







H2020-MSCA-RISE-2016 — Grant Agreement N°





































Massimiliano Razzano SB Meeting - June 25, 2018









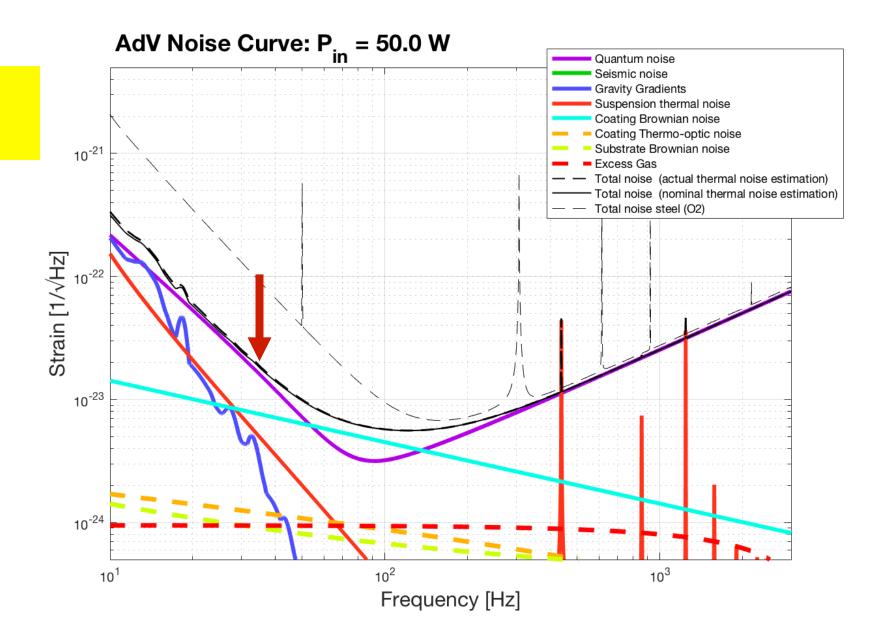


NEWS from GW Physics

- Second observing run (O2) finished successfully this summer
- Joint LIGO-Virgo run (1-25 August) yielded 2 detections
- Now moving toward O3
- Improving detectors and data analysis infrastructures
- Higher sensitivity → higher events (and noise transients rate)
- Both search pipelines and detector characterization activities
- Aiming at lowest latency possible



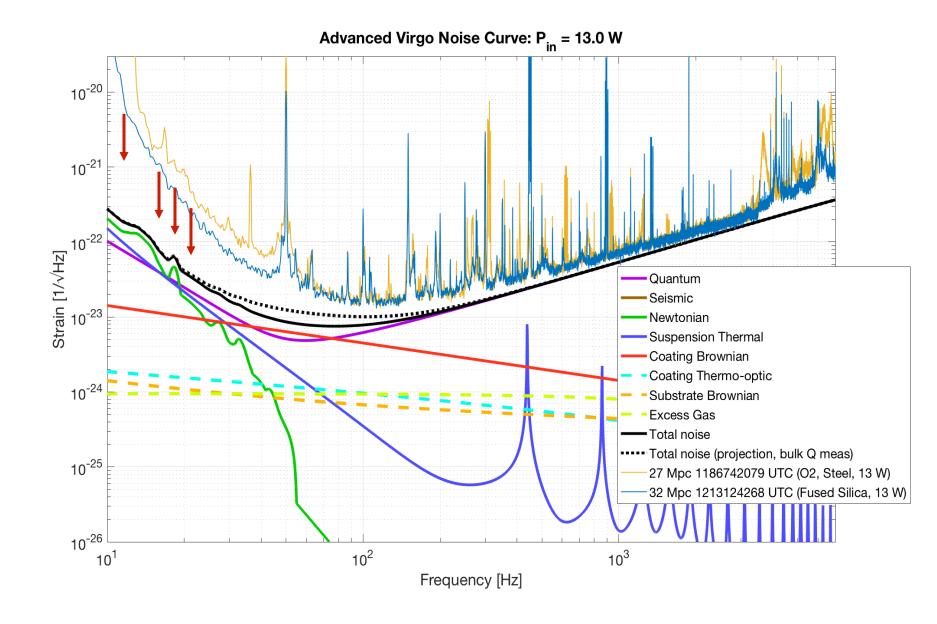
Summary O3 Target





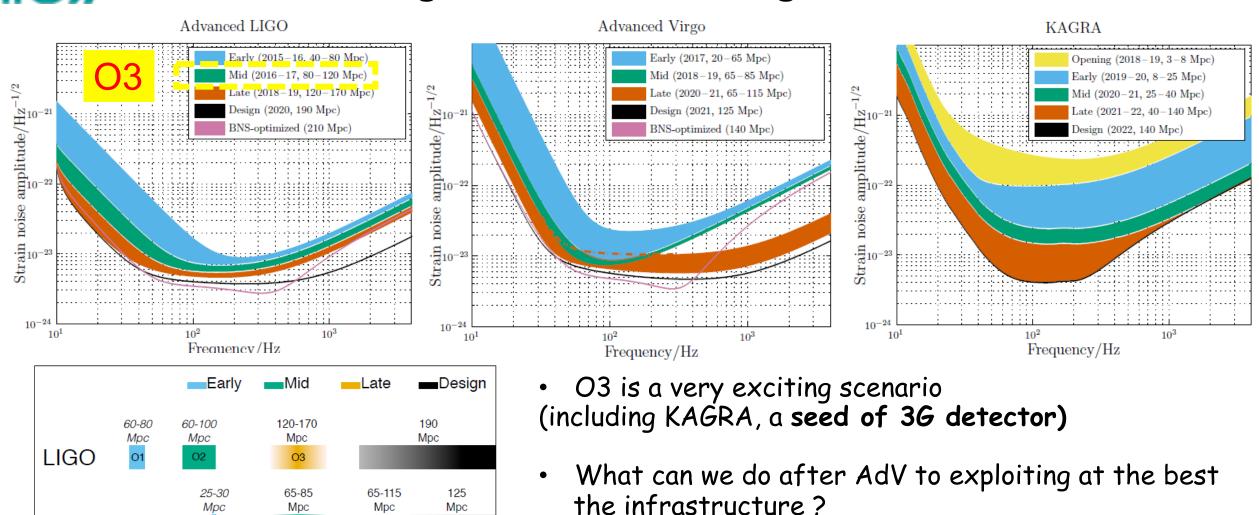
Mid June

- LF gradually improving
- Preliminary estimate of bulk TN on going
- Effect of fused silica fibres clear
- Noise hunting dedicated to:
- o Control noise
- Environmental
- Stray light





LIGO-Virgo-KAGRA observing scenario



• AdV+ vision document submitted to Virgo Council: a two-step approach, 160 Mpc/300 Mpc, 6-year

Living Rev. Relativ. (2016) 19: 1

02

40-140

2023

Virgo

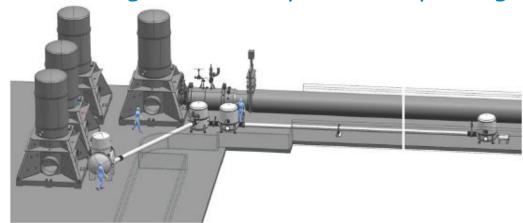
KAGRA



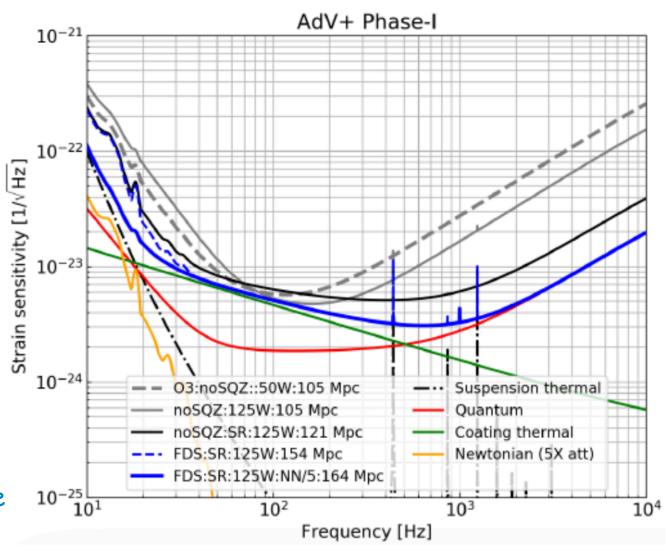
MOM AdV+ vision document, phase I: reaching AdV full target

Complete the AdV program: 200 W laser; 125W at the ITF input Signal recycling → 120 Mpc range

Frequency Dependent Squeezing 300 m-long filter cavity →150 Mpc range



Newtonian Noise Cancellation →160 Mpc range





MOJJ AdV+ vision document phase II: Thermal noise reduction large beams

Larger mirrors

Diameter: 550 mm, thickness: 200 mm, mass: 100-120 kg

Scenario 1: **ETM-only** \rightarrow 200 Mpc range

Scenario 2: Full \rightarrow 230 Mpc range

In parallel coating improvements:

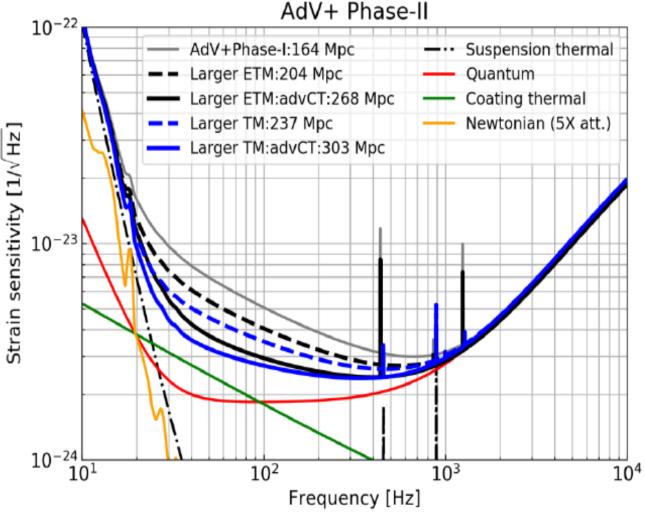
If factor 3 reduction in Coating TN:

Scenario 1: ETM-only \rightarrow 260 Mpc range

Scenario 2: full upgrade → 300 Mpc range

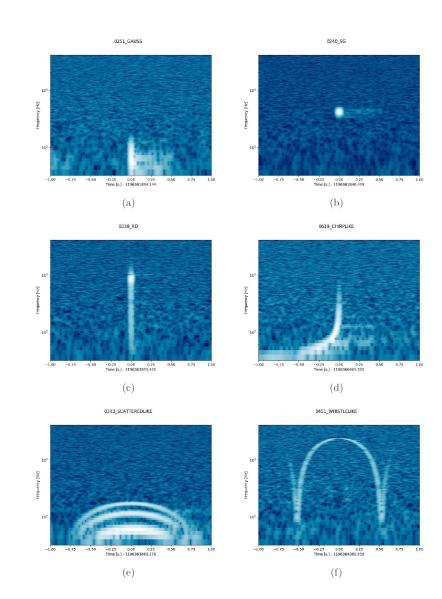
Several challenges, feasibility under study:

- Handling large masses (100kg-150kg for BS)
- Handling and integrating the payloads into present and integrating the payloads and integratin modified VACUUM system
- Folded VS marginally stable Recycling FP
- Design of very large baffles



Ongoing projects - detector characterization

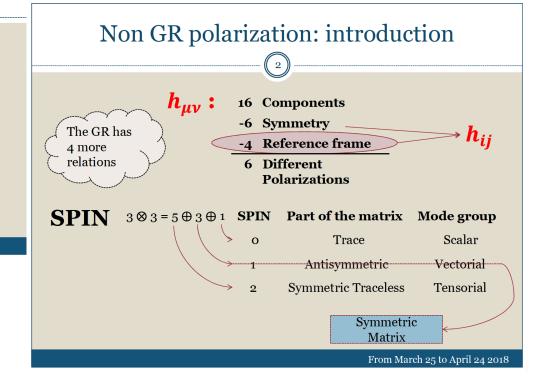
- Deep learning based pipeline implemented
- Tested on simulations
 - Details in Razzano & Cuoco, 218, CQG,35,9
- Now moving to Virgo data, in progress



US

G. Intini, 1 secondment new algorithm to observe the presence of non-GR polarization is crucial for testing the Einstein equations for General Relativity

5-vectors with non-GR polarizations



Difference with respect to Max Isi (2015, PRD 91)

5

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$
 Breathing: $\mathbf{e}_{b} = 2\mathbf{e}_{t} + \frac{2}{3}\mathbf{e}_{d}$

$$\begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$
 Longitudinal: $\mathbf{e}_{l} = \mathbf{e}_{t} - \frac{2}{3}\mathbf{e}_{d}$

- In the (b,l) base:
 - there are no undetectable polarizations
 - b and l are <u>degenerate</u> and <u>undistinguishable</u> (same antenna pattern)
 - we are mixing different spin particle
- In the (d,t) base:
 - there are no degenerate polarizations
 - t is undetectable in Virgo/LIGO but <u>may be detectable</u> in other interferometers

From March 25 to April 24 2018

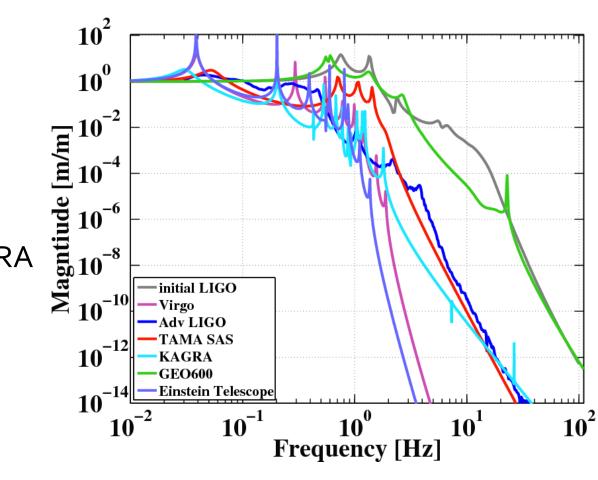
JAPAN

E. Majorana, 1 secondment from June 27° 2018 :

- 4° International KAGRA Workshop
- Sesimic Suspension performance tests for KAGRA
- Payload works
- 3° Generation detector roadmap

US

M. Razzano, 1 secondment from July 2018



Ongoing projects - Open Data portal

- Working with LIGO colleagues to new portal for GW open data
- New version of existing LOSC
- Close collaboration with Caltech researchers
- Hope to launch soon



Conclusions

- GW astronomy has begun
- Dawn of multimessenger observations
- Activities toward next run (O₃)
- Main topics
 - Detector characterization
 - Open data portal
 - Other?
- Secondments started, will continue during this year