \theta_w)^2 = 4.14 \times 10^{-4}$$
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- → In absence of new physics enhancement, Belle2 should be able to observe the SM $Y(1S) \rightarrow vv$

A signal of $Y(1S) \rightarrow invisible$ is an excess of events over the background in the M_r distribution at a mass equivalent to that of the Y(1S) (9.460 GeV/c²)

$$M_r^2 = s + M_{\pi^+\pi^-} - 2\sqrt{s} E_{\pi^+\pi^-}^{CMS}$$

$$e^{+}e^{-} \rightarrow Y(3S)$$

$$\downarrow (4.4\%)$$

$$Y(3S) \rightarrow \pi^{+}\pi^{-}Y(1S)$$

$$\downarrow$$

$$Y(1S) \rightarrow invisible$$

$$e^{+}e^{-} \rightarrow Y(2S)$$

$$\downarrow (18.1\%)$$

$$Y(2S) \rightarrow \pi^{+}\pi^{-}Y(1S)$$

$$\downarrow$$

$$Y(1S) \rightarrow invisible$$

 $\begin{array}{l} \mbox{Belle2 Simulation} \\ \mbox{Y(3S)} \rightarrow \pi^{+}\pi^{-}\mbox{Y(1S)}, \\ \mbox{Y(1S)} \rightarrow \nu\nu \end{array}$

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No signal was observed over the expected background and upper limits have been obtained: BR(Y $\rightarrow \nu\nu$) < 3x10⁻⁴ (BaBar) and BR(Y $\rightarrow \nu\nu$) < 3.0x10⁻³(Belle).

At Belle 2 one would expect to collect >200fb⁻¹ of data @ Y(3S) (ongoing discussion for Y(2S) data taking and trigger) allowing one to reconstruct between 30 and 300 events, assuming 10^{-5} (SM)<BR(Y \rightarrow invisible)< 10^{-4} (NP) and Belle efficiencies. 20

Irreducible peaking background when final states go undetected (i.e. detector supports, beampipe etc.) in the process $Y(3S) \rightarrow \pi^+ \pi^- Y(1S), Y(1S) \rightarrow undetected f.s.$

Q: Can advanced machine learning techniques separate reduce this irreducible Background? Tests are planned..

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No signal was observed over the expected background and upper limits have been obtained: BR($Y \rightarrow vv$) < 3x10⁻⁴ (BaBar) and BR($Y \rightarrow vv$) < 3.0x10⁻³(Belle).

Process	$L_{int}(ab^{-1})$	ϵ	$N(\Upsilon(1S))$	$N_{\Upsilon(1S)\to\nu\bar{\nu}}$	N_{NP}
$\Upsilon(2S) \to \pi^+ \pi^- \Upsilon(1S)$	$0.2, \Upsilon(2S)$	0.1-0.2	2.3×10^8	230-460	6900-13800
$\Upsilon(3S) \to \pi^+ \pi^- \Upsilon(1S)$	$0.2, \Upsilon(3S)$	0.1-0.2	3.2×10^7	32-64	945-1890
$\Upsilon(4S) \to \pi^+ \pi^- \Upsilon(1S)$	$50.0, \Upsilon(4S)$	0.1-0.2	5.5×10^6	5.5-11	165 - 310
$\Upsilon(5S) \to \pi^+ \pi^- \Upsilon(1S)$	$5.0, \Upsilon(5S)$	0.1 - 0.2	7.6×10^{6}	7.6-15.2	228-456
$\gamma_{ISR}\Upsilon(2S) \to (\gamma_{ISR})\pi^+\pi^-\Upsilon(1S)$	$50.0, \Upsilon(4S)$	0.1 - 0.2	1.5×10^8	150 - 300	4500-9000
$\gamma_{ISR}\Upsilon(3S) \to (\gamma_{ISR})\pi^+\pi^-\Upsilon(1S)$	$50.0, \Upsilon(4S)$	0.1-0.2	3.5×10^7	35-70	1050-2100

 $e^{+}e^{-} \rightarrow Y(3S)$ $\downarrow (4.4\%)$ $Y(3S) \rightarrow \pi^{+}\pi^{-}Y(1S)$ \downarrow $Y(1S) \rightarrow invisible$ $e^{+}e^{-} \rightarrow Y(2S)$ $\downarrow (18.1\%)$ $Y(2S) \rightarrow \pi^{+}\pi^{-}Y(1S)$ \downarrow $Y(1S) \rightarrow invisible$

 $\begin{array}{l} \mbox{Belle2 Simulation} \\ \mbox{Y(3S)} \rightarrow \pi^{+}\pi^{-}\mbox{Y(1S)}, \\ \mbox{Y(1S)} \rightarrow \nu\nu \end{array}$

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Translating $Y(1S) \rightarrow$ invisible search to dark matter limits



Extrapolation based on ArXiv: 1511.03728, 1404.6599

The L_u-L_t model in the context of dark sector searches: a dark Z'

The model is a new gauge boson, Z', which couples to $L_{\mu} - L_{\tau}$. The interaction Lagrangian is

$$\mathcal{L} = -g'\bar{\mu}\gamma^{\mu}Z'_{\mu}\mu + g'\bar{\tau}\gamma^{\mu}Z'_{\mu}\tau - g'\bar{\nu}_{\mu,\mathrm{L}}\gamma^{\mu}Z'_{\mu}\nu_{\mu,\mathrm{L}} + g'\bar{\nu}_{\tau,\mathrm{L}}\gamma^{\mu}Z'_{\mu}\nu_{\tau,\mathrm{L}}.$$

The equations for the partial widths are,

$$\begin{split} \Gamma(Z' \to \ell^+ \ell^-) &= \frac{(g')^2 M_{Z'}}{12\pi} \left(1 + \frac{2M_\ell^2}{M_{Z'}^2} \right) \sqrt{1 - \frac{4M_\ell^2}{M_{Z'}^2}} \, \theta(M_{Z'} - 2M_\ell), \\ \Gamma(Z' \to \nu_\ell \bar{\nu}_\ell) &= \frac{(g')^2 M_{Z'}}{24\pi}. \\ BR(Z' \to invisible) &= \frac{2\Gamma(Z' \to \nu_l \overline{\nu}_l)}{2\Gamma(Z' \to \nu_l \overline{\nu}_l) + \Gamma(Z' \to \mu \overline{\mu}) + \Gamma(Z' \to \tau \overline{\tau})} \end{split}$$

Partial width results and BR derived from eqn. 2.12 of Essig et al. JHEP02(2015)157, arXiv:1412.0018 [hep-ph]:

$$\Gamma(Z_D \to \bar{f}f) = \frac{N_c}{24\pi m_{Z_D}} \sqrt{1 - \frac{4m_f^2}{m_{Z_D}^2}} \left(m_{Z_D}^2 \left(g_L^2 + g_R^2 \right) - m_f^2 \left(-6g_L g_R + g_L^2 + g_R^2 \right) \right)$$

The branching fraction to one neutrino species is half of the branching fraction to one charged lepton flavour. The reason is, of course, that the Z' only couples to left-handed neutrino chiralities whereas it couples to both left- and right-handed charged leptons.

→ For
$$M_{z'} < 2M_{\mu} Br(Z' \rightarrow invisible) = 1$$

- → For $2M_{\mu} < M_{z'} < 2M_{\tau} Br(Z' \rightarrow invisible)~1/2$
- → For $M_{z'}$ >2 M_{τ} Br(Z' → invisible)~1/3

The L_{u} - L_{τ} model in the context of dark sector searches: a dark Z'



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→ Final state: µ⁺µ⁻ +nothing

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 - → For $M_{z'}$ >2 M_{τ} Br($Z' \rightarrow$ invisible)~1/3

$Z' \rightarrow$ invisible, Belle 2 event display (MC simulation)



Cross section for Z' → invisible



- Cross section provided by MadGraph for $e^+e^- \rightarrow \mu^+\mu^- Z'$, $Z' \rightarrow \nu_{\mu} \overline{\nu_{\mu}}$ and multiplied by a factor 2 to account for $Z' \rightarrow \nu_{\tau} \overline{\nu_{\tau}}$ as this is the other channel that contribute to the invisible decays of Z'.

It is assumed g' = 0.01, cross section for different values of g' can be obtained by considering that:
 ...the production process involves exactly one insertion of the g' coupling in the matrix element, then we know that the cross section is proportional to g'². Let's say you use g' = 10⁻³ to generate events in MadGraph. Then, if I want the cross section for g' = 10⁻⁴, I simply take the cross section output by MG and multiply by (10⁻¹)²...

Cross section for Z' → invisible



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 ...private conversation with B. Shuve.

Belle II, expected sensitivity for $Z' \rightarrow invisible$

- Based only on expected background and luminosity
- Expected upper limits to g' value at 90% C.L.
- Bad mass resolution on the signal at low masses affects final sensitivity



- Red band shows the preferred $(\pm 2\sigma)$ region of the parameter space assuming the muon g-2 anomaly being generated by a Z' boson.

Belle II 2018 data, event display



Belle II 2018 data, event display



We consider now a different model, in which the Z' does not couple only to e and μ , but it couples to all leptons and we allow for lepton flavour violation.

See for example arXiv:1610.08060 or ArXiv:1701.08767



final state: $e^+e^-\mu^+\mu^-$

final state: $e^+\mu^-$ + invisible (+c.c.)

See for example arXiv:1610.08060 or ArXiv:1701.08767

- Complement the search for low mass Z' and low mass dark sector
- Alternative way to look into cLFV, complementing ongoing searches
- → Work in progress at Belle II



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- Almost) background free
- Get a search for doubly charged bosons for free
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Dark photon searches in pp collisions: ATLAS, CMS, LHCb



Dark photon searches @ ATLAS





- The Higgs boson decays to dark fermions; dark fermions decay to
 - 1) A Hidden Lightest Stable Particle and a dark photon
 - 2) A Hidden Lightest Stable Particle and a dark scalar
 - The dark scalar decays to a dark photon pair



HLSP.

HLSP

Dark photon searches @ ATLAS



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- arXiv:1811.07370 [hep-ex]
 - Similar techniques as shown before
- Search for macroscopic decay lengths of scalar particles under different assumptions



• The analysis strategy includes:

1) the case of two observed isolated vertices coming from two long lived particles,

2) The case in which only one decay vertex is seen in combination with other activity in the detector. This is to increase the sensitivity (probability to observe two long lived particles inside the detector is small for decay lengths grater than 5m).

- arXiv:1811.07370 [hep-ex]
 - Similar techniques as shown before
- Search for macroscopic decay lengths of scalar particles under different assumptions



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Scalar proper lifetime (ct) [m]

- arXiv:1811.07370 [hep-ex]
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- arXiv:1811.07370 [hep-ex]
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Dark photon searches @ CMS

- Select 4-muon events
- Look for di-muon invariant mass bump at low mass in *isolated* muon pairs





Andy Haas

Search for Higgs boson decays to a dark photon pair plus X

Dark photon searches @ CMS

- Look for two mu+ mu- pairs to have the same mass
- Exclude SM H decays with ~1% BR to muon-jets





Dark photon searches @ CMS



arXiv:1506.00424 [hep-ex]

"This search constrains a large, previously unconstrained area of the parameter space. Unlike the other results in the figure, the CMS and ATLAS limits are model-dependent and only valid under the assumption that B(h \rightarrow 2n1 \rightarrow 4µ + X) \neq 0"

Dark photon searches @ LHCb $A' \rightarrow \mu^+\mu^-$, prompt and displaced searches

$$n_{\rm ex}^{A'}[m(A'),\varepsilon^2] = \varepsilon^2 \begin{bmatrix} \frac{n_{\rm ob}^{\gamma^*}[m(A')]}{2\Delta m} \end{bmatrix} \mathcal{F}[m(A')] \epsilon_{\gamma^*}^{A'}[m(A'),\tau(A')] \quad \begin{array}{l} \text{Phys. Rev. Lett. } \mathbf{120} \ \text{O61801} \\ \text{ArXiv: } \mathbf{1710.02867 \ [hep-ex]} \end{bmatrix}$$



Dark photon searches @ LHCb Preliminary studies for run 3: from ~2021

During run 3 LHCb will run in a "triggerless" Mode. Plan to search for

• low mass dark photon in the reaction:

 $D^{*0} \rightarrow D^0 \gamma$

$$\frac{\Gamma(D^{*0}) \rightarrow D^{0} A'}{\Gamma(D^{*0}) \rightarrow D^{0} \gamma} = \epsilon^{2} \left(1 - \frac{M_{A'}^{2}}{\Delta m_{D}^{2}}\right)^{3/2}$$



• Prompt and displaced vertex decays of the dark photon in the mass region[$2 m_{\mu}$, $50 GeV/c^2$] in the reaction:

 $pp \to X A', A' \to \mu^{+} \mu^{-}$ $\Gamma_{A'} \approx \Gamma(A' \to l^{+} l^{-}) = \frac{1}{3} \alpha \epsilon^{2} m_{A'} \sqrt{1 - \frac{4 m_{l}^{2}}{m_{A'}^{2}}} (1 + \frac{2 m_{l}^{2}}{m_{A'}^{2}})$







Conclusions

- The lack of solid experimental evidences for WIMPs might imply that DM belong to a more complicated dark sector with dark interactions/dark forces.
- Many diverse experiments currently searching an planning searches for dark forces/dark sector.
- Searches are of general interests and many people are joining the effort.
- The "Dark Sectors 2016 Workshop: Community Report", arXiv:1608.08632, summurise the strategies and the experimental opportunities for the next 5-10 years.

Eff. contact operators in for dark matter in $Y(1S) \rightarrow$ invisible

ArXiv: 1404.6599

Name	Interaction Structure	Annihilation	Scattering
F5	$(1/\Lambda^2) \bar{X} \gamma^\mu X \bar{q} \gamma_\mu q$	Yes	SI
F6	$(1/\Lambda^2) \bar{X} \gamma^\mu \gamma^5 X \bar{q} \gamma_\mu q$	No	No
F9	$(1/\Lambda^2) \bar{X} \sigma^{\mu u} X \bar{q} \sigma_{\mu u} q$	Yes	SD
F10	$(1/\Lambda^2) \bar{X} \sigma^{\mu\nu} \gamma^5 X \bar{q} \sigma_{\mu\nu} q$	Yes	No
S3	$(1/\Lambda^2) \imath Im(\phi^{\dagger}\partial_{\mu}\phi) \bar{q}\gamma^{\mu}q$	No	SI
V3	$(1/\Lambda^2) \imath Im (B^{\dagger}_{\nu} \partial_{\mu} B^{\nu}) \bar{q} \gamma^{\mu} q$	No	SI
V5	$(1/\Lambda)(B^{\dagger}_{\mu}B_{\nu} - B^{\dagger}_{\nu}B_{\mu})\bar{q}\sigma^{\mu\nu}q$	Yes	SD
V6	$(1/\Lambda)(B^{\dagger}_{\mu}B_{\nu}-B^{\dagger}_{\nu}B_{\mu})\bar{q}\sigma^{\mu\nu}\gamma^{5}q$	Yes	No
V7	$(1/\Lambda^2)B^{(\dagger)}_{\nu}\partial^{\nu}B_{\mu}\bar{q}\gamma^{\mu}q$	No	No
V9	$(1/\Lambda^2)\epsilon^{\mu\nu\rho\sigma}B^{(\dagger)}_{\nu}\partial_{\rho}B_{\sigma}\bar{q}\gamma_{\mu}q$	No	No

TABLE I. Effective contact operators which can mediate the decay of a $J^{PC} = 1^{--}$ quarkonium bound state. We also indicate if the operator can permit an *s*-wave dark matter initial state to annihilate to a quark/anti-quark pair; if so, then a bound can also be set by indirect observations of photons originating from dwarf spheroidal galaxies. Lastly, we indicate if the effective operator can mediate velocity-independent nucleon scattering which is either spin-independent (SI) or spindependent (SD).