

# Deep Learning Continuous Gravitational Waves

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In Collaboration with:

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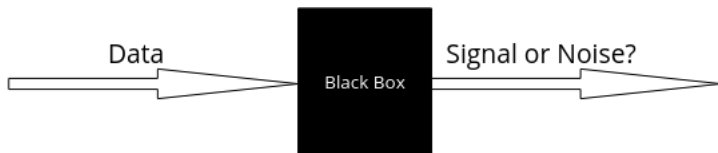
<sup>3</sup>University of Glasgow, UK

# Continuous Gravitational Waves

- ▶ Generated by spinning non-axisymmetric compact objects.
  - ▶ Eg. Neutron Stars
- ▶ Weak signals. Buried under noise.
  - ▶  $h_0 \lesssim 10^{-24}$
  - ▶  $\sqrt{S_n} \gtrsim 10^{-23} \text{Hz}^{-\frac{1}{2}}$
  - ▶  $\text{depth} \equiv \frac{\sqrt{S_n}}{h_0} \gtrsim 10 \text{Hz}^{-\frac{1}{2}}$
- ▶ Matched filtering used for the search.
  - ▶ Fully Coherent: Computationally not feasible.
  - ▶ Empirical semi-coherent methods used.

# Aim

- ▶ Find if the data contains a signal or just noise.

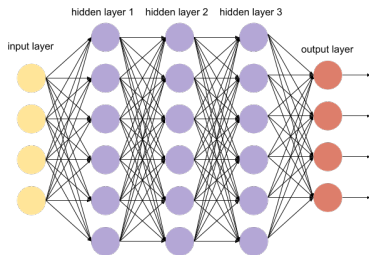
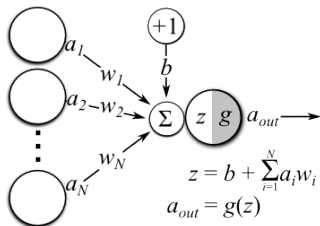


# Matched Filtering Approach

- ▶ Generate templates of signals based on the search parameters.
- ▶ Match the templates with the signal and calculate a statistic.
  - ▶ for eg  $\mathcal{F}$ -Statistic
- ▶ Fix a threshold based on false alarm rate
  - ▶ for eg: only 1% of noise have the statistic  $>$  threshold
- ▶ Data with statistic  $>$  threshold considered a signal candidate.

# Deep Neural Networks

- ▶ Algorithms which learn from examples
- ▶ Work on raw input data
- ▶ Already successful in CBC searches
  - ▶ George & Huerta, PRD(2018), Phys.Lett.B(2018)
  - ▶ Gabbard et al, PRL(2018)
- ▶ We use convolutional neural networks.
  - ▶ Using Keras + Tensorflow as our framework.



# Our approach

- ▶ Generate training examples containing noise and signals.
  - ▶ In the frequency domain.
- ▶ Train the neural network to generate a statistic.
  - ▶ Higher value for signals, lower value for noise.
- ▶ Fix the threshold based on a fixed false alarm rate.
- ▶ Compare the performance with (coherent) matched filtering.

## Organizing the data

- ▶ Dataset organized into “cases” depending on the parameters
  - ▶ frequency ranges, spin-down, duration of observation etc.
- ▶ Fix the strength ( $h_0$ ) of the signal such that.
  - ▶ Matched filtering gets detection probability of 90% at 1% false alarm rate.
  - ▶  $\text{depth} = \frac{\sqrt{S_n}}{h_0}$

$f$ (Hz)	$\dot{f}$ (Hz/s)	$T_{obs}$ (s)	depth ( $\text{Hz}^{-\frac{1}{2}}$ )
[20,20.005]	$[-3 \times 10^{-9}, 0]$	$1 \times 10^5$	10.53
[200,20.0025]	$[-1 \times 10^{-10}, 0]$	$1 \times 10^5$	10.3
[20,20.001]	$[-1 \times 10^{-10}, 0]$	$1 \times 10^6$	30.4
..	..	..	..

# Training

- ▶ Generate and store signals in a file using LALSuite<sup>1</sup>
- ▶ Generate and add Gaussian noise on the fly.
- ▶ Use noise and noise+signal examples to train the network.
- ▶ Check the performance using an independent validation set.
  - ▶ not seen by the network before.

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<sup>1</sup><https://wiki.ligo.org/DASWG/LALSuite>



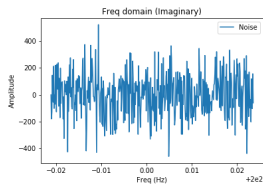
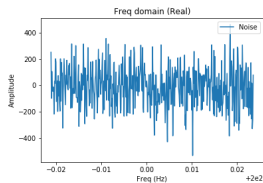
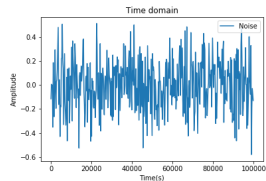
# What does the input data look like?

$$f = [200, 200.0025] \text{ Hz}$$

$$\dot{f} = [-1 \times 10^{-10}, 0] \text{ Hz/s}$$

$$T_{\text{span}} = 1 \times 10^5 \text{ s}$$

$$\text{depth} = 10.3 \text{ Hz} \frac{-1}{2}$$



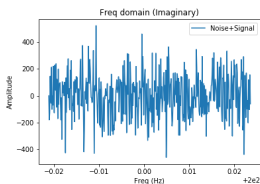
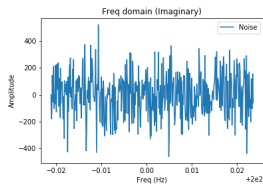
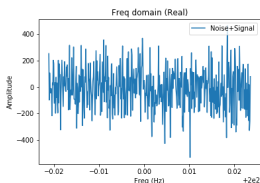
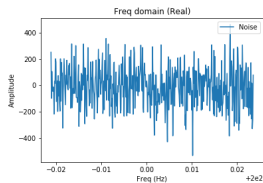
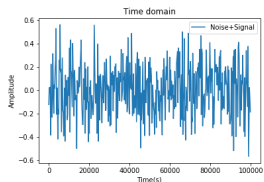
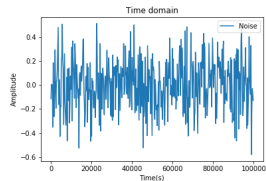
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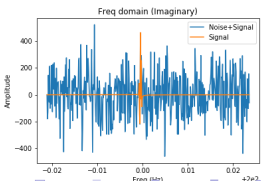
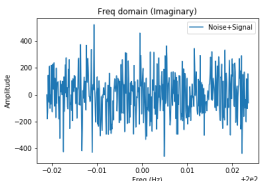
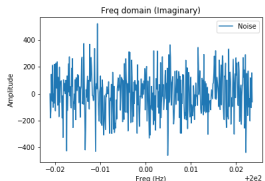
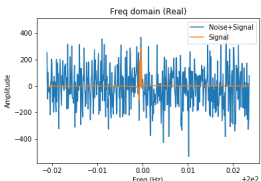
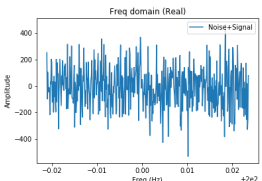
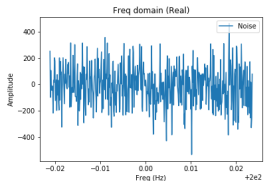
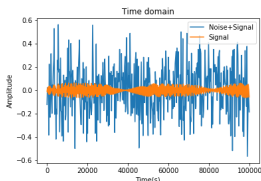
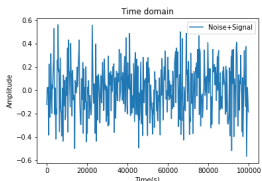
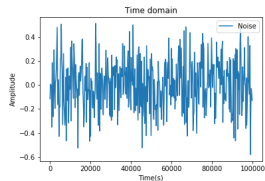
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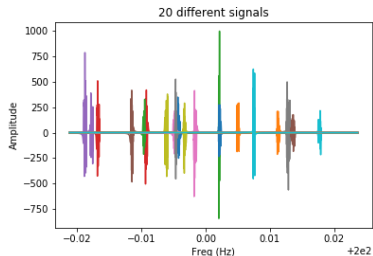


# Frequency Band

- ▶ Sky position affects the doppler shifts.
- ▶ Signals cover wide frequency range.

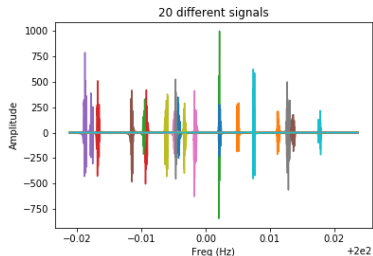
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- ▶ Sky position affects the doppler shifts.
- ▶ Signals cover wide frequency range.
- ▶ Actual signal width quite small.
- ▶ Computationally more expensive.



# Solution

- ▶ freq range kept as twice the max width of a signal.
- ▶ Smaller input. Faster training.
- ▶ Slide over the whole range in the actual search.
  - ▶ Maximum value over the slides taken as the statistic.

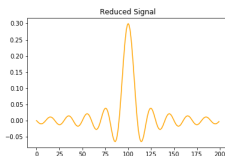
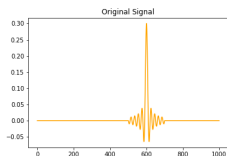
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- ▶ Example
  - ▶ max width = 200, total width = 1000



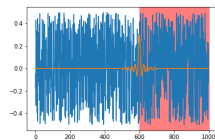
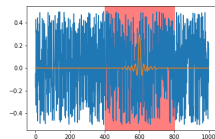
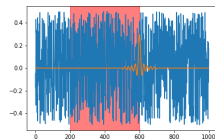
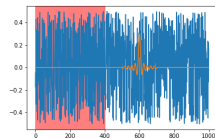
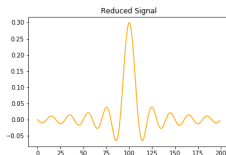
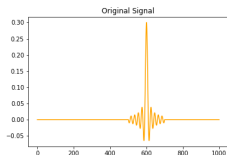
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# Results

▶  $f=[20,20.005]\text{Hz}$

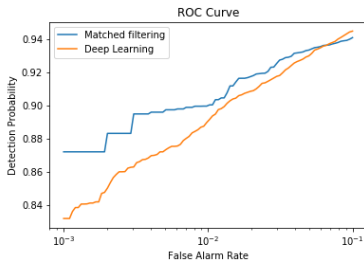
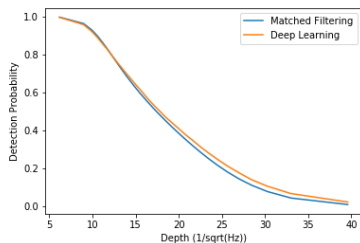
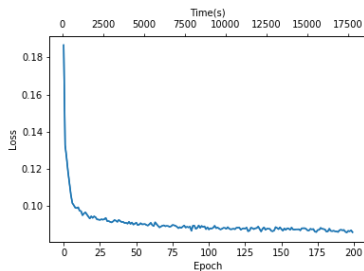
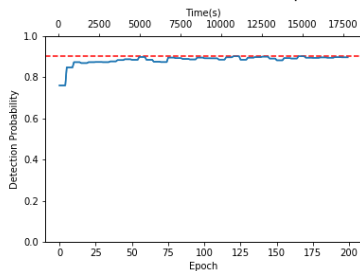
▶  $T_{\text{Span}}=1 \times 10^5\text{s}$

▶ Matched Filtering Templates =  $2.4 \times 10^6$

▶  $\dot{f}=[-3 \times 10^{-9}, 0]\text{ Hz/s}$

▶  $\text{depth}=10.57\text{ Hz}^{-\frac{1}{2}}$

▶ Matched Filtering Time = 0.52s



# Results

▶  $f=[20,20.001]\text{Hz}$

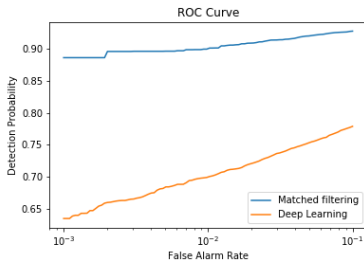
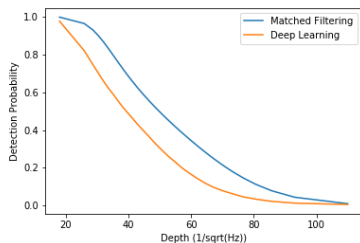
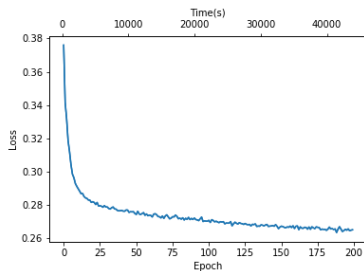
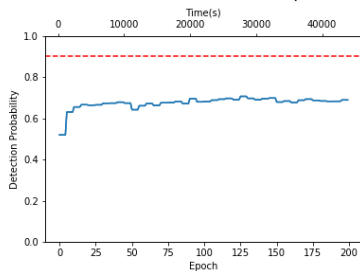
▶  $T_{\text{Span}}=1 \times 10^6\text{s}$

▶ Matched Filtering Templates =  $6.3 \times 10^7$

▶  $\dot{f}=[-1 \times 10^{-10}, 0]\text{ Hz/s}$

▶  $\text{depth}=30.4\text{ Hz}^{-\frac{1}{2}}$

▶ Matched Filtering Time = 71s



# Results

▶  $f=[100,100.001]\text{Hz}$

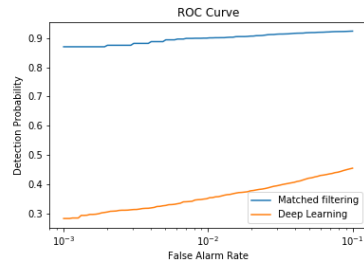
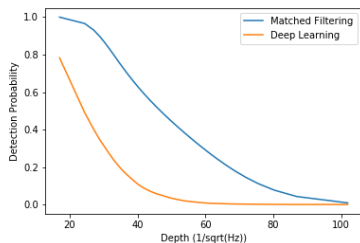
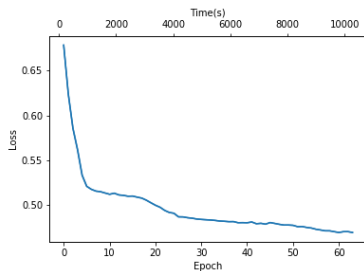
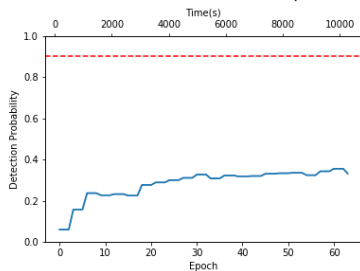
▶  $T_{\text{Span}}=1 \times 10^6\text{s}$

▶ Matched Filtering Templates =  $1 \times 10^9$

▶  $\dot{f}=[-1 \times 10^{-10}, 0]\text{ Hz/s}$

▶  $\text{depth}=28.7\text{ Hz}^{-\frac{1}{2}}$

▶ Matched Filtering Time =  $1.2 \times 10^3\text{s}$



# Future Work

- ▶ Find optimal network architecture for the cases.
- ▶ Better generalization over the frequencies.
  - ▶ Currently fixed narrow frequency range.
  - ▶ Generalizing over wider frequency bands.
- ▶ Parameter Estimation.

Thank You







# Appendix

