Calibration of the ANTARES neutrino telescope



ERLANGEN CENTRE FOR ASTROPARTICLE PHYSICS

Fritsch Ulf ICATPP Villa Olmo Como 5. - 9. October 2009



Friedrich-Alexander-Universität Erlangen-Nürnberg



Outline

- ANTARES neutrino telescope
- Time calibration
- K40 calibration
- Position calibration / Alignment
- Summary



The ANTARES neutrino telescope



Upgoing live event





Time calibration

need precise relative timing of light arrivals at Photomulitplier tubes

- individual PMT transit time spread $\sigma \approx 1.3$ ns
- optical properties of sea water: e.g. scattering $\sigma \approx 1.5$ ns for 40m light propagation
- read-out electronics

Time calibration - Hardware

- Internal clock
- Internal OM LED

transit time evolution for individual PMTs Agilent HLMP-CB15 LED (blue)

LED optical beacon (LOB)
36-LEDs on a hexagonal structure
4 LOBs per line (storey 2, 9, 15 and 21)
Hamamatsu H6780-03 PMT inside

• Laser beacon

2 NdYAG lasers with 1ns (FWHM) pulse duration on line 7 and 8

ERLANGEN CENTRE FOR ASTROPARTICL PHYSICS

Time calibration - Internal clock

 measure time offsets of each storey 20 MHz on-shore clock generator clock transceiver board in each LCM

storey answers through same optical fiber in situ measurements every 4 h show a stable resolution of 100 ps

 address absolute event time (from GPS) with 100 µs accuracy



Time calibration - Internal OM LED

- monitor transit time spread/stability for individual PMT
- Agilent HLMP-CB15 LED illuminating the back of LED the photocathode
- peak 470 nm, 15 nm FWHM





Time calibration - LED optical beacon (LOB)



1 top LED 5 lateral LEDs hexagonal structure

- 36 LEDs per beacon (synchronised to better than 0.1 ns)
- 4 LOBs per ANTARES line (2, 9, 15, 21)
- Hamamatsu H6780-03 PMT inside



Time calibration - LED optical beacon (LOB)





Arrival time distribution after LOB flashing



ERLANGEN CENTRE

ARS T0 correction

Line 1 - 12 interstorey OM-OM time difference





- special local trigger (20 min of data, once a week)
- coincidence rate from integral of events using Gaussian fit → monitor OM efficiencies
- check ARS T0 calibration from monitoring peak offset





Ulf Fritsch for the ANTARES Collaboration - ICATPP09 Como Italy

K40 coincidence rate



- mean 15.8 Hz, RMS 2.4 Hz
- no dependence on bioluminescense
- corresponding to 10% OM efficiency spread



K40 coincidence rate - time dependence



Position calibration - Hardware

per line

- 5 receiver hydrophones
- 1 receiver/emitter hydrophone
 40-60 kHz, attenuation length 0.7-1 km
- 2 sound velocimeters
- connected to conductivity, temperature and pressure probes

per storey

- compass ↔ heading
- tiltmeter \leftrightarrow 2 local tilts



Relative positioning

- triangulation of hydrophone positions (5 storeys) every 2 minutes
- perform a fit on line shape ^(E)/_(S)
 to obtain positions of
 remaining storeys





Tiltmeter/Compass system (TCS)

- TCM 2 and TCM 2.5 cards from PNI Sensor Corporation on each storey
- measurement of 2-axis tilt with 0.2° accuracy and 0.1° resolution for pitch and roll
- measurement of heading with 0.8° accuracy
- measurement of B-Field with 0.05 µT accuracy



Calibration of TCS

- B-field at the ANTARES site $B_x = 0,06 \mu T$, $B_y = 24 \mu T$
- B-fields from electronic container, miscalibration in lab





Line shape model

• buoyancy, drag forces of detector elements



• radial displacement at height z, current velocity v

• performing χ^2 fit and obtain v_x and v_y from fit



Comparison with Acoustic Doppler Current Profiler (ADCP)





Accuracy





Summary

- Time calibration done monthly with a set of Optical Beacons is in good agreement with the K40 analysis.
- Desired precision of electronics (< 0.5 ns) to the overall time resolution is achieved.
- Detector homogeneity: 10% spread in OM efficiency.
- Positioning (every 2 min): Alignment is able to reconstruct positions within 10-20 cm.



Thank you very much!

antares.in2p3.fr









Attenuation length

Measure Q(h) and fit with Q(h) = $Q_0/R^2 \exp(-R/L)$





Absolute positioning

 anchor positions obtained by LFLBL measurement within 1m accuracy after deployment.



• using HFLBL measurements positions are recalculated within 1cm accuracy.



Stability of ARS T0





Time calibration - Internal OM LED



stable within 0.5 ns

ERLANGEN CENTRE FOR ASTROPARTICLE PHYSICS

Charge calibration





Charge calibration



