

The CUORE and CUPID Experience





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Cryogenic Calorimeters



- Convert **energy** in **temperature**: $\Delta T \sim \Delta E/C$
- To obtain sizable ΔT we need a small C
- Crystals cooled at 10 mK: C ~ T³
- ΔT converted into electrical signals





CUORE

- O 988 TeO₂ crystals (~1 ton) + other 2 tons of copper at 10 mK
- O ~20 tons at different T stages



Physics Motivation

Main goal: search for **0vββ**: hypothesized, **never observed**, nuclear transition



- Forbidden by SM: it violates L (actually B-L) conservation
- O Can occur only if v is a Majorana particle
- O It creates matter (no anti-matter balancing)
- O Majorana phases: other sources of CPV?

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But large mass, low threshold, low background and Quenching~1:

- Annual modulation of nuclear recoils induced by
 WIMPs (tens of keV for ~100 GeV WIMPs)
- O Solar axions by inverse Primakoff effect: M1 transition of ⁵⁷Fe, expected at 14.4 keV



The Nightmare: Radio-Activity

- O In CUORE (but not only) the **external background** can be effectively **suppressed**
- O The detector itself becomes the main problem!
- O Challenge: such small contamination levels required, that **no screening techniques** available!

ACTIVE MATERIALS

Can often be used as detectors to assess their own contaminations

PASSIVE MATERIALS

Can be more problematic, especially when dealing with surface contaminations

Crystals

- Used by many experiments as calorimeters, scintillators...
- CUORE (TeO₂), CUPID (ZnSe and LiMoO₄), AMoRE (⁴⁰Ca¹⁰⁰MoO₄), GERDA and Majorana (Ge)...
- CRESST (CaWO₄), Edelweiss, CoGeNT (Ge)...
- Nal (DAMA, SABRE, ANAIS, DM-Ice..)
- ...and many others



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- Contaminations from ⁴⁰K, ²³⁸U, ²³²Th and daughters
- Cosmogenic activation of materials
- Selection and monitoring of all materials, tools and facilities for crystal growth
- Selection of equipment for chemical and mechanical processing, storage



An excellent result: TeO₂



Fig. 2 Radio-purity certification protocol applied for TeO₂ crystals.

- Must be <3x10⁻¹³ g/g in ²³⁸U, ²³²Th
- Contribution of I.Dafinei (INFN Roma)
- Strict protocol for each step of crystal production



- In clean room, etching and polishing
- Packaging and underground storage

doi:10.1016/j.jcrysgro.2010.06.034

Screening Technologies

Each stage of production is monitored accurately:

- ICP-MS for contaminations in ²³²Th and ²³⁸U in Te, TeO₂ and consumables and in all other reagents. Sensitivity of 10⁻¹² g/g achievable.
- HPGe to be sensitive also to broken chains and ⁴⁰K. For our purposes, 10⁻¹⁰ g/g on Th and U was enough, but 10⁻¹² g/g achievable (long run)
- SBD for surface contaminations of specific materials ad monitoring of selected components (in particular for lapping cloths, packaging...)





Cryogenic test

Bulk: <1.8x10⁻¹⁴ g/g in ²³⁸U and <5.5x10⁻¹⁴ g/g in ²³²Th

Passive Materials: Copper

Good mechanical/thermal properties, radio-pure —> structural parts of many detectors





1st problem: how to make it more radio-pure

2nd problem: how to check the purity level

Electro-formed Copper



- When electro-plating copper from a solution onto cathodes, most of the contaminants do not follow copper (different electrochemical value)
- Electrodeposition OFHC copper on stainless steel forms
- Everything done underground (reduce activation)
- Outstanding purity: 0.3 µBq/kg for ²³²Th and ²³⁸U

- Originally proposed at PNLL
- Now CES (Copper Electroforming Process) at CANFRANC
- Not (yet) able to clean surface of the materials but efficient for the purification in many materials



3D-Printed Copper



OFHC copper converted into powder and used via 3D-printers in controlled environment

Radio-purity measured at LNGS

Improvement of surface roughness at LNL

Characterization at low temperatures at Roma (N.Casali)



Radio-Assay

Neutron Activation Analysis and HPGe spectroscopy showed no traces for contaminations

	µBq/kg	g/g
²³² Th	<2	<0.5x10 ⁻¹²
238	<70	<6x10 ⁻¹²

Study of surface contaminations is difficult

doi:10.1016/j.astropartphys.2013.02.005





T1 T2 T3 Large arrays provide sensitivities of tens of nBq/cm²

For next generation detectors, we must ensure an even better purity (~ a factor of 10)

This becomes a problem also for assay

Improving Sensitivity: Bulk



- ICP-MS with pre-concentration of contaminants
- Current goal: automation and ⁴⁰K assay

Material	Th, U Detection Limits	
	μ Bq/kg	ppt
Copper (Electroformed or commercial OFHC)	<0.1	<0.01
Lead	<1	<0.1
Titanium	<1	<0.1
Stainless Steel	<1	<0.1
Polymers (PTFE, PVDF, Acrylic, Bioabsorbables, etc.)	<1	<0.1
Linear Alkyl Benzene (LAB)	<0.1	<0.01
Quartz, Fused Silica	<1	<0.1
Electronic Components: FETs, resistors, thermocouples, etc	<0.1 pg/component	
Solutions	<0.01	<0.001

Improving Sensitivity: Surface

Scintillating bolometers allow to achieve an extremely high sensitivity on surface contaminations in a few weeks of measurements.

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Advantages:

- · Simultaneously α and γ spectroscopy
- Wide active surface (hundreds of cm²)
- No dead layers
- Excellent energy resolution (< %)
- Low intrinsic background

with these features, we can achieve a ~tens of nBq/ cm² in a few weeks

http://stacks.iop.org/1748-0221/7/i=10/a=P10022



Conclusions

- A long R&D on materials allowed to reach a very low background with CUORE
- The development of pure material is crucial for many experiments in different fields
- Most of the purity levels are extremely difficult to assess with conventional techniques
 - New, versatile technologies have been proposed

M2 spectra (X-rays)

Efficiency - corrected energy spectrum



Recoiling Nuclei	Eexpected	Emeasured	QF
²⁰⁶ Pb	103.12	95.62 ± 0.24	0.927 ± 0.002 (stat.)
²¹⁸ Po	100.8	100.0 ± 0.9	0.992 ± 0.009 (stat.)
²²⁰ Rn	103.50	100.45 ± 1.21	0.971 ± 0.012 (stat.)
²¹⁴ Pb	112.13	110.92 ± 0.96	0.989 ± 0.009 (stat.)

Table 3 Expected energy, measured energy obtained from fits, and resulting quenching factors (QFs) for the selected recoiling nuclei. Only statistical uncertainties are shown in the QFs.

- Assuming coherent scattering (A²) TeO2 is interesting both for high and low masses
- Only spin-independent
- Coherent isospin-invariant coupling and the Helm model for nuclear form factors
- For WIMP velocity, standard halo model (SHM)
- WIMP local density 0.3 GeV/c², local circular velocity 220 km/s, galactic escape velocity 650 km/h, orbital
 velocity of Earth around the Sun 29.8 km/s,
- For each (mW, σ SI) we generate 100 toy MC simulations
- For each MC spectrum, maximum likelihood analysis for annual modulation and no modulation hypothesis
- Significance of modulation quoted as log-likelihood ratio of the best fits
- The likelihood is calculated using a PDF containing the target mass, the detection efficiency (BoD) and the background pdf (no temporal dependency)
- ROI: 10-28 keV to exclude peaking background (most of the signal expected at low energy)
- Uncertainty on energy scale dominated by QF (consevatively 7%)
- Sensitivity computed requiring a 90% CL in 90% of the toy experiments