

# SVD Hit Time Calibration

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Luigi Corona

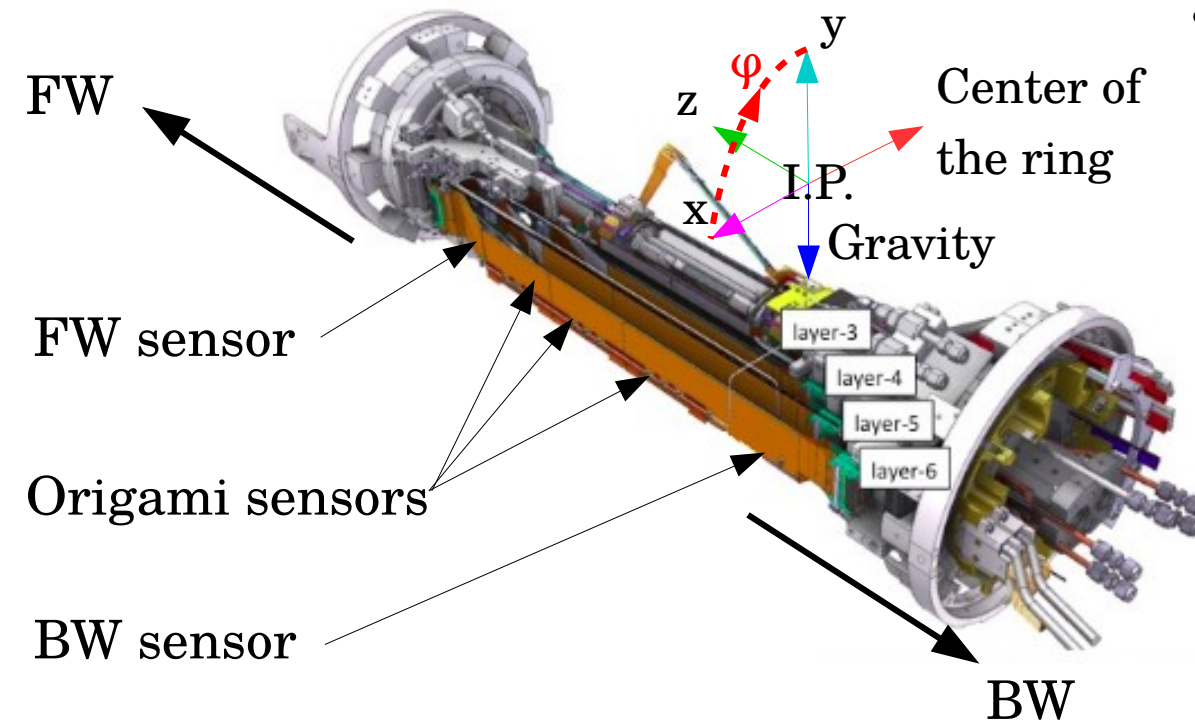
23-05-2018

# Introduction

- Hints on SVD Phase2 Configuration and signal readout
- Brief introduction on the Center of Gravity (CoG)
- How we did the CoG calibration using the timing informations of the CDC
- Final results on the resolution of the cluster time estimated by CoG
- Conclusions and future plans

# Hints on SVD Phase2 Configuration

- SVD current configuration:
  - 4 layers: L3, L4, L5, L6 (1 ladder each), positioned at  $\varphi = 0$
  - Each layer has a different number of sensors: the L4, L5 and L6 have respectively 3, 4 and 5 sensors of different types: BW + Origami and FW (slanted); the L3 has 2 smaller equal sensors

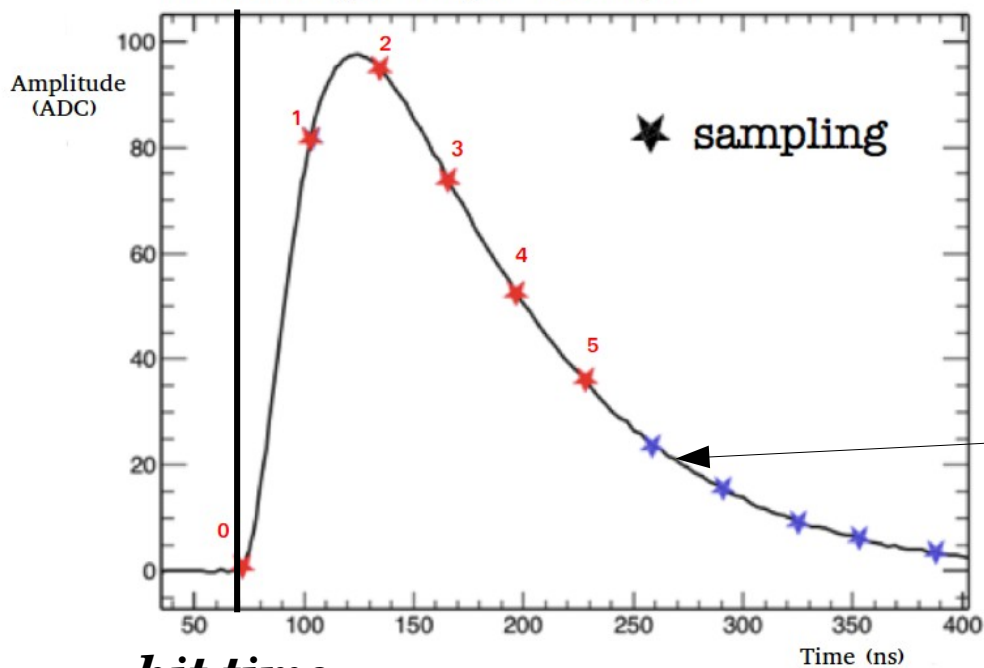


- Each sensor has 2 sides, U and V, with a different strips configuration:
  - U-side strips are of type  $p$  and measure the  $\varphi$  direction
  - V-side strips are of type  $n$  and measure the  $z$  direction

# Hints on signal readout

- The signal of the strips is read by an APV25 chip
- Each APV25 read 128 strips and sample the waveform of each strip with a clock that has a period of 31.45 ns
  - APV25 → provides 6 samples of the waveform
  - from the 6 samples we reconstruct:

APV25 Shaper Output (example)



***hit time***

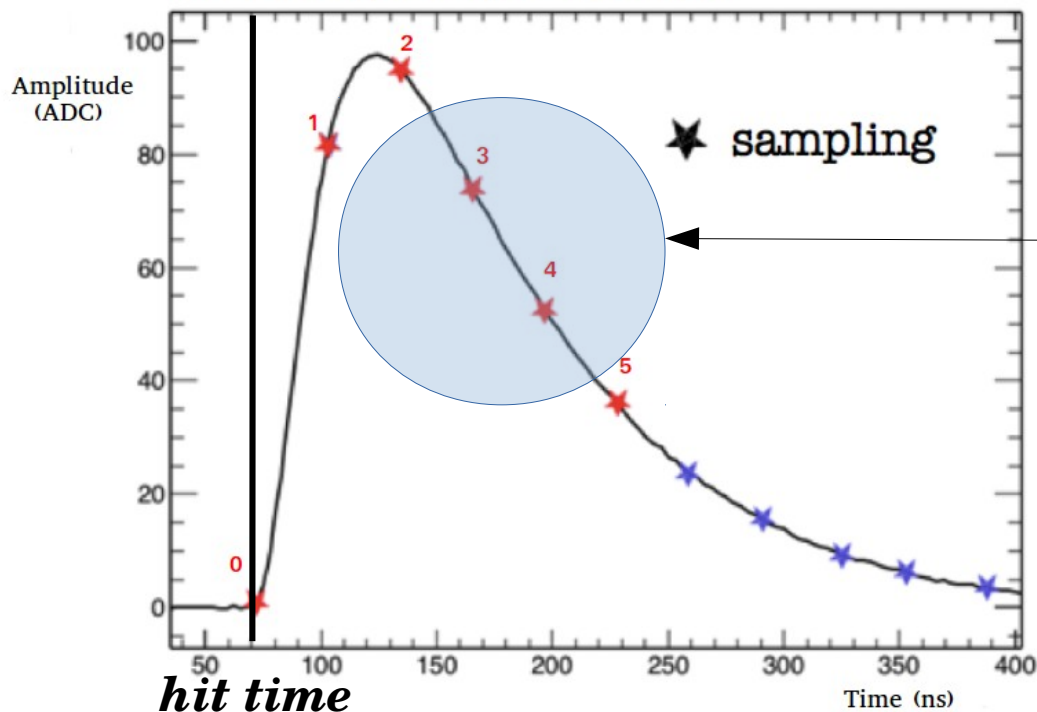
- the strip charge: the biggest charge of the 6 samples
- the strip *hit time*: it is the time at which the APV25 signal start rising

Waveform relative to a signal in a strip which is sampled by the APV25

# Center of Gravity (CoG)

- How we estimate the *hit time*? Using the **CoG**!
- The CoG is a weighted mean of the time of the 6 samples with the charge

APV25 Shaper Output (example)

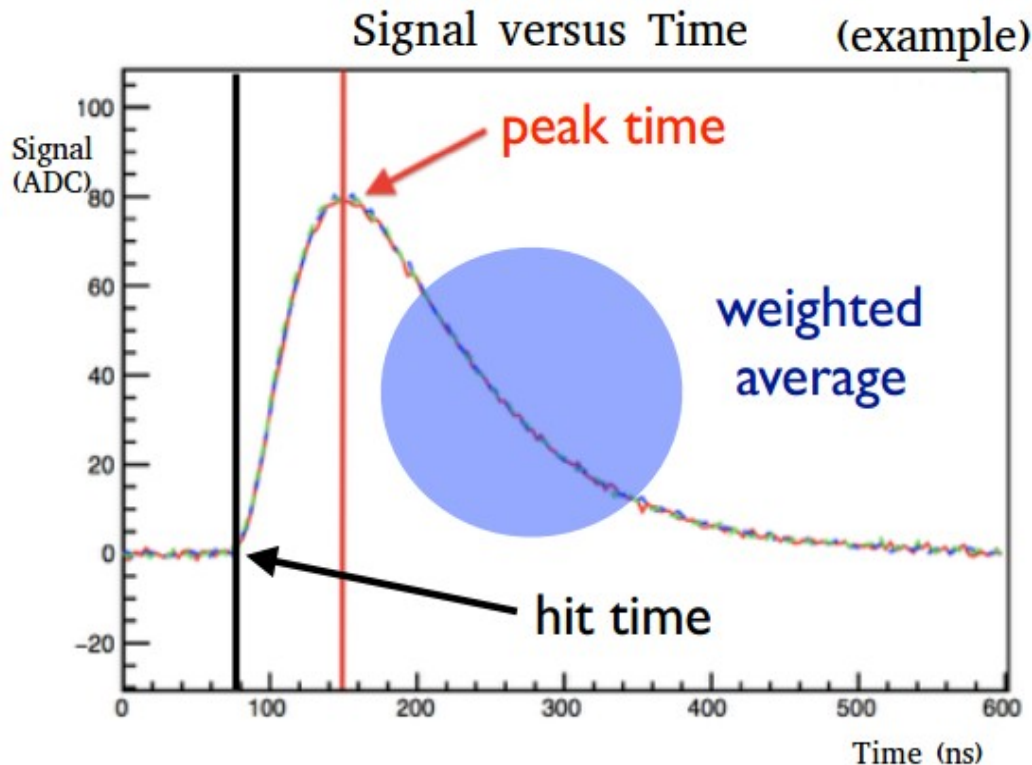


$$t_{\text{COG}}^{\text{raw}} = \frac{\sum_n A_n \cdot T_n}{\sum_n A_n}$$

- $n$ : 6 samples,  $n = [0, 1, 2, 3, 4, 5]$
- $A_n$ : Charge of each sample
- $T_n$ : Time of each sample

# Center of Gravity (CoG)

- $t_{\text{COG}}^{\text{raw}}$  estimates the *peak time*

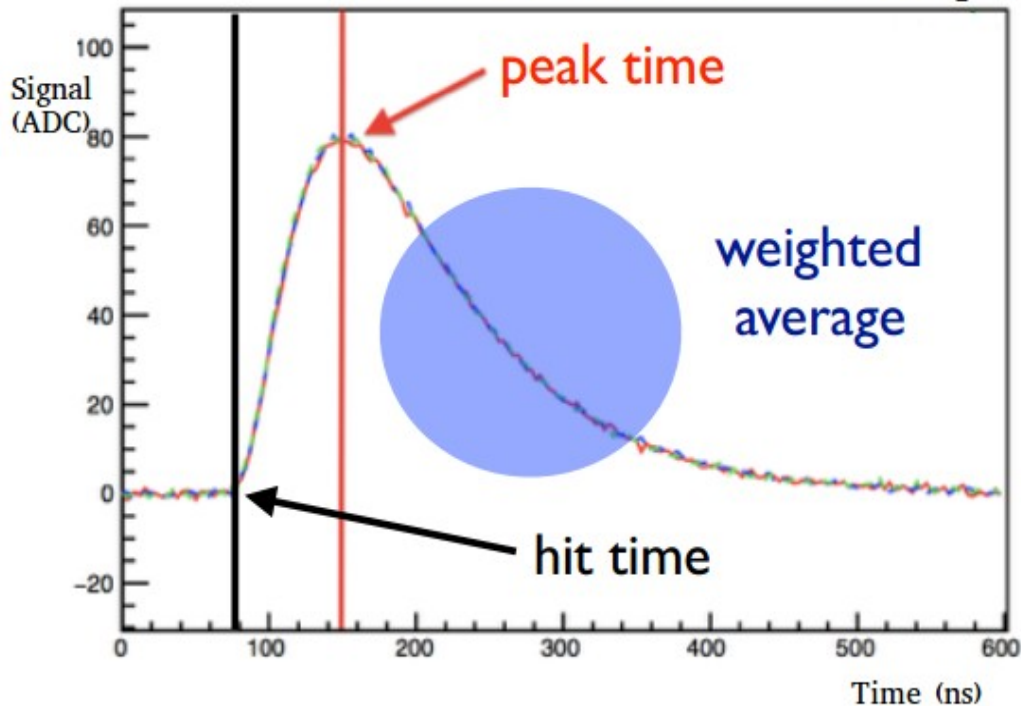


$$t_{\text{COG}}^{\text{raw}} = \frac{\sum_n A_n \cdot T_n}{\sum_n A_n}$$

- We want the *hit time*
- $t_{\text{COG}}^{\text{raw}}$  has to be corrected
- The correction that we applied is the *peaking time* correction (strip-dependent)

# Center of Gravity (CoG)

Signal versus Time (example)



$$t_{\text{COG}} = \frac{\sum_n A_n \cdot T_n}{\sum_n A_n} - p.t.$$

$t_{\text{COG}}^{\text{raw}}$

- *p.t.*  
 → *peaking time* = constants + rising time of the signal (strip-dependent) (written on DB)

- $t_{\text{COG}}$  estimates the *hit time*

# Clusters time

- SVDShaperDigits

CoG

## Selection of strips

- SNR threshold  $> 5$ : filter for noisy strips
- Strips near the tracks extrapolated on the sensors

- RecoDigits (*hit time* and amplitude informations...)

SimpleClusterizer

- Clusters (cluster dimension, cluster position, cluster time, ...)
- **Cluster time** = weighted mean of the strip times with charge

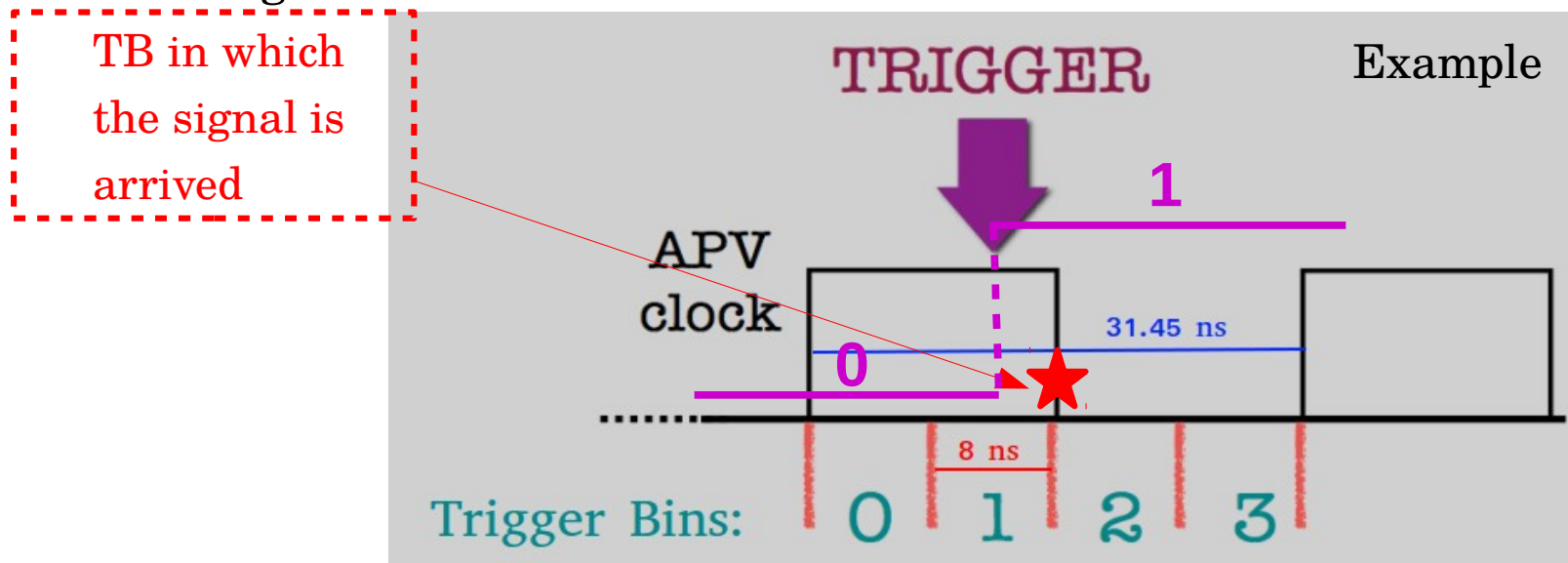


# CoG Calibration

- SVD + CDC Cosmics Run 2804 reconstruction
- Cluster time estimated by CoG (applying only the *peaking time correction*):  $t_{\text{COG}}$
- Event time-zero estimated by CDC:  $t_0$
- Scatter plot of  $t_0$  versus  $t_{\text{COG}}$  distinguishing the different Trigger Bins

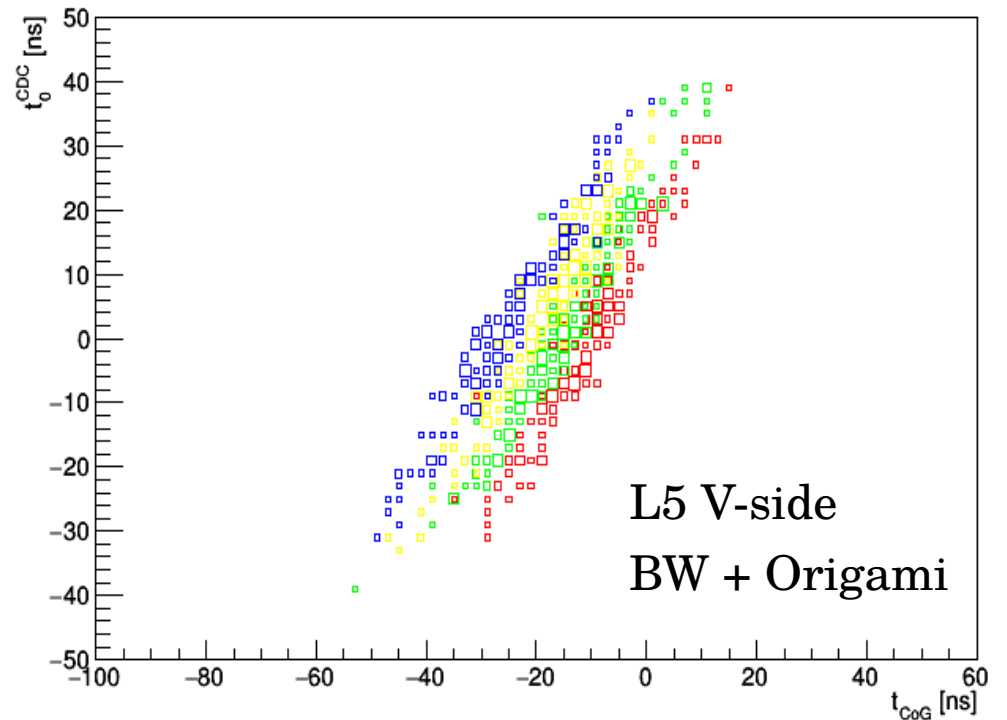
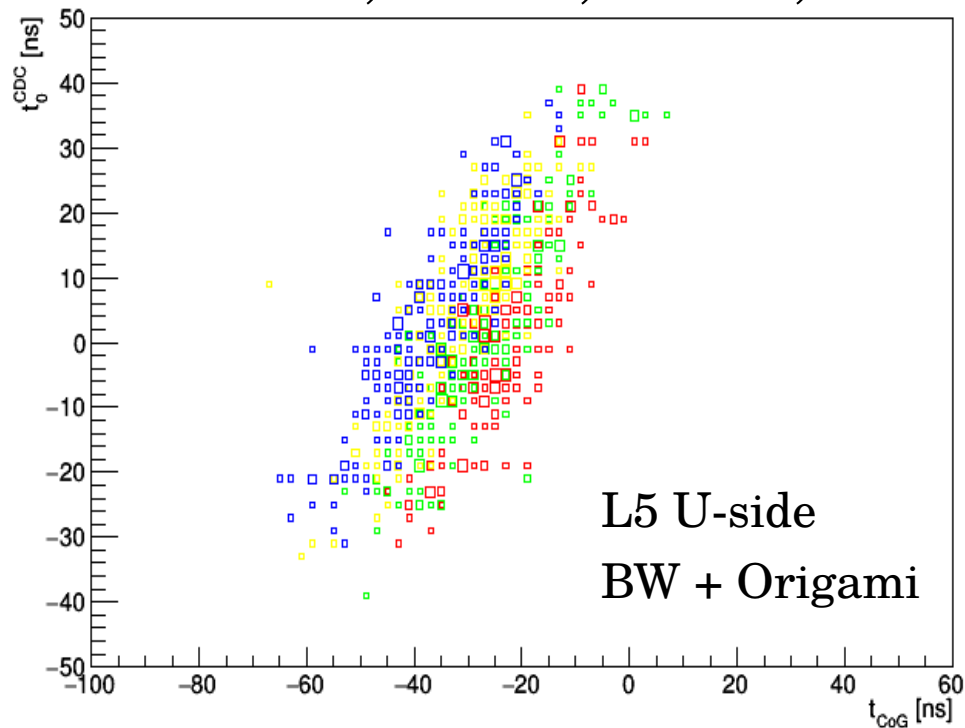
# Trigger Bins

- The APV25 clock has a period of about 31.45 ns. The “check trigger clock” has a period of about 8 ns: it is more precise than APV25 clock.
- Each Trigger Bin (TB) is the 8 ns wide window and it contains the informations about:
  - in which window the signal has arrived
  - the shift in time between the SVD origin of time (first sample) and the CDC origin of time



# CoG Calibration

- Cosmics run 2804
- $t_0$  [ns] versus  $t_{\text{CoG}}$  [ns]
- **TB = 0**, **TB = 1**, **TB = 2**, **TB = 3**

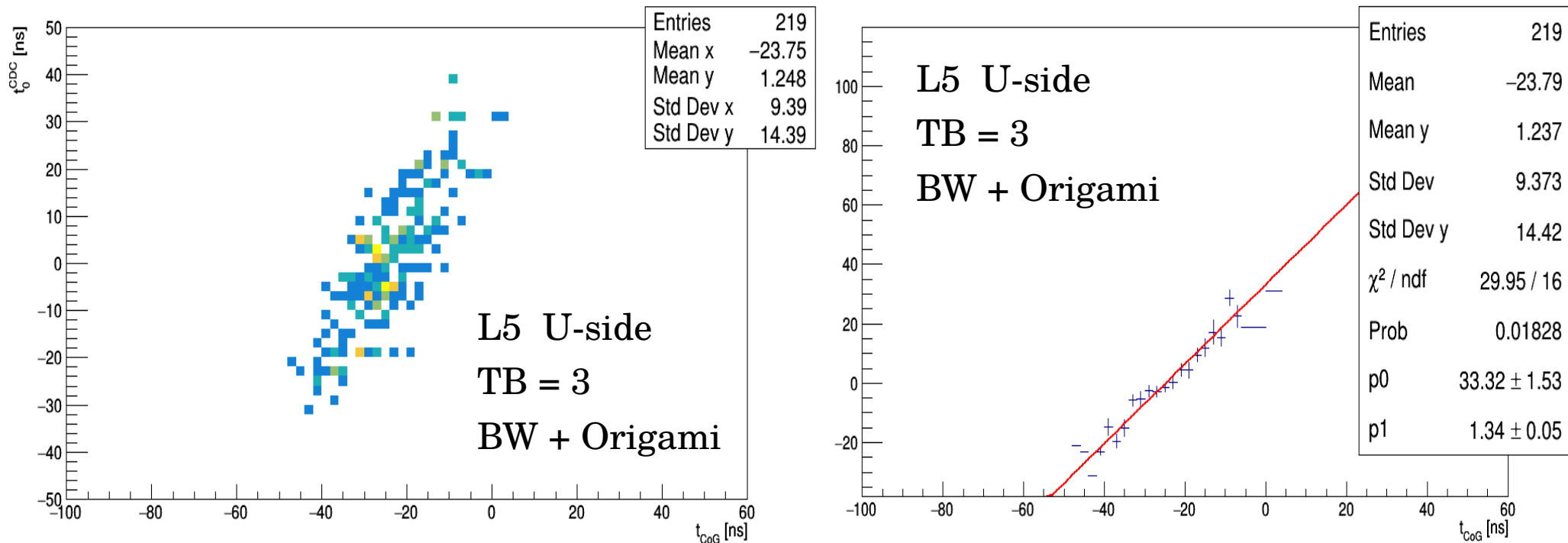


- **Linear correlation between  $t_0$  and  $t_{\text{CoG}}$ !**
- The TB are quite well separated for the V-side

# CoG Calibration

- Cosmics run 2804
- $t_0$  [ns] versus  $t_{\text{CoG}}$  [ns]
- left: Scatter plot, right: linear fit of the ProfileX of the scatter plot

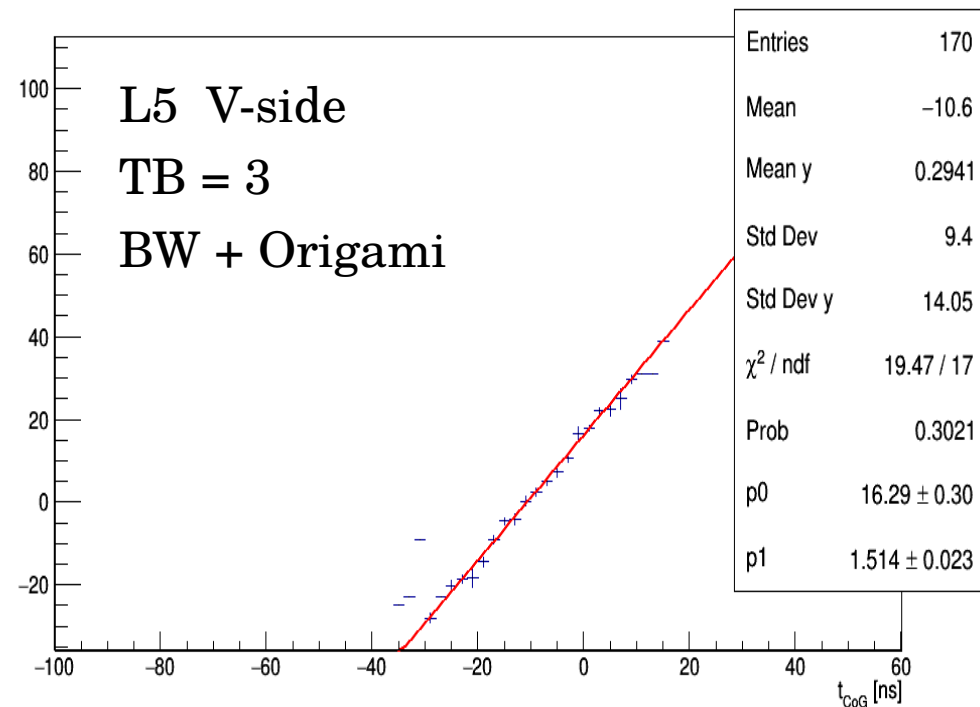
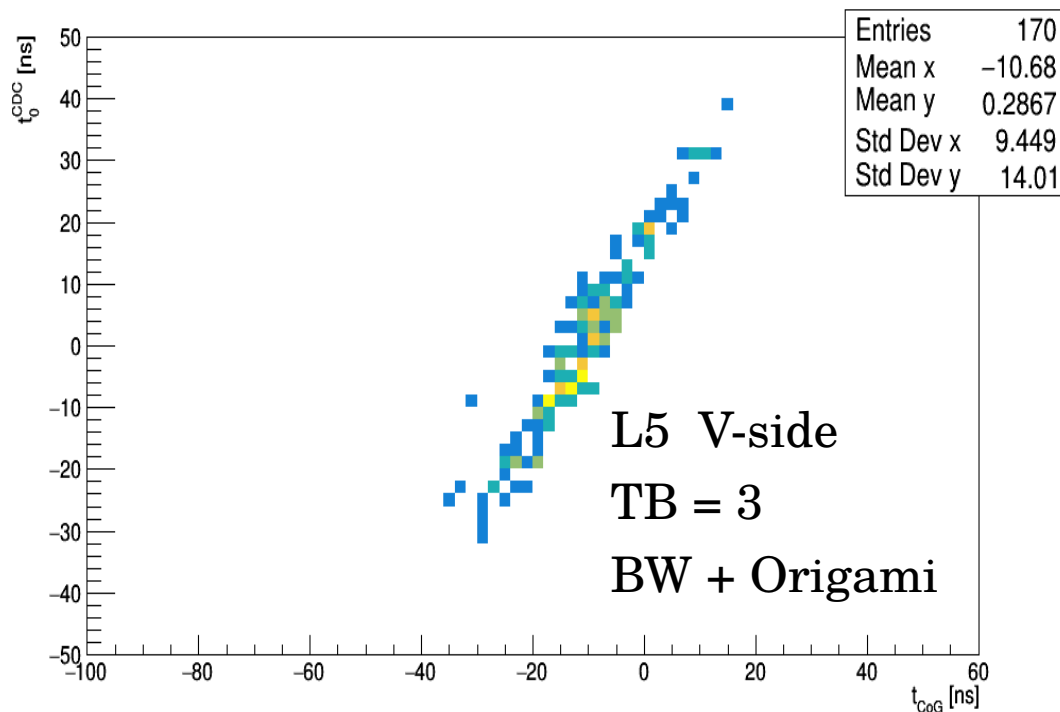
- $f(x) = m \cdot x + q$



# CoG Calibration

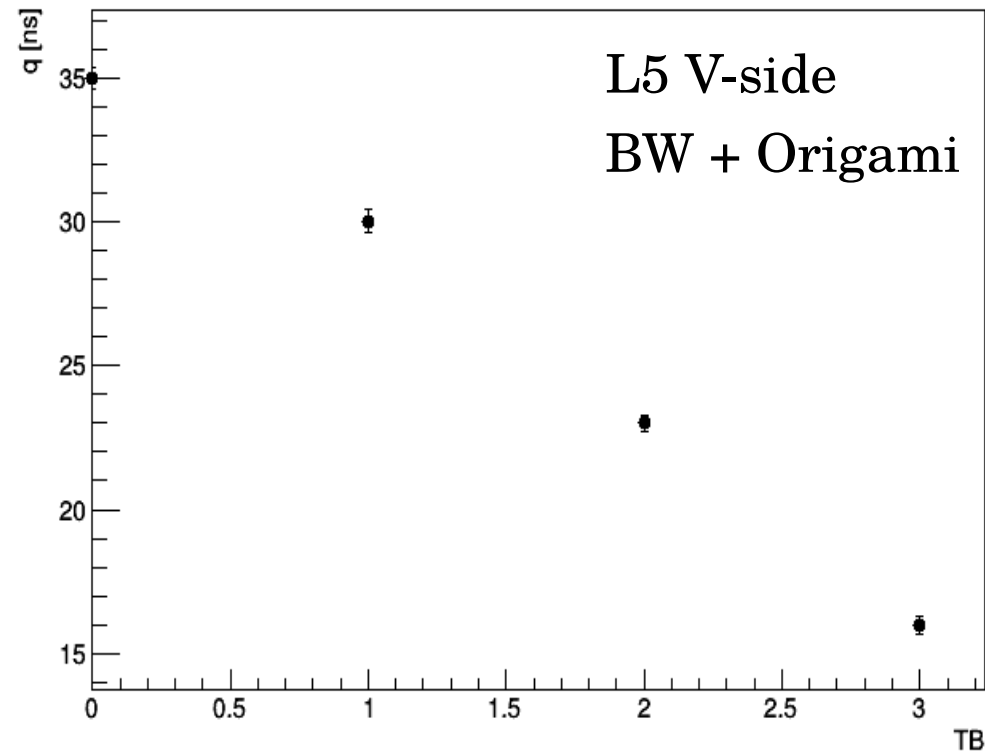
- Cosmics run 2804
- $t_0$  [ns] versus  $t_{\text{CoG}}$  [ns]
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- $f(x) = m \cdot x + q$

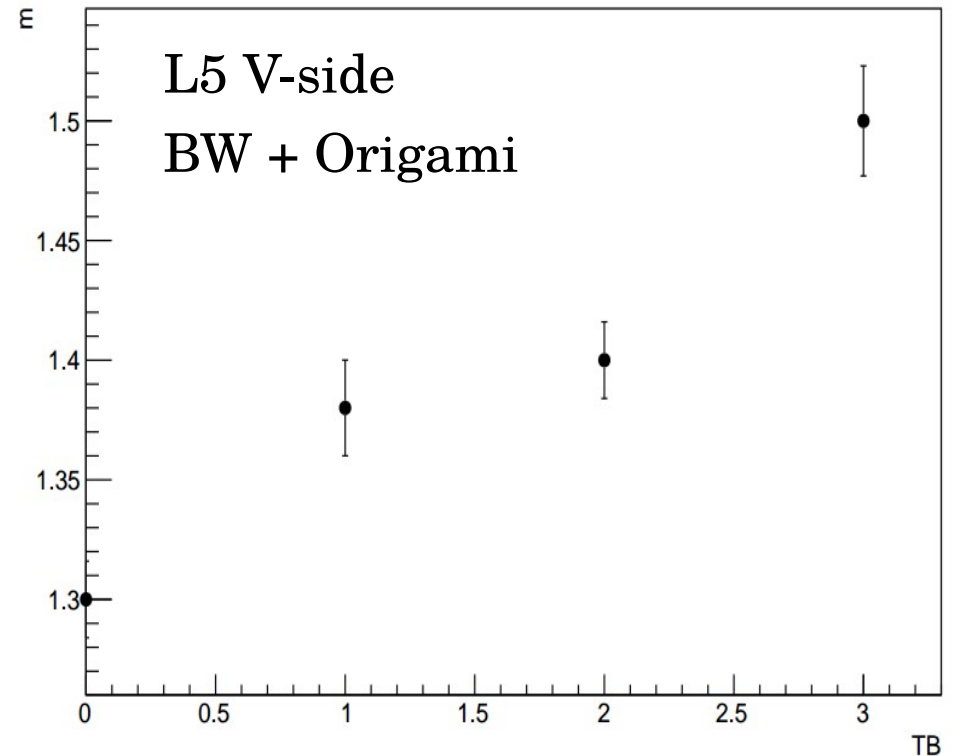


# CoG Calibration

Intercept  $q$  versus TB



Slope  $m$  versus TB

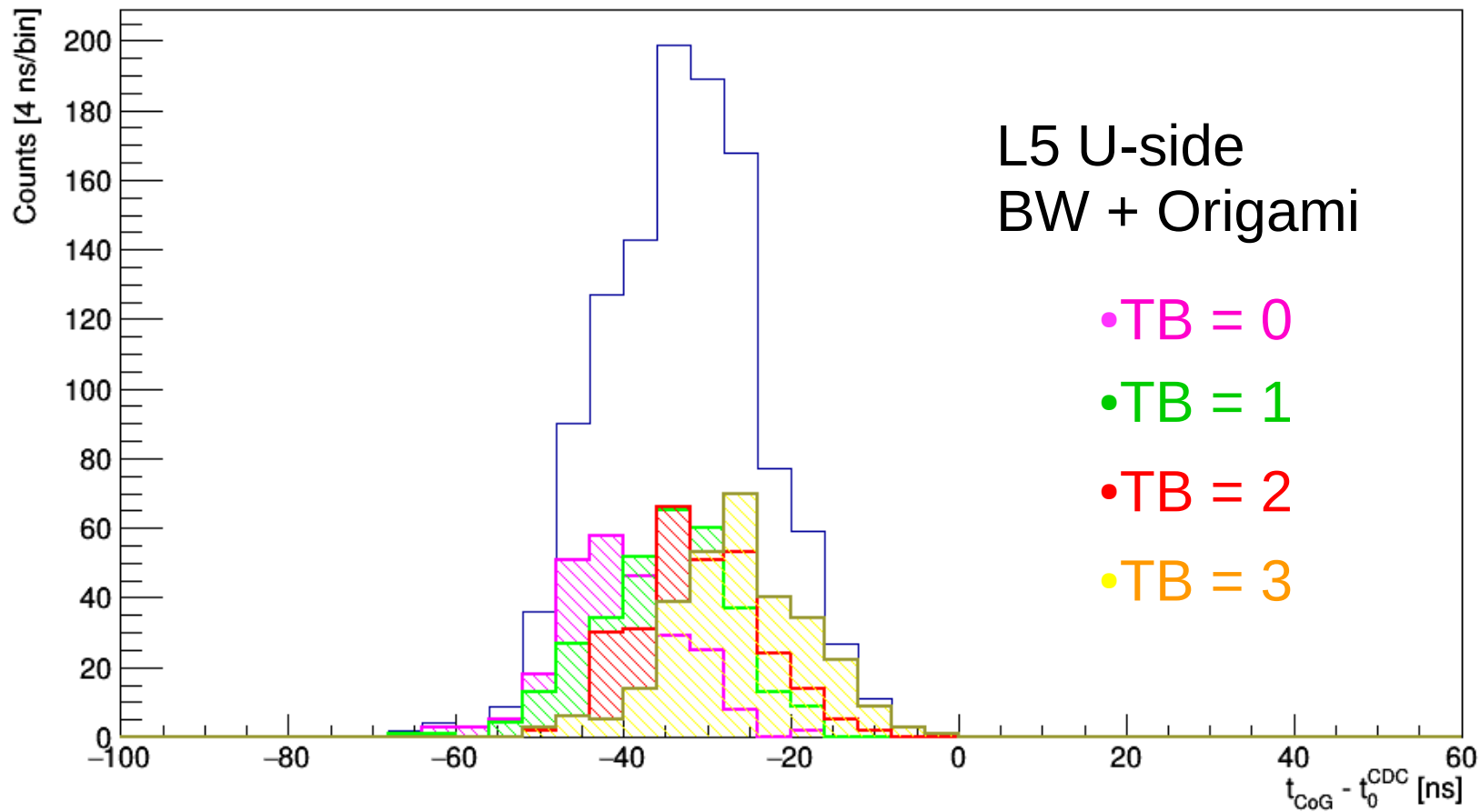


- The slope  $m$  and the intercept  $q$  are different for each TB
- General trend that we expect if the linear fit is good

# CoG Calibration

- Time resolution defined as:  $\Delta(T) = t_{\text{CoG}} - t_0$

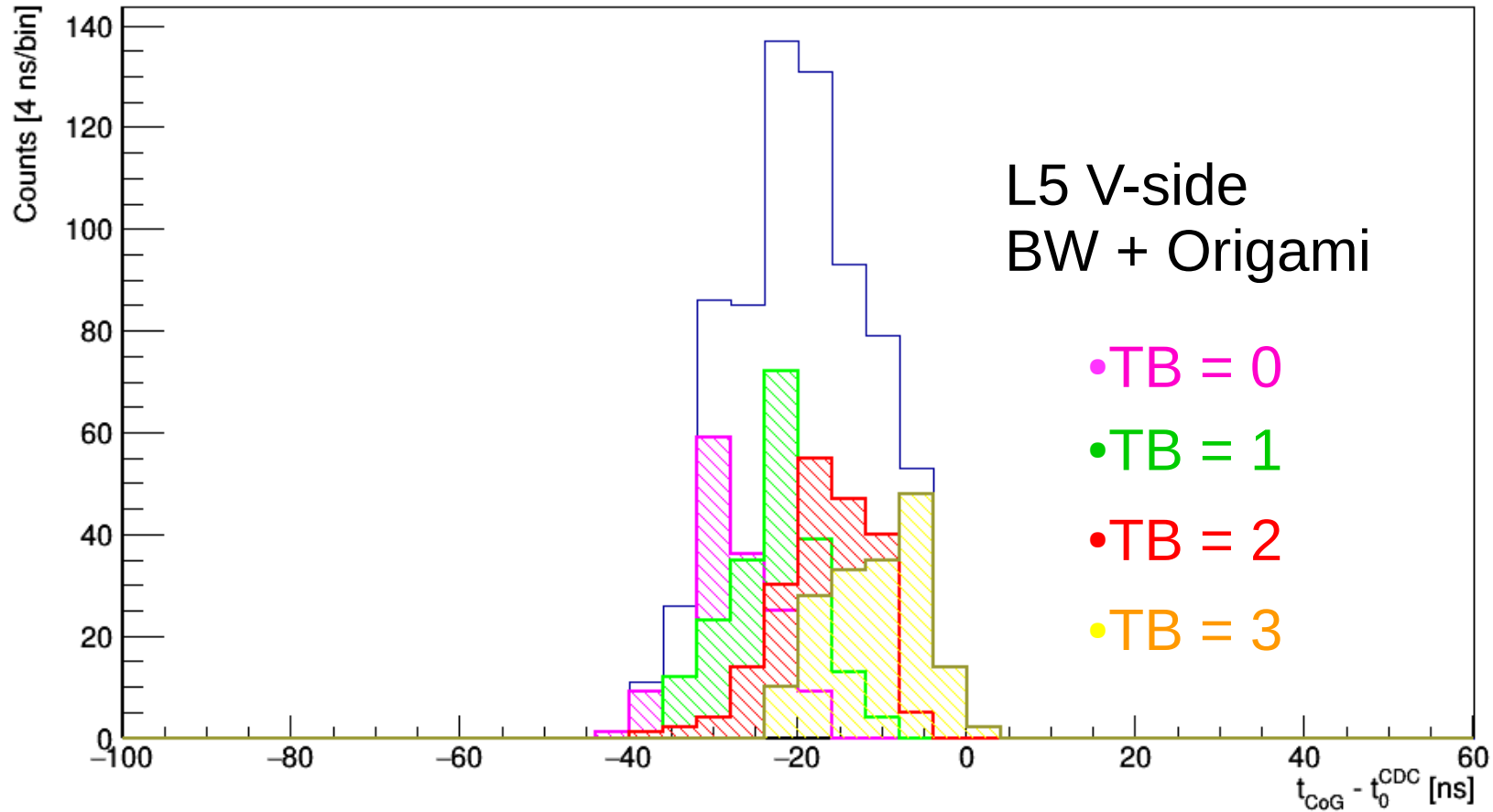
Before the correction



# CoG Calibration

- Time resolution defined as:  $\Delta(T) = t_{\text{CoG}} - t_0$

Before the correction





# CoG Calibration

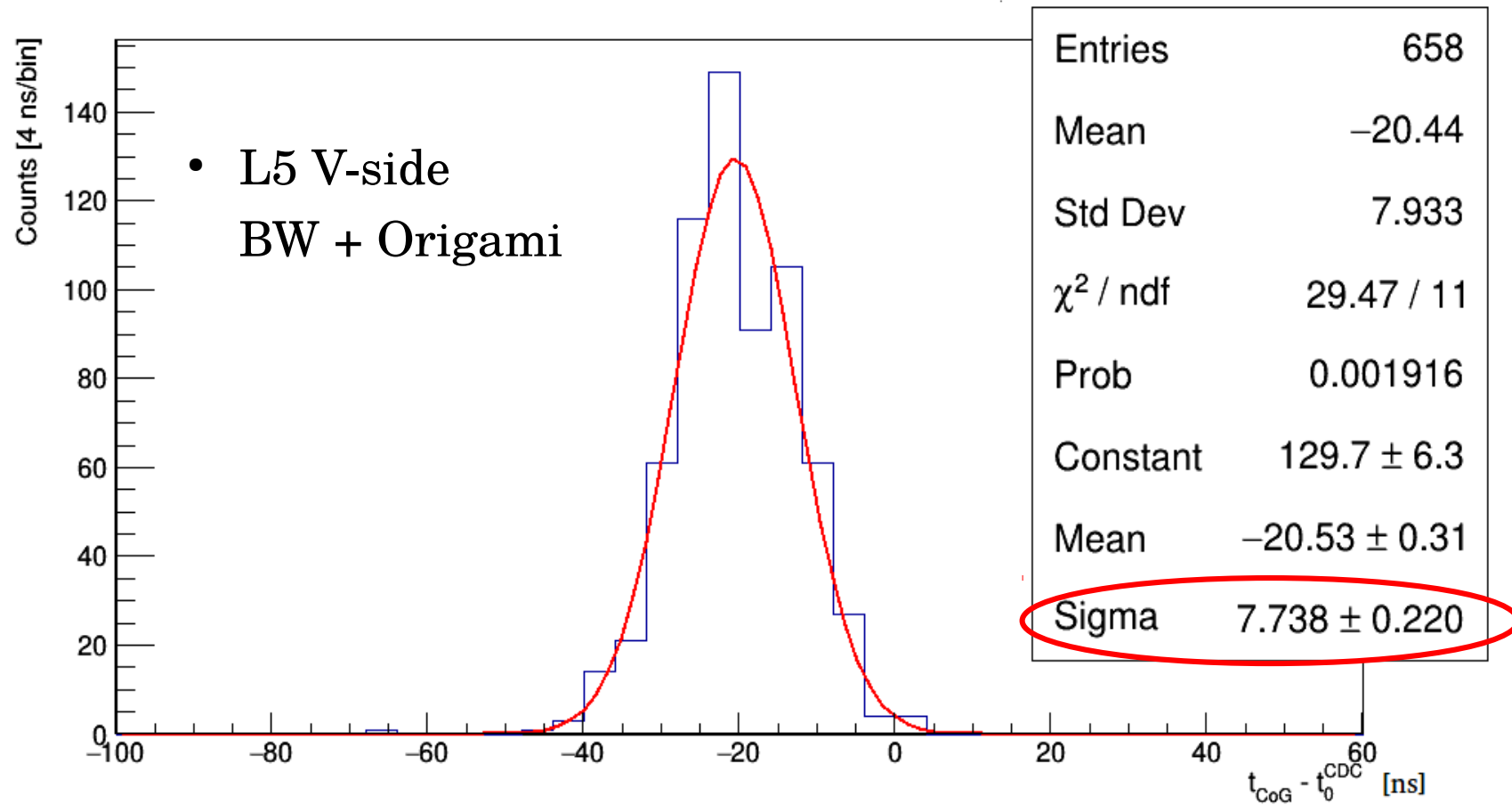
- Correction to the CoG obtained by the linear fit of the ProfileX of the scatter plot  $t_0$  versus  $t_{\text{COG}}$ :

$$t'_{\text{COG}} = m \cdot t_{\text{COG}} + q$$

- $m$  and  $q$  have been obtained from the Cosmics run 2804 (reference run) and they have been used to correct the CoG in the Cosmics run 2712 (a generic run)
- Correction has been applied to all layers and all sensors (except the FW for Layers 4, 5 and 6) and to both sensors of the Layer 3

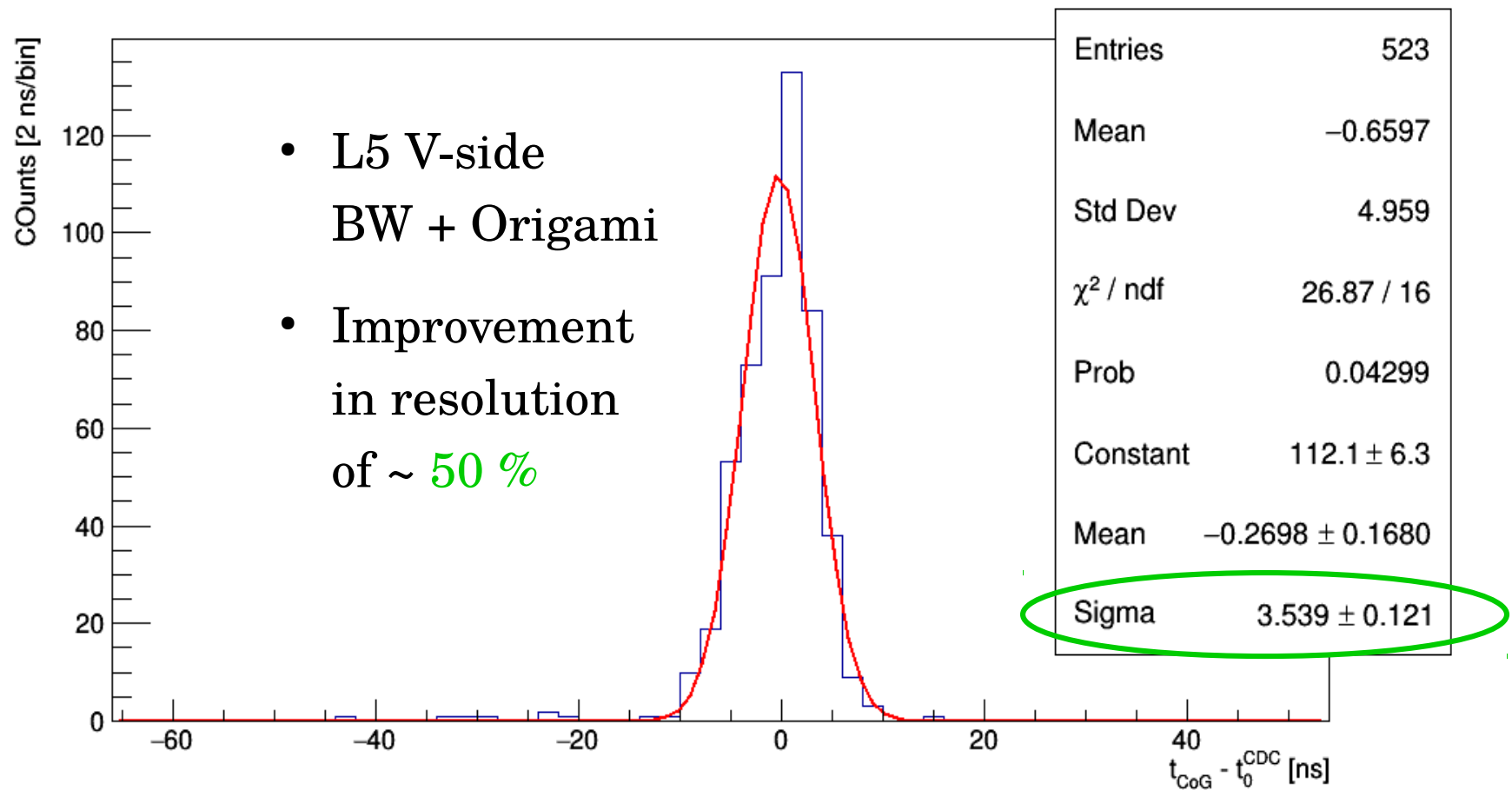
# Results

- Run 2712, CoG resolution before correction



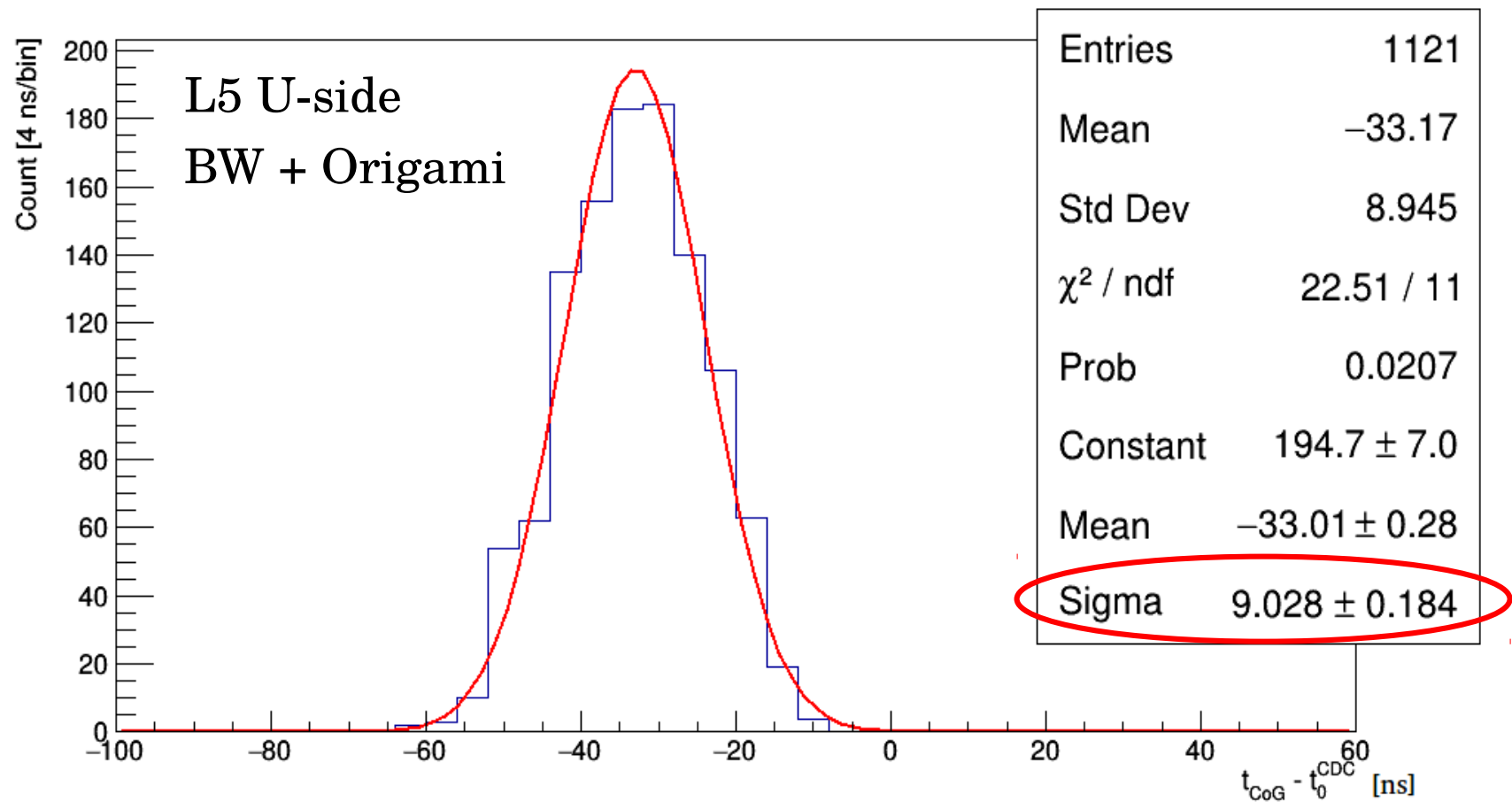
# Results

- Run 2712, CoG resolution after correction using the parameters obtained by run 2804



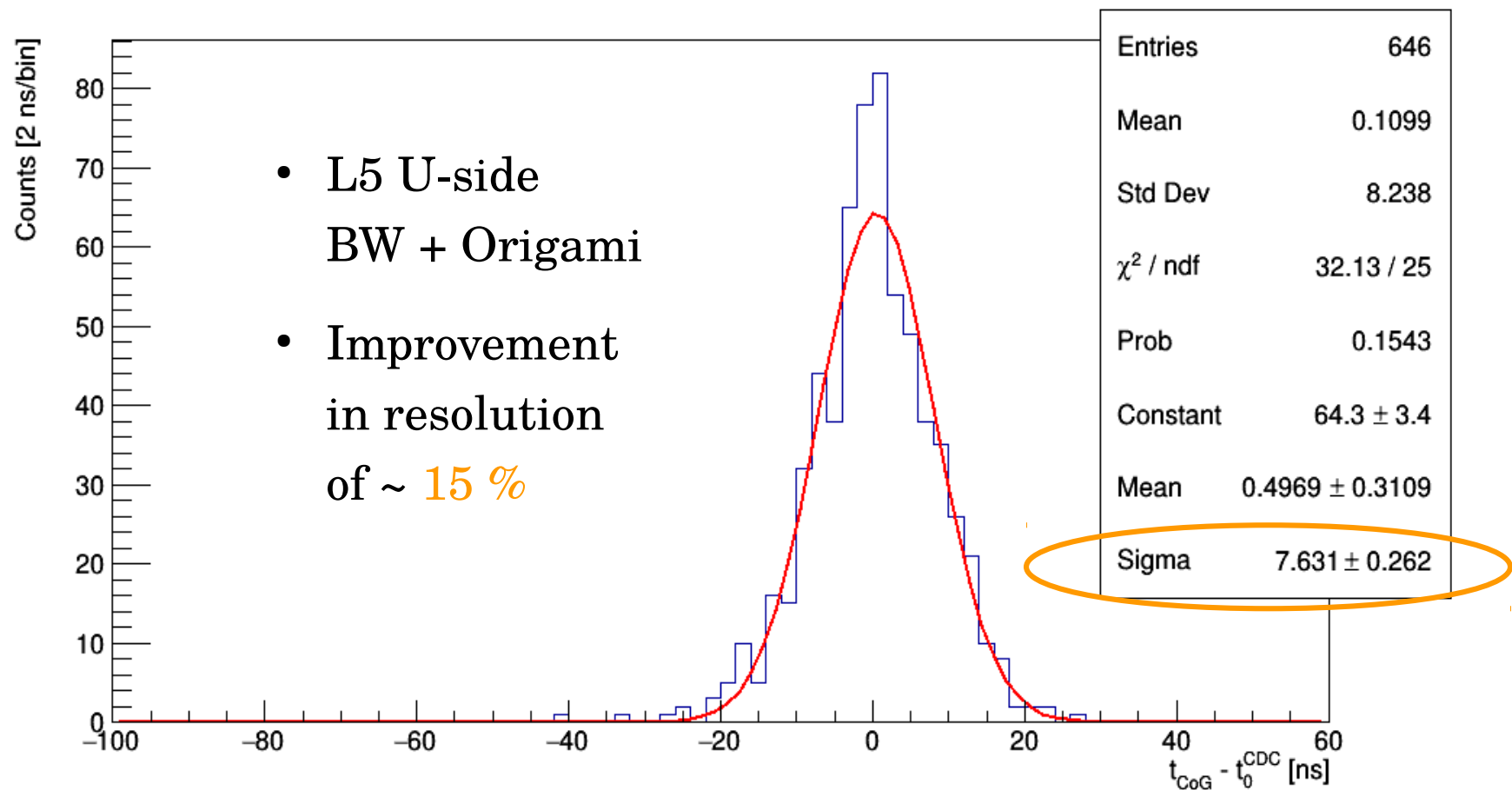
# Results

- Run 2712, CoG resolution before correction



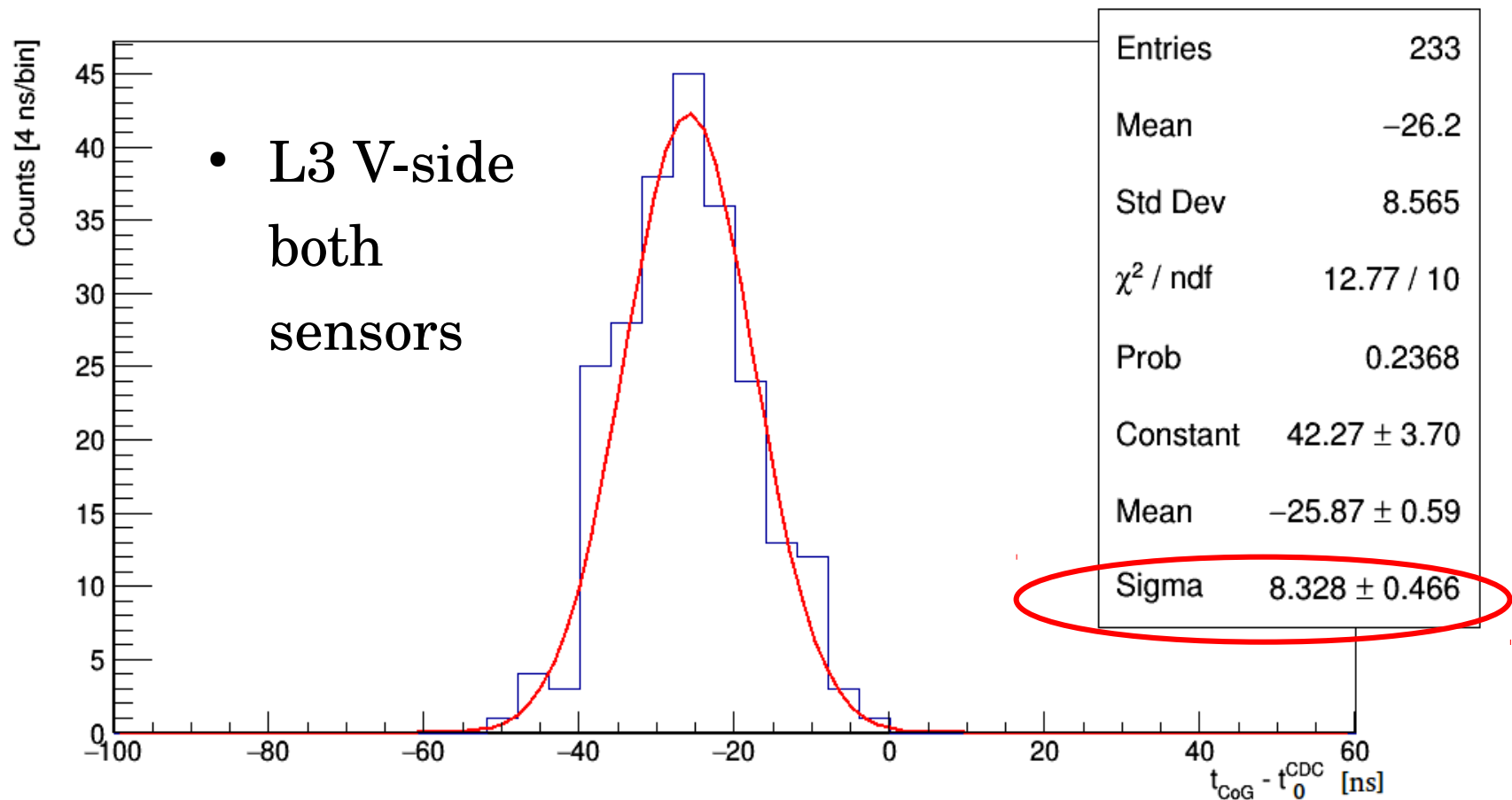
# Results

- Run 2712, CoG resolution after correction applied using the parameters obtained by the run 2804



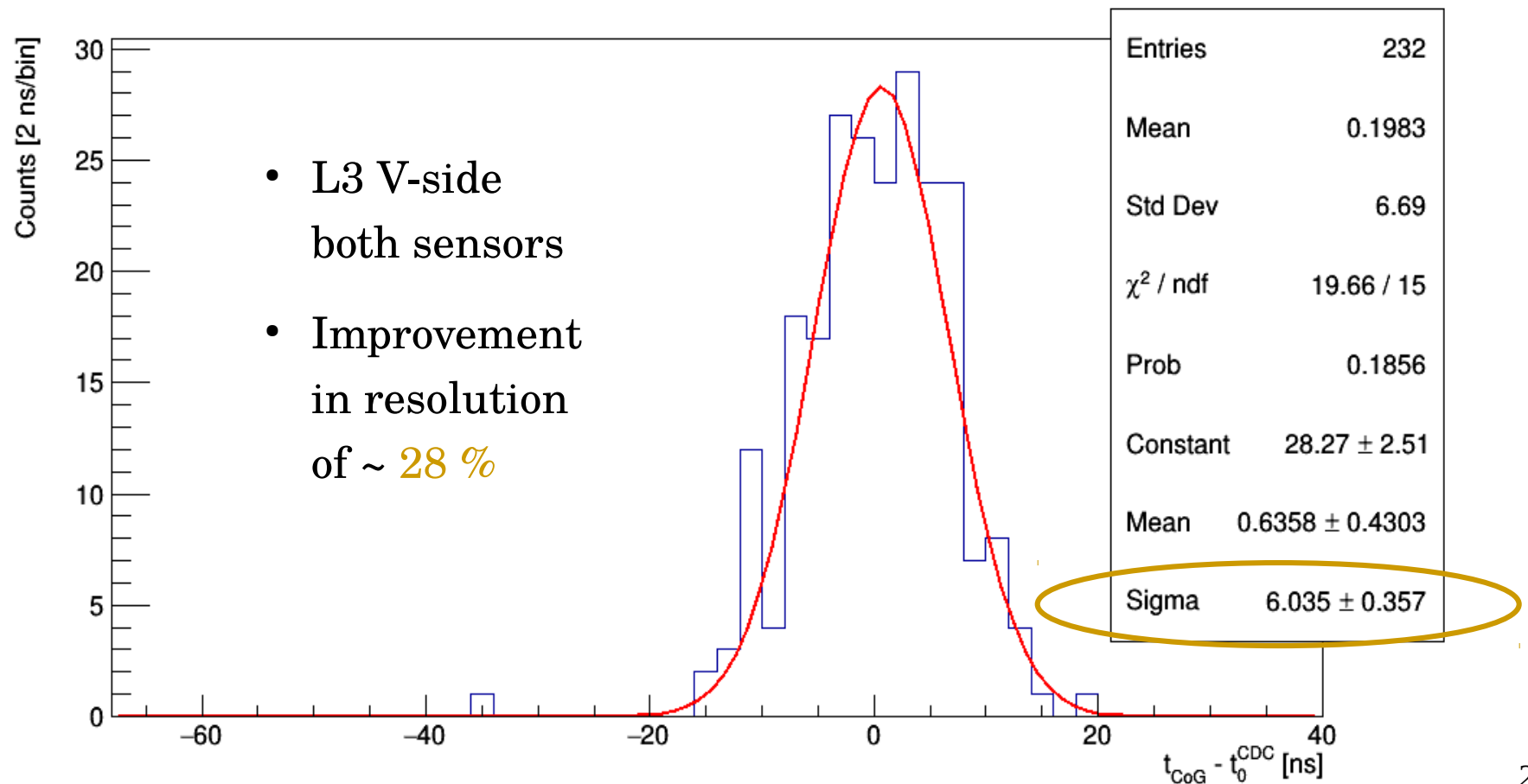
# Results

- Run 2712, CoG resolution of L3 V-side (both sensors) before correction



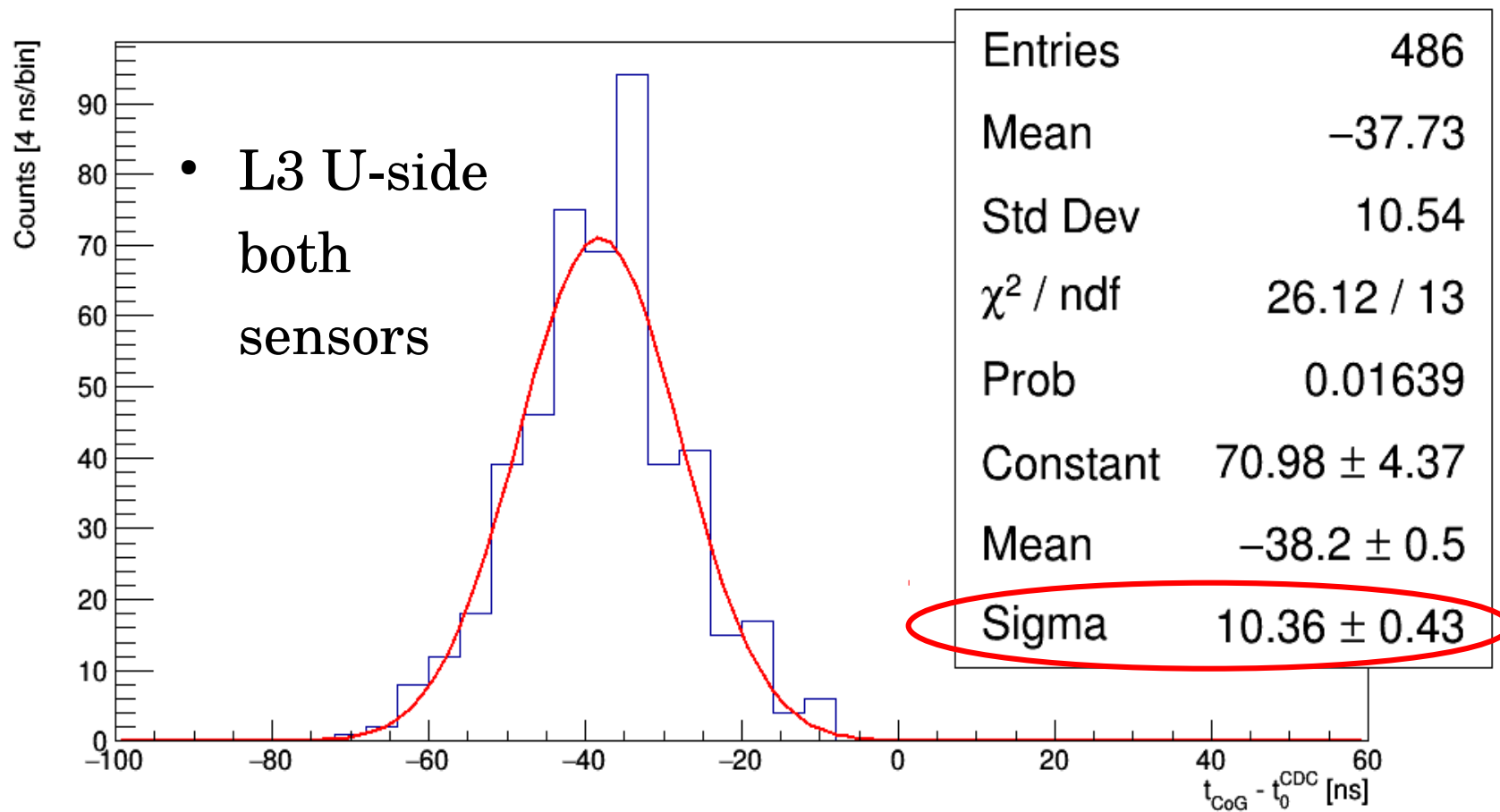
# Results

- Run 2712, CoG resolution of L3 V-side (both sensors) after correction using the parameters obtained by run 2804



# Results

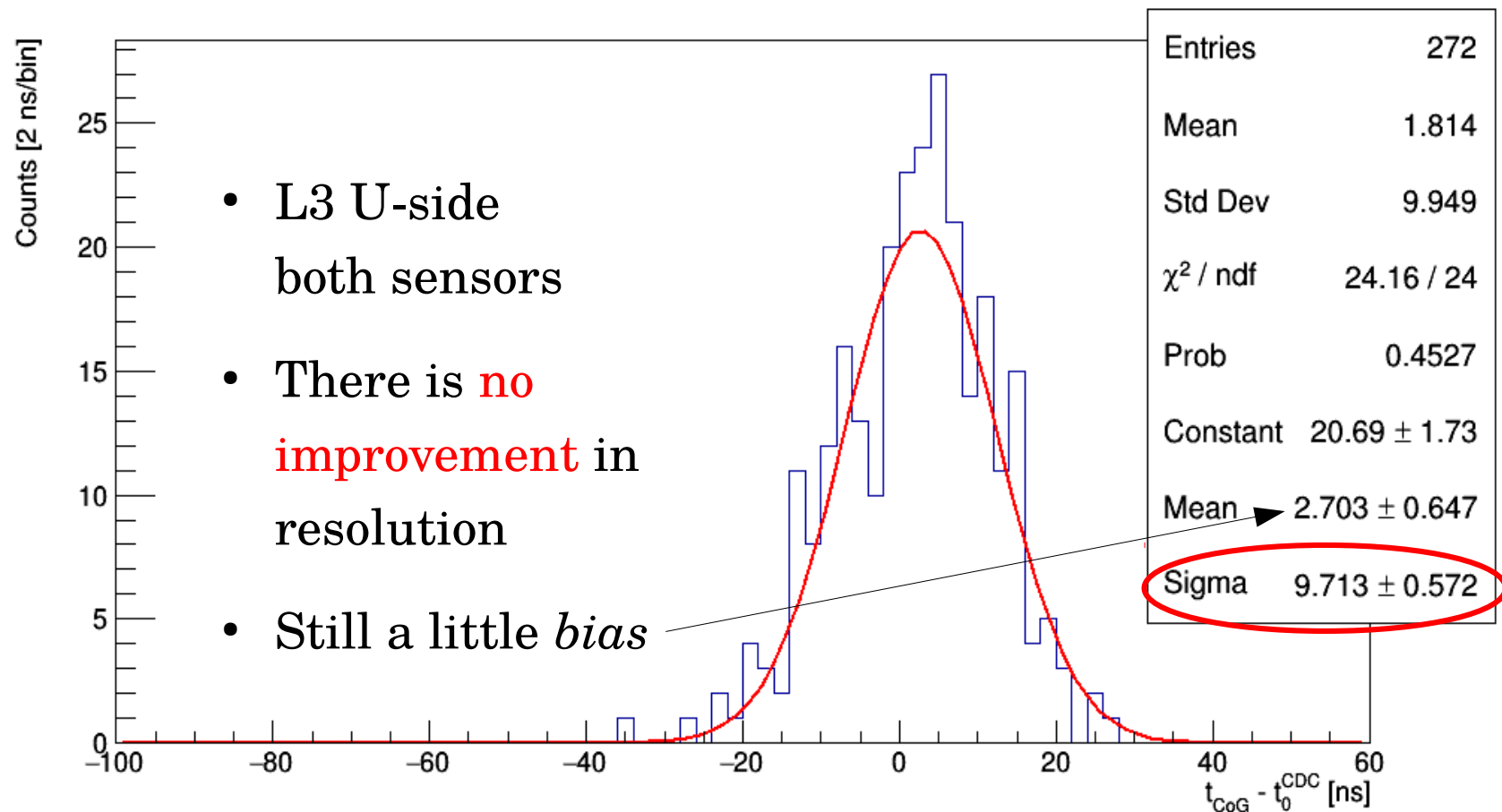
- Run 2712, CoG resolution of L3 U-side (both sensors) before correction





# Results

- Run 2712, CoG resolution of L3 U-side (both sensors) after correction using the parameters obtained by run 2804



# Conclusions

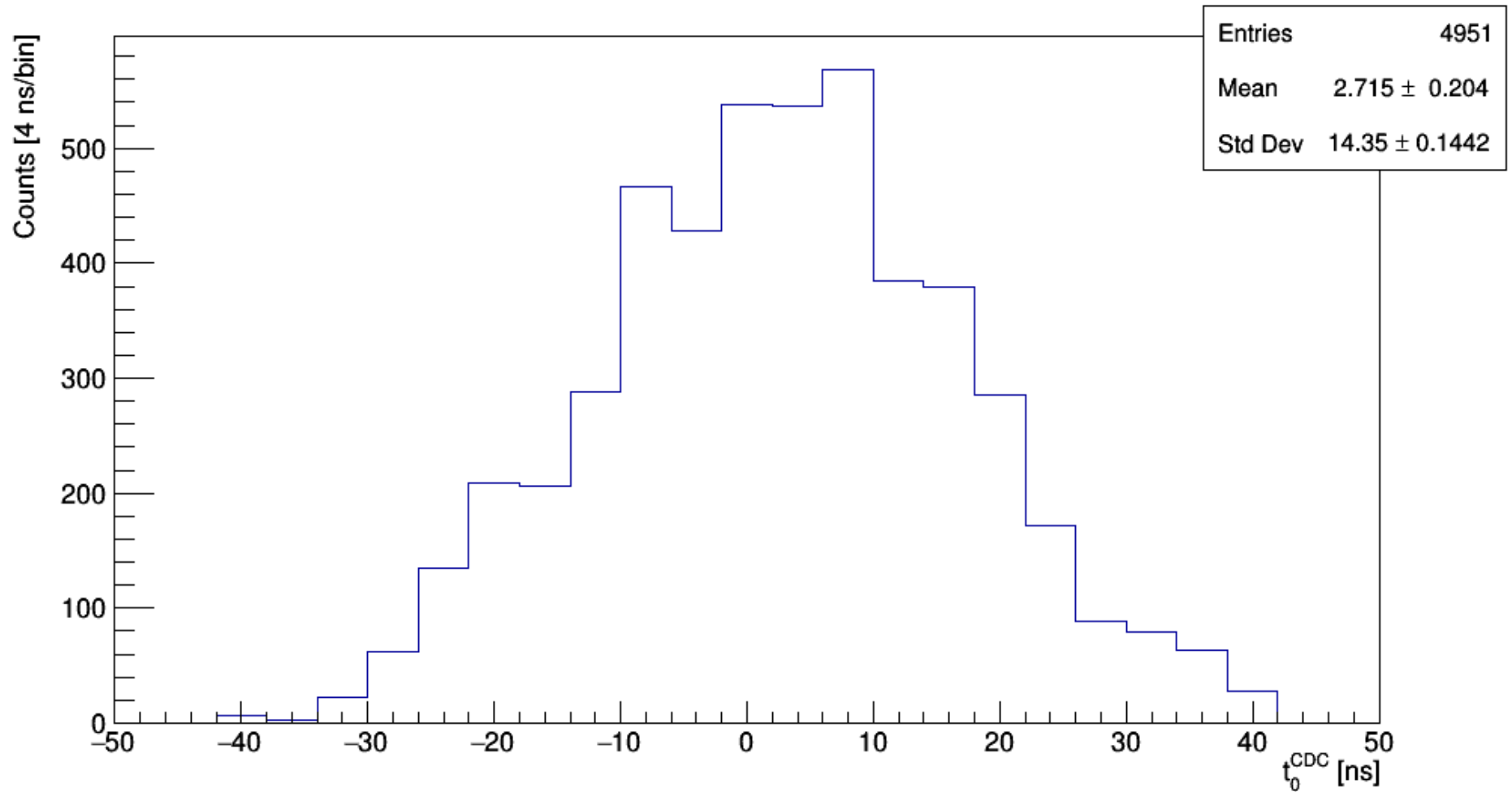
- We observe a bigger improvement of resolution in V-side than in U-side as expected
- The improvement in CoG resolution in the V-side is quite good,  $\sim 3.5$  ns for the L5 V-side and  $\sim 3.6$  ns for the L6 V-side (comparable with what we obtained from the test beam), indeed the resolution is convolved with the CDC resolution ( $\sim 2$  ns)

# Plan for the future

- Repeat the same studies for the collision runs, considering the clusters associated to the tracks

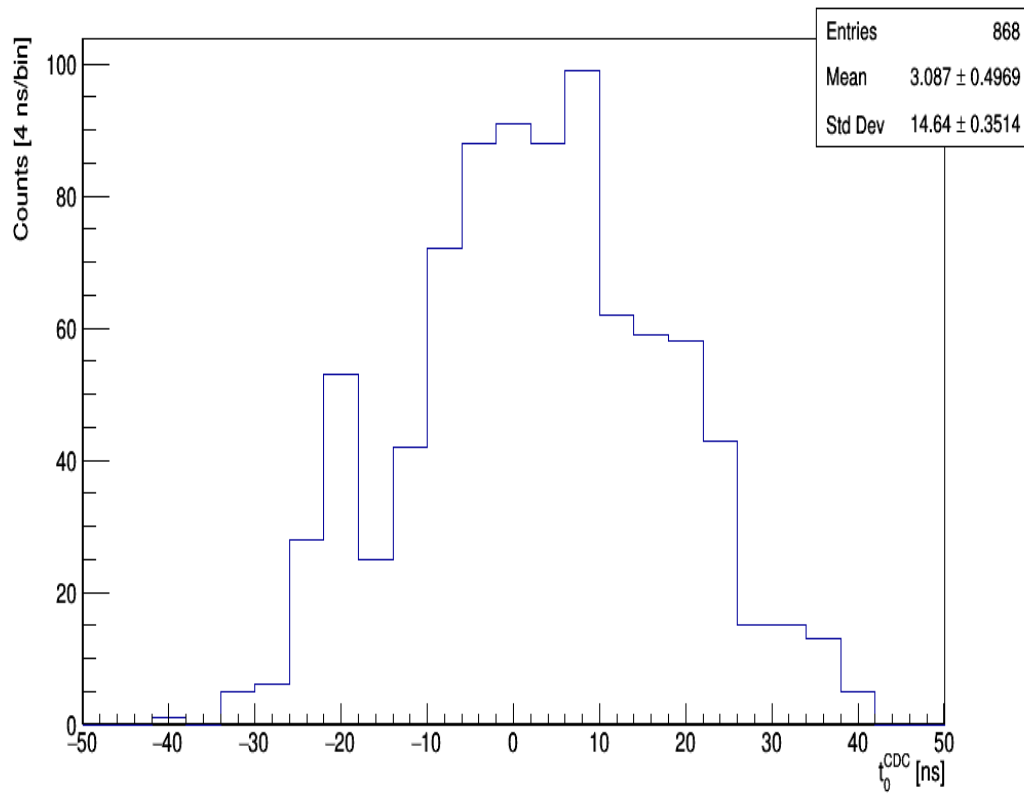
# Backup slides

# $t_0$ distribution

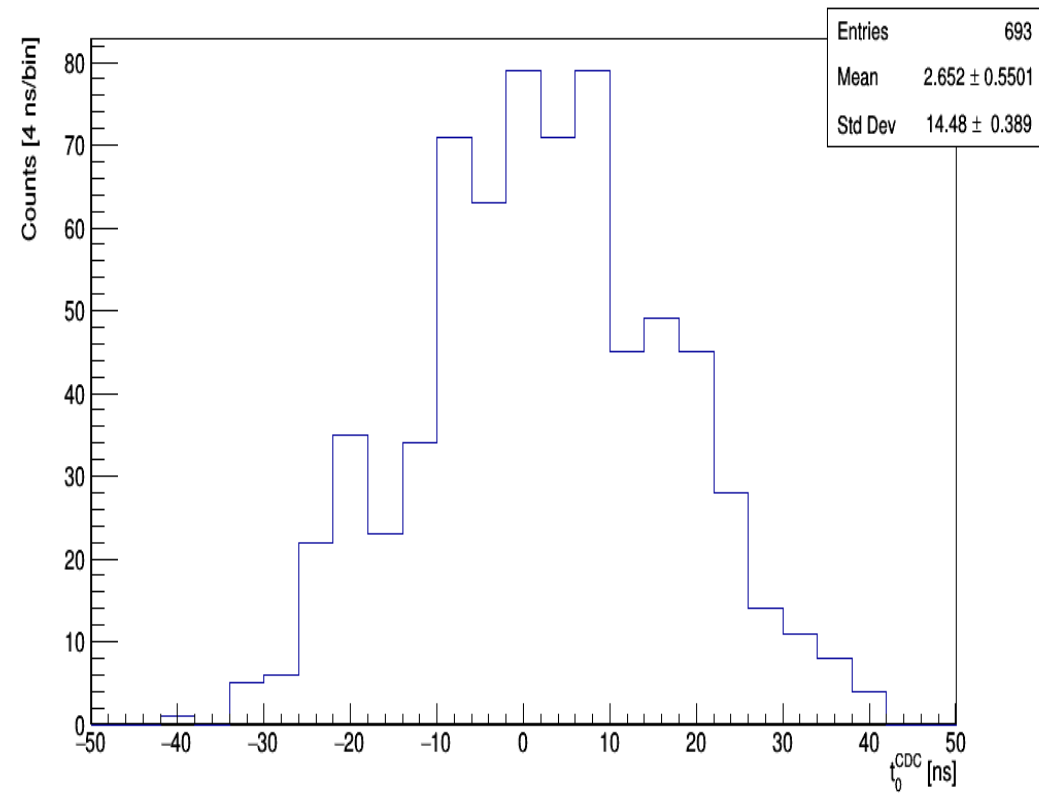


# $t_0$ distribution

L5 U-side, BW+ Origami

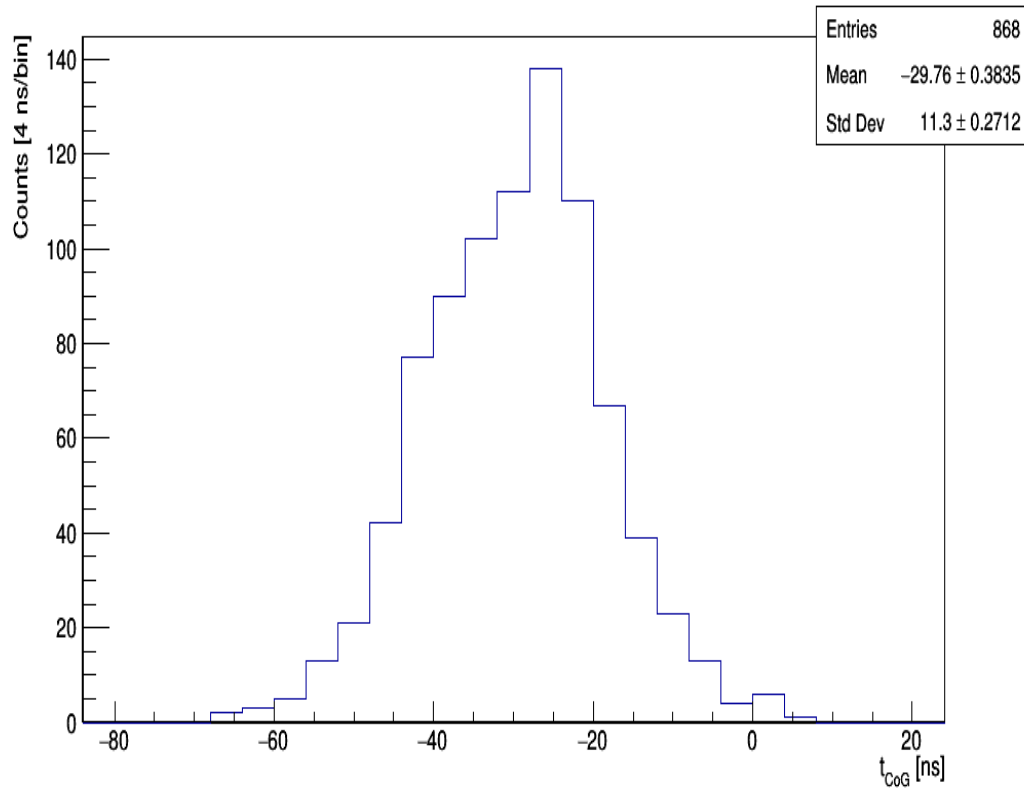


L5 V-side, BW+ Origami

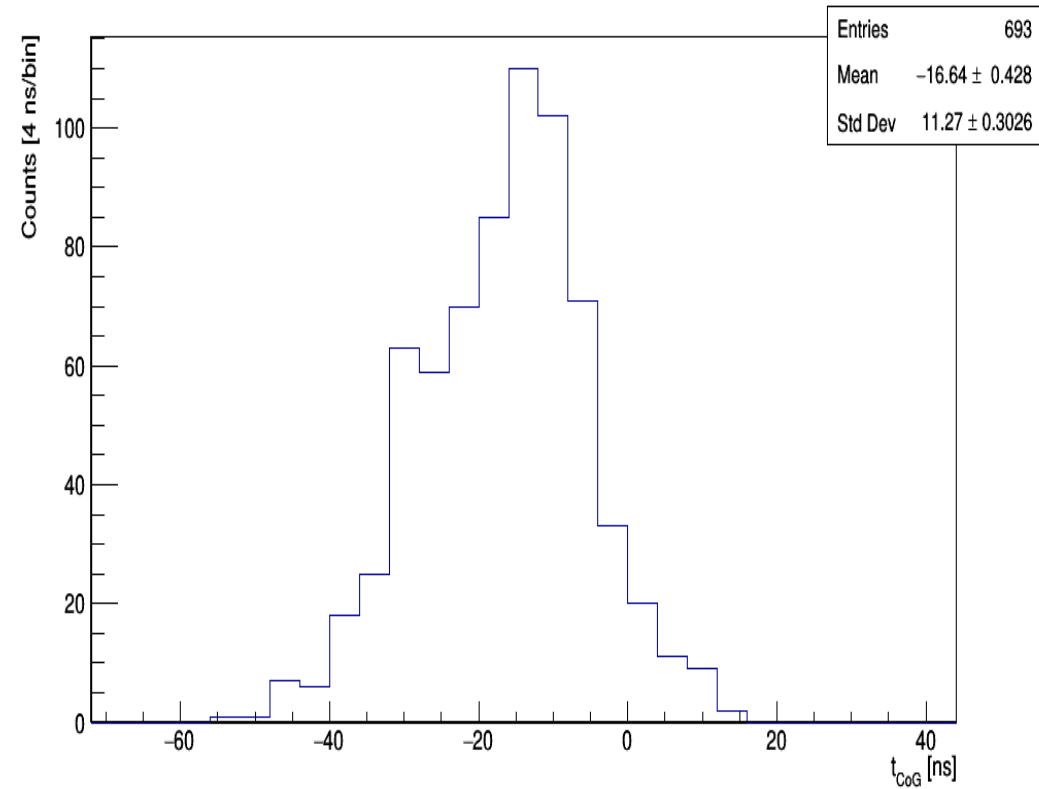


# $t_{\text{COG}}$ distribution

L5 U-side, BW+ Origami

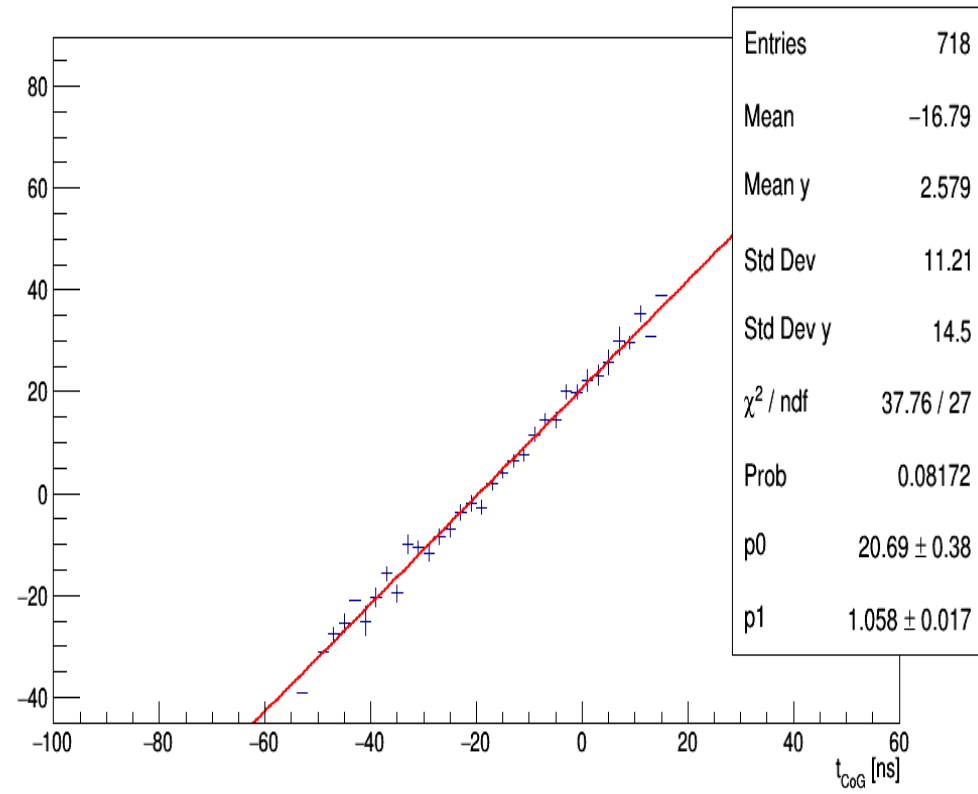
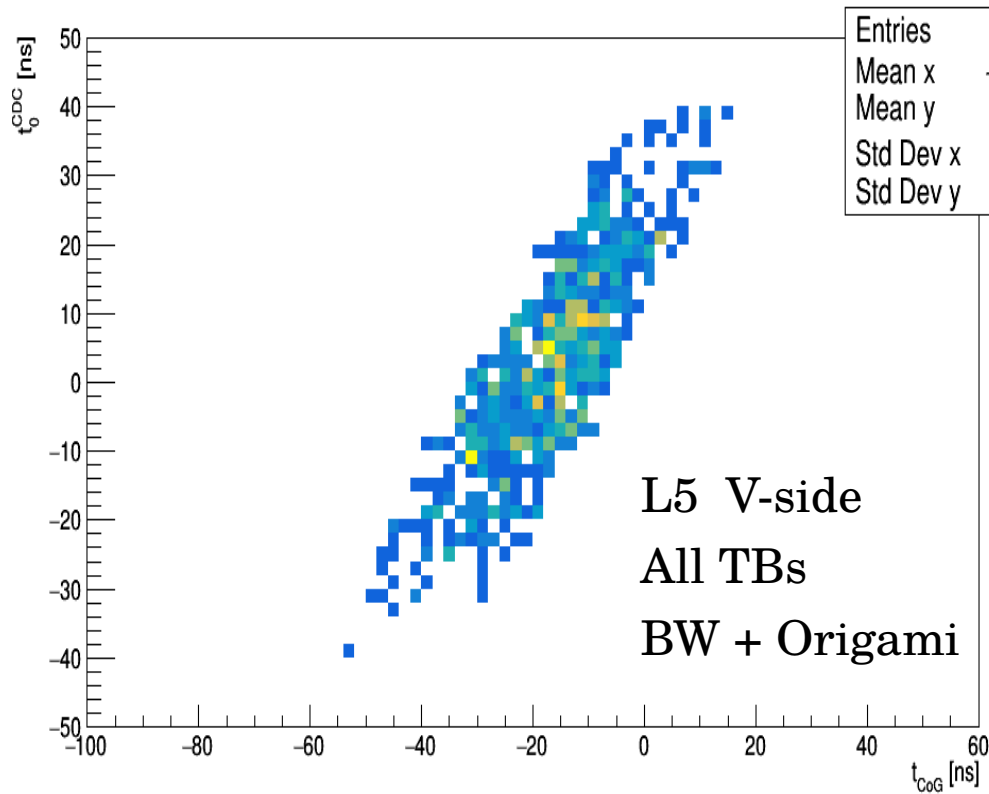


L5 V-side, BW+ Origami



# CoG Calibration

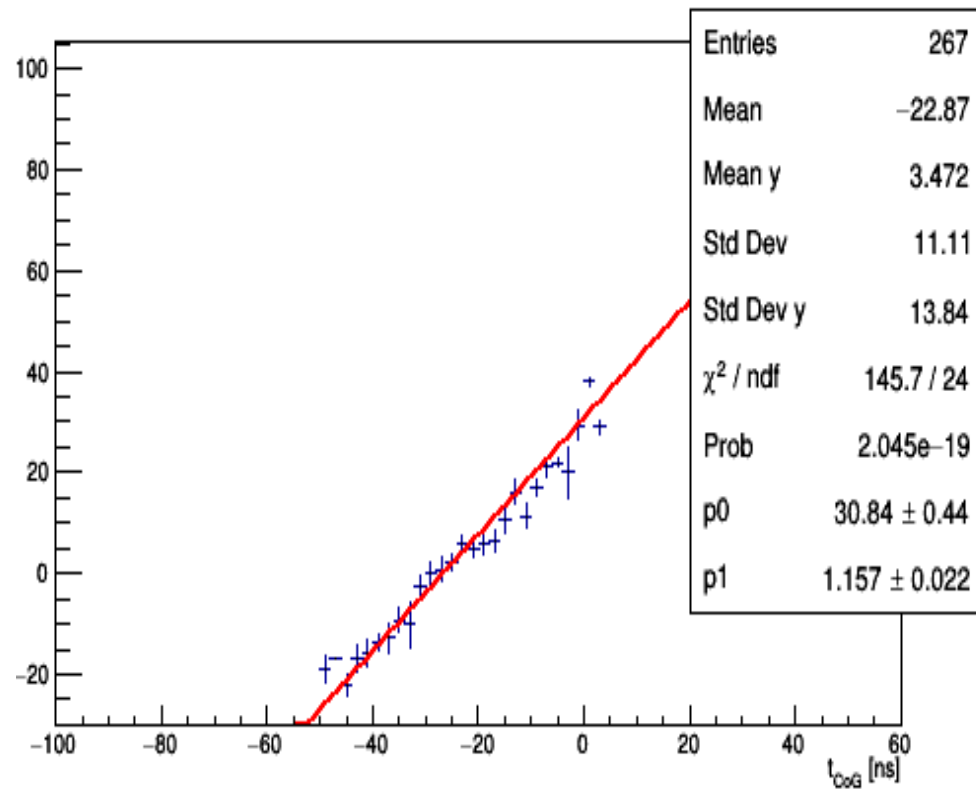
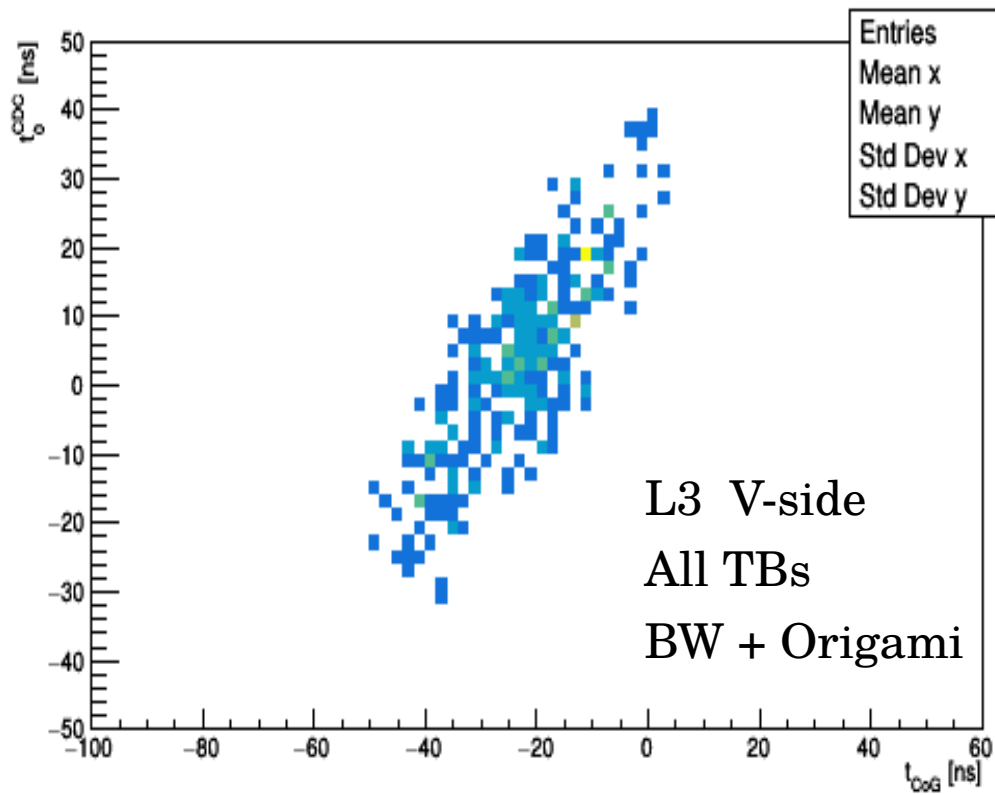
- $t_0$  [ns] versus  $t_{\text{CoG}}$  [ns],
- left: Scatter plot, right: linear fit of the ProfileX of the scatter plot





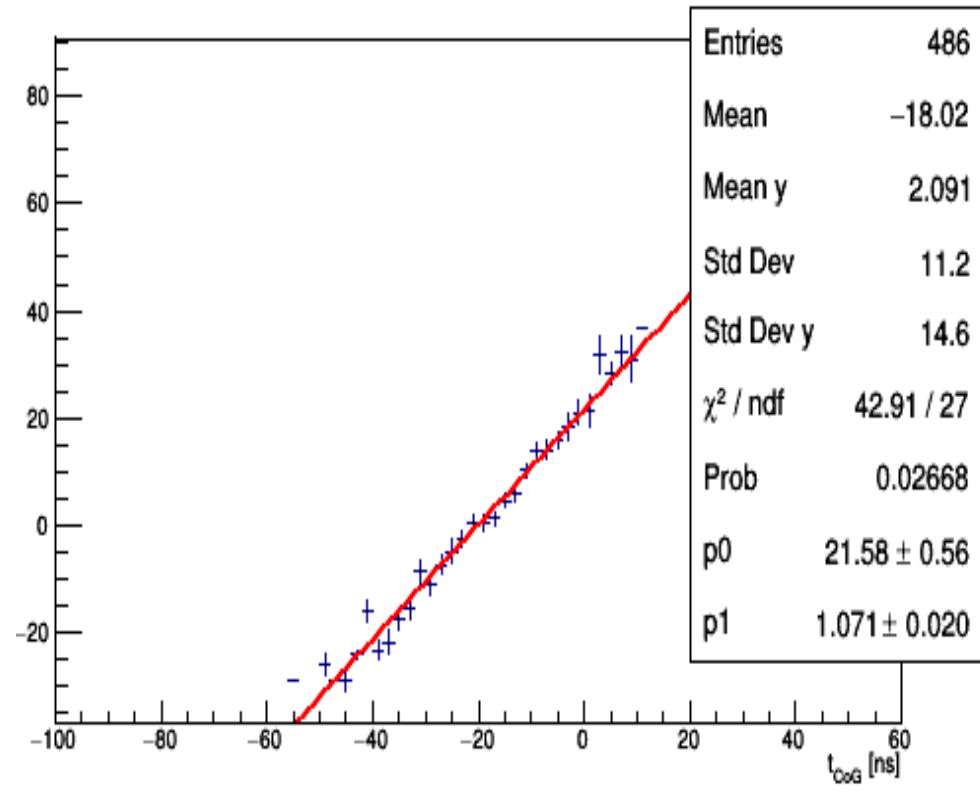
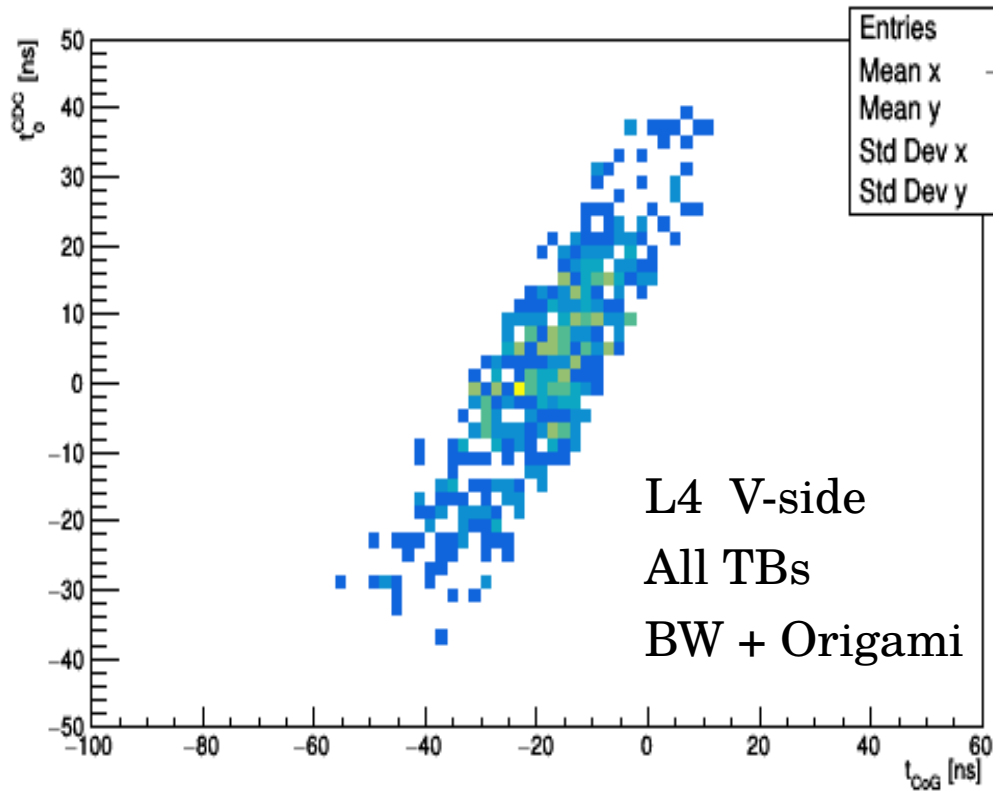
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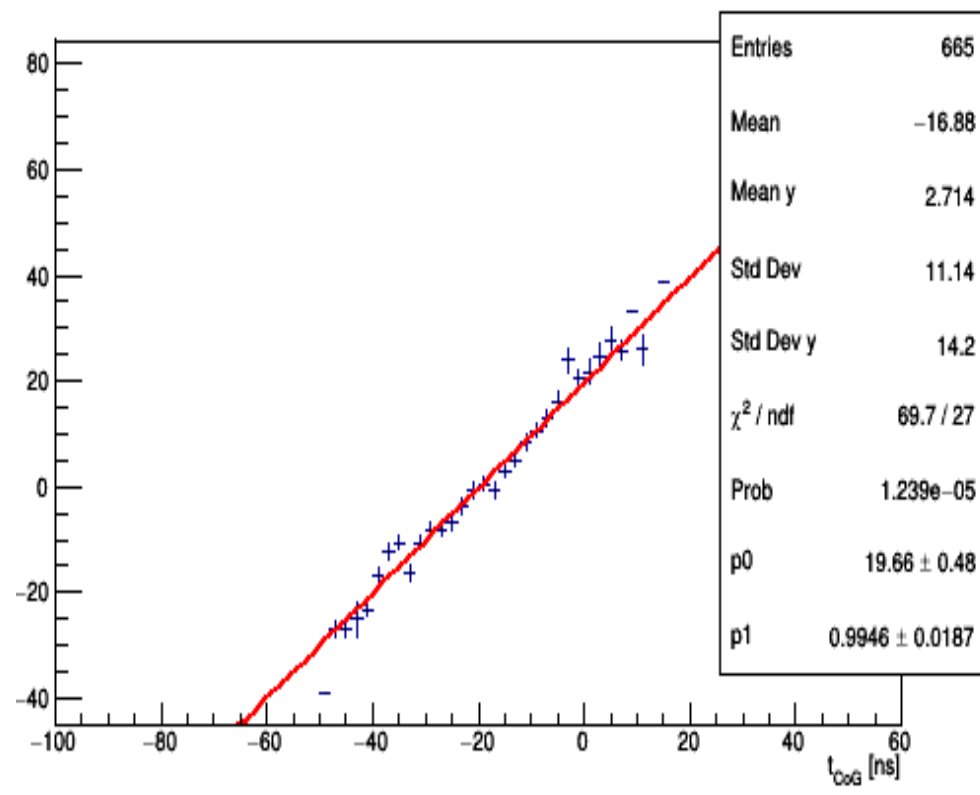
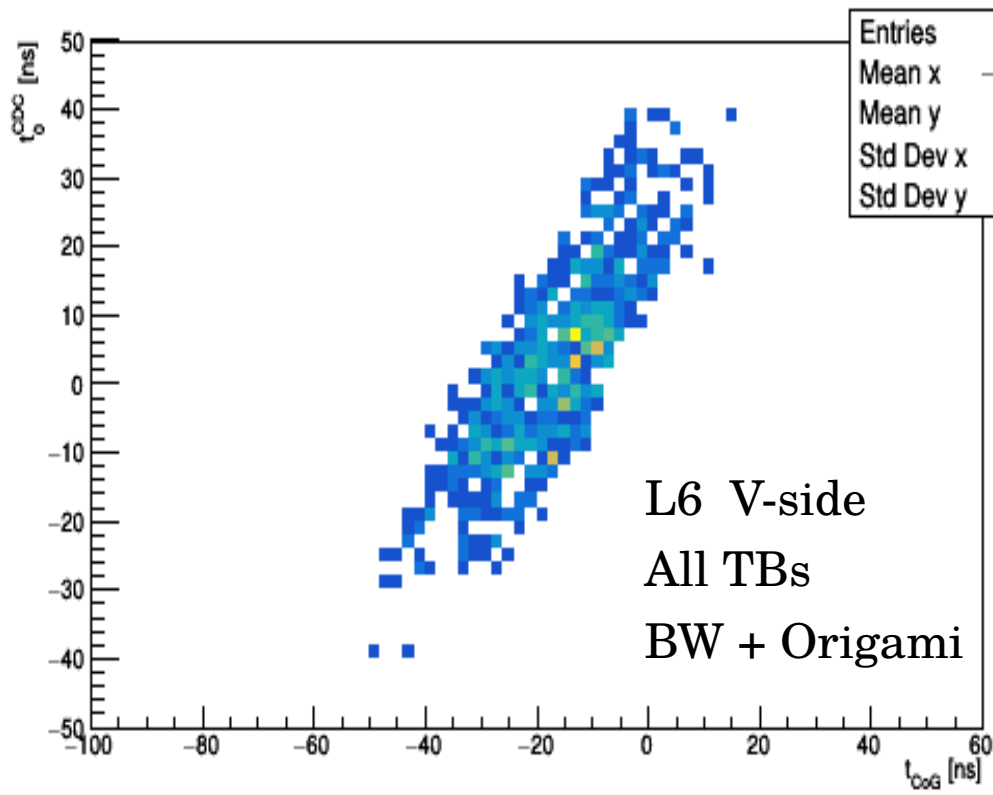
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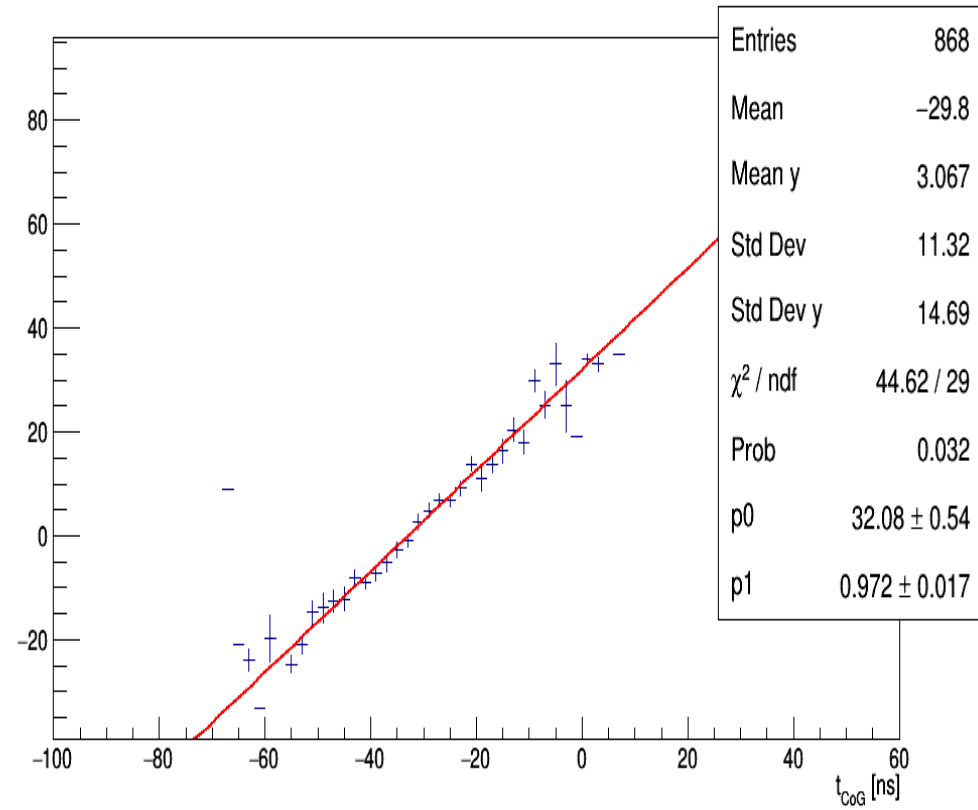
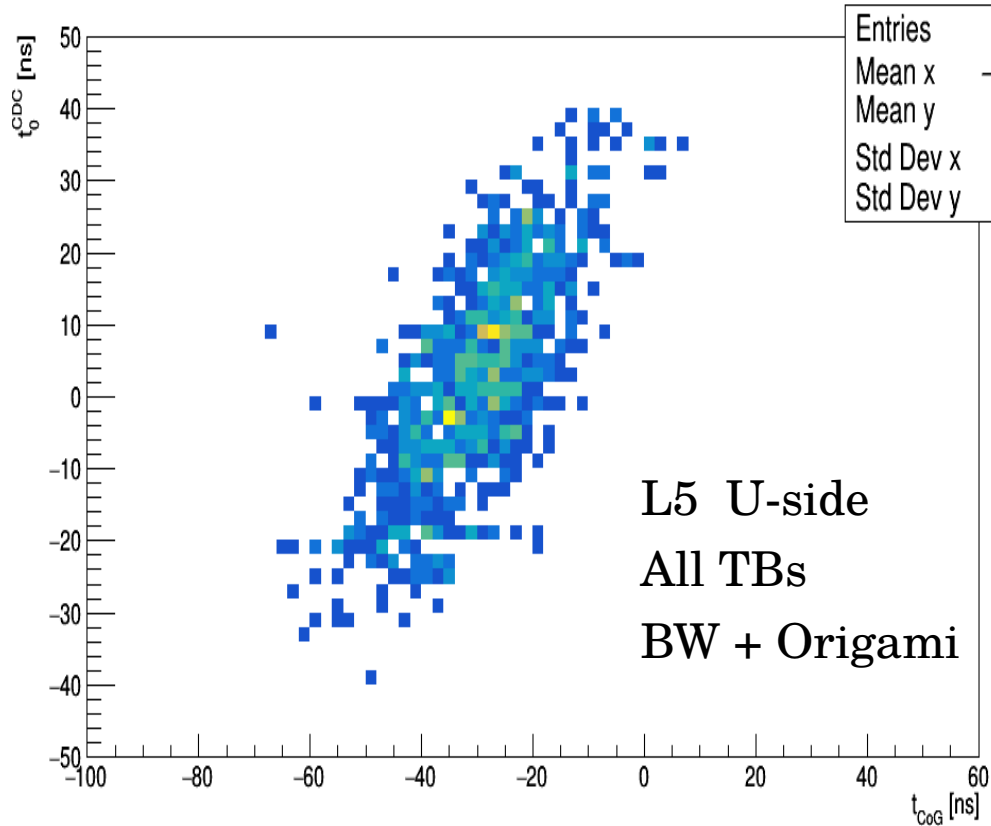
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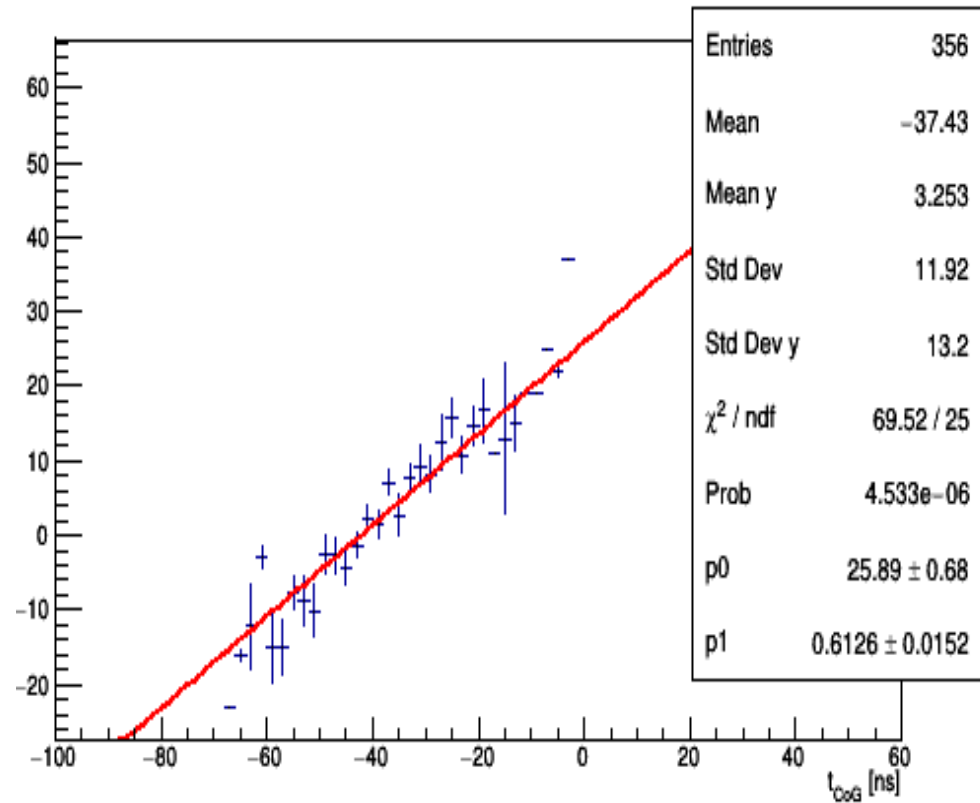
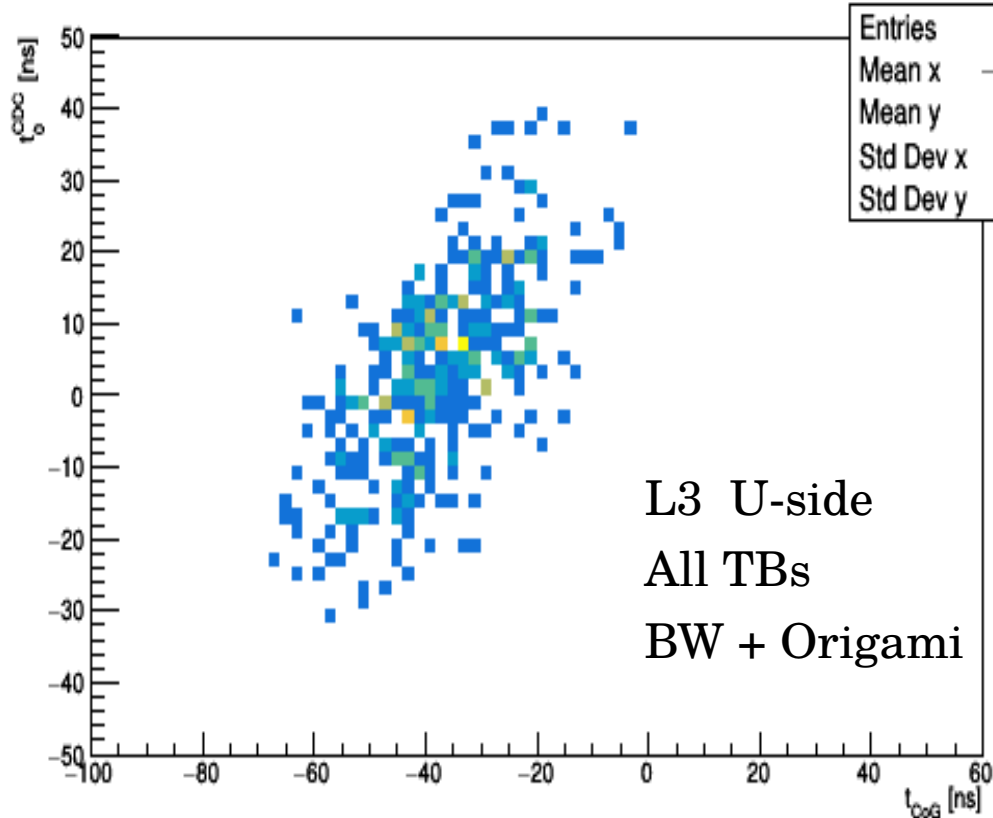
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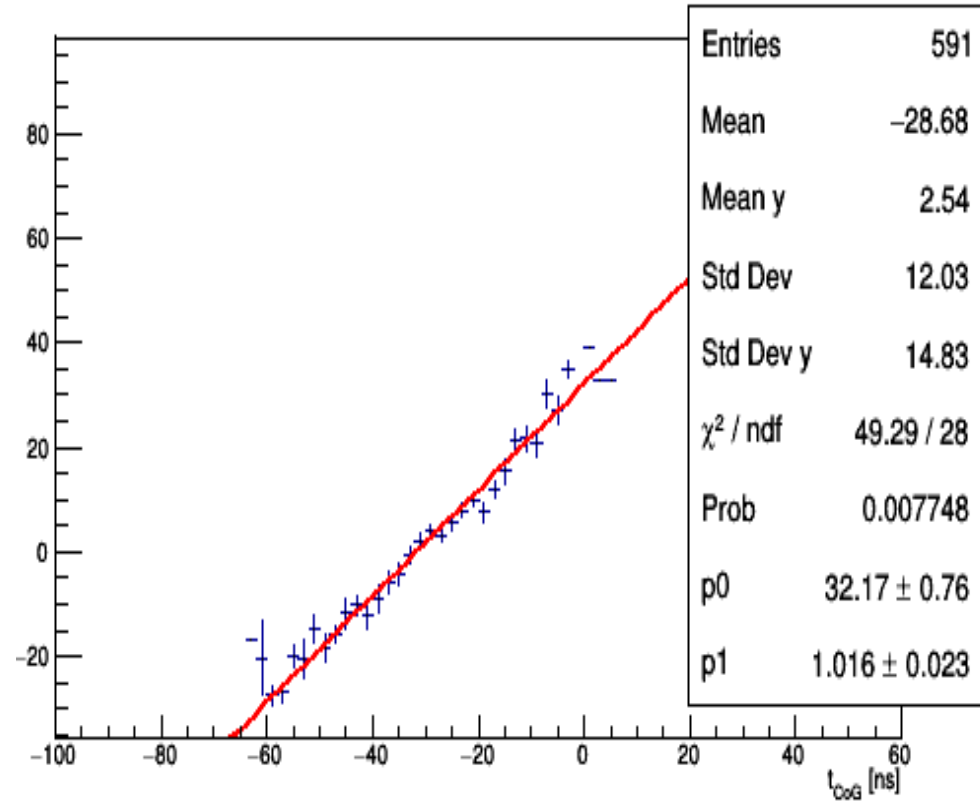
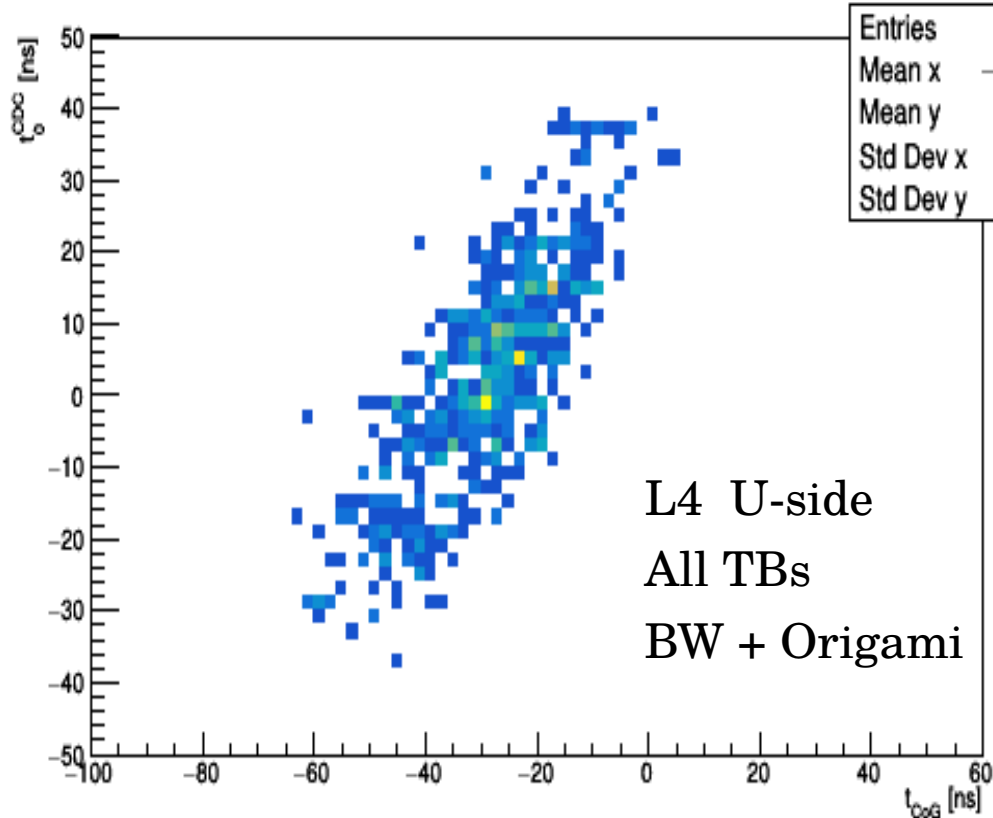
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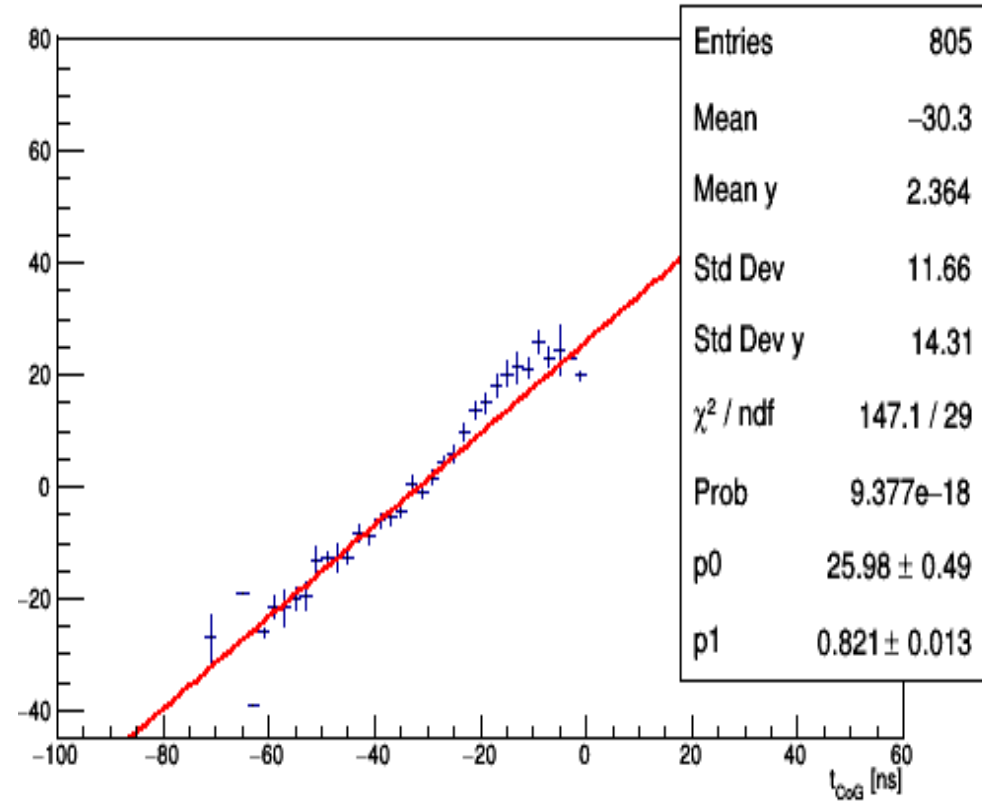
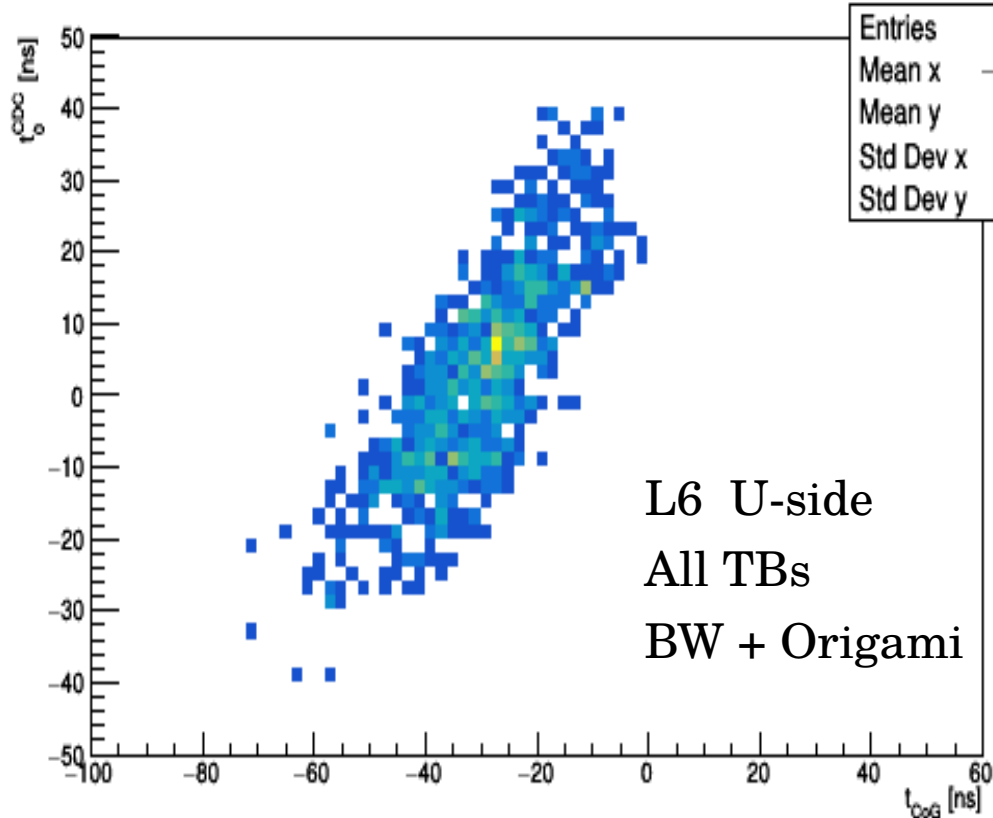
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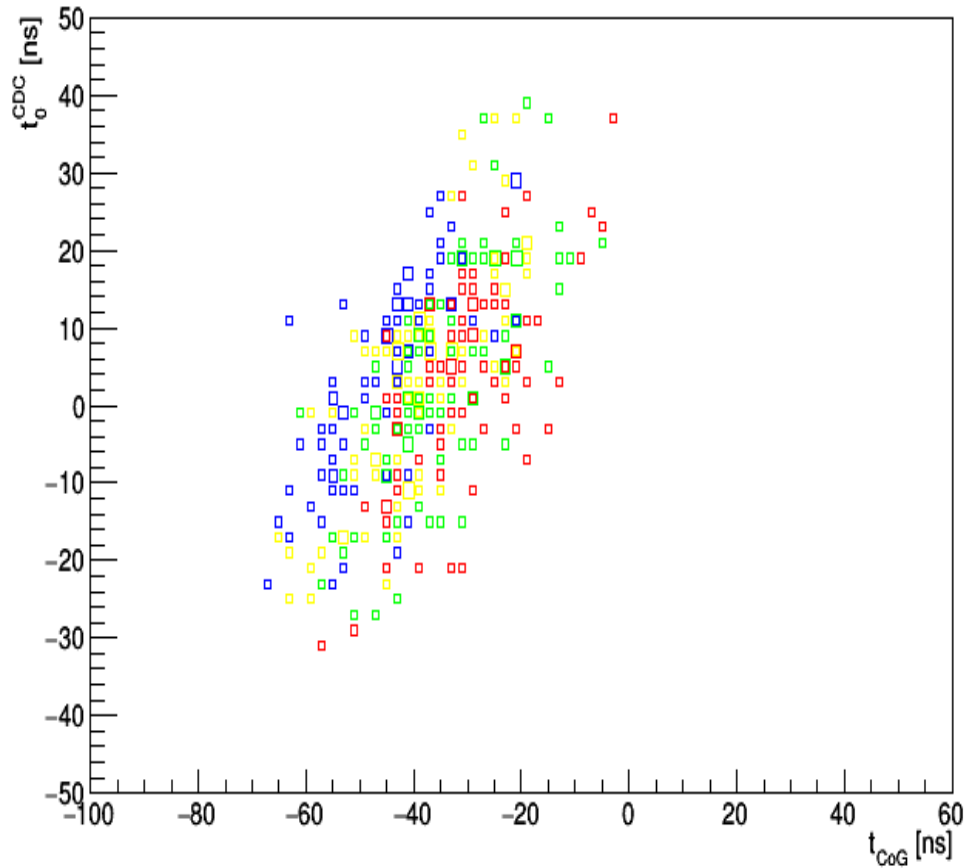
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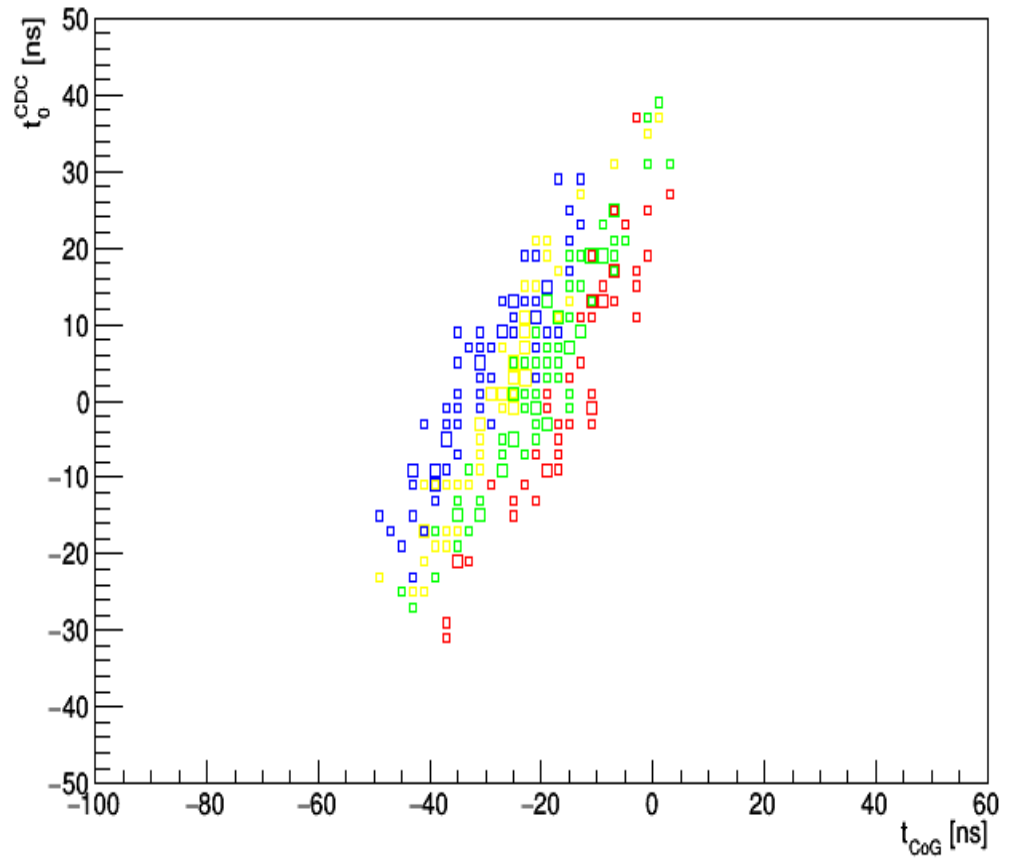


# CoG Calibration

L3 U-side, both sensors, run 2804



L3 V-side, both sensors, run 2804

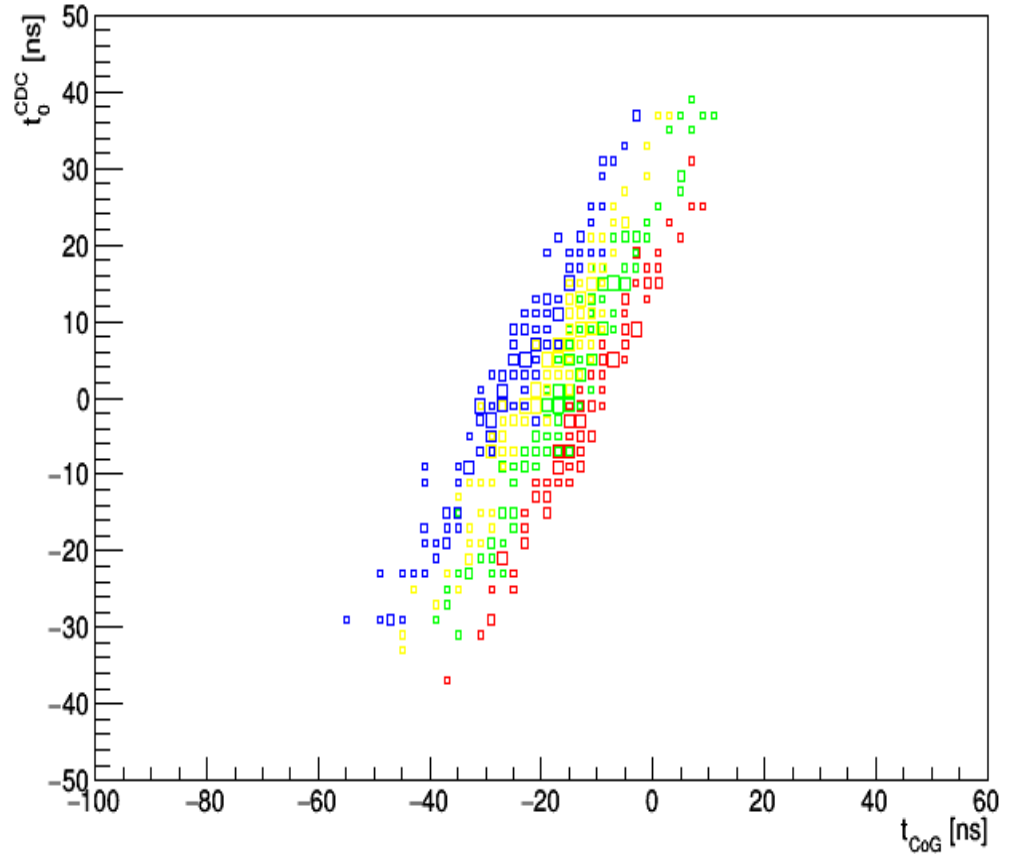
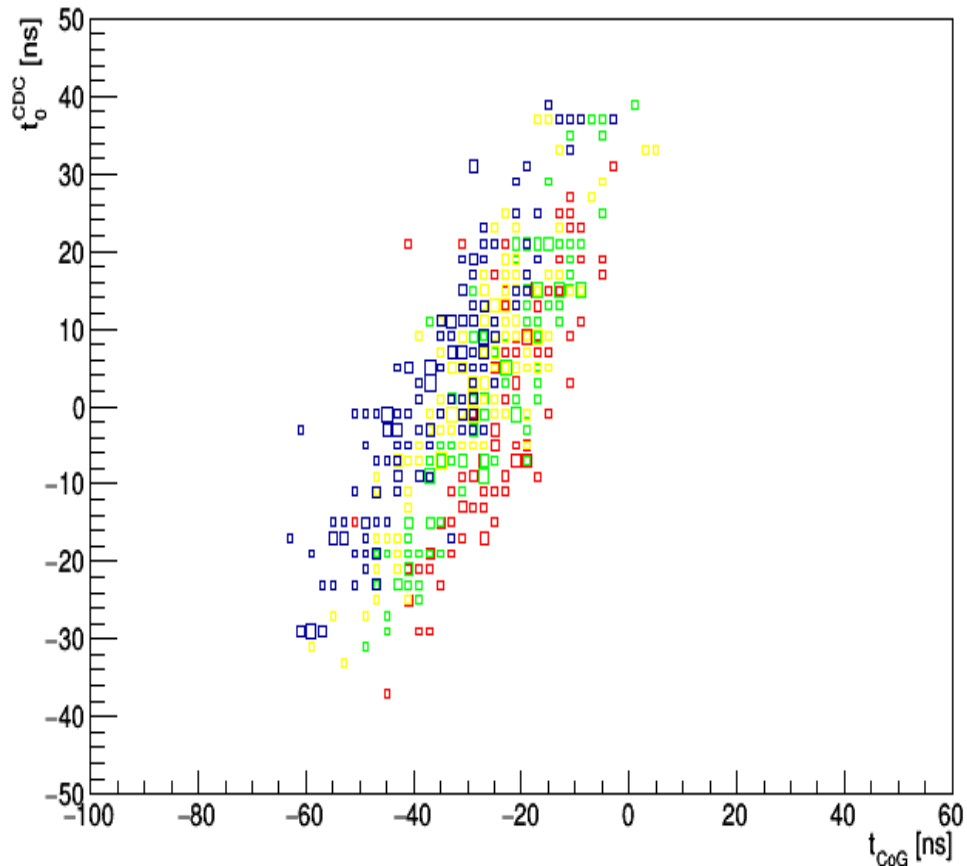




# CoG Calibration

L4 U-side, BW+ Origami, run 2804

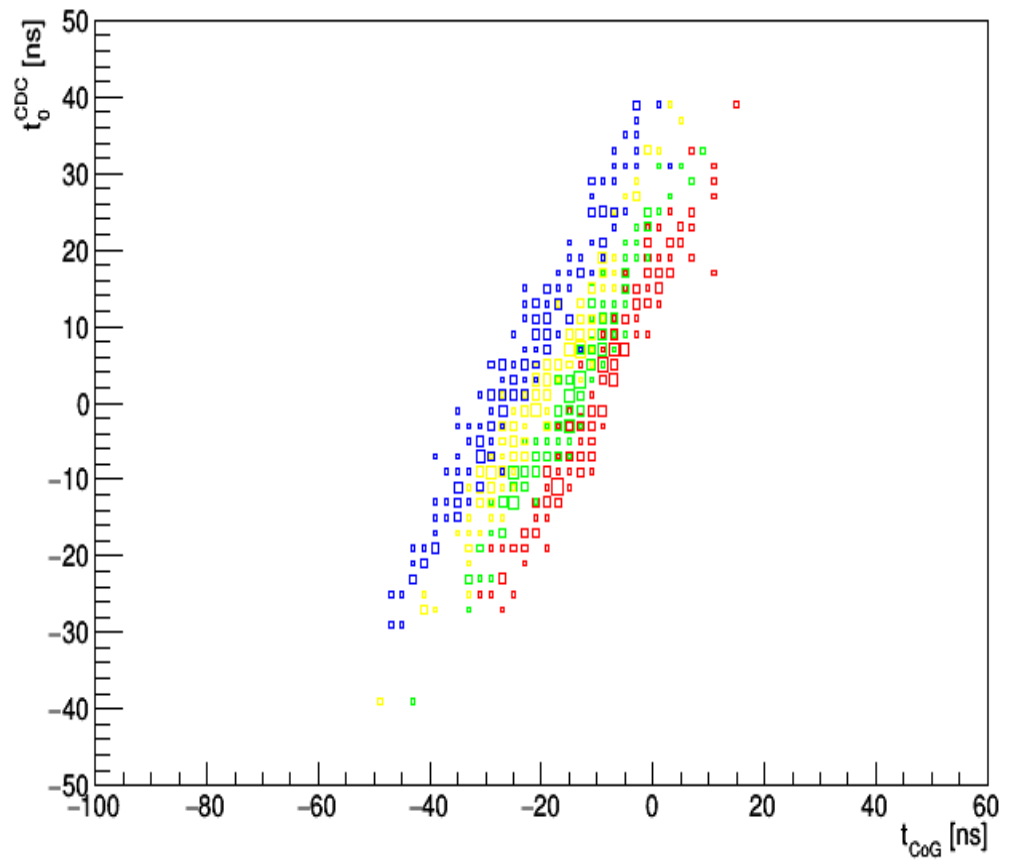
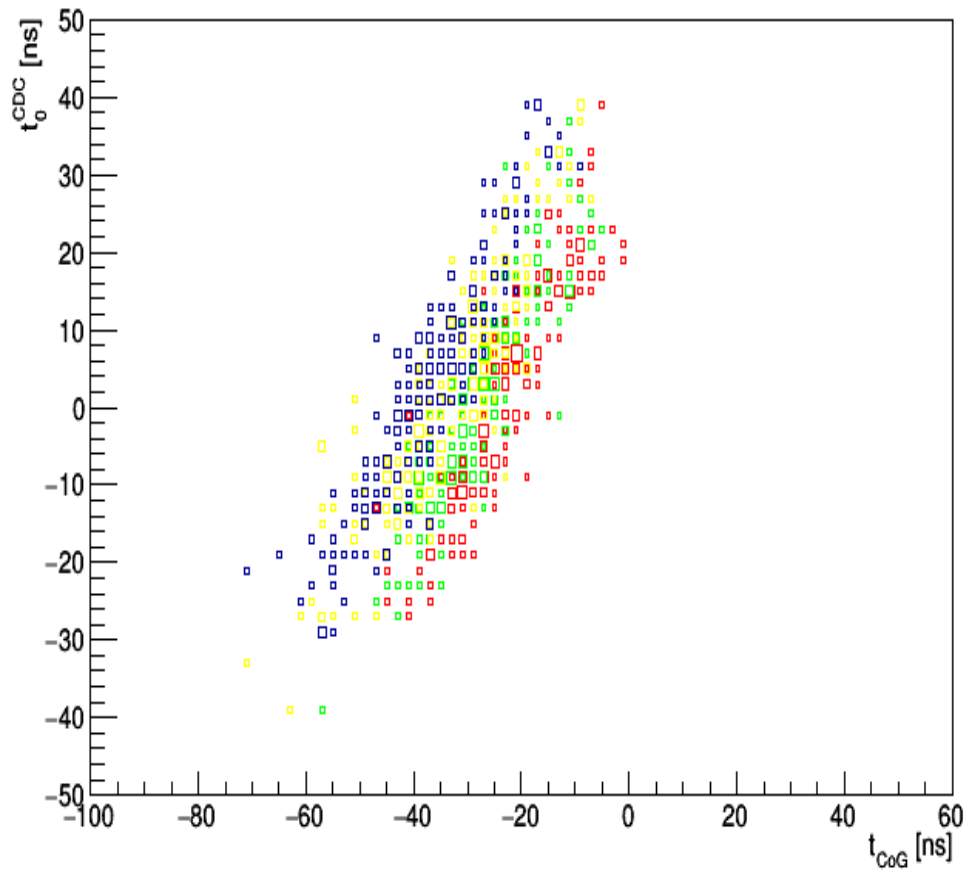
L4 V-side, BW+Origami, run 2804



# CoG Calibration

L6 U-side, BW+ Origami, run 2804

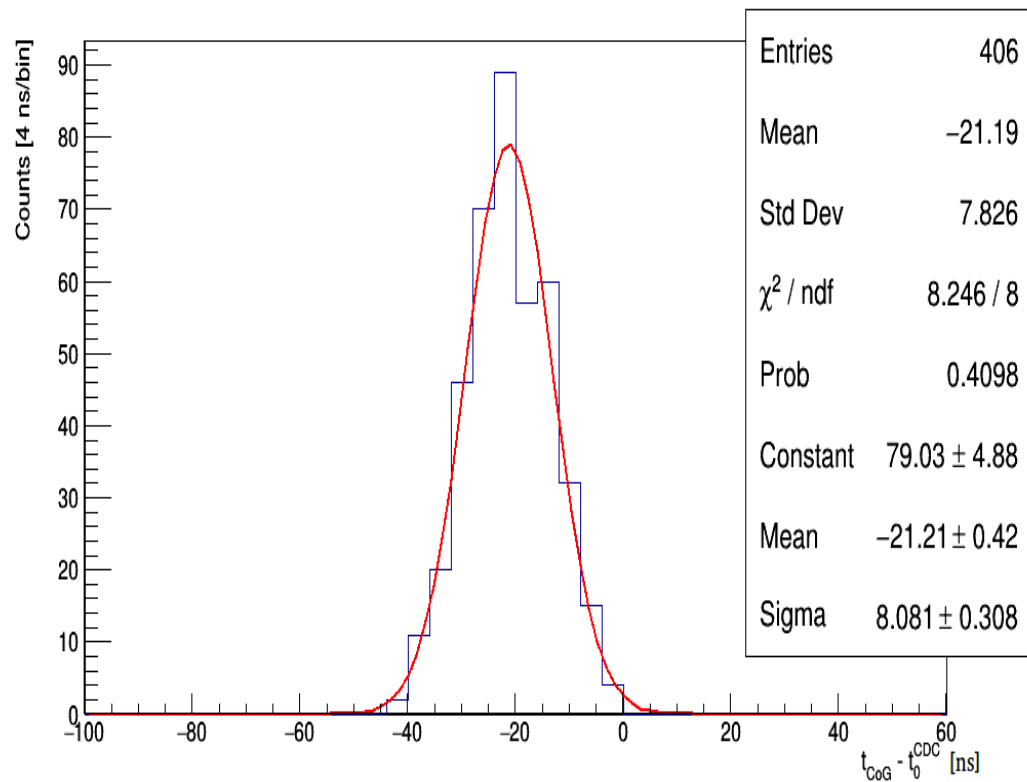
L6 V-side, BW+Origami, run 2804



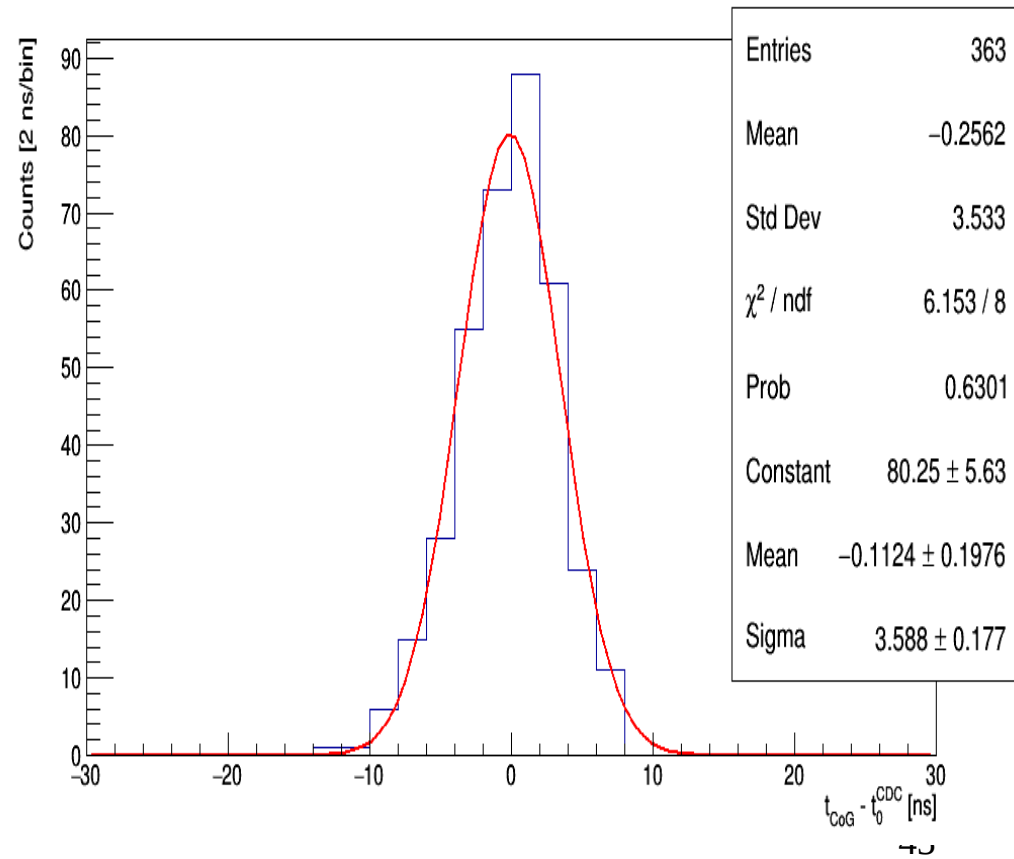
# CoG Calibration

- L4 V-side BW + Origami

before correction



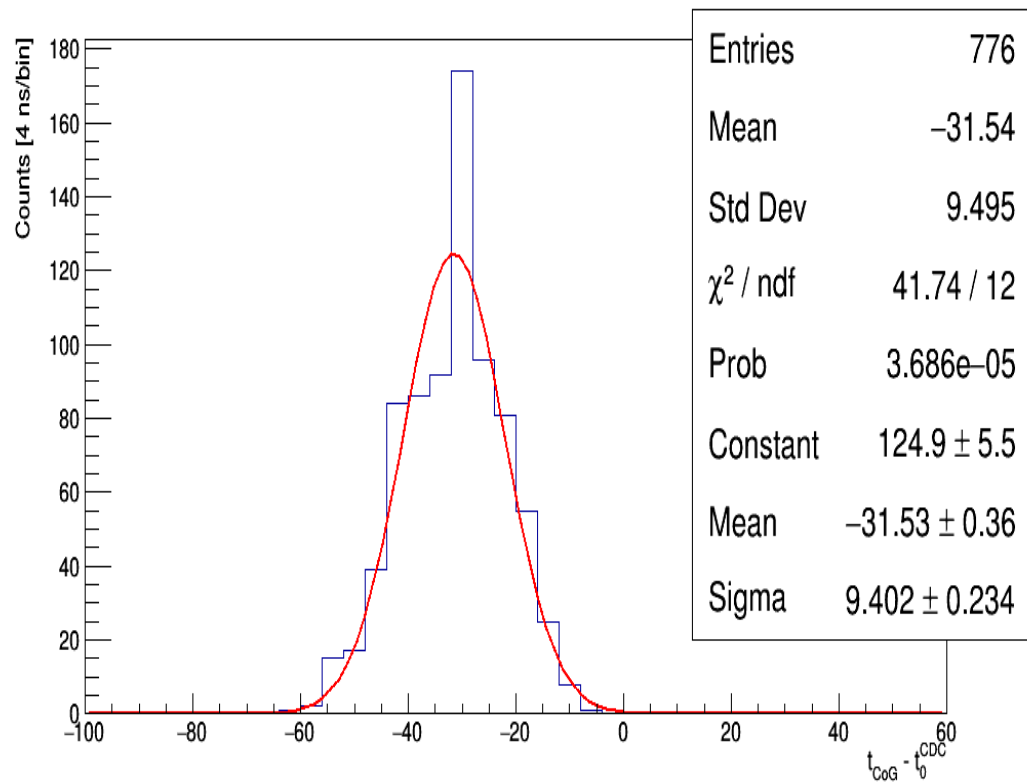
after correction



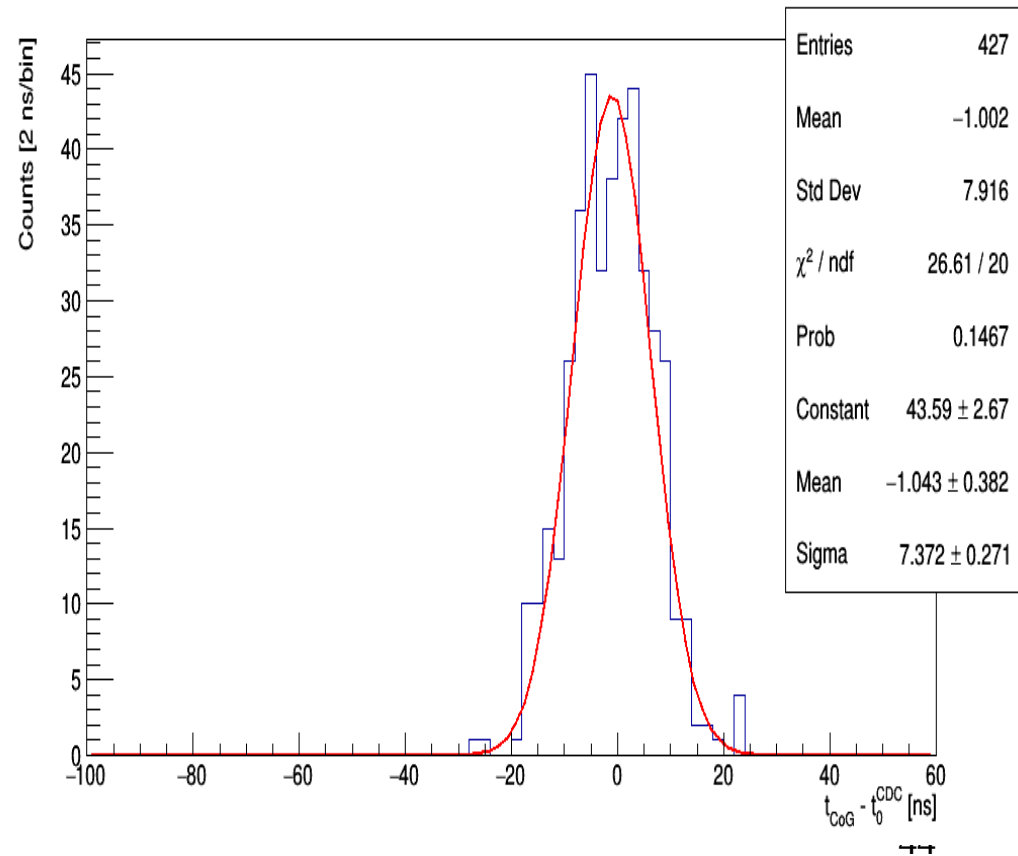
# CoG Calibration

- L4 U-side BW + Origami

before correction



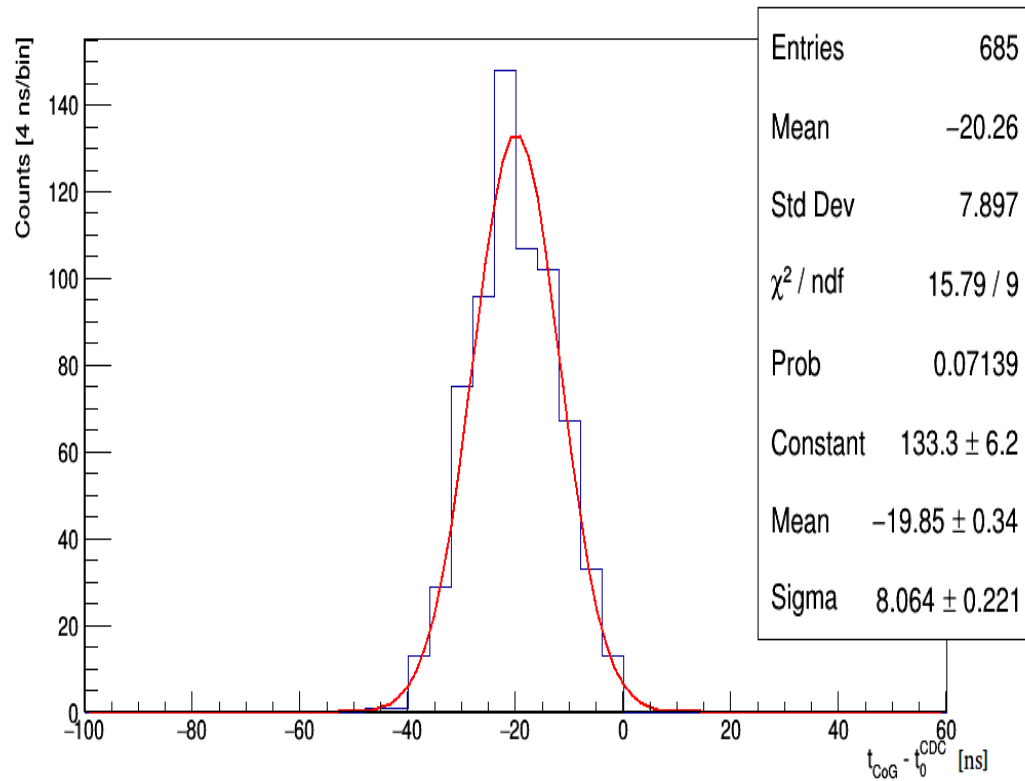
after correction



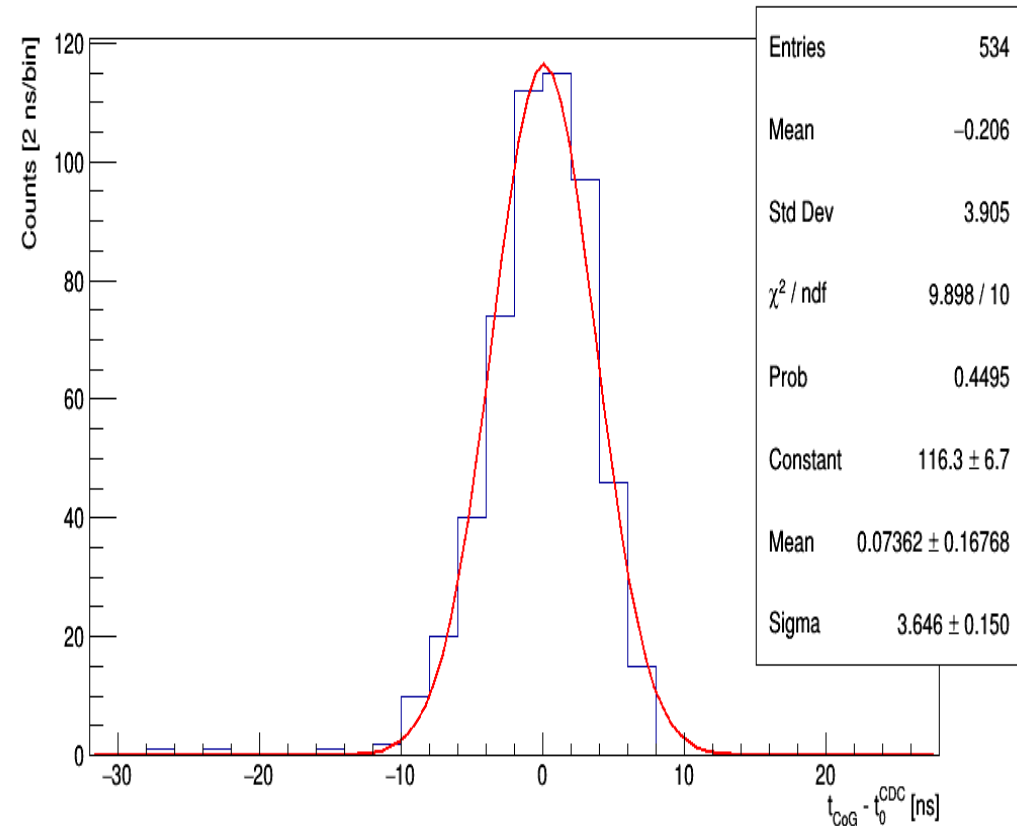
# CoG Calibration

- L6 V-side BW + Origami

before correction



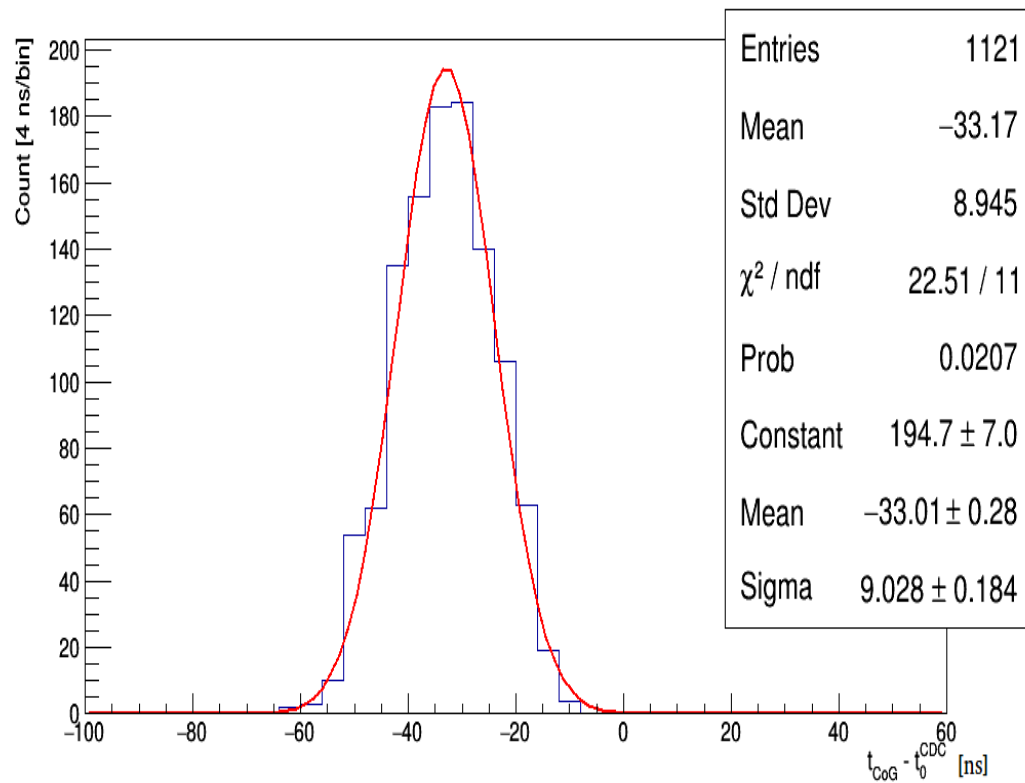
after correction



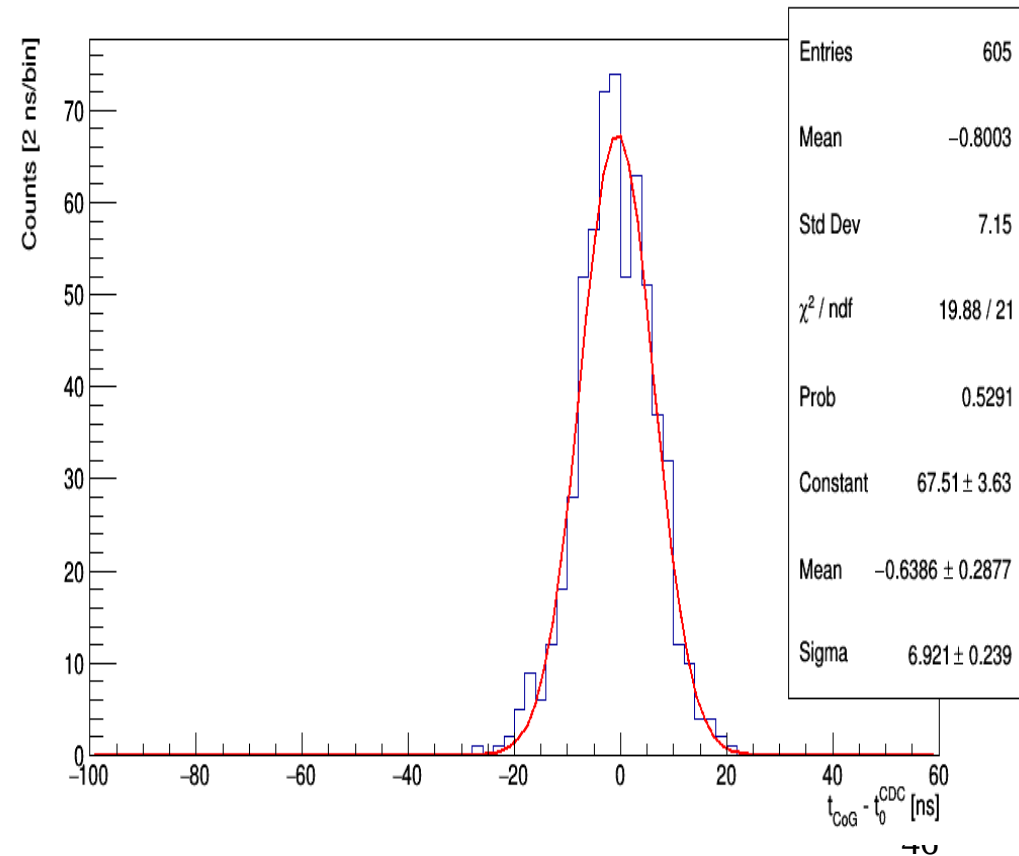
# CoG Calibration

- L6 U-side BW + Origami

before correction

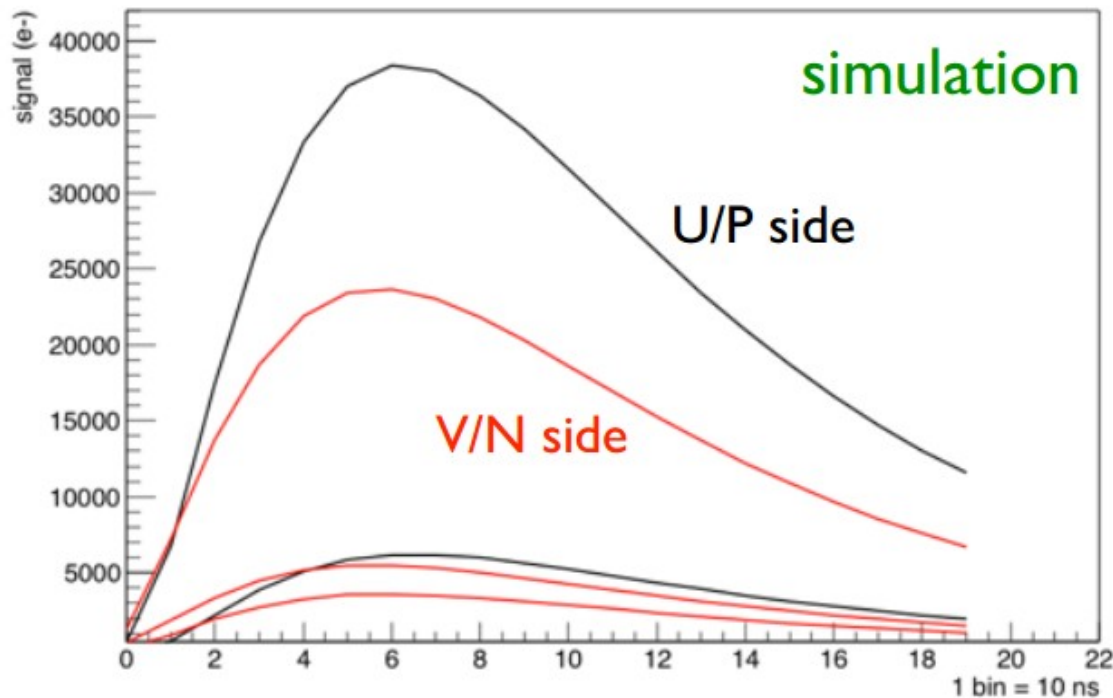


after correction



# U-side and V-side different waveform

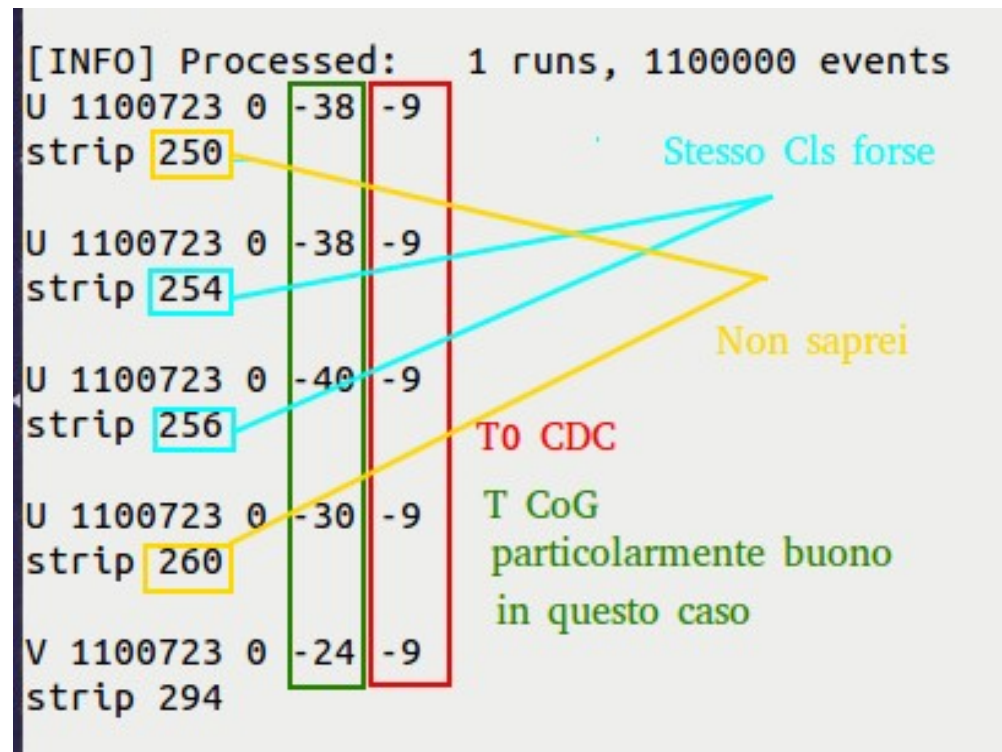
- The V-side and U-side have different waveform, due to the fact that electron are faster than holes: this could contributes to the different estimation of CoG



- The U-side resolution for the L3 U-side is influenced by the fact that we are joining both sensors but we realized that they are different

# Splitting of the Clusters

- Different entries between U and V due probably to not physics events or to signals which start in event and propagates in the following event or to splitting of clusters

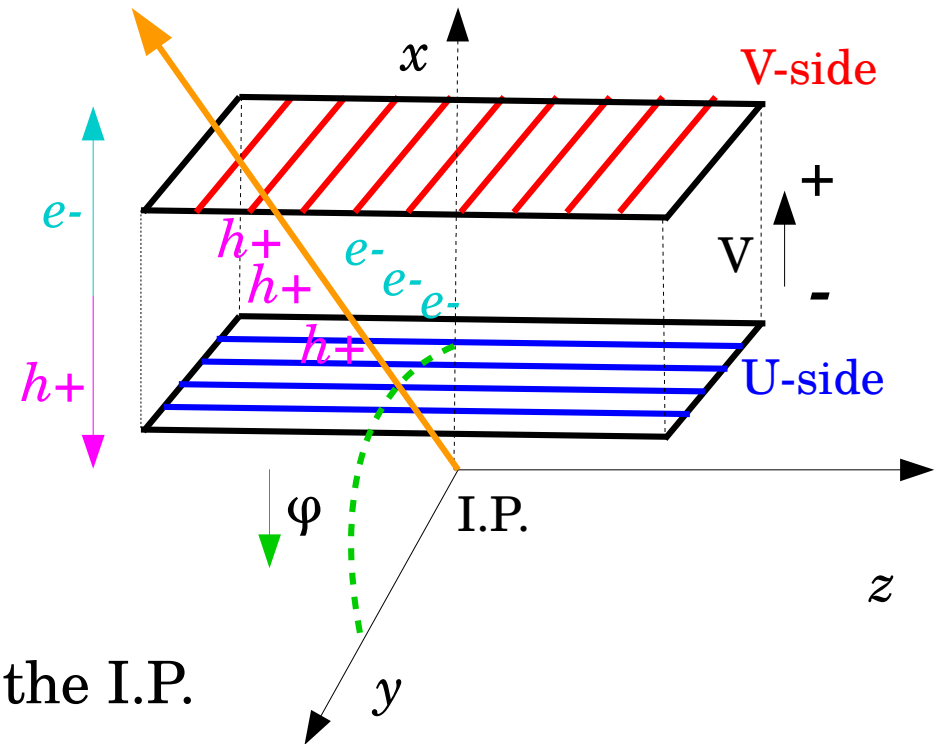
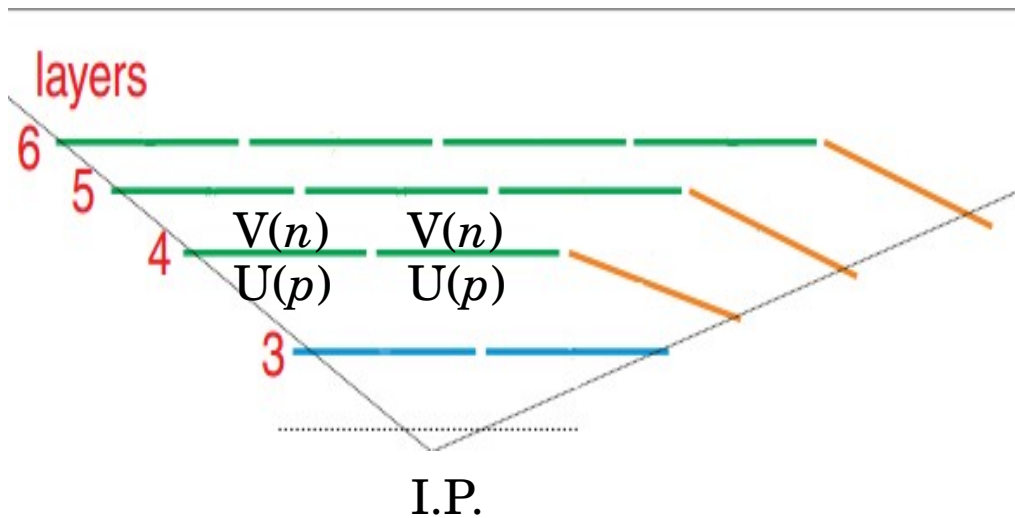


- This problem maybe will be not a problem with collision tracks because the clusters are smaller



# Hints on SVD configuration

- Each sensor has 2 sides, U and V, with a different strips configuration: the strips of the U-side are of type  $p$  and measure the  $\varphi$  direction while the strips of V-side are of type  $n$  and measure the  $z$  direction



- U-sides are those that watch through the I.P.
- V-sides are those that watch through the CDC