

# Neglected modular bootstrap equations and 3d Quantum Gravity

F. Gliozzi

Physics Department, Torino U.

SM&FT 2019

The XVIII Workshop on Statistical Mechanics  
and nonperturbative Field Theory

*Challenges in Computational Theoretical Physics*

## Quantizing 3d pure gravity?

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{g} (R - 2\Lambda)$$

- ✓ At the classical level it is “trivial”: no gravitons. Easy to quantize ?

## Quantizing 3d pure gravity?

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{g} (R - 2\Lambda)$$

- ✓ At the classical level it is “trivial”: no gravitons. Easy to quantize ?
- ✗ Despite to be trivial, actually it is unrenormalizable by power counting, since  $[G]=\text{Length}$

## Quantizing 3d pure gravity?

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{g} (R - 2\Lambda)$$

- ✓ At the classical level it is “trivial”: no gravitons. Easy to quantize ?
- ✗ Despite to be trivial, actually it is unrenormalizable by power counting, since  $[G] = \text{Length}$
- ✓ Any counterterm can be reabsorbed in a redefinition of  $\Lambda$  ,  $\Rightarrow \Lambda \neq 0$

## Quantizing 3d pure gravity?

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{g} (R - 2\Lambda)$$

- ✓ At the classical level it is “trivial”: no gravitons. Easy to quantize ?
- ✗ Despite to be trivial, actually it is unrenormalizable by power counting, since  $[G] = \text{Length}$
- ✓ Any counterterm can be reabsorbed in a redefinition of  $\Lambda$ ,  $\Leftrightarrow \Lambda \neq 0$
- ✗ Solutions with  $\Lambda > 0$  are unstable, must be studied as part of a larger system

## Quantizing 3d pure gravity?

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{g} (R - 2\Lambda)$$

- ✓ At the classical level it is “trivial”: no gravitons. Easy to quantize ?
- ✗ Despite to be trivial, actually it is unrenormalizable by power counting, since  $[G] = \text{Length}$
- ✓ Any counterterm can be reabsorbed in a redefinition of  $\Lambda$ ,  $\Leftrightarrow \Lambda \neq 0$
- ✗ Solutions with  $\Lambda > 0$  are unstable, must be studied as part of a larger system
- ✓  $\Lambda < 0$  is the only case having a chance to be consistently quantized

## Quantizing 3d pure gravity?

$$S = \frac{1}{16\pi G} \int d^3x \sqrt{g} (R - 2\Lambda)$$

- ✓ At the classical level it is “trivial”: no gravitons. Easy to quantize ?
- ✗ Despite to be trivial, actually it is unrenormalizable by power counting, since  $[G] = \text{Length}$
- ✓ Any counterterm can be reabsorbed in a redefinition of  $\Lambda$ ,  $\Leftrightarrow \Lambda \neq 0$
- ✗ Solutions with  $\Lambda > 0$  are unstable, must be studied as part of a larger system
- ✓  $\Lambda < 0$  is the only case having a chance to be consistently quantized
- ✓  $\Lambda < 0$  can be formulated in an asymptotically anti de Sitter spacetime  $\Leftrightarrow$  *AdS/CFT correspondence*

- \* According to holographic duality, solving pure quantum gravity means finding the  $2d$  CFT on the boundary of the asymptotically AdS spacetime in the large  $c$  (=central charge) limit, with

$$c = \frac{3\ell}{2G}, \quad (\Lambda = -\frac{1}{\ell^2})$$

- \* Euclidean AdS has a toroidal boundary,  $\Leftrightarrow$  *modular invariance*
- ✗ The degrees of freedom of pure gravity correspond, on the boundary, to multitraces of the stress-tensor belonging to the Virasoro module of the identity  $\Leftrightarrow$  not modular invariant
- $\Leftrightarrow$  Modular invariance of the boundary theory requires additional degrees of freedom on the bulk
- \* What is the meaning of modular invariance on the gravity side?



- \* On the boundary side, modular invariance can be obtained by summing over all the modular transformations and then regularize (Poincaré sum)

- \* On the boundary side, modular invariance can be obtained by *summing* over all the modular transformations and then regularize (Poincaré sum)
- \* In a quantum approach to gravity one expects to *sum* over different topologies of spacetime

- \* On the boundary side, modular invariance can be obtained by *summing* over all the modular transformations and then regularize (Poincaré sum)
- \* In a quantum approach to gravity one expects to *sum* over different topologies of spacetime
- ➡ Modular invariance arises from the *sum* over saddle points of the gravitational path integral

- ⇒ Modular invariance arises from the *sum* over saddle points of the gravitational path integral
- \* One saddle point is thermal AdS with periodic Euclidean time
- ✓ Others correspond to the black holes discovered by Bañados, Teitelboim and Zanelli (BTZ) : necessary degrees of freedom for a quantum description of pure gravity

- ⇒ Modular invariance arises from the *sum* over saddle points of the gravitational path integral
- \* One saddle point is thermal AdS with periodic Euclidean time
- ✓ Others correspond to the black holes discovered by Bañados, Teitelboim and Zanelli (BTZ) : necessary degrees of freedom for a quantum description of pure gravity
- \* Are they enough?

- ⇒ Modular invariance arises from the *sum* over saddle points of the gravitational path integral
- \* One saddle point is thermal AdS with periodic Euclidean time
- ✓ Others correspond to the black holes discovered by Bañados, Teitelboim and Zanelli (BTZ) : necessary degrees of freedom for a quantum description of pure gravity
- \* Are they enough?
- \* There is a one-to-one correspondence between the BTZ black holes and the primary states of the CFT on the boundary
- \* BTZ black holes can exist only above the threshold dual to the primary with  $\Delta_{BTZ} = \frac{c-1}{12} + O(1/c)$

- ⇒ Modular invariance arises from the *sum* over saddle points of the gravitational path integral
- \* One saddle point is thermal AdS with periodic Euclidean time
- ✓ Others correspond to the black holes discovered by Bañados, Teitelboim and Zanelli (BTZ) : necessary degrees of freedom for a quantum description of pure gravity
- \* Are they enough?
- \* There is a one-to-one correspondence between the BTZ black holes and the primary states of the CFT on the boundary
- \* BTZ black holes can exist only above the threshold dual to the primary with  $\Delta_{BTZ} = \frac{c-1}{12} + O(1/c)$
- ⇒ Proving that a primary with  $\Delta < \Delta_{BTZ}$  is necessary for a consistent CFT would be sufficient to argue that pure quantum gravity does not exist

- \* The modular invariance of the partition function of a CFT defined on a torus constrains the spectrum. It reads

$$Z(\tau, \bar{\tau}) = Z\left(\frac{a\tau+b}{c\tau+d}, \frac{a\bar{\tau}+b}{c\bar{\tau}+d}\right), \begin{pmatrix} a & b \\ c & d \end{pmatrix} \in G \equiv PSL(2, \mathbb{Z}), \tau \in H_+$$

it becomes particularly constraining on the special points of  $H_+$  invariant under the action of some non-trivial subgroup of  $G$

- \*  $G$  admits three of them: the cusp at  $\tau = i\infty$  is stabilized by  $T : \tau \rightarrow \tau + 1$ , the  $\mathbb{Z}_2$  elliptic point at  $\tau = i$  stabilized by  $S : \tau \rightarrow -1/\tau$ , the elliptic point at  $\tau = e^{2i\pi/3}$  stabilized by  $ST : \tau \rightarrow \frac{-1}{\tau+1}$ , the generator of a  $\mathbb{Z}_3$  subgroup of  $G$ .
- ⇒ Consistency conditions at  $\tau = i\infty$  imply integer spins  $j = h - \bar{h}$  and a constraint on the the twist gap  $\Delta - j$  at large  $j$
- \* smooth derivatives at  $\tau = i$  demand an infinite set of eq.s, dubbed modular bootstrap (Hellerman 2009):

$$(\tau \partial_\tau)^m (\bar{\tau} \partial_{\bar{\tau}})^n Z(\tau, \bar{\tau})|_{\tau=-\bar{\tau}=i} = 0 \text{ for } m+n \text{ odd}$$



- ⇒ Hellerman analytic upper bound for the scaling dimensions  $\Delta$  of the most relevant primary (very far from the threshold  $\Delta_{BTZ}$ )

$$\Delta_0 \leq \frac{c}{6} + 0.4737 \quad (c > 1)$$

- \* In spite of many analytical as well as numerical efforts on these modular bootstrap equations, small improvements have been found. Best results to date:
- \* analytic bound  $\Delta_0 \leq c/8.503 \quad (c \rightarrow \infty)$  (T. Hartman, D. Mazàc and L. Rastelli, 2019)
- \* numerical bound extrapolating large- $c$  data  
 $\Delta_0 \leq c/9.08 \quad (c \rightarrow \infty)$  (N. Afkhami-Jeddi, T. Hartman and A. Tajdini, 2019)

# The forgotten $\mathbb{Z}_3$ elliptic point at $\rho = \frac{-1}{\rho+1}$

- \* Consistency conditions: take arbitrary derivatives of  $Z(\tau, \bar{\tau}) = Z(\frac{-1}{\tau+1}, \frac{-1}{\bar{\tau}+1})$  and evaluate them at  $\rho = e^{2i\pi/3}$

# The forgotten $\mathbb{Z}_3$ elliptic point at $\rho = \frac{-1}{\rho+1}$

- \* Consistency conditions: take arbitrary derivatives of  $Z(\tau, \bar{\tau}) = Z\left(\frac{-1}{\tau+1}, \frac{-1}{\bar{\tau}+1}\right)$  and evaluate them at  $\rho = e^{2i\pi/3}$

$$\Rightarrow \partial_{\tau}^n Z|_{\tau=\rho} - \sum_{m=1}^n \rho^{m+n} \frac{n!}{m!} \binom{n-1}{m-1} \partial_{\tau}^m Z|_{\tau=\rho} = 0$$

# The forgotten $\mathbb{Z}_3$ elliptic point at $\rho = \frac{-1}{\rho+1}$

- \* Consistency conditions: take arbitrary derivatives of  $Z(\tau, \bar{\tau}) = Z(\frac{-1}{\tau+1}, \frac{-1}{\bar{\tau}+1})$  and evaluate them at  $\rho = e^{2i\pi/3}$

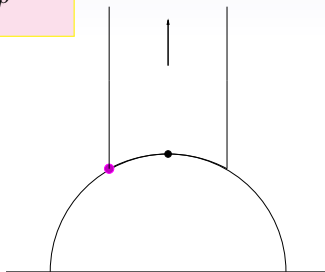
$$\Rightarrow \partial_{\tau}^n Z|_{\tau=\rho} - \sum_{m=1}^n \rho^{m+n} \frac{n!}{m!} \binom{n-1}{m-1} \partial_{\tau}^m Z|_{\tau=\rho} = 0$$

- \* Set  $\tau = -\frac{1}{2} + i\frac{\beta}{2\pi}$ ,  $\bar{\tau} = -\frac{1}{2} - i\frac{\bar{\beta}}{2\pi}$

$$Z = \sum_{h, \bar{h}} q^{h-\frac{c}{24}} \bar{q}^{\bar{h}-\frac{c}{24}}, \quad (\Delta = h + \bar{h}, j = h - \bar{h})$$

$$q \equiv e^{2i\pi\tau} = -e^{-\beta}, \quad \bar{q} \equiv e^{-2i\pi\bar{\tau}} = -e^{-\bar{\beta}}$$

- $\Rightarrow$  Z real, even spin contributions positive, odd spin contributions negative



## New modular bootstrap equations:

$$\partial_\beta Z|_{\beta=\bar{\beta}=\beta_c} = 0, \quad \left( \partial_\beta^4 + \frac{2\sqrt{3}}{\pi} \partial_\beta^3 \right) Z|_{\beta=\bar{\beta}=\beta_c} = 0,$$

$$\partial_\beta^2 Z|_{\beta=\bar{\beta}=\beta_c} = 0, \quad \left( \partial_\beta^5 - \frac{10}{\pi^2} \partial_\beta^3 \right) Z|_{\beta=\bar{\beta}=\beta_c} = 0,$$

$$\left( \partial_\beta^7 + \frac{525}{\pi^4} \partial_\beta^3 + \frac{7\sqrt{3}}{\pi} \partial_\beta^6 \right) Z|_{\beta=\bar{\beta}=\beta_c} = 0,$$

$$\left( \partial_\beta^8 - \frac{1470\sqrt{3}}{\pi^5} \partial_\beta^3 - \frac{98}{\pi^2} \partial_\beta^6 \right) Z|_{\beta=\bar{\beta}=\beta_c} = 0,$$

..... ,

$$\beta_c = \sqrt{3} \pi$$

$Z$  can be expanded in Virasoro characters. If  $c > 1$  and the theory is unitary, the modules of the Virasoro algebra are the identity degenerate module  $\chi_0(q)$  and a family of non-degenerate modules  $\chi_A(q)$

$$\chi_0(q) = \frac{q^{-\frac{c-1}{24}}}{\eta(\tau)} (1 - q), \quad \chi_A(q) = \frac{q^{h_A - \frac{c-1}{24}}}{\eta(\tau)} \quad (q = -e^{-\beta}, \bar{q} = -e^{-\bar{\beta}})$$

$$\Leftrightarrow Z = \chi_0(q)\chi_0(\bar{q}) + \sum_A N_A \chi_A(q)\chi_A(\bar{q}) \quad (N_A \in \mathbb{N})$$

$$\ast \sqrt{\tau - \bar{\tau}} \eta(\tau)\eta(\bar{\tau}) = Z_b^{-1} \text{ is modular invariant}$$

$$\Leftrightarrow Z/Z_b = Z_{vac} + \sum_A N_A Z_A \text{ modular invariant}$$

$$Z_{vac} = \sqrt{\beta + \bar{\beta}} e^{\beta \frac{c-1}{24}} e^{\bar{\beta} \frac{c-1}{24}} (1 + e^\beta)(1 + e^{\bar{\beta}}),$$

$$Z_A = \sqrt{\beta + \bar{\beta}} e^{\beta \frac{c-1}{24}} e^{\bar{\beta} \frac{c-1}{24}} e^{-\beta h_A} e^{-\bar{\beta} \bar{h}_A}$$

$\ast$  Applying the new bootstrap equations  $\Leftrightarrow \Delta_A = h_A + \bar{h}_A$  always appear in the combination  $\Delta_A - \Delta_+$  with

$$\Delta_+ = \frac{c-1}{12} + \frac{1}{2\sqrt{3}\pi}$$

## resulting sum rules:

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} (\Delta_A - \Delta_+) = v^2 \Delta_+ - uv,$$

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} j_A^2 = -u,$$

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} \left( (\Delta_A - \Delta_+)^2 - \frac{1}{6\pi^2} \right) = -v^2 \Delta_+^2 + 2uv\Delta_+ - u(1 + u) + \frac{v^2}{6\pi^2},$$

$$u = 2e^{-\beta_c}, \quad v = 1 + e^{-\beta_c}, \quad \beta_c = \sqrt{3}\pi$$

## large- $c$ limit

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} \left( \frac{\Delta_A - \Delta_+}{\Delta_+} \right)^n = v^2 (-1)^{n+1} + O\left(\frac{1}{c}\right)$$



## large- $c$ limit

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} \left( \frac{\Delta_A - \Delta_+}{\Delta_+} \right)^n = v^2 (-1)^{n+1} + O\left(\frac{1}{c}\right)$$

⇒ *Theorem*: There is at least a primary of odd spin with

$$\Delta < \Delta_+ \equiv \frac{c-1}{12} + \frac{1}{2\sqrt{3}\pi}$$

## large- $c$ limit

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} \left( \frac{\Delta_A - \Delta_+}{\Delta_+} \right)^n = v^2 (-1)^{n+1} + O\left(\frac{1}{c}\right)$$

⇒ *Theorem*: There is at least a primary of odd spin with

$$\Delta < \Delta_+ \equiv \frac{c-1}{12} + \frac{1}{2\sqrt{3}\pi} \text{ remarkably close to } \Delta_{BTZ} = \frac{c-1}{12}$$

## large- $c$ limit

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} \left( \frac{\Delta_A - \Delta_+}{\Delta_+} \right)^n = v^2 (-1)^{n+1} + O\left(\frac{1}{c}\right)$$

⇒ *Theorem*: There is at least a primary of odd spin with

$$\Delta < \Delta_+ \equiv \frac{c-1}{12} + \frac{1}{2\sqrt{3}\pi} \text{ remarkably close to } \Delta_{BTZ} = \frac{c-1}{12}$$

✱ This theorem can be proved by *reductio ad absurdum*, i.e. negating the theorem implies no solution of the above equations (see FG arXiv:1908.00029 for a complete proof)

## large- $c$ limit

$$\sum_A (-1)^{j_A} N_A e^{-\beta_c \Delta_A} \left( \frac{\Delta_A - \Delta_+}{\Delta_+} \right)^n = v^2 (-1)^{n+1} + O\left(\frac{1}{c}\right)$$

⇒ *Theorem*: There is at least a primary of odd spin with

$$\Delta < \Delta_+ \equiv \frac{c-1}{12} + \frac{1}{2\sqrt{3}\pi} \text{ remarkably close to } \Delta_{BTZ} = \frac{c-1}{12}$$

\* This theorem can be proved by *reductio ad absurdum*

**Hint:** Assume that the sum of the first two terms has the same sign of the total sum, i.e. ( $a$ : even spin term,  $b$ : odd spin term,  $k > 0$  multiplicity ratio)

- 1  $a > k b$
- 2  $a^2 < k b^2$
- 3  $a^3 > k b^3$

if  $b > 0$  ⇒

$$1 \times 3 = a^4 > k^2 b^4$$

while

$$2 \times 2 = a^4 < k^2 b^4$$

# Conclusions

- 1 There is an infinite set of modular bootstrap equations so far neglected
- 2 They generate an upper bound  $\Delta_0$  for the first non-trivial primary which is remarkably close to the BTZ black hole threshold
- 3 Even a modest improvement of the upper bound could push  $\Delta_0$  down  $\Delta_{BTZ}$  implying that pure Einstein gravity in  $AdS_3$  do not exist as a quantum theory
- 4 Further study of the new set of equations would give important information on the spectrum of primaries in a general  $2d$  CFT