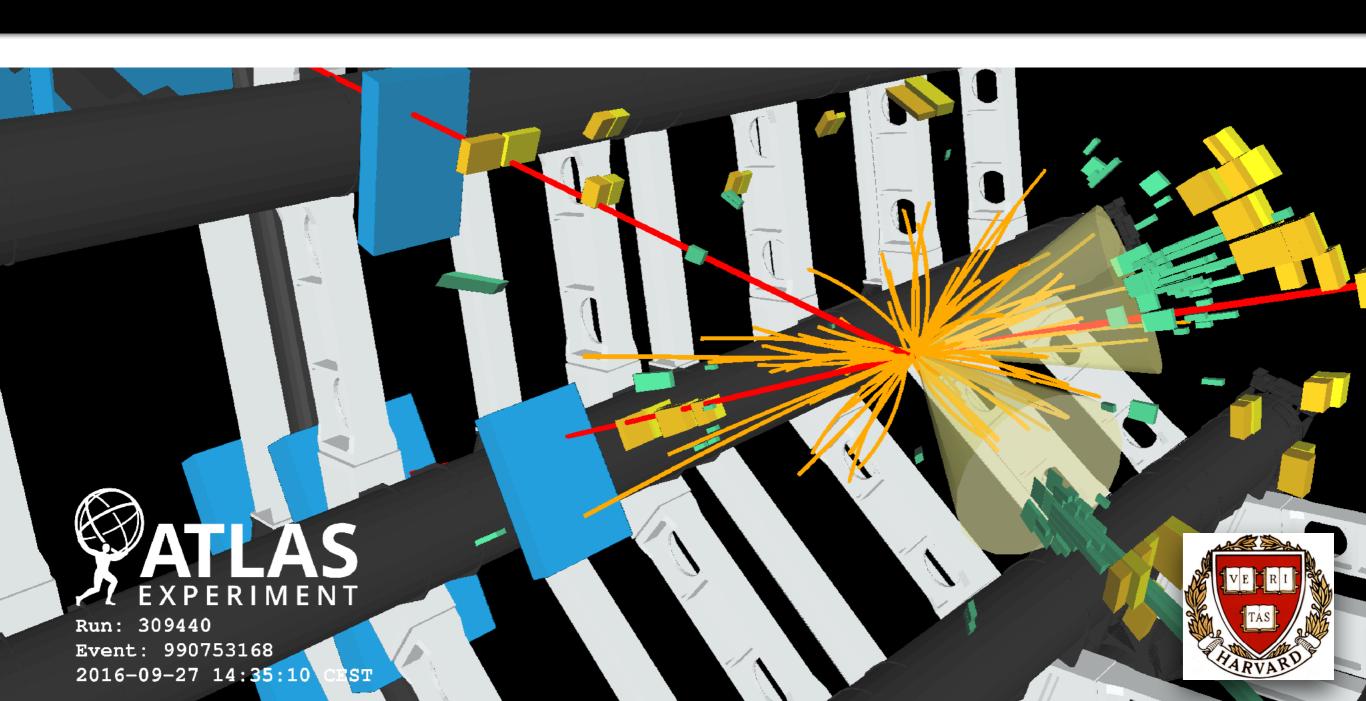
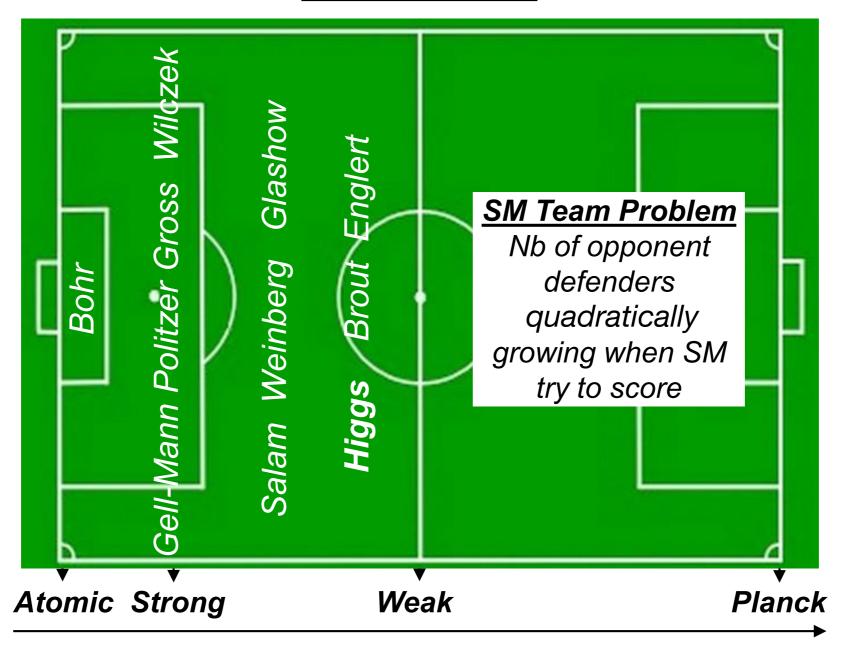
Stefano Zambito, Harvard University

Hunting for Supersymmetry at the LHC A Dive Into Naturalness... And Beyond



Quantum Field



E-scale

Quantum Field





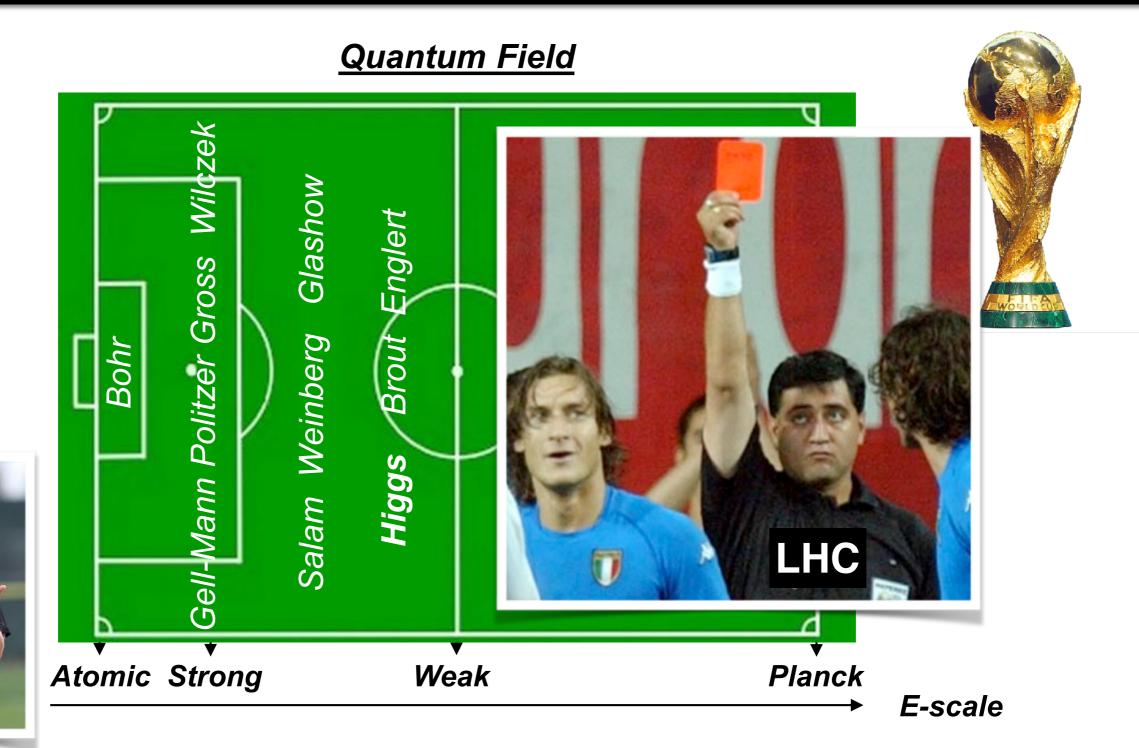


Atomic Strong Weak **Planck**

E-scale

SUSY coach (formed in Poincarè Academy in France*): knows how to annihilate defenders!

Doubling the number of SM team players...

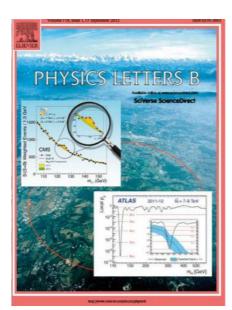


SUSY coach (formed in Poincarè Academy in France*): knows how to annihilate defenders!

Doubling the number of SM team players... Will the LHC referee allow that?

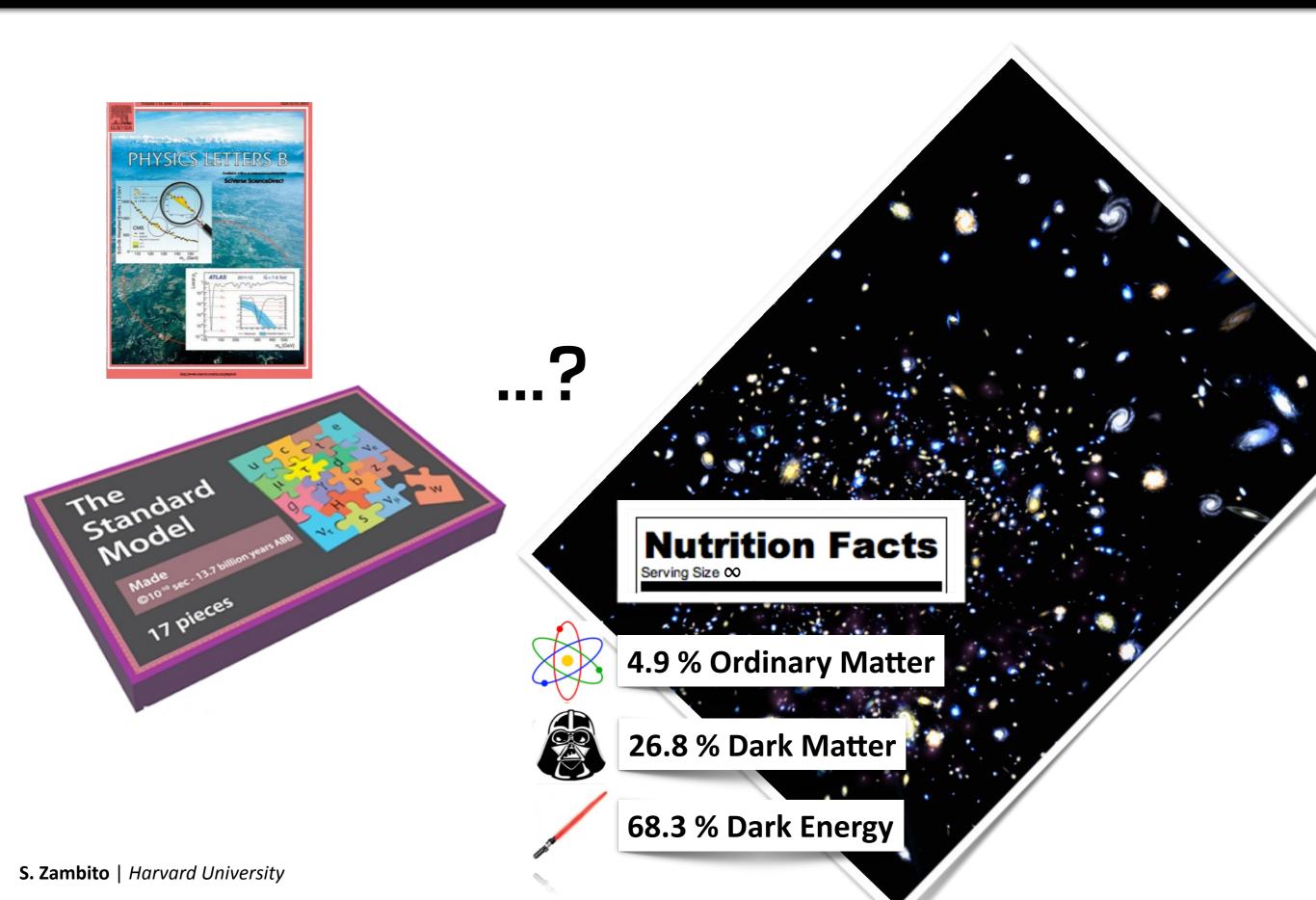
SUSY

The Standard Model, After The Higgs





The Standard Model, After The Higgs



The Standard Model, After The Higgs



Living On The Edge - The Tough Life Of A 125 GeV H Boson

Effective potential

Ginzburg-Landau: $V(\Psi) = \alpha |\Psi|^2 + \beta |\Psi|^4$

Higgs Potential: $V(\phi) = m_h^2 |\phi|^2 + \lambda |\phi|^4$

Dynamic

cooper pairs, etc...

333

Living On The Edge - The Tough Life Of A 125 GeV H Boson

Effective potential

Ginzburg-Landau: $V(\Psi) = \alpha |\Psi|^2 + \beta |\Psi|^4$

Higgs Potential: $V(\phi) = m_h^2 |\phi|^2 + \lambda |\phi|^4$

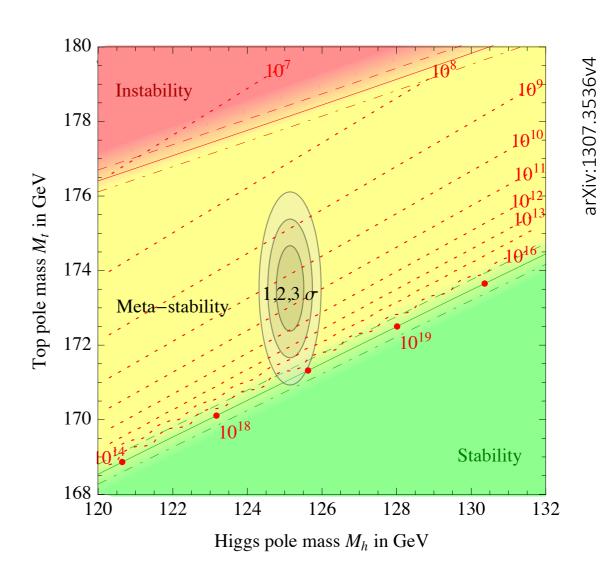
200 Instability Non-perturbativity Stability Stability 0 0 50 100 150 200

Higgs pole mass M_h in GeV

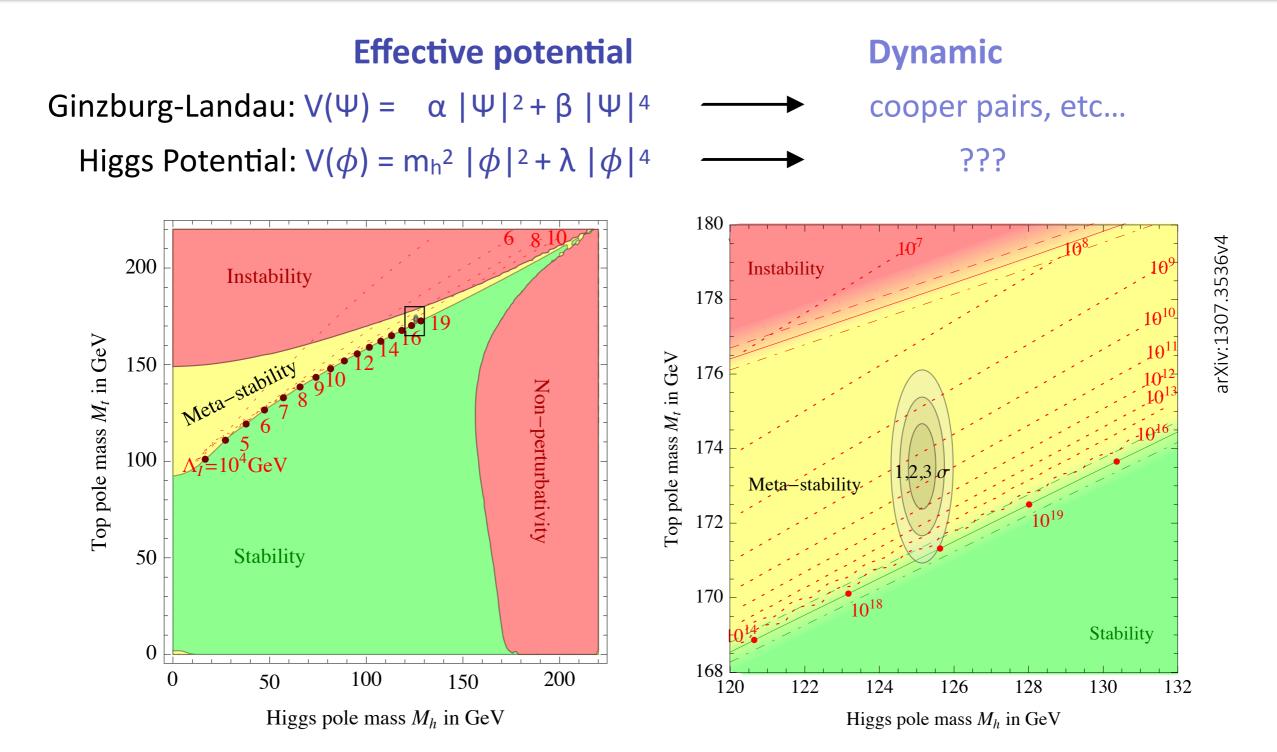
Dynamic

cooper pairs, etc...

???



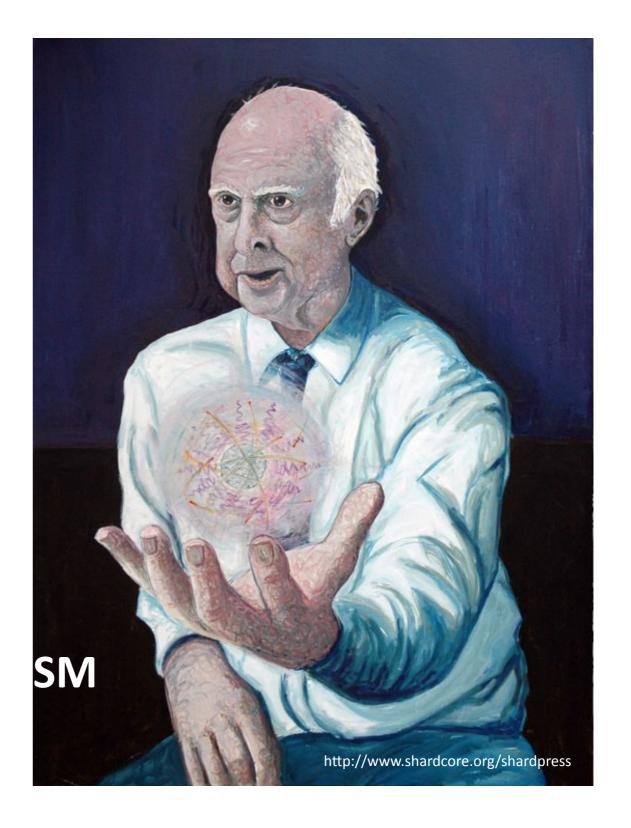
Living On The Edge - The Tough Life Of A 125 GeV H Boson

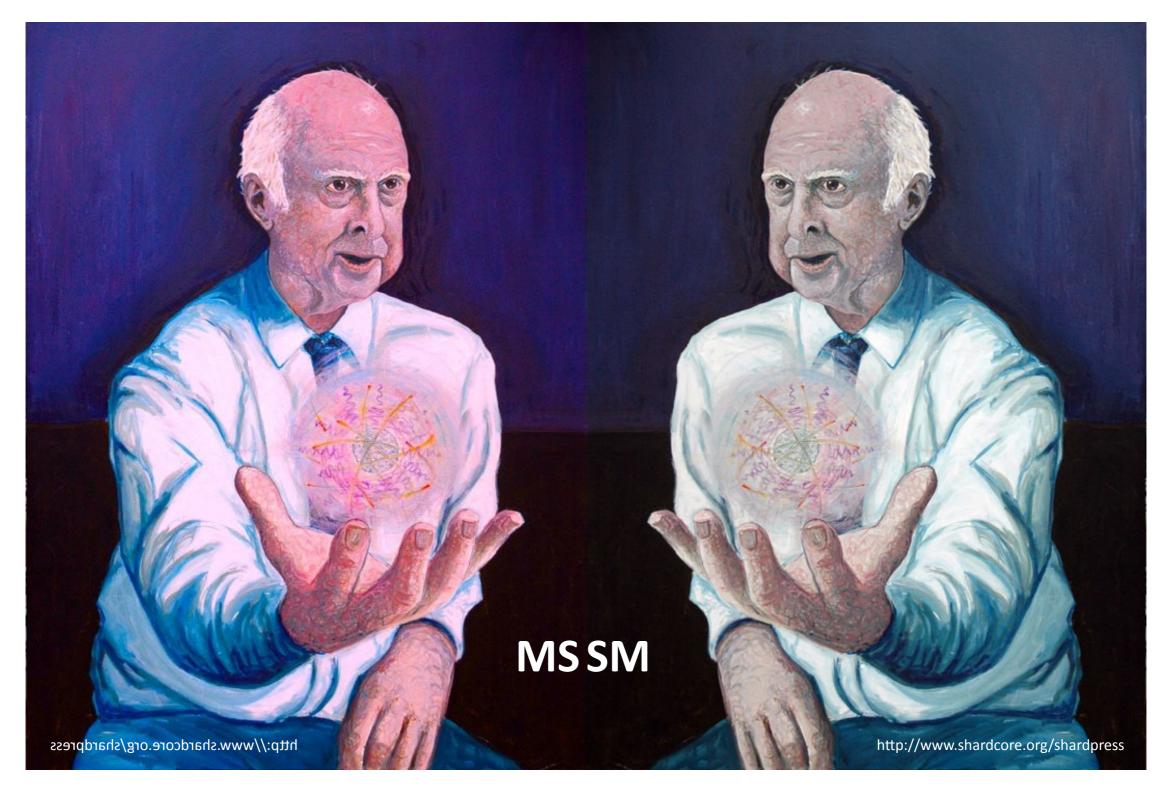


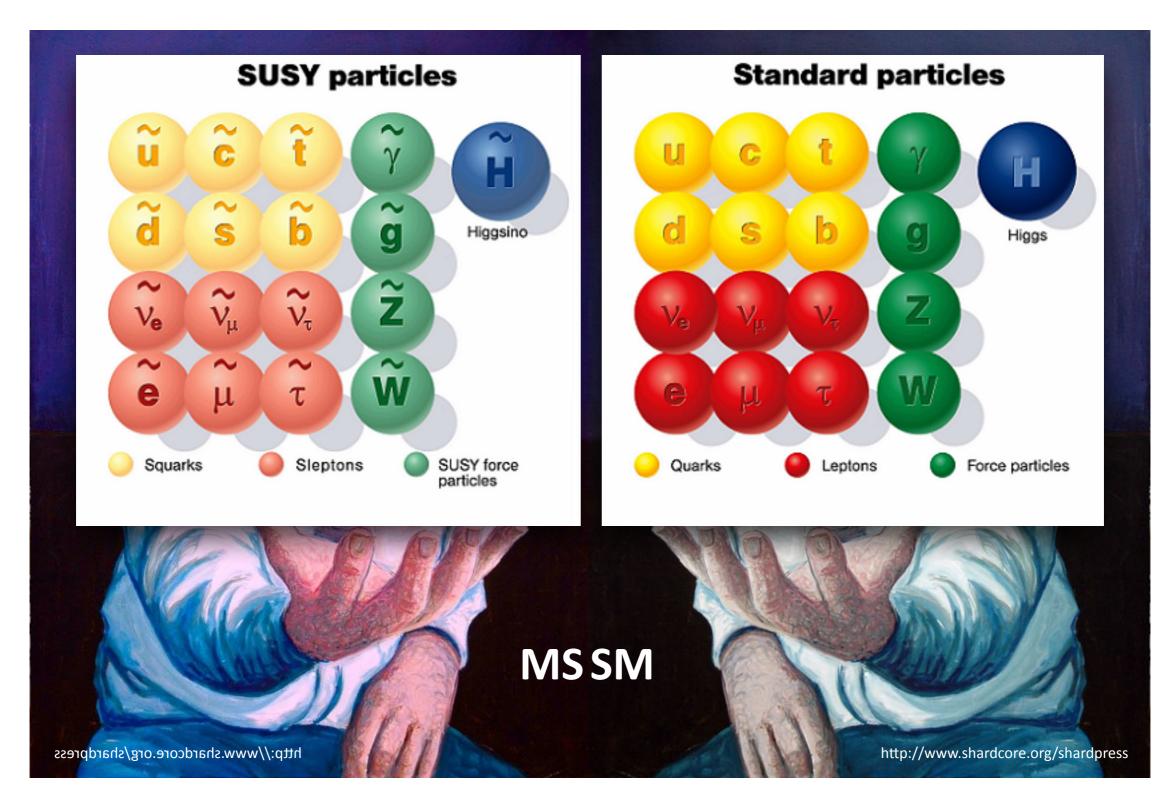
Why is the Higgs sitting so close to criticality?

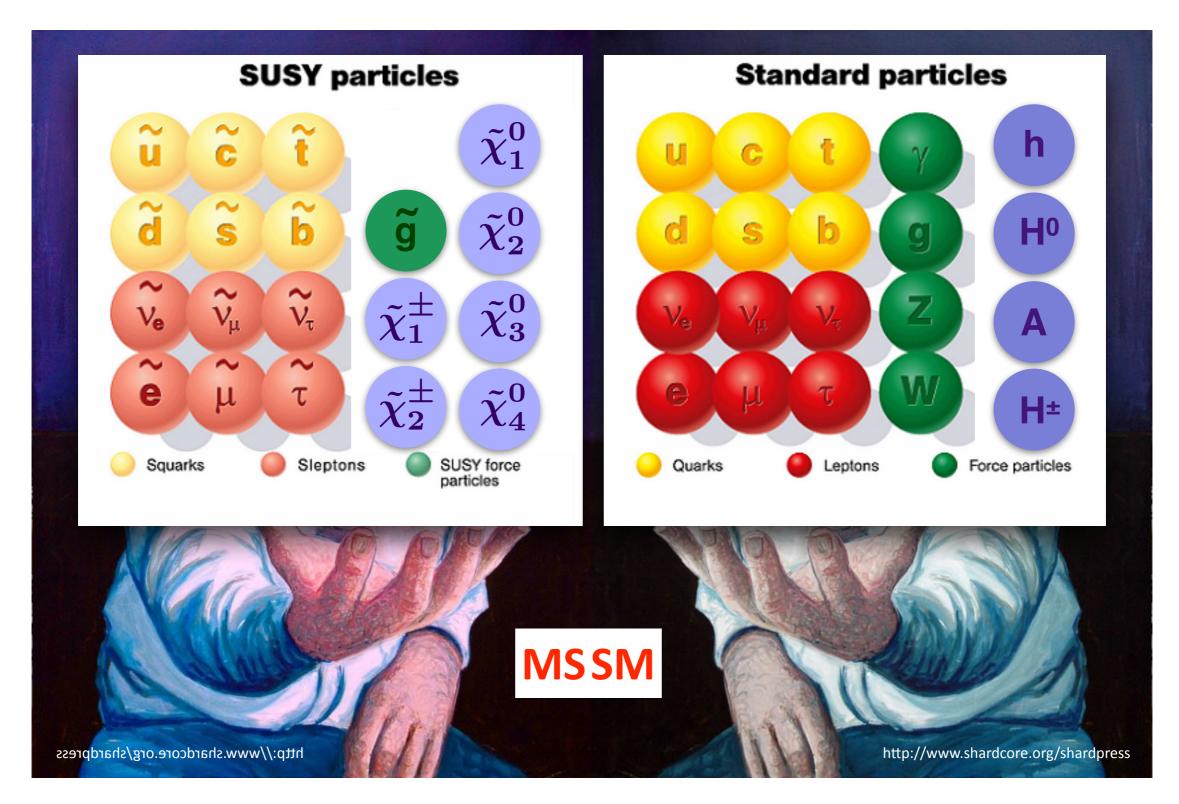
In the lab, critical systems are inevitably dragged away from critical point/line...

Why the Higgs so close to criticality?

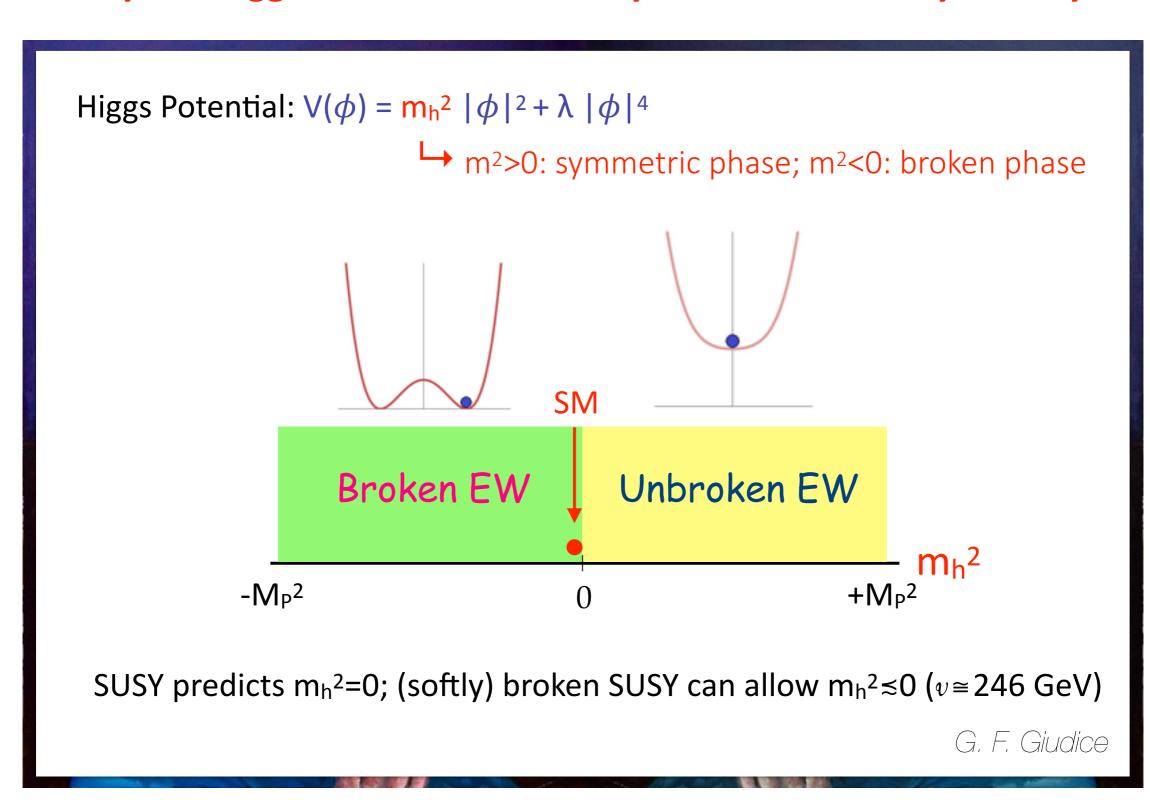


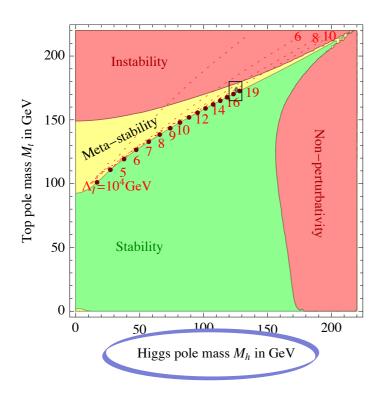






Higgs Naturalness As "Criticality" Condition



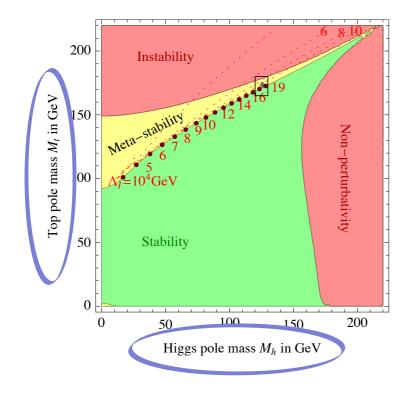


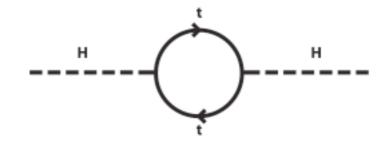
physical m_h: 125 GeV

$$m_h^2 \approx m_{h\,0}^2 - \frac{\lambda_f^2}{8\pi^2} N_c^f \int^{\Lambda} \frac{d^4p}{p^2} + \dots \approx m_{h\,0}^2 + \frac{\lambda_f^2}{8\pi^2} N_c^f \Lambda^2 + \dots$$

bare mass 1-loop correction

ultraviolet cutoff





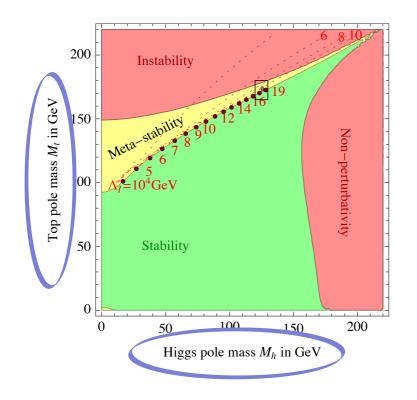
main "SM" term: top's loop (λ_t≈1)

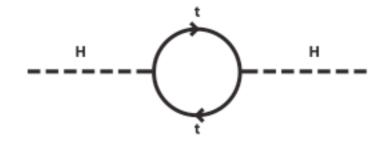
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bare mass 1-loop correction

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main "SM" term: top's loop ($\lambda_t \approx 1$)

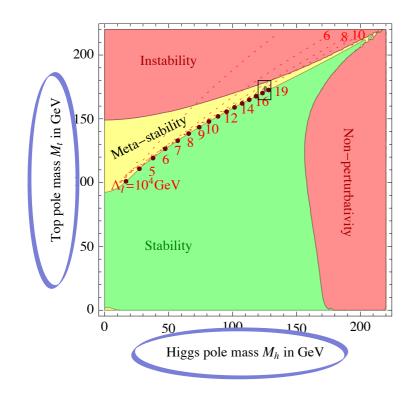
physical *m_h*: 125 GeV

$$m_h^2 \approx m_{h\,0}^2 - \frac{\lambda_f^2}{8\pi^2} N_c^f \int^{\Lambda} \frac{d^4p}{p^2} + \dots \approx m_{h\,0}^2 + \frac{\lambda_f^2}{8\pi^2} N_c^f \Lambda^2 + \dots$$

bare mass 1-loop correction

ultraviolet cutoff

Fine tuning: if $\Lambda \approx$ plank mass, need cancellation between bare mass and corrections across many orders of magnitude to get 125 GeV!



Hierarchy problem: SUSY's solution



stop's loop (opposite sign)

+

main "SM" term: top's loop ($\lambda_t \approx 1$)

physical m_h: 125 GeV

$$m_h^2 \approx m_{h\,0}^2 - \frac{\lambda_f^2}{8\pi^2} N_c^f \int^{\Lambda} \frac{d^4p}{p^2} + \dots \approx m_{h\,0}^2 + \frac{\lambda_f^2}{8\pi^2} N_c^f \Lambda^2$$

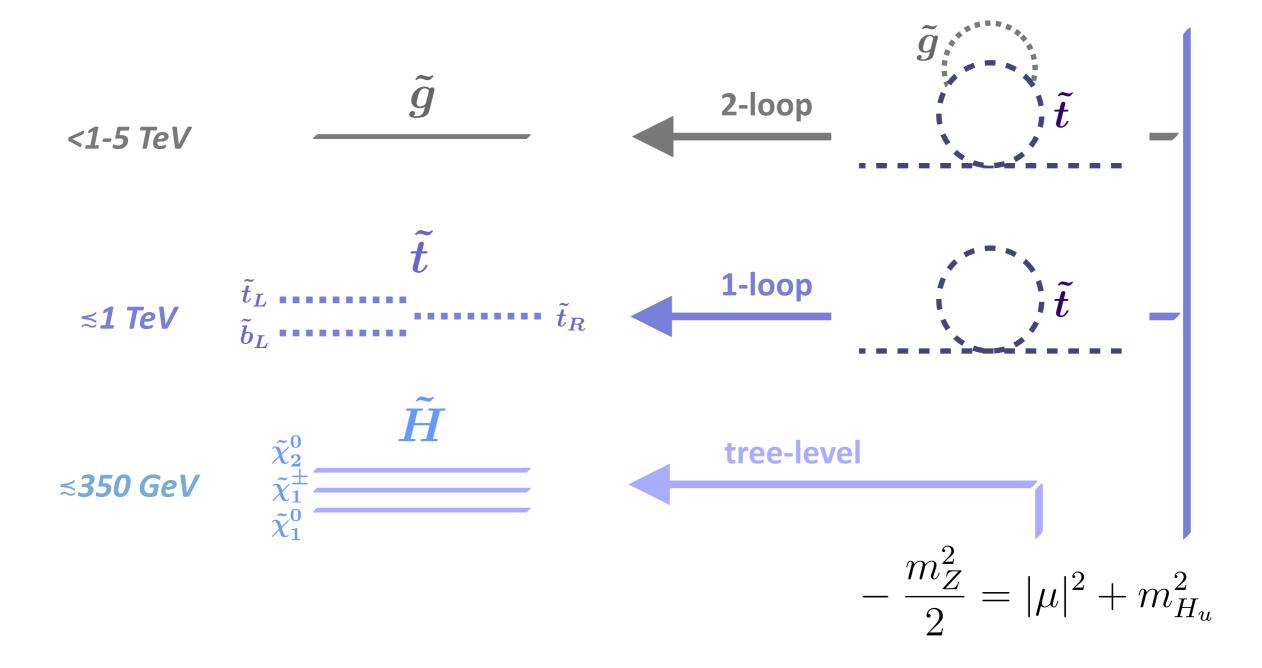
1-loop bare mass correction ultraviolet cutoff

natural cancellation (*)

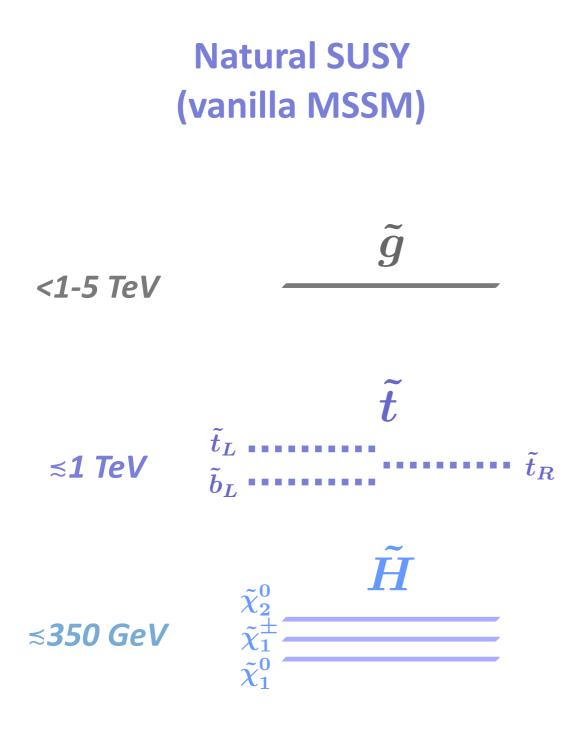
Fine tuning: if $\Lambda \approx plank mass ...$

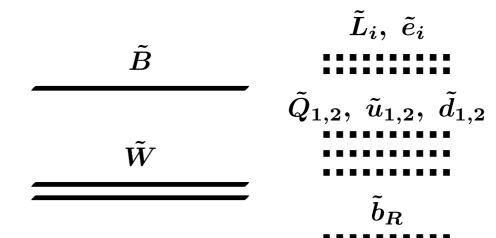
(*) provided a light stop!

Natural SUSY (vanilla MSSM)

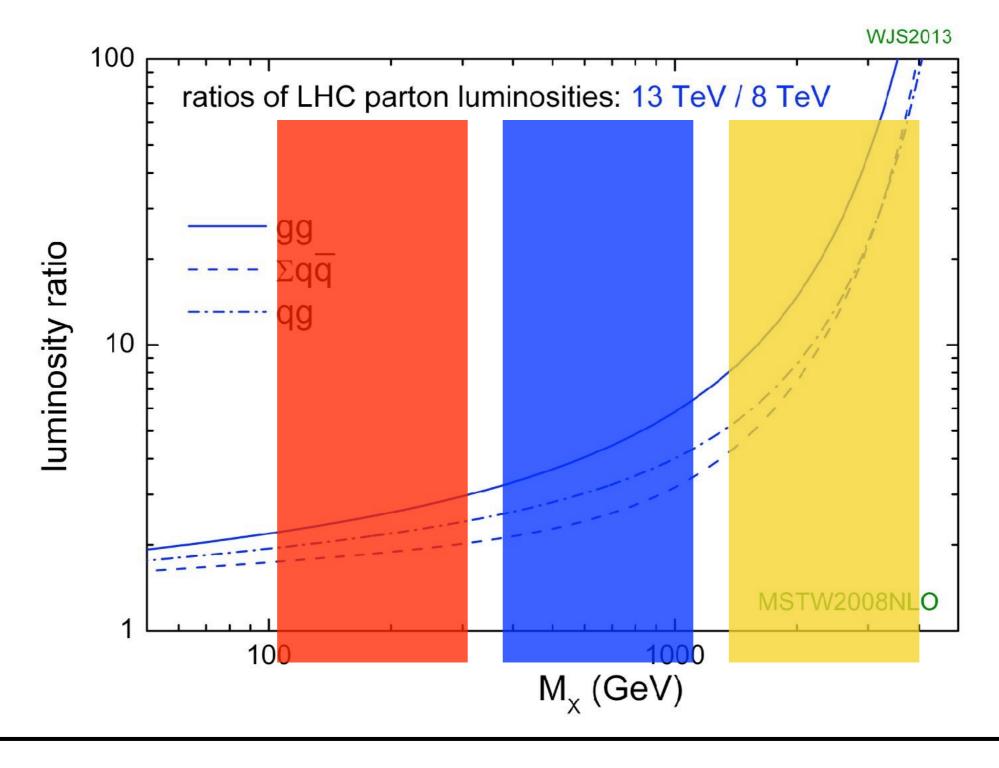


S. Zambito | *Harvard University*





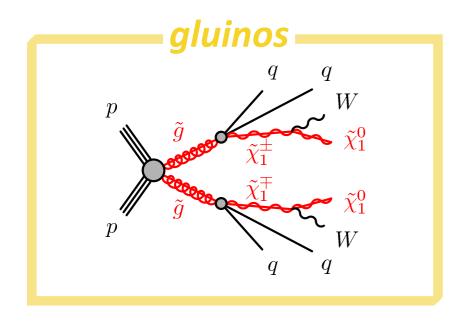
Decoupled SUSY (vanilla MSSM)

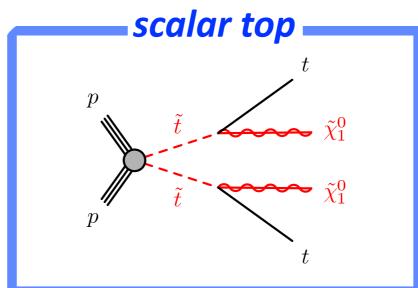


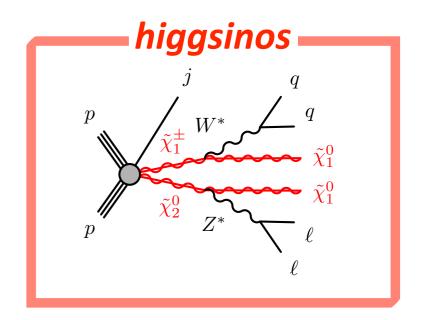
higgsinos / EWK-inos

scalar top gluinos / squarks

SUSY After $\sqrt{s}=7 \& 8$ TeV LHC



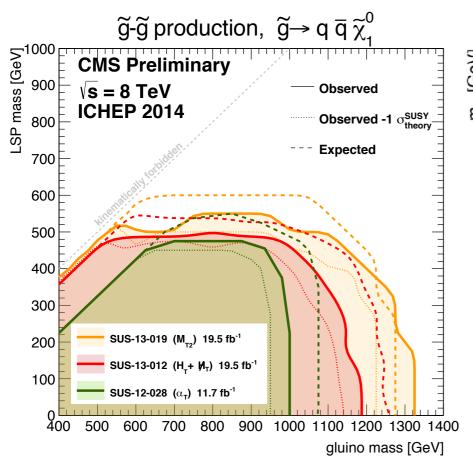


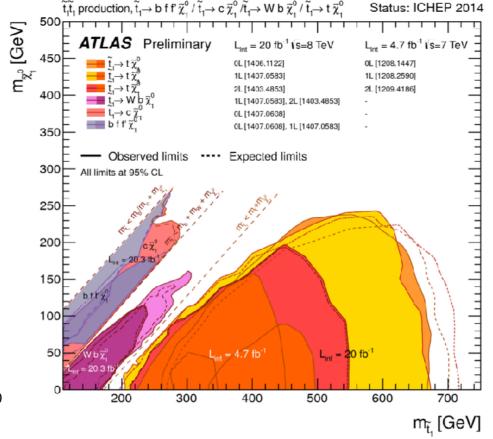


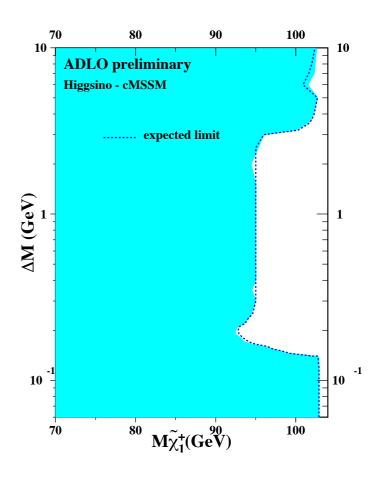
excluded m<1.3 TeV

excluded m<650 GeV

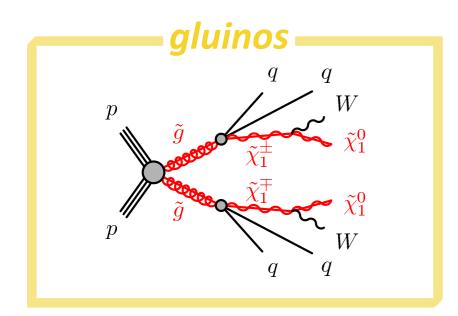
excluded m<110 GeV

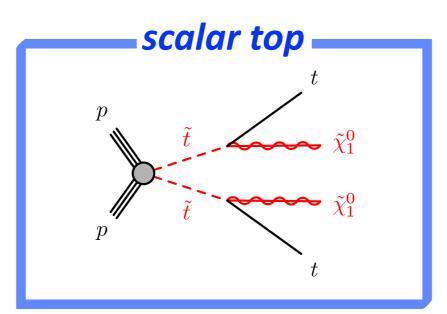


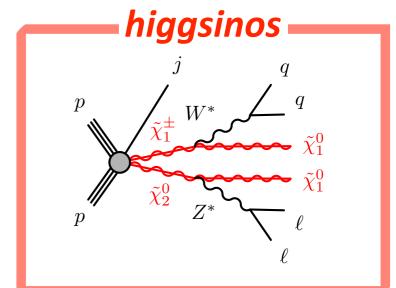




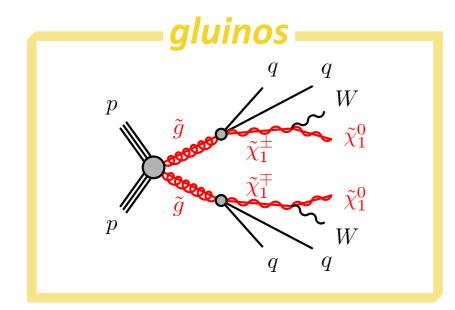
S. Zambito | *Harvard University*

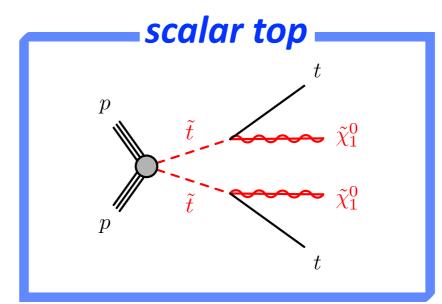


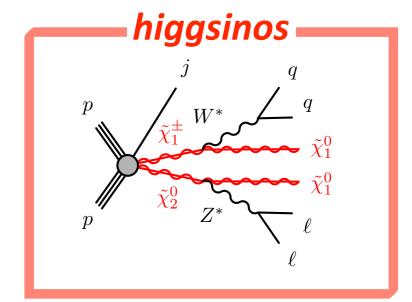




Simplified models to cope with complex, model-dependent phenomenology

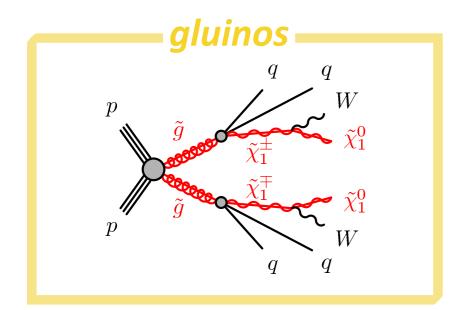


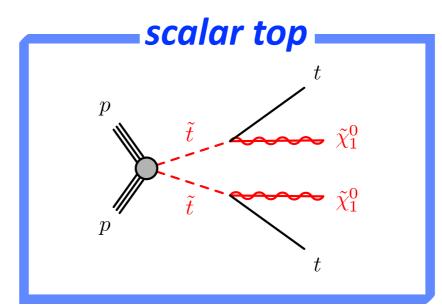


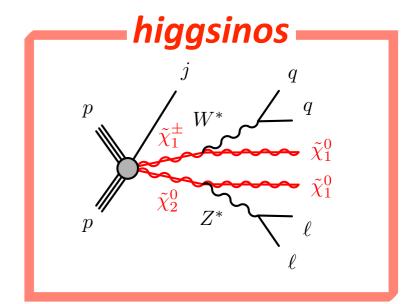


Simplified models to cope with complex, model-dependent phenomenology

branching ratios in decay vertices set to 100% (except for SM particles)







Simplified models to cope with complex, model-dependent phenomenology

branching ratios in decay vertices set to 100% (except for SM particles)

"R-parity" conservation assumed: SUSY particles produced in pairs, and lightest neutralino doesn't decay: dark matter candidate!

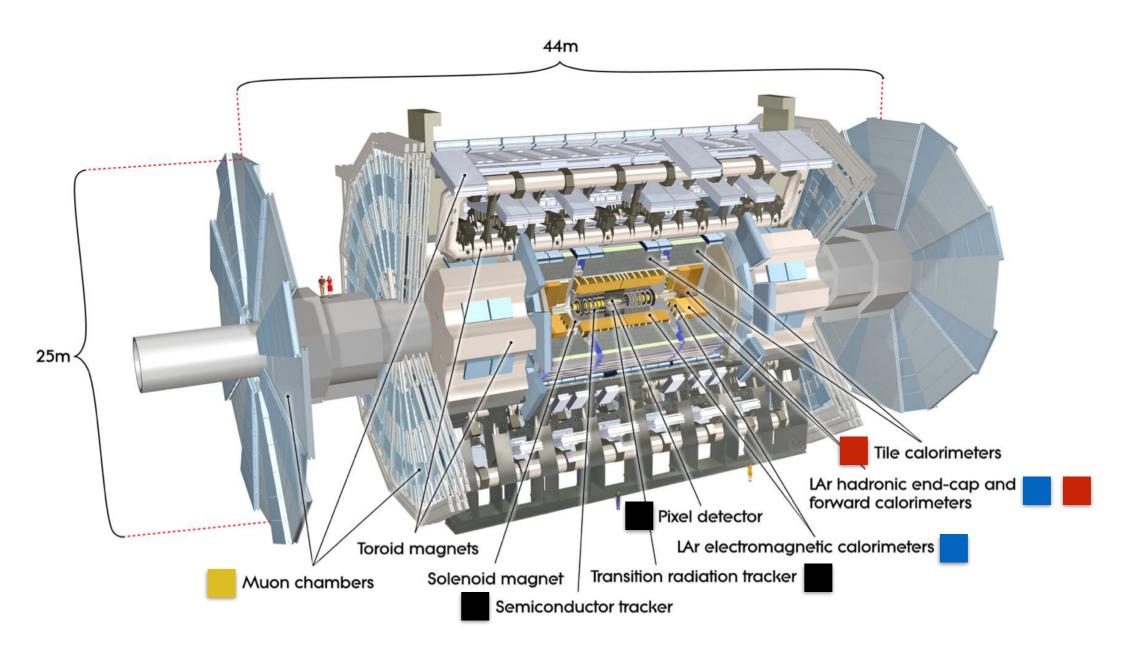
(models with "R-parity" violation not covered in this talk)

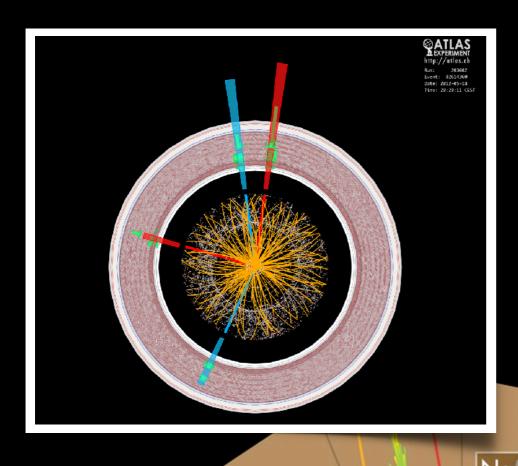
Inner detector (ID): tracks → charged particles & vertices

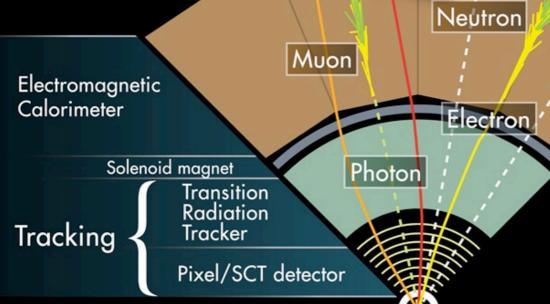
EM calorimeter: energy depositions → electrons and photons

Hadronic calorimeter: energy depositions → jets of hadrons

Muon spectrometer (MS): tracks muons

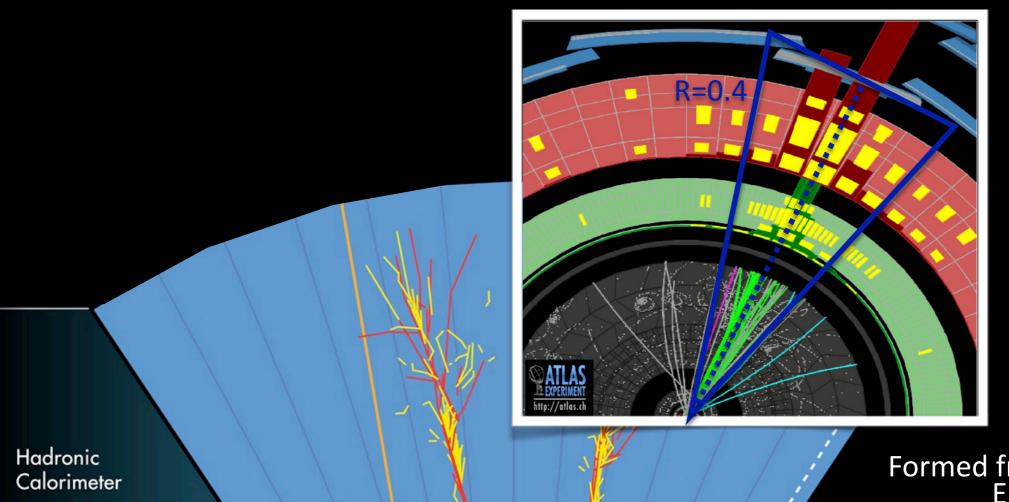






Electrons

Track in ID + E cluster in EM calorimeter Identified via shower shape + radiation in TRT Calibrated offline using standard candles (J/ Ψ ,Z)



Proton

Photon

Muon

Electromagnetic

Solenoid magnet

Transition

Radiation

Pixel/SCT detector

Tracker

Calorimeter

Tracking

Neutron

Electron

Neutrino

Jets of hadrons

Formed from energy clusters in EM + HAD calorimeters (anti- k_T R=0.4) Calibrated offline

Electrons

Track in ID + E cluster in EM calorimeter Identified via shower shape + radiation in TRT Calibrated offline using standard candles (J/ Ψ ,Z)

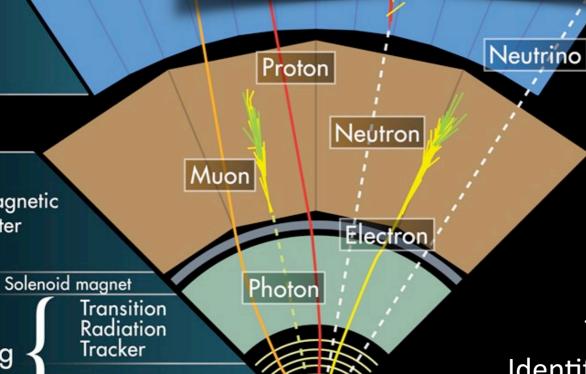
Muons

Survive calorimeters - E_{loss}≈3 GeV Mostly formed as ID +MS tracks Calibrated offline (J/Ψ,Z)

Jets of hadrons

Formed from energy clusters in EM + HAD calorimeters (anti-k_T R=0.4)

Calibrated offline



Muon

Hadronic

Calorimeter

Electromagnetic

Calorimeter

Tracking

Pixel/SCT detector

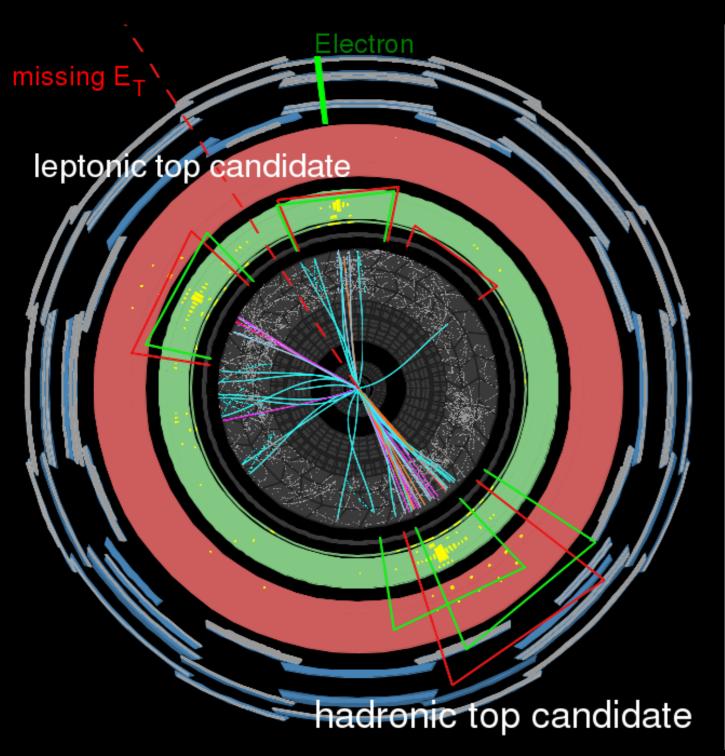
Spectrometer

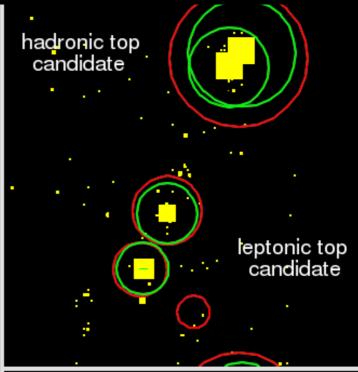
Electrons

Track in ID + E cluster in EM calorimeter Identified via shower shape + radiation in TRT Calibrated offline using standard candles (J/ Ψ ,Z)

Missing transverse energy: $\mathcal{E}_T = |\vec{p}_{T,miss}|$

negative vector sum of transverse momenta of all reconstructed & identified particles + all remaining tracks





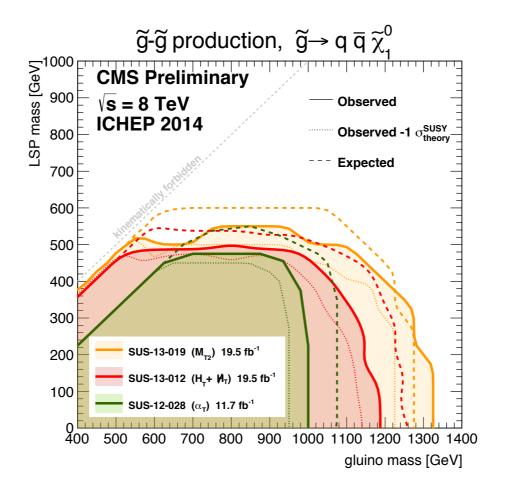


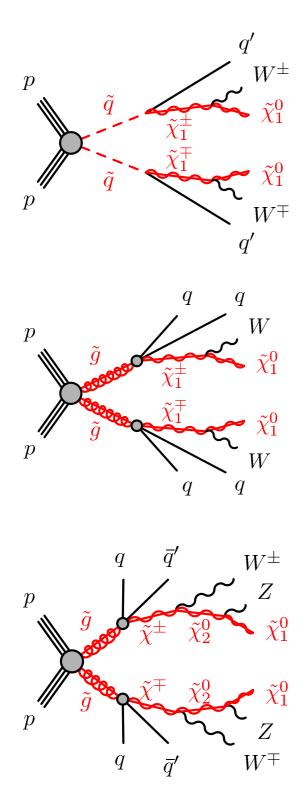
Run Number: 180144, Event Number: 43671503

Date: 2011-04-22 09:46:15 EDT

Hunting for Natural SUSY at the LHC Part I - The Gluinos

Excluded up to 1.3 TeV after the LHC run1

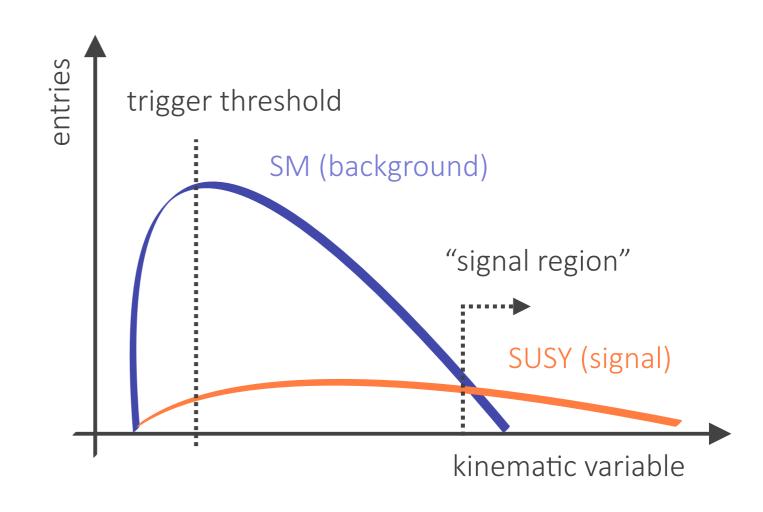




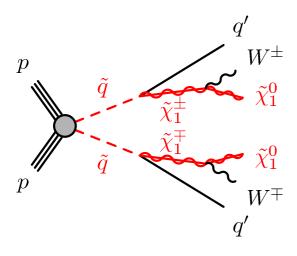
increasing number of expected jets

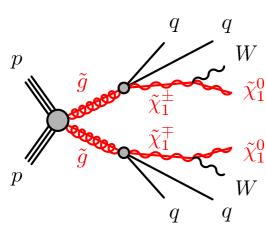
Select interesting events:

one electron or muon from $W \rightarrow \ell \nu$ decay many jets from gluinos/squarks decay chain large $\not\vdash$ from undetected neutralinos

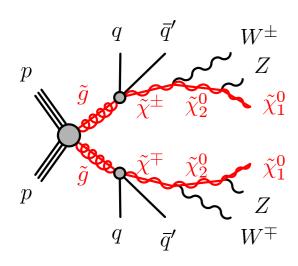


Hunting Gluinos & Squarks





ncreasing number of expected jets

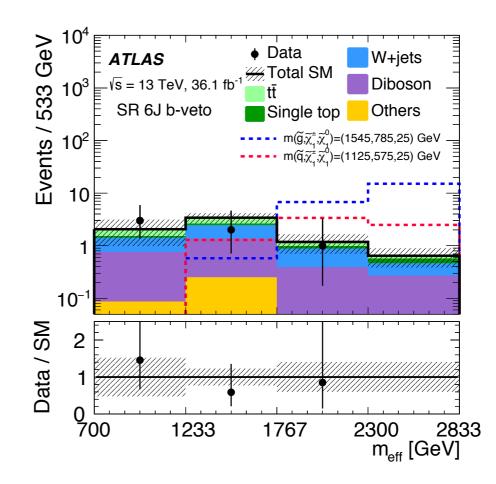


Select interesting events:

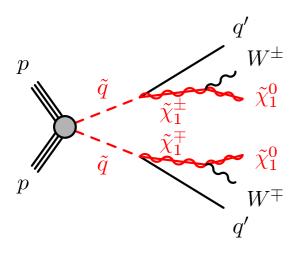
one electron or muon from $W \rightarrow \ell \nu$ decay many jets from gluinos/squarks decay chain large $\not\vdash$ from undetected neutralinos

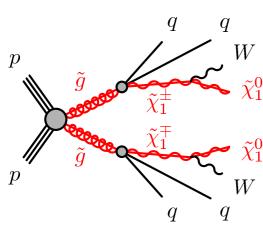
effective mass
$$m_{\rm eff}^{incl} = H_T + E_{\rm T}^{\rm miss} = \sum p_T^{\ell} + \sum p_T^{jet} + E_{\rm T}^{\rm miss}$$

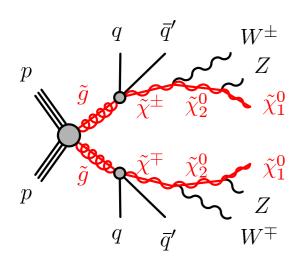
sensitive to SUSY mass scale



Hunting Gluinos & Squarks







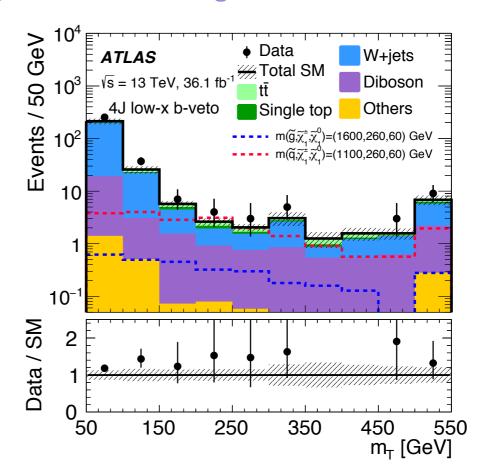
ncreasing number of expected jets

Select interesting events:

one electron or muon from $W \rightarrow \ell \nu$ decay many jets from gluinos/squarks decay chain large $\not\vdash$ from undetected neutralinos

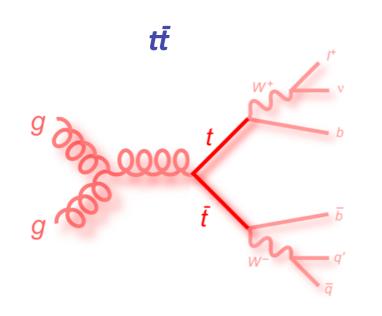
transverse mass
$$m_{\rm T} = \sqrt{2 \cdot p_T^{\ell} \cdot E_{\rm T}^{\rm miss} \cdot (1 - cos(\Delta \phi(\ell, E_{\rm T}^{\rm miss})))}$$

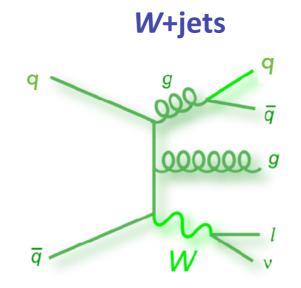
edge at mw for backgrounds with on-shell W

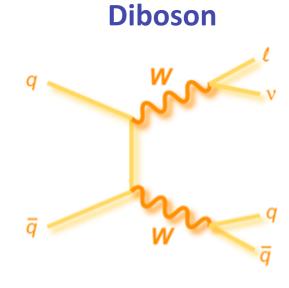


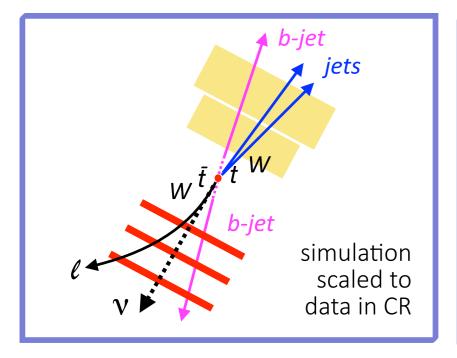
Experimental signature involves large £, many jets and 1 isolated lepton

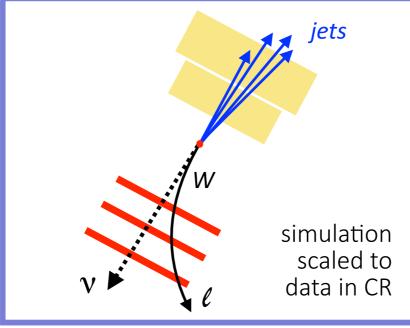
Control (CR) and validation (VR) regions used to extract / x-check background predictions

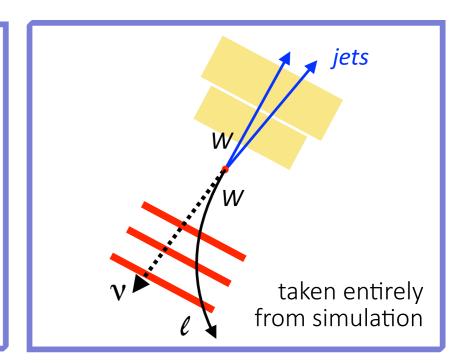






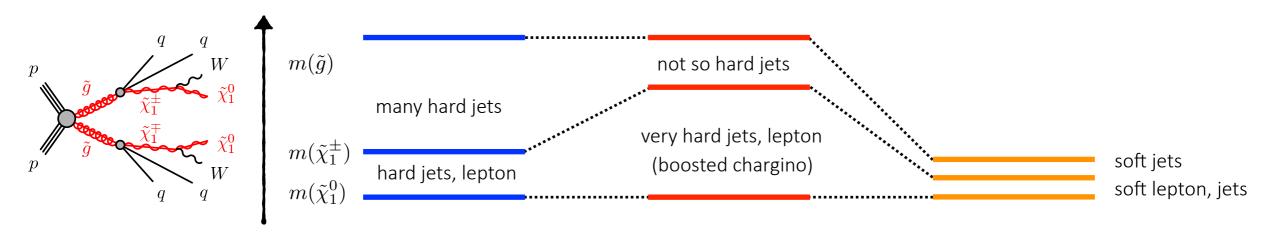


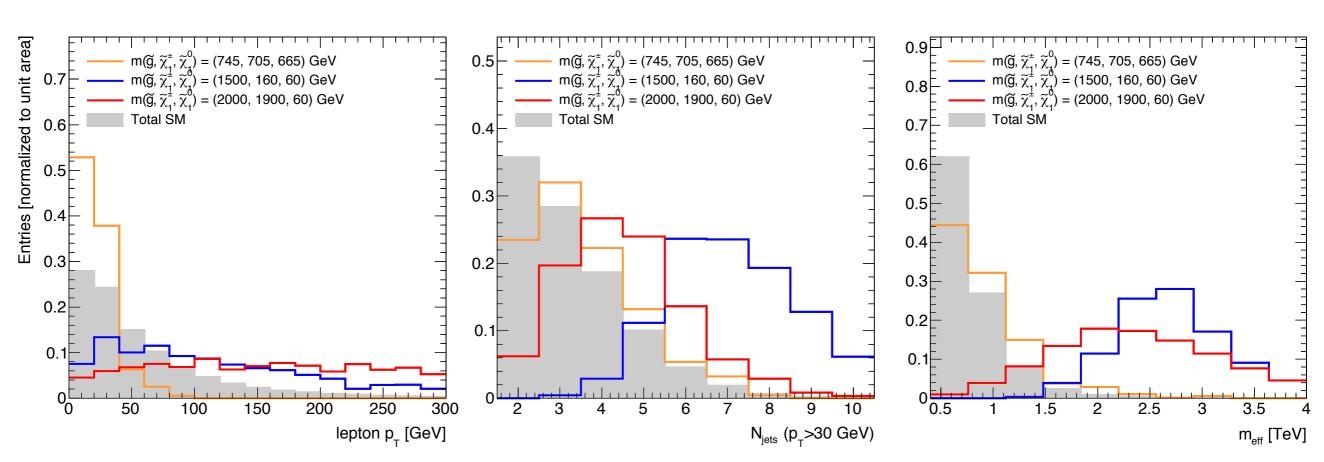




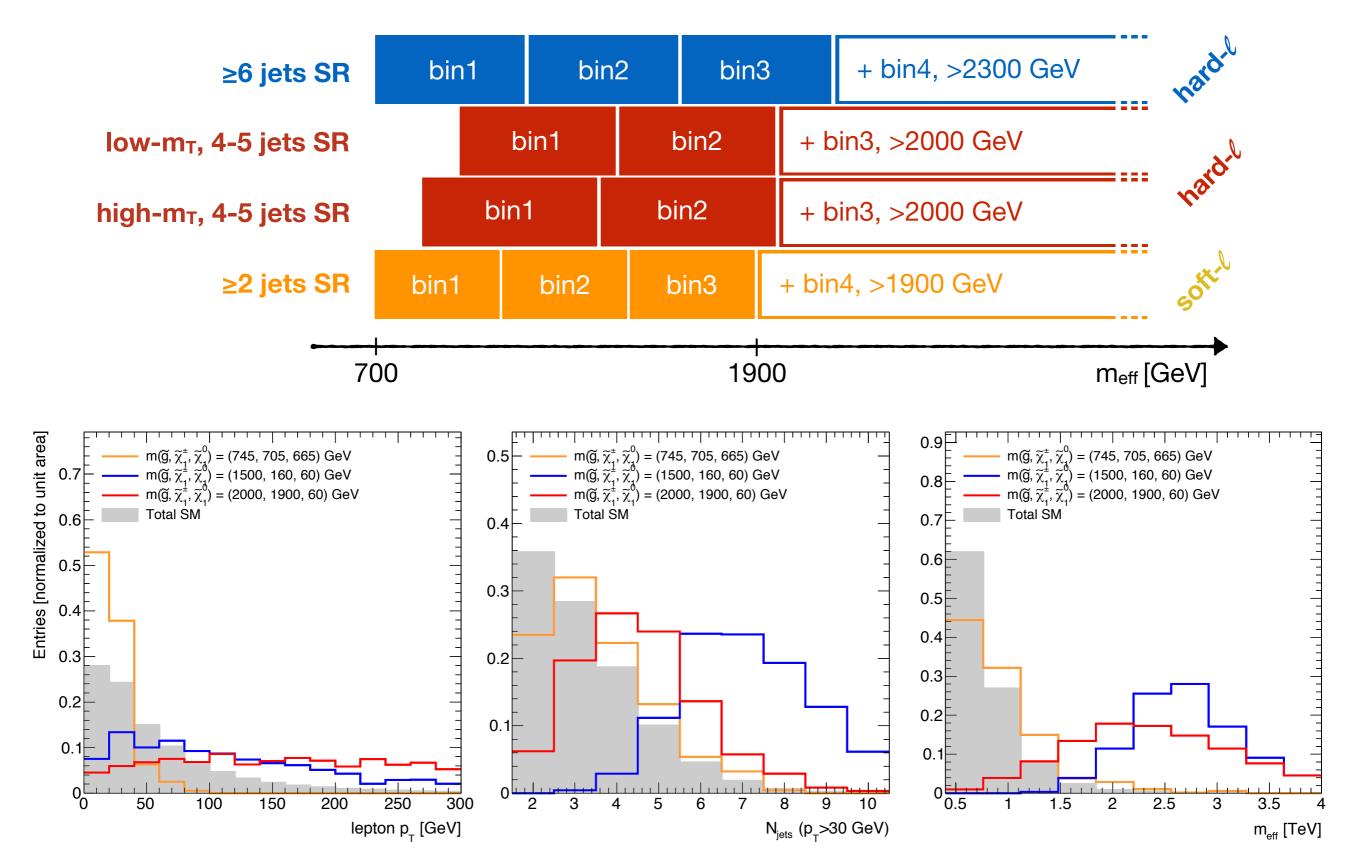
The challenge: signal kinematics strongly depend on sparticles' masses!

targeting whole parameter space in one analysis is very complicated...





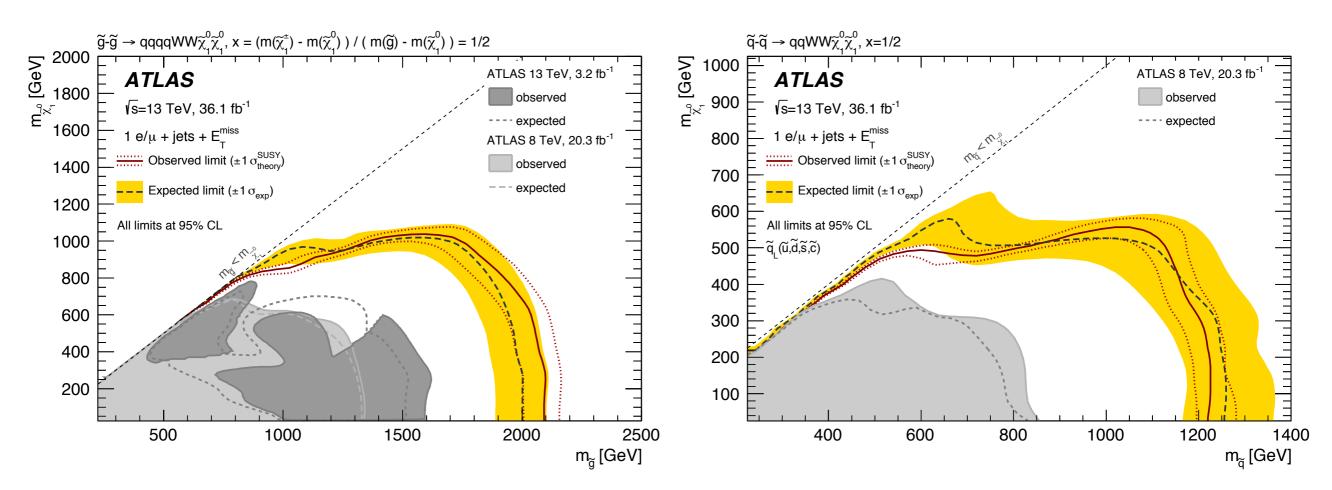
Signal Selection: Multi-Bin Fit



Observed event yields (simplified SR for model-independent limits)

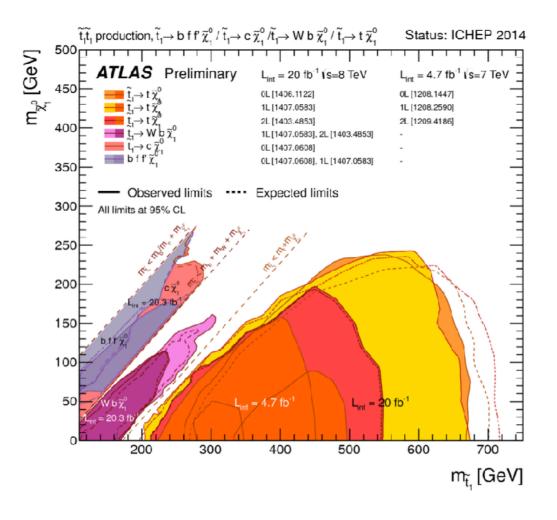
SR _{disc}	2J	4J high-x	4J low-x (gluino)	4J low-x (squark)	6J (gluino)	6J (squark)	9J
Observed events	80	16	24	50	0	28	4
Fitted bkg events	67 ± 6	17.7 ± 2.7	17.2 ± 3.2	47 ± 7	2.6 ± 0.6	23.4 ± 3.1	3.1 ± 1.6
$S_{\rm exp}^{95}$	$21.6^{+9.2}_{-5.6}$	$10.8^{+3.7}_{-3.0}$	$11.8^{+4.8}_{-2.7}$	$19.9^{+7.5}_{-5.6}$	$4.5^{+1.8}_{-1.0}$	$12.7^{+5.0}_{-4.0}$	$6.0^{+2.2}_{-1.2}$
p(s=0)	0.10	0.50	0.10	0.35	0.50	0.21	0.34

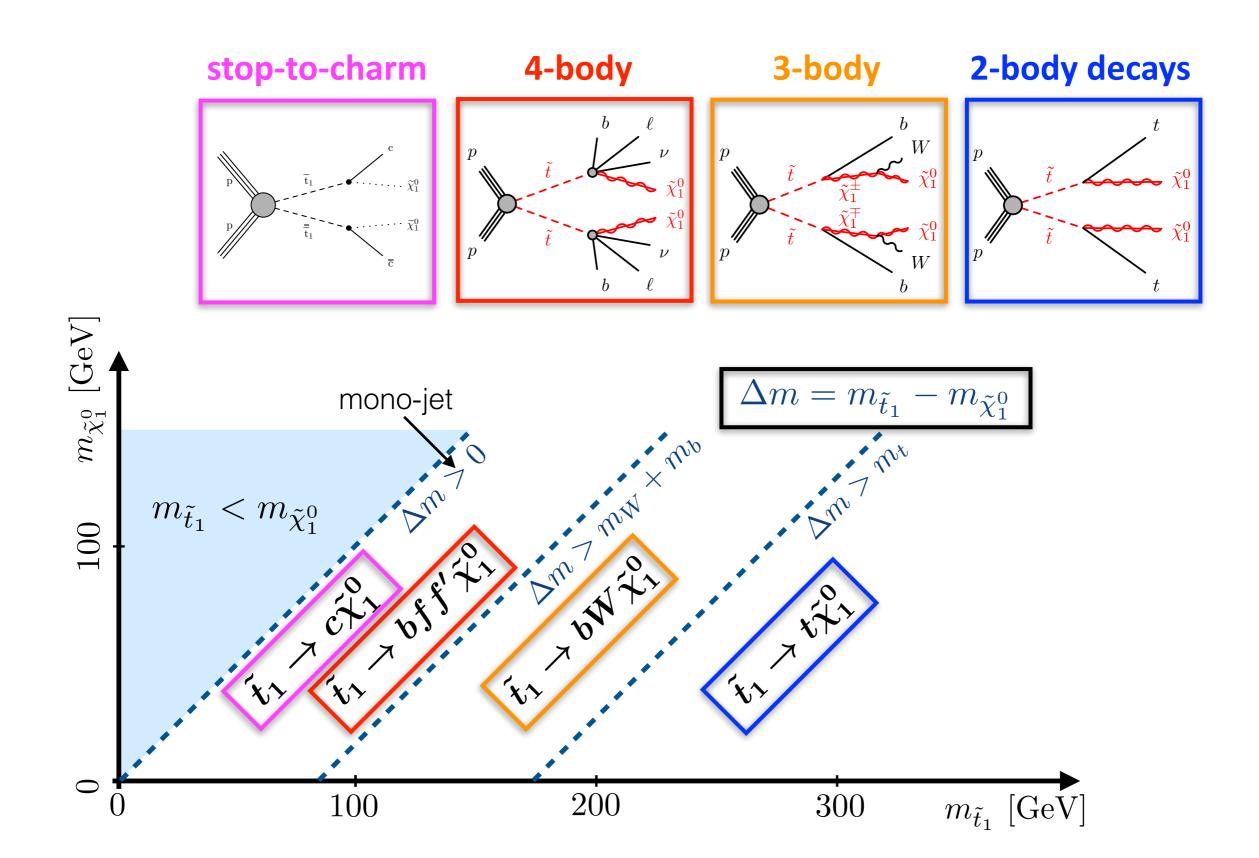
Exclusions: simulated SUSY signal rejected, or not, via fit to pseudo/observed data



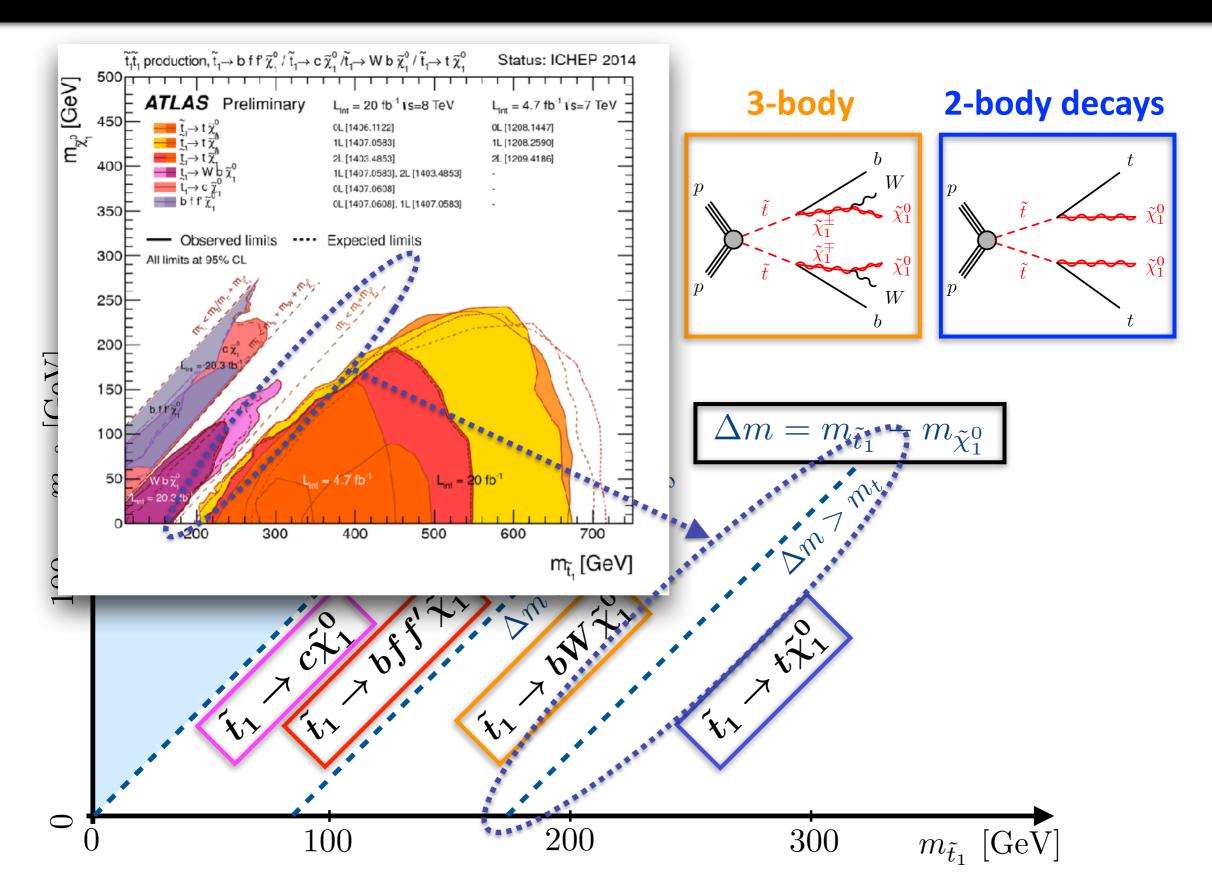
Hunting for Natural SUSY at the LHC Part II - The Stops

Excluded up to 650 GeV after the LHC run1





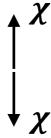
Hunting Stops: Decays



Hunting Diagonal Stops

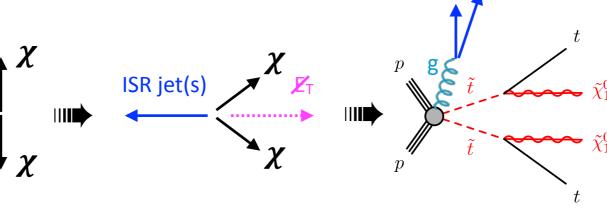
$m_{ ilde{t}}-m_{ ilde{\chi}_1^0}pprox m_t: ilde{t} ilde{t}$ kinematics close to $tar{t}$

need ISR activity to "misalign" $\chi\chi$ and get tangible contribution to \not{E}_{T} contribution to E_T

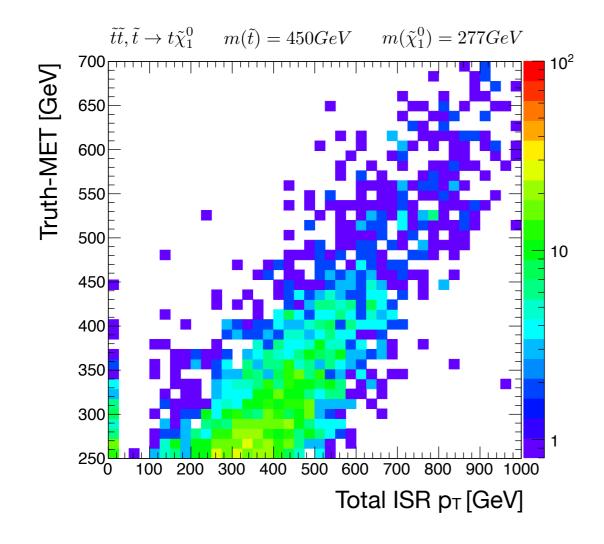




need ISR activity to "misalign" $\chi\chi$ and get tangible contribution to \mathcal{E}_{T}



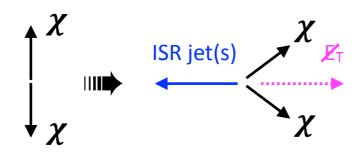
ISR jet(s)

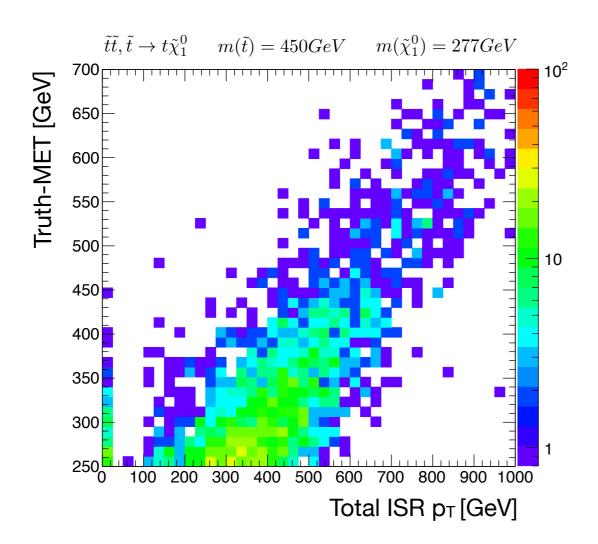


Hunting Diagonal Stops

$m_{ ilde{t}}-m_{ ilde{\chi}_1^0}pprox m_t: ilde{t} ilde{t}$ kinematics close to $tar{t}$

need ISR activity to "misalign" $\chi\chi$ and get tangible contribution to \not{E}_{T}





plethora of pheno. papers:

$$p_{T,ISR} \approx -(p_T(\tilde{t}_1) + p_T(\tilde{t}_2))$$

$$\frac{\mathbb{Z}_T}{p_{T,ISR}} \approx \frac{m(\tilde{\chi}_1^0)}{m(\tilde{t})}$$

K. Hagiwara et al, 2015: arXiv:1307.1553v3

H. An et al, 2015: arXiv:1506.00653v2

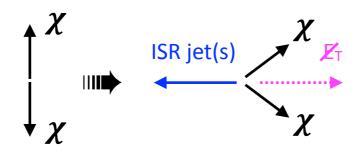
S. Macaluso et al, 2015: arXiv:1506.07885.

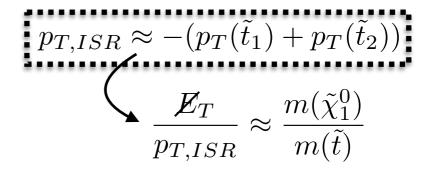
H. Cheng et al, 2016: arXiv:1604.00007v1

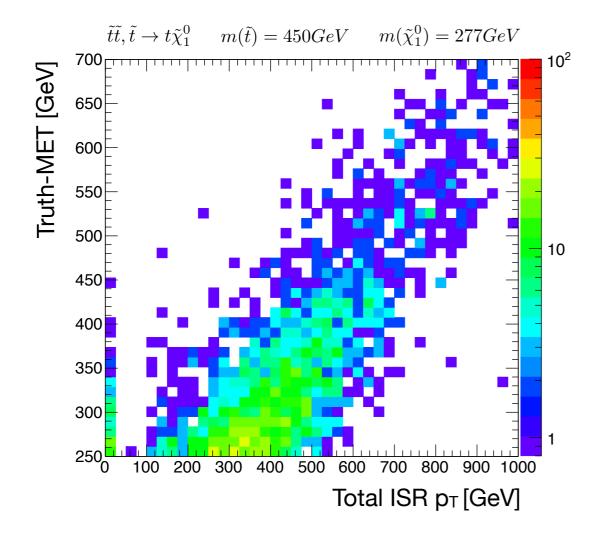
etc...

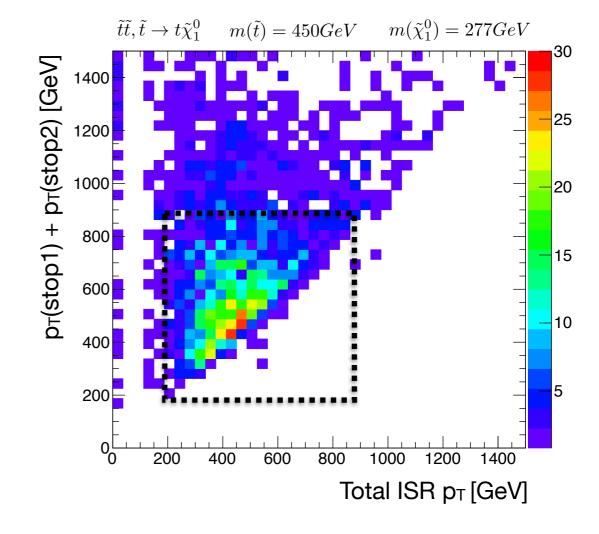
$m_{ ilde{t}}-m_{ ilde{\chi}_1^0}pprox m_t: ilde{t} ilde{t}$ kinematics close to $tar{t}$

need ISR activity to "misalign" $\chi\chi$ and get tangible contribution to \not{E}_{T}

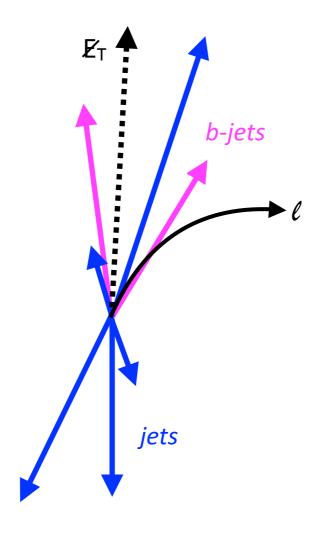








easier said than done... among the many jets, which ones are from ISR?

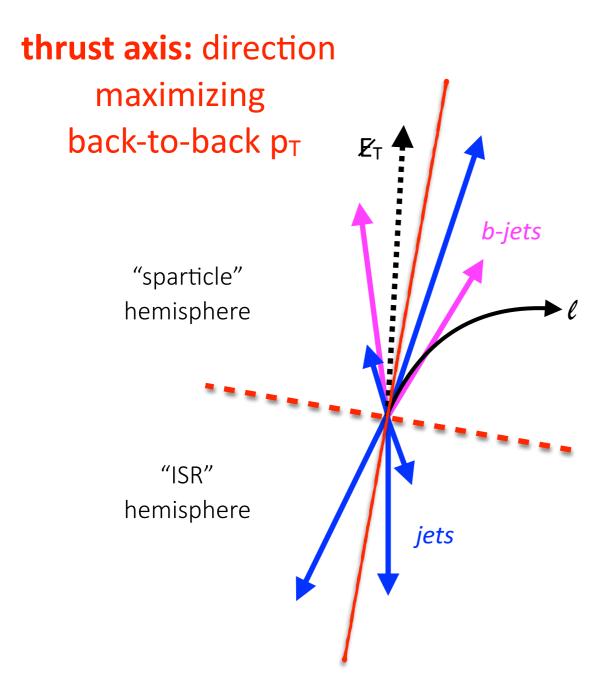


$$p_{T,ISR} \approx -(p_T(\tilde{t}_1) + p_T(\tilde{t}_2))$$

$$\frac{E_T}{p_{T,ISR}} \approx \frac{m(\tilde{\chi}_1^0)}{m(\tilde{t})}$$

Thrust-Based ISR Identification

easier said than done... among the many jets, which ones are from ISR?

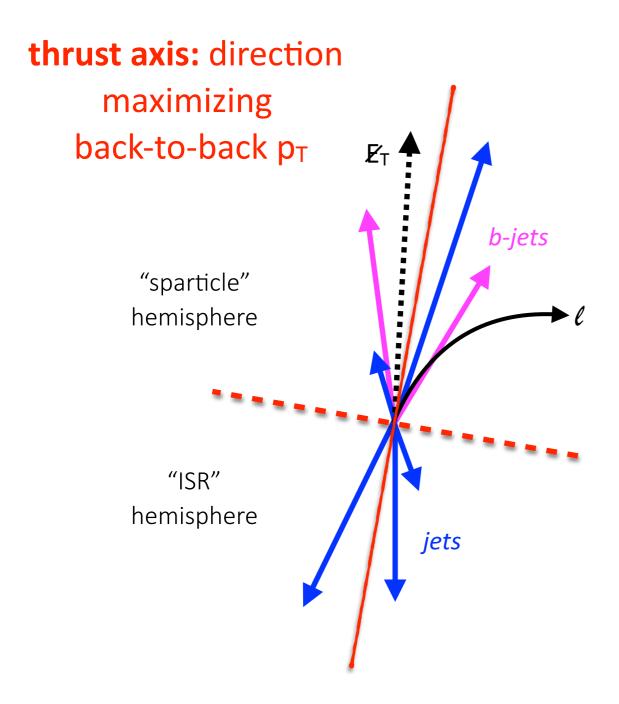


$$p_{T,ISR} \approx -(p_T(\tilde{t}_1) + p_T(\tilde{t}_2))$$

$$\frac{\mathbb{Z}_T}{p_{T,ISR}} \approx \frac{m(\tilde{\chi}_1^0)}{m(\tilde{t})}$$

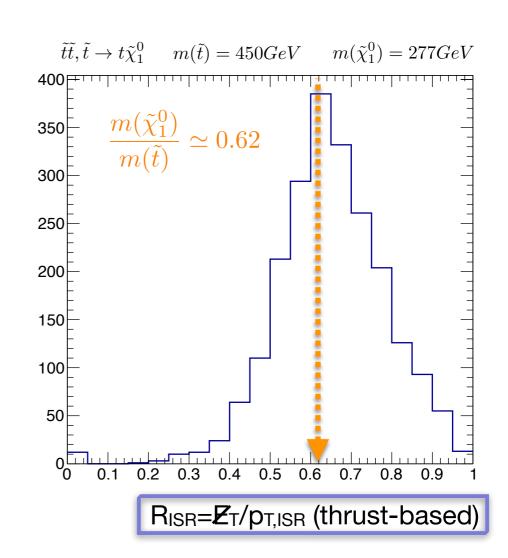
Thrust-Based ISR Identification

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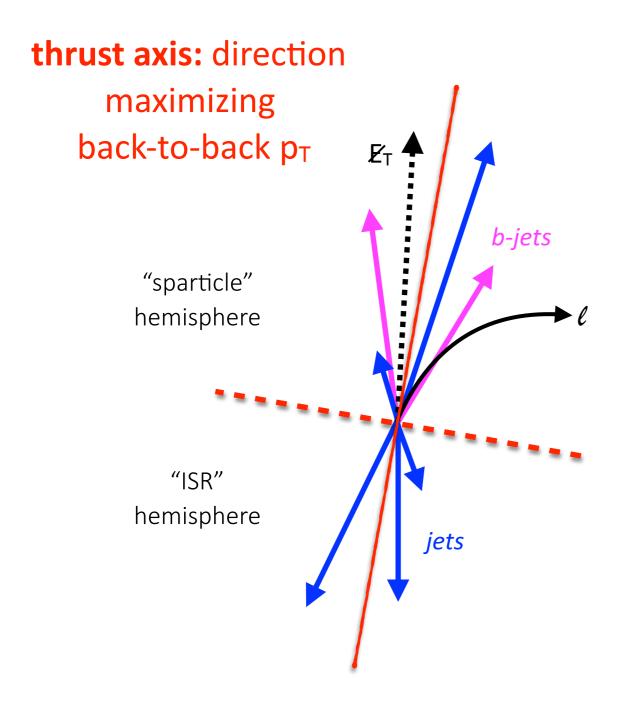
$$p_{T,ISR} \approx -(p_T(\tilde{t}_1) + p_T(\tilde{t}_2))$$

$$\frac{\mathbb{Z}_T}{p_{T,ISR}} \approx \frac{m(\tilde{\chi}_1^0)}{m(\tilde{t})}$$



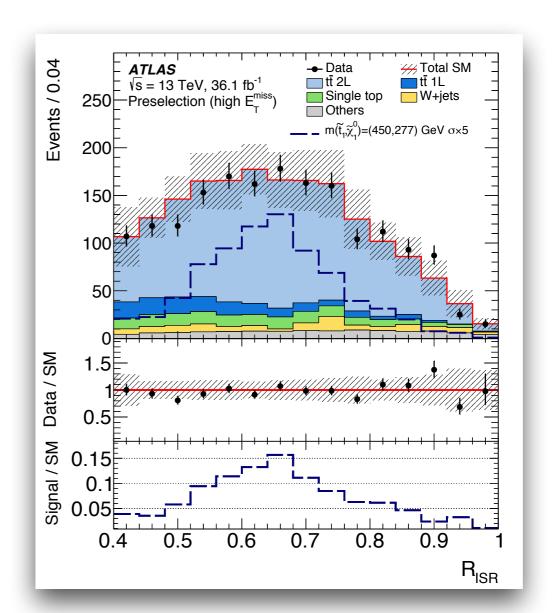
Thrust-Based ISR Identification

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$$p_{T,ISR} \approx -(p_T(\tilde{t}_1) + p_T(\tilde{t}_2))$$

$$\frac{\mathbb{Z}_T}{p_{T,ISR}} \approx \frac{m(\tilde{\chi}_1^0)}{m(\tilde{t})}$$



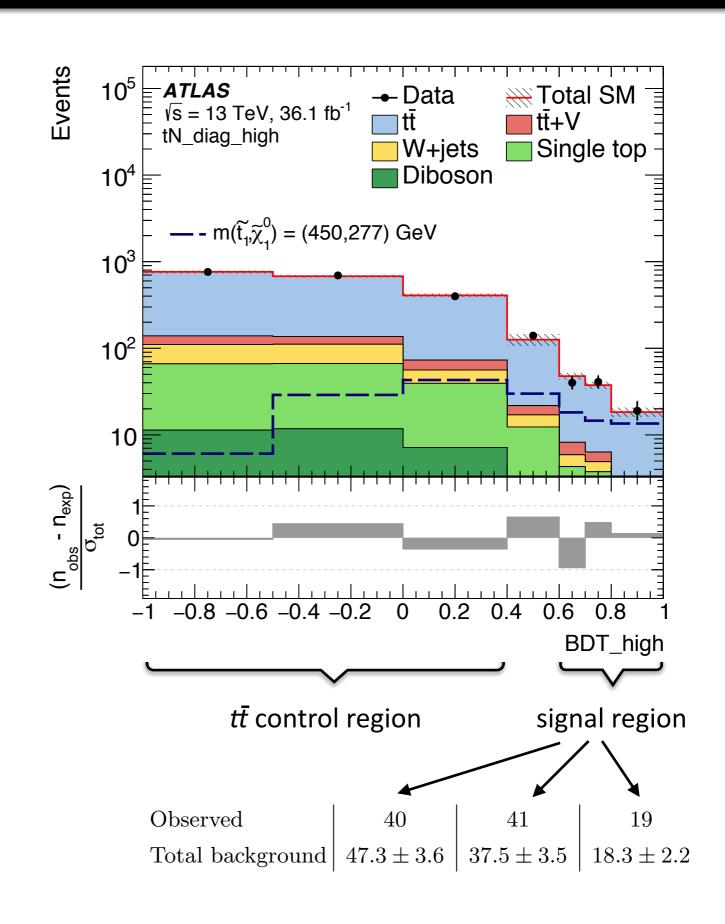
Signal Selection: Boosted Decision Tree (BDT)

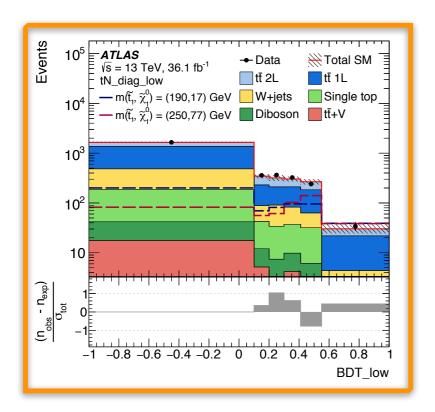
SR based on BDT discriminant trained to select stop and reject $t\bar{t}$

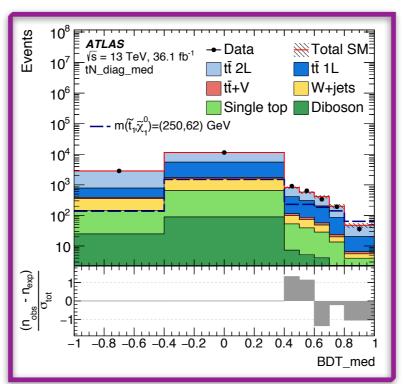
10 variables, some computed using thrust-based ISR reconstruction

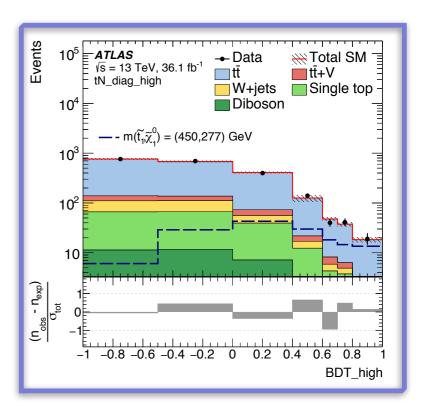
only well-modeled and understood variables enter the BDT

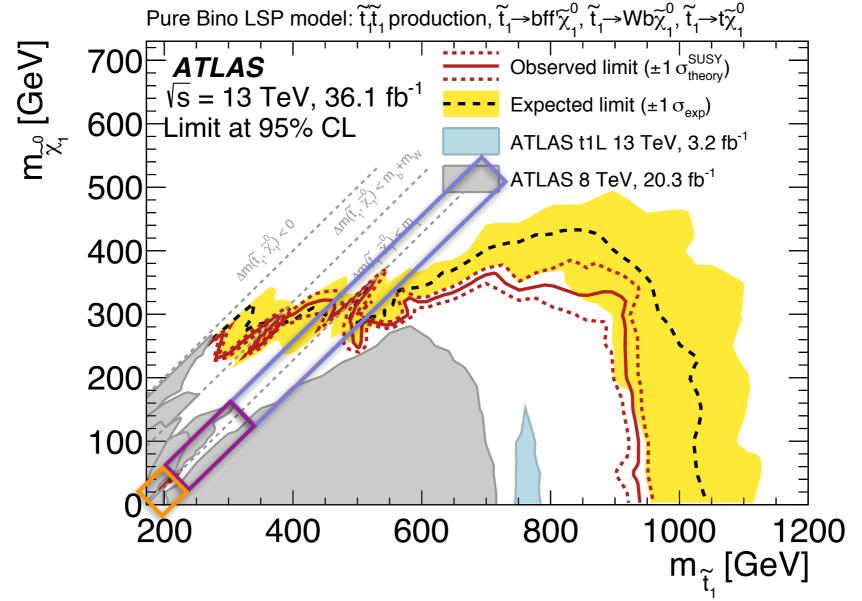
lots of validation performed to make sure the exploited correlations are understood





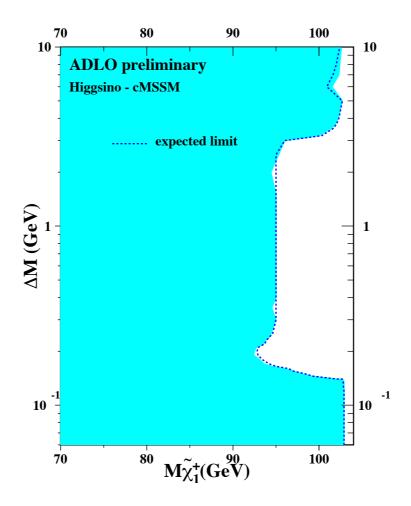






Hunting for Natural SUSY at the LHC Part III - The Higgsinos

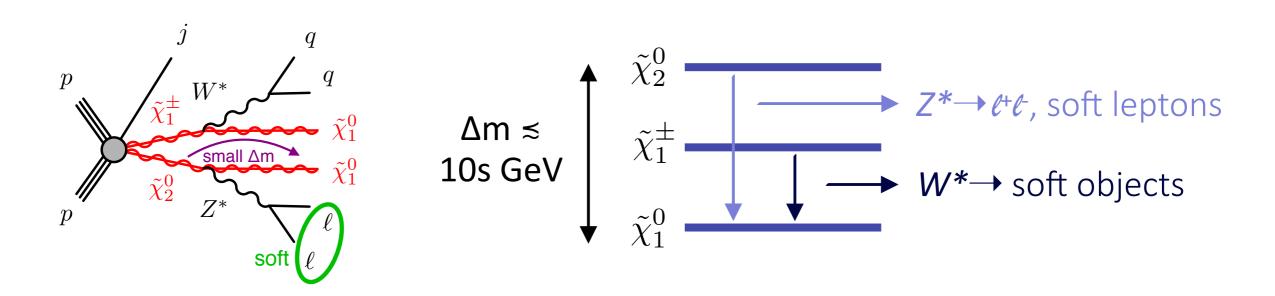
Excluded up to 110 GeV at LEP



As for "diagonal" stop, sensitivity driven by ISR-induced MET

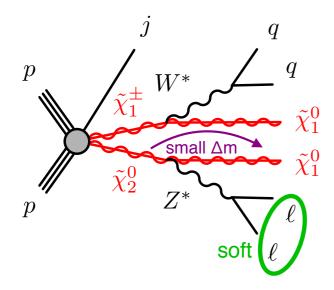
Higgsinos mix with other EWKinos:

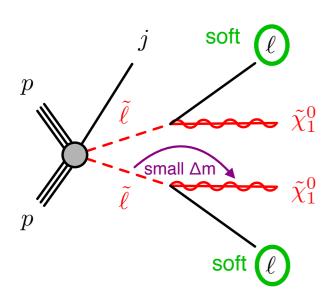
multiplets of neutralinos and charginos



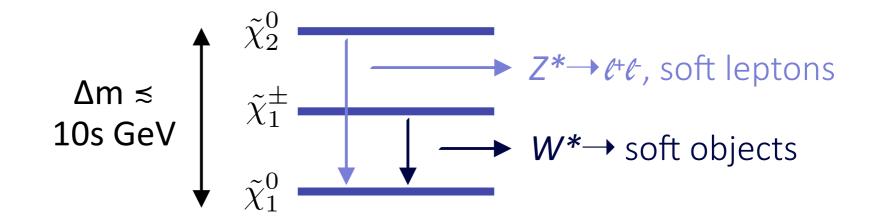
Main analysis challenge: soft leptons intra-higgsinos decay products often too soft to be reconstructed

As for "diagonal" stop, sensitivity driven by ISR-induced MET





Higgsinos mix with other EWKinos: multiplets of neutralinos and charginos

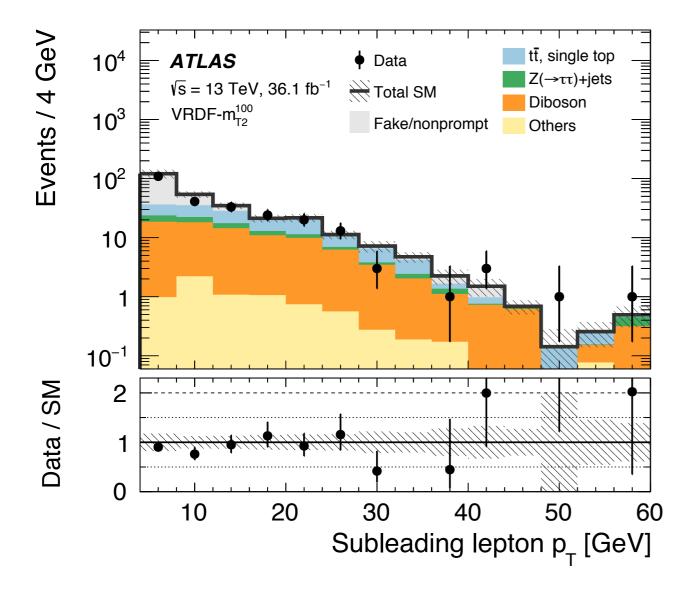


Main analysis challenge: soft leptons intra-higgsinos decay products often too soft to be reconstructed

Low-p_T Leptons: Performance

Sensitivity impacted by soft ℓ **performance**

key to signal acceptance, background rejection (lepton mis-ID)

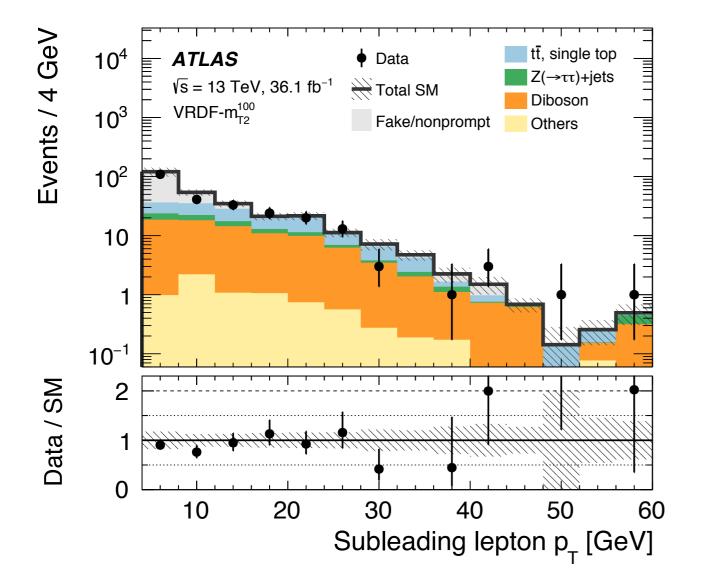


also connected to largest uncertainty!

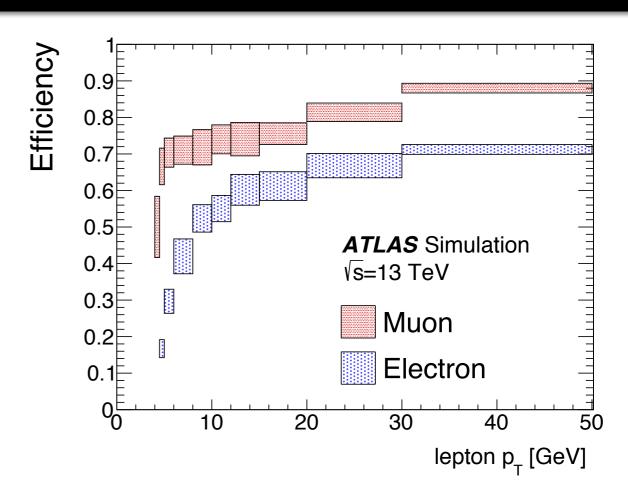
Low-p_T Leptons: Performance

Sensitivity impacted by soft ℓ performance

key to signal acceptance, background rejection (lepton mis-ID)



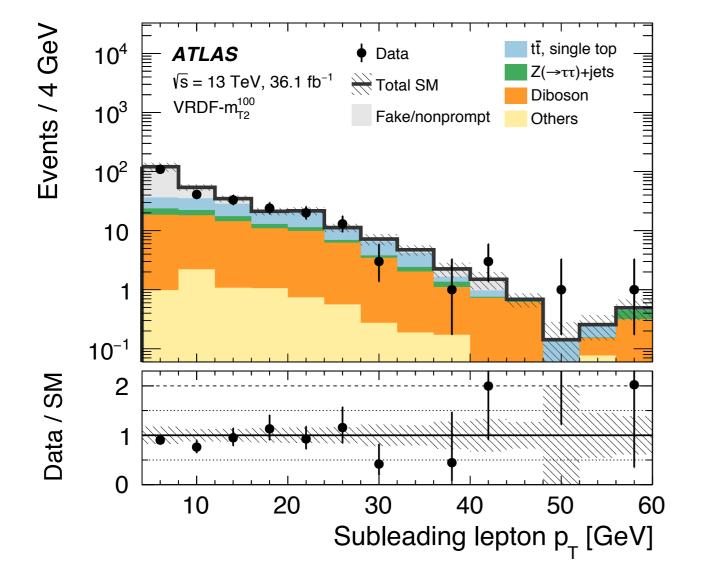
also connected to largest uncertainty!



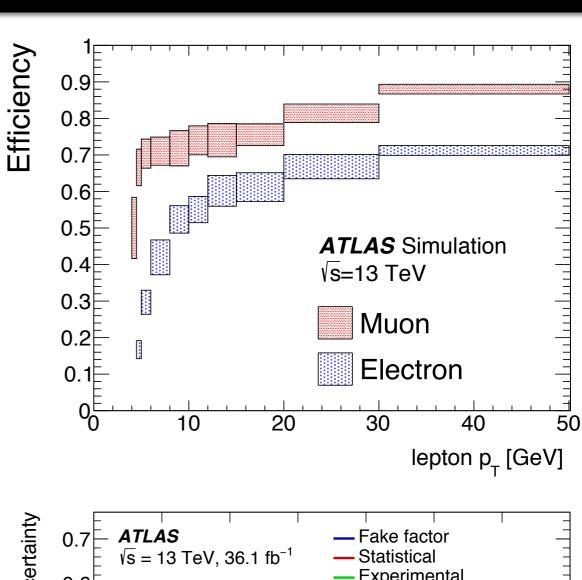
Low-p_T Leptons: Performance

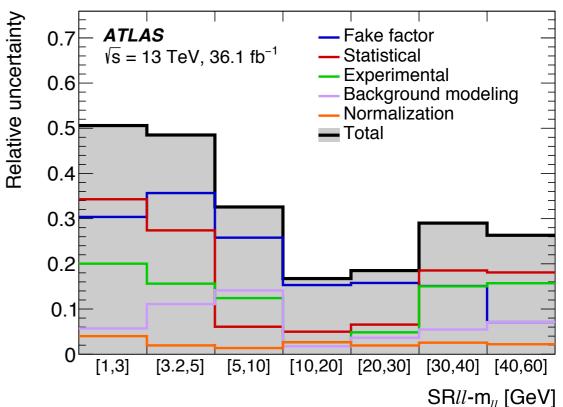
Sensitivity impacted by soft ℓ performance

key to signal acceptance, background rejection (lepton mis-ID)



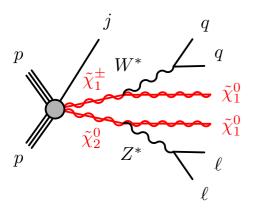
also connected to largest uncertainty!



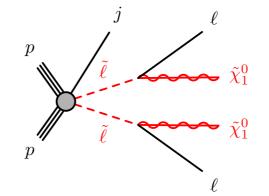


Key variables: $m_{\ell\ell}$ (higgsinos) or m_{T2} (sleptons) and $\mathcal{L}_T/H_{T,lep}$

Backgrounds tend to have lower 上 and harder leptons ➡ smaller 上 / H_{T,lep}



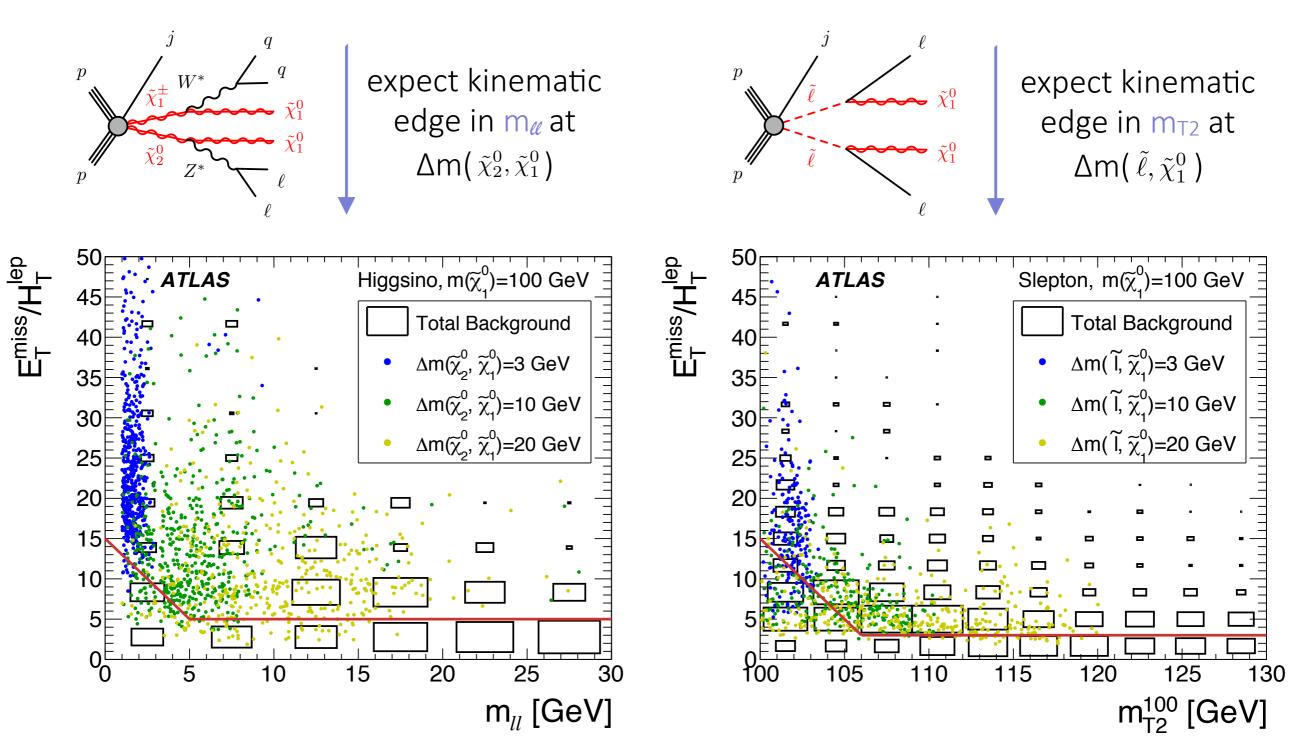
expect kinematic edge in m_{ℓ} at $\Delta m(\tilde{\chi}_{2}^{0}, \tilde{\chi}_{1}^{0})$



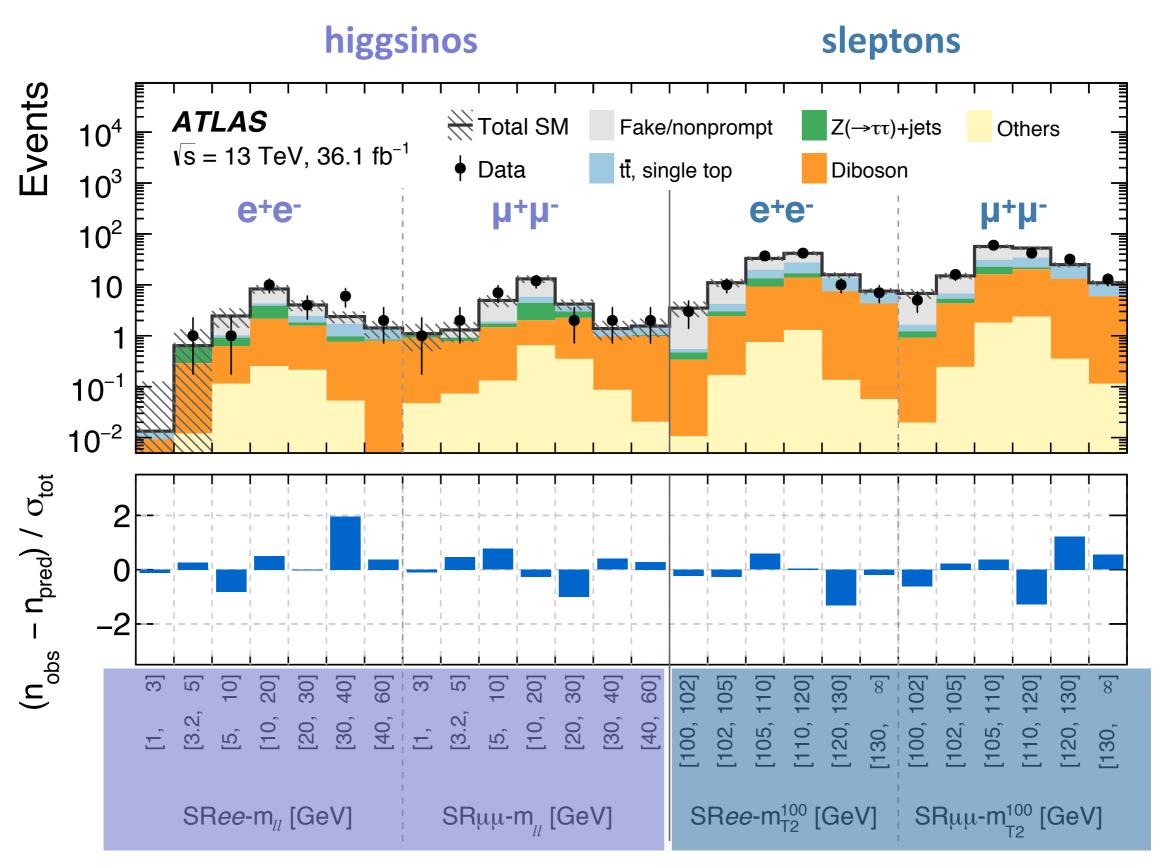
expect kinematic edge in m_{T2} at $\Delta m(\tilde{\ell}, \tilde{\chi}_1^0)$

Key variables: $m_{\ell\ell}$ (higgsinos) or m_{T2} (sleptons) and $\mathcal{L}_T/H_{T,lep}$

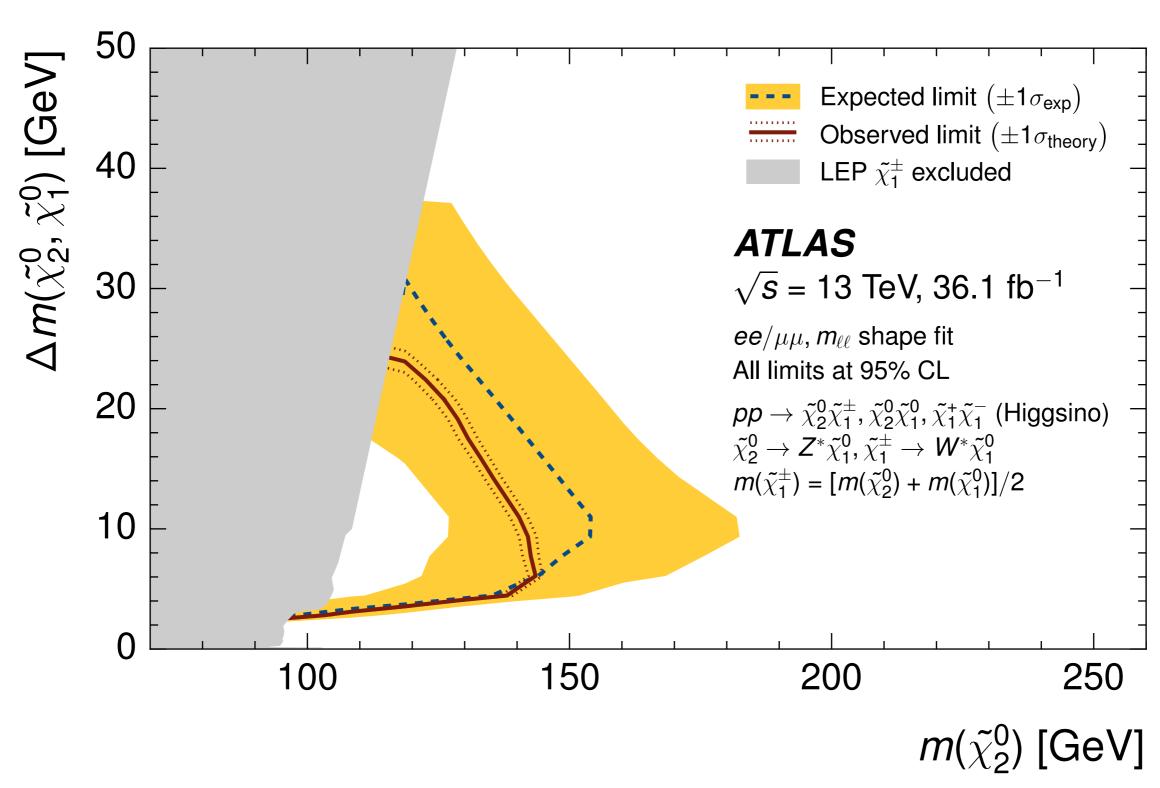
Backgrounds tend to have lower 上 and harder leptons ➡ smaller 上 / H_{T,lep}



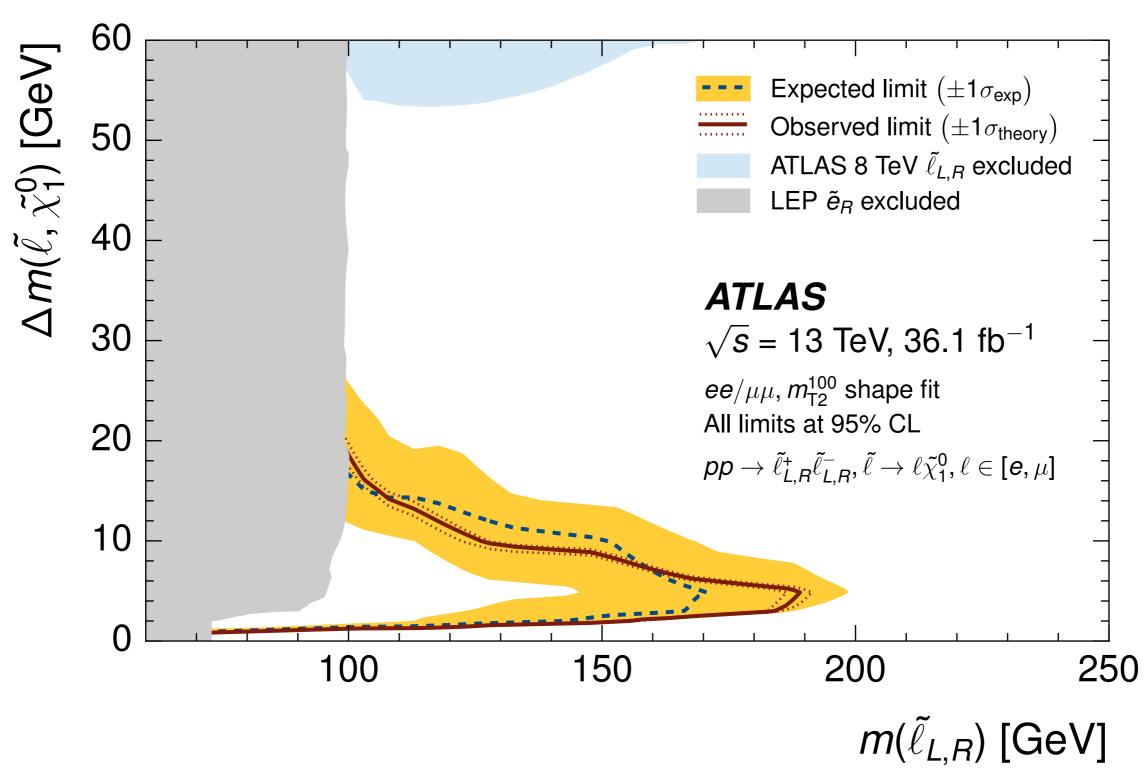
Results: Observed Event Yields



Exclusion limits on higgsinos exploiting multi-bin fit on ma



Exclusion limits on sleptons exploiting multi-bin fit on m_{T2}



A Bit Of Fortune-Telling... Part IV - What next?



ATLAS Preliminary

ATLAS-CONF-2017-039

1501.07110

1405.5086

1507.05493

LHC: Stringent Limits, No New Physics...

ATLAS SUSY Searches* - 95% CL Lower Limits

2-3 e, μ

 e, μ, γ

 $4e, \mu$

0-2 iets

0-2h

Yes

Yes

Yes

Yes

36 1

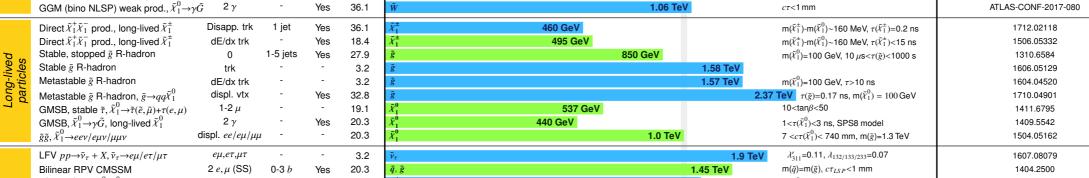
20.3

20.3

20.3

 \tilde{W}

\sqrt{s} = 7, 8, 13 TeV December 2017 e, μ, τ, γ Jets $E_{\mathrm{T}}^{\mathrm{miss}} \int \mathcal{L} \, dt [\mathrm{fb}^{-1}]$ Reference Model **Mass limit** \sqrt{s} = 13 TeV \sqrt{s} = 7, 8 TeV 1712.02332 2-6 jets Yes 36.1 $m(\tilde{\chi}_1^0)$ <200 GeV, $m(1^{st} \text{ gen. } \tilde{q})$ = $m(2^{nd} \text{ gen. } \tilde{q})$ $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ 0 710 GeV $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed) mono-iet 1-3 jets Yes 36.1 $m(\tilde{q})-m(\tilde{\chi}_1^0)<5 \text{ GeV}$ 1711.03301 2-6 jets Yes 36.1 $m(\tilde{\chi}_1^0)$ <200 GeV 1712 02332 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$ 0 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{\pm} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0}$ 2-6 iets Yes 36 1 $m(\tilde{\chi}_1^0) < 200 \text{ GeV}, m(\tilde{\chi}^{\pm}) = 0.5(m(\tilde{\chi}_1^0) + m(\tilde{g}))$ 1712.02332 0 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\tilde{\chi}_1^0$ $ee, \mu\mu$ 2 jets Yes 14.7 $m(\tilde{\chi}_1^0)$ <300 GeV, 1611.05791 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow qq(\ell\ell/\nu\nu)\tilde{\chi}_1^0$ $3e, \mu$ 4 jets 36.1 1706.03731 $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$ 0 7-11 jets Yes 36.1 1.8 TeV $m(\tilde{\chi}_1^0)$ <400 GeV 1708.02794 GMSB ($\tilde{\ell}$ NLSP) $1-2 \tau + 0-1$ 0-2 jets Yes 3.2 1607.05979 GGM (bino NLSP) 2.15 TeV cτ(NLSP)<0.1 mm 2γ Yes 36.1 ATLAS-CONF-2017-080 GGM (higgsino-bino NLSP) 2 jets γ Yes 36.1 **2.05 TeV** $m(\tilde{\chi}_1^0) = 1700 \text{ GeV}, c\tau(NLSP) < 0.1 \text{ mm}, \mu > 0$ ATLAS-CONF-2017-080 Gravitino LSP 865 GeV $m(\tilde{G}) > 1.8 \times 10^{-4} \text{ eV}, m(\tilde{g}) = m(\tilde{g}) = 1.5 \text{ TeV}$ 0 mono-jet Yes 20.3 1502.01518 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}\tilde{\chi}_1^0$ 0 3hYes 36.1 1.92 TeV $m(\tilde{\chi}_1^0)$ <600 GeV 1711.01901 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0$ 0-1 e, μ 3 b Yes 36.1 **1.97 TeV** $m(\tilde{\chi}_1^0)$ <200 GeV 1711.01901 1708.09266 $\tilde{b}_1\tilde{b}_1,\,\tilde{b}_1{\rightarrow}b\tilde{\chi}_1^0$ 0 2 b Yes 36.1 $2e, \mu$ (SS) 36.1 275-700 GeV $m(\tilde{\chi}_{1}^{0}) < 200 \text{ GeV}, \ m(\tilde{\chi}_{1}^{\pm}) = m(\tilde{\chi}_{1}^{0}) + 100 \text{ GeV}$ 1706.03731 $\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{\chi}_1^{\pm}$ 1 *b* Yes $\tilde{t}_1\tilde{t}_1,\,\tilde{t}_1{\rightarrow}b\tilde{\chi}_1^{\pm}$ $0-2e, \mu$ Yes 4.7/13.3 117-170 GeV 200-720 GeV $m(\tilde{\chi}_{1}^{\pm}) = 2m(\tilde{\chi}_{1}^{0}), m(\tilde{\chi}_{1}^{0}) = 55 \text{ GeV}$ 1209.2102, ATLAS-CONF-2016-077 1-2 b $\tilde{t}_1\tilde{t}_1, \, \tilde{t}_1 {\rightarrow} Wb\tilde{\chi}_1^0 \text{ or } t\tilde{\chi}_1^0$ 0-2 e, μ 0-2 jets/1-2 b Yes 20.3/36. 90-198 GeV 0.195-1.0 TeV $m(\tilde{\chi}_1^0)=1 \text{ GeV}$ 1506.08616, 1709.04183, 1711.11520 $\tilde{t}_1\tilde{t}_1, \, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$ 0 mono-jet Yes 36.1 90-430 GeV $m(\tilde{t}_1)-m(\tilde{\chi}_1^0)=5 \text{ GeV}$ 1711.03301 $\tilde{t}_1\tilde{t}_1$ (natural GMSB) $2e, \mu(Z)$ 1 b Yes 20.3 150-600 GeV $m(\tilde{\chi}_1^0) > 150 \text{ GeV}$ 1403.5222 $\tilde{t}_2\tilde{t}_2,\,\tilde{t}_2{ ightarrow}\tilde{t}_1+Z$ Yes 36.1 $3 e, \mu (Z)$ 1 *b* 290-790 GeV $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ 1706.03986 $\tilde{t}_2\tilde{t}_2, \, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$ 1-2 e, μ 4 h Yes 36.1 320-880 GeV $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ 1706.03986 $2e, \mu$ $\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$ Yes 36.1 90-500 GeV $m(\tilde{\chi}_1^0)=0$ ATLAS-CONF-2017-039 0 $\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow \tilde{\ell} \nu(\ell \tilde{\nu})$ $2e, \mu$ 0 Yes 36 1 $m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^{\pm})+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2017-039 $\begin{array}{c} \ddot{\chi}_{1}^{\perp} \ddot{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{1} \nu \tilde{\ell}_{1} \tilde{\chi}_{1}^{0} \rightarrow \tilde{\tau}_{1} \\ \ddot{\chi}_{1}^{\pm} \dot{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{1} \nu \tilde{\ell}_{1} \ell (\tilde{\nu} \nu), \ell \tilde{\nu} \tilde{\ell}_{1} \ell (\tilde{\nu} \nu) \\ \ddot{\chi}_{1}^{\pm} \dot{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} Z \tilde{\chi}_{1}^{0} \\ \vdots \\ \ddot{\zeta}_{n}^{0} \end{array}$ $\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp}/\tilde{\chi}_{2}^{0}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau}\nu(\tau\tilde{\nu}), \tilde{\chi}_{2}^{0} \rightarrow \tilde{\tau}\tau(\nu\tilde{\nu})$ Yes 2 τ 36.1 760 GeV $m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$ 1708 07875 $3e, \mu$ 0 Yes 36 1 ATLAS-CONF-2017-039 1.13 TeV $m(\tilde{\chi}_{1}^{\pm})=m(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$



270 GeV

115-370 GeV

580 GeV

635 GeV

 $m(\tilde{\chi}_1^{\pm})=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \tilde{\ell}$ decoupled

 $m(\tilde{\chi}_1^{\pm})=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \tilde{\ell}$ decoupled

 $m(\tilde{\chi}_{2}^{0})=m(\tilde{\chi}_{3}^{0}), m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_{2}^{0})+m(\tilde{\chi}_{1}^{0}))$

 $\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W\tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow eev, e\mu\nu, \mu\mu\nu$ 1.14 TeV $4e, \mu$ Yes 13.3 $m(\tilde{\chi}_1^0) > 400 \text{GeV}, \ \lambda_{12k} \neq 0 \ (k = 1, 2)$ ATLAS-CONF-2016-075 $\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W \tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau \tau \nu_e, e \tau \nu_{\tau}$ 450 GeV $3e, \mu + \tau$ 20.3 $m(\tilde{\chi}_{1}^{0}) > 0.2 \times m(\tilde{\chi}_{1}^{\pm}), \lambda_{133} \neq 0$ 1405.5086 $\tilde{g}\tilde{g}, \, \tilde{g} \rightarrow qq\tilde{\chi}_1^0, \, \tilde{\chi}_1^0 \rightarrow qqq$ 0 4-5 large-R iets -36.1 $m(\tilde{\chi}_{1}^{0})=1075 \text{ GeV}$ SUSY-2016-22 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qqq$ 8-10 jets/0-4 h $1 e, \mu$ 36.1 **2.1 TeV** $m(\tilde{\chi}_{1}^{0}) = 1 \text{ TeV}, \lambda_{112} \neq 0$ 1704 08493 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$ $m(\tilde{t}_1)=1 \text{ TeV}, \lambda_{323}\neq 0$ 8-10 iets/0-4 b 36.1 $1e, \mu$ 1704.08493 2 jets + 2 b 100-470 GeV 480-610 GeV $\tilde{t}_1\tilde{t}_1, \, \tilde{t}_1 {\rightarrow} bs$ 36.7 1710 07171 0 $\tilde{t}_1\tilde{t}_1,\,\tilde{t}_1{\rightarrow}b\ell$ $2e, \mu$ 0.4-1.45 TeV BR($\tilde{t}_1 \rightarrow be/\mu$)>20% 2 b 36.1 1710.05544 Yes 20.3 510 GeV 1501.01325 Other Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$ $m(\tilde{\chi}_1^0)$ <200 GeV

 $\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W\tilde{\chi}_{1}^{0}h\tilde{\chi}_{1}^{0}, h \rightarrow b\bar{b}/WW/\tau\tau/\gamma\gamma$

GGM (wino NLSP) weak prod., $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$ 1 $e, \mu + \gamma$

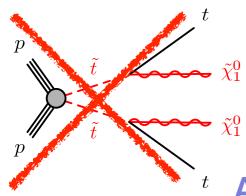
 $\tilde{\chi}_{2}^{0}\tilde{\chi}_{3}^{0}, \tilde{\chi}_{2,3}^{0} \rightarrow \tilde{\ell}_{R}\ell$

^{*}Only a selection of the available mass limits on new states or phénomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.



Outline

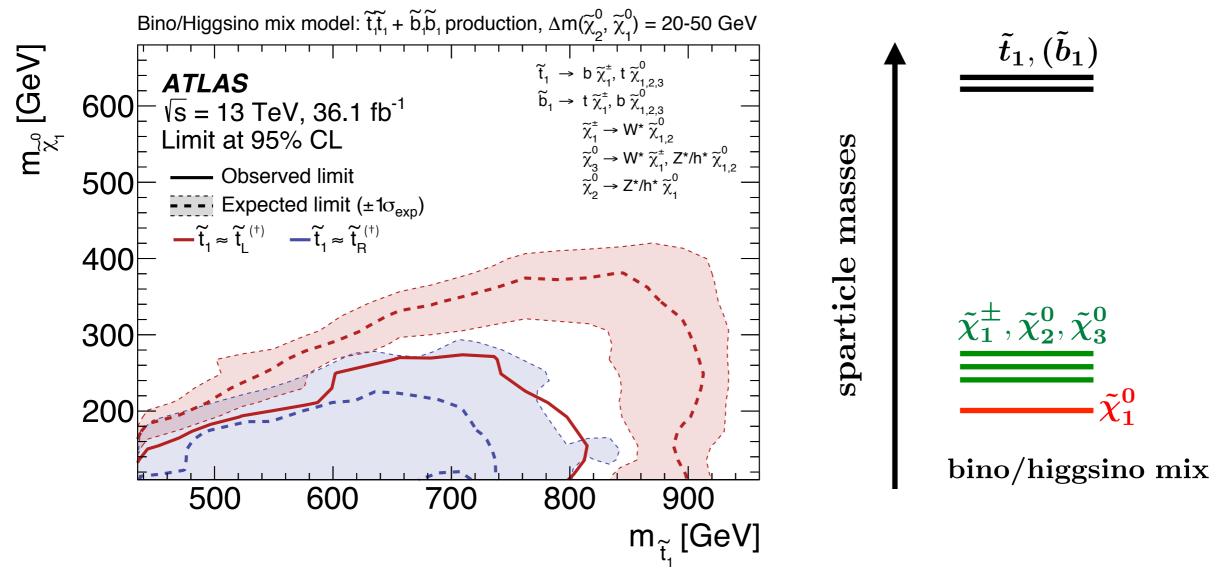
Hunting For Vanilla Stops



Stop search reinterpreted in terms of a pMSSM-inspired model:

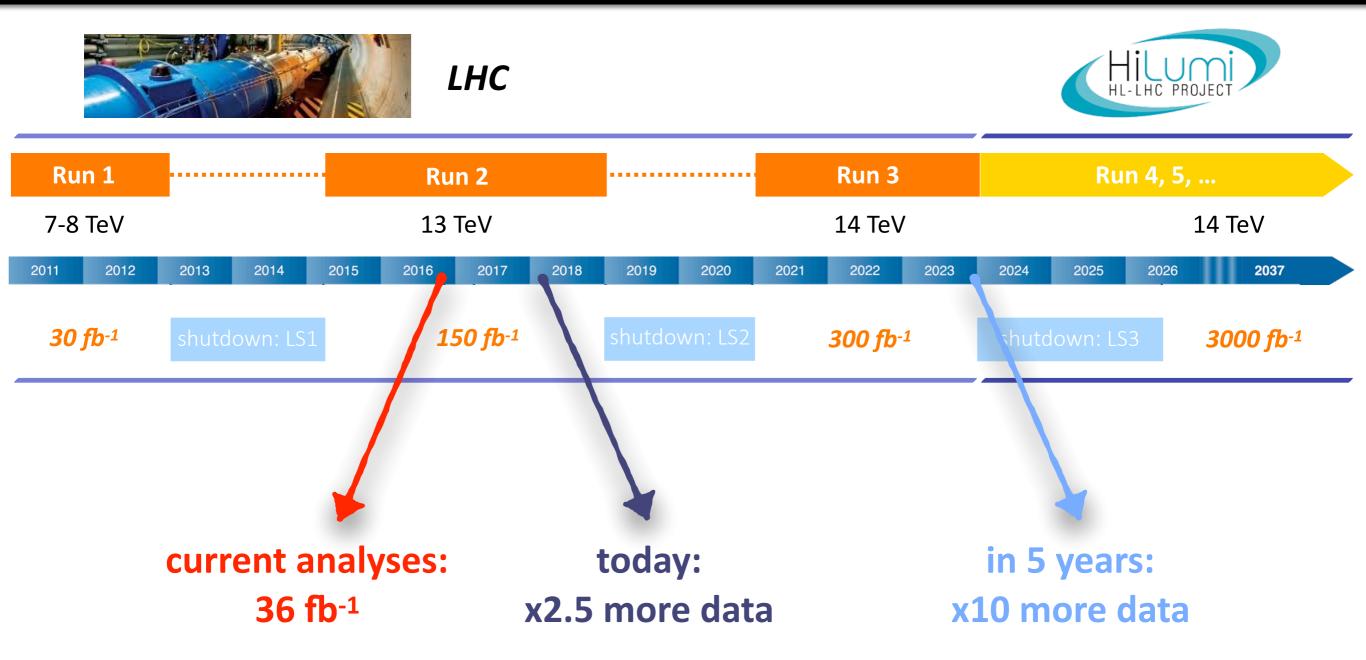
- \rightarrow correct dark matter relic abundance: 0.10<Ω h^2 <0.12
- → natural, compressed EWKinos mass spectrum

A light stop with a light higgsino-like LSP might still be allowed



^{(†) &}quot;handiness" of scalar top comes from couplings (SM partner, top, is a fermion)

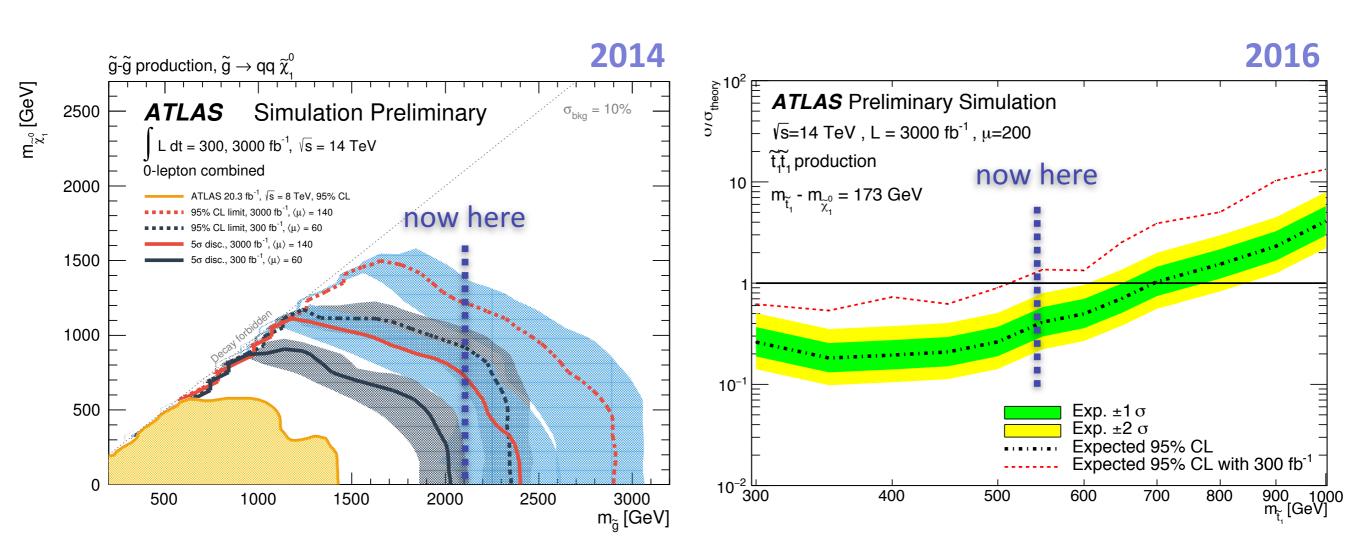
LHC Long-Term Schedule

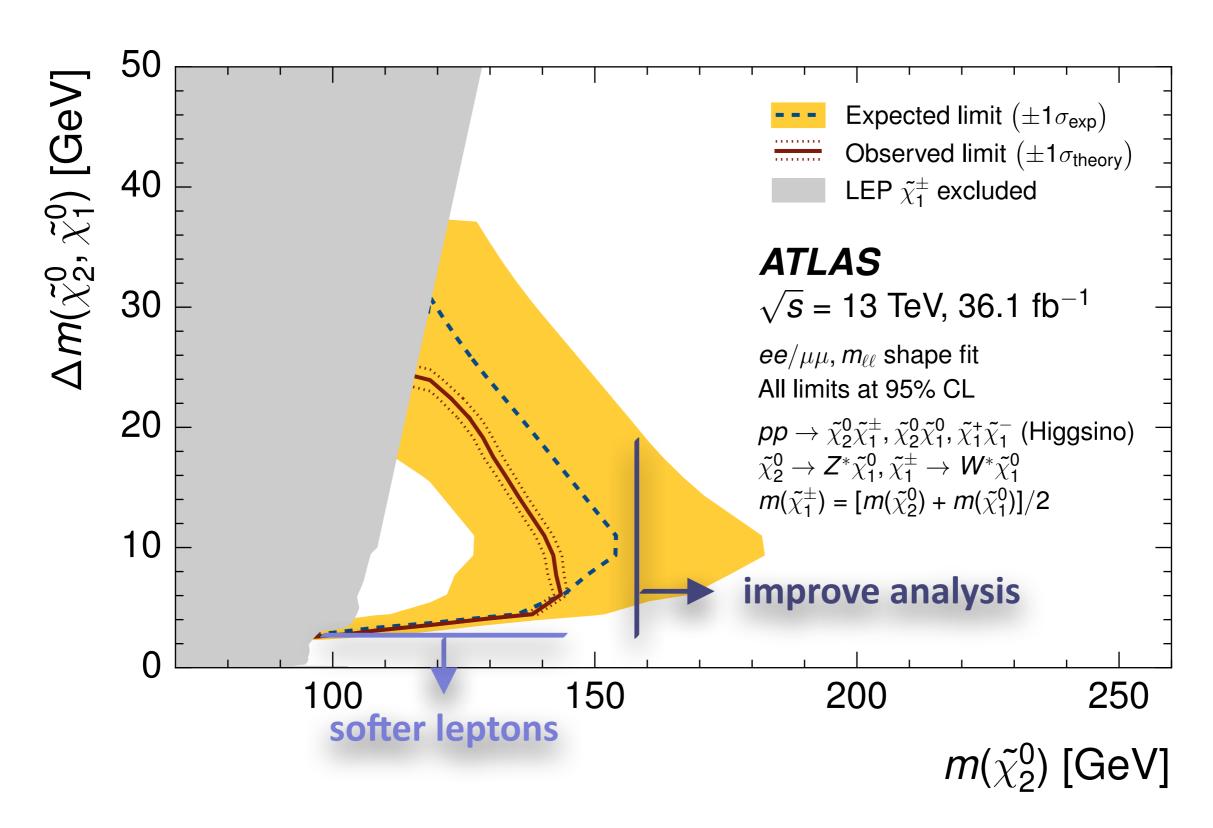


100x data expected to bring <50% improvements in mass reach

extrapolations based on simple analyses → probably over-conservative...
...however, the main message stands!

searches will move toward more <u>sophisticated analysis techniques</u> and more <u>complex signatures</u>





Improving Reach Of Higgsino Analysis

Softer leptons → lower Δm

designed new PID level for soft muons

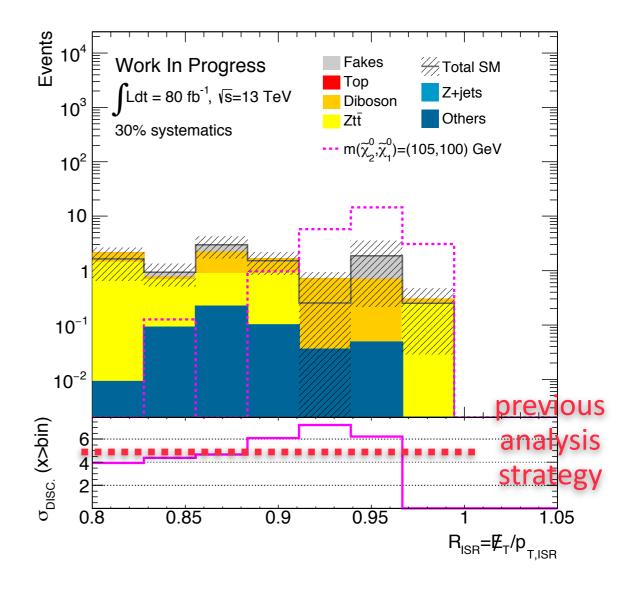
recovers efficiency for p_T <6 GeV tracks, improves π/K rejection for p_T >6 GeV

Rate w.r.t. reconstructed ID tracks [%] √s=13 TeV, tt̄ MC Medium PID 3.0<p_<5.0 [GeV] 20 _owPt PID new Prompt muons 00 80 40 20 0.5 1.5 2.5 lηl

Improve analysis → larger *m* reach

adopt thrust-based ISR identification

very similar to the diagonal stop case, large improvements for small mass splittings

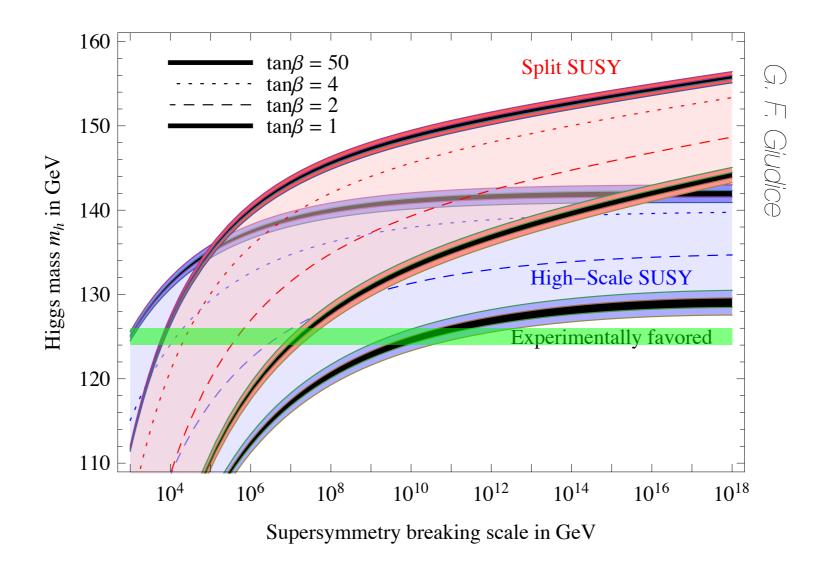


Long-Lived And/Or Highly-Ionizing Particles

Long-lived and/or highly-ionizing particles ("R-hadrons") well motivated

Hideyuki Oide (Genova) and Larry L. Lee (Harvard) leading current efforts in ATLAS

Interest can only rise if we will progressively abandon natural SUSY!



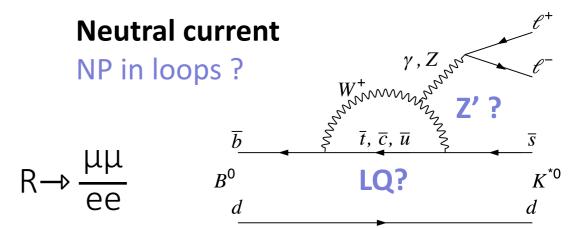
New physics can show up first from indirect constraints (LHCb - intensity frontier)

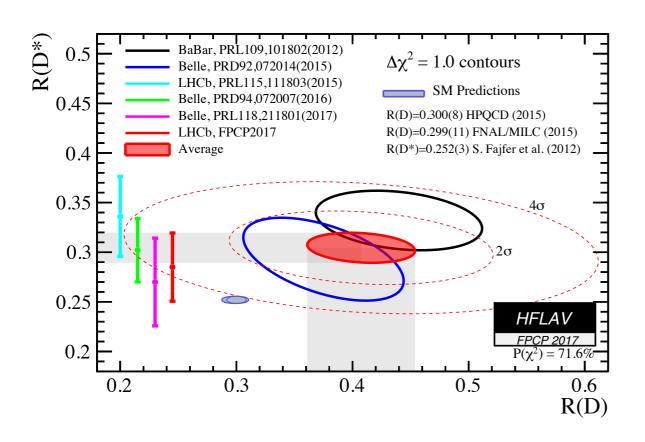
first evidences of lepton flavor universality violation in b→c and b→s transitions?

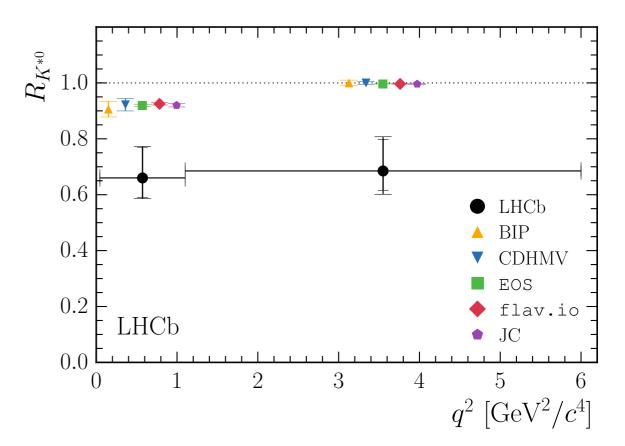
Charged current

NP at tree level?

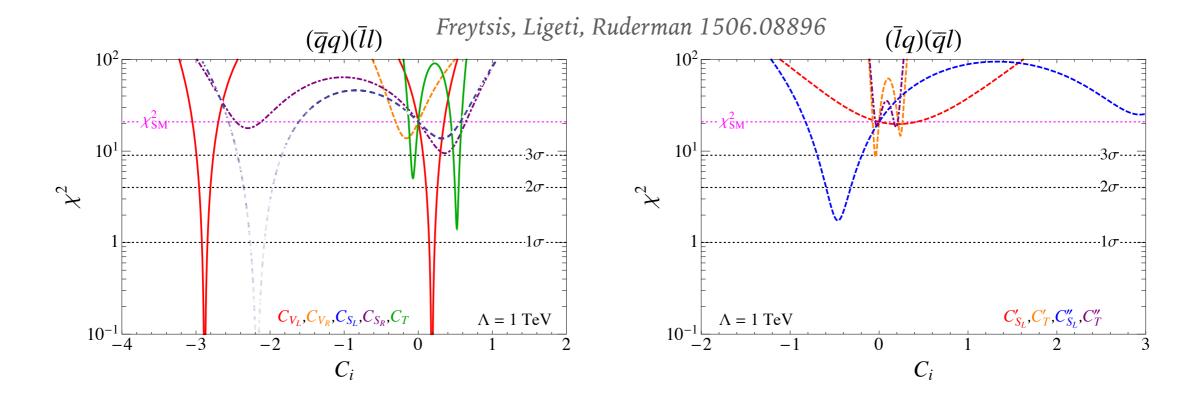
$$R \rightarrow \frac{\tau v_{\tau}}{\mu v_{\mu}} \quad \overline{B} \left\{ \begin{array}{c} b \\ \overline{q} \end{array} \right\} D^{(*)}$$





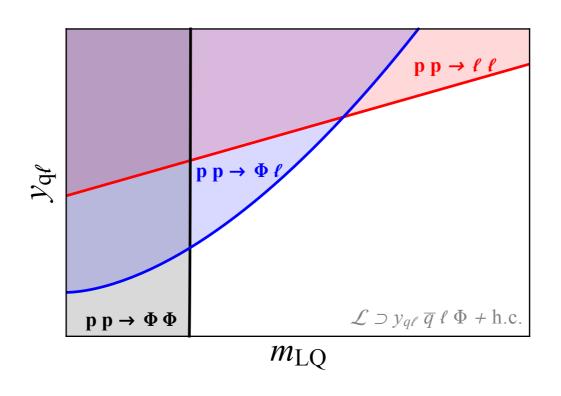


"Flavor Anomalies"



Favored interpretation: *leptoquark?*

If real, though for SUSY but great for SUSY searches!

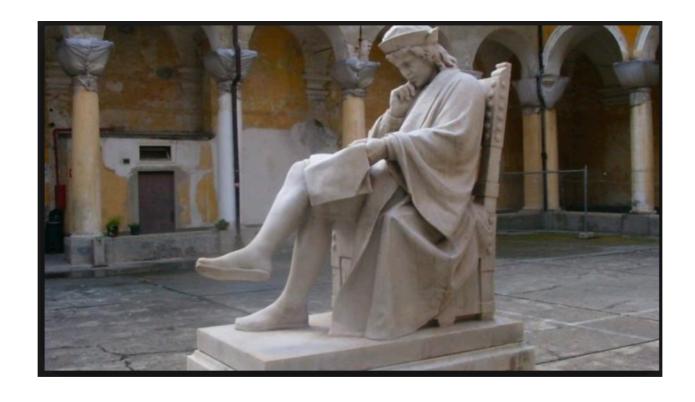


E.g. colored scalar, decaying into:

Colombo left Genova with an idea: he sailed to India... and found America!

We also left for a long journey seeking new physics, but no land at the horizon yet...

Will we find SUSY? Will we find our Americas?

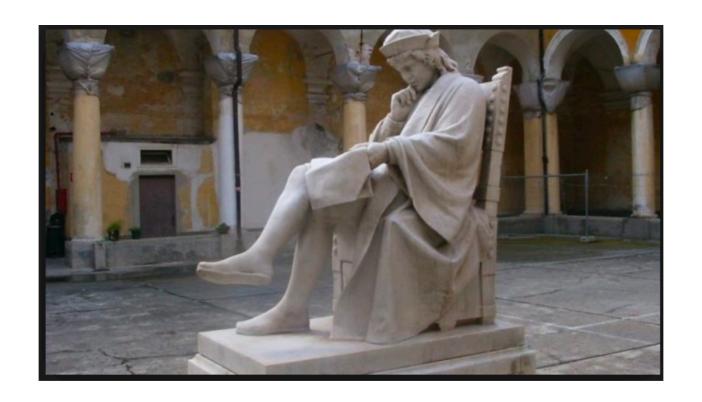


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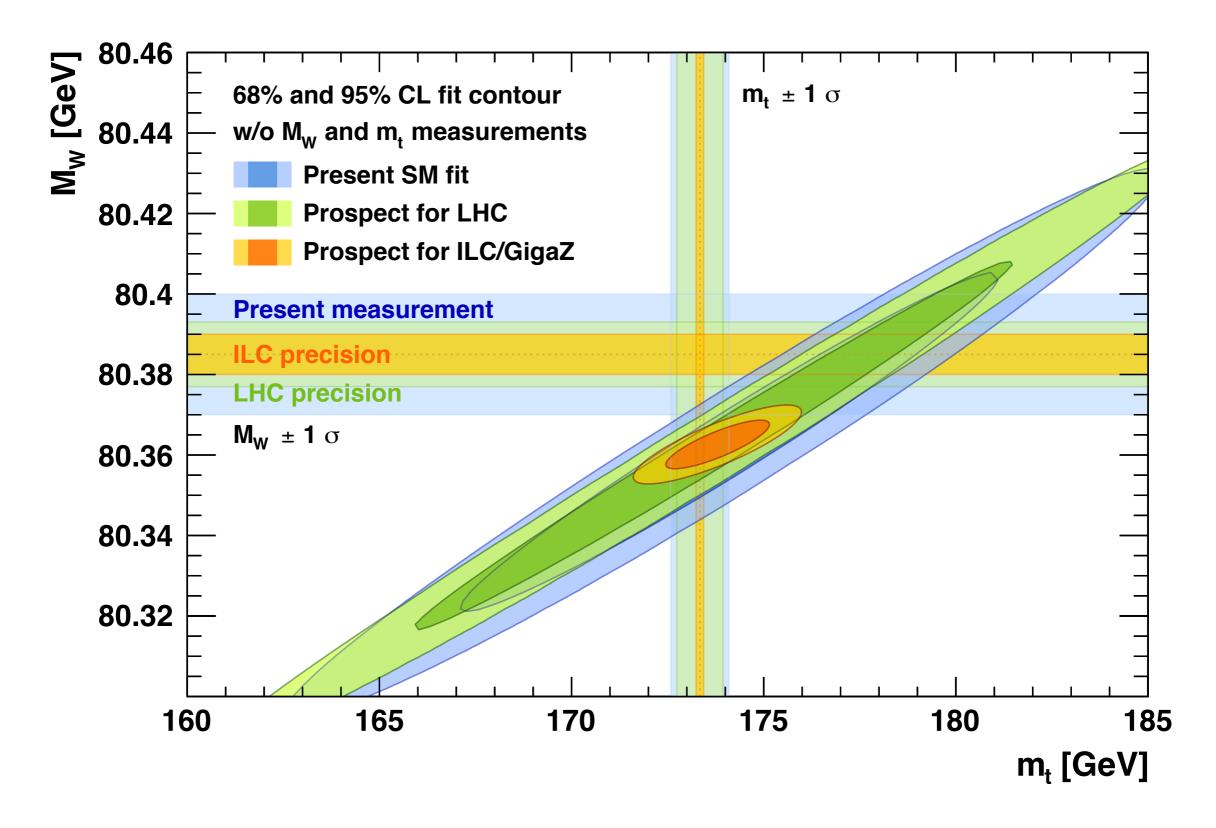
Will we find SUSY? Will we find our Americas?

One thing is for sure: no matter what, we will have a much richer cartography, which will dictate the direction of next generation of HEP experiments





Precision Measurements Of SM Parameters

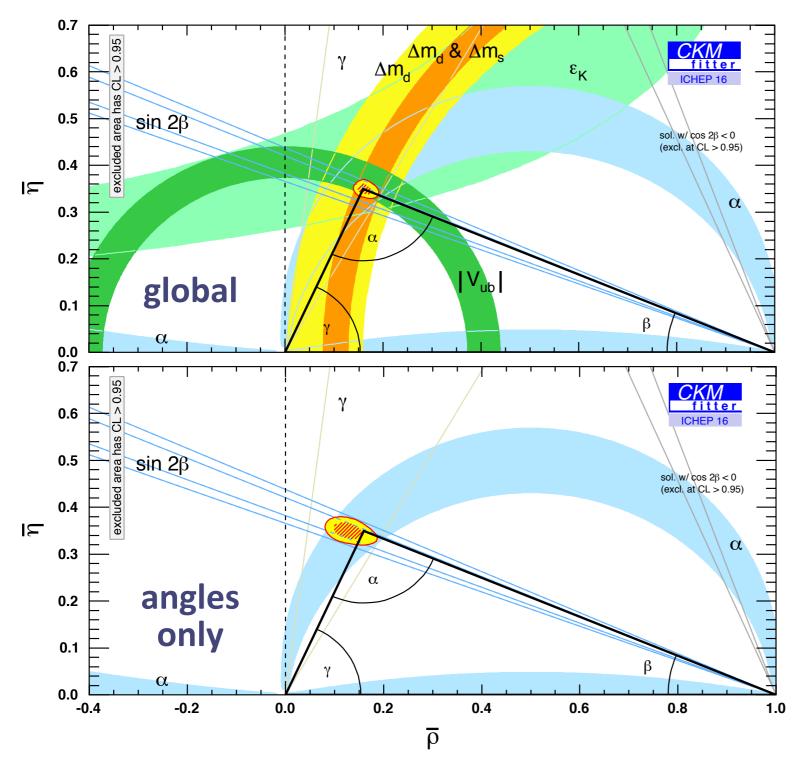


CKM Matrix: Measuring γ

near future

intensity: LHCb

γ : least constrained angle of unitary triangle!



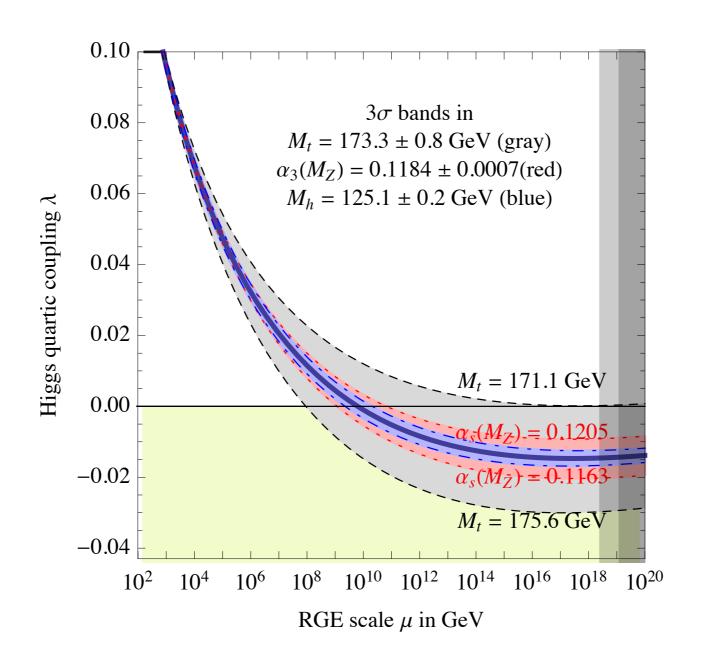
Insensitive to "pollution" from BSM physics: perfect reference to seek for SM deviations using global fits

Accessible in B→DK decays involving tree-level processes with ≈ no hadronic uncertainties

Currently knows within a ≈5° accuracy: beating the 1° barrier is one of LHCb goals

Investigating the near-criticality of the Higgs boson

Dario Buttazzo a,b , Giuseppe Degrassi c , Pier Paolo Giardino a,d , Gian F. Giudice a , Filippo Sala b,e , Alberto Salvio b,f , Alessandro Strumia d

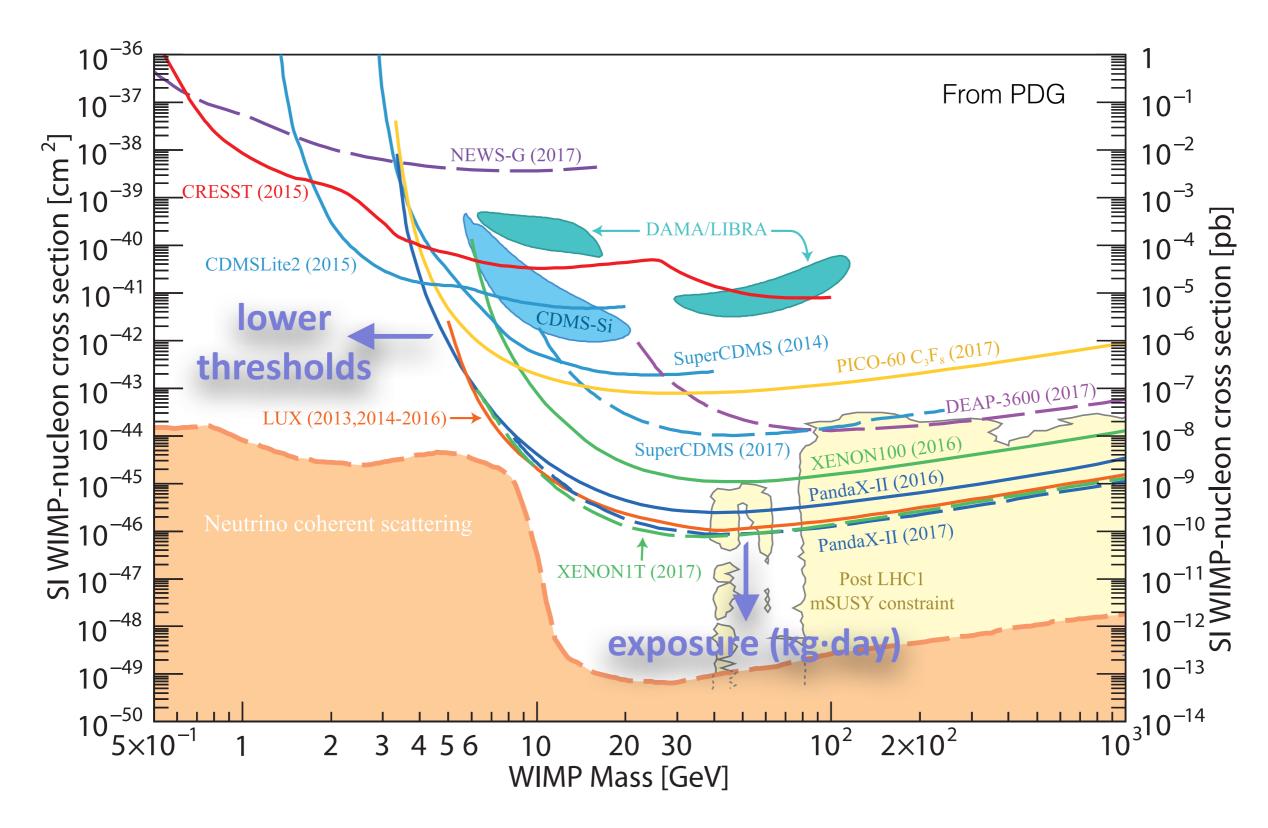


Dark Matter: Direct Detection

near future

intensity: LHCb

cosmic/neutrino frontier



Dark Matter: Directional Direct Detection

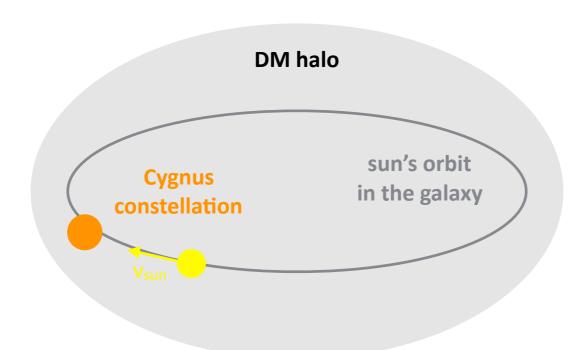
near future

intensity: LHCb

cosmic/neutrino frontier

Anisotropic signal

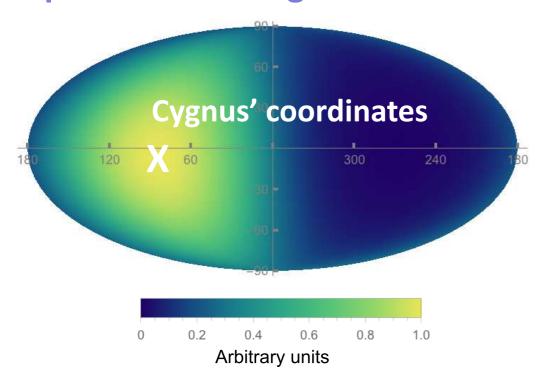
"wind" of WIMPs in the direction of sun's motion (≈Cygnus)



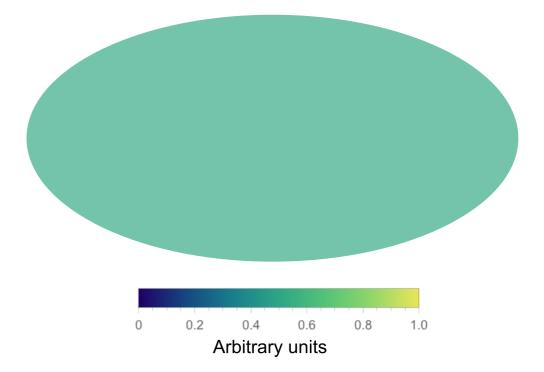
Annual modulation

Earth's motion induces (small) annual modulation

Expected WIMP signal distribution



Expected background distribution



S. Zambito | *Harvard University*

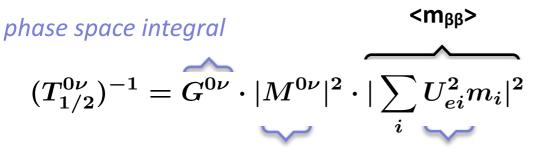
near future

intensity: LHCb

cosmic/neutrino frontier

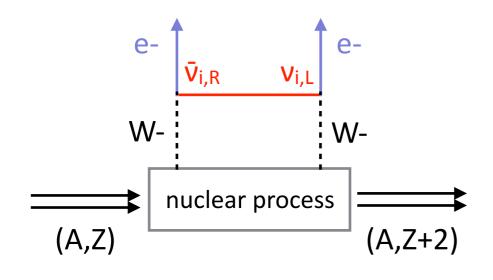
Are neutrinos Dirac or Majorana particles?

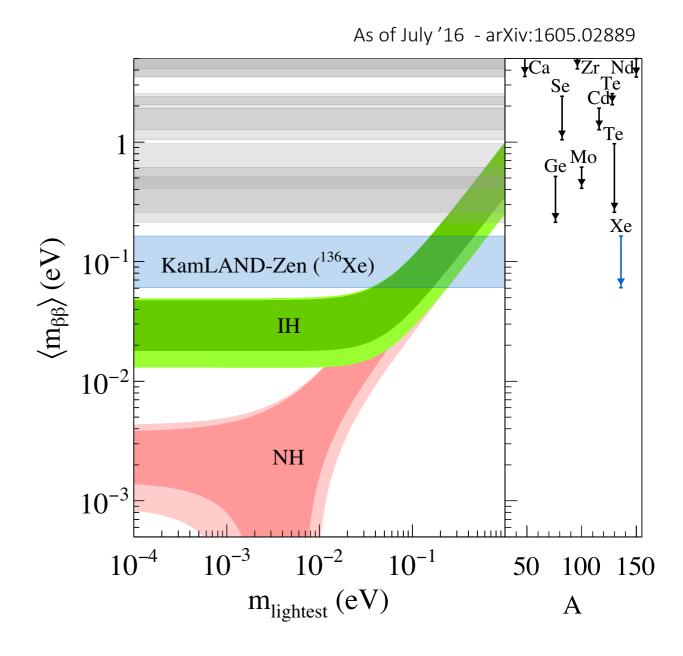
Ovββ aim to address this question, measure v mass scale, test lepton number conservation can probe new physics contributions: need multi-isotope comparative analysis!



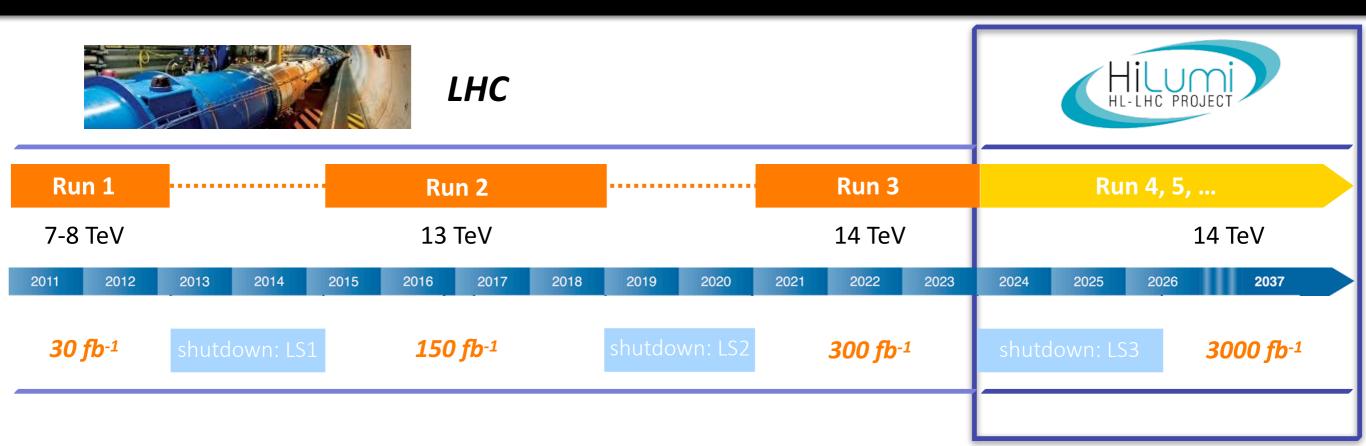
nuclear matrix element

PMNS matrix

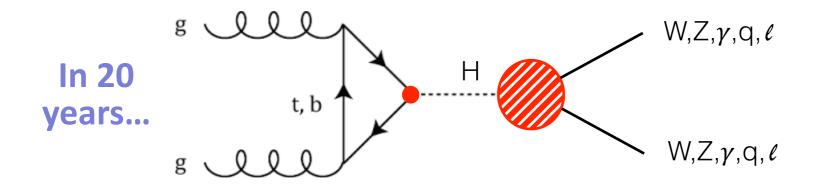




LHC Long-Term Schedule

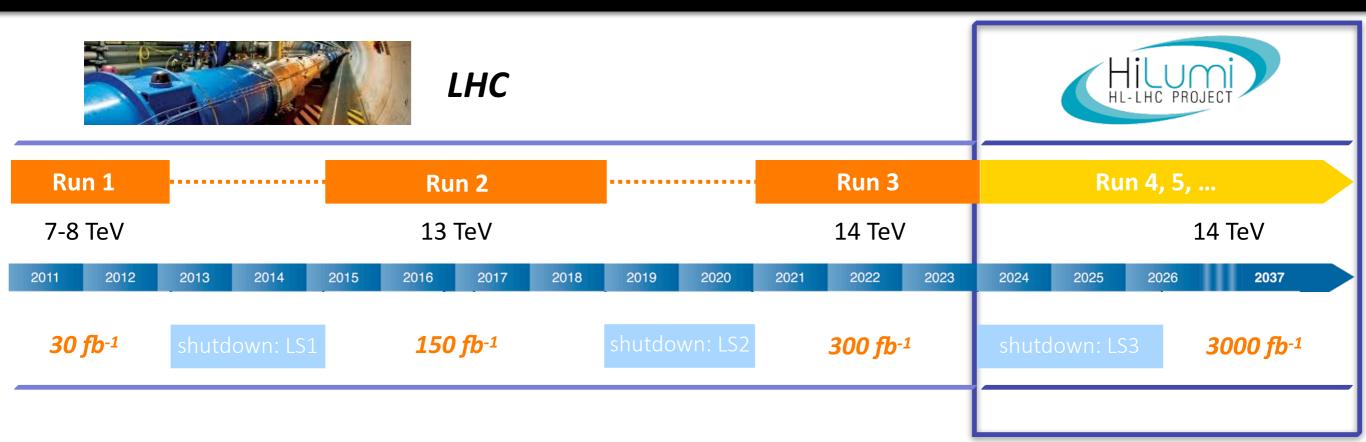


High intensity @ HL-LHC: shifting toward precision physics e.g. fingerprinting Higgs sector

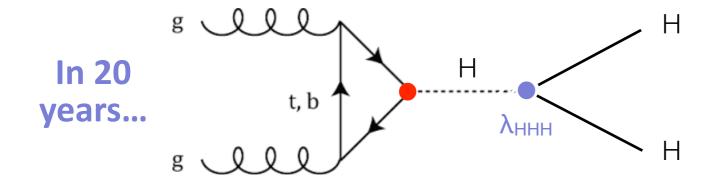


≈5% precision on Higgs couplings: SM-like or not?

LHC Long-Term Schedule



High intensity @ HL-LHC: shifting toward precision physics e.g. fingerprinting Higgs sector



≈5% precision on Higgs couplings: SM-like or not?

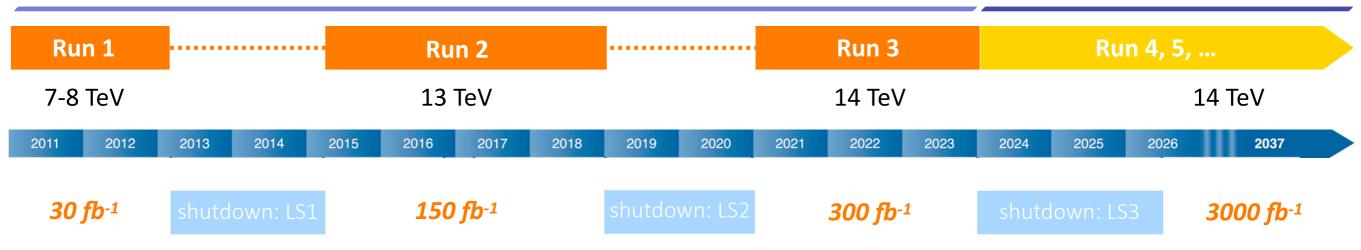
 $1 \le \lambda/\lambda_{SM} \le 9$ at 95% CL: not enough to resolve details of Higgs potential

LHC Long-Term Schedule + Future Colliders?



LHC





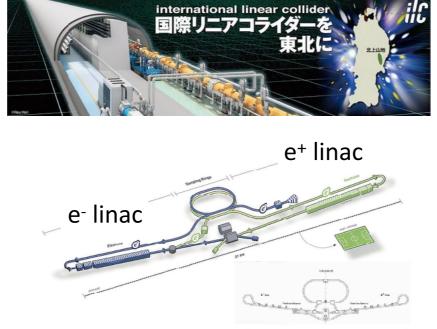
FCC - reach \sqrt{s} =100 TeV

search/study new physics

Schematic of an 80 - 100 km long tunnel Aravis

ILC - mature technology

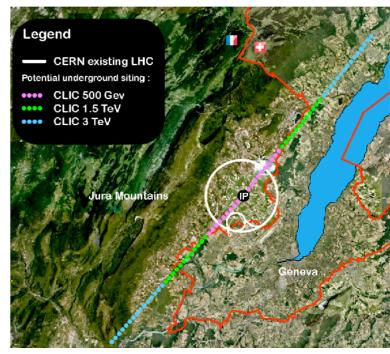
characterize the Higgs



31 km

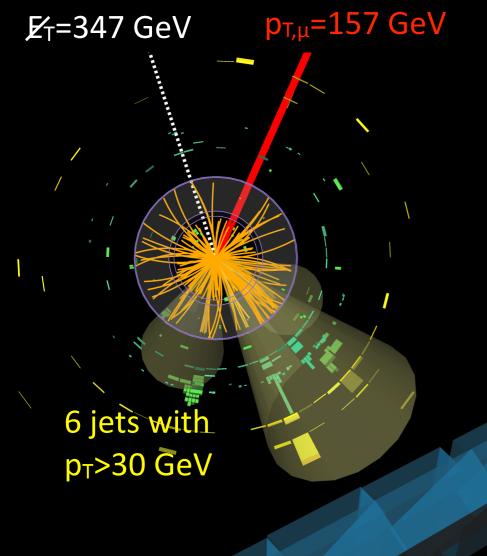
CLIC - new technology

Higgs and new physics?



S. Zambito | *Harvard University*

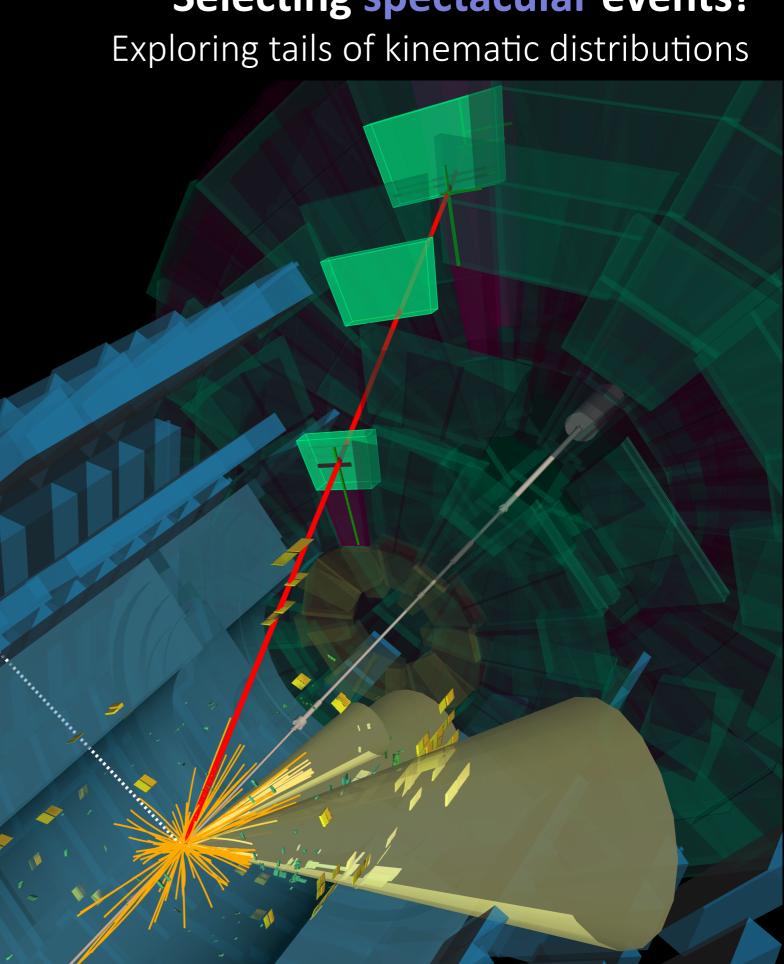
Selecting spectacular events!



Run: 279598 Event: 929301935

2015-09-17 09:53:11 CEST





Gluinos: Background Estimation

Two well-understood and well-modeled variables define CR/VR/SR plane

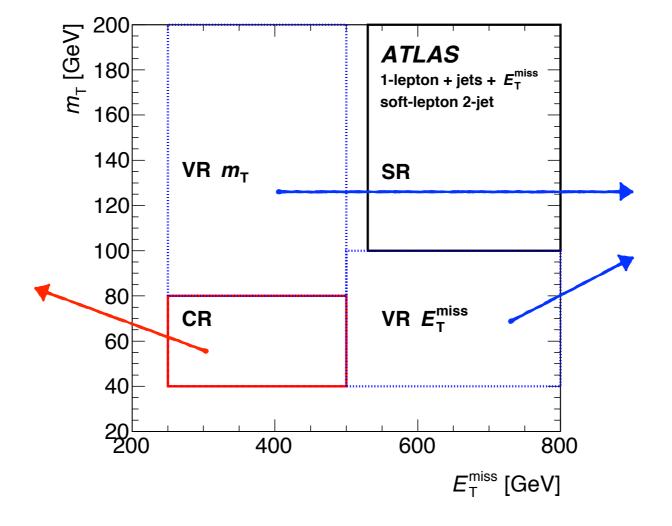
Control Regions (CR) → normalize simulated backgrounds to data

 \hookrightarrow select or veto b-jets to enrich in $t\bar{t}$ or W+jets

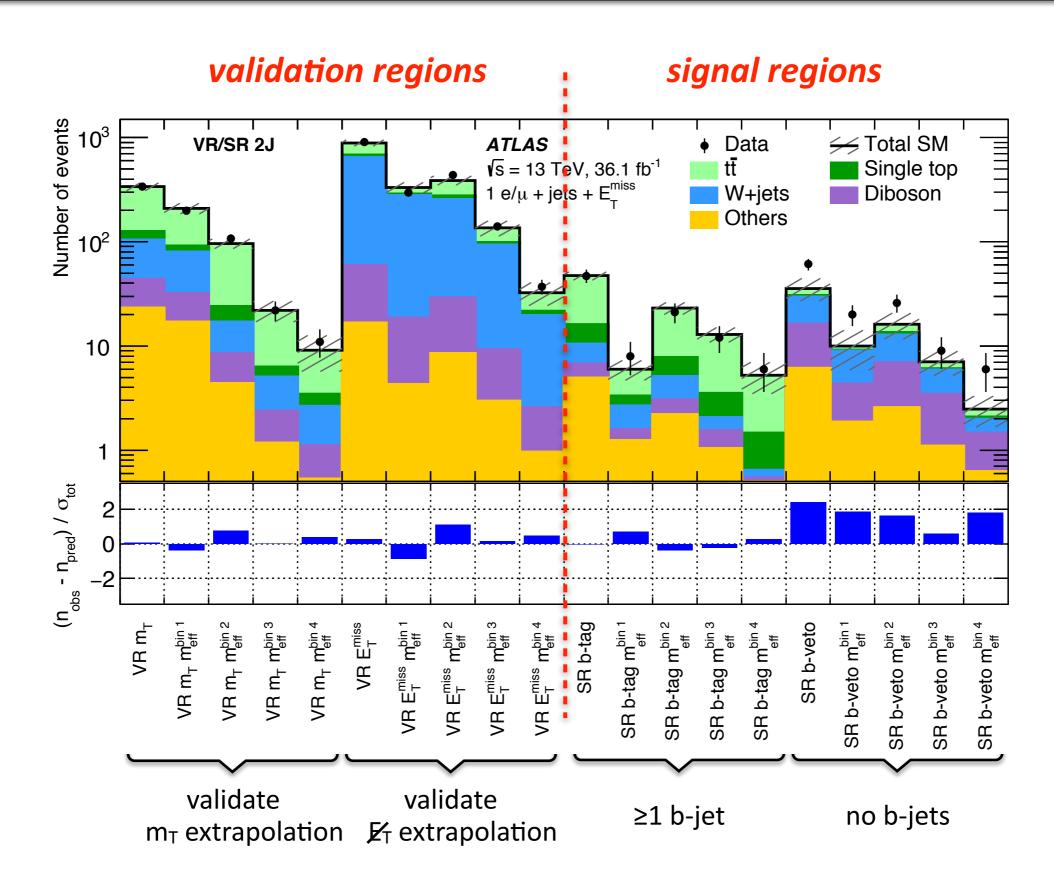
→ extrapolate to SR using MC-based transfer factors

Validation Regions (VR) → validate background estimates against data

kinematic region close to SR, with negligible signal contamination

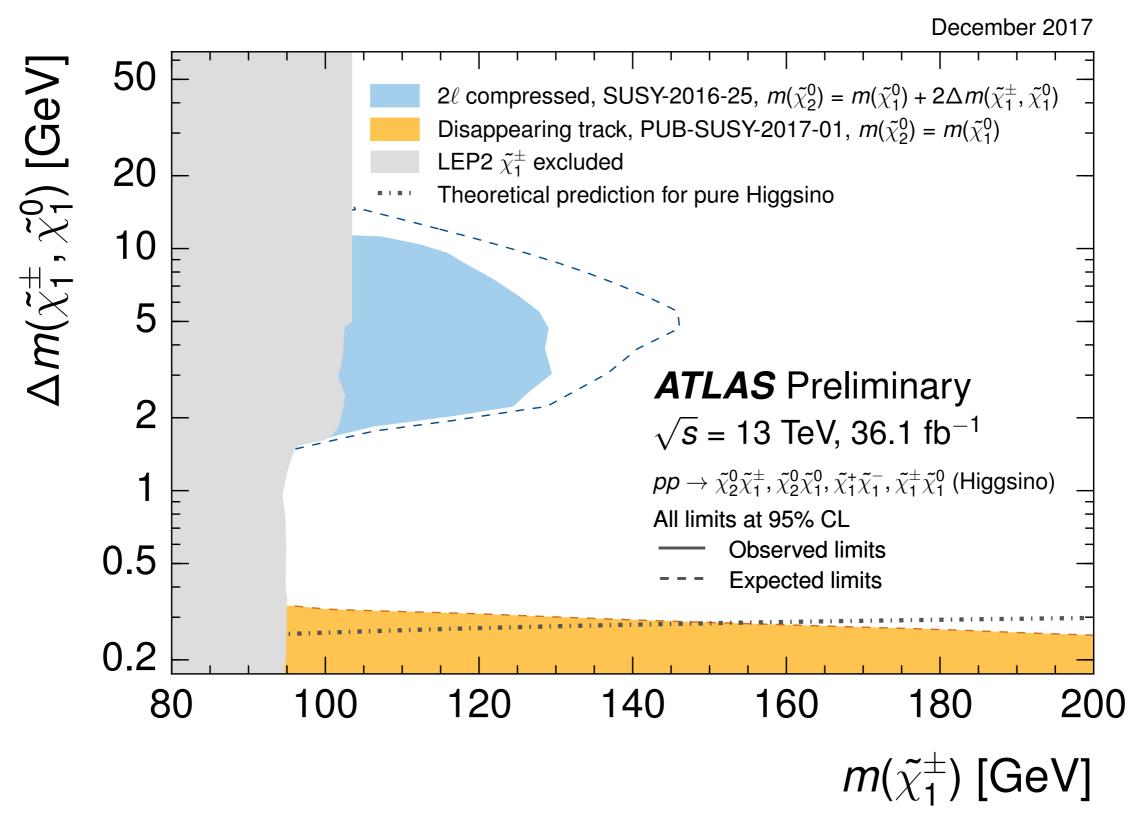


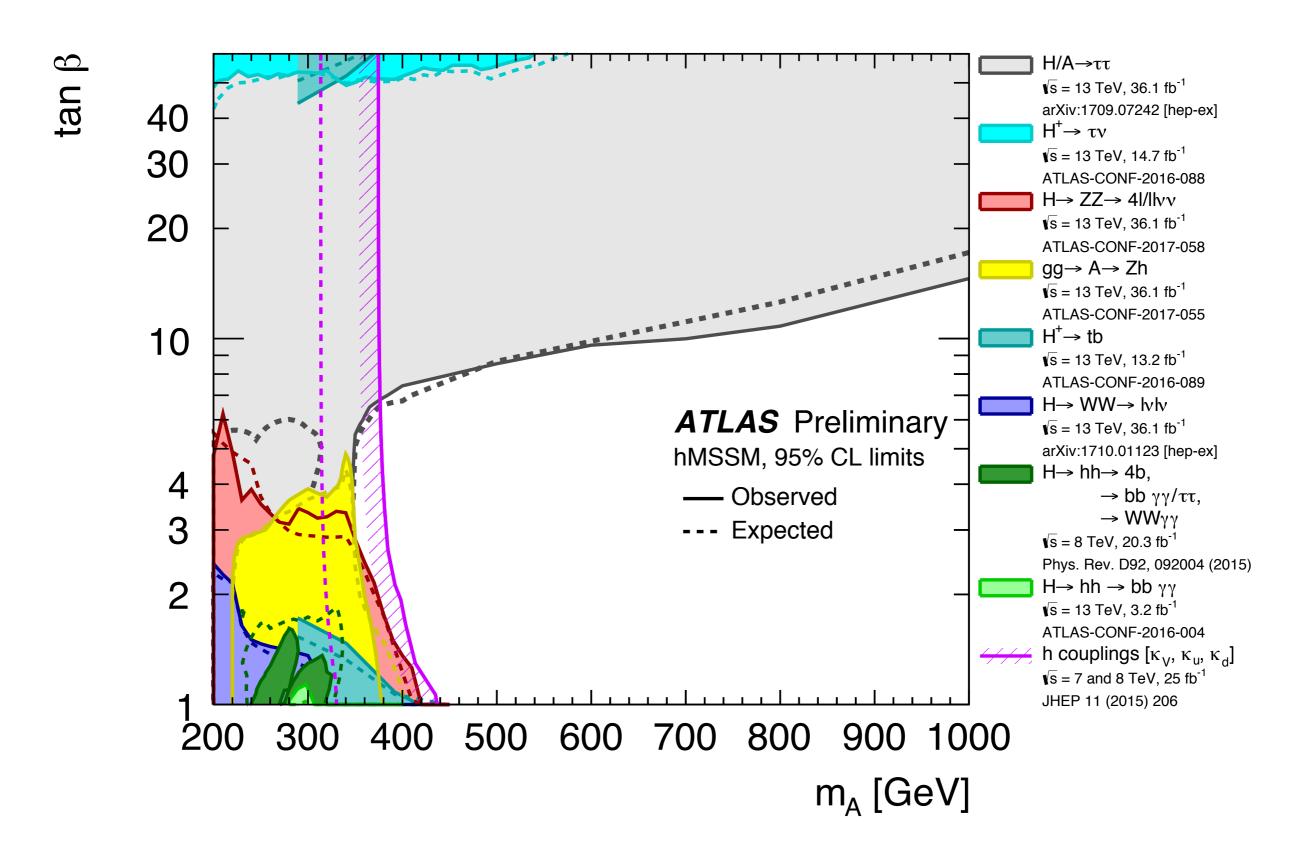
validate CR → SR extrapolation along each variable, separately



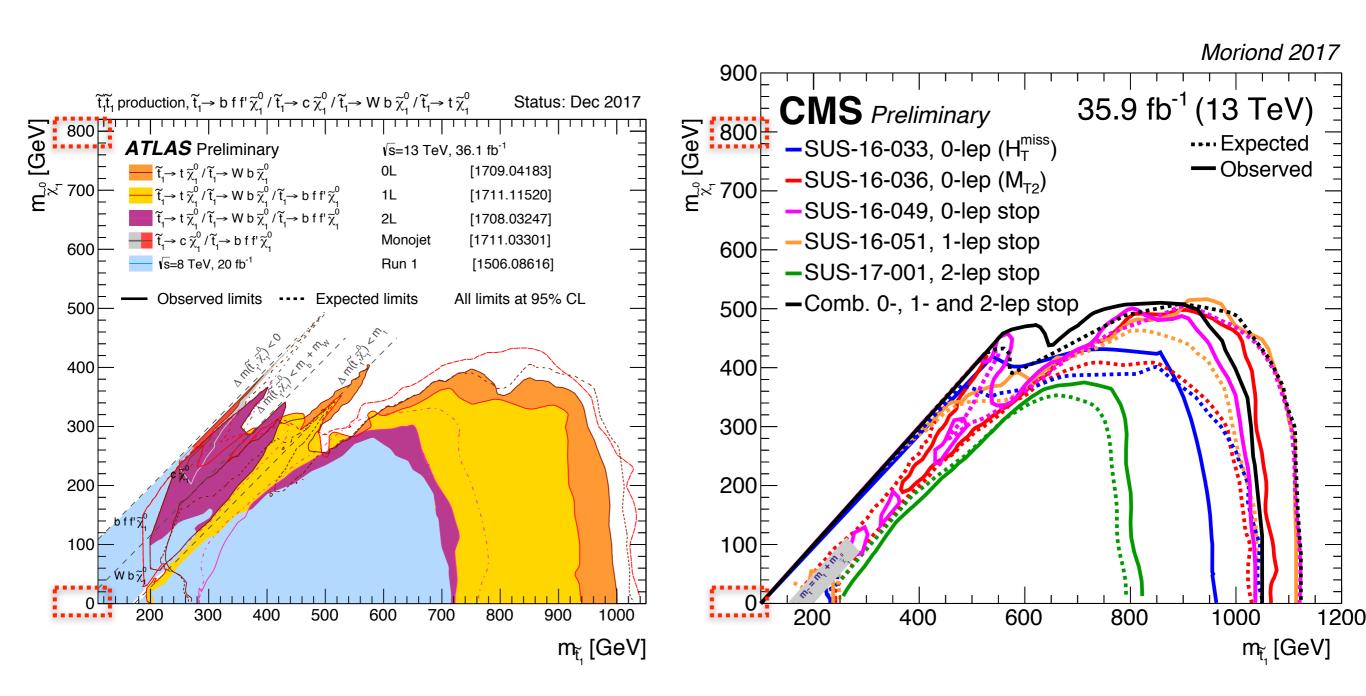
We Need To Get Smarter!

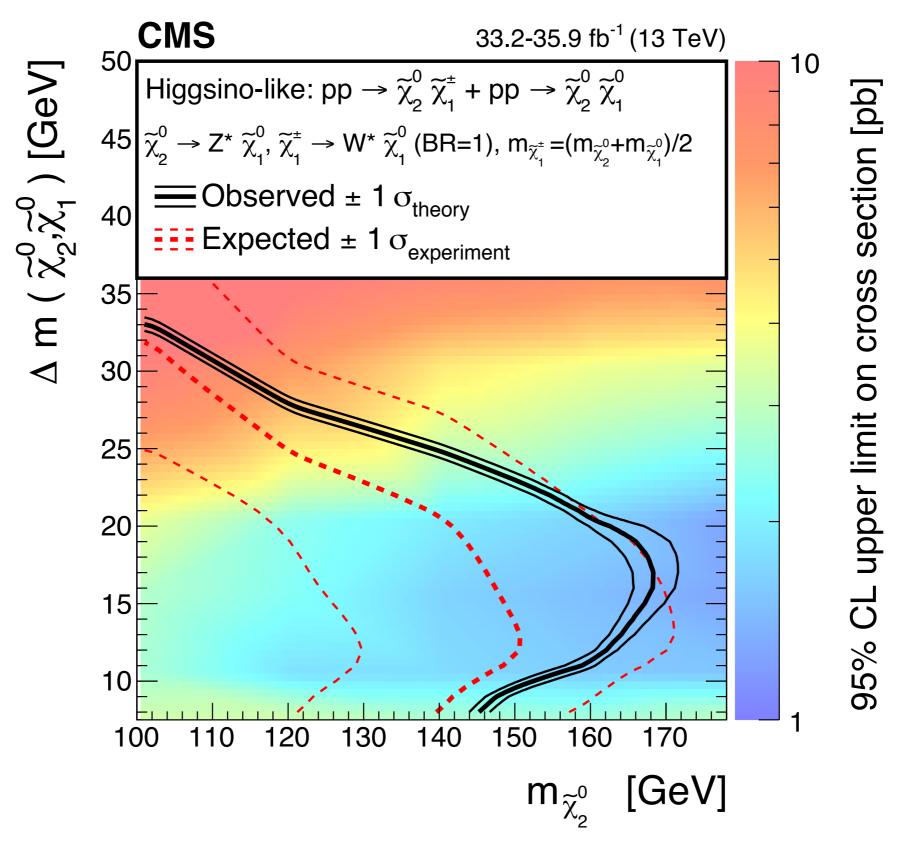
near future





Stop Searches - Atlas Vs CMS





DM @ The LHC

