

# Going Relativistic with Lattice Boltzmann Methods

### A. Gabbana, D. Simeoni, S. Succi, R. Tripiccione

Università degli studi di Ferrara Bergische Universität Wuppertal University of Cyprus

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### Describing a fluid: from micro to macro scale

Scale Microscopic (Å) Mesoscopic  $(nm - \mu m)$ Macroscopic (> mm) **Equations** Newton Boltzmann (BE) Navier-Stokes

Unknowns  $\mathbf{x}(t), \mathbf{v}(t)$  $f(\mathbf{x}, \mathbf{v}, t)$  $\rho, \mathbf{u}, T$ 

Lattice Boltzmann Methods exploit the dynamic at the mesoscopic scale to provide a correct description of the fluid at the macroscopic level



### Boltzmann equation

Boltzmann equation with BGK collisional model:

$$\partial_t f + \mathbf{v} \cdot \nabla f = \Omega(f, f^{eq}) = -\frac{1}{\tau}(f - f^{eq})$$

Relaxation to Maxwell-Boltzmann distribution:

$$f^{eq} \propto \exp\left(-rac{m}{2k_BT}(\mathbf{v}-\mathbf{u})^2
ight)$$

Macroscopic quantities as the first moments of f

$$\begin{split} \rho(\mathbf{x},t) &= \int f(\mathbf{x},\mathbf{v},t) d\mathbf{v} \\ \mathbf{u}(\mathbf{x},t) &= \frac{1}{\rho(\mathbf{x},t)} \int \mathbf{v} f(\mathbf{x},\mathbf{v},t) d\mathbf{v} \end{split}$$

m 
u particle velocity m u macroscopic fluid velocity au relaxation time ho fluid density

### Lattice Boltzmann Methods

#### Velocity Discretization

Employ a Gauss Quadrature rule to replace  $\boldsymbol{v}$  with a (small) set of velocities  $\boldsymbol{c}_i$  such that:

- **c**<sub>i</sub> connects two sites in a Cartesian grid (perfect streaming)
- Equality holds for the following:



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## RLBM: Relativistic Lattice Boltzmann Method

Relativistic extension of Lattice Boltzmann Methods to simulate relativistic hydrodynamics

- Preserves all advantages of classic LB methods
- Perfect streaming: measures not affected by interpolation schemes
- Equipped to simulate fluids in 1, 2, 3 spatial dimensions



## Relativistic Boltzmann Equation

▶ Relativistic Boltzmann Equation with Anderson-Witting collisional model:

$$p^{lpha} rac{\partial f}{\partial x^{lpha}} = \Omega(f, f^{eq}) = -rac{p^{\mu} U_{\mu}}{c^2 au} (f - f^{eq})$$

Maxwell Juettner distribution:

$$f^{eq} \propto \exp\left(-rac{p^{\mu}U_{\mu}}{k_BT}
ight)$$

Macroscopic quantities as the first moments of f

$$N^{\alpha} = c \int f p^{\alpha} \frac{d\mathbf{p}}{p_0} \qquad \qquad T^{\alpha\beta} = c \int f p^{\alpha} p^{\beta} \frac{d\mathbf{p}}{p_0}$$

 $\begin{array}{c|c|c|c|c|c|c|} \blacktriangleright & \text{Relativistic Parameter } \zeta = \frac{mc^2}{k_BT}: \\ & & \text{ultra-relativistic} & \text{mildly relativistic} & \text{non-relativistic} \\ & & \zeta \ll 1 & & \zeta \sim 1 & & \zeta \gg 1 \end{array}$ 

 $p^{lpha}$  particle four-momentum  $U^{lpha}$  macroscopic four-velocity au relaxation time

## RLBM: Relativistic Lattice Boltzmann Method

 Expand f<sup>eq</sup> in a basis of orthonormal polynomials J<sup>(k)</sup> up to an order N capable of recovering the hydrodynamic quantities of interest:

$$f^{eq}(\boldsymbol{\rho}^{\mu}, \boldsymbol{U}^{\mu}, t) = \omega \sum_{k=0}^{N} \mathbf{a}^{(k)}(\boldsymbol{U}^{\mu}, t) \cdot \mathbf{J}^{(k)}(\boldsymbol{\rho}^{\mu})$$

2. Using the following Gauss Quadrature rule

$$\int \omega(\mathbf{p}) \mathbf{J}^{(k)}(\mathbf{p}^{\mu}) \mathbf{J}^{(\ell)}(\mathbf{p}^{\mu}) \frac{d\mathbf{p}}{\mathbf{p}_{0}} = \delta_{\ell k} = \sum_{i} w_{i} \mathbf{J}^{(k)}(\mathbf{p}_{i}^{\mu}) \mathbf{J}^{(\ell)}(\mathbf{p}_{i}^{\mu})$$

determine a discrete set of momenta  $p_i^{\mu}$  that preserves the hydrodynamic quantities of interest:

$$N^{\alpha} = c \int f p^{\alpha} \frac{d\mathbf{p}}{p_0} = \sum_{i} w_i f_i p_i^{\alpha} \qquad T^{\alpha\beta} = c \int f p^{\alpha} p^{\beta} \frac{d\mathbf{p}}{p_0} = \sum_{i} w_i f_i p_i^{\alpha} p_i^{\beta}$$

3. Derive the discrete relativistic Boltzmann equation via an Euler scheme:

$$f_i(\mathbf{x} + \mathbf{v}_i \Delta t, t + \Delta t) - f_i(\mathbf{x}, t) = -\Delta t \frac{p_i^{\mu} U_{\mu}}{p^0 \tau} (f_i - f_i^{eq})$$

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## Numerical Validation: The Riemann Problem

- Commonly used benchmark in hydrodynamics
  - It has a semi-analytical solution in the ultra-relativistic & non-dissipative case
- Possible to compare against other numerical methods in the mildly-relativistic / dissipative case



#### A few comparisons



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#### Connection between scales

- Chapman-Enskog's (CE) and Grad's expansions connect mesoscopic to macroscopic scale
- ▶ They link macroscopic transport coefficients  $\lambda$ ,  $\mu$ ,  $\eta$  to the kinetic ones, in our case the relaxation time  $\tau$

$$\lambda = c^2 k_B n f_1(\zeta) \tau$$
  $\eta = P f_2(\zeta) \tau$   $\mu = P f_3(\zeta) \tau$ 

They lead to divergent results in relativistic theories

#### CE analysis

Multi-scale expansion based on weak departure from local equilibrium:

$$f = f^{eq} + \phi(\mathbf{x}, t) f^{eq}$$
 with  $\phi(\mathbf{x}, t) \sim O(Kn)^1$   $Kn = \frac{\lambda}{L}$ 

#### Grad's method

Expansion of the particle distribution f in Hilbert space of the momentum vector  $p^{\alpha}$  around the equilibrium  $f^{eq}$ 

$$f = f^{eq} (1 + a_{\alpha} p^{\alpha} + a_{\alpha\beta} p^{\alpha\beta} + ...)$$

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#### Analytical comparison between the two methods: $\eta$



#### Numerical comparison between the two methods: $\eta$



#### Comparison between the two methods: $\eta$ , $\mu$ , $\lambda$

The analysis has been extended to all transport coefficients in all spatial dimensions



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► Free electrons in graphene behave like a plasma of charged particles with a *photon-like* relativistic dispersion relation:  $E_k = \hbar v_f k$ 



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- Experimental observation of anomalous voltage drop near current injectors which originate the formation of whirlpools in the electron flow



- Free electrons in graphene behave like a plasma of charged particles with a *photon-like* relativistic dispersion relation:  $E_k = \hbar v_f k$
- Electron-electron collisions dominant over electron-phonon scattering even at room temperature
- Experimental observation of anomalous voltage drop near current injectors which originate the formation of whirlpools in the electron flow
- Theoretical study of pre-turbulent regimes at low Reynolds



# Outlook: Quark Gluon Plasma



# Sum up and Conclusions

Sum up:

- Developed a fast numerical scheme for the simulation of Relativistic Hydrodynamics
- ► Validated the code with the Riemann problem Benchmark
- Created a firm connection between the kinetic and the hydrodynamic layer via the Chapmann-Enskog expansion
- Correctly reproduced experiments of electron flows in graphene

#### Future steps:

- Introduce external magnetic field in graphene simulations to check for early occurrence of turbulent regimes
- Start using RLBM for Quark Gluon Plasma simulations

## Some bibliography

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# Thank you for your attention!

### RLBM research group:

- ► Alessandro Gabbana<sup>1,2</sup>
- Daniele Simeoni<sup>1,2,3</sup>
- Sauro Succi<sup>4,5</sup>
- Raffaele Tripiccione<sup>1</sup>

- 1. Universita' di Ferrara and INFN-Ferrara, Ferrara, Italy
- 2. Bergische Universitaet Wuppertal, Wuppertal, Germany
- 3. University of Cyprus, Nicosia, Cyprus
- 4. Center for Life Nano Science @ La Sapienza, Italian Institute of Technology, Roma, Italy
- 5. Istituto Applicazioni del Calcolo, National Research Council of Italy, Roma, Italy



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Daniele Simeoni