



Status of the frequency dependent squeezing experiment at TAMA

K. Arai, N. Aritomi, Y. Aso, M. Barsuglia, **E. Capocasa**, M. Eisenmann, R. Flaminio, Y. Guo, R. Lee, M. Leonardi, H. Lück, L. Pinard, E. Polini, P. Prat, R. Schnabel, E. Schreiber, K. Somiya, M. Tacca, R. Takahashi, D. Tatsumi, A. Tomura, H. Valbruch, M. Vardaro, C.Wu, Y. Zhao

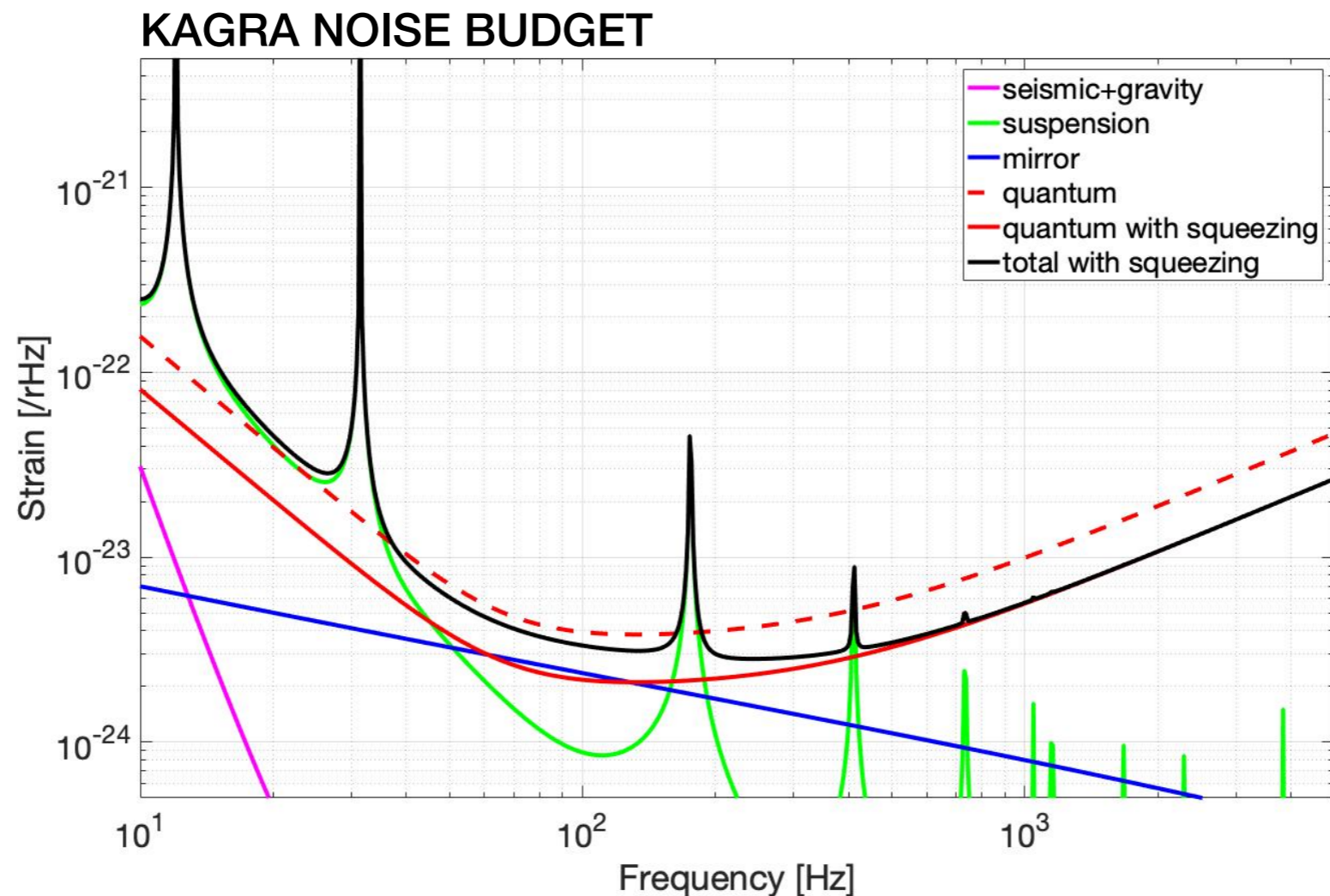
Supported by the JSPS Grant-in-Aid for Scientific Research (Grant code 15H02095 & 18H01235)



GWADW2019 - Gravitational-Wave Advanced
Detector Workshop - From Advanced
Interferometers to Third Generation Observatories

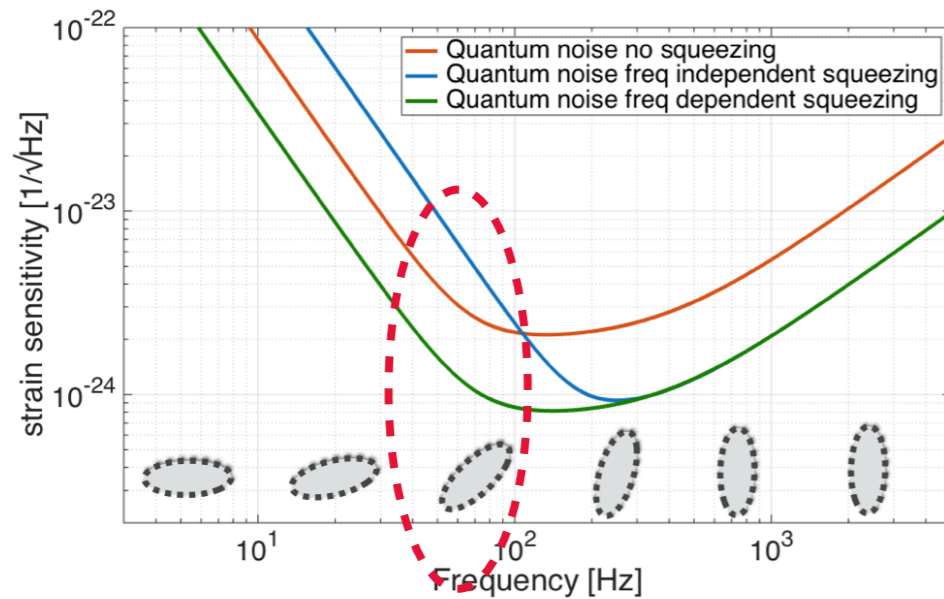
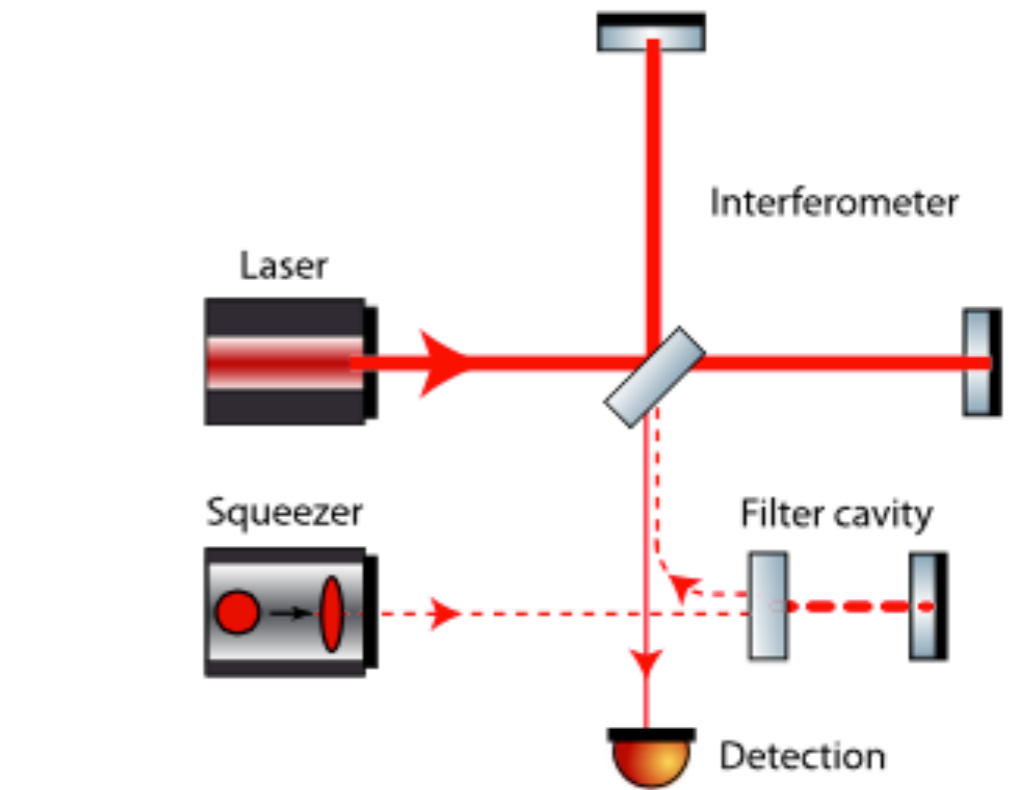
Motivation and context

- Frequency dependent squeezing goal: **broadband quantum noise reduction**
- 300 m filter cavities planned for A+ and AdVirgo+ in O4
- Considered among the upgrades of KAGRA+

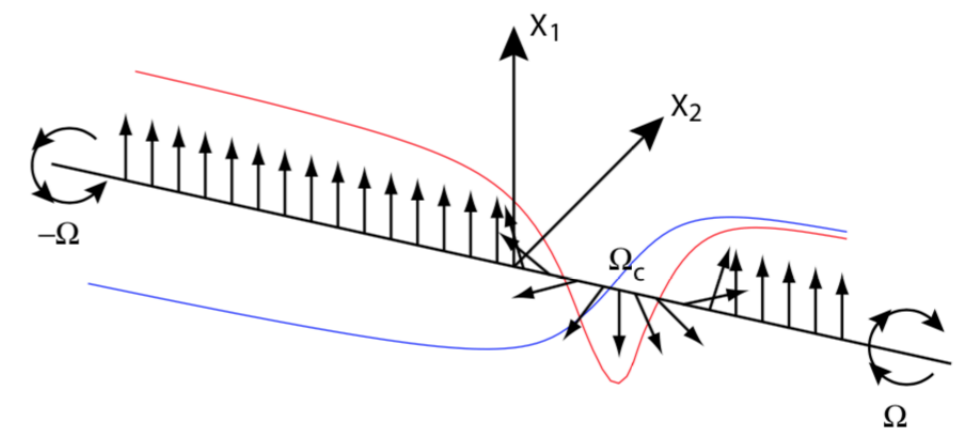


How to produce frequency dependent squeezing?

- Reflect frequency independent squeezing off a detuned cavity
- Rotation frequency depends on the cavity line-width
- Performances strongly affected by round trip loss per unity length

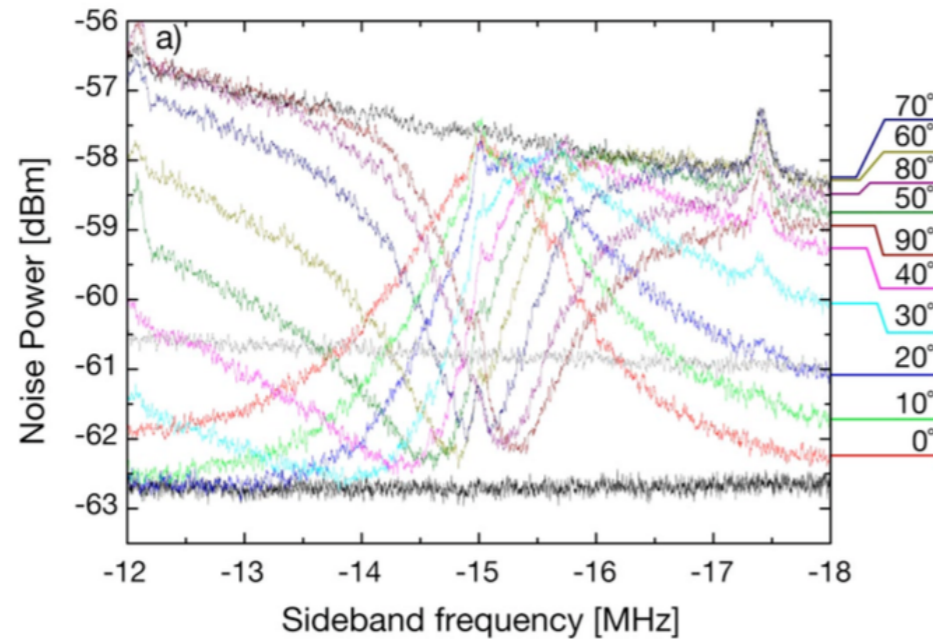


Optimal rotation frequency
between 40 and 70 Hz



Squeezing angle rotation already demonstrated

@ MHz frequency

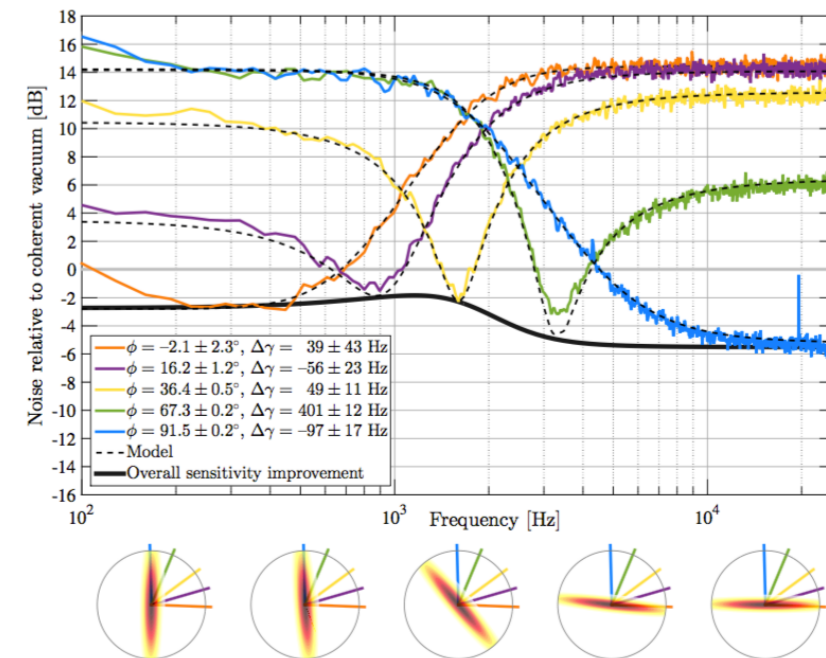


PHYSICAL REVIEW A 71, 013806 (2005)

Experimental characterization of frequency-dependent squeezed light

Simon Chelkowski, Henning Vahlbruch, Boris Hage, Alexander Franzen, Nico Lastzka, Karsten Danzmann, and Roman Schnabel

@ kHz frequency



PRL 116, 041102 (2016)

PHYSICAL REVIEW LETTERS

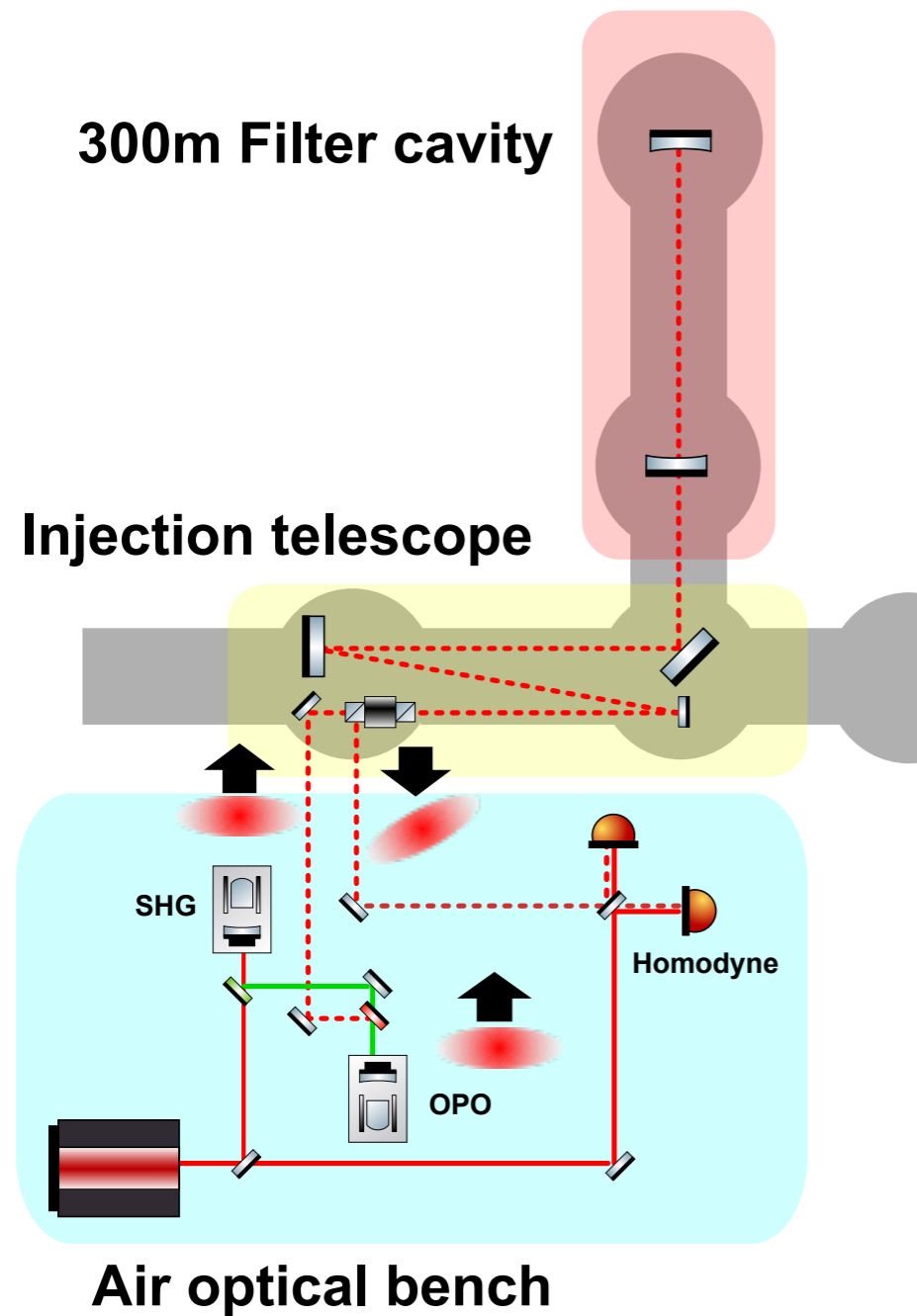
week ending
29 JANUARY 2016

Audio-Band Frequency-Dependent Squeezing for Gravitational-Wave Detectors

Eric Oelker, Tomoki Isogai, John Miller, Maggie Tse, Lisa Barsotti, Nergis Mavalvala, and Matthew Evans*
Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA
(Received 20 August 2015; revised manuscript received 10 December 2015; published 29 January 2016)

Our goal: **full scale filter cavity** prototype to demonstrate frequency dependent squeezing with **rotation around 70 Hz**

Experiment overview



FILTER CAVITY

- Installed @ TAMA facility at NAOJ
- Finesse 4400
- Round trip losses 45-85 ppm

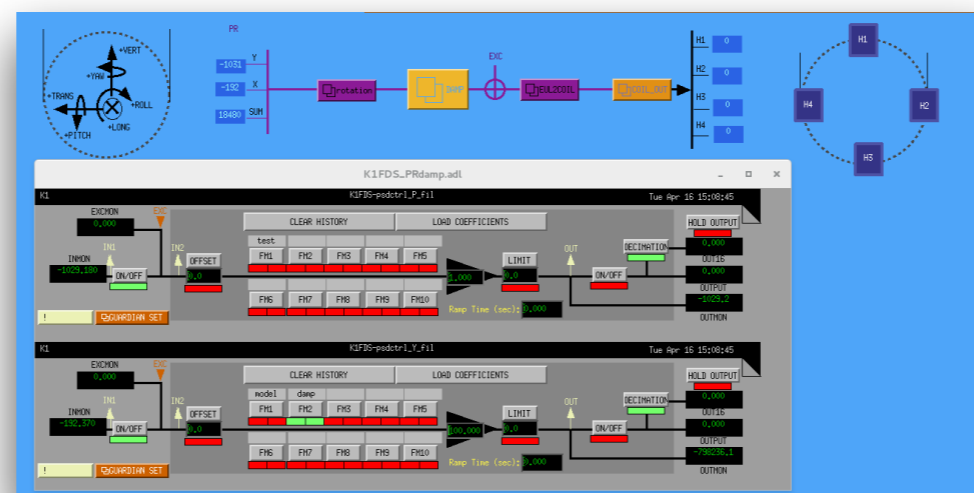
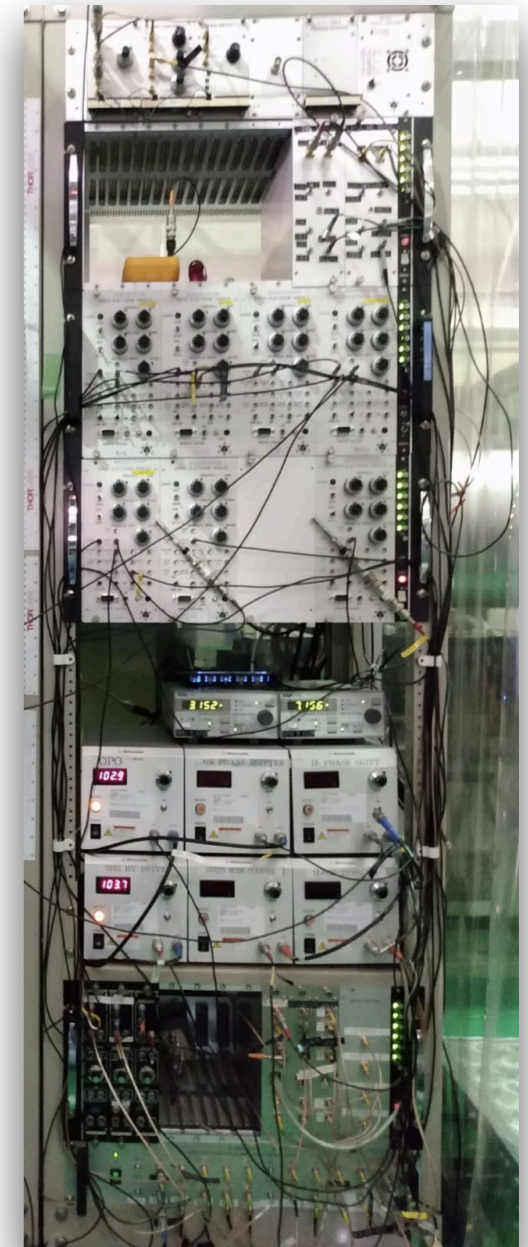
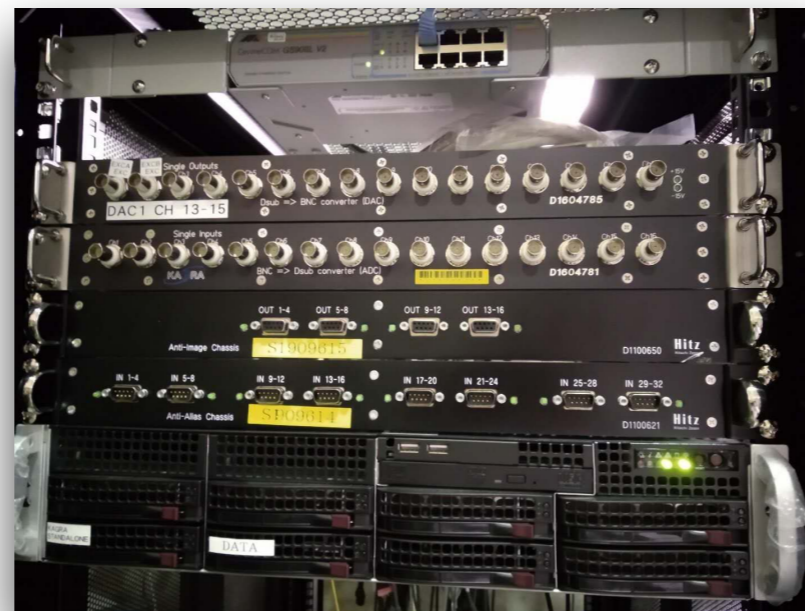
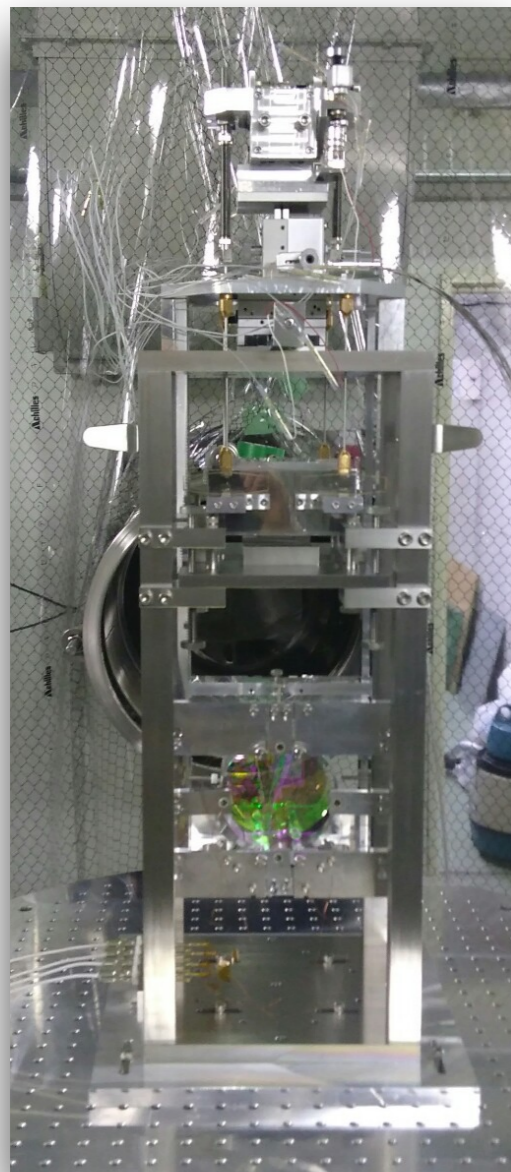
FREQ. IND. SQUEEZED VACUUM SOURCE

- Based on AEI design
- Target 9dB from 10 Hz

More details on Yuhang Zhao's poster

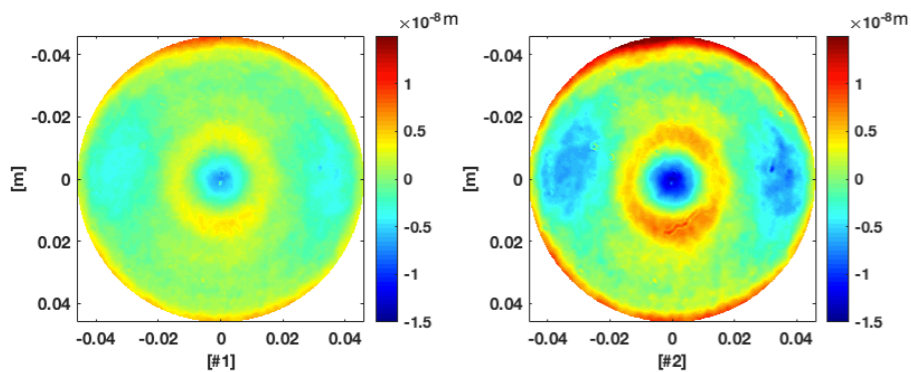
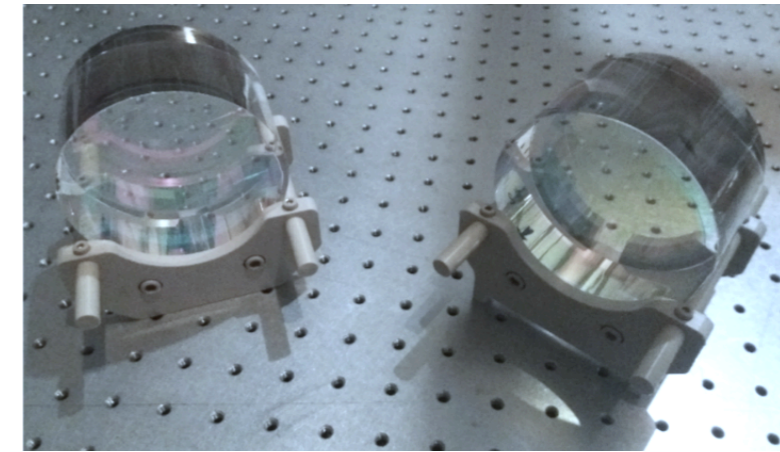
Filter cavity: suspensions and digital control system

- Double pendulum suspensions (so-called KAGRA Type C)
- Digital system for local control recently changed from old TAMA LabVIEW to KAGRA DGS



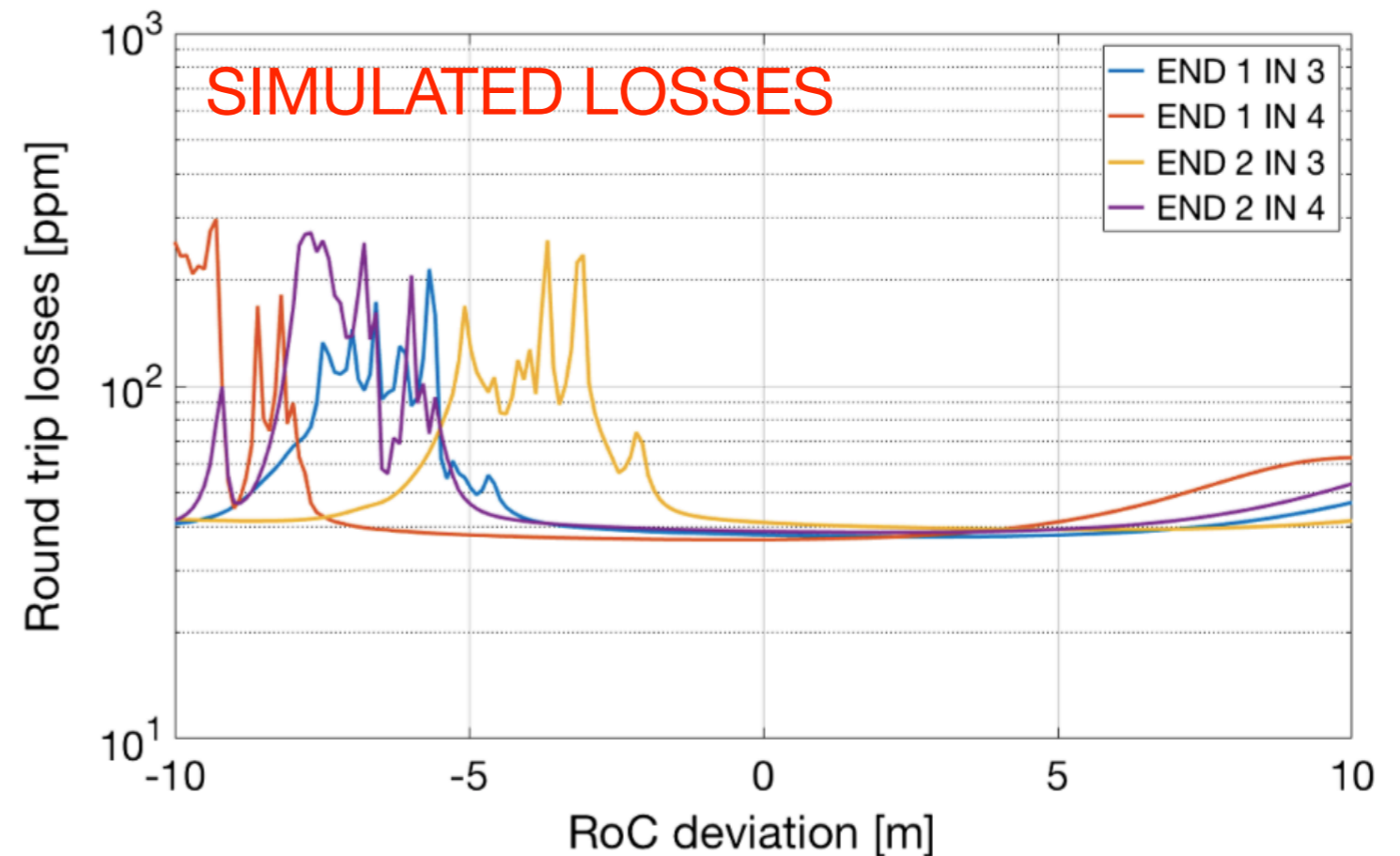
Filter cavity: the mirrors

- Tama size: 10 cm diameter, 6 cm thickness
- Beam radius: ~ 1 cm
- Requirement on surface quality set to have **80 ppm of round trip losses**



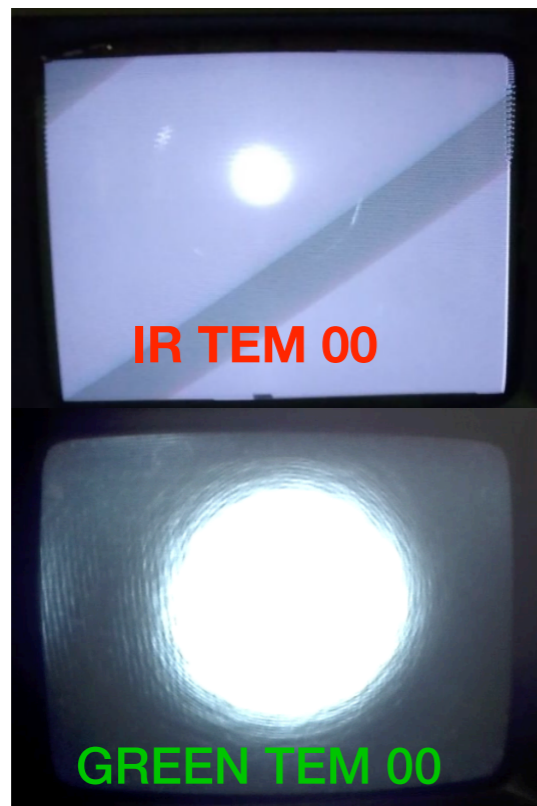
Mirror	diameter 0.05 m		diameter 0.02 m	
	RMS (nm)	PV (nm)	RMS (nm)	PV (nm)
#1	1.96	11.5	0.52	3.28
#4	1.94	14.8	0.48	3.28

Substrates coated and characterized at LMA

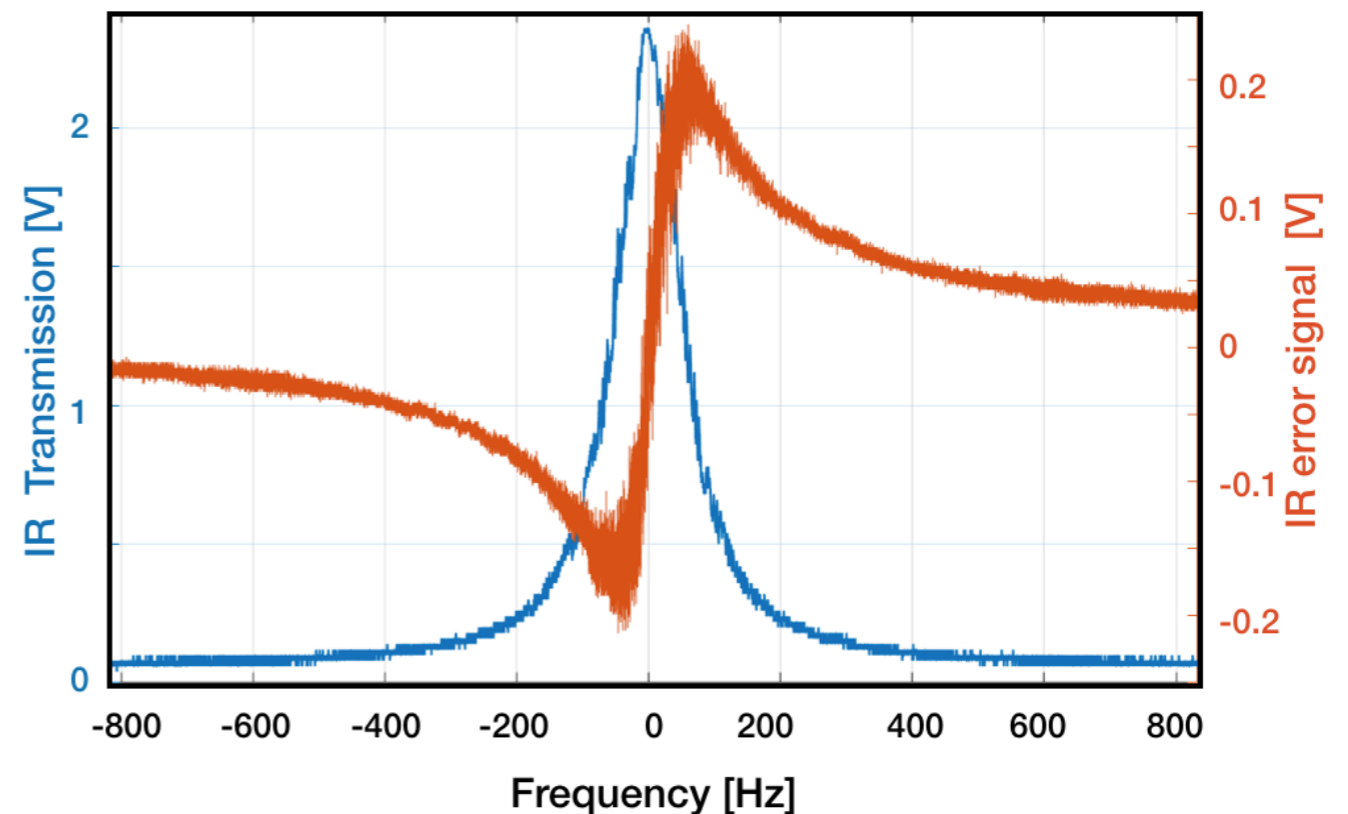


Filter cavity: the control

- Laser frequency locked on the cavity length
- Auxiliary green beam used (finesse ~ 170)
- IR detuning controlled with AOM on green path
- Alternative control strategy with coherent control sidebands is being studied (see Naoki Aritomi's poster)

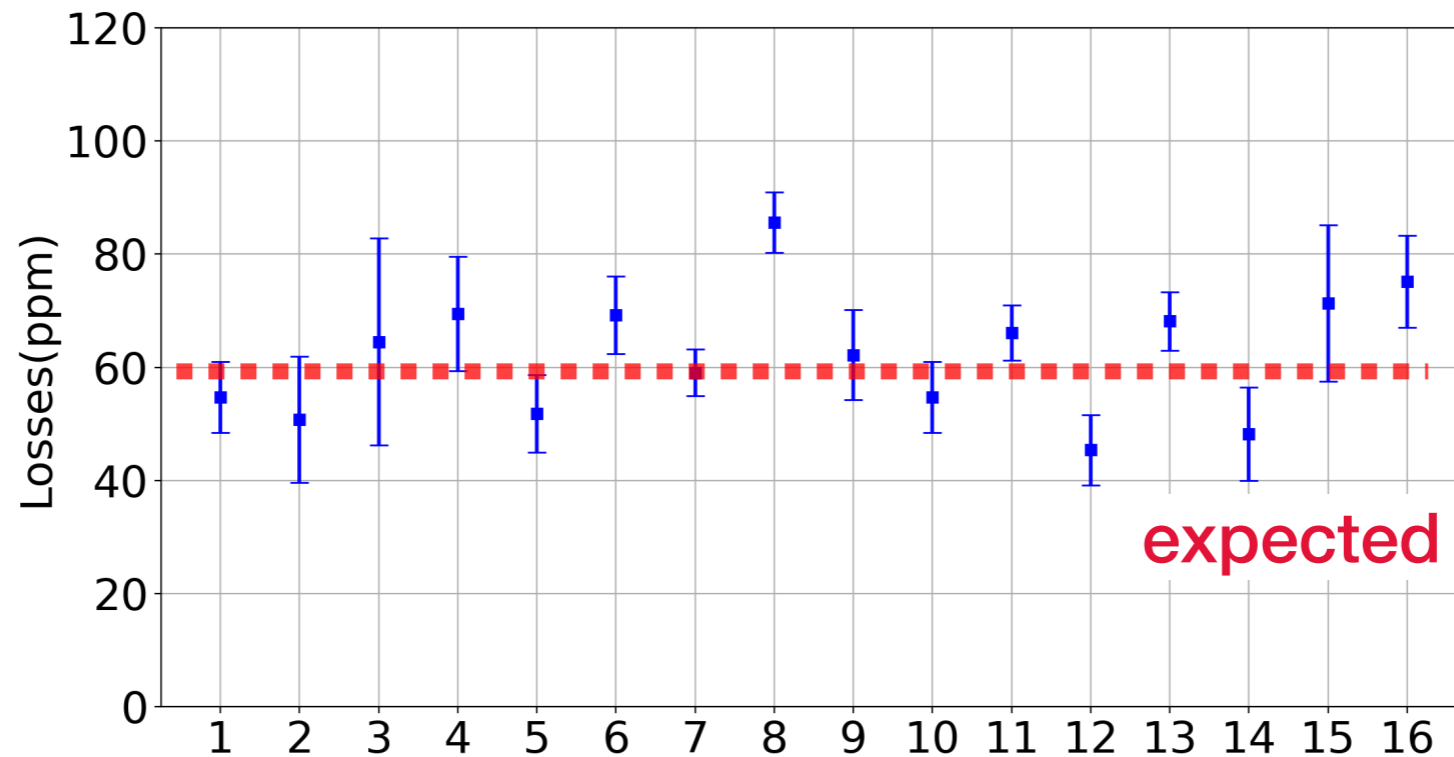


IR resonance crossing by driving AOM

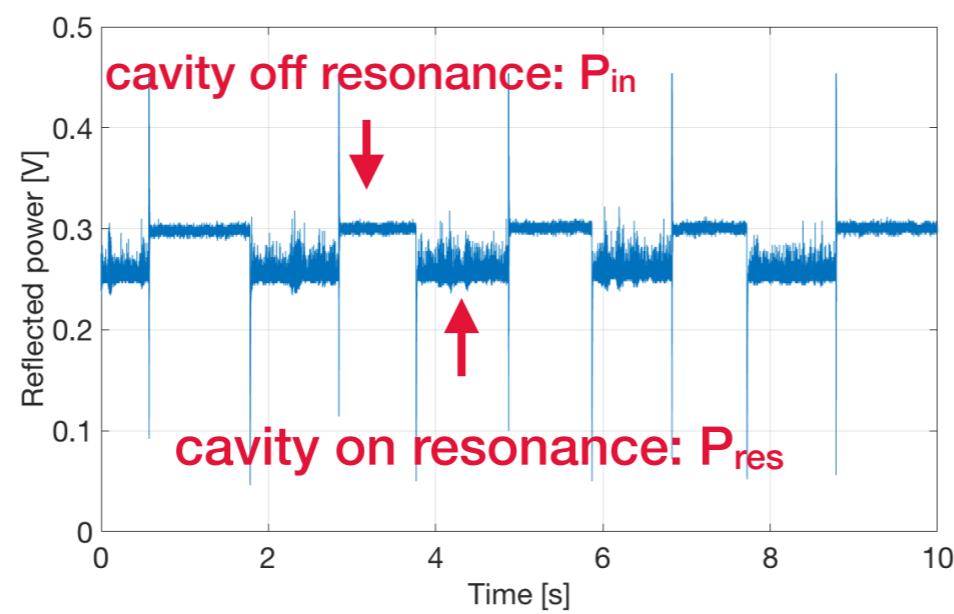


Filter cavity: round trip loss characterization

- Round trip losses between 45 and 85 ppm



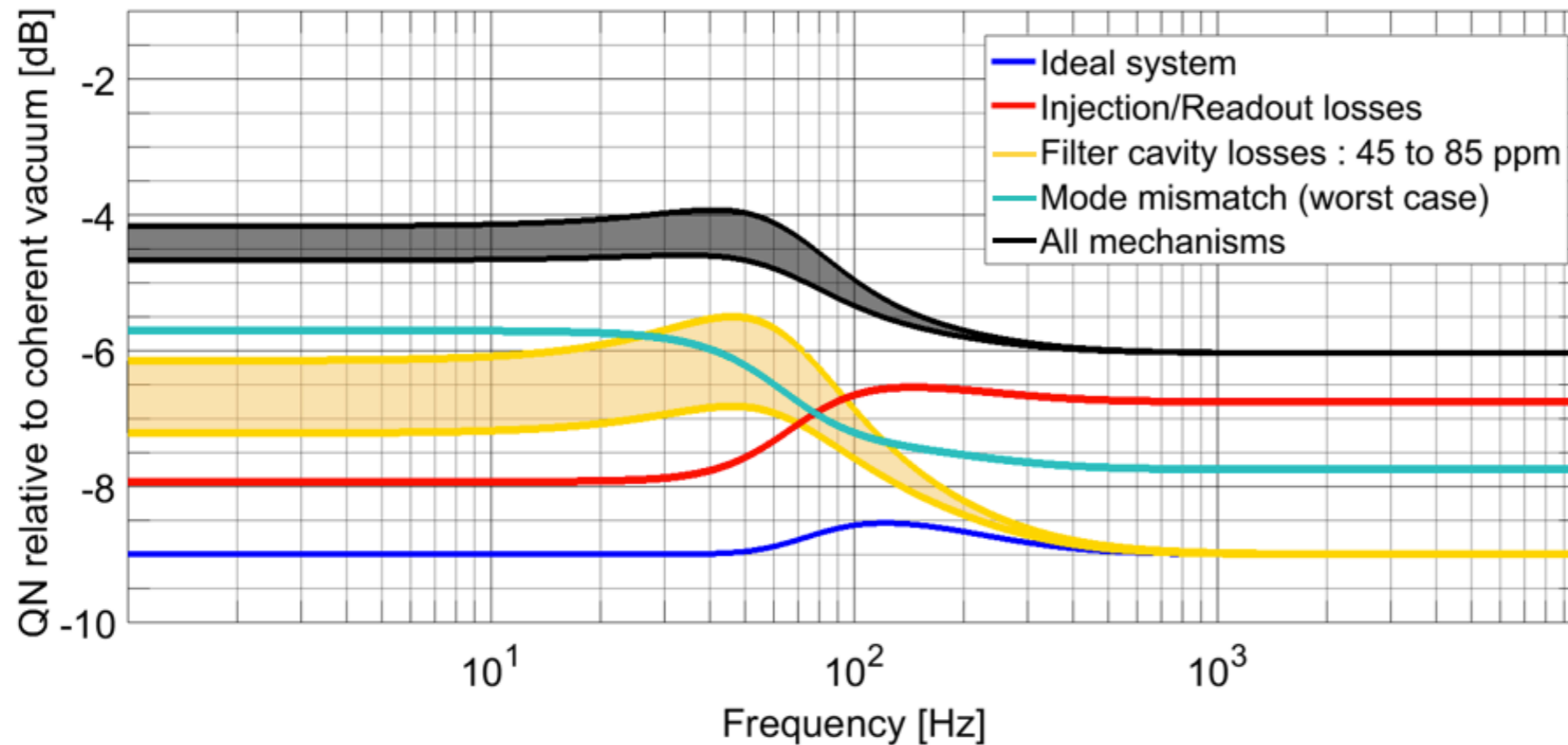
- Round trip losses extracted from reflectivity measurements



$$L \sim \frac{T_1}{2} \frac{1 - P_{\text{res}}/P_{\text{in}}}{1 + P_{\text{res}}/P_{\text{in}}}$$

Expected squeezing level

- Round trip losses below 85 ppm allows for ~ 4 dB of squeezing at low frequency (assuming 9 dB of injected squeezing)



injection losses	5%
readout losses	5%
squeezer-filter cavity mismatch	2%
squeezer-local oscillator mismatch	5%
δL (rms)	0.3 pm

PHYSICAL REVIEW D **98**, 022010 (2018)

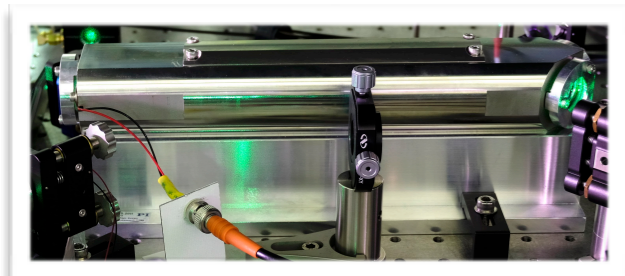
Measurement of optical losses in a high-finesse 300 m filter cavity for broadband quantum noise reduction in gravitational-wave detectors

Eleonora Capocasa,^{1,2,*} Yuefan Guo,³ Marc Eisenmann,⁴ Yuhang Zhao,^{1,5} Akihiro Tomura,⁶ Koji Arai,⁷ Yoichi Aso,¹ Manuel Marchiò,¹ Laurent Pinard,⁸ Pierre Prat,² Kentaro Somiya,⁹ Roman Schnabel,¹⁰ Matteo Tacca,¹¹ Ryutaro Takahashi,¹ Daisuke Tatsumi,¹ Matteo Leonardi,¹ Matteo Barsuglia,² and Raffaele Flaminio^{4,1}

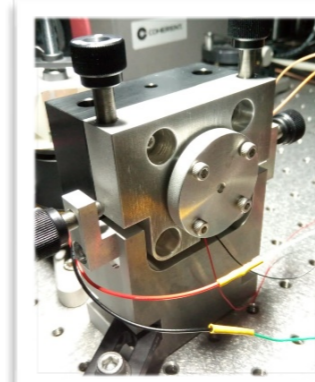
(Received 27 May 2018; published 31 July 2018)

Squeezed vacuum source: integration completed

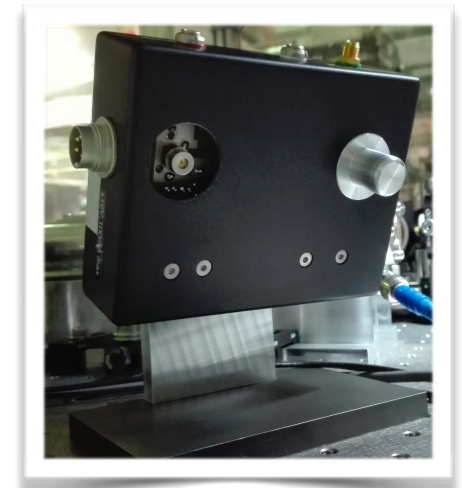
Green mode cleaner



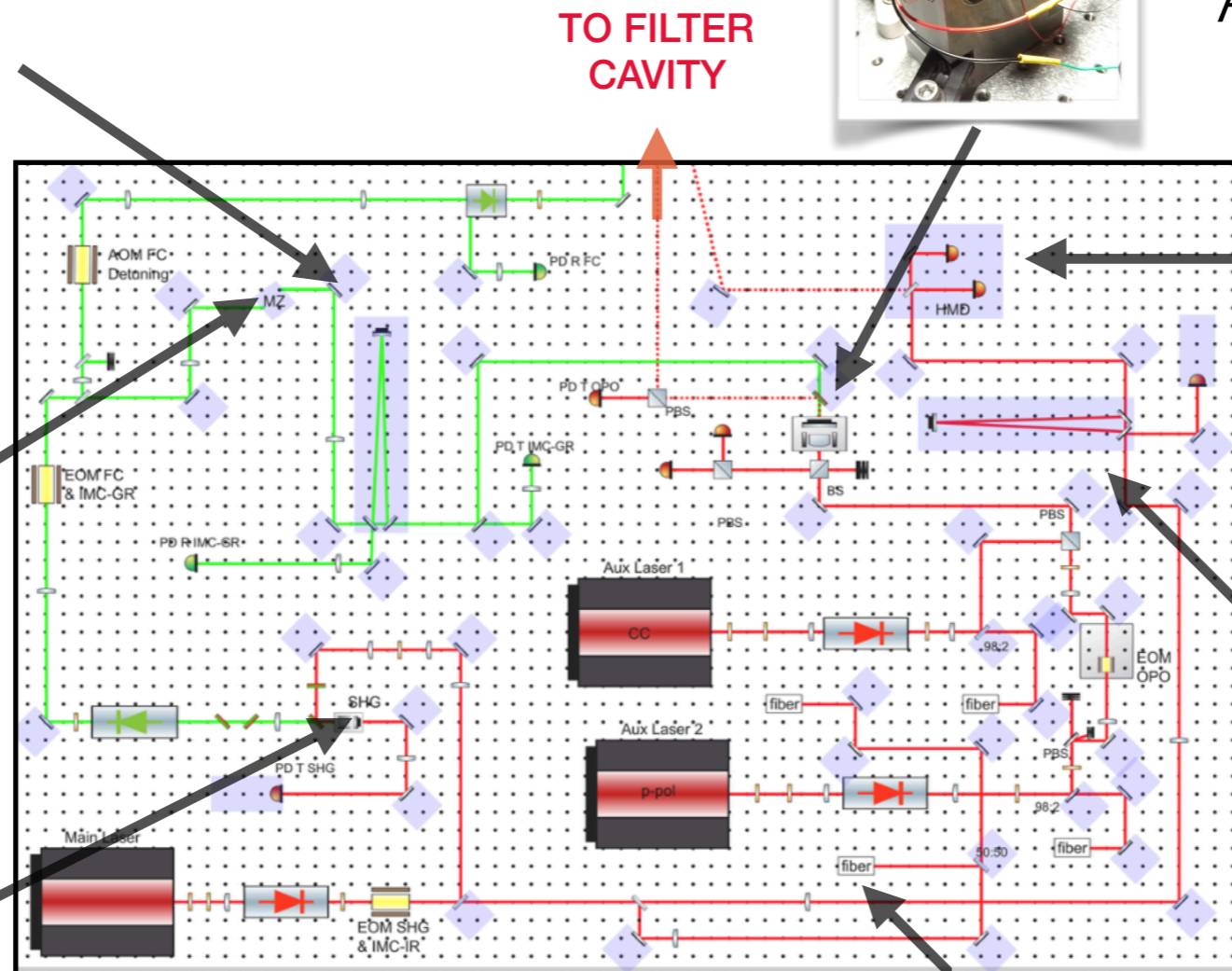
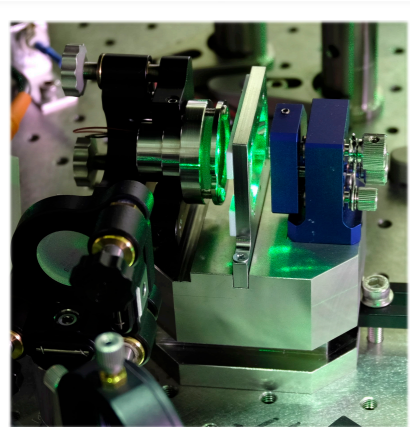
OPO to produce squeezed vacuum



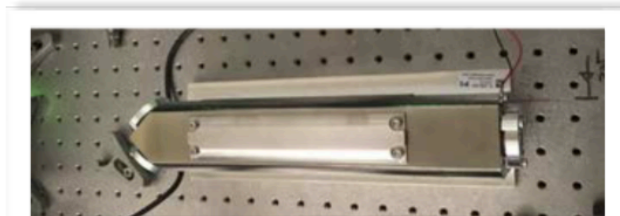
Homodyne detector (in collaboration with AEI, Hannover)



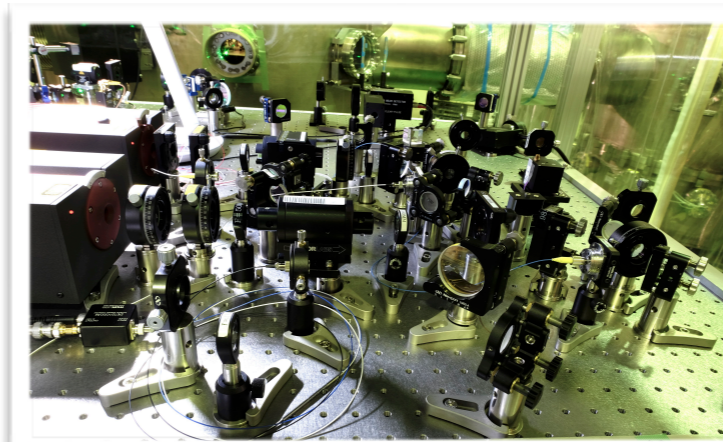
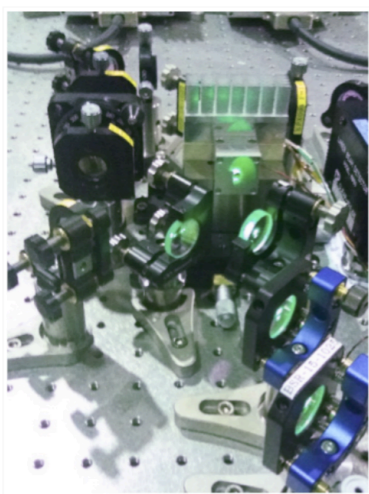
Mach-Zehnder to stabilize and tune the green pump



IR mode cleaner



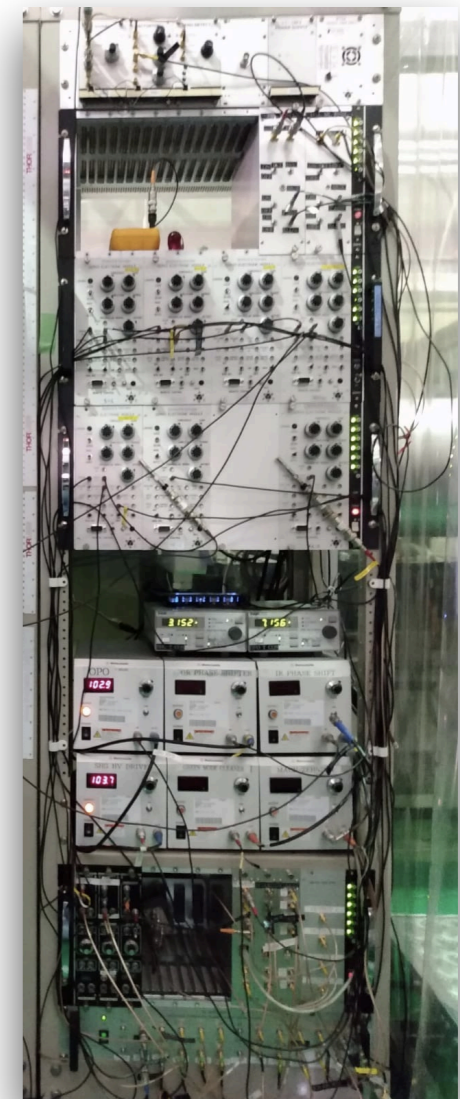
Second Harmonic Generator to produce green pump



PLL between main laser and auxiliary lasers

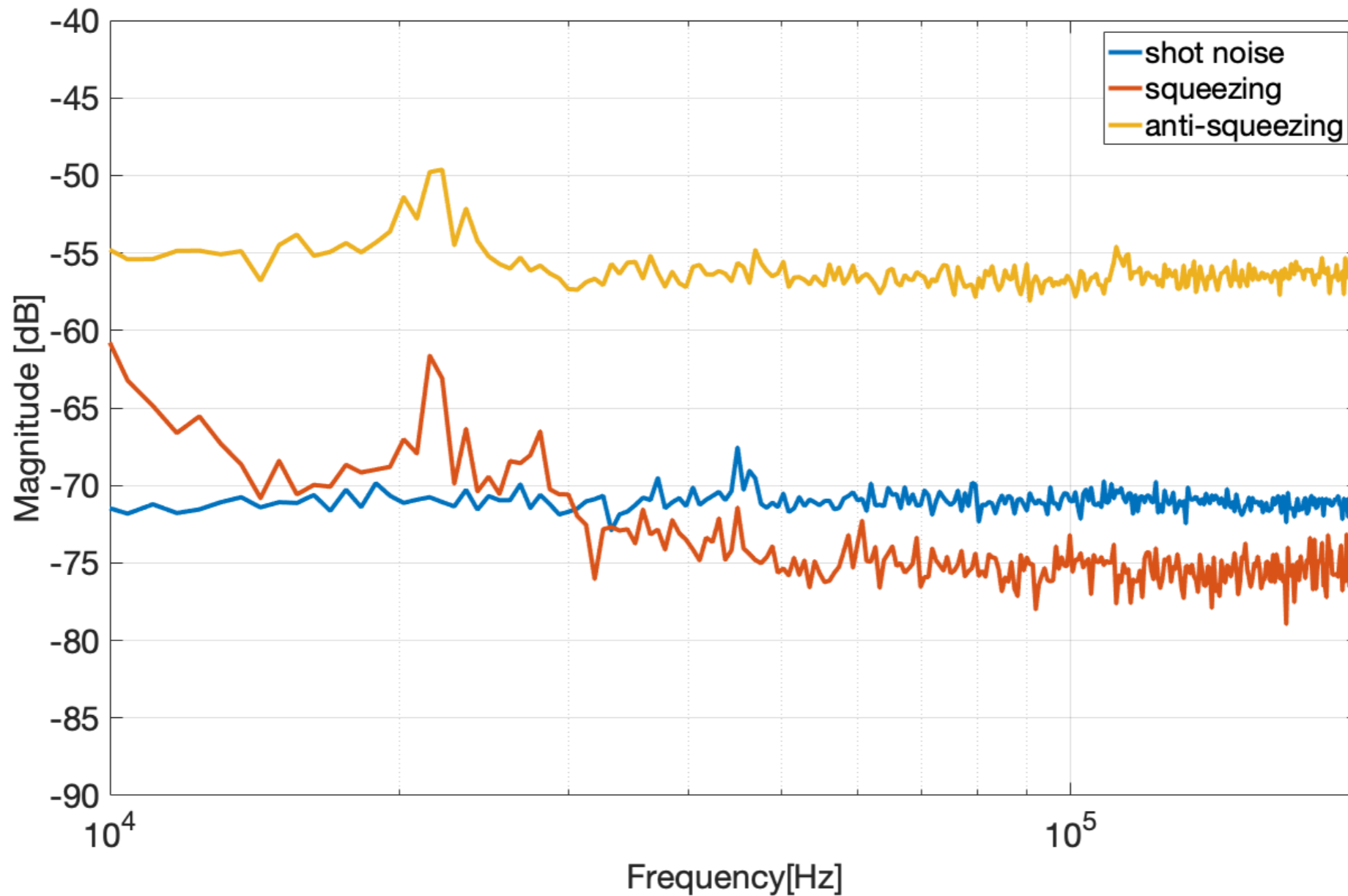
Squeezed vacuum source: analog control servos

- Developed in collaboration with APC
- 7 servos: 4 optical cavity, 1 Mach-Zehnder, 2 Coherent control loops
- Hierarchical auto-lock implemented
- Parameters controlled on the front panel or remotely



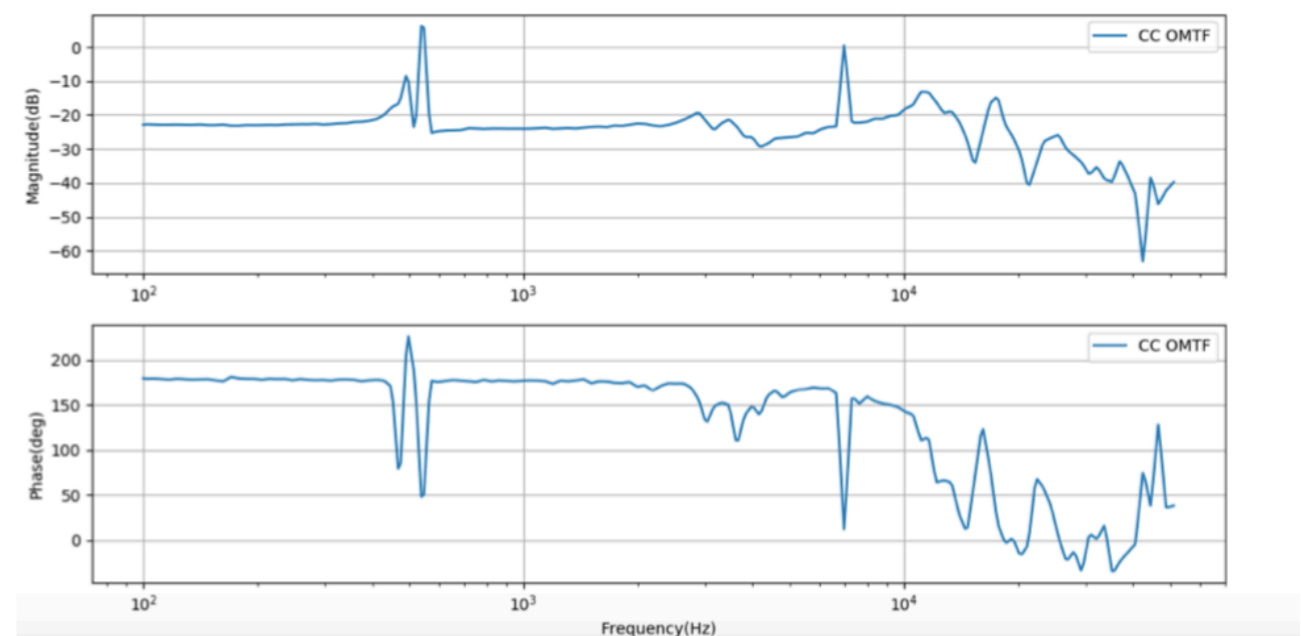
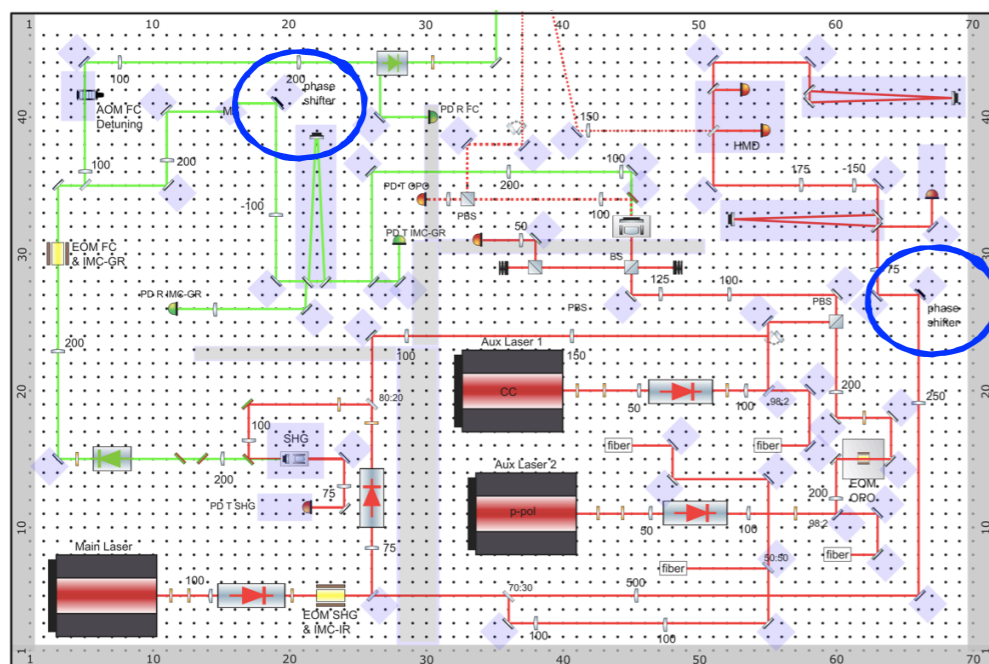
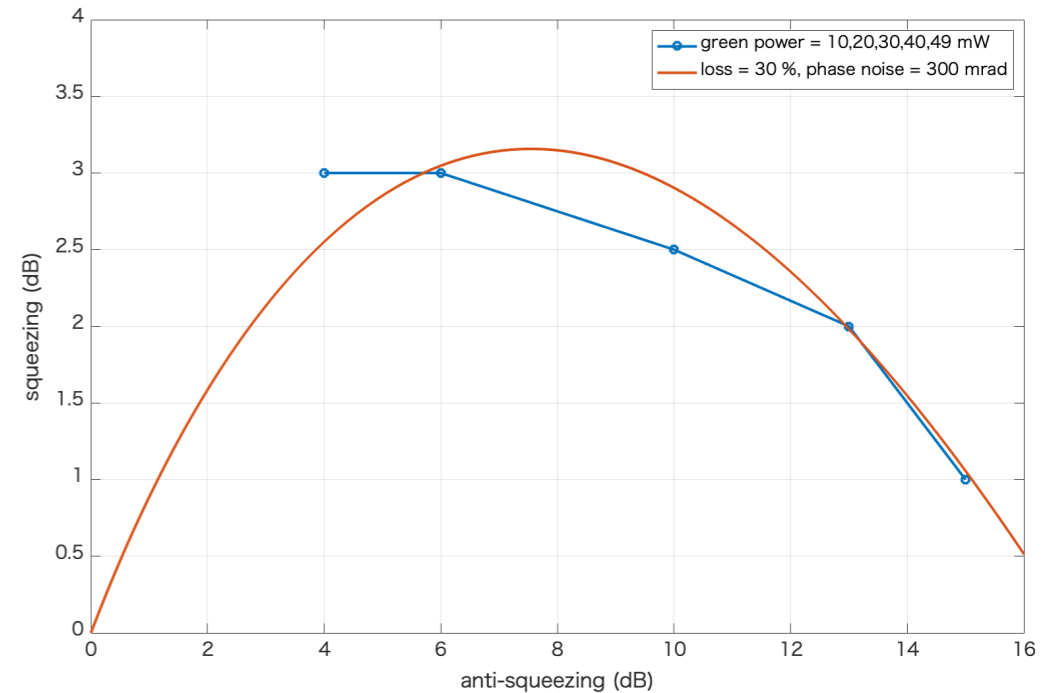
Current squeezing performance

- About 4.5 dB of squeezing (and 15 dB of anti-squeezing) down to ~30 kHz



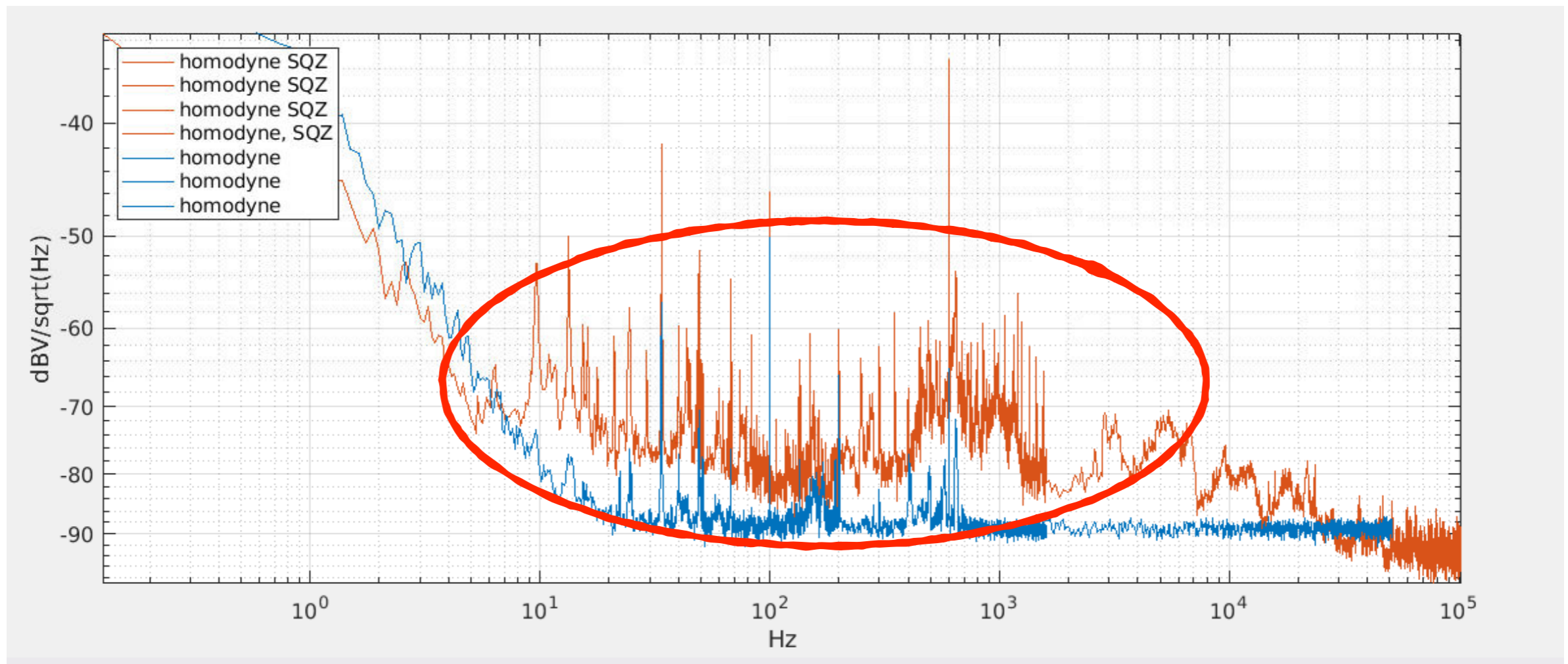
Phase noise reduction

- Since measured squeezing is lower when we increase the pump power, we believe that it is **limited by phase noise**
- Coherent control optimization is our first priority
- Coherent control loop bandwidth is limited to 80 Hz because of mechanical resonances of the piezo phase shifter -> stiffer mounts are being installed



Low frequency noise hunting

- We start investigation on low frequency peaks, looking for correlations with mechanical resonances and environmental sensors
- Preliminary actions to be taken :
 - improve bench seismic isolation
 - Improve bench acoustic isolation



Summary

- Integration of squeezed vacuum source in TAMA completed
 - ~4.5 dB of squeezing down to ~30 kHz
 - Current performances limited by phase noise
- Filter cavity installed, controlled and characterized:
 - Cavity round trip losses between 45 and 85 ppm, in agreement with the requirement

Next steps

- Increase the squeezing level and reach the low frequency
 - Phase noise control loop improvement
 - Intensive low frequency noise-hunting
- Upgrade of the filter cavity control
 - Automatic alignment implementation
 - Test of new control scheme
- Preparation for squeezing injection into the filter cavity
- Design for filter cavity integration in KAGRA

Visitors and collaborations

- Fruitful collaboration with groups from Europe and Taiwan
- Many contributions given from visitors



Emil Schreiber
GEO600/AEI



Shu-Rong Wu
Tsing Hua University



Marco Vardaro
Padova University



Matteo Barsuglia
APC/CNRS



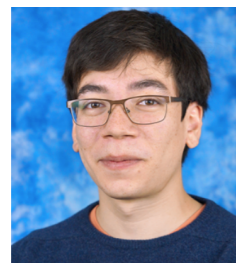
Eleonora Polini
La Sapienza Roma



Matteo Tacca
Nikhef



Federico Paoletti
INFN-Pisa



Marc Eisenmann
LAPP/CNRS



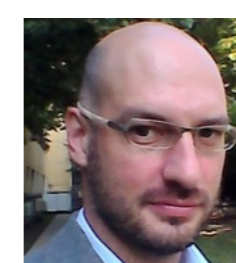
Yuefan Guo
Nikhef



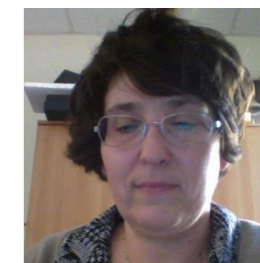
Chien-Ming Wu
Tsing Hua University



Pierre Prat
APC/CNRS



Marco Banzan
Padova University



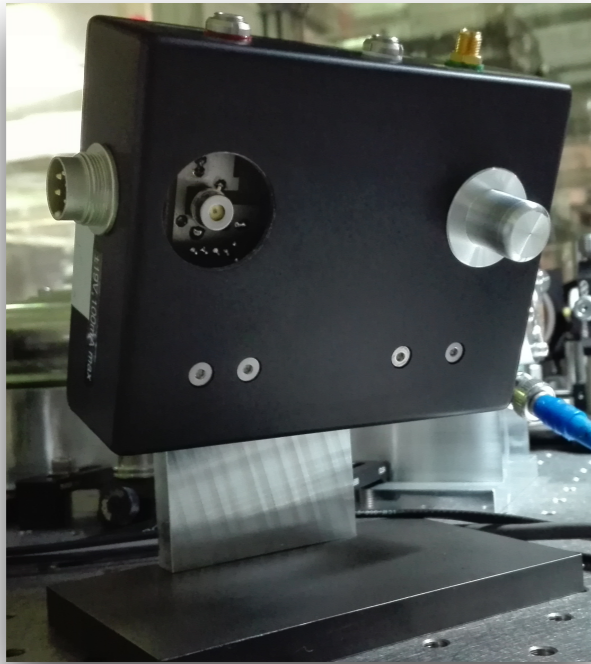
Irene Fiori
EGO

visits supported by JSPS Core-to-Core Program, A.Advanced Research Networks EU under the Framework Program 7 (FP7)
'People' project ELiTES (Grant Agreement No. 295153)

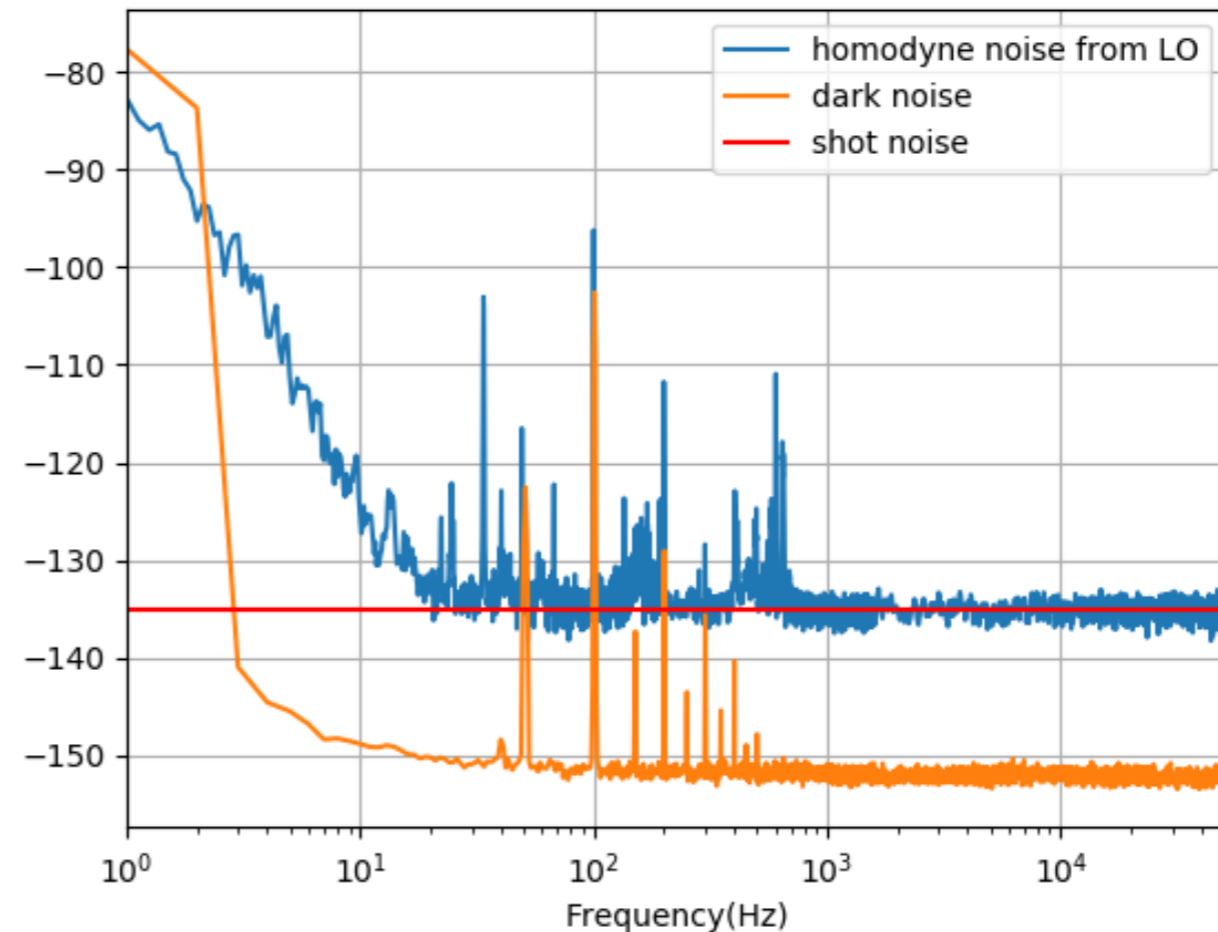
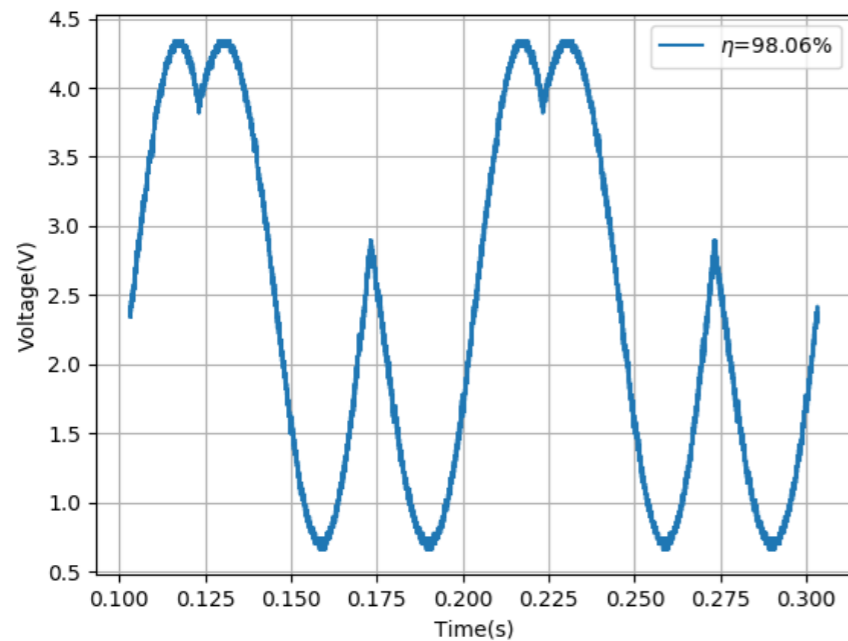
EU Horizon 2020 Research and Innovation Programme under the Marie Skłodowska-Curie Grant Agreement No.734303

EXTRA SLIDES

Homodyne detector

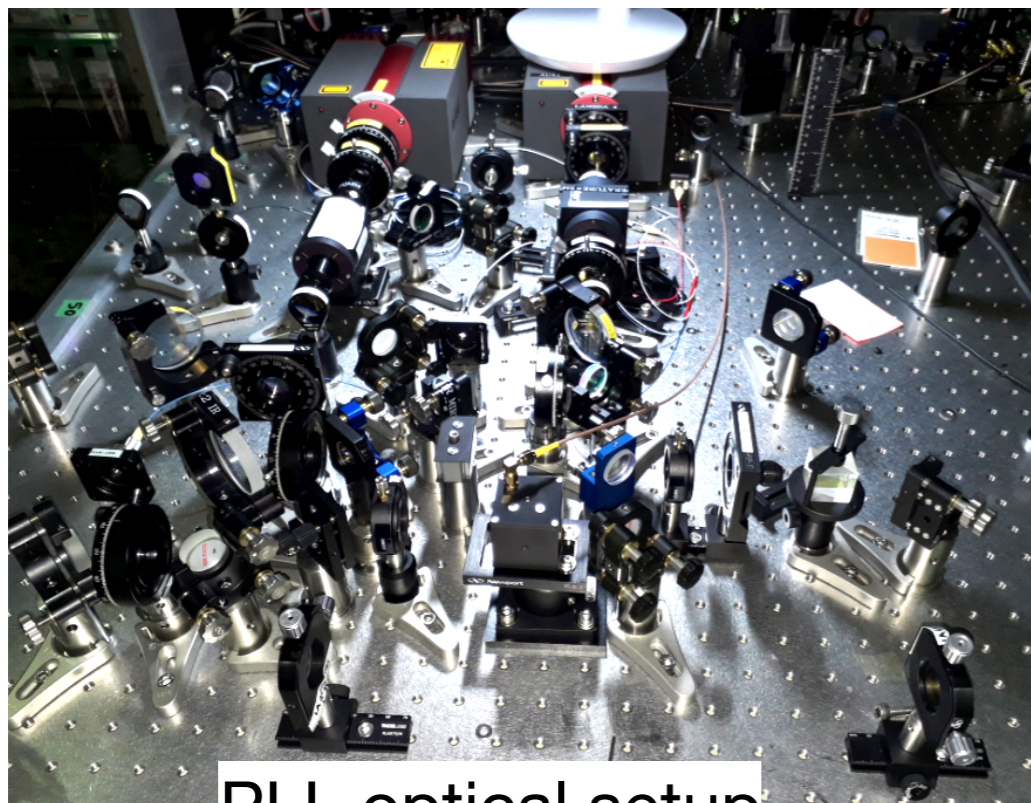


- Developed at AEI
- Clearance >15 dB
- Visibility $\sim 99\%$
- Shot noise limited down to 20 Hz
- Investigation on residual peaks and structures ongoing

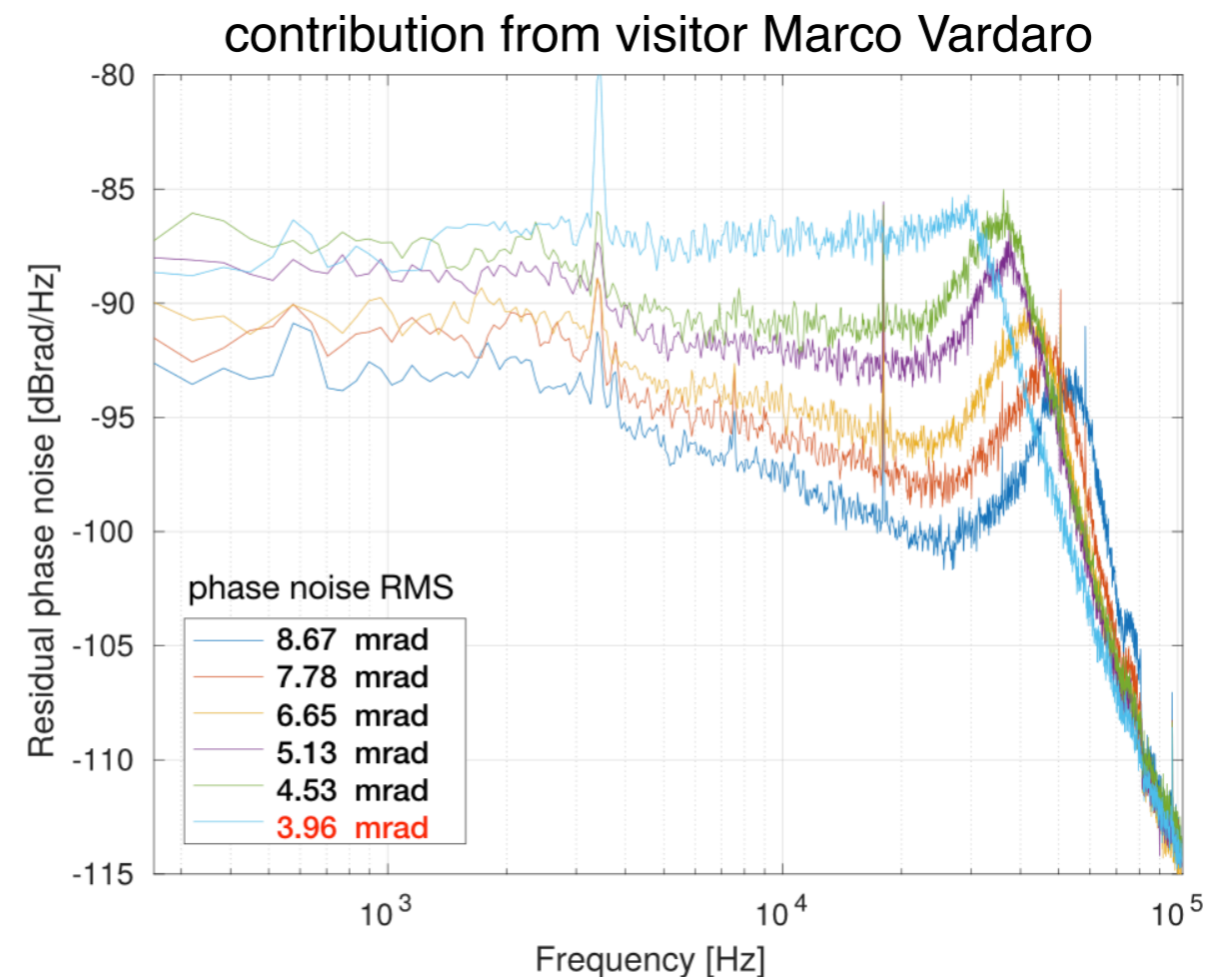


Phase lock loop (PLL) for AUX lasers

- Two AUX lasers are used for OPO length lock and control of the squeezing field phase (coherent control). They need to be phase locked to the main laser
- PLLs realized with fibered beam splitter and fiber coupled photodiode
- Electronics based on commercial Phase Frequency Detector (ADF4002)
- Residual phase noise ~ 4 mrad RMS between 100 Hz and 100 kHz



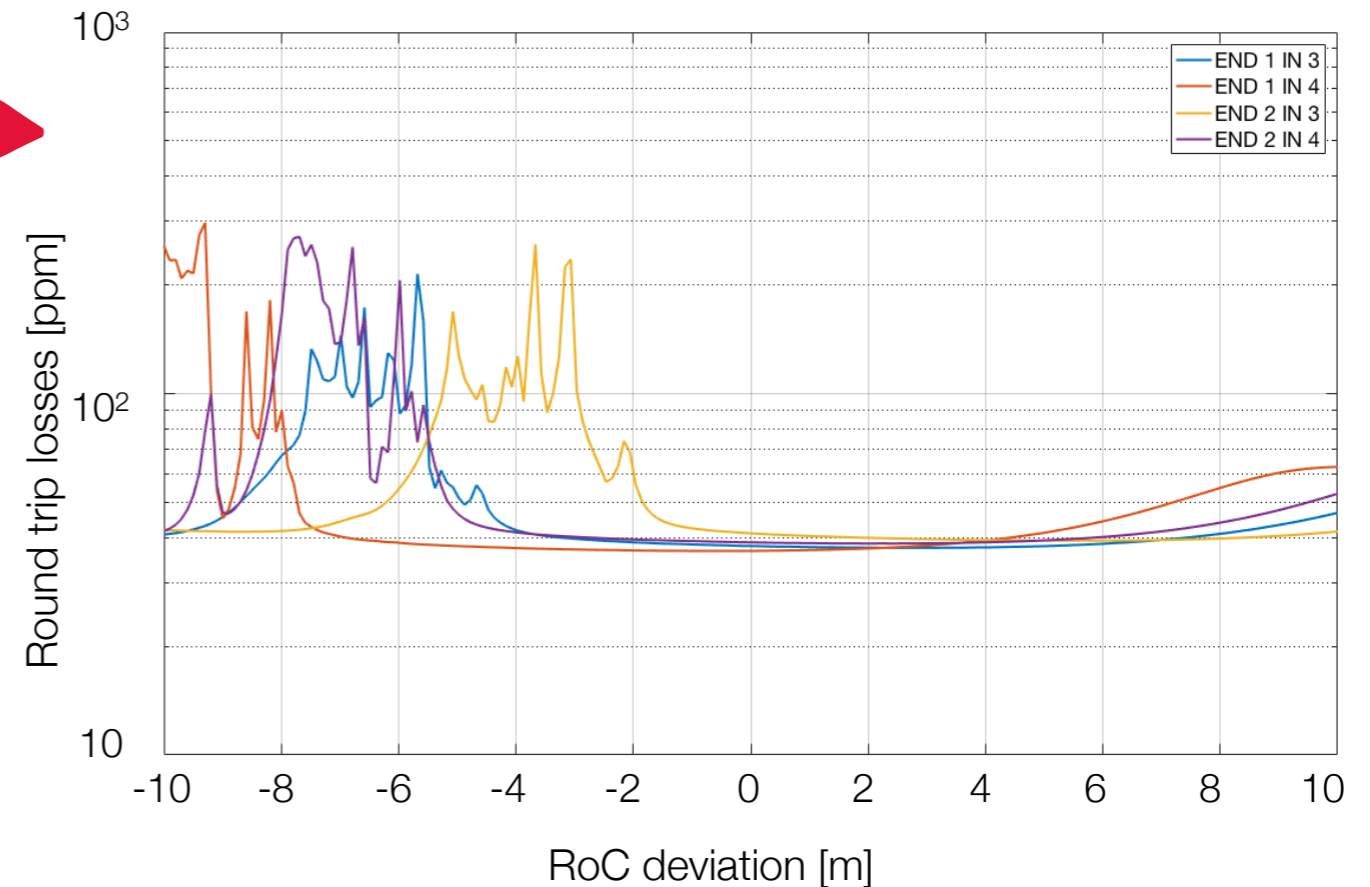
PLL optical setup



Round trip losses budget

RTL REQUIREMENT : 80 ppm

- ~ **40 ppm** from flatness (simulation)
- ~ **15 ppm** from roughness and point defect (measured)
- ~ **5 ppm** from absorption and transmission (measured)



TOTAL EXPECTED RTL : ~ 60 ppm

Why ppm/meter are important?

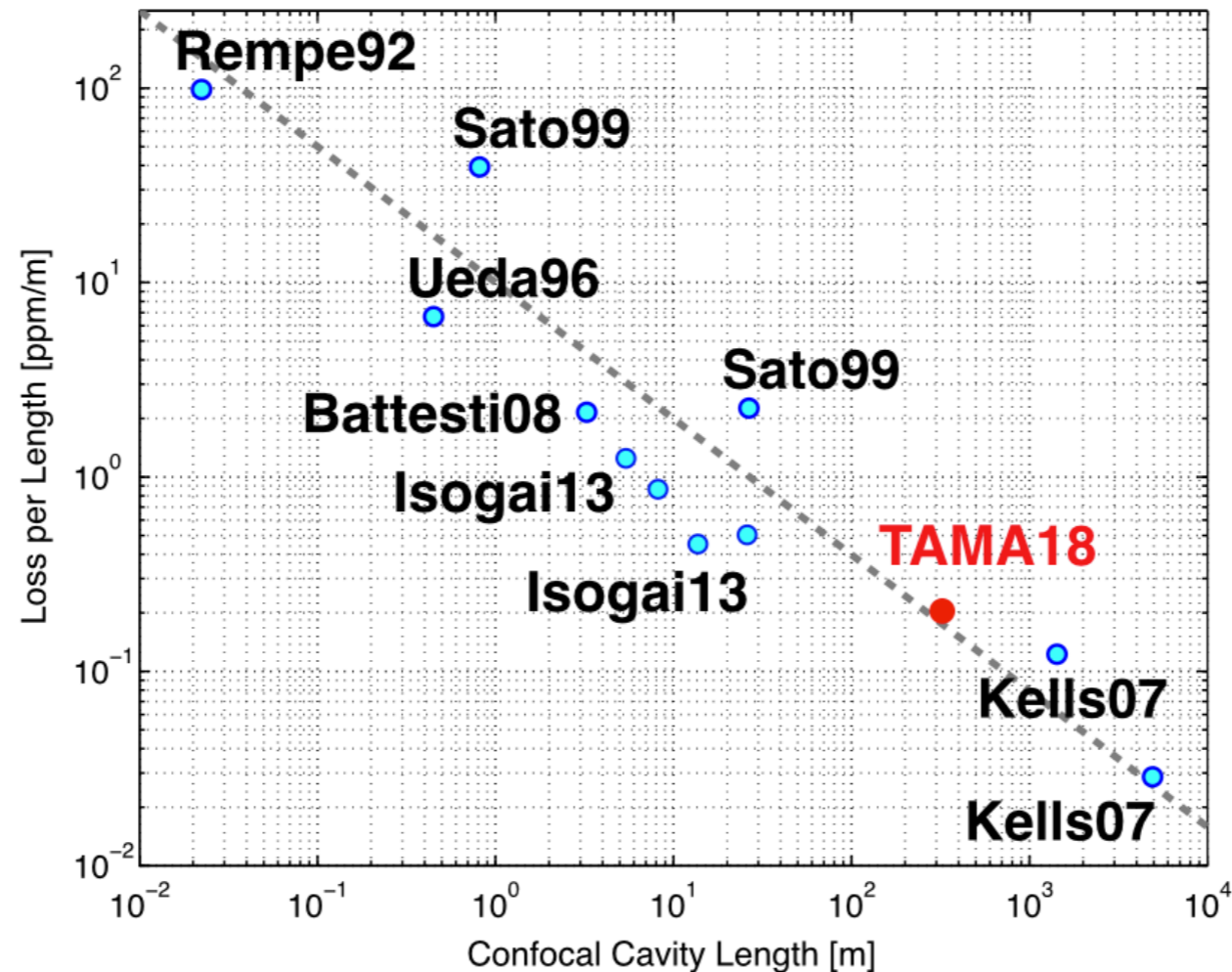
Total losses: RTL per number of round trip $N \sim 1/T_f$

$$\mathcal{E} \approx \frac{\epsilon}{T_f}$$

Optimal rotation: filter Cavity bandwidth comparable with ITF bandwidth γ

$$T_f \approx \frac{4\gamma L_f}{c} \quad \longrightarrow \quad \mathcal{E} \approx \frac{c\epsilon}{4\gamma L_f} \propto \frac{\epsilon}{L_f}$$

Comparison with round trip losses in literature



$$L_{\text{rt}}(\mathcal{L}_{\text{confocal}}) = 10 \text{ ppm} \cdot \left(\frac{\mathcal{L}_{\text{confocal}}}{1 \text{ m}} \right)^{0.3}$$

$\mathcal{L}_{\text{confocal}}$ is the length of the confocal cavity which has the same spot size at its mirrors as the cavity whose losses are reported

PHYSICAL REVIEW D **88**, 022002 (2013)

Realistic filter cavities for advanced gravitational wave detectors

M. Evans, L. Barsotti, and P. Kwee

Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

J. Harms

INFN, Sezione di Firenze, Sesto Fiorentino 50019, Italy

H. Miao

California Institute of Technology, Pasadena, California 91125, USA

(Received 9 May 2013; published 29 July 2013)

How to measure round trip losses?

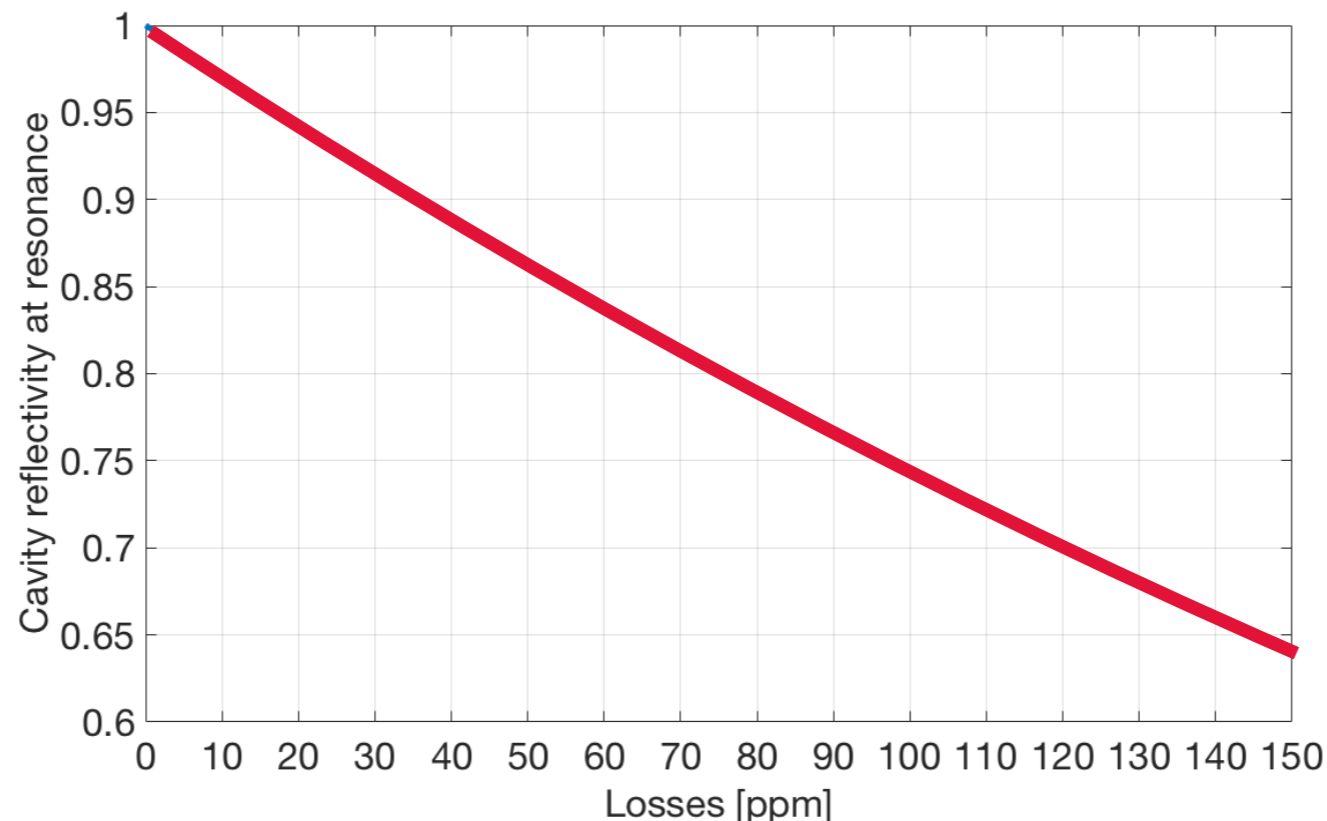
- From the cavity reflectivity at resonance
- Reflectivity is less affected from the input mirror transmissivity (with respect to finesse or decay time)

$$R_{\text{cav}} = \left[\frac{r_1 - r_2}{1 - r_1 r_2} \right]^2$$

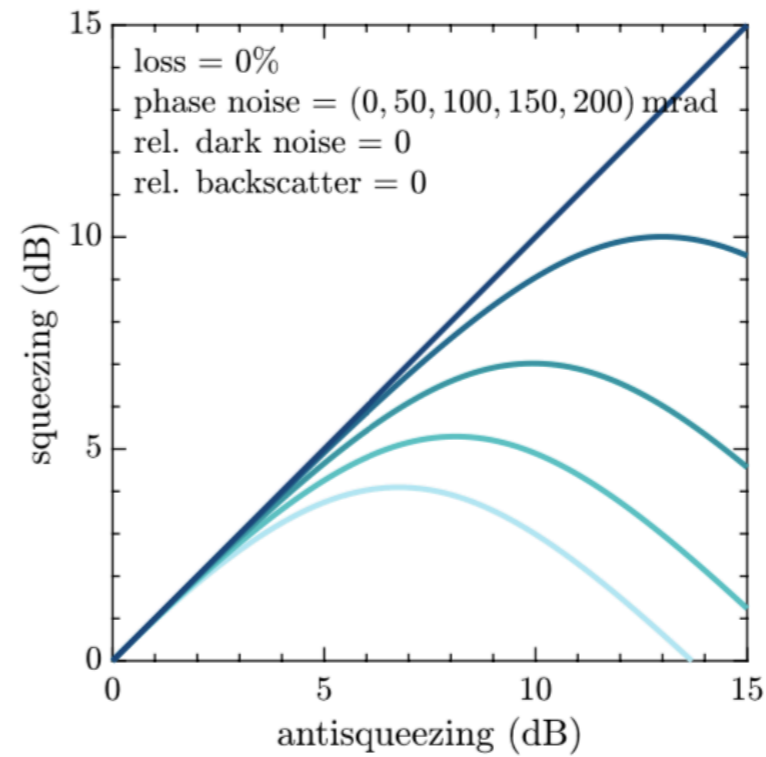
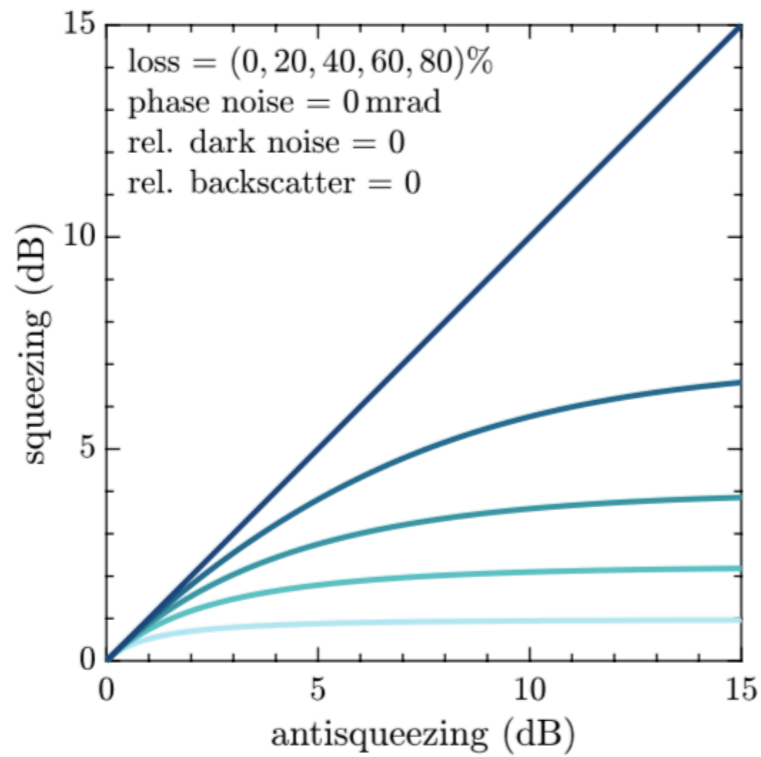


$$r_2 = \sqrt{1 - T_2 - L} \simeq \sqrt{1 - L}$$

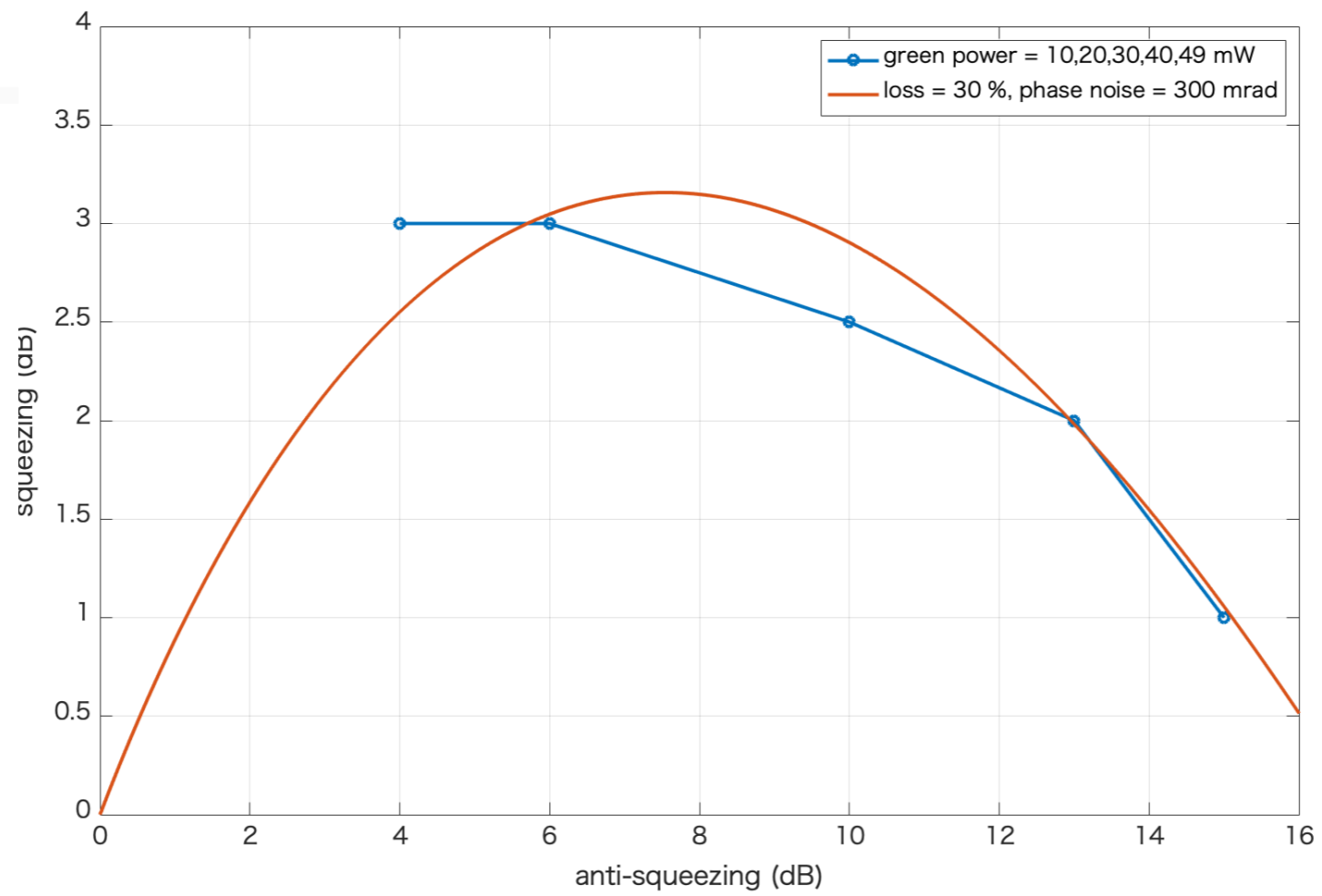
$$L = \frac{T_1}{2} \cdot \frac{1 - R_{\text{cav}}}{1 + R_{\text{cav}}}$$



Phase noise vs optical losses

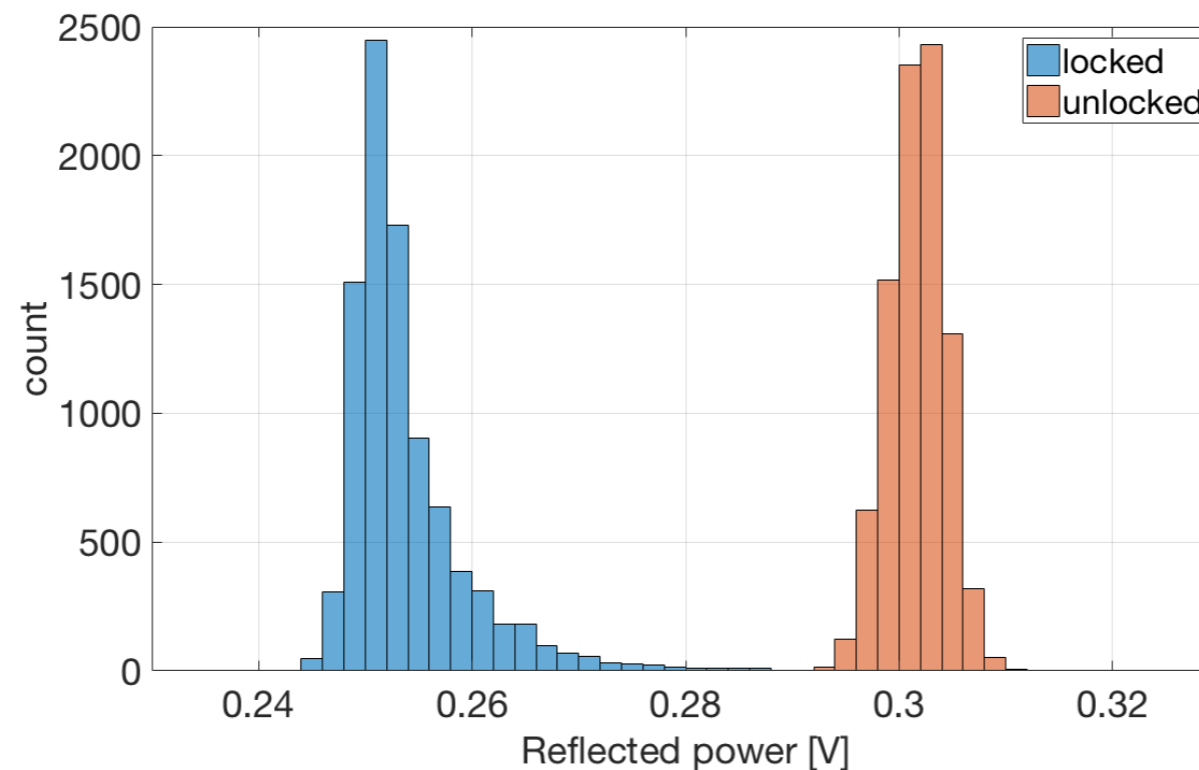


From E.Schreiber
thesis



How to estimate the reflected power

- Reflected power has some fluctuations which show different features if the cavity is locked or unlocked



- Cavity unlocked: Gaussian histogram. Main influence: input power fluctuations
- Cavity unlocked: asymmetric distribution. Influence of the input power fluctuations, cavity alignment fluctuations, finite lock accuracy