



Novel technologies for direct dark matter detection

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Outline



Not a review, but a (very) personal view

Directionality - a tool to reject background

- Iow density gas detector
- anisotropic targets (CNT)
- Sub-GeV dark matter
 - detecting electrons

Direct detection: the name of the game

- No-one knows *what* a dark matter particle is
- WIMP model: non relativistic 10-1000 GeV particles with cross section much larger than solar neutrino weak cross section



DAMA/Libra new results

Cerc European Research Council Endelshed by European Convension

WIMP "wind"

- Year-modulated signal in Nal crystal target (ton-y exposure)
- Confirmed in phase 2 data
- All the interpretation in terms of background not convincing so far.

DAMA/LIBRA-phase1+DAMA/LIBRA-phase2 (2.17 ton × yr)



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Directionality

- WIMP must appear as coming from CYGNUS
- Nuclear recoils must reflect this feature (*dipole* distribution)
- Radioactive background is isotropic
- Solar **neutrinos** comes from the **Sun**!



One m³ low pressure gas TPC



The DRIFT detector (Boulby)

Astropart.Phys. 91 (2017) 65-74





 Measuring
 Head-tail of the recoiling nuclei

- Very small mass: gas mixture CS₂ : CF₄ : O₂ (30-10-1 torr) amplification with wires
- Full electron rejection up to 30 keV_{ee}
- Negative ion drift, reduced diffusion
- Fiducialization with multiple charge carries

Low mass WIMPs (still wiMps??)

- Lowering threshold in cryogenic detectors (CRESST, CDMS, DAMIC)
- Lowering
 mass of the target nucleus
- Gamma rejection difficult at low WIMP masses: new techniques ?







He based detector: CYGN-O

- Atmospheric pressure gas TPC (He:CF4)
- He recoils, more sensitive to low masses
- Amplification with GEM
- High resolution Optical readout
- Aim at very low threshold (~ keV)





CYGNO detector concept





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- Read the **light from the GEM** discharge with a low noise s**CMOS camera**
- 2D X-Y resolution at 100 µm
- No interference with gas
- Use PMT as trigger
- Combine PMT and camera for directionality

Can actually "image" the recoiling particles



R&D, CYGNUS - 10 liter prototypes









A snapshot of radioactivity

TPC with electron drift in He-CF4 (60-40) gas mixture

 Use 2D projection to separate *electron recoils* from *nuclear recoil* (specific ionisation)







Electron mip tracks



450 MeV electron tracks producing ionisation cluster, drifting to GEM, flash of lights during GEM discharge



HeCF₄(60/40) by Garfield simulation

dE/dx @ 400 MeV = 0.23 keV/mm about ~ 7 primary e⁻/ mm V_{drift} = 4.5 cm/µs @ 640 V/cm



Resolution for a high energy electron track is 30-60 μm

D.Pinci et al., PoS EPS-HEP2017 (2017) 077



Energy resolution



Small prototype 1cm drift gap exposed to **5.9 keV** (⁵⁵Fe) gamma rays and relativistic electrons (**2.3 keV/cm**)





Measurements with neutrons

Small prototype exposed to an Am-Be source: **1-10 MeV neutrons** along with 4 MeV and 60 keV photons. A 0.2 T magnetic field within the drift field (a permanent magnet).

Low energy electrons due to X rays





MeV electrons due to 4 MeV γ

He nuclear recoils (**α**)







Ionization density measured by counting photons per unit length/area







Towards a large volume DM detector ?



- Optical readout of large surface seems promising
 - Exploiting GEM high gain
 - Large market of cameras, higher resolution and lower prices in the future ?
- Need to address key points
 - Low radioactive material
 - Fiducialization (possible use of negative ion drift ? add SF₆?)
- An innovative fast neutron detector (measure background at LNGS)



Building a collaboration for CYGNO, a demonstrator in the context of the CYGNUS-TPC collaborative network towards a 1000 cubic meter multi-site nuclear recoil observatory



Solid target: CNT



- Idea: WIMP scatters on a *anisotropic* target as *aligned* carbon nanotubes.
- Nuclear recoils are exiting the target only when along the CNT axis - otherwise, absorbed!

collaboration University of Mons, Belgium



length: 100 μ m (can be increased) ext. diameter: (20 ± 4) nm aspect ratio: $5x10^4$



length: 75 μm ext. diameter: (13 ± 4) nm aspect ratio: 0.6 x10⁴ detector side

absorbing substrate



C ion moving within the array



A prediction for an "acceptance" channeling angle of 35 deg is made



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Investigating CNT structure







Scanning Electron Microscopy (SEM) at sub- μ m scale CNIS lab. @ Sapienza





Raman spectroscopy at sub- μ m scale Phys. Dept. @ Sapienza



IR inelastic scattering: vibrational structure of the carbon lattice, defects, ...







- Spatially resolved, can asses the quality of the CNT bonds at various heights of the target
- Light can be focused at various depths in the interior of the target (up to few 10 µm)



Ar⁺ ion beam on CNT



▶ 5 keV Ar+ beam onto a CNT at *different* angles with respect to CNT axis



Number of **defects** measures the **penetration** of Ar ions When Ar beam **aligned** with CNT, defects present **at all the heights** When Ar beam **orthogonal** to CNT, defects present **only on the surface**





Towards a prototype detector

Indications that low energy Ar ion might be "channeled" in CNT

- More tests going-on, different beam intensity, different kinetic energy (< 1 keV)
- Try to confirm the prediction of 35 deg critical angle for channeling (ion trapped within the interstices among CNT)
- If confirmed, try to detect a single scattered C ions
 - use a relativistic electron beam to scatter C out of the CNT



WIMP Mass $[GeV/c^2]$





What if dark matter is not so massive ? Scattering over the target *electrons*

Sub-GeV dark matter





US Cosmic Visions: New Ideas in Dark Matter 2017: Community Report : https://arxiv.org/abs/1707.04591

Electron emission from a cathode



What about a DM particle scattering off an electron ? a dark-cathode ?

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Work function of CNT is > 4 eV All these effects are suppressed: room temperature is low enough, UV photon efficiently screened, E field < 100 V/μm

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Electron emitted from aligned CNT



Electron extracted by a DM scattering

Few eV energy electrons are recoiling off



graphene interactions are **suppressed** at this energy (compare *e* wavelength) Electron collected by an external electric field *E*

- electrons can be transmitted, reflected *absorbed* by a graphene sheet
- absorption $\sim 10^{-3}$ (but no good data available)

Directionality





Different rate at different angles θ_w

 $\theta_w\,{\sim}90$ preferred by graphene electron wave function

A rate asymmetry can be measured by comparing two CNT target orientation

With an exposure of 100g * 200 day a 5σ non null asymmetry can be measured

Sensitivity region



G.Cavoto et al, Phys.Lett. B776 (2018) 338-344



Two arrays of dark PMT (~10⁴ units, 10mg dark cathode mass each)





- Build a dark PMT prototype out of a CNT as a cathode, an electric field and an electron sensors (an avalanche photodiode in vacuum).
 - A few eV single electron detector
- Synergy with PTOLEMY

(PonTecorvo Observatory for Light Early universe Massive neutrino Yield).

- PTOLEMY's concept uses a target of tritium doped graphene layers for relic neutrinos
- Need to asses the radioactivity of graphene based structure: a dark PMT could be a tool for it
- Electron graphene interactions must be studied.

My conclusion (for direct searches)



- WIMP is a great paradigm, but no firm experimental evidence so far.
 - DAMA/LIBRA result yet unconfirmed with different target nuclei.
 - "Several ton-year "is the new frontier in the high mass (> 10 GeV) region
- Need to prepare to dig into the neutrino floor
 - directional tools need to be explored now (anisotropic targets, low density gas and large volume)
- Look in other mass range, well below GeV
 - electron recoils

Besides brute force, we (desperately?) need some brave (experimental) idea.

And - possibly - some luck (since nobody really knows where exactly to look).

Back-up slides



Energy spectrum of modulation

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- Lower PMT threshold (higher Q.E. PMT) can shed light on different models of the WIMP distributions (halo, stream, etc.)
- Diurnal modulation might be investigated (large exposure)

If an anisotropic scintillation crystal is found(ZnWO?), measure the WIMP direction!

Some effect of the nuclear form factor

But other scheme are possible, interaction with electron, inelastic interaction...

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Gas TPC DRIFT: negative ion drift TPC

- Limited diffusion ever over long drift distance (<0.5mm on 0.5m)
- If anode segmented, a "ion recoil track" can be reconstructed: direction
- Head-tail information is valuable as well (might be enough for discrimination)

DRIFT fiducialization

Astropart.Phys. 91 (2017) 65-74

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Light readout

Light production in a GEM discharge readout with low noise CMOS camera
For a min, track we predict (Carfield)

Camera calibration, pedestal is less than *2* photon per pixel (*"photon" noise equivalent*)

Very good signal to noise ratio! For a m.i.p. track we predict (Garfield) ~ 7e⁻/mm (primary ionisation electrons) For a m.i.p. track we **measure** ~ 1000 photons/mm that means ~150 photons/primary electron

X-RAYS FROM A 55FE SOURCE

The light response to 5.9 keV photons from a ⁵⁵Fe was

We used a Hough transform to individuate spots and measure their dimension and light yield.

X-RAYS FROM A 55FE SOURCE

Light yield distributions were studied as a function of the mixture and the to the GEM voltages

By means of an X-ray tube, electron yield was also studied for different mixtures and different voltages.

Primary light

Still very far from being useful for full fiducialization (but we used sub-optimal Q.E. PMT and limited light collection) It might be good to reject cathode events anyway!

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R&D NITEC

NITEC: Negative Ion Time Expansion Chamber

3 x 3 x 5 cm³ 0.045 Liters Triple thin GEMs Timepix pixel charge readout

- Small gas volume to study gas properties
- Operation in pure SF₆ at low pressure
- Operation with gas mixture
- GEM pixelated charge readout

SF₆ negative ion drift

 Measured ion mobility in SF₆ based gas mixture at 600 torr. (from drift velocity of negative ions generated by the 45 MeV BTF electron beam)

- Need to understand the "light" gain of gas mixture with SF₆
- Hope to see signal of minority carriers (SF₅, SF₄, etc.)

E.Baracchini et al.JINST 13 (2018) no.04, P04022

COMBINED LIGHT READOUT

JINST 13 (2018) no.05, P05001

Sensitive gap parallel to the beam Sensitive gap tilted w.r.t. the beam

1 cm in 140 ns => drift velocity 7.2 cm/ μ s in agreement with Garfield expectation of 7.3 cm/ μ s.

PMT+CMOS COMBINED READOUT

Single cluster 3D position reconstruction can be obtained by comparing the light profile along the track (X, Y) and the PMT waveform (t);

A peak finding algorithm was used to highlight the main cluster signals;

By means of the measured drift velocity, Z coordinate was evaluated;

Residual distribution to a 3D fit allows to compute a resolution on Z of 100 μ m.

Near-term results for MeV Dark Matter expected with 2e- thresholds in Si (DAMIC/ SENSEI/SuperCDMS)

→Significant overlap in sensitivity to follow up with Directional Detection (PTOLEMY-G³)

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Carbon nanotubes

Scattering on a carbon nanotube

- CNT are "empty"
 - no electrons along the carbon ion path
- Large aspect ratio:
 ~10 nm diameter vs.
 ~100 µm height
- "target" mass on the CNT surface

CNT as a potential well

- Transverse energy is conserved
- 6+C ion scattered off \mathcal{E}_{\perp} the CNT are channeled in the CNT
- Little effect of electrons on CNT surface

Aligned and oriented CNT "brush"

- Recoiling C ions are emerging from CNTs with different rates depending on CNTs orientation.
- When C ions are not channeled they are absorbed within the brush
- Effect of rechanneling or inter-CNT trapping
 NOT included HERE

Channeling of an ion

Demonstrate ~10-100 KeV C ions are trapped. Trapping has a larger effective θ_c ~ 35 deg

Functionalization

Figure 8. Different approaches to chemical modification of carbon nanotubes. (a) substitutional doped single-walled nanotubes (either during synthesis or by post-growth ion-implantation), (b,c) nanotube bundles intercalated with atoms or ions, (d,e) peapods: SWNTs filled with fullerenes (other endohedral fillings are possible), (f) fluorinated tubes, (g) covalently functionalised tubes and (h) functionalised nanotubes via *n*-stacking of the functionality and the tubes.

- CNT can be very efficiently **doped**
- Alkali metal can be bonded to CNT surface (Na,Cs,...) or F.
- WIMP can scatter on Na, Cs, ... and these ions can then be channeled

Even with large spread of ionization, the range measurement might help to identify the signal

Energy [KeV]

Carbon ion (10 KeV)

in 100 mbar Ar

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10²

Range / Straggling [µm]

How to pack enough mass ?

SEM image of NanoLab aligned CNT on mm Si substrate

First test shows <10⁻³ contamination of Fe ("seed") and O.

- Simple-man calculation for 1 m² layer - 100 layers detector
- ρ =5 nm and h =300 μ m
- number CNT per layer (single wall)
 ~ 2 m²/a²
- Surface density of a graphene layer: 1/1315 g/m²

$a \; [\mathrm{nm}]$	CNT detector mass [kg]
11	11.8
30	1.6
45	0.7
58	0.4

First look at CNT

Color is drift time (along Z) for different Z position (X-Y projection)

CNT are good conductor, modifying the field cage electric field. Different field configuration to be tried for the conceptual demonstration