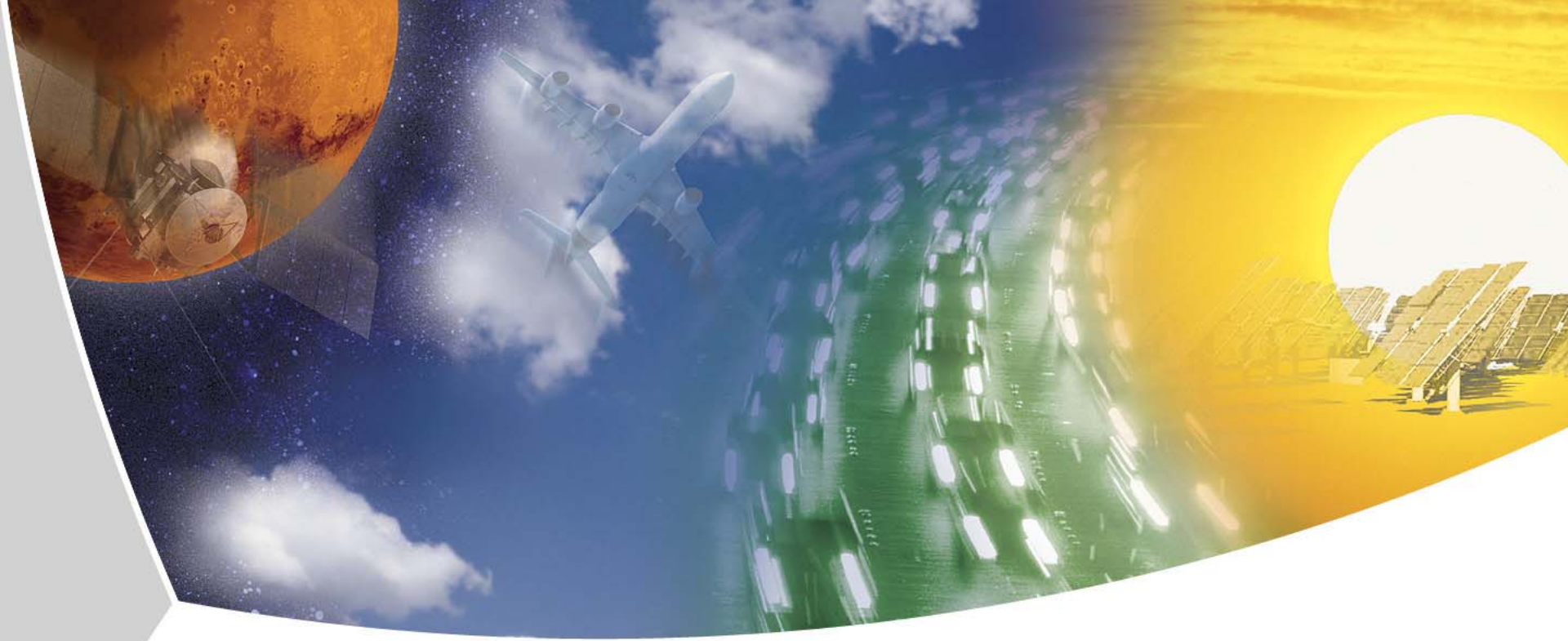




Astroparticle Physics, Cosmic Ray Physics and Space Situational Awareness

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Inputs by Medipix Consortium

S. Pospisil CTU Prague & K. Kudela IEPSAS Kosice



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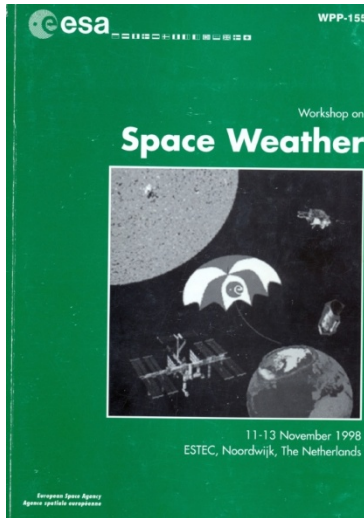


Content

- 1) Space Situational Awareness (SSA)
- 2) High Energy Galactic & Solar Cosmic Ray Data / Models
- 3) Forecast of SW storms and
Microphysics Detector Technology Applications in SSA



1. Space Situational Awareness



1998 ... 2008:
ESA / ESTEC



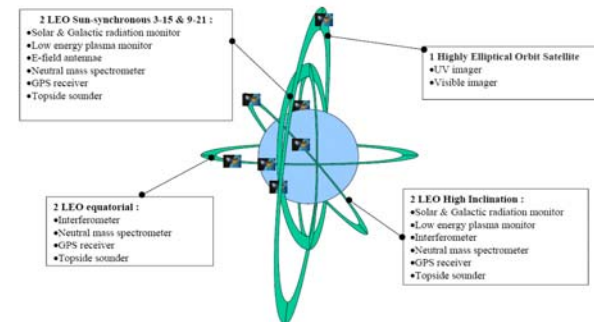
2000: <http://www.swissre.com/>



ESA Space Weather Studies



FULL Scale Scenario : Ionosphere/Thermosphere Monitors



Alcatel Space Industries

ESA Study for SPACE WEATHER PROGRAMME
© Copyright Alcatel Space 2001

December 17th 2001 - OSM/IE - Page : 12 / 17

2000 / 2001 : SW Feasibility Studies



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1. Space Situational Awareness



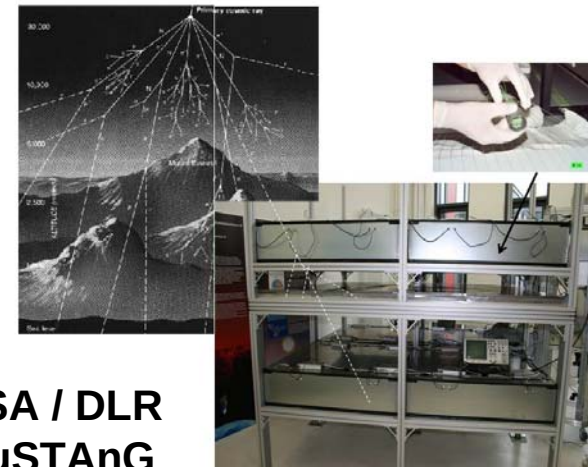
2003: ESP
White Paper



2007 / 2009: EU
/ FP6 SWEETS
DVD



2002 / 2003: EU / FP 5 SWE
(Space Weather and Europe)



2004 - 2006: ESA / DLR
MuSTAnG



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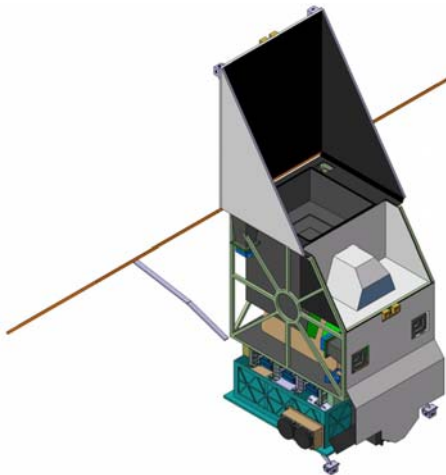


1. Space Situational Awareness

2008 - 2018: EU – ESA SSA



1) NEOs

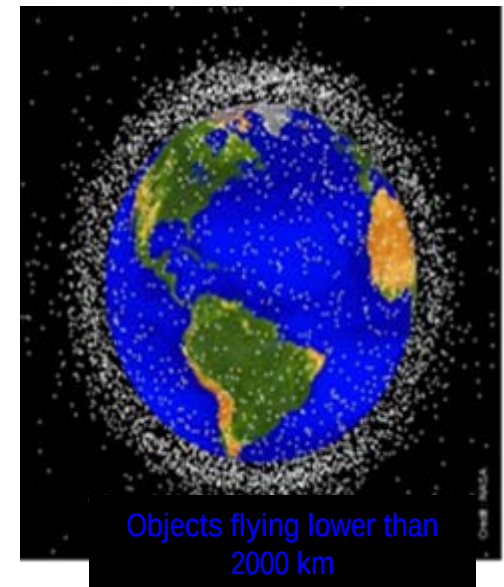


DLR Bremen
AsteroidFinder
Phase A & B

2) Space Weather



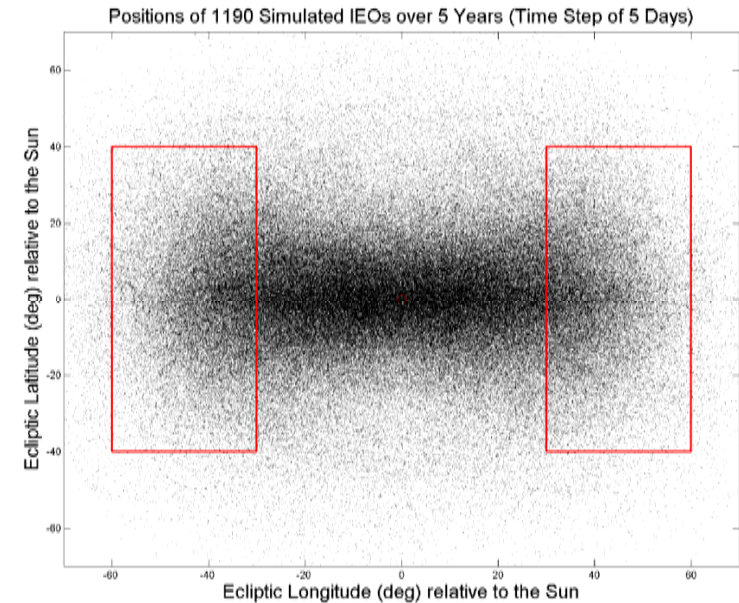
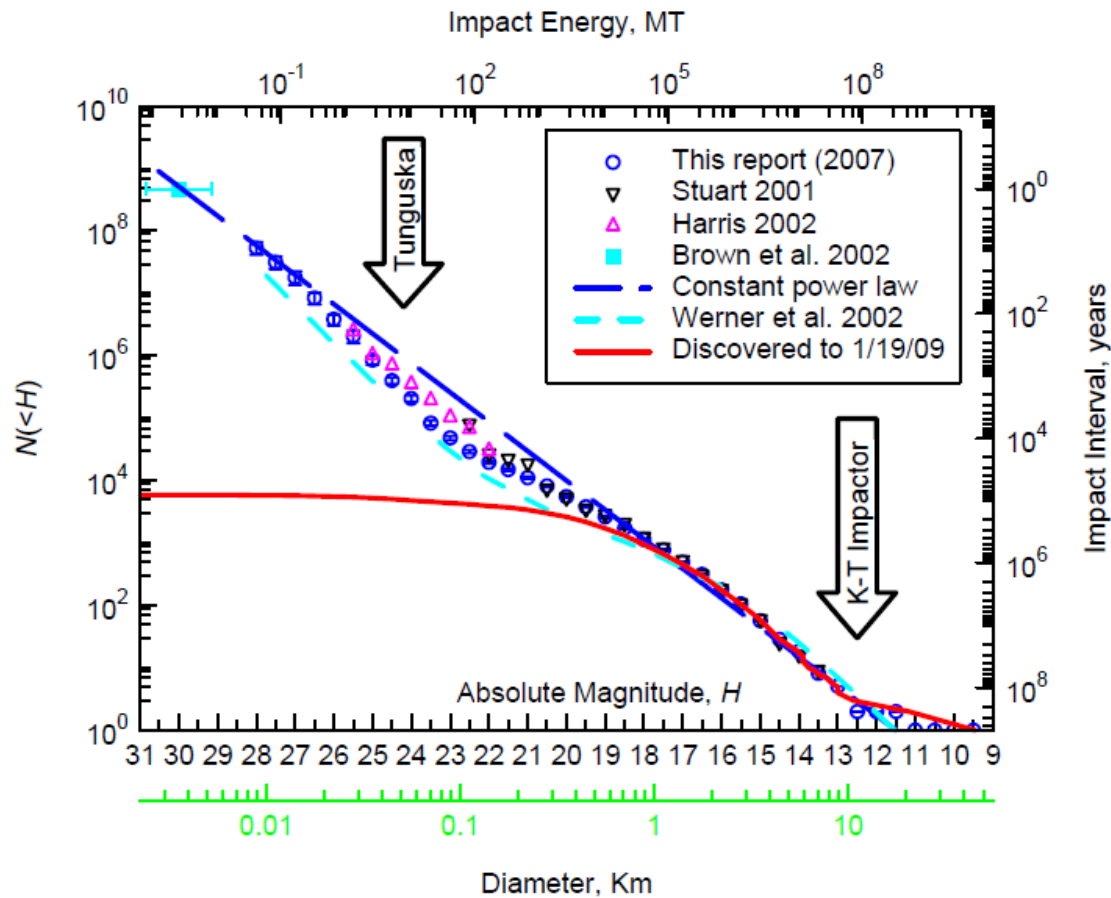
3) Space Debris



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1. Space Situational Awareness: NEO / IEO

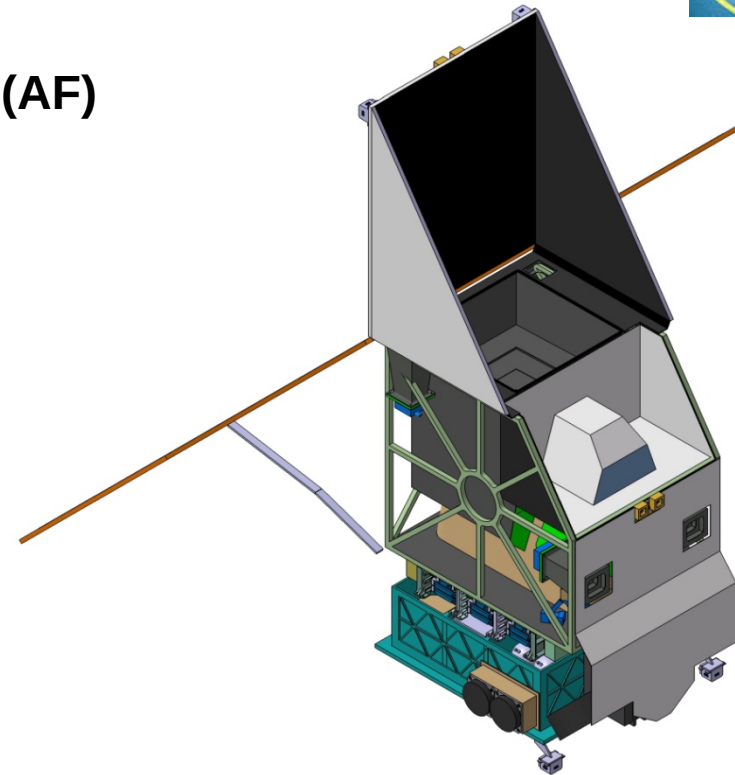
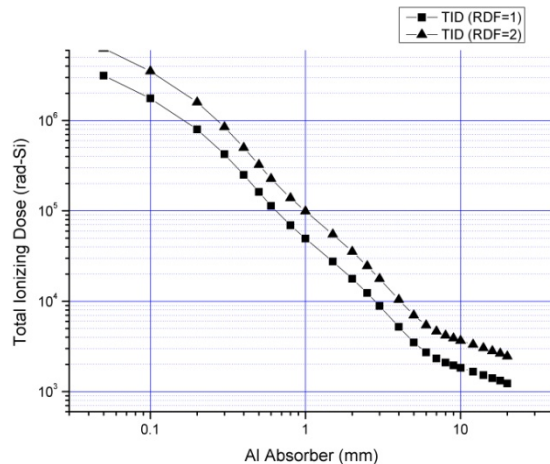


1. SSA and NEO: DLR AF



Compact satellite AsteroidFinder (AF)

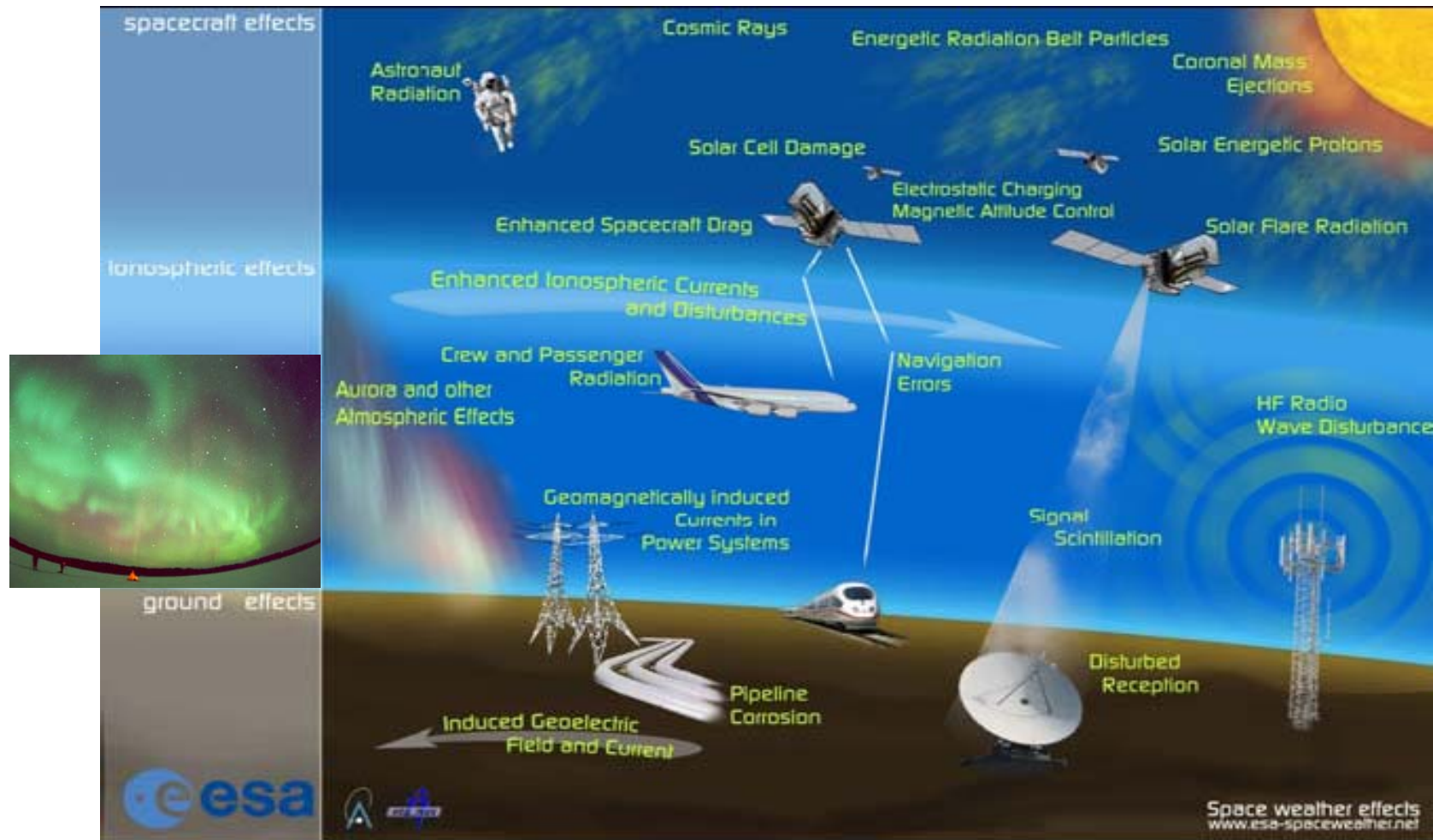
- satellite bus based on BIRD (TET)
- phase A finished
- phase B in 2009
- launch in 2012
- to be operated by GSOC



- CREME 96 + SPENVIS
- TID (trapped + solar particles)
- 3 years at 850 km orbit



1. Space Situational Awareness



SW effects

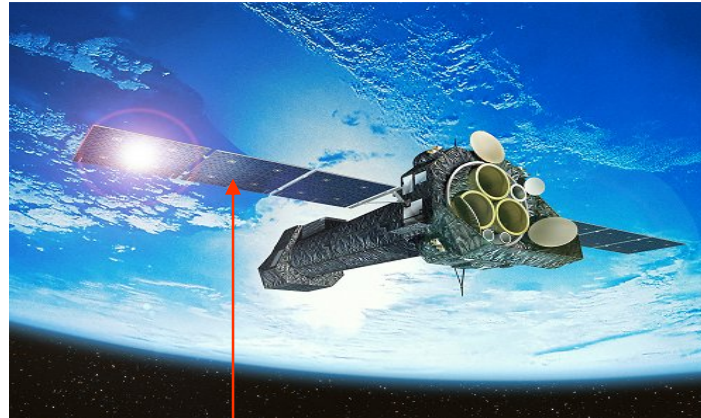


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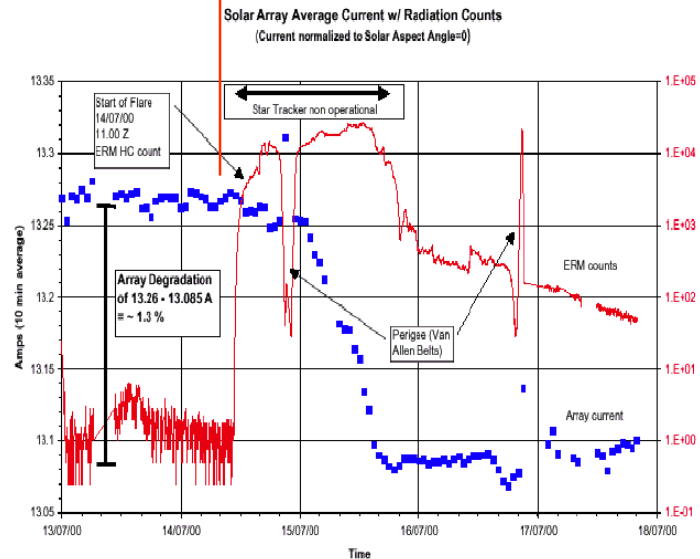


11th ICATPP Como 2009

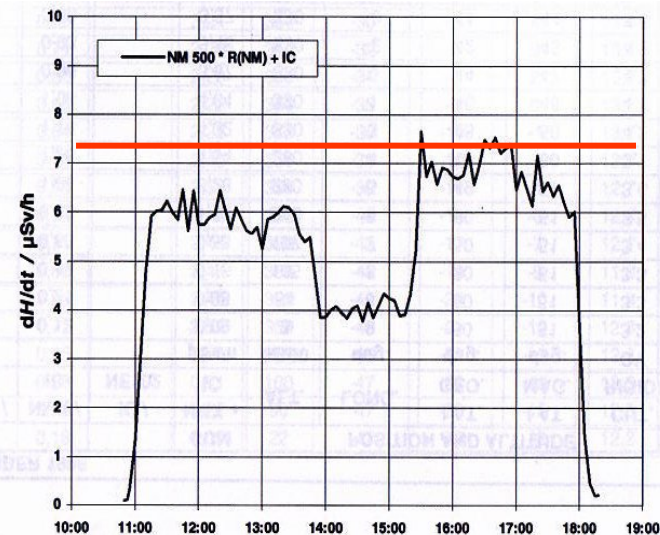
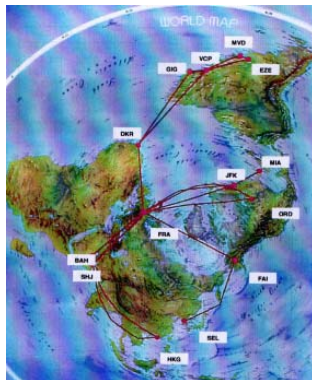
1. Space Situational Awareness: Space Effects



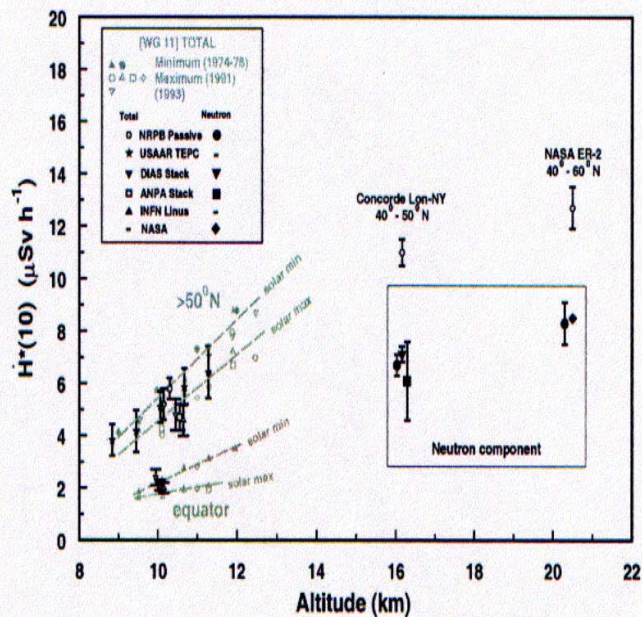
- 31 failures and **losses** (12 total losses)
Meteosat, ERS-1, XMM ...
Equator-S, Anik 1&2, Telstar 401, ASCA ...
- losses in 4 years more than **500 Millionen USD**



1. Space Situational Awareness: Aircraft Effects

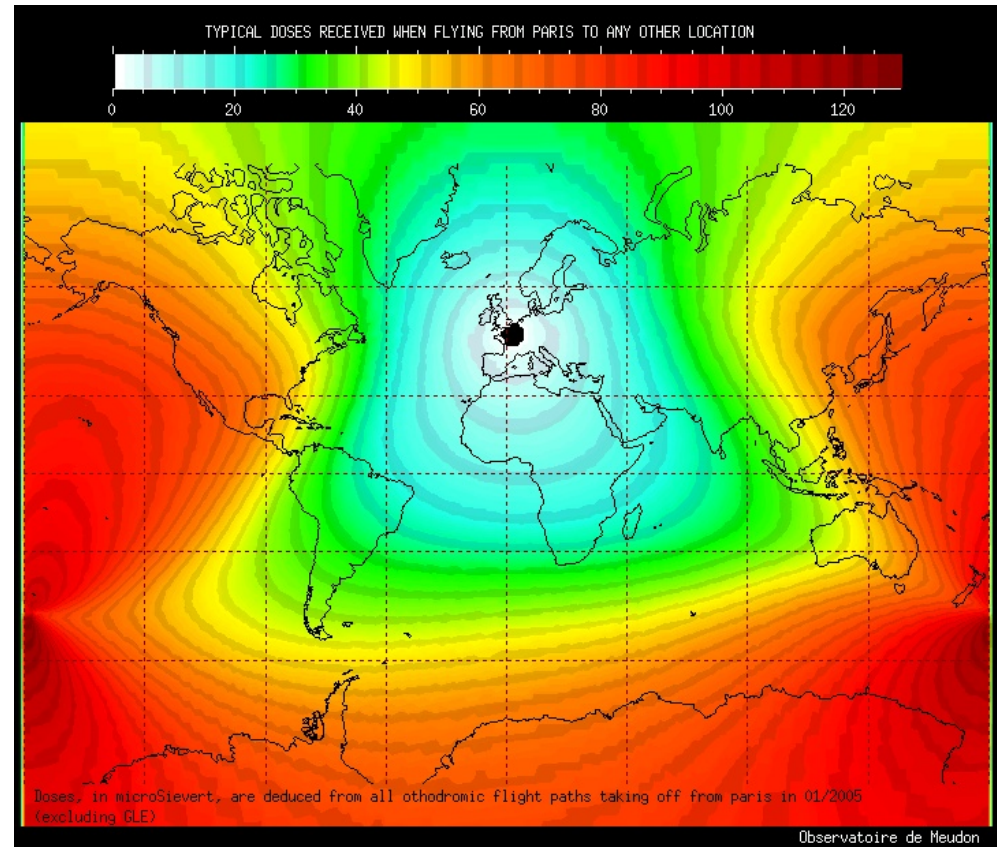


flight Chicago – Frankfurt 4 Oct 98 (PTB)

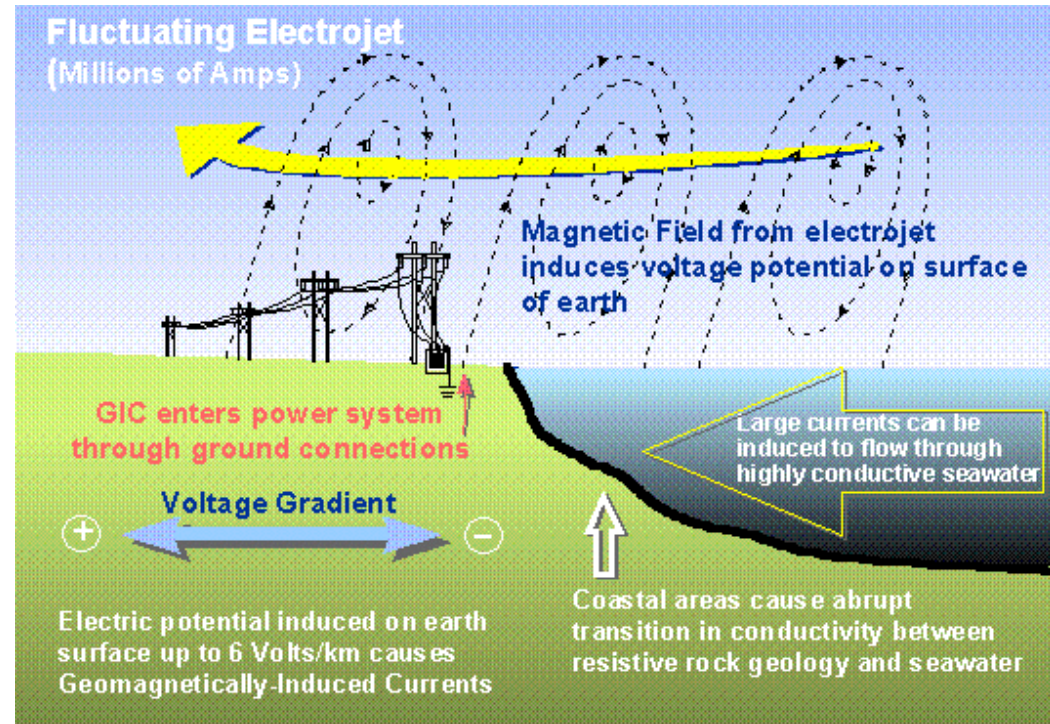
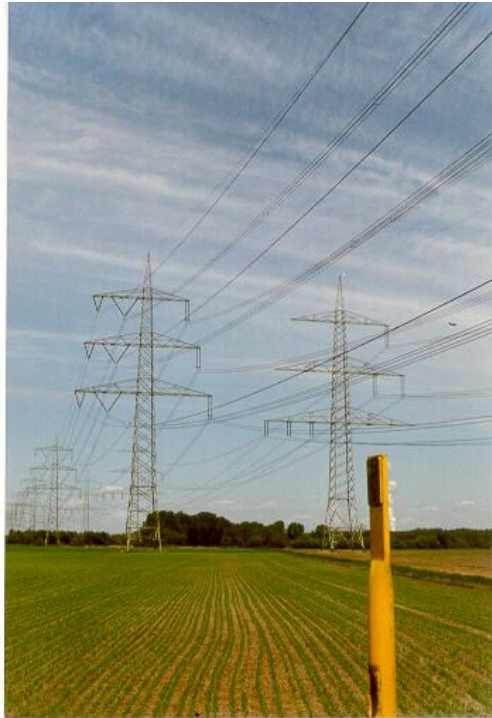


- European regulation
- in Germany new radiation law (BfS, BMU) since 1 June 2001:
airline have to estimate the radiation exposure (effective dose 1 mSv, upper limit 20 mSv per year)

1. Space Situational Awareness: Your CR Dose



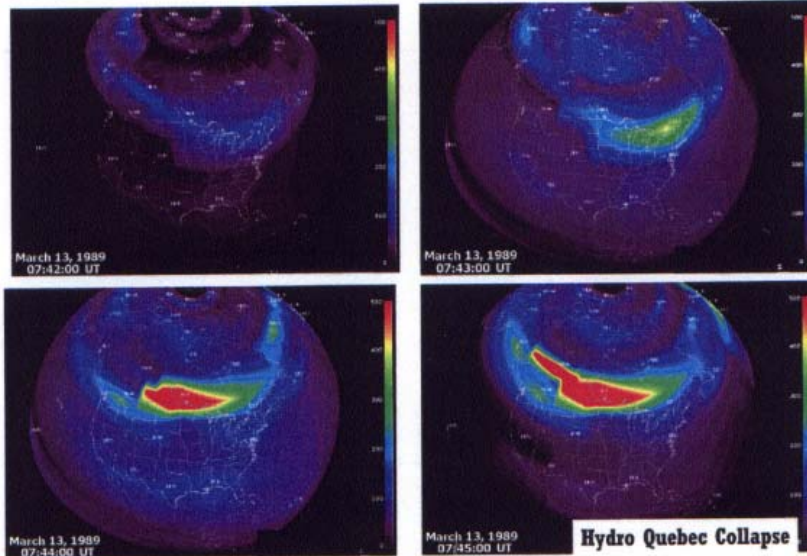
1. Space Situational Awareness: GICs



1. Space Situational Awareness: Power Cut

Power supply: 13 March 1989 – just before the blackout

Four Minutes of a Super Storm - March 13, 1989



- Hydro Quebec: collapse time 90 sec ! (400nT/min) (9 h power blackout for milion people in winterr)
- 865 nT/min measured in Minnesota!

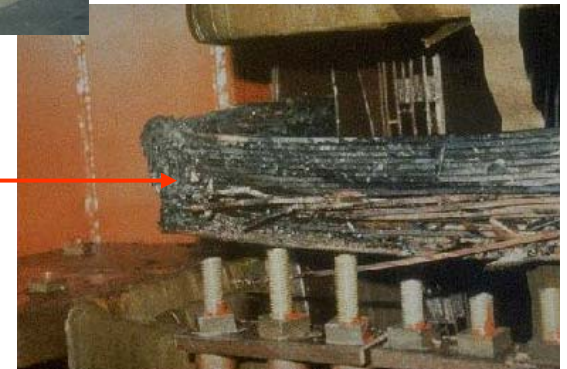
10 Milliarden USD



New Jersey



- normally 3000 A
- several amps => unsufficient working regime of the transformer
- 200 A measured in trafo.

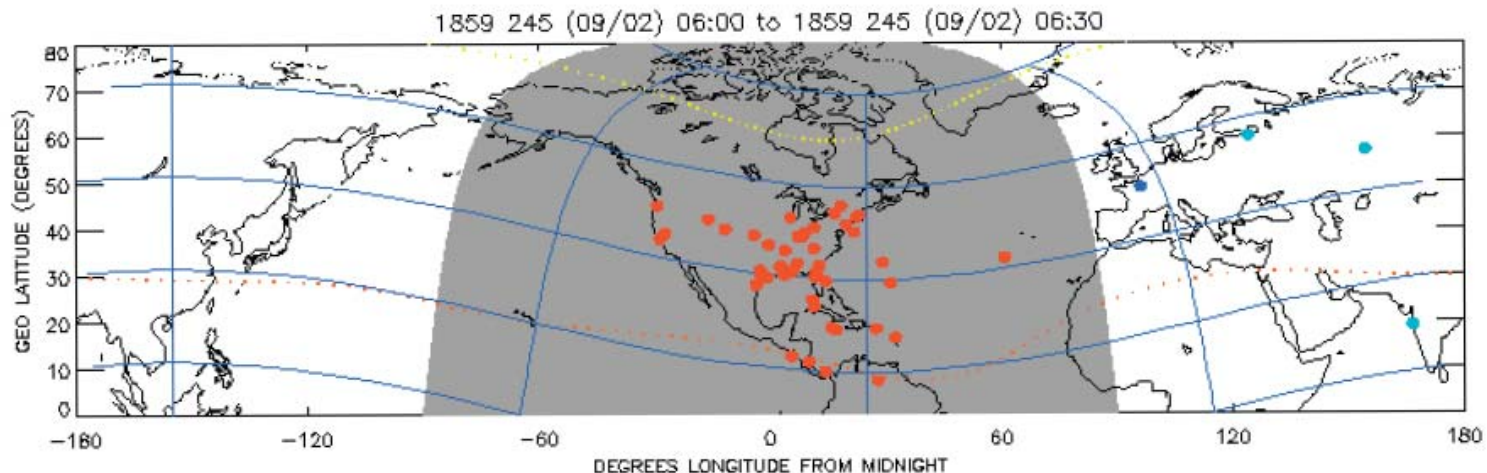


- Rep.: 10 Mill. USD (up to one year, here only 6 weeks)
- Sweden: 6 130- kV-I power line
- Cu-coils up to 3000 A
- April 94: 5 transformoren in Chicago
- Finnland: 143 amps measured in power lines



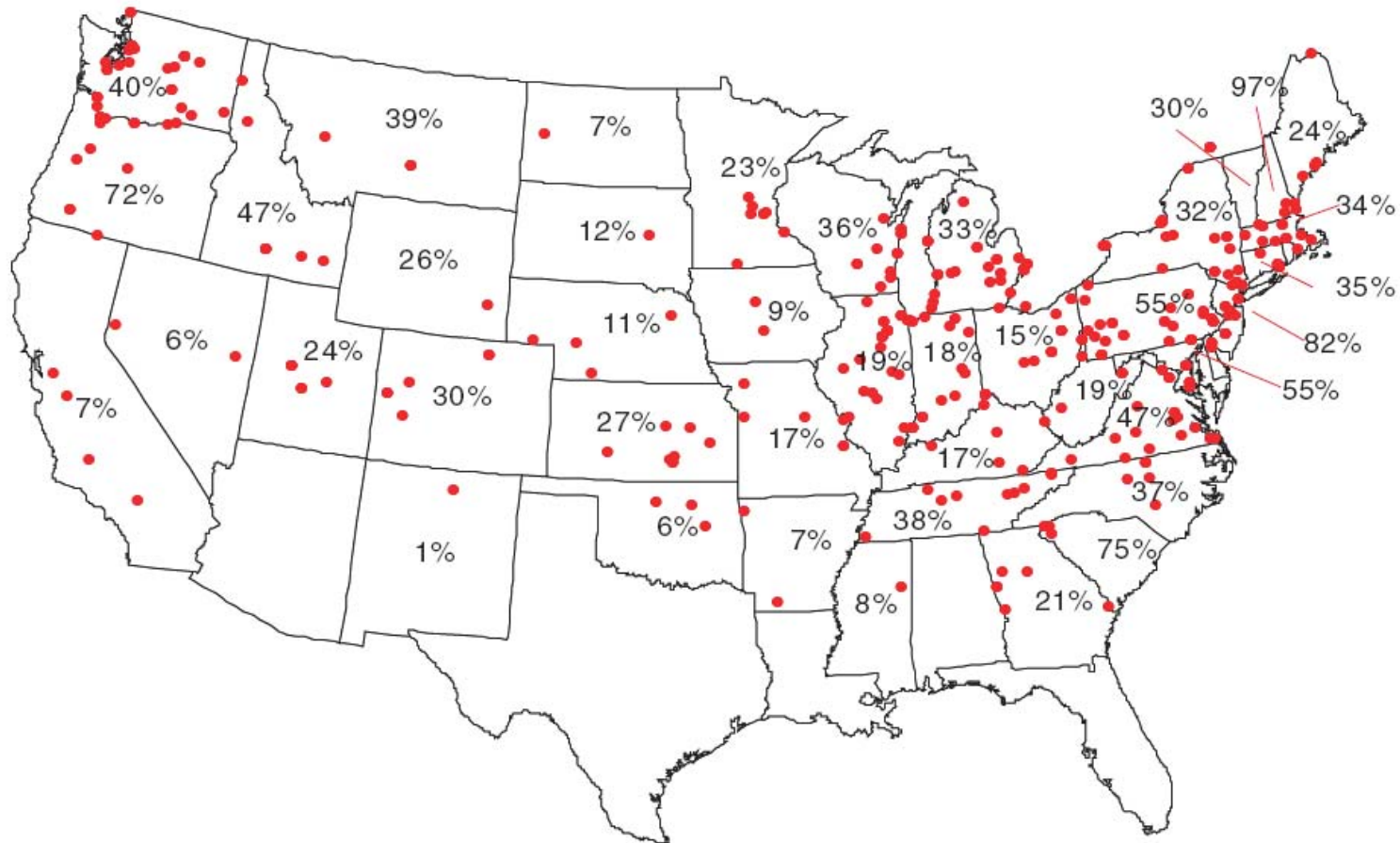
1. Space Situational Awareness: Space Katrina

2 September 1859 – and today?

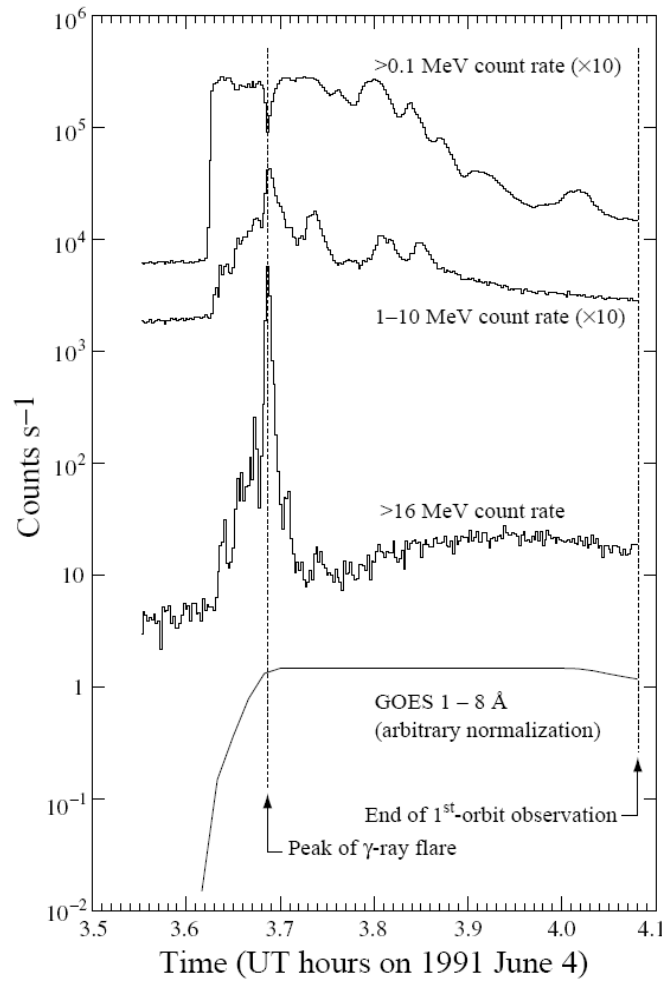


- Katrina: 81 – 125 Bill. US \$
- sw super storm sturm: 1 – 2 Trillions US \$ in the first year
- 365 transformers in danger (4800nT/min)
- regions in %, with power cuts (4 – 10 years)
- in the US: over 130 Millionen inhabitants

1. SSA: Space Catrina in the US



2. Solar Gamma Rays

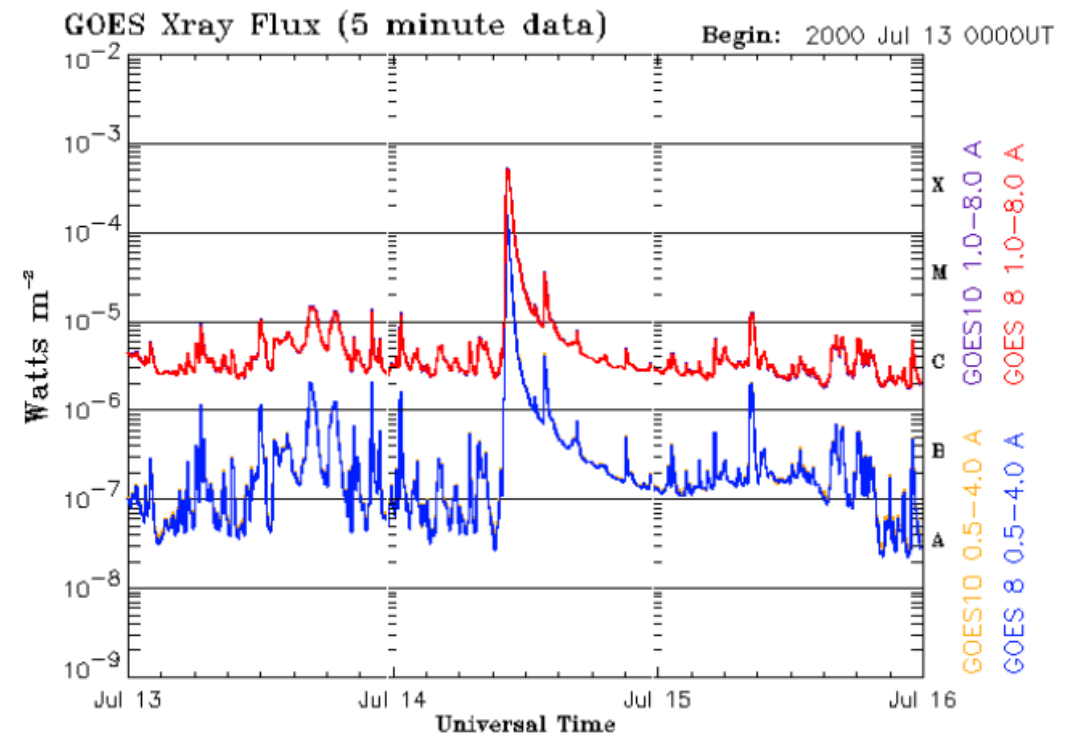
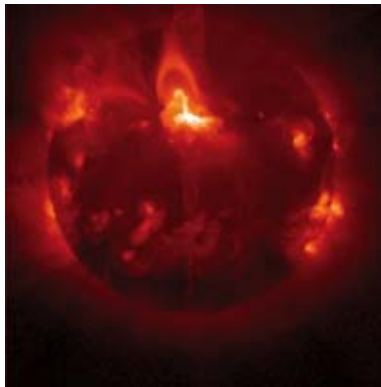
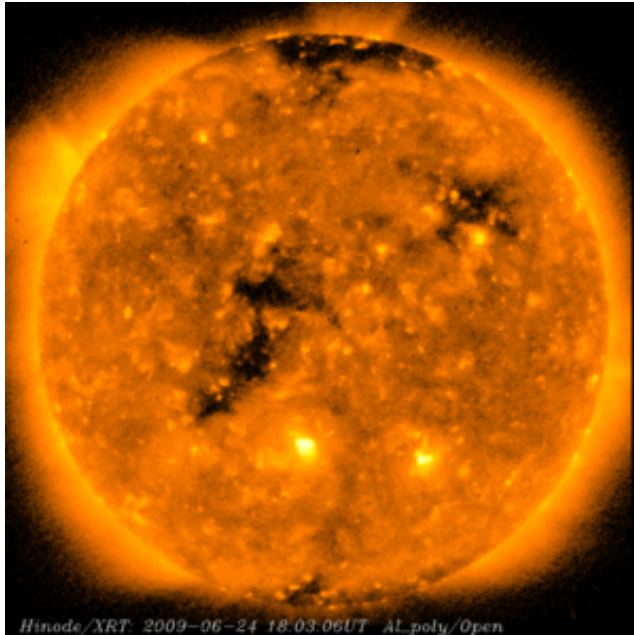


First solar gamma ray flare (nuclear reactions from ions during continuously acceleration for about 2h).

**LOW NUMBER OF
SOLAR GAMMA RAY
IMAGES UNTIL NOW!**

2. Solar Environment – X ray Images / Movies and Spectra

Hinode, Yohkoh, SOHO, GOES



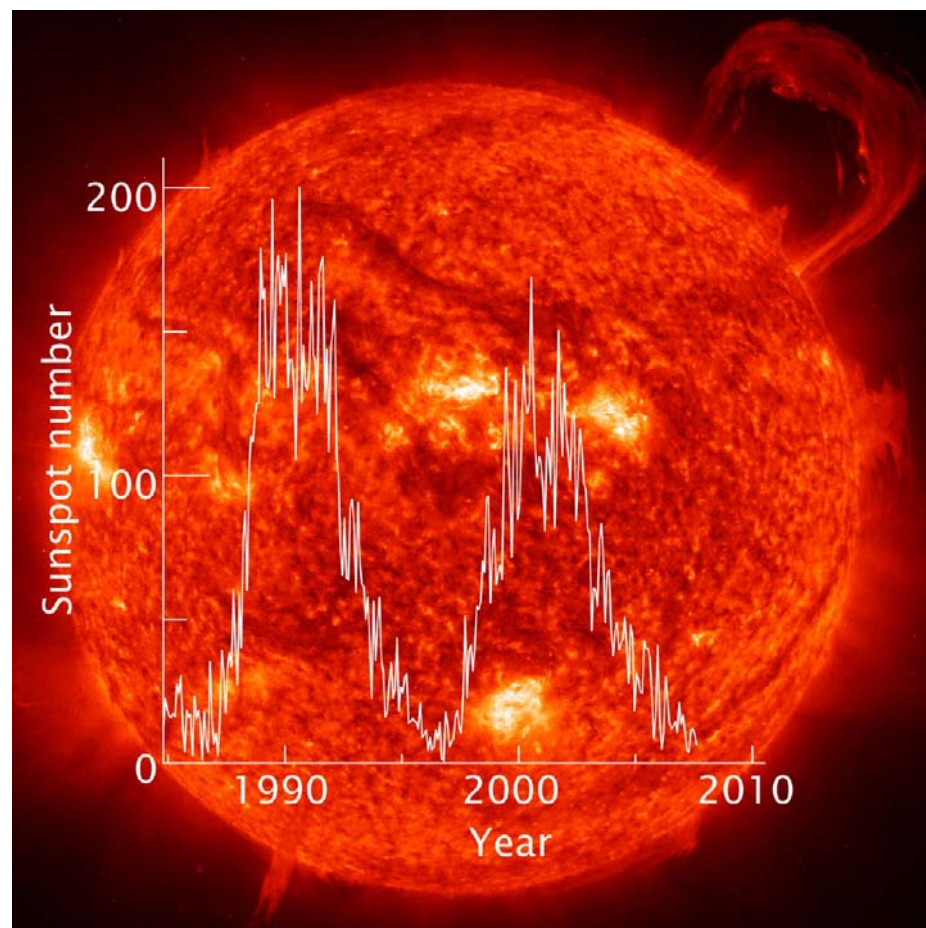
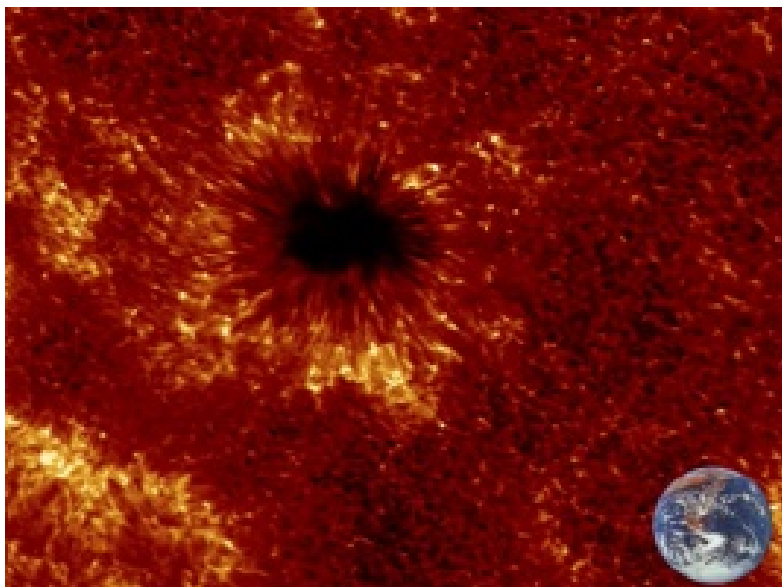
Updated 2000 Jul 15 23:59:04

NOAA/SEC Boulder, CO USA

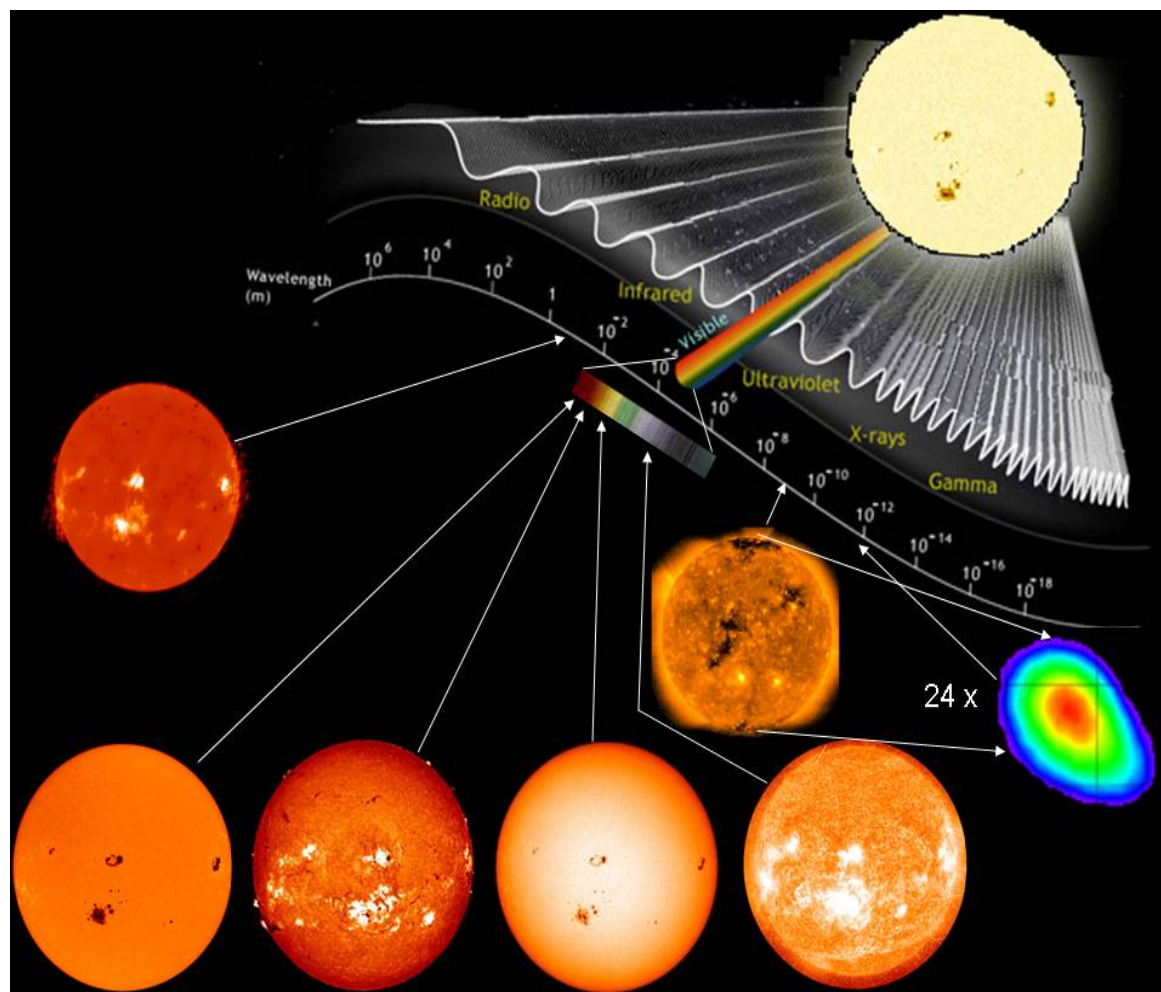


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2. Solar Environment – Sunspots



2. Sun Imaging Summary



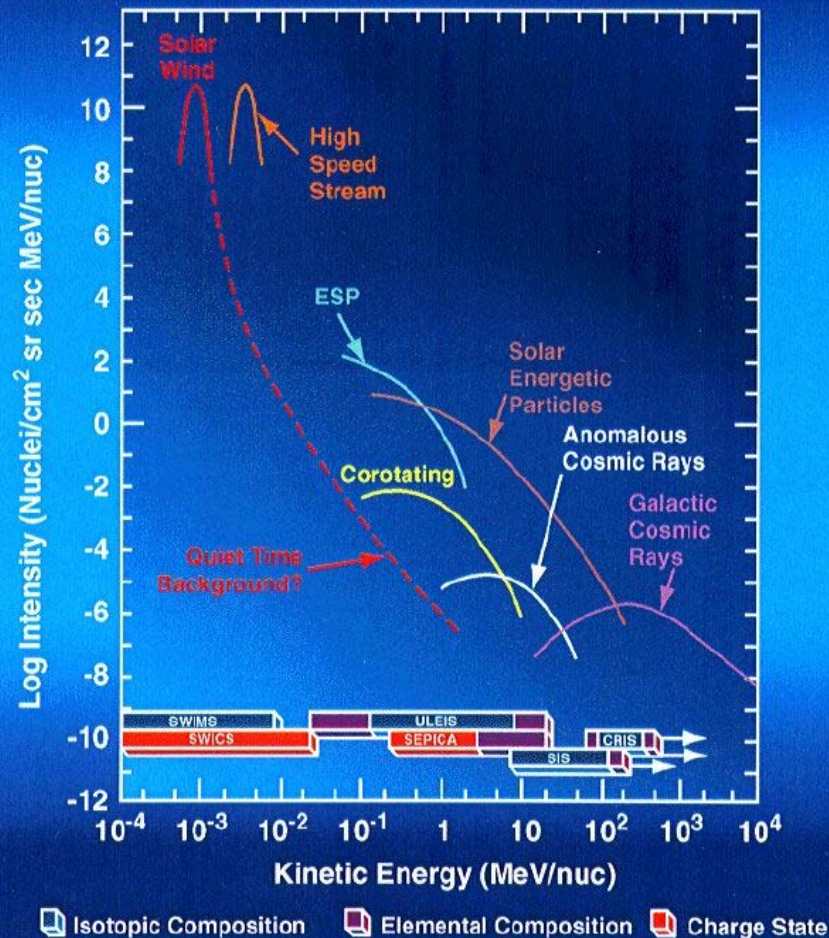
28 October 2003

High resolution
solar gamma ray
imaging is needed!

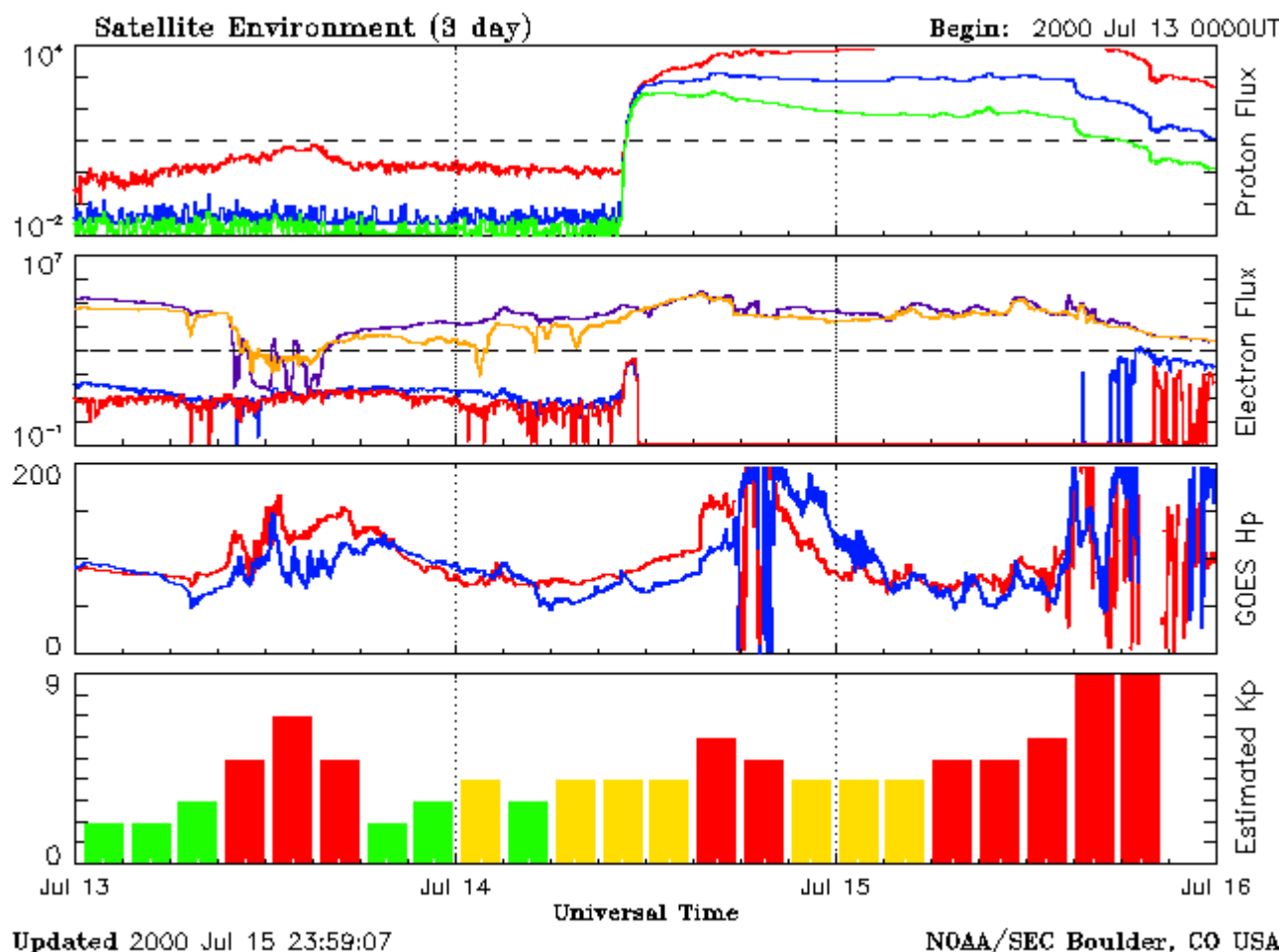


2. High Energy Galactic & Solar Cosmic Rays

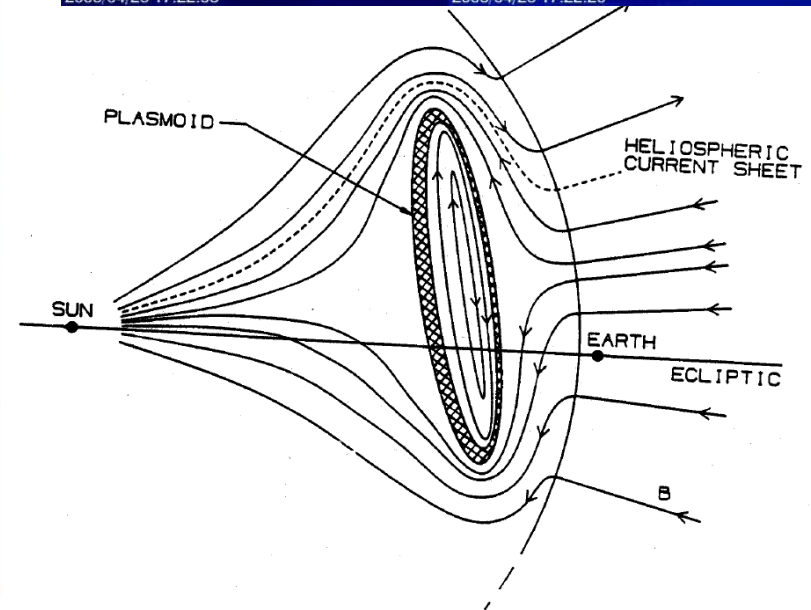
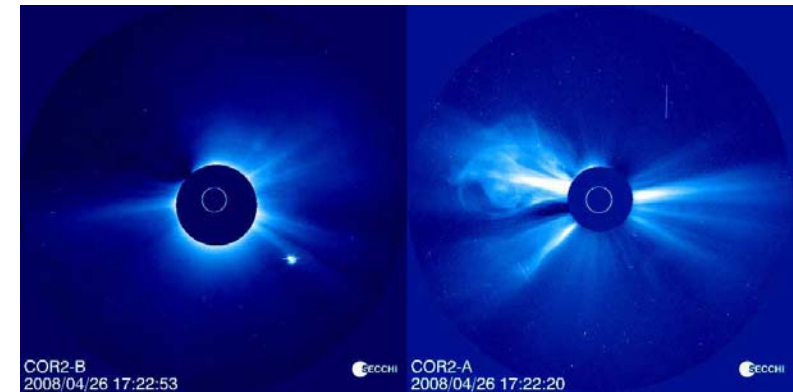
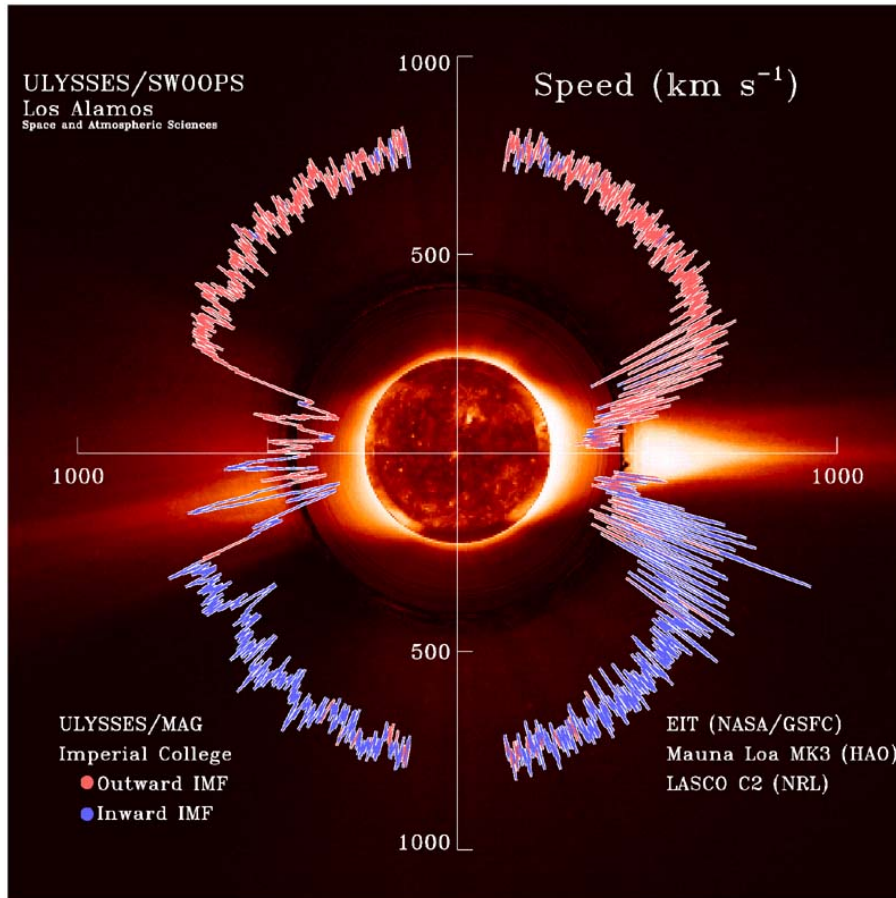
Spectra of Energetic Oxygen Nuclei



2. Solar Energetic Particles (SEP) nearby Earth

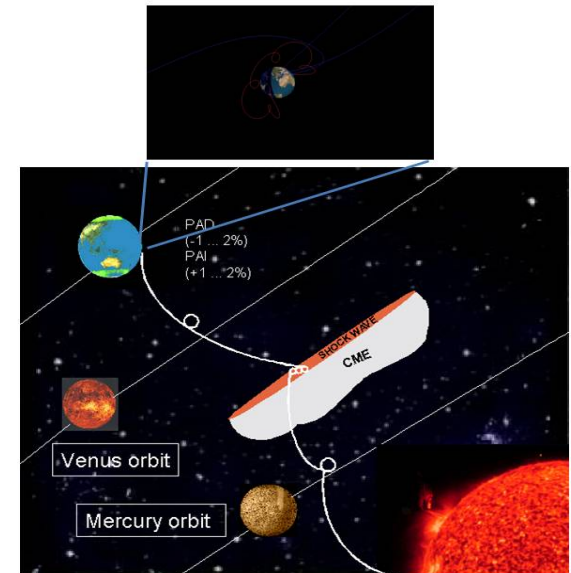
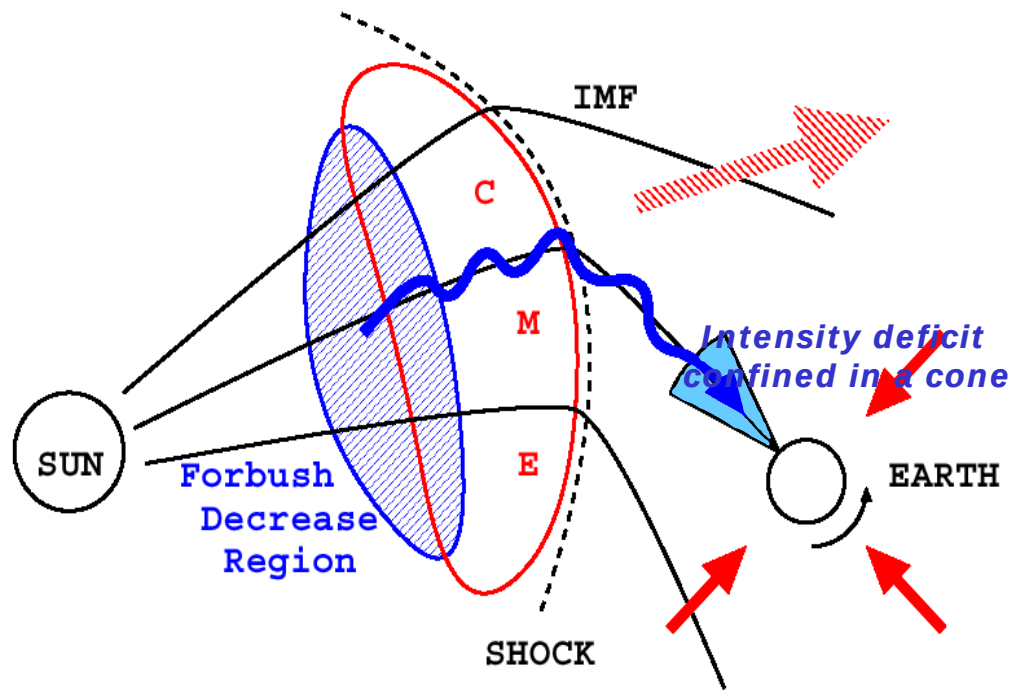


2. Solar Wind and CMEs



Gosling, AGU Geophys. Monogr. 1990

2. CME Detection via CRs



Galactic CRs interact with CME / shock, propagate from FD region to the upstream Earth with the speed of light overtaking the CME shock ahead => isotropic versus anisotropic distribution on ground and in near Earth environment.



2. Principle: Muon Telescopes on Ground

Extra-terrestrial space

Primary cosmic rays(Protons)

Upper atmosphere

Pions

Electrons

Neutrinos

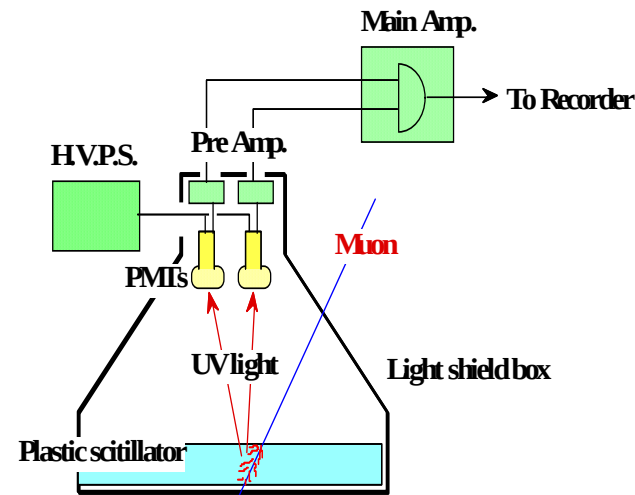
Muons

Muons

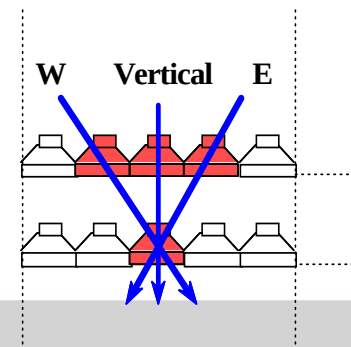
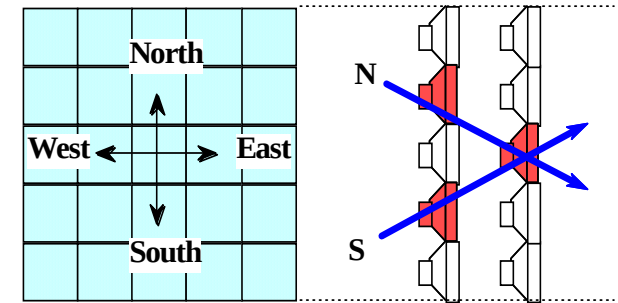
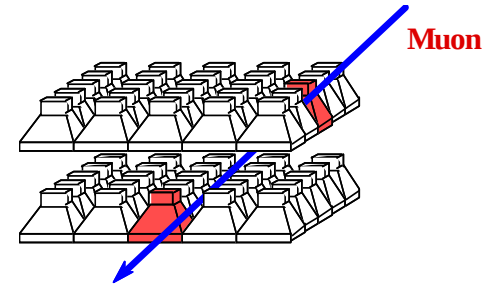
Surface

Underground

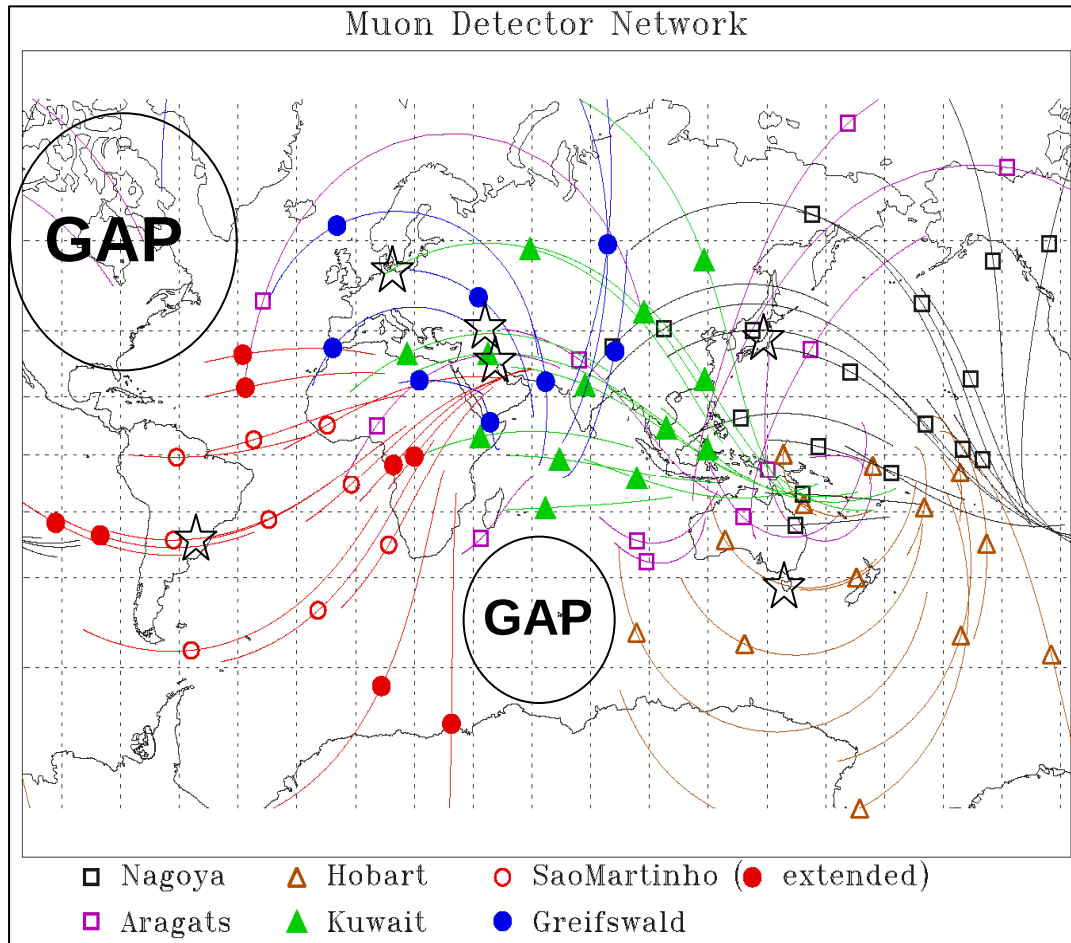
Typical energy of primary ~ 50 GeV for galactic cosmic rays (surface muon detector)



The detector of the Muon Telescopes at SSO / Brazil



2. GMDN: Viewing, Overlapping, Forecast Time



Observing directions:

Nagoya, Japan 17

Hobart, Australia 13

Greifswald, Germany 9

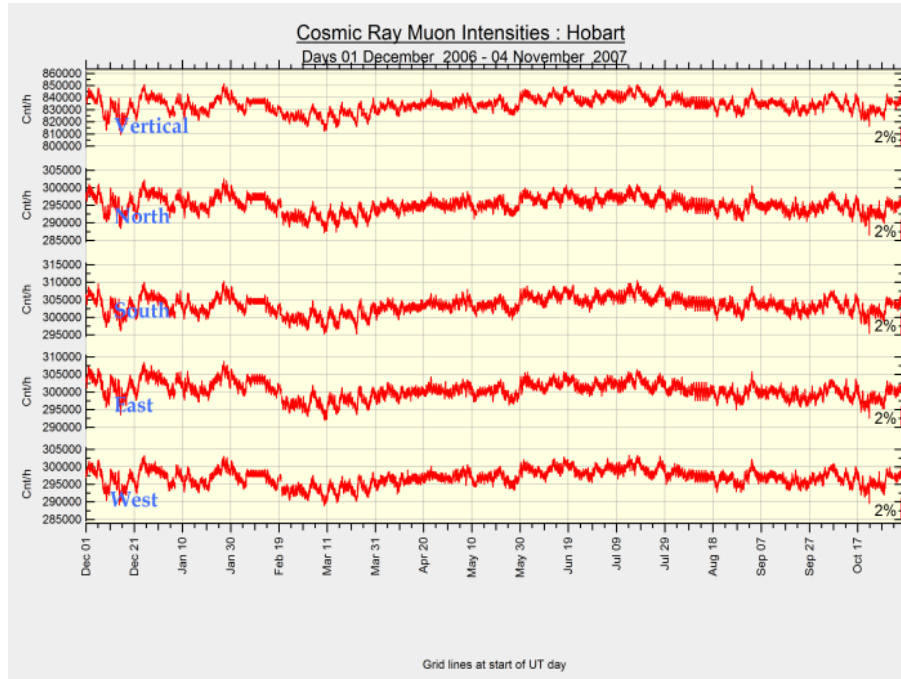
Aragats, Armenia 11

Kuwait, Kuwait 13

São Martinho da Serra, Brazil 17

!!! 10 – 20 hours
in advance the
CME detection !!!

2. GMDN: Counts / Anisotropy



$$I_{i,j}^{cal}(t) = c_{0i,j}^0 I^0(t) + \xi_x^{GEO}(t) (c_{li,j}^l \cos \omega t_i - s_{li,j}^l \sin \omega t_i) \\ + \xi_y^{GEO}(t) (s_{li,j}^l \cos \omega t_i + c_{li,j}^l \sin \omega t_i) \\ + \xi_z^{GEO}(t) c_{li,j}^0$$

$$I^0(t), \xi_x^{GEO}(t), \xi_y^{GEO}(t), \xi_z^{GEO}(t)$$

$$\omega = 2\pi / 24$$

CR muon data: K.Munakata 1, M. Duldig 2, J. Bieber 3, J. Kuwabara 3, E. Flückiger 4, I. Sabbah 5, N.J.Schuch 6, F. Jansen 7

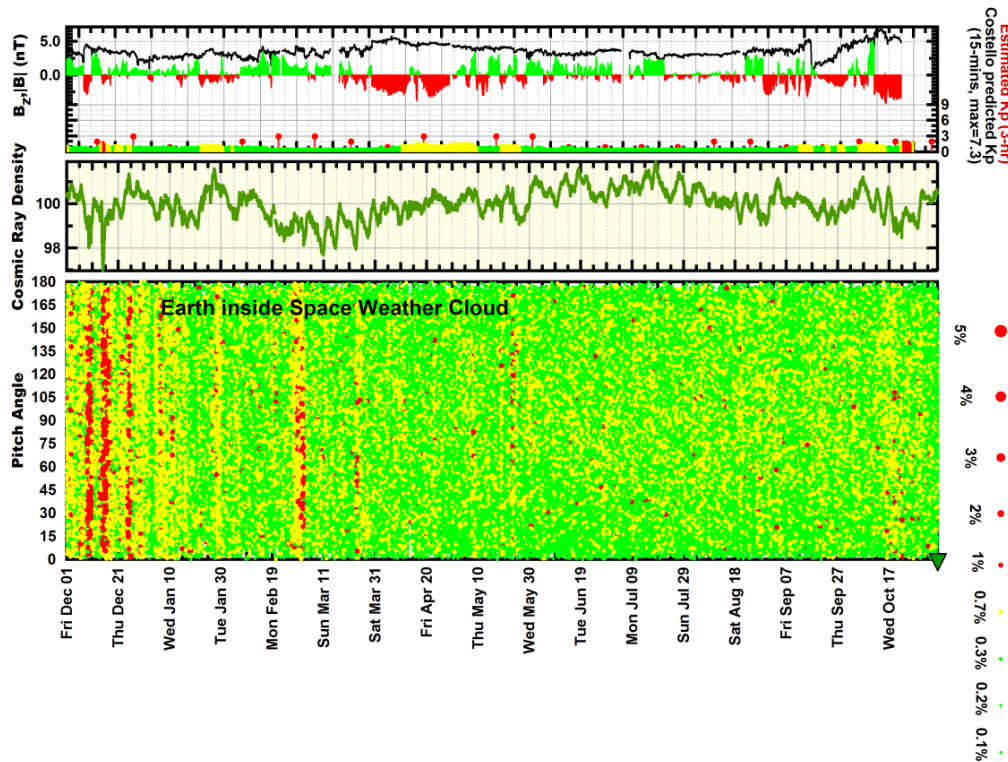
1Shinshu University Japan, 2 Australian Antarctic Division Australia, 3University of Delaware, U.S.A.

4 University of Bern Switzerland, 5 University of Kuwait,

6 Brazilian Southern Space Observatory, 7 DLR Bremen

2. GMDN: SW Service from CR Anisotropy

Real Time Space Weather Cloud Warning (Muon Telescope Network)
Days 01 December 2006 - 04 November 2007



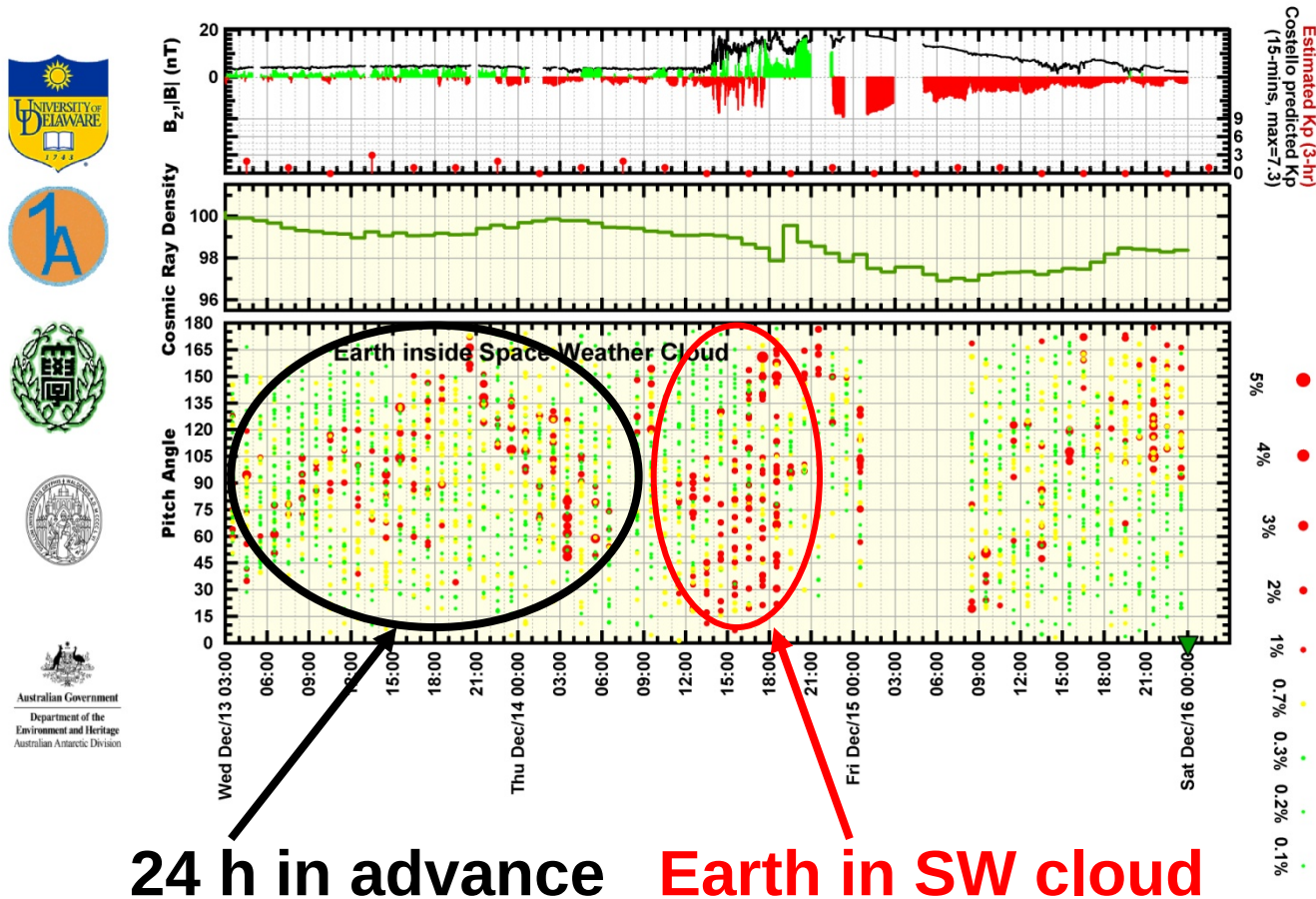
- CR muon anisotropy (%) is a measure of interplanetary CME
- data from CR muon anisotropy: HST, NST, SMST, KPC
- space weather operator friendly displays for Bz, Kp and CR muon anisotropy
- real time service: ACE, CR, CME



2. GMDN: 13 – 15 December 2006

Real Time Space Weather Cloud Warning (Muon Telescope Network)

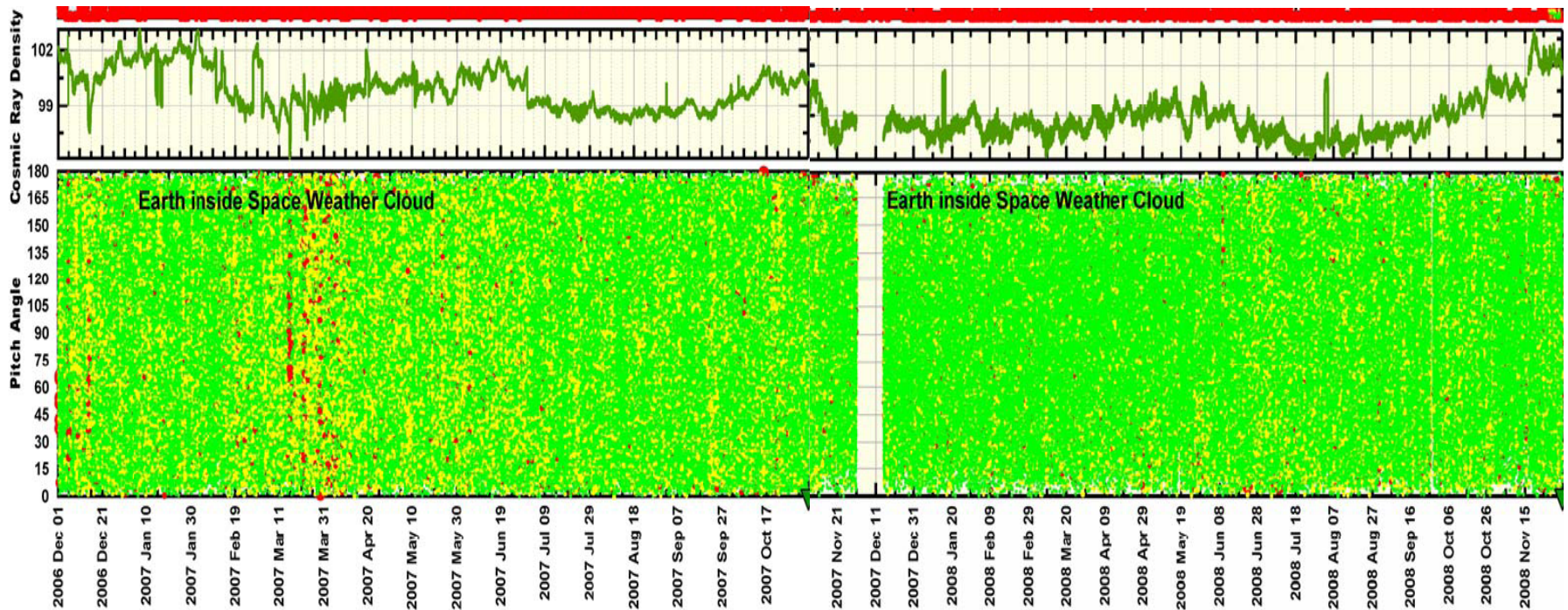
Last 3 days 13 December 2006 - 15 December 2006



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2. Service with GMDN

- user friendly service: real time, early, monthly, 27, 3 days and others plots from GMDN
- University of Delaware / Bartol and DLR Institute of Space Systems / Bremen



3. GMDN: No CME Images - But with Space Based Technology!



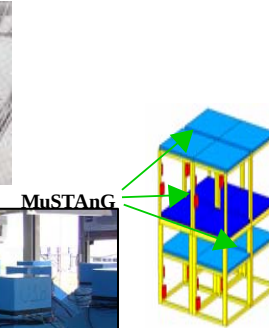
MPC



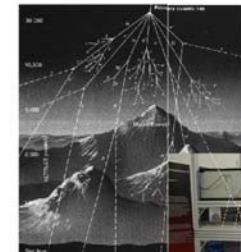
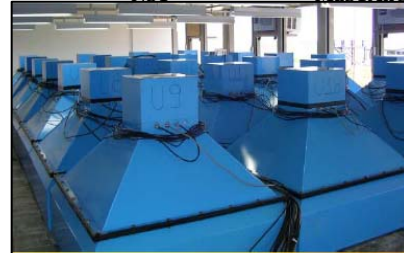
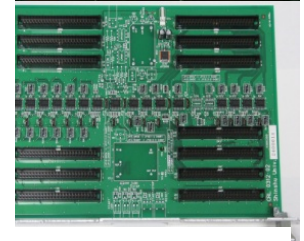
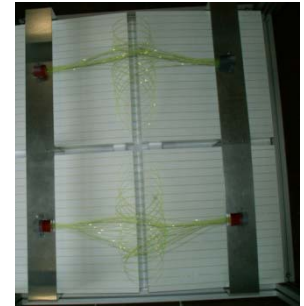
HST



NST



MuSTAnG



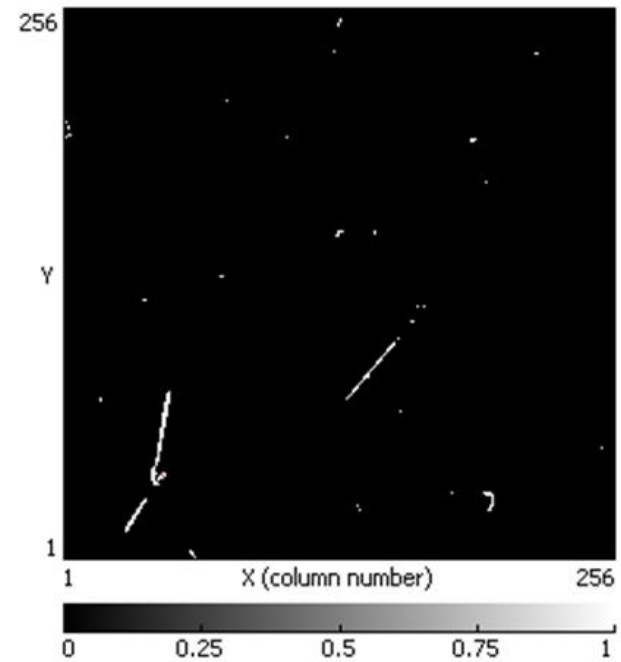
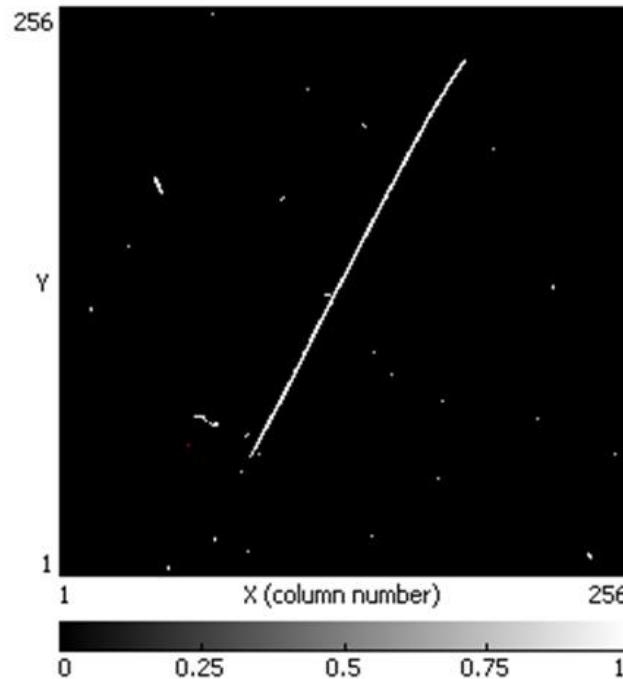
Change from UV telescopes
technology towards CR
telescopes based on highly
miniaturized detector
technology:

Medipix / Timepix

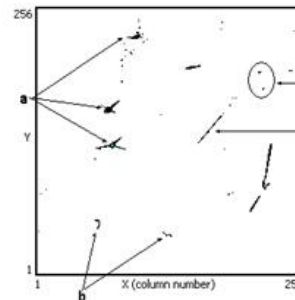


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3. Medipix 2 and CR Balloons Flight



a – heavy charged particles
b – slow light charged particles

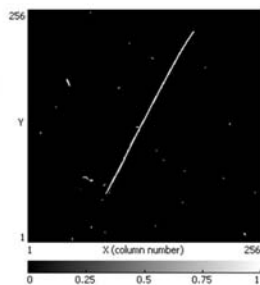
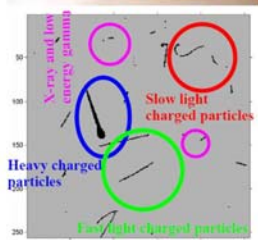
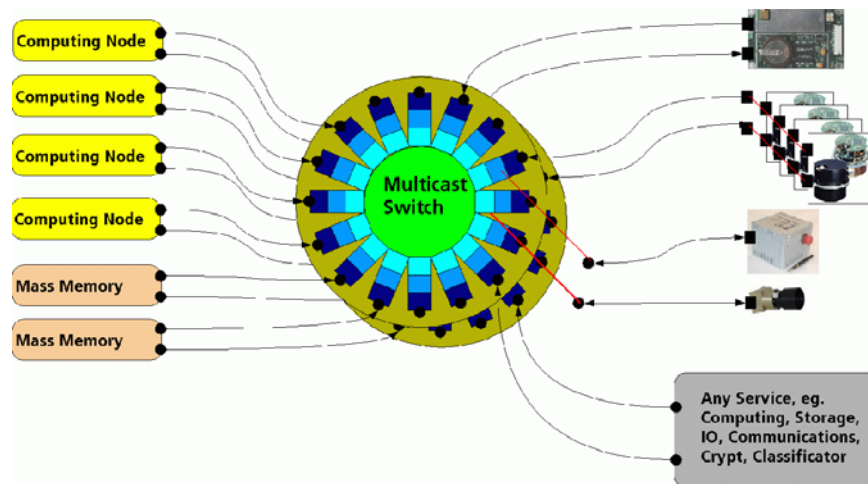


c – fast light charged particles
d – x and low energy gamma rays

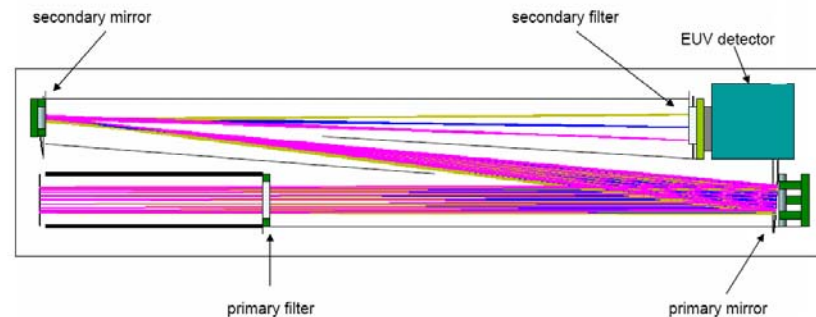
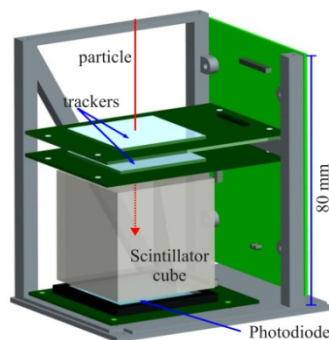
3. NeSTec (New Space Technology) – EU / FP6 Space Proposal



cosine



Medipix



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Conclusions



2010 Bremen: 38th COSPAR Scientific Assembly, SSA session !

THANK YOU



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