

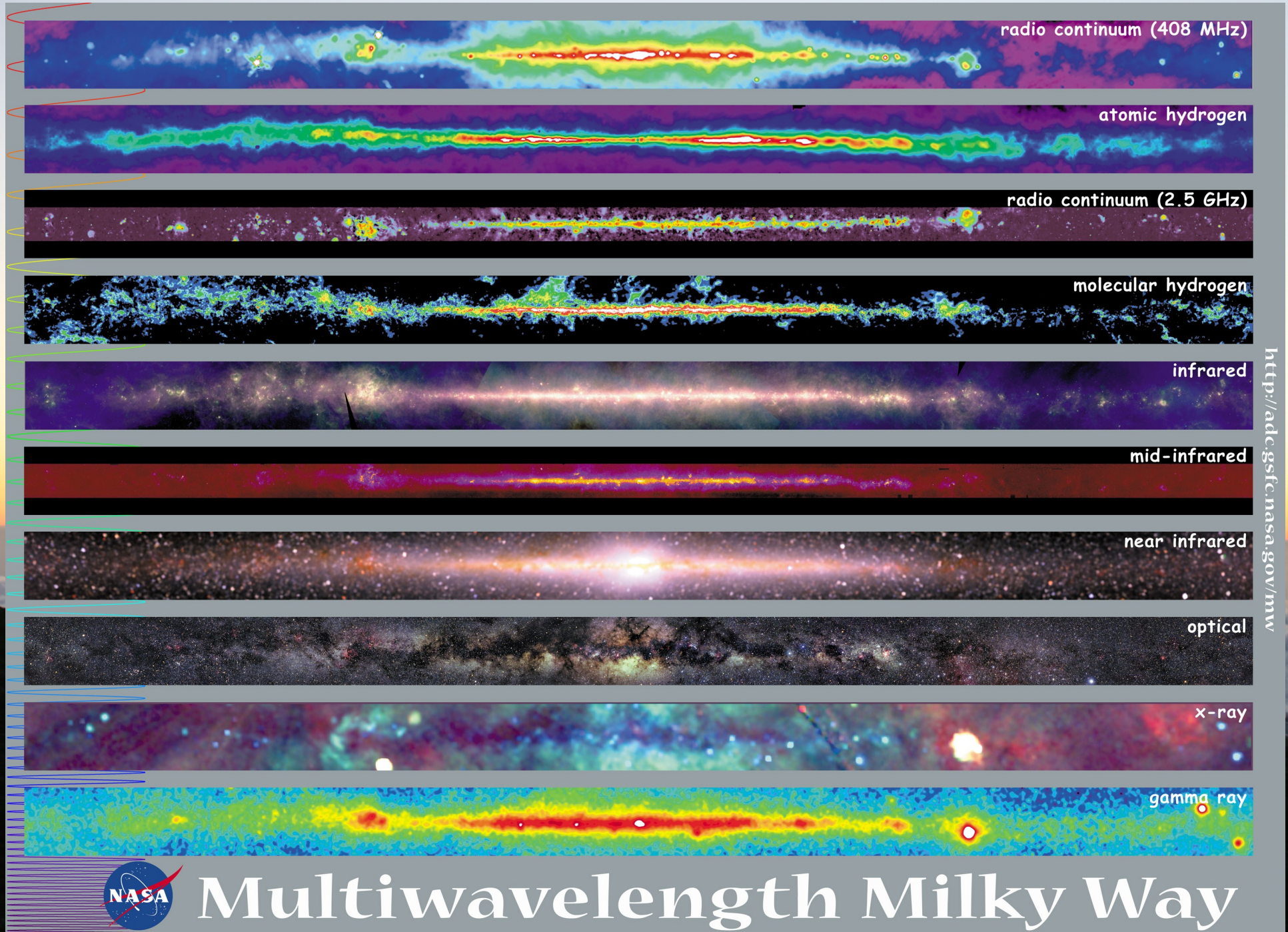
MAGIC...

...una finestra
sull'Universo

"seconda parte"

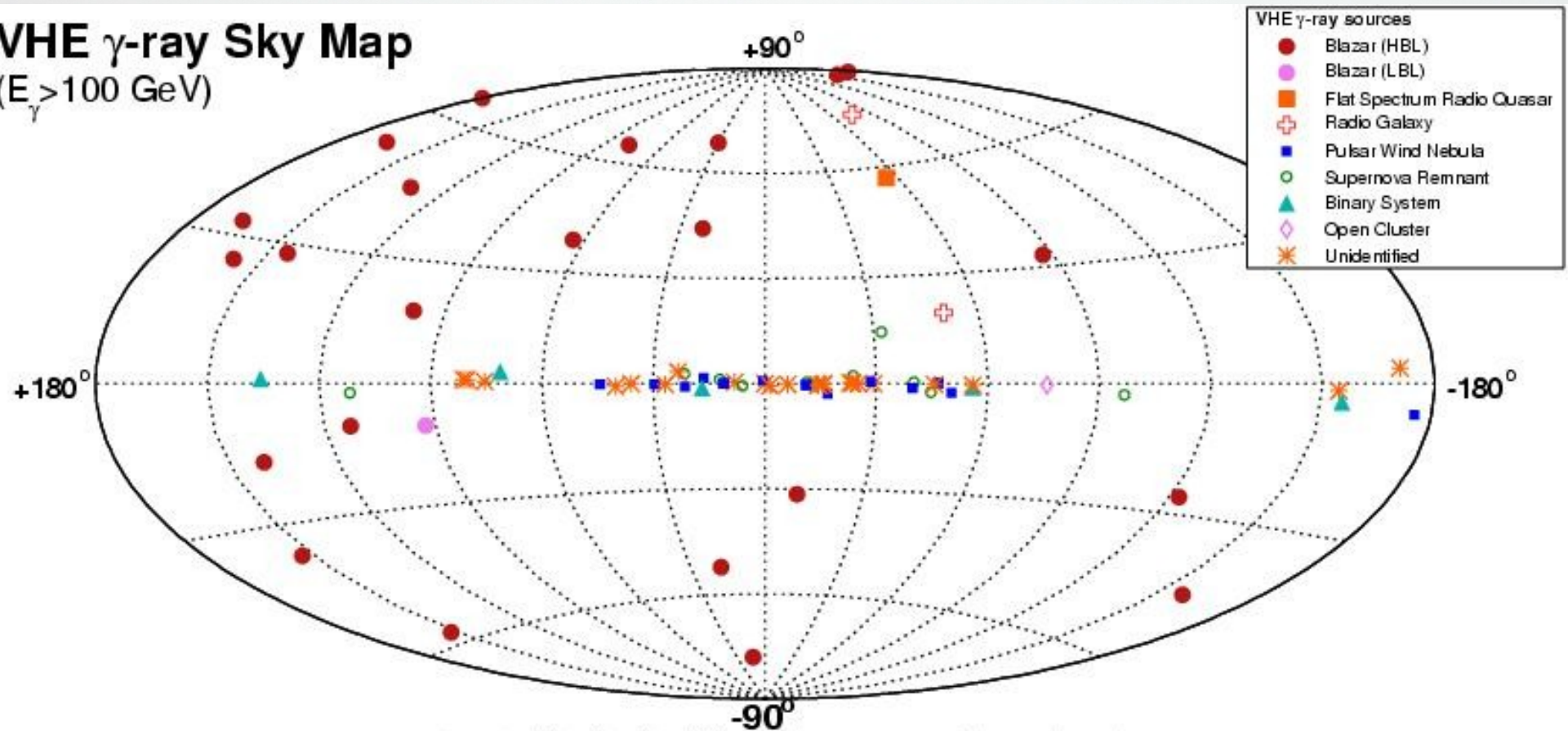


L'emissione della Via Lattea



Il cielo ad altissima energia

VHE γ -ray Sky Map ($E_{\gamma} > 100$ GeV)



2009-06-15 - Up-to-date plot available at <http://www.mppmu.mpg.de/~rwagner/sources/>

Come sono prodotti i raggi gamma?

Emissione termica (equilibrio) $\rightarrow$ esempio luce (IR) di un corpo caldo, la luce delle stelle

Emissione non termica (condizioni di NON equilibrio, spesso dovute a fenomeni relativistici, tipo alta velocità delle particelle)



I RAGGI GAMMA PROVENGONO DA OGGETTI
ASTROFISICI CON CONDIZIONI MOLTO
PARTICOLARI (intensi campi magnetici, flussi di
particelle ultra-relativistiche)

Come sono prodotti i raggi gamma?

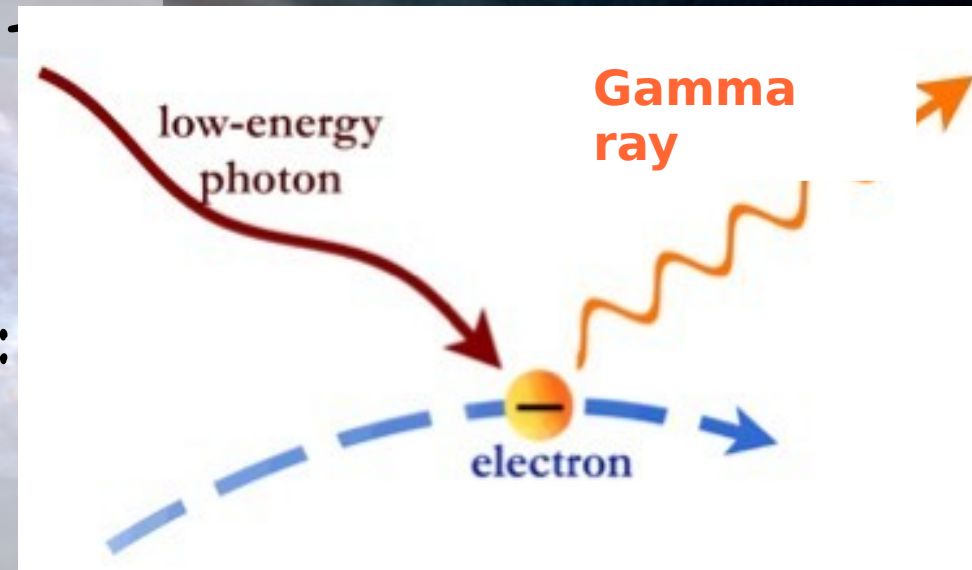
INVERSE COMPTON SCATTERING

Particella carica veloce
campo magnetico

emissione (onde radio):
sincrotrone

onda radio urta particella

RAGGIO GAMMA

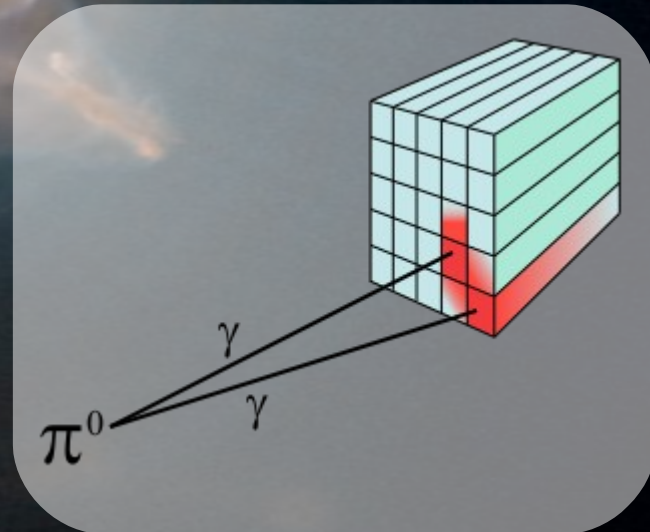


Come sono prodotti i raggi gamma?

NEUTRAL PION DECAY

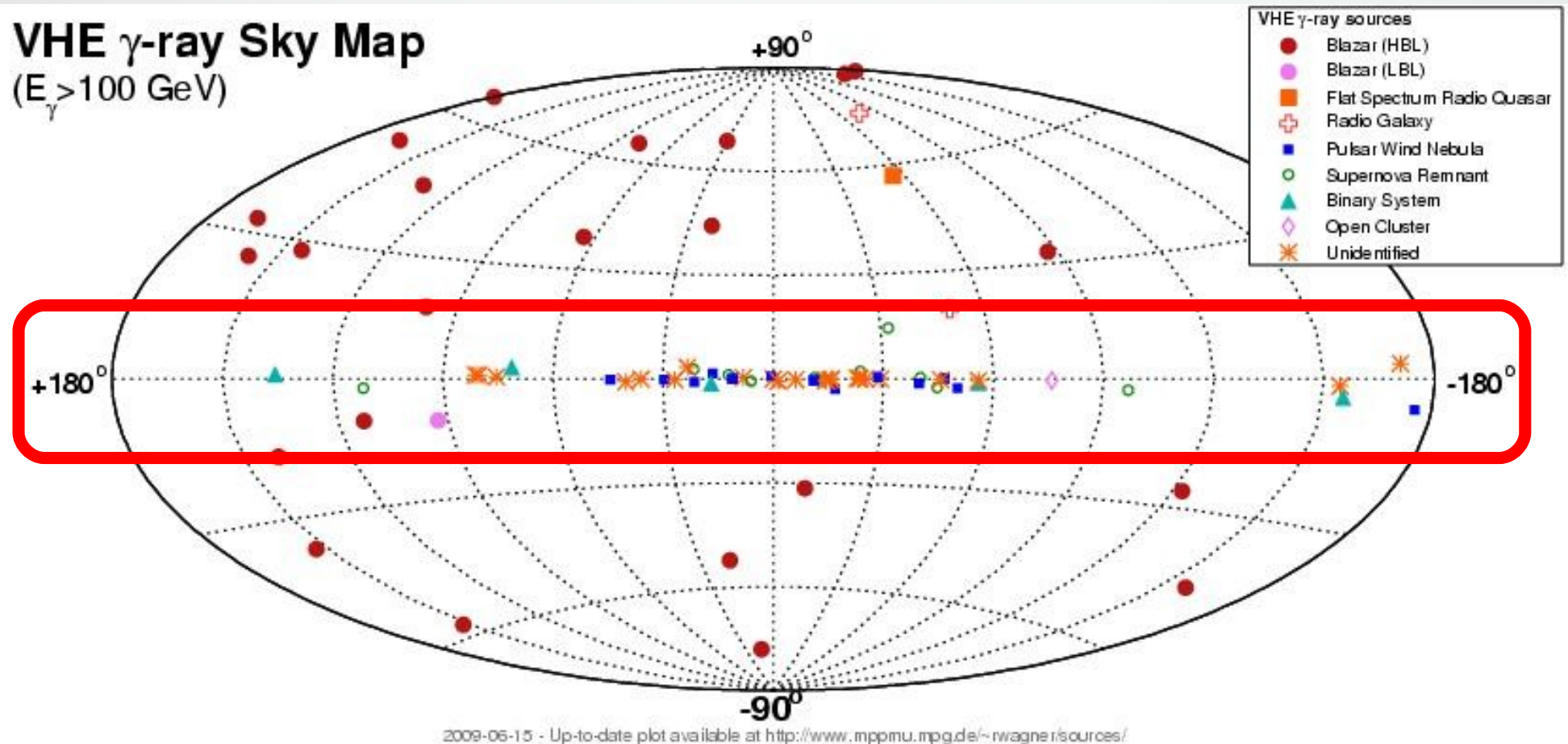


Pione Neutro (particella SUBATOMICA) ha una vita media molto breve e decade principalmente **gamma-gamma**

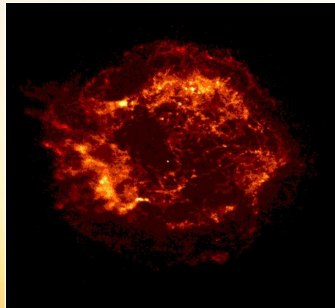


Il cielo ad altissima energia

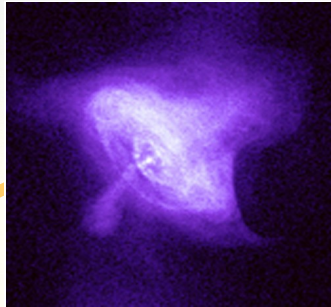
VHE γ -ray Sky Map ($E_{\gamma} > 100$ GeV)



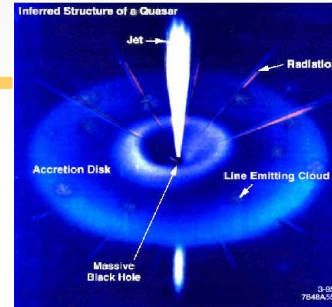
Possibili sorgenti per il Telescopio MAGIC



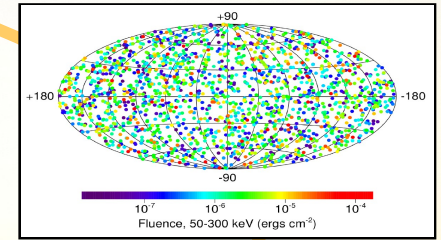
SNRs



Pulsars

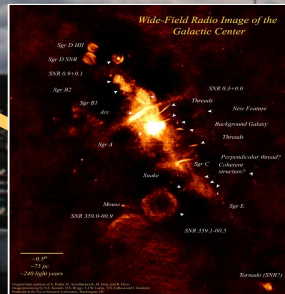


AGNs

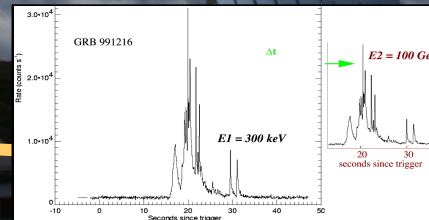


GRBs

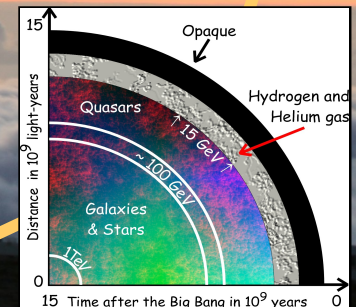
Origin of
Cosmic Rays



Cold Dark
Matter



Quantum Gravity
effects

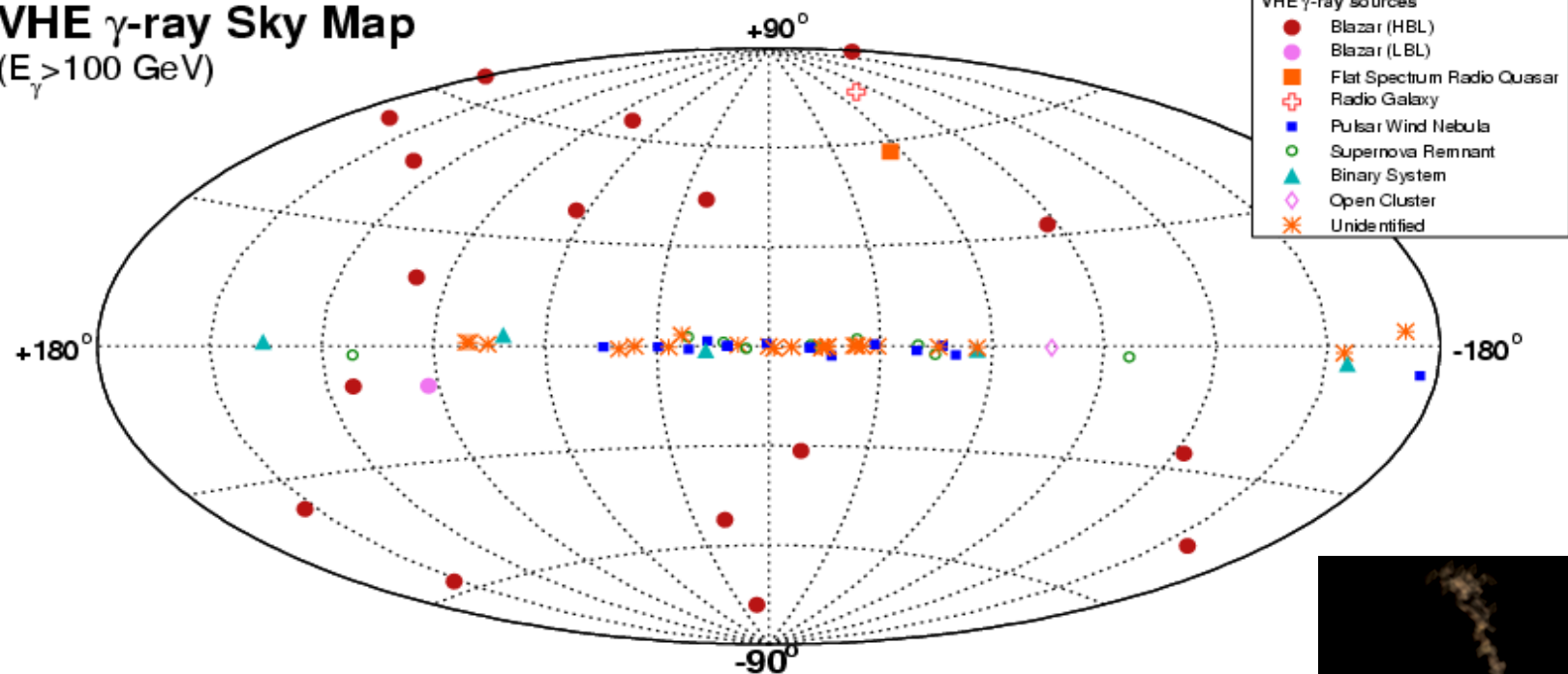


cosmological
γ-Ray Horizon

Sorgenti gamma osservate

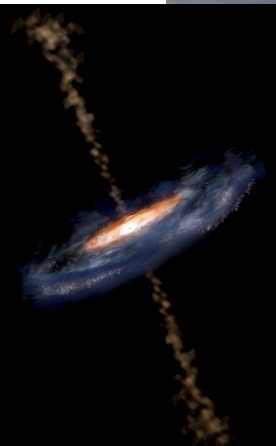
galattiche...

VHE γ -ray Sky Map
($E_\gamma > 100$ GeV)



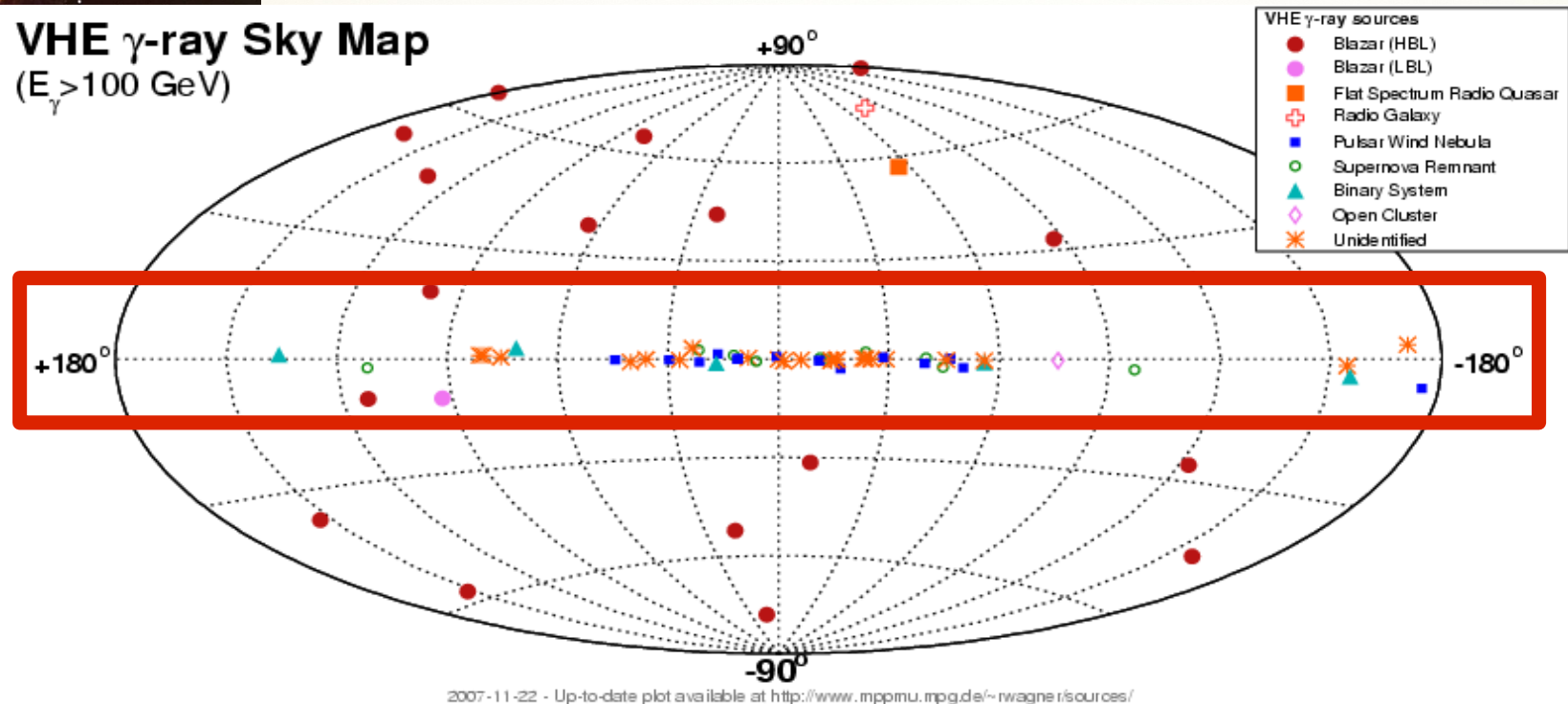
2007-11-22 - Up-to-date plot available at <http://www.mppmu.mpg.de/~rwagner/sources/>

... ed
extragalattiche



Le Sorgenti Galattiche

VHE γ -ray Sky Map
($E_\gamma > 100$ GeV)

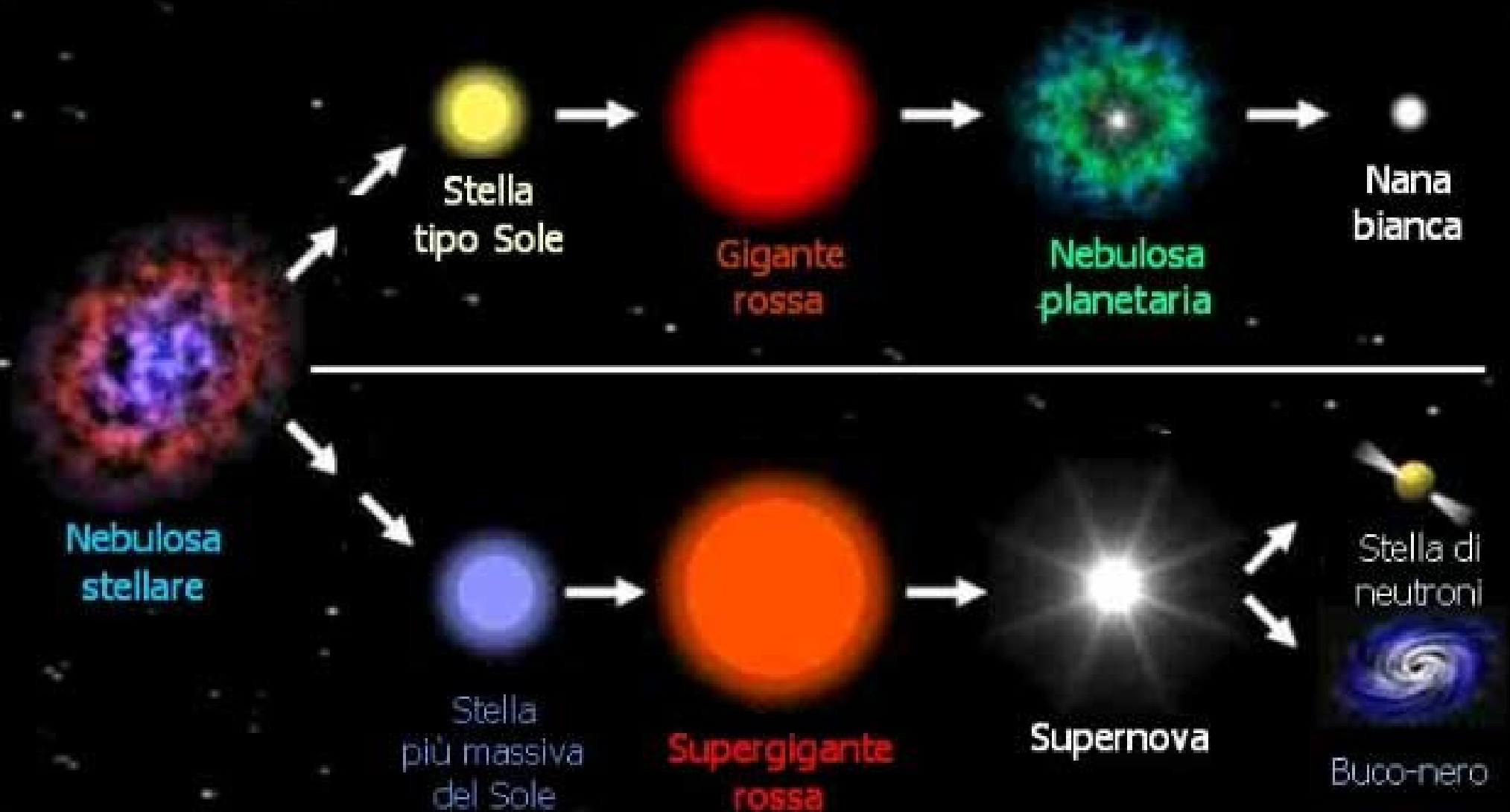


Via Lattea

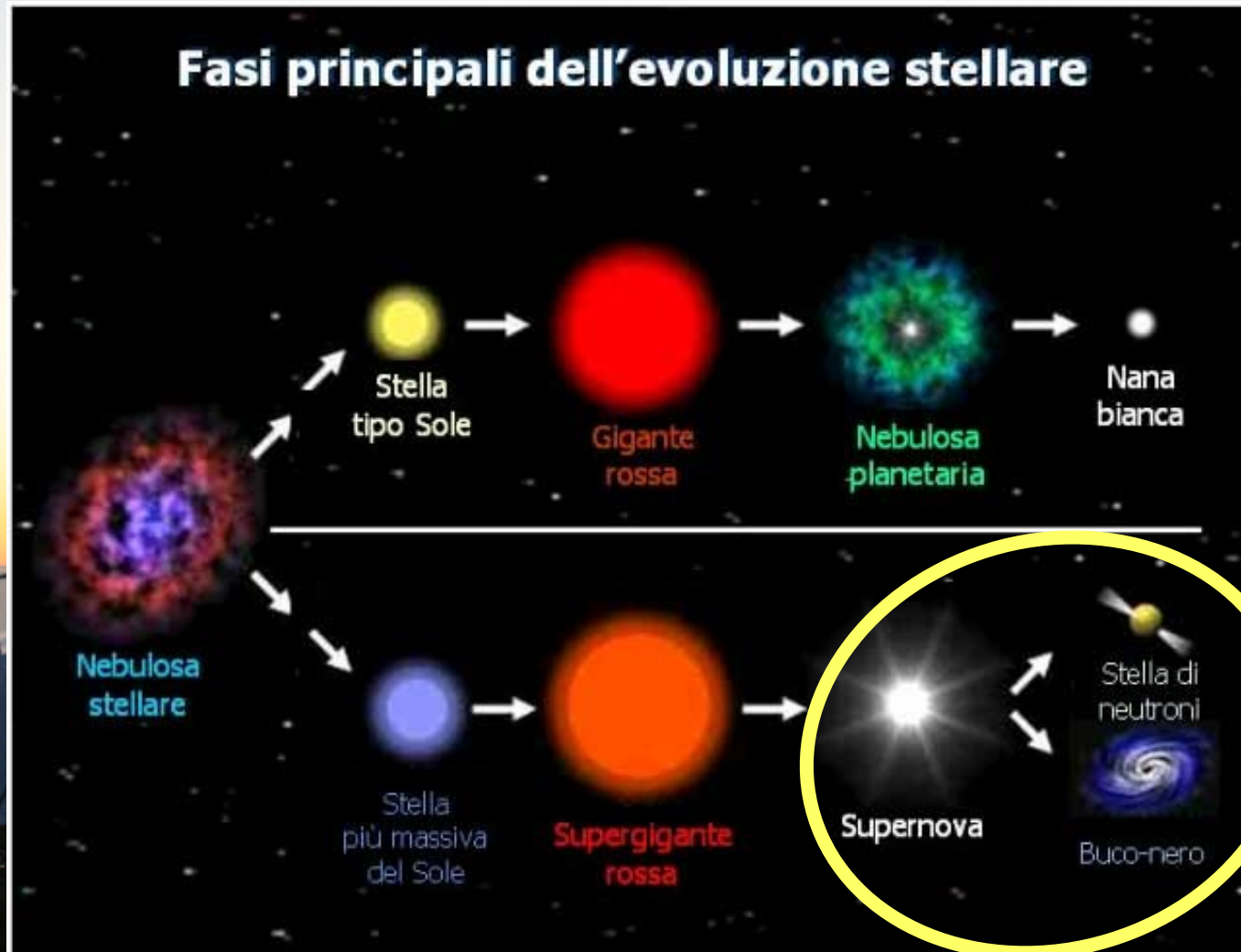


Att! Ricostruzione immaginaria!!

Fasi principali dell'evoluzione stellare



Raggi gamma ultra-energetici dalle stelle

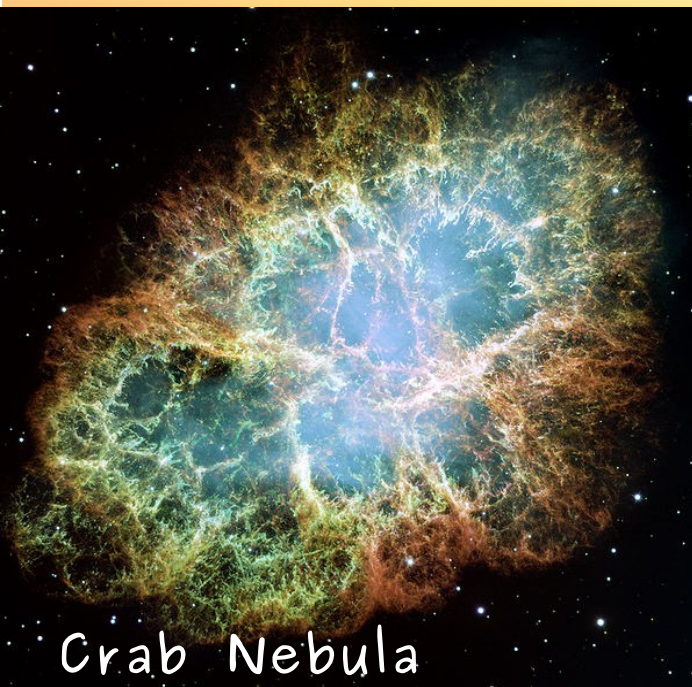


Particelle accelerate --> raggi gamma!

Supernova remnants (SNRs)



Tycho Nebula



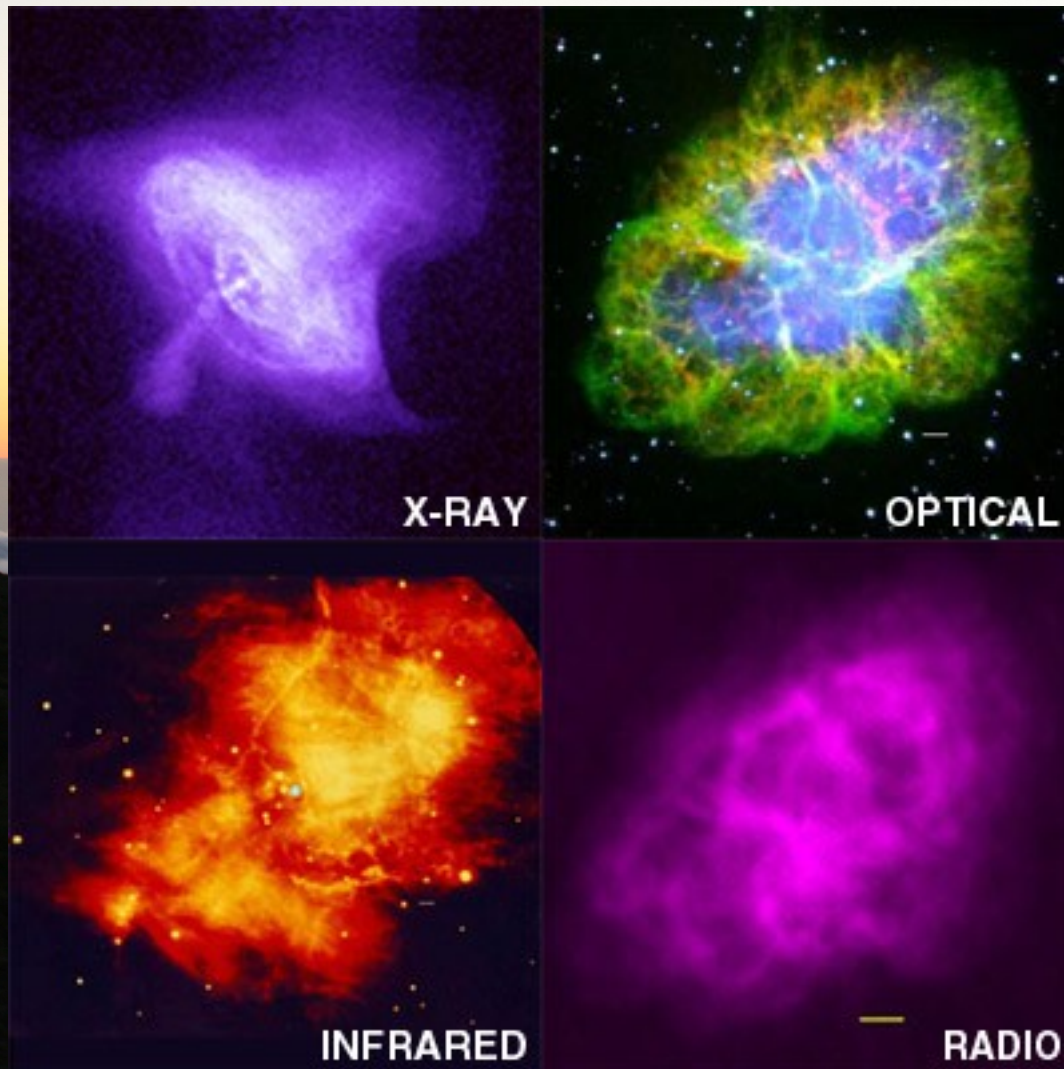
Crab Nebula

- Sono cio' che rimane da un'esplosione stellare
- Catalogate in "shell type" e "plerions (pulsar wind nebulae)"
- Onde di shock molto energetiche accelerano le particelle cariche

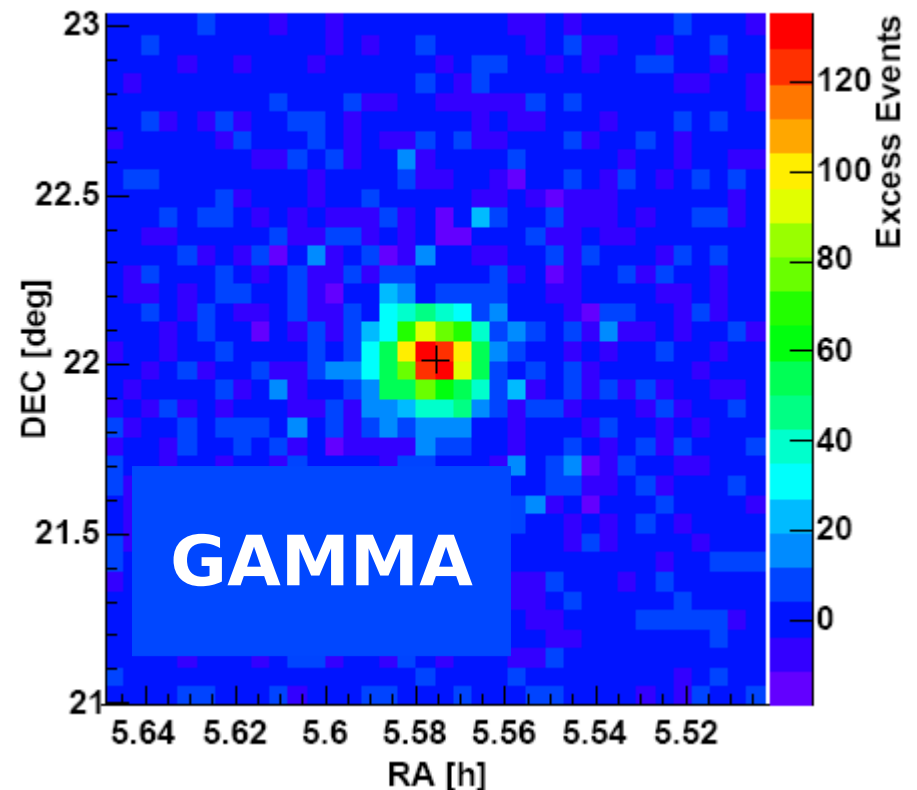


Crab Nebula: check up completo!

Ognuno fa la sua parte!
Così possiamo studiare
cosa avviene

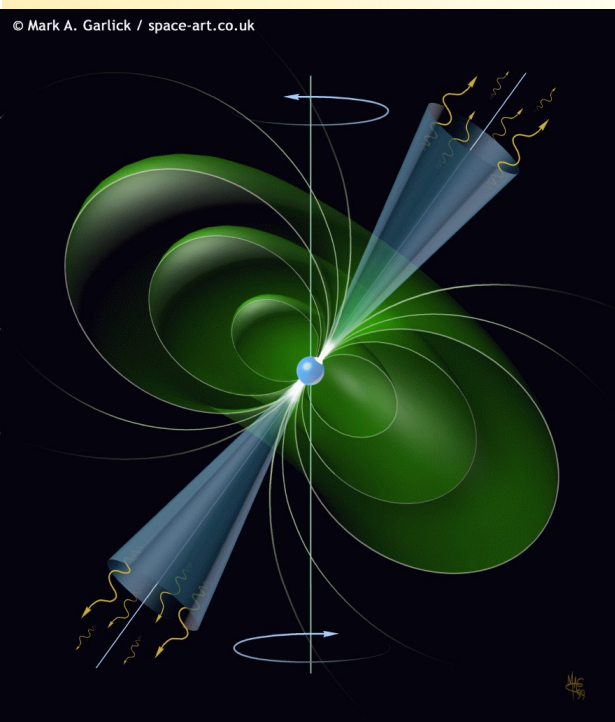
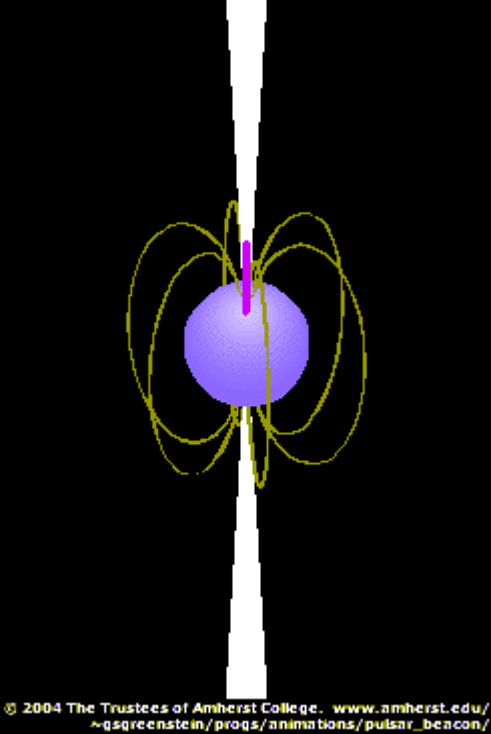


Gamma ray

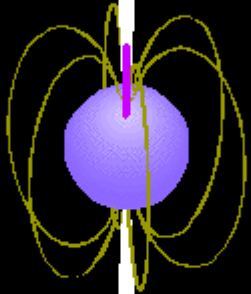


Pulsar

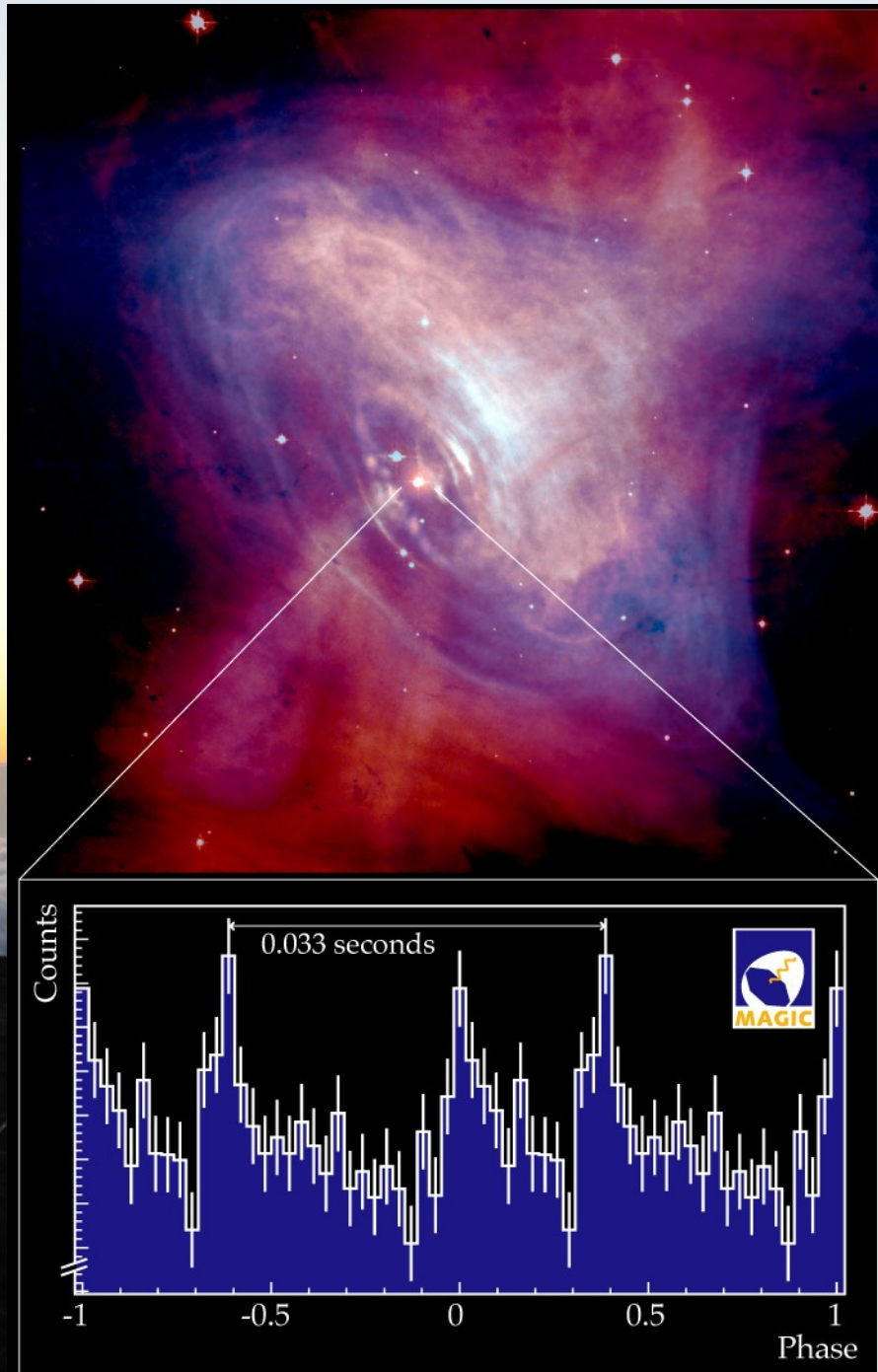
- All'interno delle SNRs in alcuni casi ci sono stelle di Neutroni che ruotano rapidamente emettendo un segnale pulsato molto rapido
- MAGIC ha scoperto un segnale gamma VHE (40 GeV) da una di queste sorgenti



Pulsar



© 2004 The Trustees of Amherst College. www.amherst.edu/~gsgrinstein/progs/animations/pulsar_beacon/



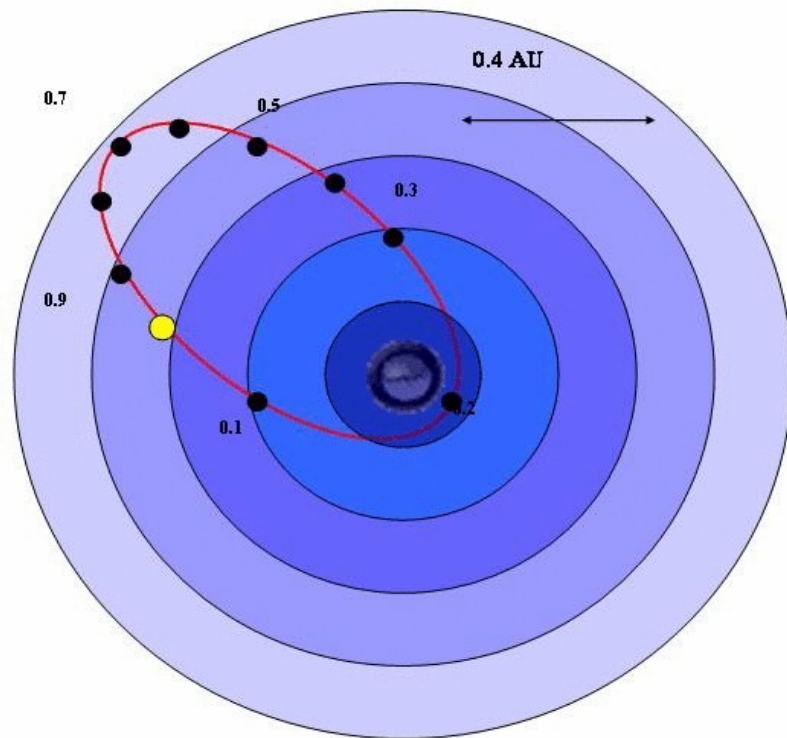
Credits: NASA/Chandra X-Ray Center/Hubble Space Telescope/Arizona State Univ./J. Hester et al.

Micoquasars

- sistemi binari formati da un buco nero (generato dalla morte di una stella massiva) e da una gigante rossa che lo alimenta

- I buchi neri sono circondati da un disco di accrescimento e da due jet in cui vengono accelerate particelle ad energie ultrarelativistiche

La Sorgente LSI 61+303



Fermiamoci un attimo e ricapitoliamo!





Al di Fuori dalla Via Lattea...

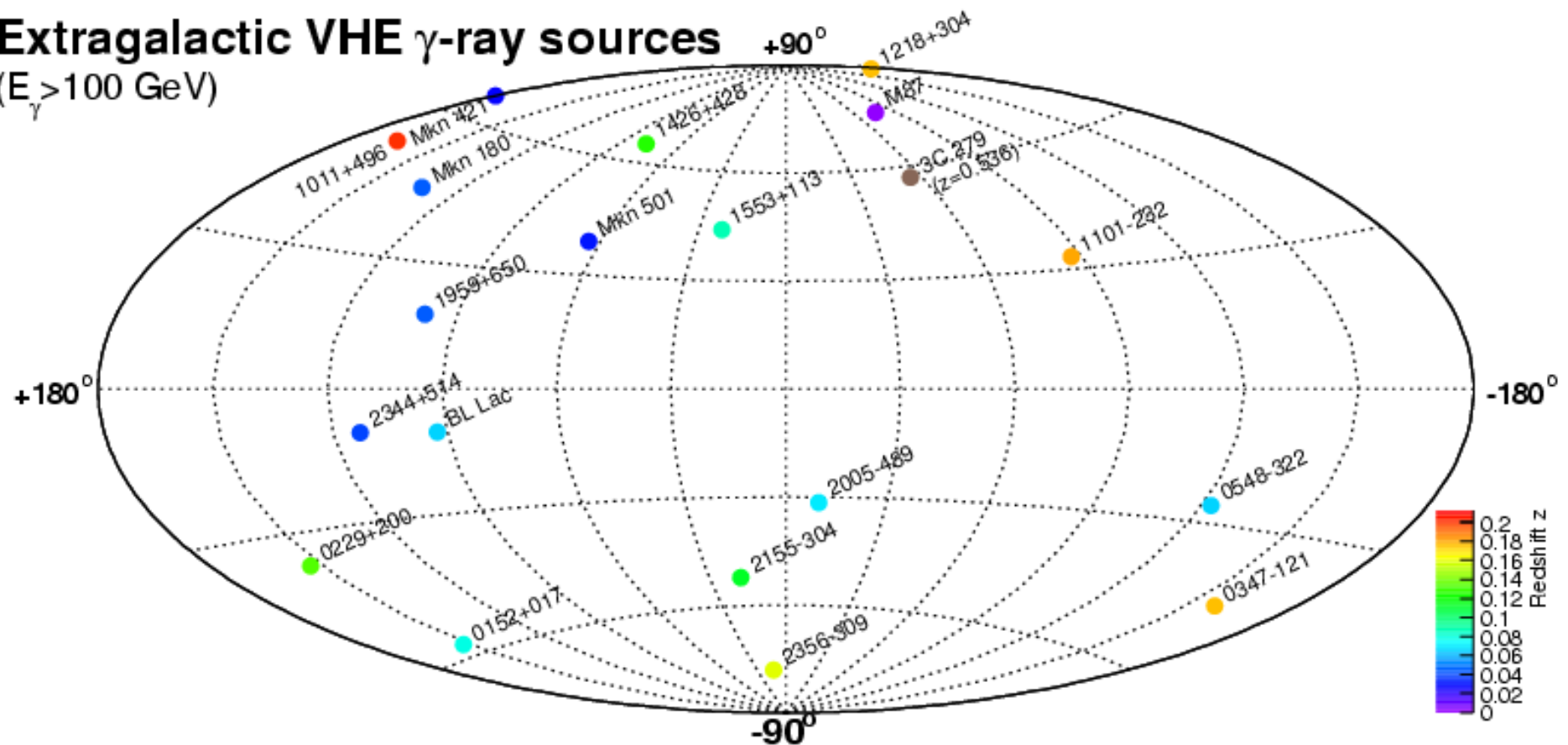
...guardiamo sorgenti
ancora piu' energetiche!

Fotografia in ottico del telescopio Hubble

Sorgenti Extragalattiche

Extragalactic VHE γ -ray sources

($E_{\gamma} > 100$ GeV)

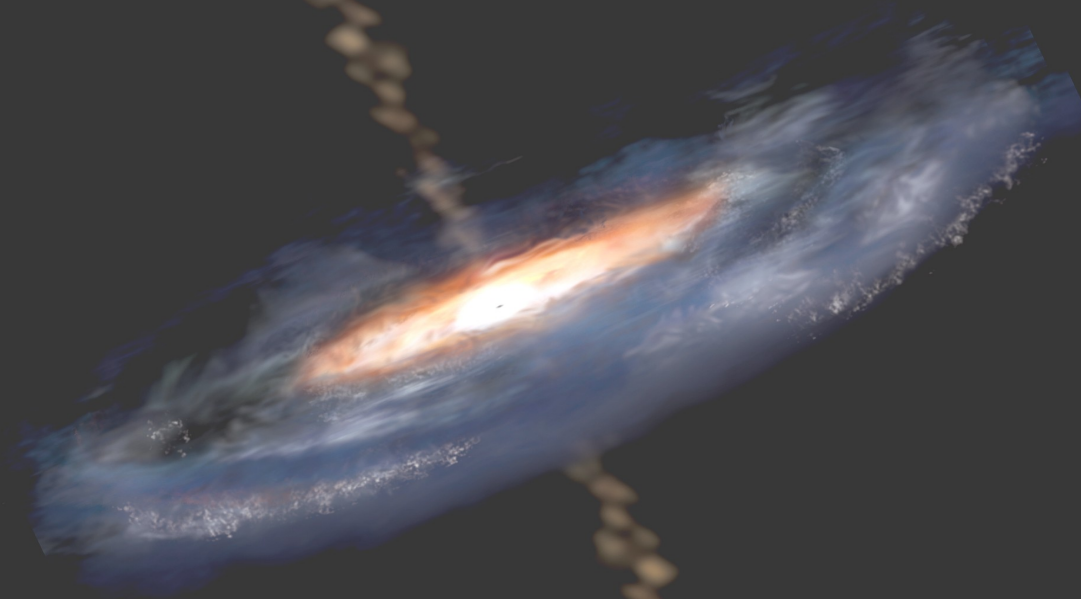


2007-11-22 - Up-to-date plot available at <http://www.mppmu.mpg.de/~rwagner/sources/>

TUTTE LE SORGENTI
EXTRAGALATTICHE SONO NUCLEI
GALATTICI ATTIVI (AGN)

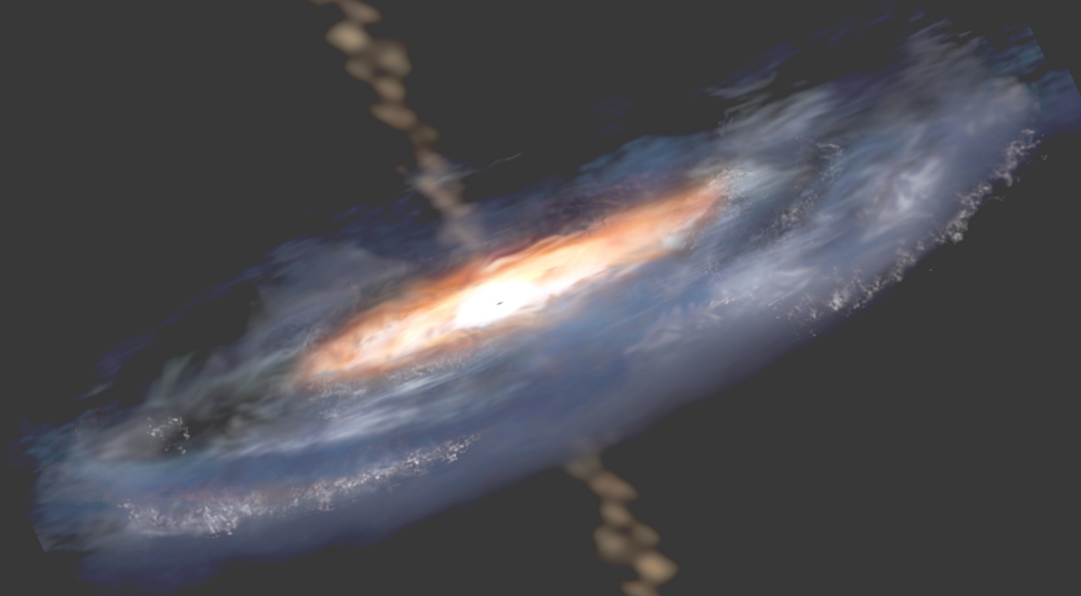
Redshift z $\rightarrow$
In astrofisica
è un
indicatore
della distanza

Nuclei Galattici Attivi



- BUCO NERO SUPERMASSIVO
- Disco di accrescimento
- Due jet di particelle ultra-relativistiche

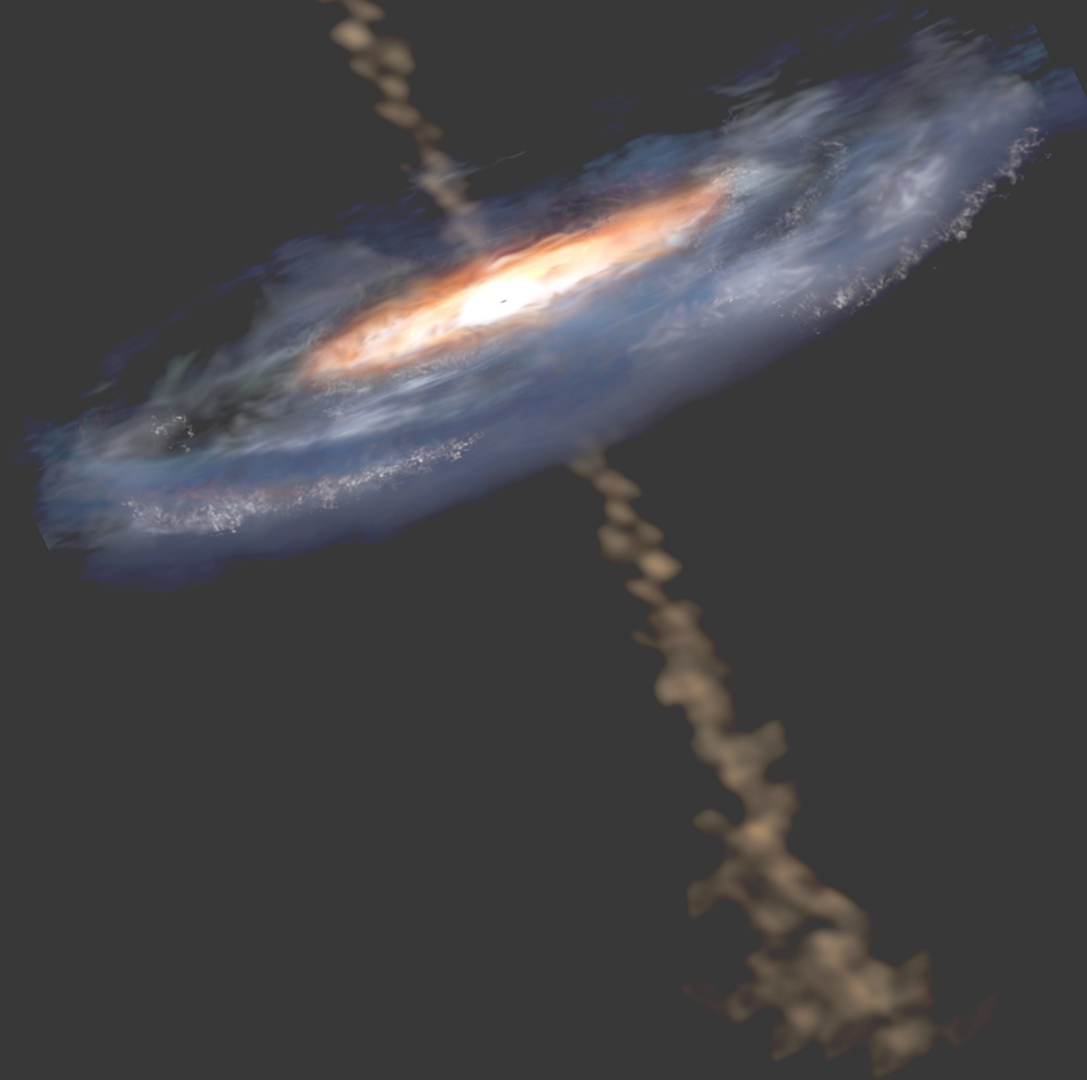
Nuclei Galattici Attivi



- Si pensa che questi oggetti estremi esistano al centro di ogni galassia...

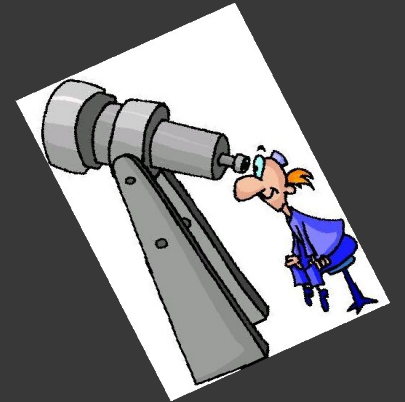
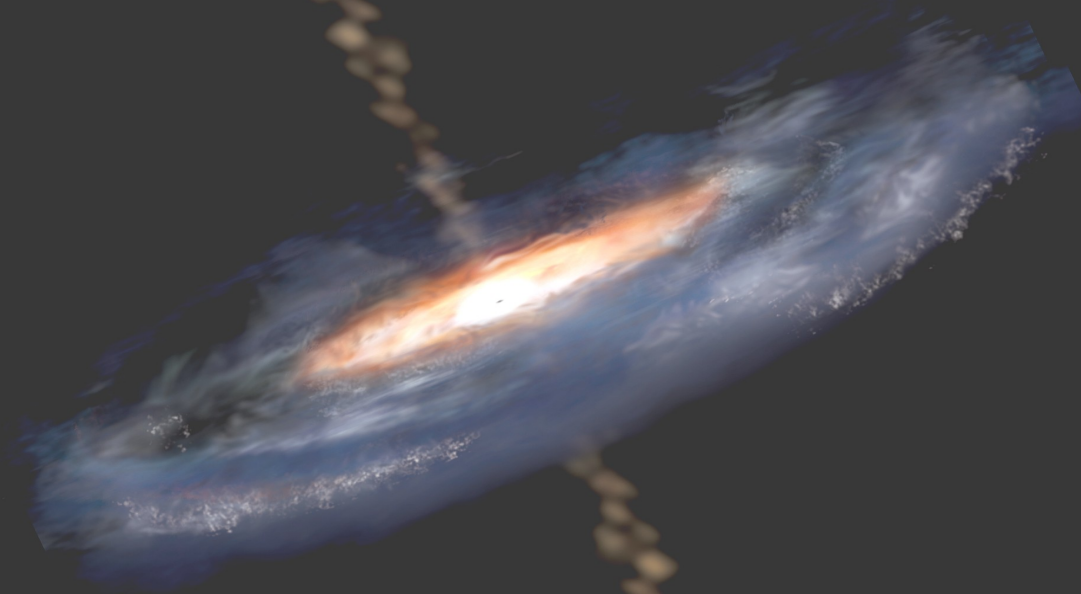
Nuclei Galattici Attivi

- Quello che osserviamo dipende dall' orientazione che il Nucleo Attivo ha rispetto a noi!



Nuclei Galattici Attivi

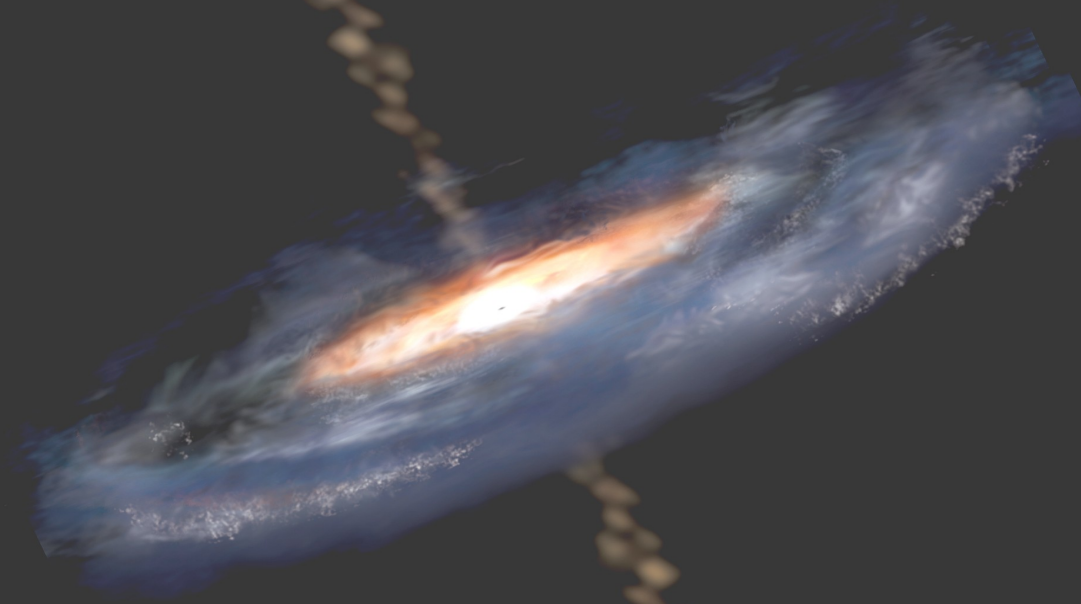
➤ Quello che osserviamo dipende dall' orientazione che il Nucleo Attivo ha rispetto a noi!



➤ **RADIOGALASSIE!**

Nuclei Galattici Attivi

➤ Quello che osserviamo dipende dall' orientazione che il Nucleo Attivo ha rispetto a noi!

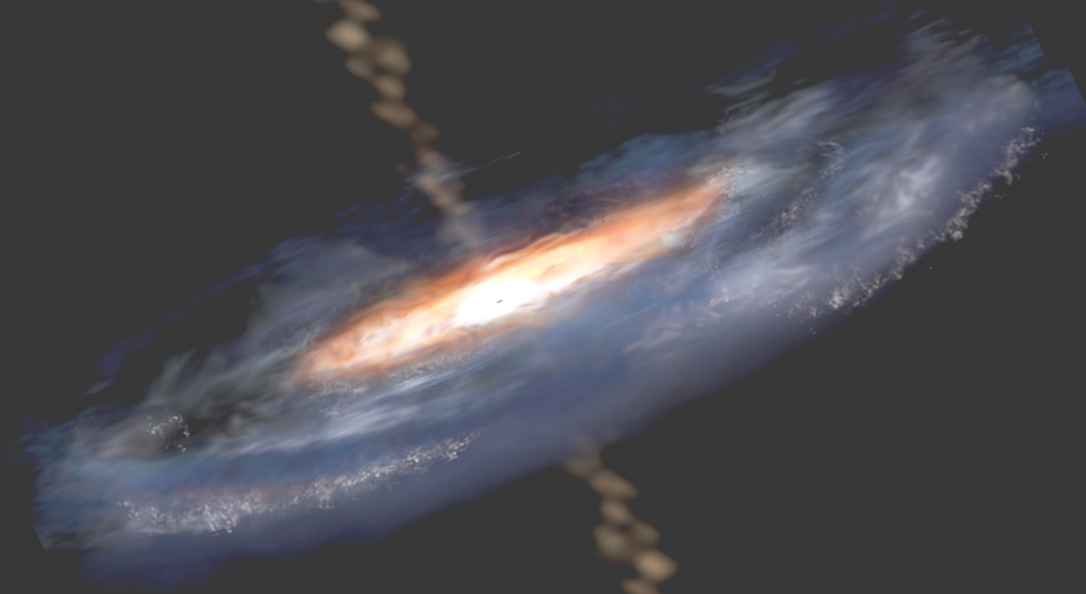


➤ QUASAR (QUASi stellar object)

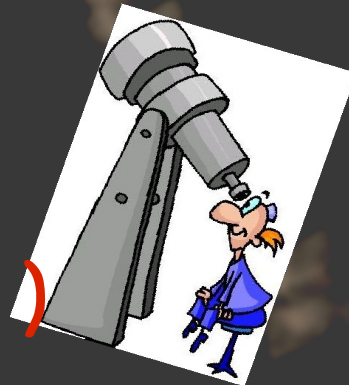


Nuclei Galattici Attivi

➤ Quello che osserviamo dipende dall' orientazione che il Nucleo Attivo ha rispetto a noi!



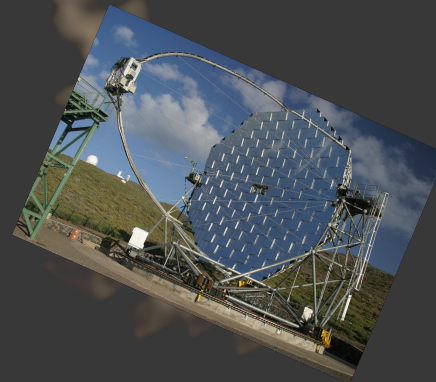
➤ BLAZAR! -->
particelle del jet
emettono raggi gamma
(anche ultraenergetici!)



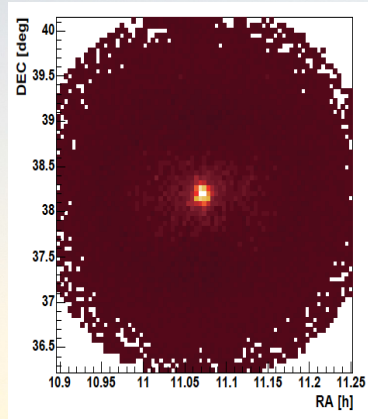
Nuclei Galattici Attivi

➤ Quello che osserviamo dipende dall' orientazione che il Nucleo Attivo ha rispetto a noi!

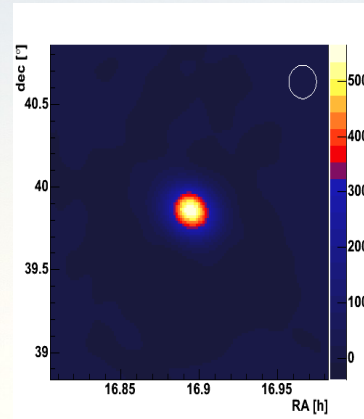
➤ BLAZAR! --> particelle del jet emettono raggi gamma (anche ultraenergetici!) -> MAGIC



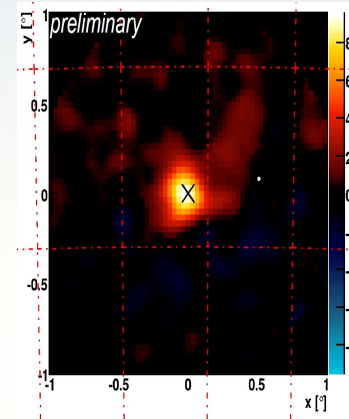
AGN osservati alle VHE



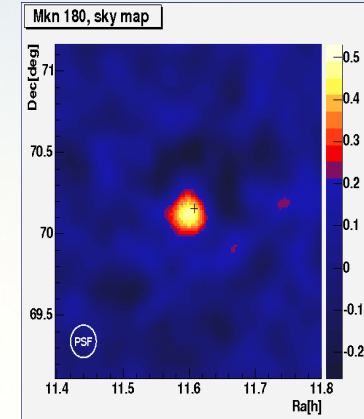
Mrk421 (0.031)



Mrk501 ($z=0.034$)

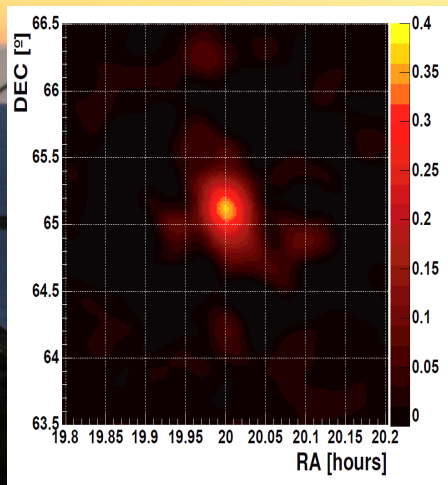


1ES2344 ($z=0.044$)

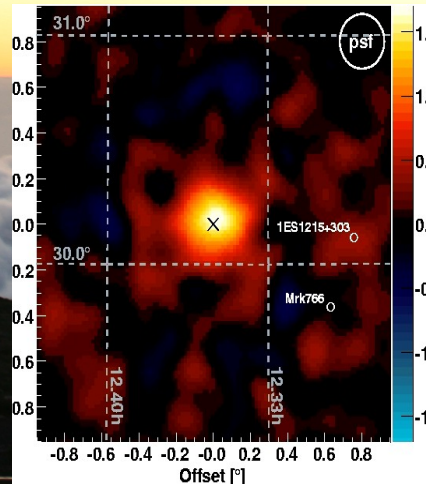


Mrk180 (0.045)

New source

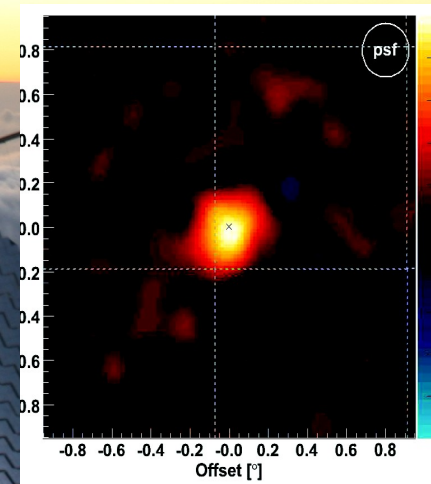


1ES1959 (0.047)



1ES1218 ($z=0.18$)

New Source



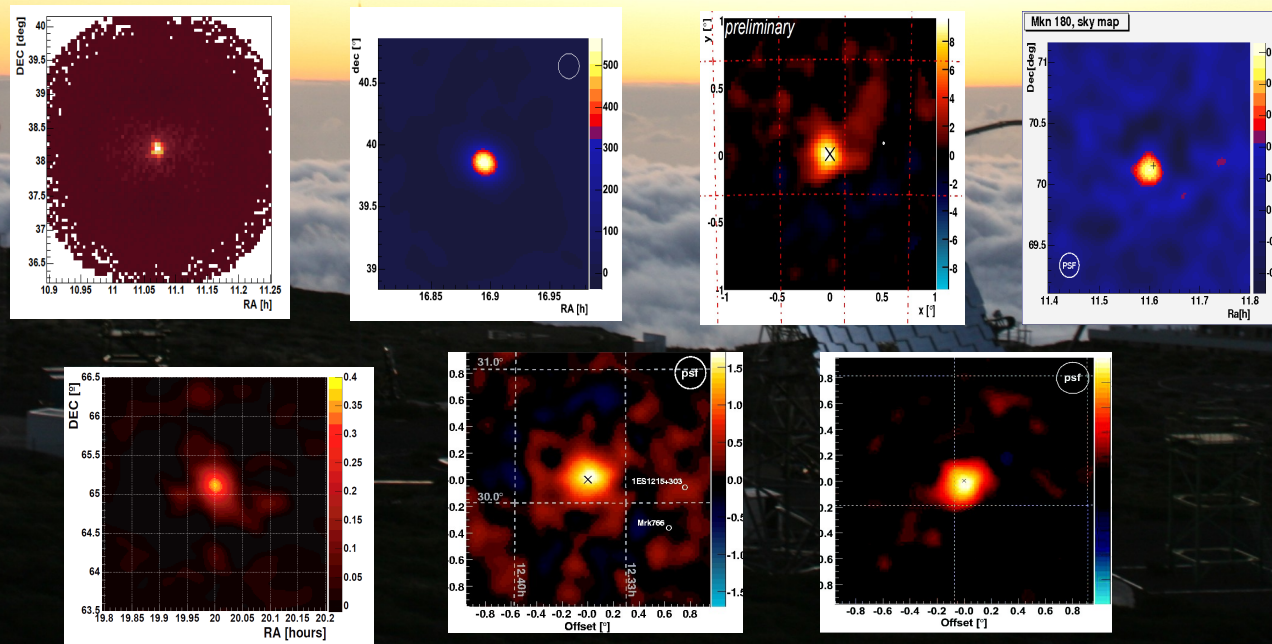
PG 1553 ($Z>0.25$)

New source

AGN osservati alle VHE

Grazie a queste osservazioni abbiamo iniziato a comprendere la natura dei jet negli AGN, anche se ancora esistono molti modelli diversi e quindi servono altre osservazioni per capire qual'è il modello giusto

MA... lo studio è appena iniziato... c'e' ancora moltissima strada da percorrere...



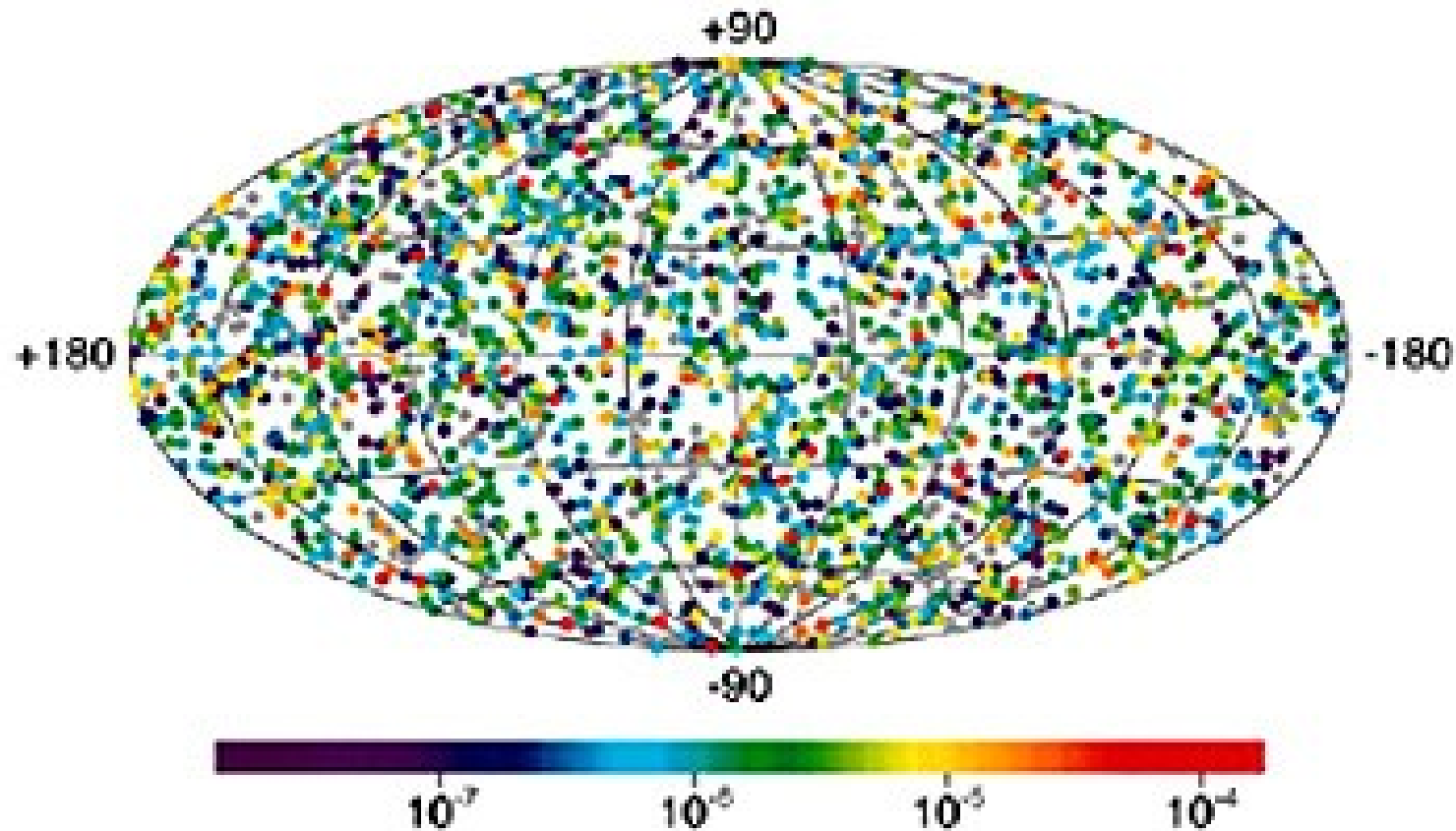
Ricerca di frontiera...

... ovvero dove si cercano
raggi gamma VHE ma finora
non si sono osservati



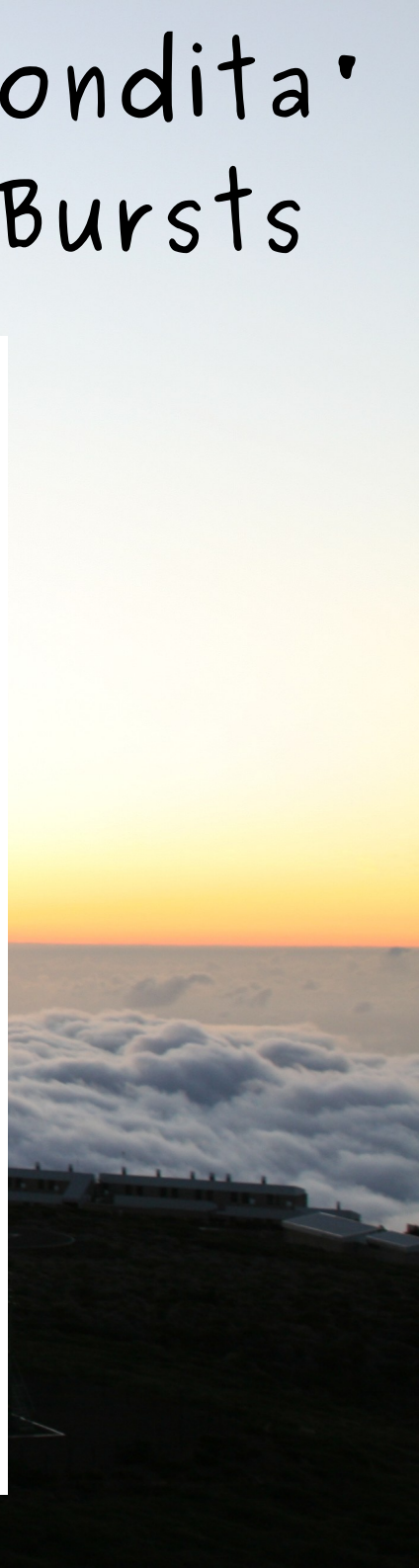
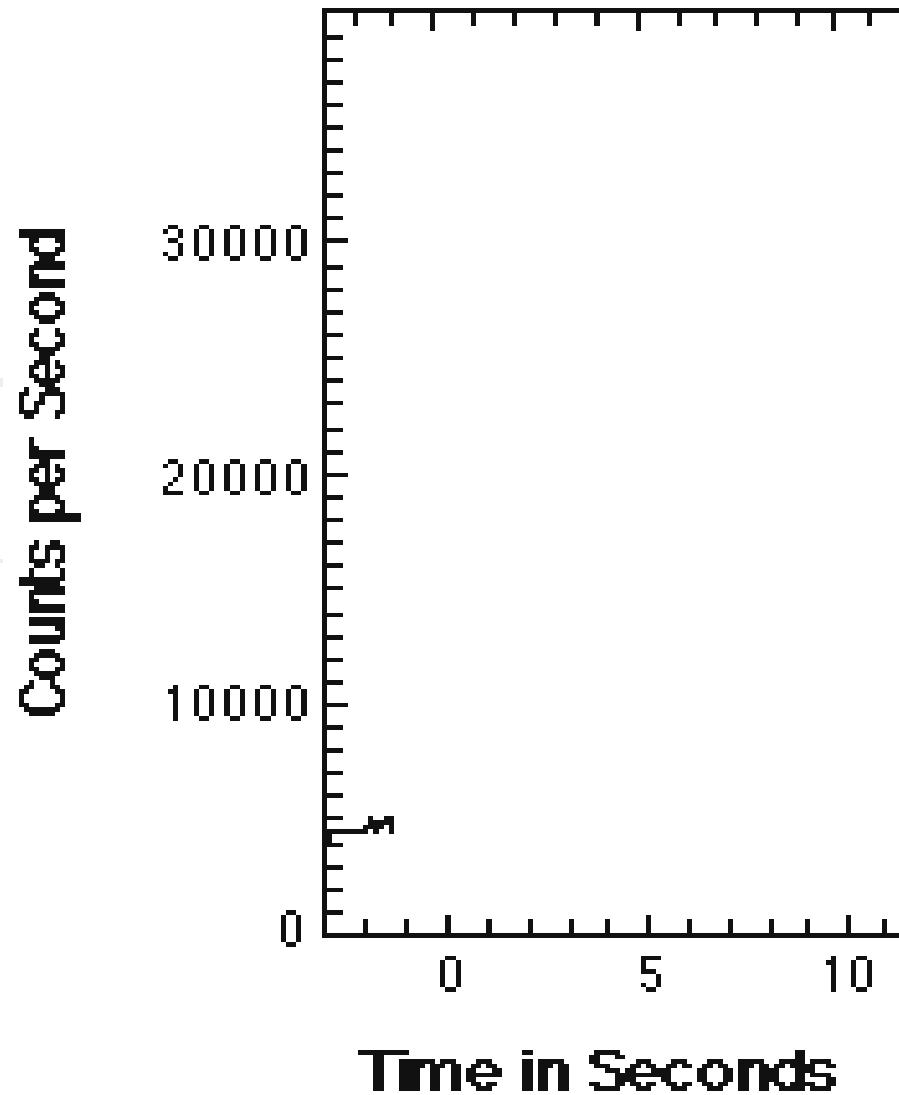
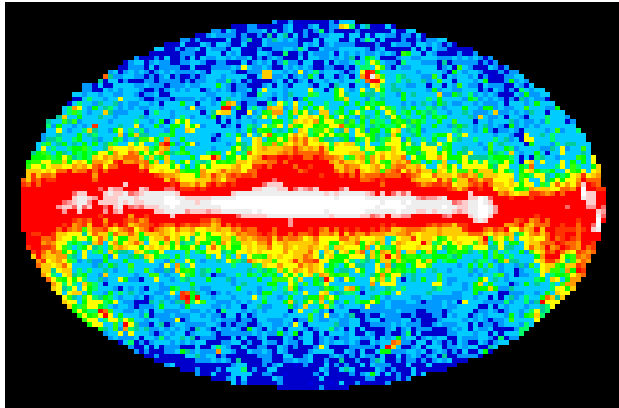
Lampi Gamma-soft dalle profondità
dell'Universo... Gamma Ray Bursts

2512 BATSE Gamma-Ray Bursts

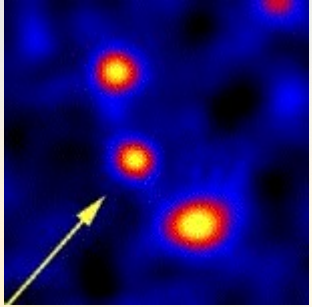


mapa dei GRB

Lampi Gamma-soft dalle profondità dell'Universo... Gamma Ray Bursts



Lampi Gamma-soft dalle profondità dell'Universo... Gamma Ray Bursts



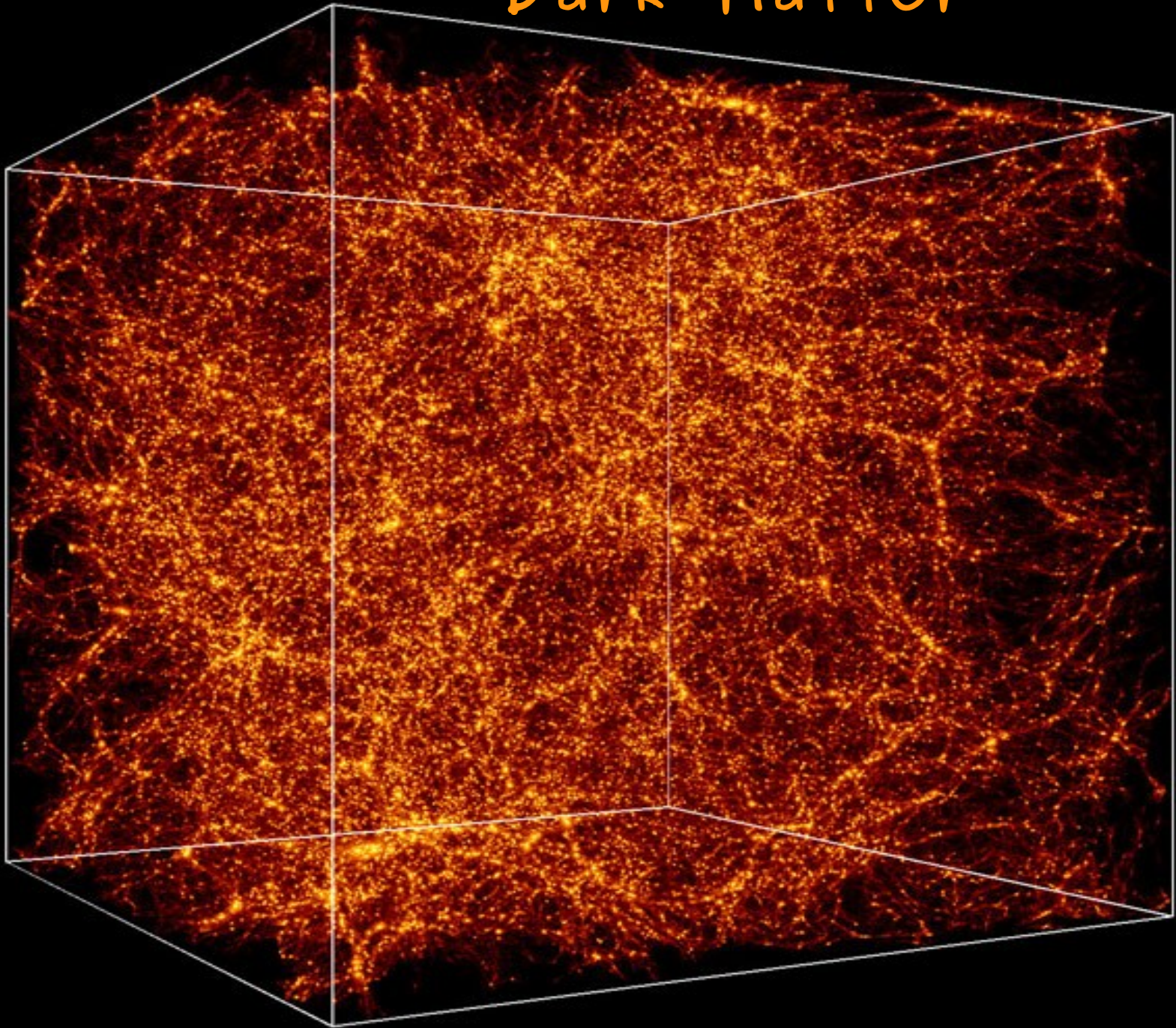
l'origine di questi fenomeni è tuttora sconosciuta!

domanda aperta: emettono VHE??

--> è quello che MAGIC vorrebbe scoprire!



Dark Matter



Dark Matter

SAPPIAMO CHE ESISTE

NON sappiamo quale sia però la particella...

?neutralino?

?assione?

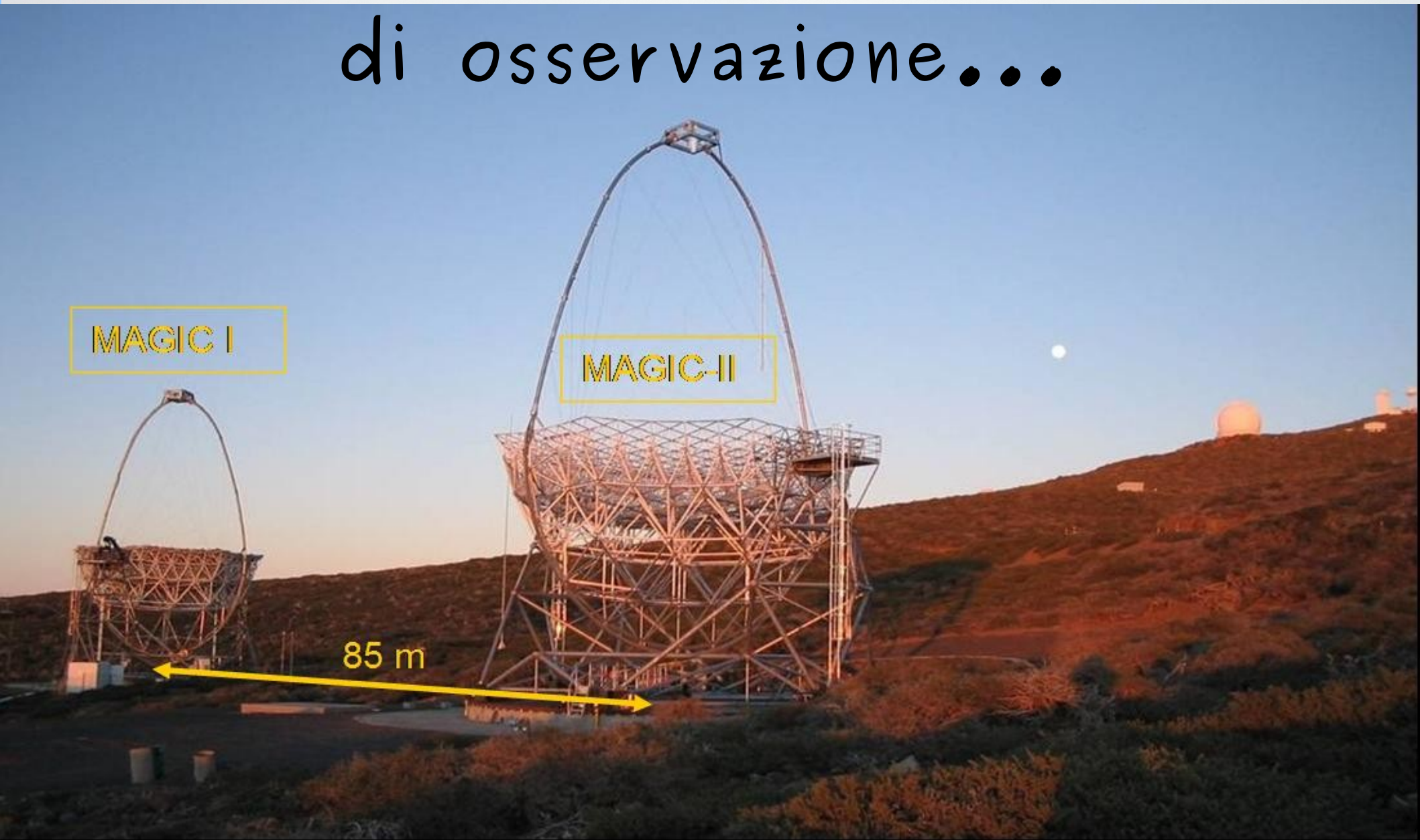
E' possibile, secondo alcune teorie, che la dark matter crei dei fotoni VHE...

... allora noi la cerchiamo!!!

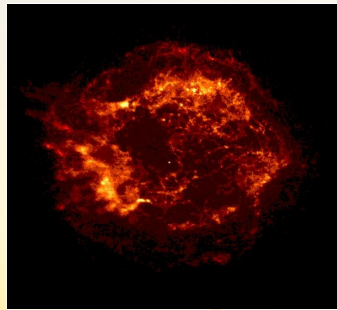
Ora non ci resta che
continuare la ricerca...



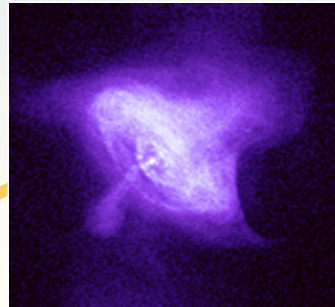
...cercando di migliorare
sempre di piu' gli strumenti
di osservazione...



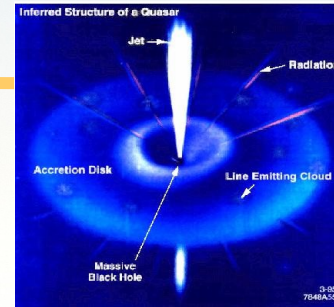
...e riuscire a "popolare" l'universo gamma!



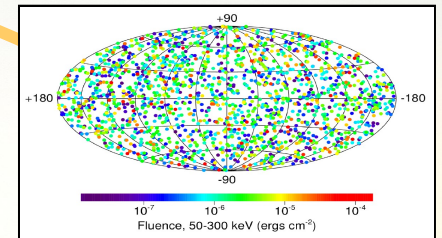
SNRs



Pulsars

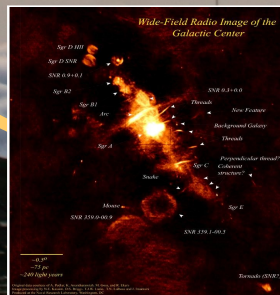


AGNs

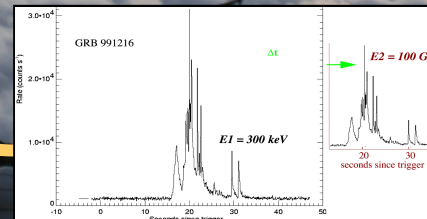


GRBs

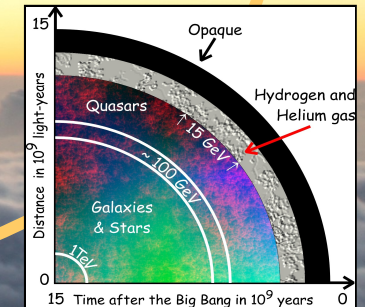
Origin of
Cosmic Rays



Cold Dark
Matter



Quantum Gravity
effects



cosmological
 γ -Ray Horizon

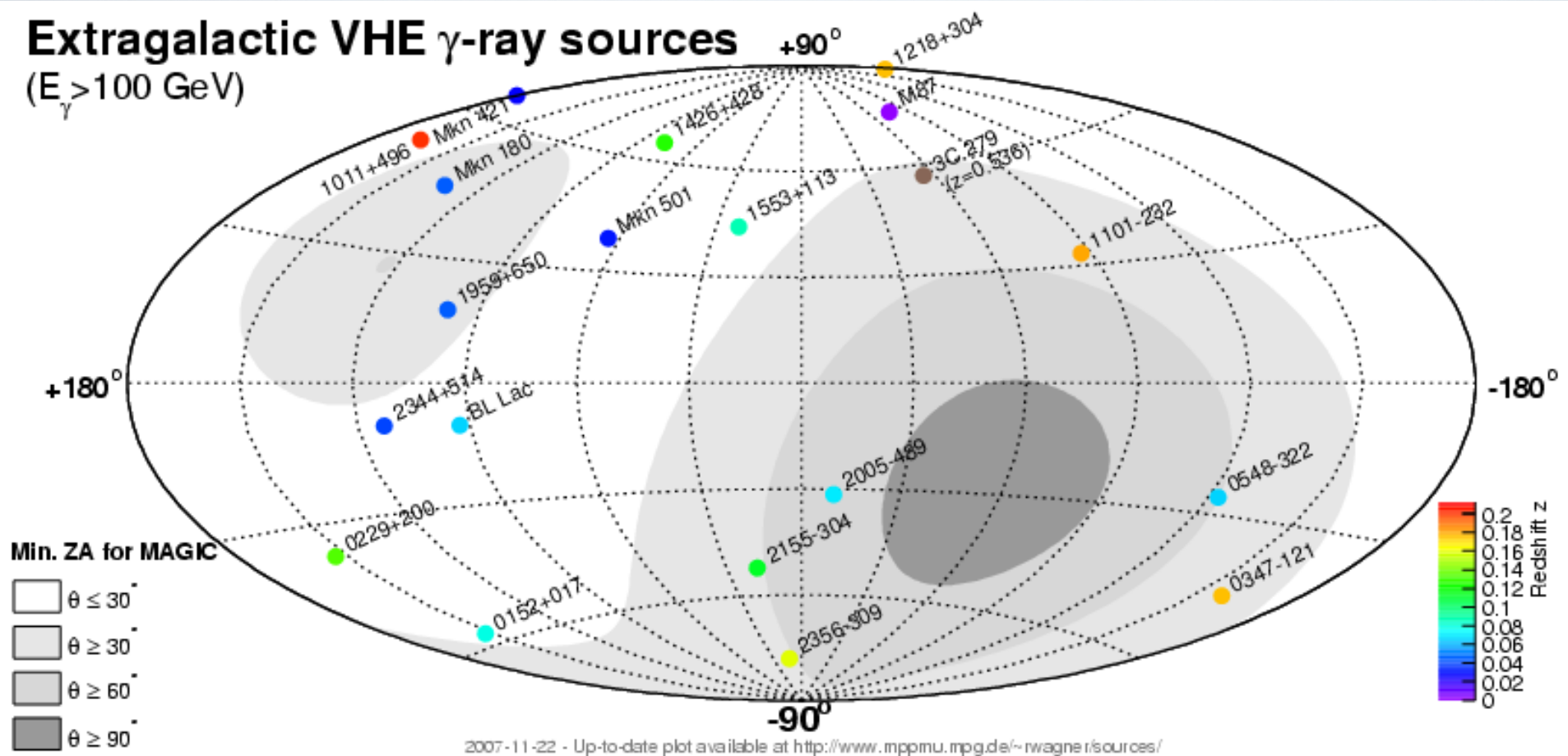
BACKUP SLIDES



Sorgenti Extragalattiche

Extragalactic VHE γ -ray sources

($E_\gamma > 100$ GeV)

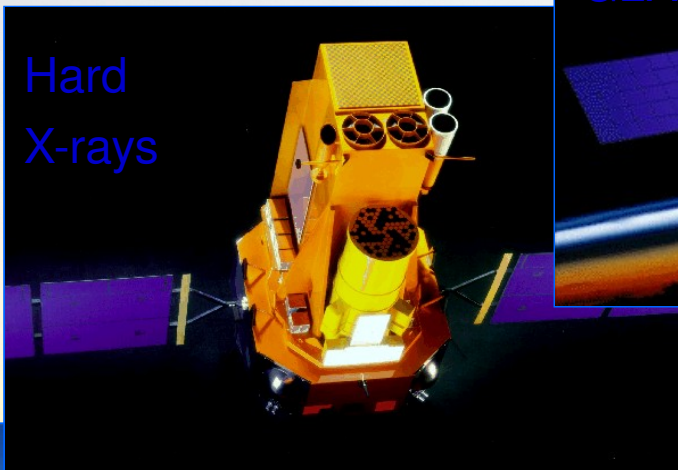


**TUTTE LE SORGENTI EXTRAGALATTICHE
SONO NUCLEI GALATTICI ATTIVI (AGN)**

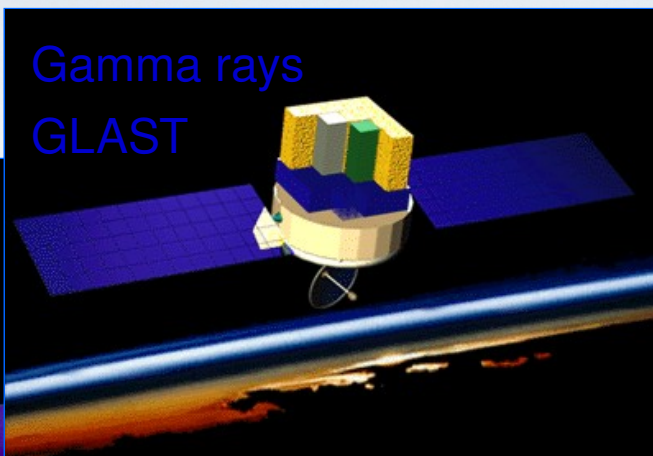
**Redshift --> In
astrofisica è un
indicatore della
distanza**

Multiwavelength / multimessenger astronomy

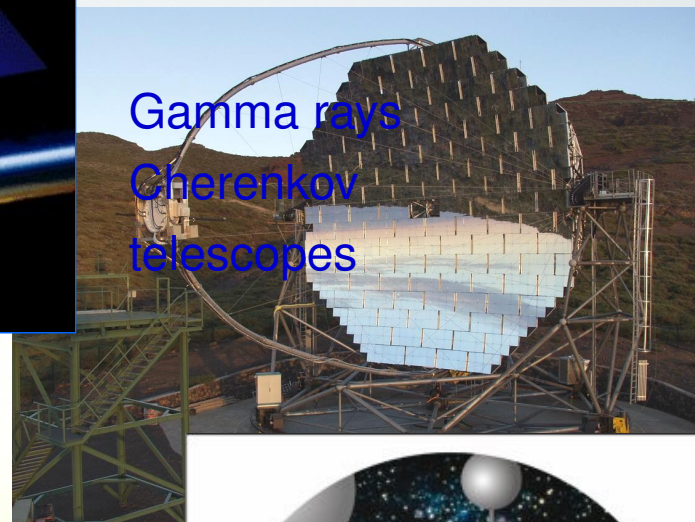
Hard
X-rays



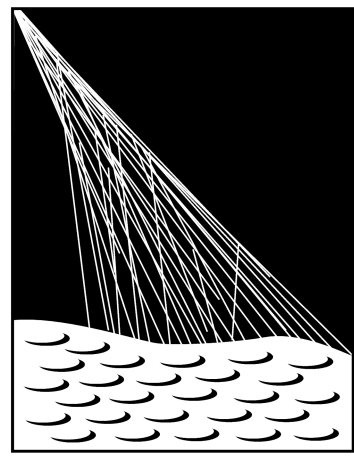
Gamma rays
GLAST



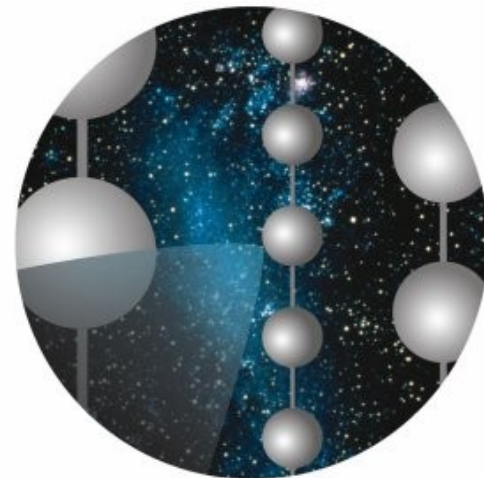
Gamma rays
Cherenkov
telescopes



X-rays



Radio
KAT,...



M3NeT

KM3NeT



Cosmic microwave background, ~3 mm

1 EeV 10^{18} eV 1 PeV 10^{15} eV 1 TeV 10^{12} eV 1 GeV 10^9 eV 1 MeV 10^6 eV 1 keV 10^3 eV 1 eV 10^{-3} eV 10^{-6} eV 10^{-9} eV 10^{-12} eV 10^{-15} eV

1 mm

UHE

VHE

HE
 γ -rays

X

UV
visible

IR

radio

10^{-18} m 10^{-15} m 10^{-12} m 10^{-9} m 10^{-6} m 10^{-3} m 1 m 10^3 m 10^6 m
1 am 1 fm 1 pm 1 nm 1 μ m 1 mm 1 m 1 km 1 Mm

300 GHz

300 MHz

300 kHz

satellites



balloons

rockets

100 km

40 km

20 km

γ

EAS

Cherenkov
telescopes

fluorescence
detectors

particle detectors

optical
Telescopes

Earth surface

N_2, O_2, O_3, O

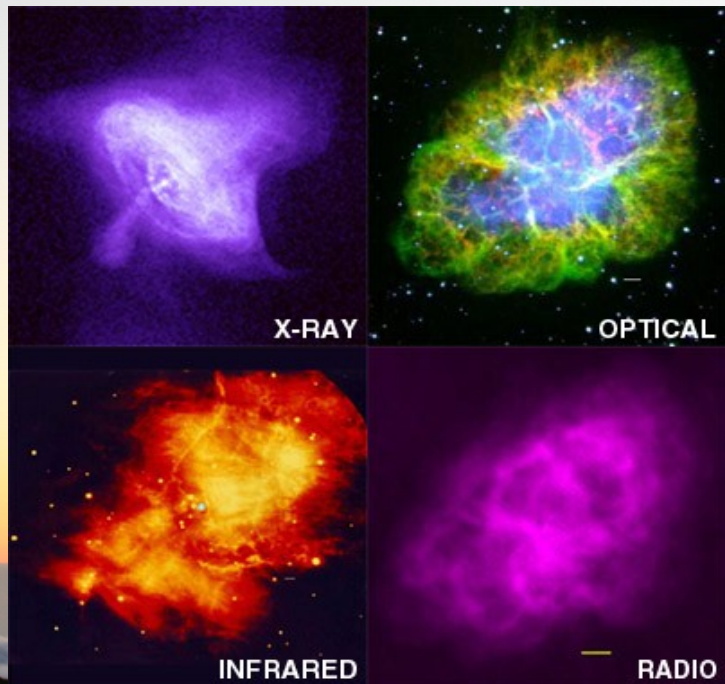
H_2O

CO_2

radio telescopes

50% of incident
radiation absorbed

La risoluzione angolare...



Gamma ray: MAGIC

