Dark sector @ Phase 2: dark photon et al.

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Different possible portals between **Dark Matter** and **Standard Model** depending on the **dark mediator X**:

 $\begin{array}{l} \mbox{Vector portal} \rightarrow \mbox{Dark Photon} \\ \mbox{Scalar portal} \rightarrow \mbox{Dark Higgs/Scalars} \\ \mbox{Pseudoscalar portal} \rightarrow \mbox{Axion-Like Particles} \\ \mbox{Neutrino portal} \rightarrow \mbox{Sterile Neutrinos} \end{array}$

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Different possible portals between Dark Matter and Standard Model depending on the **dark mediator X**:

Vector portal \rightarrow Dark Photon Scalar portal \rightarrow Dark Higgs/Scalars

Pseudoscalar portal \rightarrow Axion-Like Particles

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Competitive studies

with Phase 2 data!









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Signal signature:

- a single, mono-chromatic, high-E photon (ISR photon)
- a bump in the recoil mass:



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peak in E_{CMS} (horizontal band)



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Single photon trigger



Trigger logic	L1 rate at full luminosity	
$E > 1 \; GeV$	4 kHz (barrel)	
+ 2 nd cluster E $<$ 300 MeV	7 kHz (endcaps)	
$E > 2 \; GeV$	5 kHz (barrel)	
+ Bhabba & $\gamma\gamma$ vetoes		

Rates OK for Phase 2: (max 8 kHz) Max. L1 rate for Phase 3: < 30 kHz Sustainable for Phase 3?



Single photon trigger



News from the Technical Board of this mornining (by Nakazawa-san)

Trigger for Dark

- Based on ECL trigger, number of isolated cluster and energy sum triggers
 - Number of isolated clusters
 - No timing, energy, position information
 - Energy sum in lab frame
 - Tracking triggers as veto
- Activated since this Monday runs (r2001-) with prescale factor 1.

2

Bits

		Rate (Hz)
c1hie	clst=1 & energy > 1GeV	78.4
c1lume	clst=1 & energy > 3GeV	16.2
n1hie	clst=1 & energy > 1GeV & track=0	78.1
n1lume	clst=1 & energy > 3GeV & track=0	16.1
c3hie	clst=3 & energy > 1GeV	53.7
c3lume	clst=3 & energy > 3GeV	2.6
n3hie	clst=3 & energy > 1GeV & track=0	53.4
n3lume	clst=3 & energy > 3GeV & track=0	2.6

3

L1 rate = 378 Hz

B2TB 20180523, Dark Trigger

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Dark Photon: invisible decay (sensitivity)





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Dark Photon: the KLM veto









"Cone" angle between MCPhotons (with no ECLClusters associated) and KLMCluster associated



Cone angle between MCParticle and (=1) KLMClusters with no ECLClusters associated

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Dark Photon: the KLM veto



MCPhotons with >0 KLMClusters & 0 ECLClusters associated

MCPhotons with 0 ECLClusters associated

Efficiency: θ vs. φ - No ECLClusters associated

Efficiency: 0 vs. 0 - No ECLClusters associated







Strategies to measure the KLM efficiency with Phase 2 data: - mu mu (g) events

- g g events, with 1 g not fully reconstructed in ECL

A (little) concern:

quality of scintillators data during Phase 2;

N.B.: the scintillators are CRUCIAL for this measurement!



Axion-Like Particles



Axion-Like Particles (ALPs) are pseudo-scalars and couple to bosons. Unlike QCD Axions, ALPs have no relation between mass and coupling.

I will focus on the coupling to photons: $\mathcal{L} \supset -\frac{g_{a\gamma\gamma}}{4} a F_{\mu\nu} \tilde{F}^{\mu\nu} \qquad \underbrace{\text{N.B.}}_{\bullet} \tau \sim 1 \ / \ \mathbf{g_{a\gamma\gamma}}^2 \mathbf{M_a}^3$

Belle II will study the **ALP-strahlung** case (low sensitivity to photon fusion production)



Axion-Like Particles (signal)





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Magnetic monopoles





Interesting predictions* for $\mathbf{g} \sim \mathbf{e}$ and $\mathbf{m} = \mathbf{4.5} \ \mathbf{GeV}...$

Minimal magnetic charge from Dirac quantization: $g_p = 68.5e$

Lower magnetic charge is not ruled out (and not covered at ~GeV scale)

* arXiv:1707.05295

... but not-relativistic at Belle II:

- \rightarrow no $1/\beta^{\rm 2}$ term in dE/dx for magnetic charges
 - \rightarrow few hits in the CDC
 - \rightarrow needed a dedicated tracking

(also because: "non-standard" tracks)

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Other dark sector and exotic searches

Belle II

- Visible Dark Photon decays
- Off-shell Dark Photon decays
- Long-lived neutral particle decays
- Dark Scalar:
 - $e^+\,e^- \rightarrow \tau^+\,\tau^-\,S$; $S \rightarrow I^+\,I^-$
- Invisible $\Upsilon(1S)$ decays via: $\Upsilon(3S) \rightarrow \Upsilon(1S) \ \pi^+ \pi^-$

Muonic Dark Force: *e⁺ e⁻ $\rightarrow \mu^{+} \mu^{-} Z'$; Z' \rightarrow insivible e⁺ e⁻ $\rightarrow \mu^{+} \mu^{-} Z'$; Z' $\rightarrow \mu^{+} \mu^{-}$ LFV: *e⁺ e⁻ $\rightarrow e^{+} \mu^{-} Z'$; Z' \rightarrow invisible *e⁺ e⁻ $\rightarrow e^{+} \mu^{-} Z'$; Z' $\rightarrow e^{+} \mu^{-}$

> * More details later in Enrico's talk!

Thank you for your attention!

Piramide Cestia



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Electromagnetic Calorimeter (ECL)







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Beam background





Effects from beam background:

- \rightarrow degrades calorimeter resolution.
- \rightarrow radiation damage.
- \rightarrow pile-up and event size.
- \rightarrow physics background

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20

40

0

0

26

 θ_{ID}

60



Dark Photon: visible decay





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 $\mathbf{27}$





Belle II: ALPs below 200 MeV?

- For ALP masses below ~200 MeV, the decay photons are reconstructed as one ECL cluster even in offline analysis. Currently under study:
 - Untagged (electrons not seen) ALP fusion production has a much higher cross section and produces ALPs with less boost (difficult to trigger).
 - Shower shapes for merged cluster are different, MVA based reconstruction has better separation power (but events have to pass L1 trigger).
 - Pair conversion of one decay photon costs statistics, but yields a distinctive four particle final state.



Pro: resolved clusters**Con:** very low-energy photons

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