

A detailed 3D rendering of the IXPE (Insight to X-ray Polarization Experiment) satellite in space. The satellite is shown from a perspective that highlights its large, flat, rectangular panels, which are likely part of the polarimetry instrument. The satellite's main body is dark, with a prominent circular opening at the bottom. In the background, a vibrant, colorful nebula in shades of purple, pink, and red is visible against a starry black sky. The Earth's horizon is partially visible in the bottom left corner, showing blue oceans and green landmasses. The overall scene is set in a deep space environment.

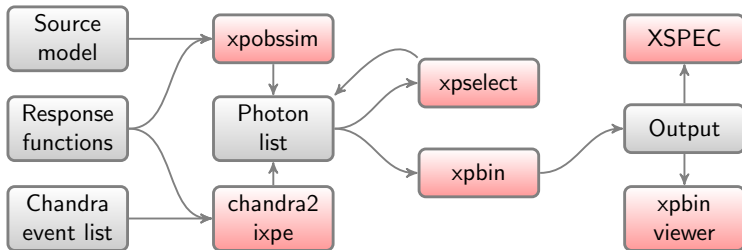
ixpeobssim

IXPE observation simulator

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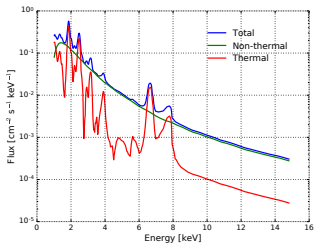
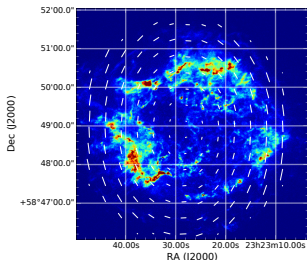
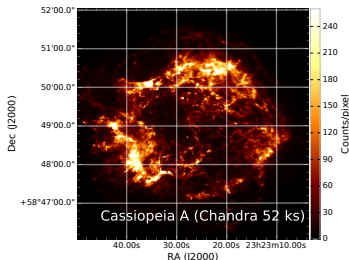
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- ▷ Generate photon lists corresponding to specific IXPE observations, based on:
  - ▷ A set of instrument response functions;
  - ▷ An arbitrary source model including morphological, temporal, spectral and polarimetric information;
  - ▷ OR a Chandra event file to be converted.

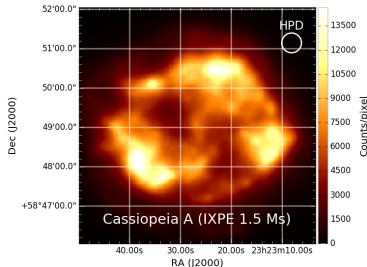
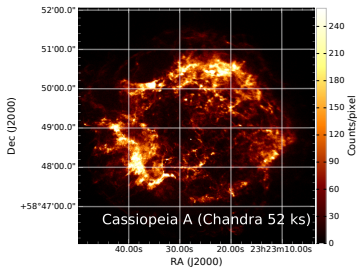
- ▷ Based on the Python programming language and the SciPy stack;
  - ▷ Details and installation instructions on the doc page.
- ▷ All I/O is FITS based.
  - ▷ The format of the response files is OGIP compliant;
  - ▷ Output photon lists and associated response files can be fed into xspec.
- ▷ ixpeobssim repository: <https://bitbucket.org/ixpesw/ixpeobssim>
- ▷ ixpeobssim documentation: <http://bigfoot.iaps.inaf.it:8080/xwiki/wiki/ixpeglobal/view/Main/IXPE%20Software%20Documentation/?Software%20Documentation>

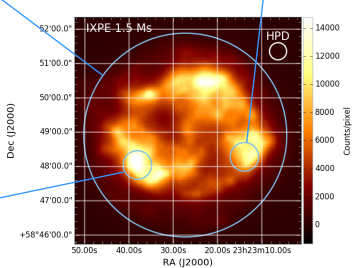
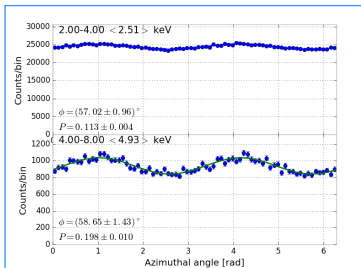
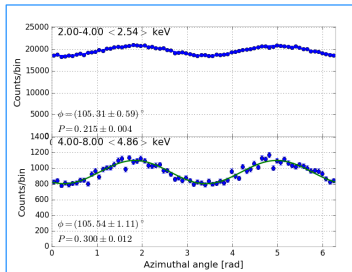
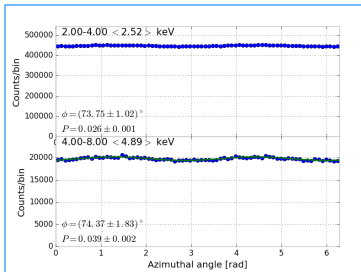


- ▷ Need four functions:
  - ▷ Morphology: point sources, disks, extended sources;
  - ▷ Energy spectrum;
  - ▷ Polarization model (degree and angle, or Stokes parameters).
- ▷ Support for phase-dependent periodic source;
- ▷ Can overlay an arbitrary number of components in the same model;

Simulate an observation starting from an arbitrary source model:

- ▷ Calculate the expected number of events based on the source spectrum and the effective area; extract the event times;
- ▷ Extract the true energies and positions in the sky and smear them with the energy dispersion and the PSF;
- ▷ Generate the angular distribution of the photoelectrons according to the polarization model.

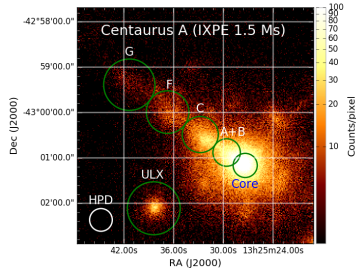
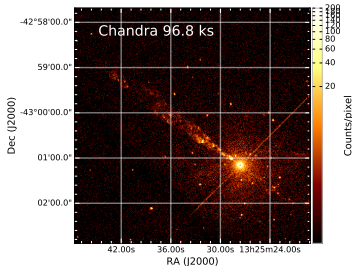




Process a Chandra photon list and produce an IXPE simulation:

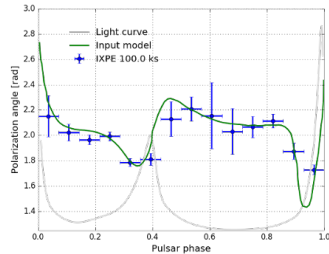
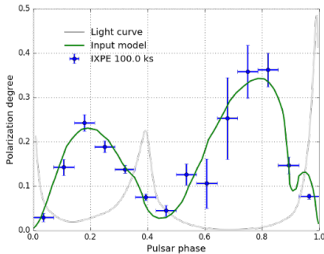
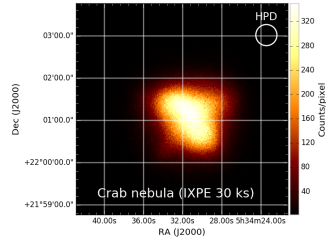
- ▷ Chandra measured energies, times and positions taken as MC truth;
- ▷ Events are down-sampled and smeared with the provided response functions;
- ▷ Generate the angular distribution of the photoelectrons according to the polarization model.

Preserve the full correlation between the morphology and the energy spectrum.



Basic tools available to:

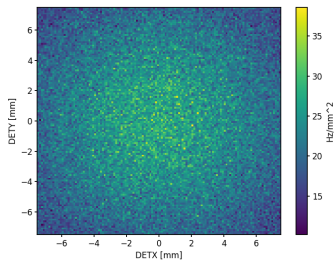
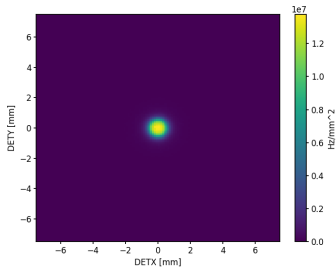
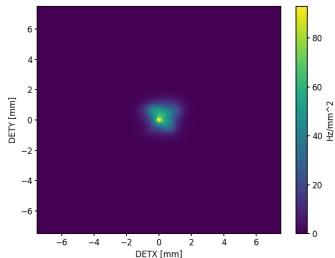
- ▷ Select subsamples of photons based on event energy, direction, time or phase;
- ▷ Bin and fit the simulated data, producing count maps, spectra, phasograms, light and modulation curves.





New binning method in order to:

- ▷ Calculate the density rate on the GPD for every sources;
- ▷ Evaluate the flux in case of Sun observation (quiet or flaring).



- ▷ Up and running, doing useful things, sophisticated in many respects.
- ▷ Large flexibility in terms of input models:
  - ▷ Point and extended sources;
  - ▷ Steady, variable and periodic sources;
  - ▷ Time and energy-dependent polarization signatures;
  - ▷ Arbitrary number of sources (and components) in the field of view.
- ▷ It surely has rough edges:
  - ▷ Essentially a product of four persons, so far;
  - ▷ Not really extensively used beyond the limited number of examples in the showcase.
- ▷ Need more core developers—please volunteer!
- ▷ Need more users providing feedback.

- ▷ A number of major changes in the works:
  - ▷ Splitting the event generation into the three detector units (with separate response functions);
  - ▷ Updating the format of the output photon lists (Level 2);
  - ▷ Updating the format of the response functions (CALDB);
  - ▷ Switch from polarization degree/angle to stoke parameters altogether.
- ▷ In this phase expect some issues:
  - ▷ Examples not fully working;
  - ▷ Documentation not updated.
- ▷ This should not prevent you from starting working now;
  - ▷ Don't wait for things to be perfect.
- ▷ You can actually help make this happen.