

## The orbital parameters of GW170817 and GW150914

1 Figure 1 shows the frequency-time maps produced by the two LIGO detectors and by Virgo for the binary neutron star event GW170817 [1].

1a Extract<sup>1</sup> a set of data  $(f_i, \Delta t_i)$  from one of the panels (LIGO-Livingston provides the clearest signal), and given the analytic relation:

$$f_{\text{GW}}^{-8/3} = \frac{(8\pi)^{8/3}}{5} \left( \frac{G\mathcal{M}}{c^3} \right)^{5/3} (t_0 - t), \quad (1)$$

fit the observed points to obtain the chirp mass of the binary  $\mathcal{M}$ .

1b Compute the value of the chirp mass for different values of the frequency using eq. (1). Is  $\mathcal{M}$  constant during the inspiral? In which range?

1c Knowing that  $\mathcal{M} = \nu^{3/5} M$ , where  $\nu = m_1 m_2 / M^2$ , and  $M = m_1 + m_2$ , find a lower bound on the total mass. [hint: plot  $M(\mathcal{M}, m_i)$  vs one of the two masses  $m_i$ .]

1d Assuming circular motion, compute the separation and the orbital velocity of the two bodies during the inspiral.

2 The gravitational strain measured by LIGO for the first binary black hole event is shown in Fig. 2. The 6 points on the ascissa correspond to  $t = (0, 0.009, 0.02, 0.0275, 0.034, 0.04)$ .

2a Estimate the GW frequency at  $t_{\text{AB}} = (t_{\text{A}} + t_{\text{B}})/2$  and  $t_{\text{DE}} = (t_{\text{D}} + t_{\text{E}})/2$ , and assuming eq. (1) compute the chirp mass of the system [hint: remember you also need to derive  $t_0$ ].

2b As done for the points 1c and 2d, derive the lower mass of the binary, and assuming a Newtonian evolution compute an upper bound on the compactness of the 2 coalescing objects, and their velocity at the largest frequency.

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<sup>1</sup>You can use one of the free codes available online, like `GrapiClick` or `WebPlotDigitizer`.

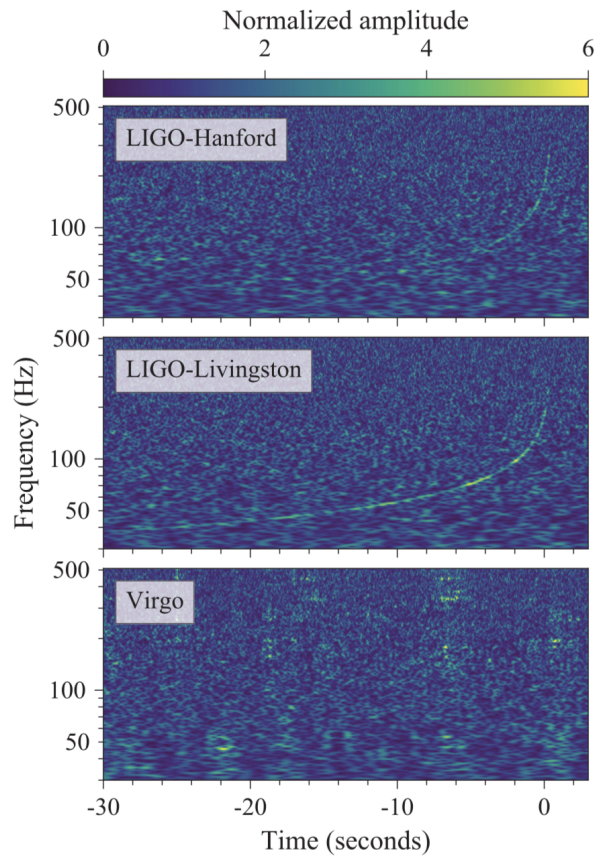


Figure 1: GW time-frequency maps for the GW170817 event, obtained from the 3 ground based interferometers LIGO/Virgo. Taken from [1].

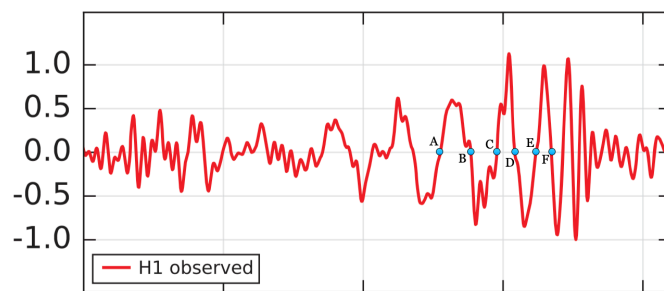


Figure 2: Gravitational wave strain measured for the GW150914 event by the LIGO Hanford interferometer. Adapted from [2].

# Bibliography

- [1] B. Abbott et al. GW170817: Observation of Gravitational Waves from a Binary Neutron Star Inspiral. *Phys. Rev. Lett.*, 119(16):161101, 2017.
- [2] B. P. Abbott et al. Observation of Gravitational Waves from a Binary Black Hole Merger. *Phys. Rev. Lett.*, 116(6):061102, 2016.