# CYGNUS TPC module with Optical readout

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#### **Galactic Nuclear Recoil Observatory** identify with **directional** sensitivity clear **dark matter** candidate and **coherent neutrino** scattering from the Sun and other sources

- 1. to establish the science case for CYGNUS, working with external experts as required
- 2. to establish the **feasibility** and **technology** choices for CYGNUS, coordinating R&D activities, resources and joint publications as necessary
- 3. to form an Institute **Board** including remit to prepare an organisational structure in readiness for launch of the collaboration
- 4. to write an **experiment LOI** as basis for formation of the collaboration based on (1-3)
- 5. to launch and follow the **international collaboration**



### **CYGNUS** collaboration objective





G. Mazzitelli, 2018 IEEE Nuclear Science Symposium, Sydney, Australia



# The CYGNUS TPC challenging to goal

- Energy threshold 1 keVee
- Target mass 100-1000 kg (F, He)
- Zero neutron background

  - -ceramics; almost no internal electronics
- x, y, z fiducialisation and radon rejection
  - -either negative ion drift or other technique
  - -material selection and scrubbing is not enough
- Gamma discrimination below 10 keVee
- Directional sensitivity



-no steel (vacuum) vessel (acrylic?)



# ...approaching the neutrino floor



preliminary estimation for CYGNUS HeSF6 very



## DM detection strategy

directionality



particle Identification (signal, background)

Ge, Si: SuperCDMS Ge: EDELWEISS

gassous low (KeV) threshold ideas to increase sensitivity and scalability C<sub>3</sub>F<sub>8</sub>, CF<sub>3</sub>I: PICO Ge: CoGeNT, CDEX SI: DAMIC CF<sub>4</sub>: DRIFT, DMTPC, MIMAC, Newage

LXe: XENON, LUX, PandaX LAr: ArDM, DarkSide-50

Charge

#### solid, cryogenic very low (10-100 eV) threshold limited mass and scalability

CaWO<sub>4:</sub> CRESST

Nal: DAMA/LIBRA Csl: KIMS

LXe: XMASS LAr: DEAP-3600

liquid, cryogenic medium (MeV) threshold high sensibility and scalability

Light

picture M. Messina



# Why gaseous TPC?

#### **Time Projection Chambers** provide:

- 3D tracking (position and direction);
- total released energy measurement;
- dE/dx profile (pid, head-tail);
- reduced readout channel number;

**gas** represents an interesting target:

- nuclei free path can be long enough to be reconstructed;
- low mass gases allow an efficient momentum transfer from light DM;
- avalanche mechanism allows a **sensitivity** to single primary electrons (i.e. energy release of 30-40 eV);

#### Nuclear recoil in gas





#### CYGNO: a CYGNUs Collaboration 1 m<sup>3</sup> Module with Optical **Readout for Directional Dark Matter Search**

#### objectives:

- demonstrating the validity of technology
- understanding noise and background due to radioactivity
- performing neutron LNGS flux measurements and contribute to directional dark matter upper limit measurements with a different systematics

... and moreover participating to a new international rising community for sub GeV DM candidate measurement







# The CYGNO demonstrator strategy

- high granularity (CMOS+PMT) optical read out:
  - threshold
  - discrimination
  - directionality;
  - x, y, (z) fiducialisation
  - electronics decoupling
- atmospheric pressure He gas mixture:
  - low target density (low threshold)





Multiple GEM structures can be used to share the gain and make more stable detectors.



- high sensitivity;
- operation and lower gas contamination);
- suitable lens allow to acquire large surfaces with small sensors;

### **Optical Read Out**



During the multiplication process, photons are produced along with electrons by the gas through atomic and molecular de-excitation;

optical sensors are able to provide high granularities along with very low noise level and

optical coupling allows to keep sensor out of the sensitive volume (no interference with HV



## CYGNO Roadmap @ LNGS



• OPT readout

• 1 cm drift

• 0 resolution

- 3D printing
- 20 cm drift
- PID
- tracking
- drift resolutions
- •
- 50 cm drift
- drift scaling
- materials test
- gas

















CMOS: low noise; high granularity (2048×2048 pixel /  $125 \times 125 \ \mu m^2$ ) and high sensitivity equipped with optics f/0.95-25 mm



pair created in the tail of an electromagnetic shower (0.2 T magnetic field)

Electron

Positron

electron from natural radioactivity

(thanks to nature...)



#### (thanks to Frascati Beam Test Facility BTF)



#### AmBe Neutron source (thanks to FISMEL)





## XY resolution in ORAnGE



About 330 detected photons per track millimetre (for VGEM = 440V), i.e. 230 eV released in gas (from Garfield). XY resolution =  $35 \,\mu m$ 









## Z resolution in ORAnGE







**Z** resolution

0





# Energy resolution <sup>55</sup>Fe (5.9 keV)











9078 ph / 5900 eV -> **1.5 ph/eV**.

60

pedestal jitter 120 ph @ 5 sigma 600 ph -> 600/1.5 -> Th = 400 eV

(pedestal based on average of the run within Fe source in the detector)

Energy resolution ~ 24% ~ 1.5 keV





# Energy resolution <sup>55</sup>Fe (5.9 keV)

An energy resolution between 20% and 30% is achieved for releases of 5.9 keV;



With 60/40 at 440 V, a resolution of 1.3 keV was measured for <sup>55</sup>Fe signals; For 1 keV energy releases a resolution of hundreds of eV is achieved.



## PID performances







## Range and recoil energy released

alpha particles path in He gas detector prototype





alpha range seems to be "determined" only by CF<sub>4</sub> and to decrease linearly with its amount

AmBe Orange GEM 440/2kVcm - drift 1kV/cm - 2000 ms









### LEMOn prototype



# LEMOn prototype design

A new prototype with 7 litre sensitive volume (LEMOn: Large Elliptical Module Optically readout) was built in 2017 tested on electron beam in July.

MANANN



20x24 cm<sup>2</sup> GEMs



Elliptical field cage with semi-transparent cathode





CMOS camera

Design: S. Tomassini, acknowledgment: T. Napolitano & F. Angeloni Servizio Progettazione e Costruzioni Meccaniche LNF-INFN G. Cooradi and D. Tagnani Servizio Elettronico e Automazione LNF-INFN A. Orlandi e M. Iannarelli

G3 ###



### XY and energy resolution

![](_page_22_Figure_1.jpeg)

![](_page_22_Figure_2.jpeg)

resolution obtained with a drift field to 0.6kV

![](_page_22_Picture_5.jpeg)

![](_page_22_Figure_6.jpeg)

### diffusion in HeCF4 for electrons

![](_page_23_Figure_1.jpeg)

![](_page_23_Picture_4.jpeg)

![](_page_23_Figure_5.jpeg)

#### **LEMON** data @ Frascati BTF

![](_page_23_Picture_8.jpeg)

## Z resolution in LEMOn

Electron diffusion in the drift gap can be exploited to evaluate the Z of the event. The transverse light profile and the PMT signal waveform are expected to become lower and larger as long as the event is far from the GEM;

Since the amplitude (A) decreases and the width (S) increases with Z, their ratio  $\eta = S/A$  increases (independently from the amount of produced light);

![](_page_24_Figure_3.jpeg)

both methods gives 10% precision:  $\sigma z \sim 2 \text{ cm} @ 20 \text{ cm}$ 

![](_page_24_Picture_6.jpeg)

![](_page_24_Figure_7.jpeg)

# Energy resolution <sup>55</sup>Fe (5.9 keV)

![](_page_25_Figure_1.jpeg)

![](_page_25_Picture_3.jpeg)

pedestal jitter 15 ph -> 75/0.12 -> **Th = 625 eV** 

(pedestal based on average of the run within Fe source in the

Energy resolution ~ 33% ~ 2 keV

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![](_page_25_Figure_9.jpeg)

![](_page_25_Picture_10.jpeg)

![](_page_25_Picture_11.jpeg)

#### 5s exposure @ 2.45 MeV neutrons Frascati Neutron Generator

![](_page_26_Picture_2.jpeg)

test beam 18-20 June 2018 (tanks to FNG)

![](_page_26_Picture_5.jpeg)

![](_page_26_Picture_6.jpeg)

#### 0.1s exposure @ 2.45 MeV neutrons Frascati Neutron Generator

![](_page_26_Figure_8.jpeg)

Longitudinal light profile shows a typical Bragg peak shape

## LEMOn @ ENEA FNG

![](_page_27_Picture_1.jpeg)

![](_page_27_Picture_2.jpeg)

550 [ph/mm] / 0.06 [ph/eV] —> 9.2 keV/mm (125 µm resolution) ~ 0.575 MeV recoil He energy

![](_page_27_Picture_6.jpeg)

### **CYGNO Detector**

#### Conceptual design of CYGNO

#### Author List

June 28, 2018

#### Abstract

The design of the project named CYGNO (a CYGNUS 1m<sup>3</sup> module with Optical readout) is described here. CYGNO aims to make significant advances in the technology of single phase gas-only time projection chambers for the specific application of rare scattering events detection. In particular it will focus on a read-out technique based on the GEM amplification of the ionization and on the visible light collection with a sub-mm position resolution sCMOS camera. This type of readout - in conjunction with a fast light detection - will allow to reconstruct 3D images of the recoiling particles, offering new ways to distinguish the electron and nuclear recoils. The final goal is to deliver a high resolution 1 cubic meter detector for underground neutron flux measurements that - with proper shielding and accurate choice of the materials - can be a prototype for a dark matter (DM) detector. The recoil direction resolution is also being investigated as a further tool to reject background in the detection of Galactic DM particles. This project is part of the world-wide effort of the CYGNUS collaboration to define an optimal DM detection scheme sensitive to DM direction, towards a one-ton gas TPC nuclear recoils observatory.

A total of 72 10<sup>6</sup> readout 165 x 165  $\mu$ m<sup>2</sup> pixels.

18 cameras monitoring 330\*330 mm each with 160 mµ resolution and a sensitivity of ~ 1 ph / 20 eV released in gas

![](_page_28_Picture_9.jpeg)

# **CYGNO Physics**

a 1.6kg-1m<sup>3</sup> of HeCF4 with a threshold of few 1keVe could exploit some interesting area in sub GeV DM search, but before that, CYGNO have to dimostrate with an emerging technology a full compression of: **background; materials; gas purification; stability; scalability and reliability** 

E interval	Thermal Neutron Flux (10 <sup>-6</sup> cm <sup>-2</sup> s <sup>-1</sup> )							
(eV)	Ref. [21]	Ref. [22]	Ref. [23]	Ref. [24]				
0 - 0.05	$5.3 \pm 0.9$	$1.08\pm0.02$	$0.54 \pm 0.13$	$0.32 \pm 0.09$				
		$(1.07 \pm 0.05)$						
0.05 - 1000		$1.84 \pm 0.20$						
		$(1.99 \pm 0.05)$						

Table 2: Thermal and epithermal (top) and fast (bottom) neutron flux measurements at the Gran Sasso laboratory reported by different authors. In analyzing their experimental data with Monte Carlo simulations, Belli et al. [22] have used two different hypothetical spectra: flat, and flat plus a Watt fission spectrum. This leads to the upper and lower data sets shown for ref.[22] respectively.

 			-				
E interval	Fast Neutron Flux (10 <sup>-6</sup> cm <sup>-2</sup> s <sup>-1</sup> )						
(MeV)	Ref. [25]	Ref. [26]	Ref. [22]	Ref. [21]	Ref. [27]	Ref. [28]	
0.1 - 1			$0.54 \pm 0.01$				
1 - 2.5		$0.14{\pm}0.12$	$(0.53 \pm 0.08)$				
2.5 - 3		$0.13 \pm 0.04$	$0.27 \pm 0.14$				
3 - 5			$(0.18 \pm 0.04)$			$2.56 \pm 0.27$	
5 - 10		$0.15 \pm 0.04$	$0.05 \pm 0.01$				
			$(0.04 \pm 0.01)$	$3.0 {\pm} 0.8$	$0.09 \pm 0.06$		
10 - 15	$0.78 \pm 0.3$	$(0.4 \pm 0.4) \cdot 10^{-3}$	$(0.6 \pm 0.2) \cdot 10^{-3}$				
			$((0.7 \pm 0.2) \cdot 10^{-3})$				
15 - 25			$(0.5 \pm 0.3) \cdot 10^{-6}$				
			$((0.1 \pm 0.3) \cdot 10^{-6})$				

![](_page_29_Figure_6.jpeg)

on the other hand CYGNO can provide a low noise **seasonal measurement of the neutron flux** intensity and spectrum at LNGS, presently highly desirable in order to check the current results and to **reduce the experimental errors.** 

![](_page_29_Figure_9.jpeg)

## in the meantime... LIME & MANGO

**LIME**: Long Imaging ModulE

- **50 cm** long drift gap, 25 litre sensitive volume
- studied material choice
- performing a detailed study, minimisation and simulation of radioactive background;
- gas re-circulation and purification.
- optimisation of **PMT/SiPM** readout and trigger.
- HV Test

Tomassini

S

atelli &

Ros

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5 cm drift gap • THGEM test • 4 GEM test (lower HV with the same gain) Negative lon test

31

![](_page_30_Picture_13.jpeg)

**MANGO** - Multipurpose Apparatus for Negative ion studies with GEM and Optical readout

![](_page_30_Picture_16.jpeg)

## Conclusion 1/2

- TPC based on **optically readout GEM** demonstrated very interesting performance:
  - X-Y resolution around 100 µm;
  - 20%-30% precision on the evaluation of released energy already in the keV range; • effect of electron diffusion can be exploited to determine the track depth with a 12%
  - uncertainty;
- use of a **PMT** for a combined allows:
  - the 3D reconstruction of single clusters with 100 µm resolution;
  - a concurrent evaluation of the event Z with 12% uncertainty;
  - detection of light produced by the particle crossing to get the t0 of the event.
- tracking and PID looks to be very promising tools to lower the energy threshold and identify signal respect to background

![](_page_31_Picture_12.jpeg)

### Conclusions 2/2

- identify background;
- detector exploring the **optical readout** performance;
- 30-100 m<sup>3</sup> where physics is very interesting and competitive.

 The international community CYGNUS is studying and promoting a third class of Dark Matter and neutrino detector able to explore the keV region, taking in to account the **directionality** of the signals and exploiting the PID characteristic to

• 1 m<sup>3</sup> prototypes has been funded around the world, and CYGNO is the one investigating the possibility to realise a very high spatial and energy resolution

• CYGNO demonstrator has a sensitivity on the paper already of interest for Dark Matter and Neutron Flux measurements, with the aim to open a roadmap to

![](_page_32_Picture_9.jpeg)

## CYGNUS (CYGNO) Collaboration

![](_page_33_Picture_1.jpeg)

Istituto Nazionale di Fisica Nucleare, Laboratori Nazionali di Frascati, Frascati (RM), Italy; Gran Sasso Science Institute, L'Aquila, Italy; Sapienza Università di Roma, Dipartimento di Fisica, Rome, Italy; Australian National University, ACT, Canberra, Australia; Istituto Nazionale di Fisica Nucleare, Sezione di Roma, Rome, Italy; University of Sheffield, Sheffield, United Kingdom; University of New Mexico, Albuquerque, New Mexico, United States of America; Museo Storico della Fisica e Centro Studi e Ricerche Enrico Fermi, Rome, Italy; University of Hawaii, Honolulu, Hawaii, United States of America; Kobe University, Kobe, Japan; Istituto Nazionale di Fisica Nucleare, Sezione di Roma TRE, Rome, Italy; ENEA, Frascati, Frascati (RM), Italy; Università di Roma TRE, Dipartimento di Matematica e Fisica, Rome, Italy; Sapienza Università di Roma, Dipartimento di Ingegneria Chimica Materiali Ambiente, Rome, Italy; Istituto Nazionale di Fisica Nucleare, Laboratori Nazionali del Gran Sasso, L'Aquila, Italy

–Johnny Appleseed

#### "spare"

## **Electrical and light GAIN**

![](_page_35_Figure_1.jpeg)

![](_page_35_Figure_3.jpeg)

#### **MIP LEMON** GEM 455V(?)/2kVcm - drift 0.6kV/cm - 30 ms

![](_page_36_Figure_1.jpeg)

![](_page_36_Figure_4.jpeg)

![](_page_36_Picture_5.jpeg)

#### **FNG LEMON** GEM 440V/2kVcm - drift 0.6kV/cm - 100 ms

![](_page_37_Figure_1.jpeg)

![](_page_37_Picture_4.jpeg)

#### FNG LEMON GEM 440V/2kVcm - drift 0.6kV/cm - 100 ms

![](_page_38_Figure_1.jpeg)

#### 550 [ph/mm] / 0.12 [ph/eV] —> 4.6 keV/mm (125 µm resolution)

![](_page_38_Picture_5.jpeg)

![](_page_39_Figure_1.jpeg)

-> 0.12 ph/eV.

### Fe@440 for FNG comparison

![](_page_39_Figure_5.jpeg)

![](_page_39_Picture_7.jpeg)

### **Detection vs Identification**

#### no signal over threshold

![](_page_40_Figure_2.jpeg)

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**PID - 4D reconstruction** 

![](_page_40_Figure_5.jpeg)

![](_page_40_Picture_7.jpeg)

![](_page_40_Picture_8.jpeg)