

Accelerated observers and Planck-scale kinematics

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Upon introduction of a **“brick wall”** regulator obtain **entropy density** $\sim 1/L_p^2$

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The common view is that **quantum properties of spacetime** are responsible for a **finite horizon entropy density** in the same way quantization of the electromagnetic field leads to a finite **black-body entropy**

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Work in collaboration with Master's student M. Laudonio

(Phys. Rev. D **97**, no. 8, 085004 (2018))

Accelerated observers

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Four-velocity of observer with **acceleration** α

$$U^\mu = (\cosh \alpha\tau, \sinh \alpha\tau, 0, 0)$$

Lorentz boost by $\eta = \alpha\tau$ of four-velocity of static Minkowski observer $U^\mu = (1, 0, 0, 0)$

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Using a **dilation** generated by $D = -i x^\mu \partial_\mu$: a finite transformation of parameter δ

$$(t, x) \rightarrow (t', x') = e^\delta(t, x) \implies \boxed{\alpha \rightarrow \alpha' = e^{-\delta}\alpha}$$

Rindler space

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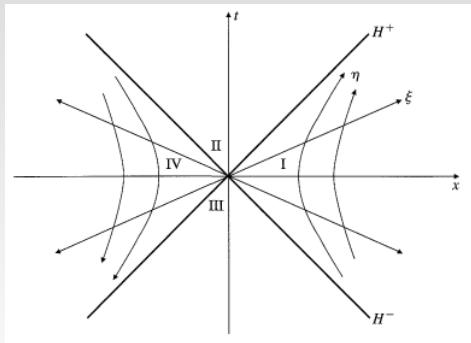
Define spatial **Rindler coordinate** ξ in terms of the **dilation parameter** $\delta = a\xi$
(with $1/a = [\text{length}]$)

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Accelerated observers and the Weyl-Poincaré group

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Weyl-Poincaré algebra in 1+1 dimensions

$$\begin{aligned}[P_t, P_x] &= 0, & [D, N] &= 0 \\ [N, P_t] &= iP_x, & [N, P_x] &= iP_t \\ [D, P_t] &= iP_t, & [D, P_x] &= iP_x\end{aligned}$$

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Besides usual reps in terms of $P_{t,x} = i\partial_{t,x}$ we have an **alternative reps** in terms of Rindler coordinates ξ, η

$$P_\xi = aD = i\partial_\xi, \quad P_\eta = aN = i\partial_\eta$$

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Note the role of acceleration scale a in order to get the right dimensions for P_ξ and P_η

Aside: the Unruh effect without space-time

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The simplest non-abelian Lie algebra

$$[D, P] = iP$$

relationship between **boundary** and **thermal effects** from representation theory
(work with Kowalski-Glikman: arXiv:1804.10550)

Rindler coordinates and reps of the Weyl-Poincaré algebra

Representation of the Weyl-Poincaré generators in terms of $(\partial_\xi, \partial_\eta)$

$$P_\xi = i\partial_\xi$$

$$P_\eta = i\partial_\eta$$

$$P_t = ie^{-a\xi}(\cosh a\eta \partial_\eta - \sinh a\eta \partial_\xi)$$

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Rindler mass shell obtained from mass Casimir

$$\mathcal{C} = P_0^2 - P_x^2 = e^{-2a\xi}(P_\eta^2 - P_\xi^2),$$

Doppler shift and Rindler horizon

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- At $\xi = -\infty$ i.e. on the light cone $x = |t|$ the photon's frequency will appear **infinitely blueshifted** (in analogy with Schwarzschild horizon)

The **Rindler horizon** is described by an *infinite contraction* generated by D

Mode counting in Minkowski space

Number of modes of a massless scalar field in a $3D$ box of size L

- **Nodes** on the wall of box

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by trivial integration leads to the known result for **density of states in Minkowski space**

$$n(E) = \frac{L^3 E^3}{6\pi^2}$$

used to derive all thermodynamical properties of a free massless field.

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Density of states: 3+1 Minkowski vs. Rindler

$$dn_M = \frac{L^3}{2\pi^2} k^2 dk \quad \text{vs} \quad dn_R = \frac{L^2 k_{\perp} k_{\xi}}{2\pi^2} dk_{\perp} d\xi$$

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With a “**brick wall**” at ξ_{min} we obtain **density of states** as function of energy

$$n_R(E) = \frac{E^3 L^2}{6\pi^2} \int_{\xi_{min}}^{\log(aR)/a} d\xi e^{-2a\xi} = \frac{E^3 L^2}{12\pi^2} \frac{1}{a} \left[e^{-2a\xi_{min}} - \frac{1}{(aR)^2} \right].$$

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$$S = \frac{\pi^2}{45} \frac{L^2}{a\beta^3} \left[e^{-2a\xi_{min}} - \frac{1}{(aR)^2} \right] = S_{wall} + \text{IR box contribution},$$

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For $\beta \sim 1/T_U \sim 2\pi/a$, ξ_{min} can be fixed \Rightarrow **Bekenstein-Hawking entropy density**

$$\sigma_{wall} = S_{wall}/L^2 = \frac{1}{4L_p^2}$$

The covariant recipe for counting states

State counting can be recast in a covariant framework by considering **state space of a massless relativistic field**

$$dn \sim \underbrace{2E dt dx^3 \delta(t)}_{\text{covariant space volume}} \times \underbrace{d^4 p \delta(C) \theta(E)}_{\text{covariant momentum-space volume}}$$

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$$dn \sim \underbrace{2E dt dx^3 \delta(t)}_{\text{covariant space volume}} \times \underbrace{d^4 p \delta(C) \theta(E)}_{\text{covariant momentum-space volume}}$$

In Minkowski space, integrating over a spatial volume $V = L^3$

$$n_M(E) = \frac{V}{(2\pi)^3} \int_E d^4 p 2p_0 \delta(p^2) \theta(p_0) = \frac{L^3 E^3}{6\pi^2}$$

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For a **Rindler field**

$$n_R(E) = \frac{V_{\perp}}{(2\pi)^3} \int_{\mathbb{R}} d\xi \int dp_{\eta} dp_{\xi} dp_{\perp}^2 2p_{\eta} e^{-2a\xi} \delta(\mathcal{C}) \theta(p_{\eta}) = \frac{V_{\perp}}{(2\pi)^3} \frac{4\pi}{3} E^3 \int_{-\infty}^{\infty} d\xi e^{-2a\xi}$$

The twisted Weyl-Poincaré algebra

The Weyl-Poincaré algebra $\mathfrak{ptw}(3,1)$ in 3 + 1 dimensions

$$[P_\mu, P_\nu] = 0, \quad [P_\mu, M_{\rho\nu}] = i(\eta_{\mu\rho}P_\nu - \eta_{\mu\nu}P_\rho)$$

$$[M_{\mu\nu}, M_{\rho\sigma}] = i(\eta_{\mu\sigma}M_{\nu\rho} - \eta_{\mu\rho}M_{\nu\sigma} + \eta_{\nu\rho}M_{\mu\sigma} - \eta_{\nu\sigma}M_{\mu\rho})$$

$$[D, P_\mu] = iP_\mu, \quad [D, M_{\mu\nu}] = 0$$

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Consider deformation by **Jordanian twist** (Aschieri, Borowiec and Pachol, JHEP **1710**, 152 (2017))

$$W^{\mathcal{F}} = \bar{f}^\alpha(W)\bar{f}_\alpha, \quad W \in \mathfrak{pw}(3, 1)$$

where $\mathcal{F} = f^\alpha \otimes f_\alpha = \exp(-iD \otimes \sigma)$, $\sigma = \log(1 + \ell P_0)$ and $\mathcal{F}^{-1} = \bar{f}^\alpha \otimes \bar{f}_\alpha$

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The resulting *twisted generators*

$$P_\mu^{\mathcal{F}} = \frac{P_\mu}{1 + \ell P_0}, \quad M_{\mu\nu}^{\mathcal{F}} = M_{\mu\nu}, \quad D^{\mathcal{F}} = D.$$

very similar to the “non-linear” redefinition of translation generators used by Magueijo and Smolin in their early DSR model (Phys. Rev. Lett. **88**, 190403 (2002))

The twisted Weyl-Poincaré algebra (continued)

In terms of the *twisted commutator*

$$[W^{\mathcal{F}}, V^{\mathcal{F}}]_{\mathcal{F}} = W^{\mathcal{F}} V^{\mathcal{F}} - (\bar{R}^{\alpha}(V))^{\mathcal{F}} (\bar{R}_{\alpha}(W))^{\mathcal{F}}.$$

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$$[D^{\mathcal{F}}, P_{\mu}^{\mathcal{F}}] = iP_{\mu}^{\mathcal{F}} - i\ell P_{\mu}^{\mathcal{F}} P_0^{\mathcal{F}}$$

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while all other commutators remain undeformed.

The mass Casimir $\mathcal{C} = P_{\mu} P^{\mu}$ in terms of the twisted translation generators $P_{\mu}^{\mathcal{F}}$ becomes

$$\mathcal{C}^{\mathcal{F}} = \frac{(P_{\mu} P^{\mu})^{\mathcal{F}}}{(1 - \ell P_0^{\mathcal{F}})^2}.$$

At the algebraic level this is all we need to go and play the “DSR game”

Twisted DSR finite boosts in the 1-direction

From the deformed algebra we have

$$\begin{aligned}
 \frac{d\omega}{d\phi} &= -i[N_1, \omega] = k_1(1 - \ell\omega) & \omega(\phi) &= \frac{\omega^0 \cosh \phi + k_1^0 \sinh \phi}{A} \\
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where $A = 1 - \ell\omega^0 + \ell\omega^0 \cosh \phi + \ell k_1^0 \sinh \phi$

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Boosts saturate at the Planck scale!

$$\lim_{\phi \rightarrow \infty} \omega(\phi) = \frac{1}{\ell}, \quad \lim_{\phi \rightarrow \infty} k_1 = \frac{1}{\ell}, \quad \lim_{\phi \rightarrow \infty} k_i = 0$$

Twisted dilations

The same procedure can be used to derive the **twisted dilation transformation**

$$\begin{aligned} \frac{d\omega}{d\delta} &= -i[D, \omega] = \omega(1 - \ell\omega) \\ \frac{dk_i}{d\delta} &= -i[D, k_i] = k_i(1 - \ell\omega) \end{aligned} \quad \implies \quad \begin{aligned} \omega(\delta) &= \frac{\omega^0}{\omega^0\ell + (1 - \omega^0\ell)e^{-\delta}} \\ k_i(\delta) &= \frac{k_i^0}{\omega^0\ell + (1 - \omega^0\ell)e^{-\delta}} \end{aligned}$$

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For $\delta \rightarrow \infty$ **dilation transformations saturate at the Planck scale!**

$$\lim_{\delta \rightarrow \infty} \omega(\delta) = \frac{1}{\ell} \quad \lim_{\delta \rightarrow \infty} k_i(\delta) = \frac{k_i^0}{\ell\omega^0}$$

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Rindler translation generators associated to accelerated observers in the 1-direction

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Deformed blueshift = inverse dilation

$$\omega_\xi = \frac{\omega}{\ell\omega + (1 - \ell\omega)e^{a\xi}}.$$

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Finite blueshift at the accelerated horizon $\xi \rightarrow -\infty$!

Could this act as a "covariant brick wall"?

Counting modes in twisted Poincaré

Warm up: mode counting for a field in twisted Poincaré

$$n(E) = \frac{V}{(2\pi)^3} \int_E d\mu(p) 2p_0 \delta(C) \theta(p_0),$$

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To LIV or not to LIV? (Gubitosi and Magueijo, *Class. Quant. Grav.* 33, no. 11, 115021 (2016))

Breaks Lorentz


$$d\mu(p)_{\text{LIV}} = d^4 p$$

Covariant under deformed boosts


$$d\mu(p)_C = \frac{d^4 p}{(1 - \ell p_0)^5}$$

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The resulting density of states are

$$n(E)_{\text{LIV}} = \frac{2V}{(2\pi)^2} \left[\frac{E^3}{3} - \frac{\ell E^4}{2} + \frac{\ell^2 E^5}{5} \right]$$

$$n(E)_C = \frac{V}{(2\pi)^2} \frac{1}{\ell^3} \left[\frac{\ell E(3\ell E - 2)}{(1 - \ell E)^2} - 2 \log(1 - \ell E) \right].$$

Finite density of states for LIV measure

Boosts saturate at $1/\ell$, **maximal energy**, what about density of states?

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However using the LIV measure

$$\lim_{E \rightarrow 1/\ell} n(E)_{\text{LIV}} = \frac{V}{(2\pi)^2} \frac{1}{15\ell^3},$$

we have a **finite number of states all the way up to the Planck scale**

Deformed Rindler: brick-wall from twist?

Look at **twisted generalization** of

$$n(E) = \frac{V_{\perp}}{(2\pi)^3} \int_{\mathbb{R}} d\xi \int dp_{\eta} dp_{\xi} dp_{\perp}^2 2p_{\eta} e^{-2a\xi} \delta(C) \theta(p_{\eta}).$$

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Sparing you the **gruesome details** the **final result** is

$$n(E) = \frac{V_{\perp}}{(2\pi)^3} \int_{\mathbb{R}} d\xi \int d\mu(p_{\eta}, p_{\xi}, p_{\perp}) \frac{p_{\eta} (1 - \ell p_{\eta})^2 e^{2a\xi} \delta(p_{\eta} - \omega_p)}{(\ell p_{\eta} + (1 - \ell p_{\eta}) e^{a\xi})(p_{\eta} e^{a\xi} + \ell p_{\xi}^2 (1 - e^{a\xi}))} \theta(p_{\eta})$$

where ω_p = on-shell energy obtained from **deformed Rindler dispersion relation**

$$C^{\mathcal{F}} = \frac{e^{-2a\xi}}{(1 - \ell P_{\eta}^{\mathcal{F}})^2} \left[-(P_{\eta}^{\mathcal{F}})^2 + (P_{\xi}^{\mathcal{F}})^2 + (P_{\perp}^{\mathcal{F}})^2 (P_{\eta}^{\mathcal{F}} \ell + (1 - \ell P_{\eta}^{\mathcal{F}}) e^{a\xi})^2 \right]$$

Brick-wall from twist? Only if LIV!

Calculate density of states in fully **covariant picture**

$$n(E)_C = \frac{V_{\perp}}{(2\pi)^2} \frac{e^{-2a\xi_{min}}}{a} \left[\frac{E^3}{3} - \frac{\ell E^4}{2} + \frac{\ell^2 E^5}{5} \right],$$

still need a “brick-wall” regulator ξ_{min} ...

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still need a “brick-wall” regulator ξ_{min} ...

If we LIV we get a **finite density of states!**

$$n(E)_{LIV} = -\frac{V_{\perp}}{(2\pi)^2} \frac{1}{6a} \frac{1}{\ell^3} \log(1 - \ell E),$$

a bitter win...

Bekenstein-Hawking entropy from “twisted brick wall”

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Evaluate the entropy associated to $n(E)_{\text{LIV}}$

Consider the logarithm of the **deformed partition function**

$$\log Q = \beta \int_0^{1/\ell} dE \frac{n(E)}{e^{\frac{\beta E}{1-\ell E}} - 1}$$

(deformed B-E distribution natural choice determined by the twist map $P_\mu^{\mathcal{F}} = \frac{P_\mu}{1+\ell P_0}$)

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Setting $\beta = \frac{1}{T_U} = \frac{2\pi}{a}$ we obtain

$$S_{\text{LIV}} = \frac{V_\perp}{144\pi\ell^2} - \frac{V_\perp a}{\ell} \frac{\zeta(3)}{32\pi^4} + \frac{V_\perp a^2}{2160\pi} + \mathcal{O}(\ell)$$

at leading order, with $\ell = \frac{L_P}{6\sqrt{\pi}}$ leads to the **Bekenstein-Hawking entropy**...

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- Planckian aspects of **Unruh and (Hawking)** quantum radiance.