

G S

S I

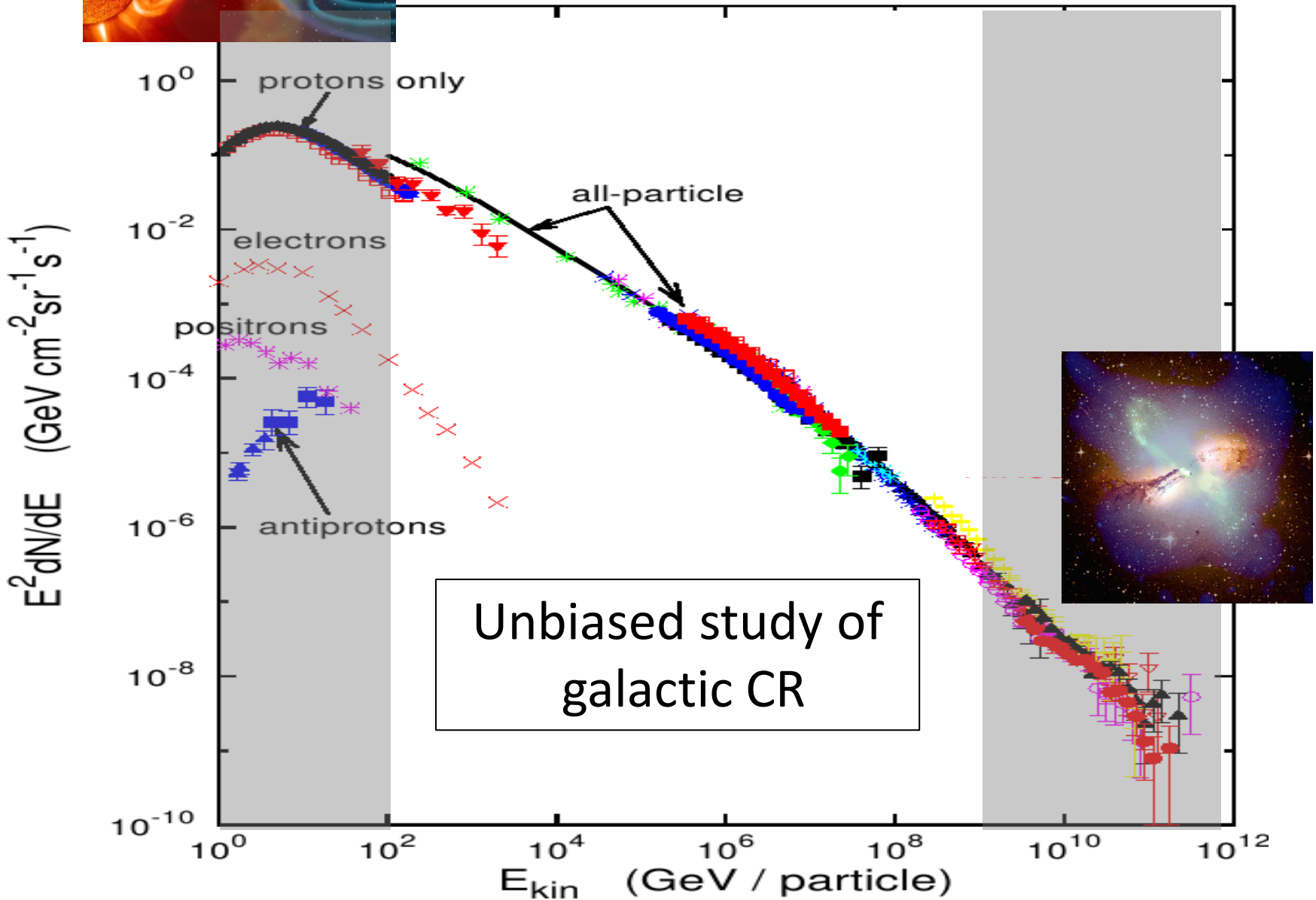
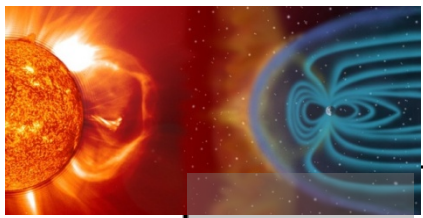


The HERD space mission: Galactic CR science and use of SiPM

Ivan DE MITRI

Gran Sasso Science Institute (GSSI) & INFN

On behalf of the HERD collaboration



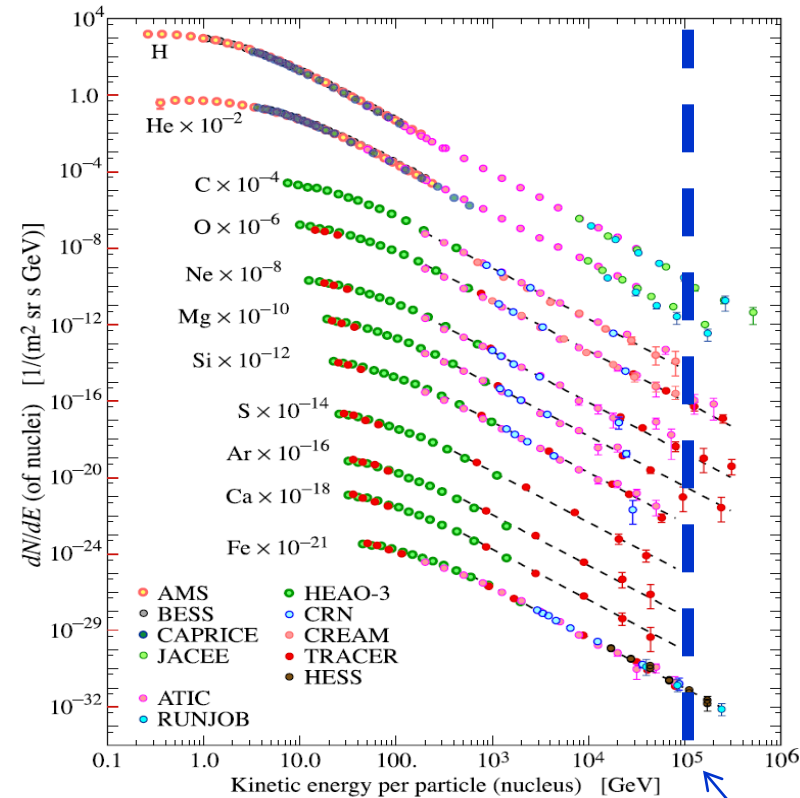
What we have...

Direct measurements

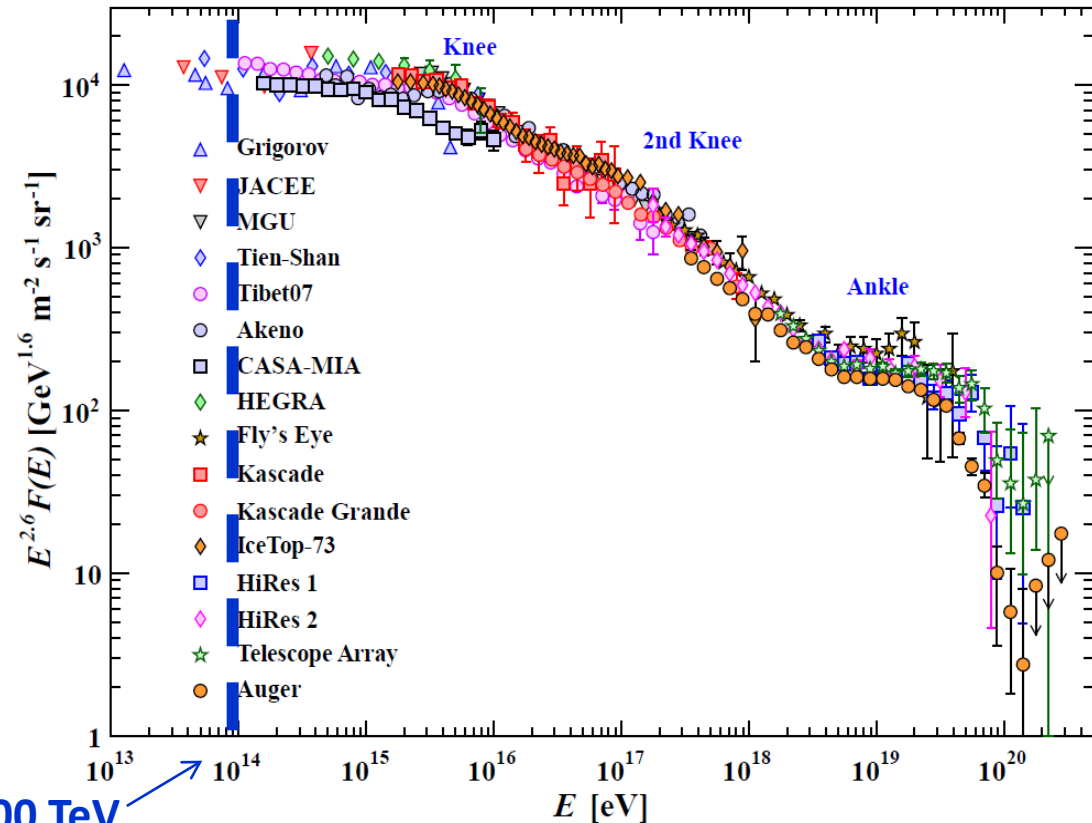
- High precision
- fluxes of single components
- (acceptance) limited in energy

Indirect measurements

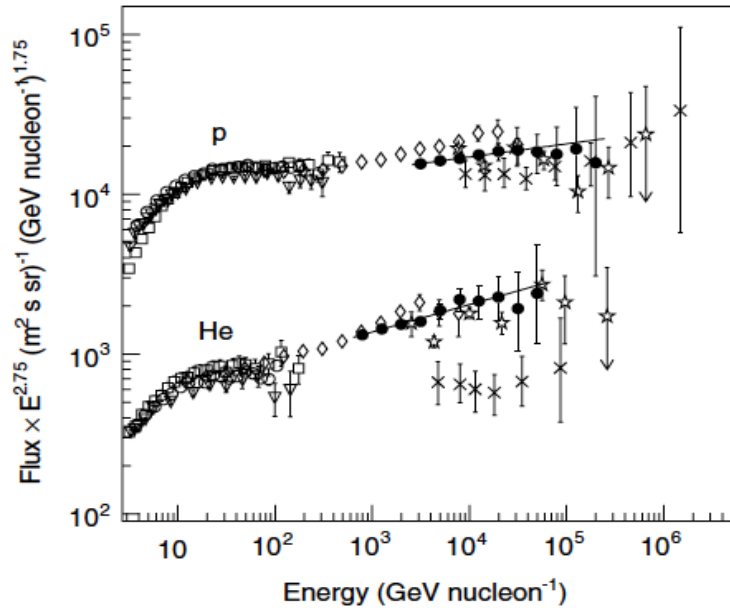
- Larger systematics
- Difficult composition measurements
- Can go to the highest energies



100 TeV

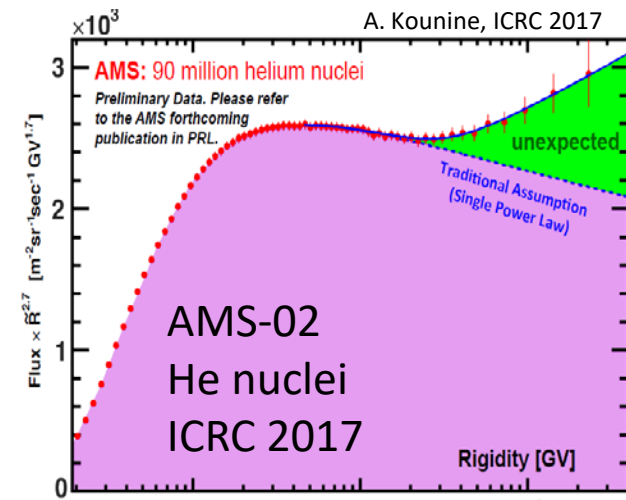
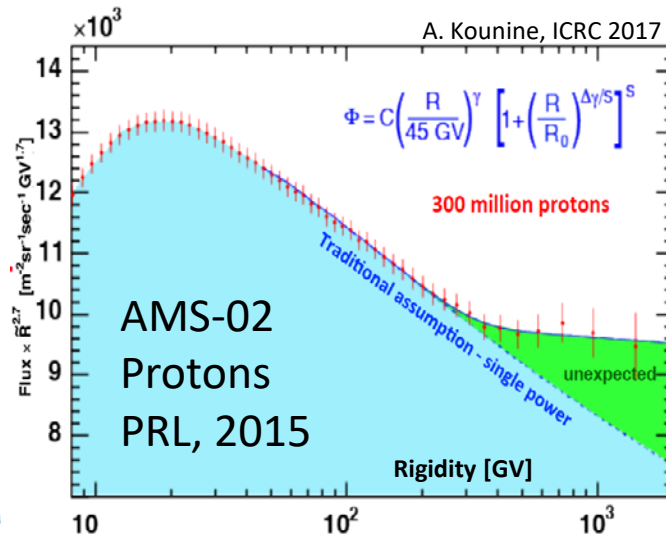
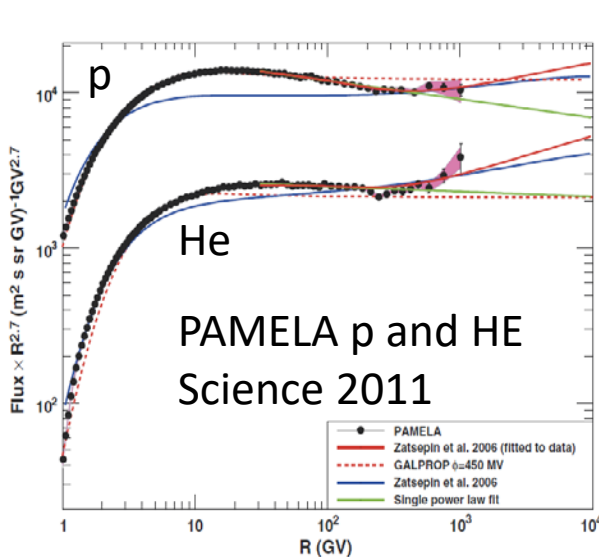
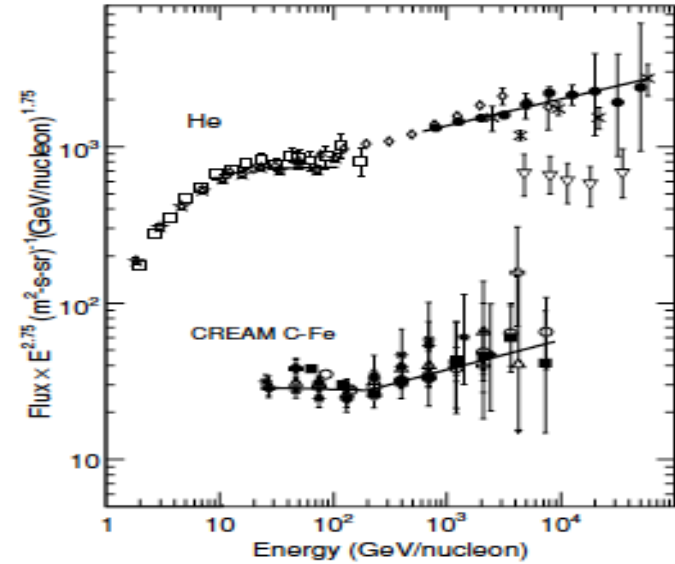


p and He spectral hardenings



CREAM
First hints for
Hardenings.

PAMELA and AMS
Direct detection fo
the break at about
250GeV/n



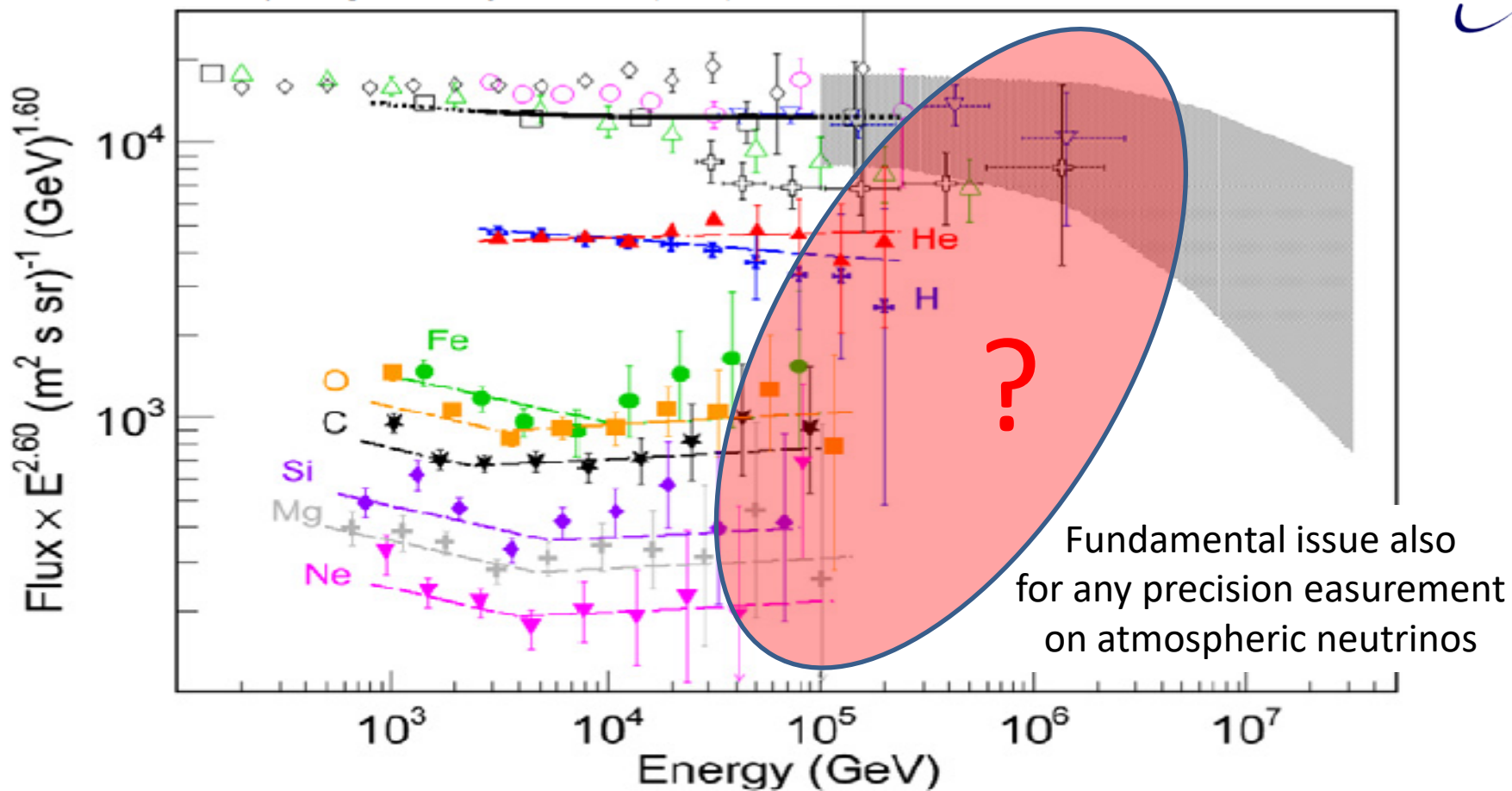


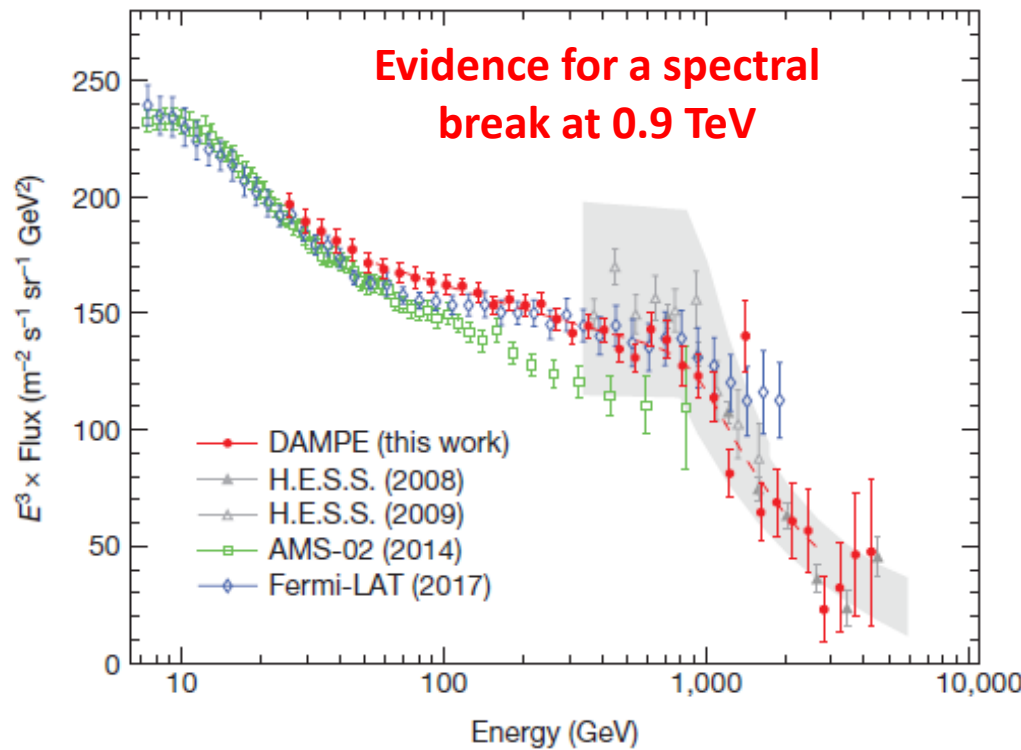
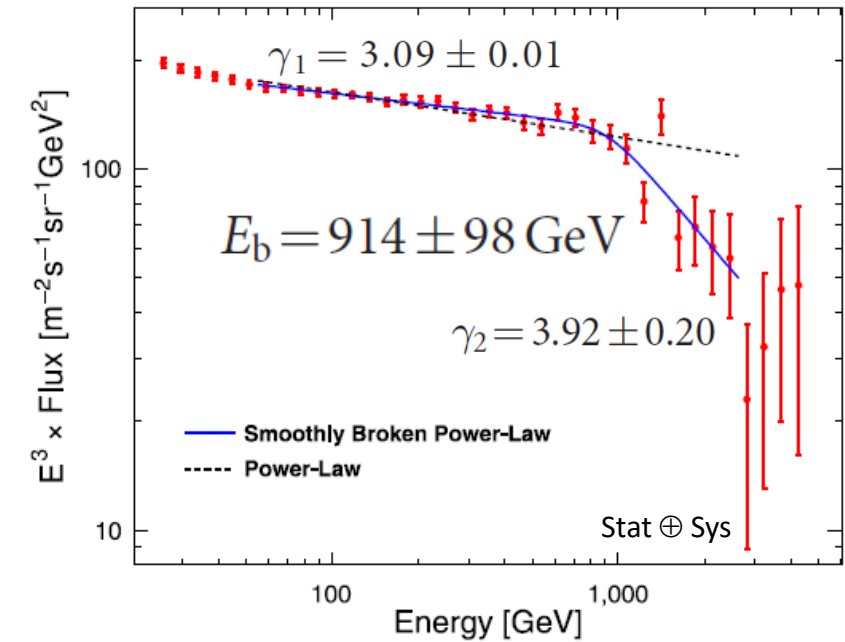
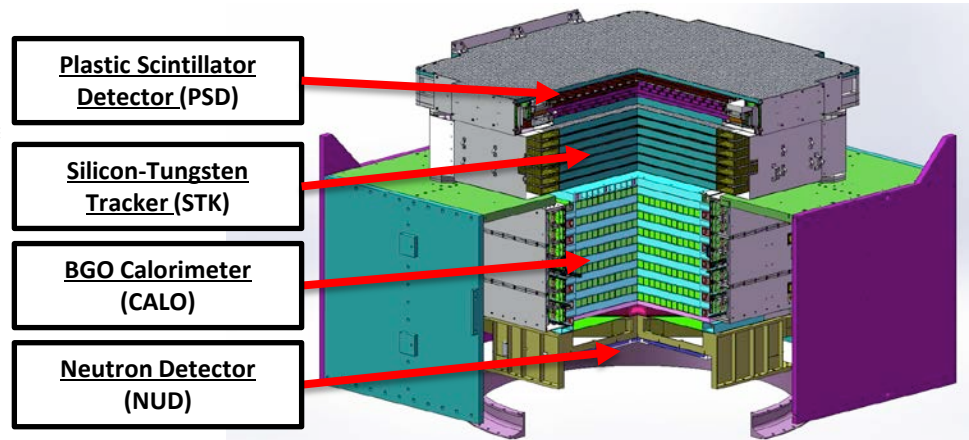
Fig. 11. The all-particle spectrum (black solid curve) obtained by summing up CREAM elemental spectra from p to Fe (filled symbols) is compared with previous measurements (open symbols): ATIC-1 [35], black squares; JACEE, blue downward triangles; RUNJOB, black crosses; Ichimura et al. [71], green upward triangles; SOKOL [72], pink circles. The gray shaded area indicates ground based indirect measurements. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

The DAMPE ($e^+ + e^-$) spectrum



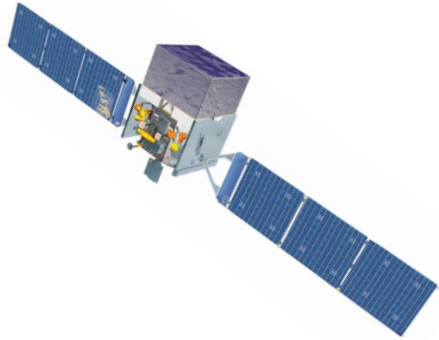
Direct detection of a break in the teraelectronvolt cosmic-ray spectrum of electrons and positrons

DAMPE Collaboration*



- 530 days
- 2.8 billions CR events
- 1.5 million CREs above 25 GeV

Space/balloon vs Ground based



Direct measurements

Requirements:

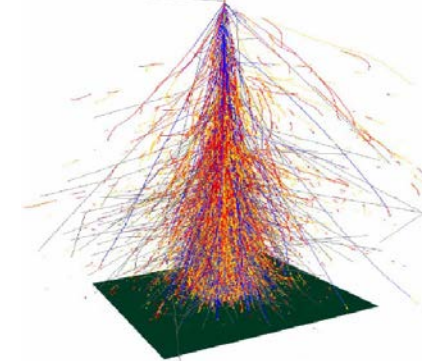
- Calorimetry vs Spectrometry
- Large acceptances
- <20% resolutions

Output:

Fully explore the sub-PeV region

Limitations:

- Surface/weight limited
- Hard to reach the all-particle knee
- Need high technology



Indirect measurements

Requirements:

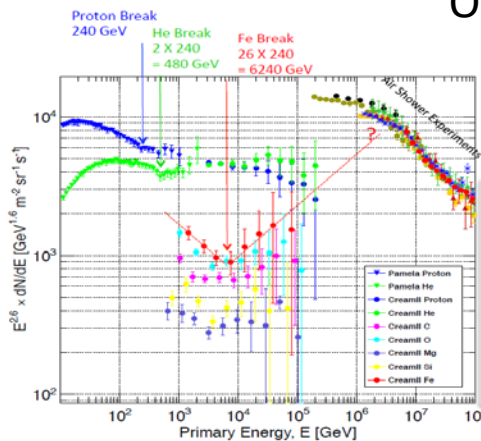
- Multi-Hybrid approach
- Operate at (not too) high altitude
- Large surfaces / samplings

Output:

Reach the highest energies

Limitations:

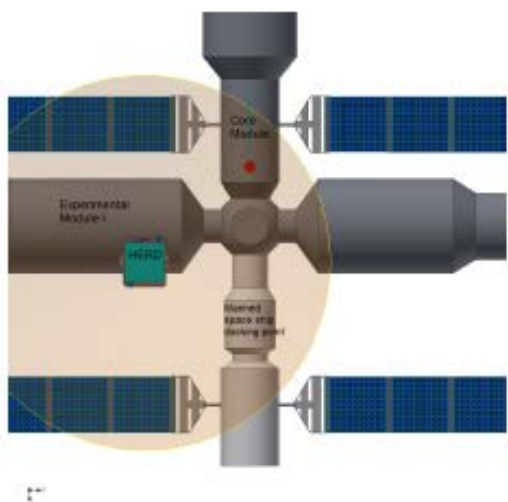
- Poor mass resolution
- Intrinsically limited by systematics
- Large model dependence



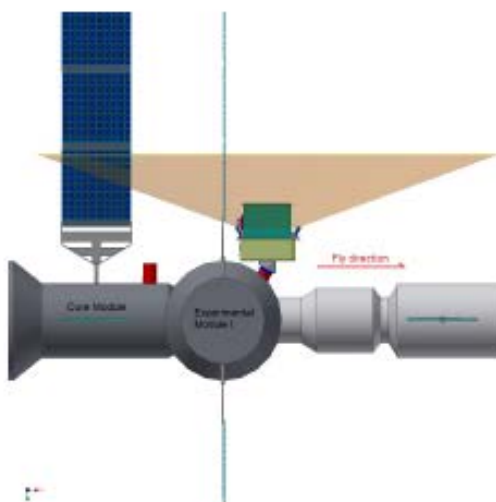
HERD

High Energy cosmic-Radiation Detection

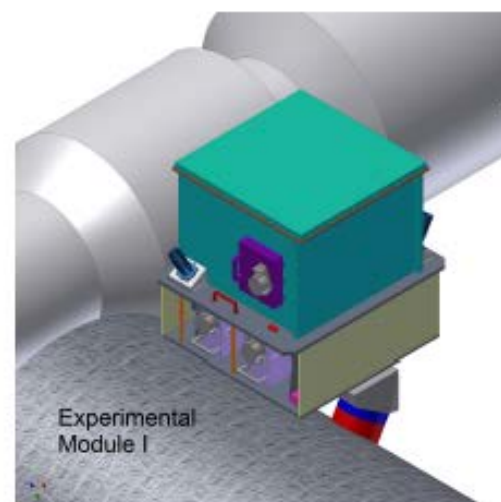
- HERD: a cosmic ray experiment onboard the China's Space Station (CSS)
- Science:
 - Precise cosmic ray spectra and composition up to the “knee”
 - Gamma-ray astronomy and transient studies e.m. follow
 - Electrons spectra (and anisotropy) up to tens of TeV
 - Indirect dark matter searches with high sensitivity



(a) Top view



(b) Side view

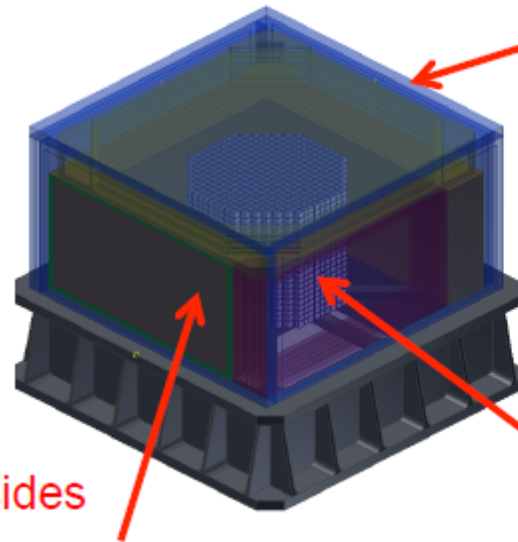
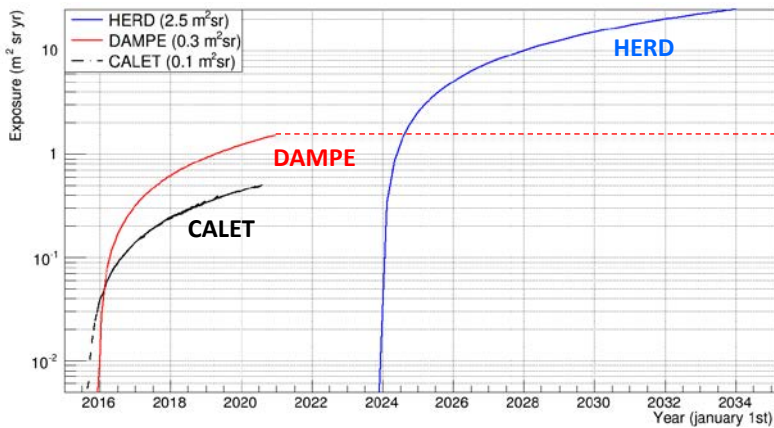


(c) Oblique view

The HERD payload

Large acceptance, deep, 3D calorimeter, equipped with silicon tracker (STK) and plastic scintillators (PSD) for primary identification, onboard CSS for a long duration mission.

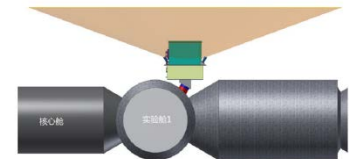
One order of magnitude jump in exposure wrt current generation CR experiment: 15 m² sr yr



PSD, five sides
low energy
Gamma Id
Charge

3D CALO
e/G/CR energy
e/p discrimination

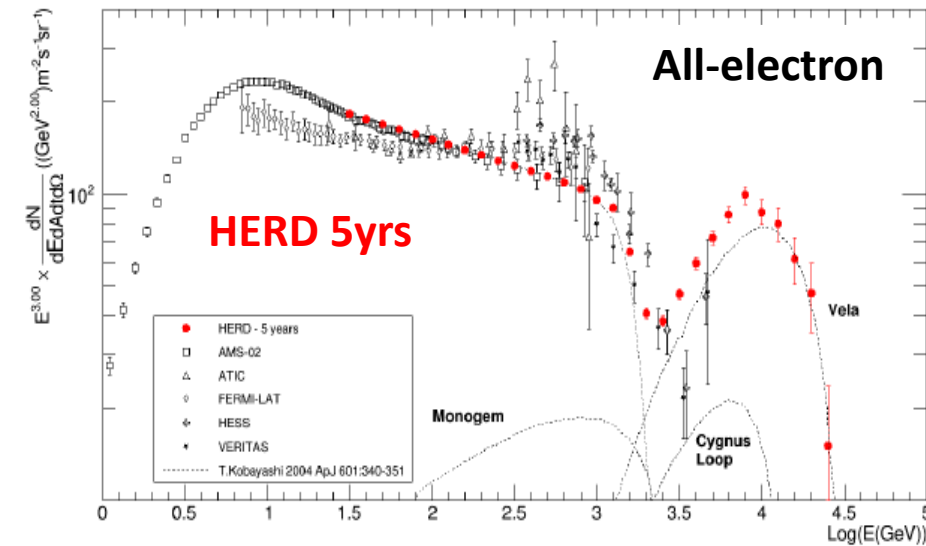
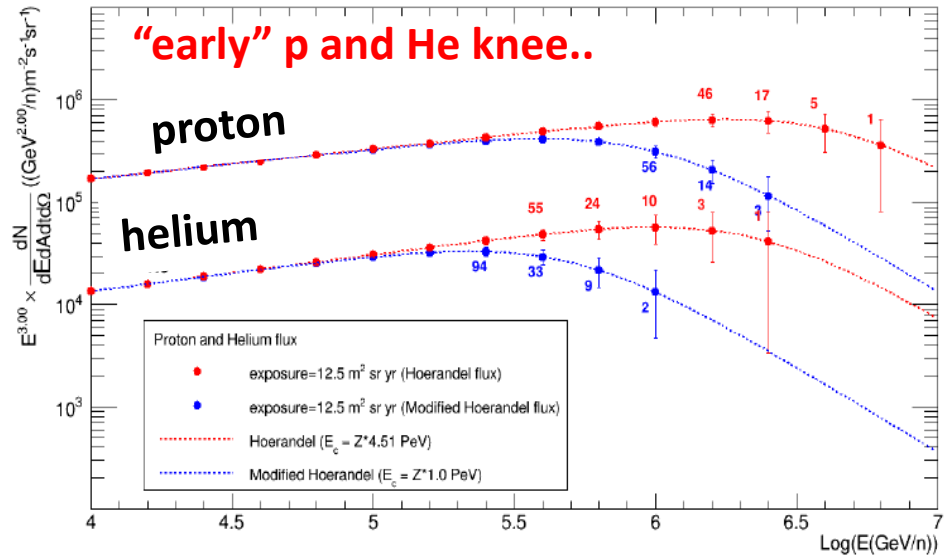
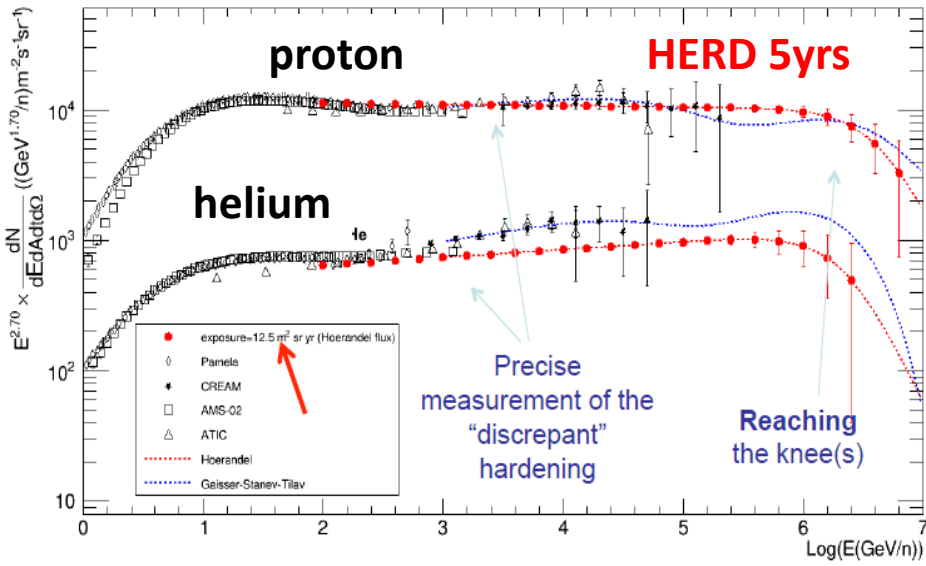
STK(SSD+W), five sides
Charge
Trajectory
Gamma tracking



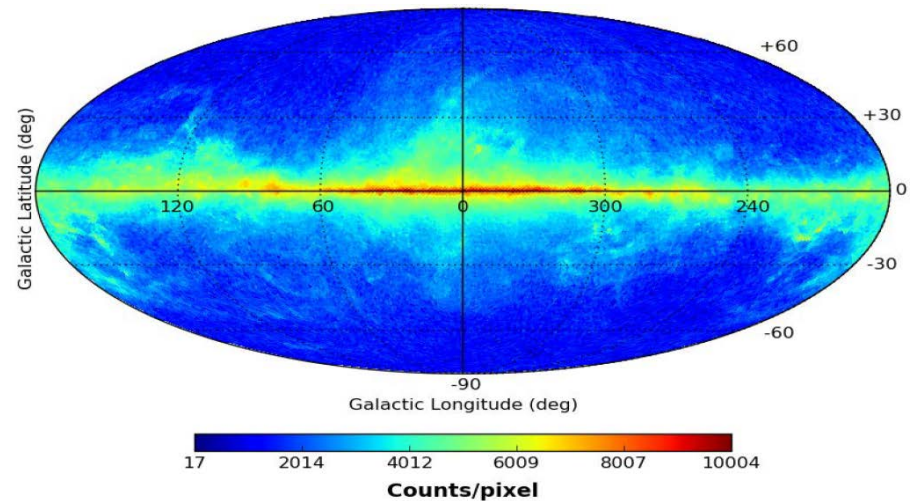
HERD specifications

Item	Value
Energy range (e/ γ)	10 GeV-100 TeV(e); 0.5 GeV-100 TeV (γ)
Energy range (CR)	30 GeV—3 PeV
Angle resolution	0.1 deg.@10 GeV
Charge measurement resolution	0.15-0.2 c.u
Energy resolution (e)	1-2%@200 GeV
Energy resolution (p)	20-30%@100 GeV – PeV
e/p separation	$\sim 10^{-6}$
G.F. (e)	$>3 \text{ m}^2\text{sr}@200 \text{ GeV}$
G.F. (p)	$>2 \text{ m}^2\text{sr}@100 \text{ TeV}$
Pointing	Zenith
Field of View	$\pm 70 \text{ deg}$ (targeting $\pm 90 \text{ deg}$)
Measur. accuracy of attitude	$<0.1 \text{ deg}$
Measur. accuracy of angular speed	$<0.005 \text{ deg/s}$
Lifetime	$>10 \text{ years}$

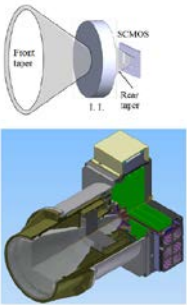
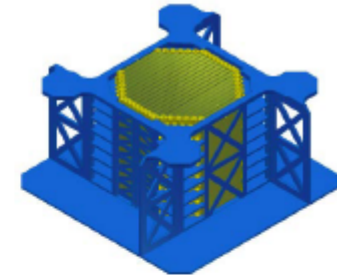
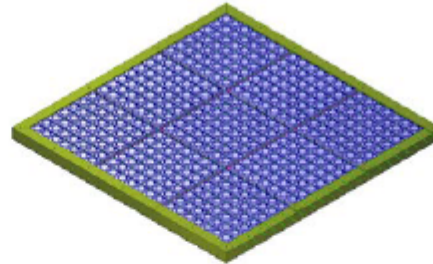
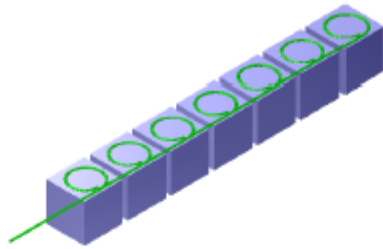
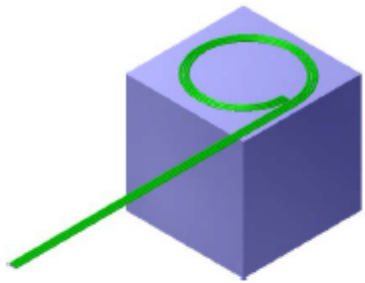
HERD: some performance plots



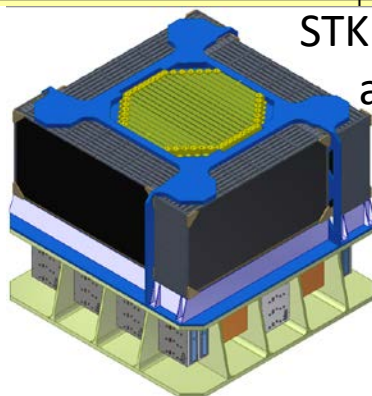
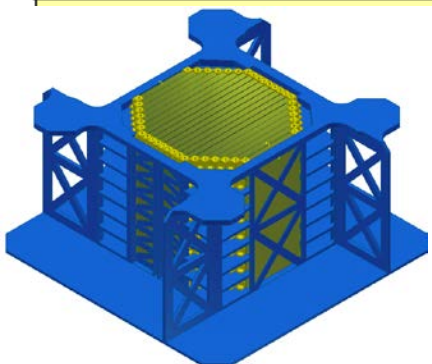
HERD 5yrs, photon map, $E_\gamma > 1\text{GeV}$



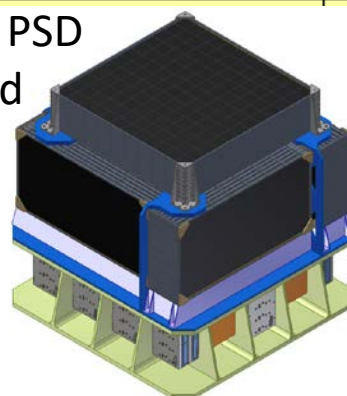
The HERD Calorimeter



item	Value	Note
Type of crystal	LYSO	
Nuclear Interaction Length	3 (55 X_0)	~ 21 LYSO crystals
Number of crystals	~7500	
Crystal dimension	3cm*3cm*3cm	
Fiber readout	3 WLSF/crystal	Low range, high range & trigger

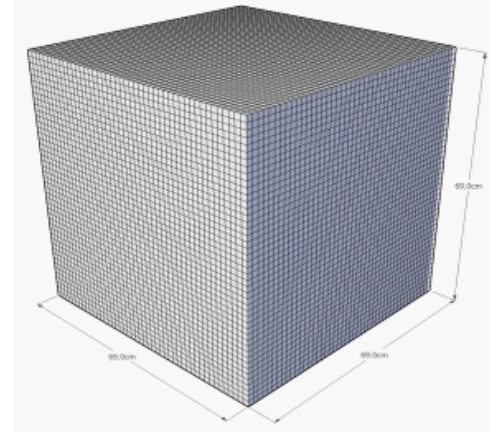
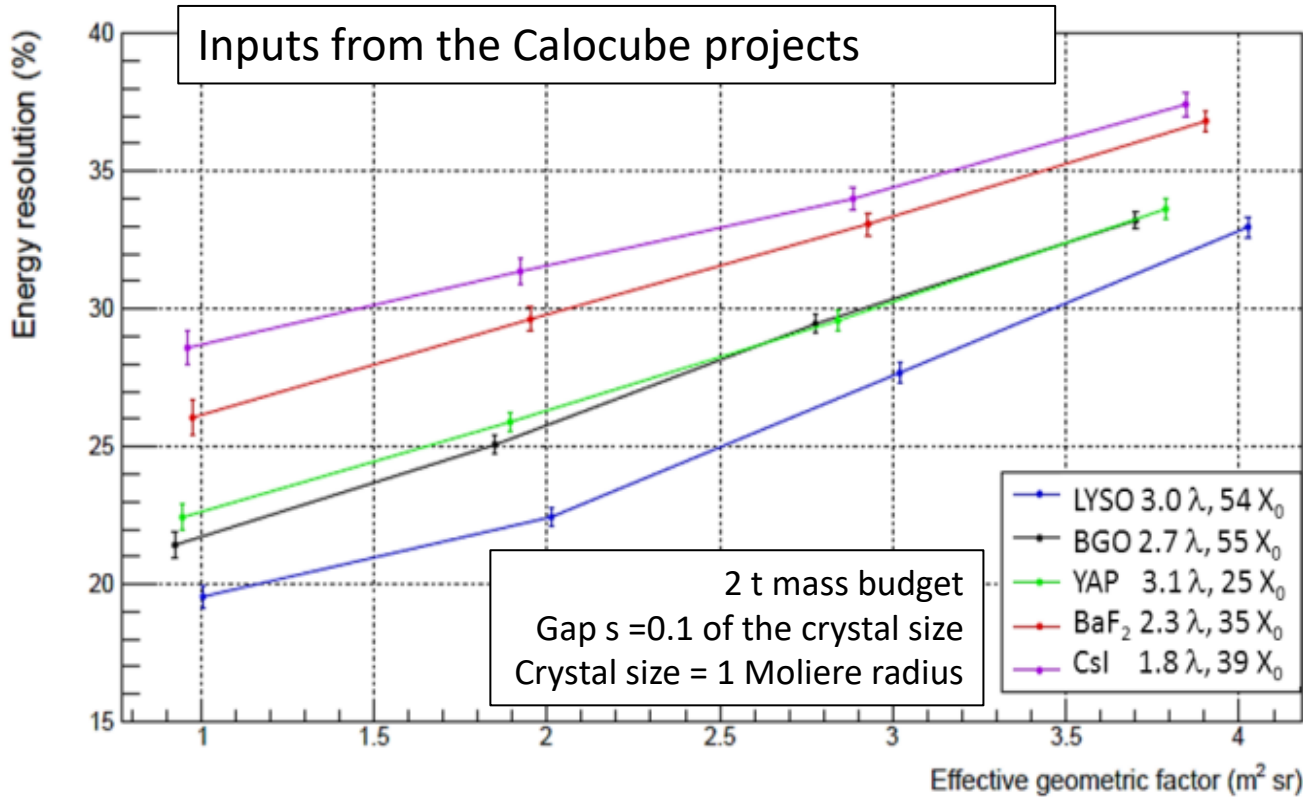


STK and PSD added

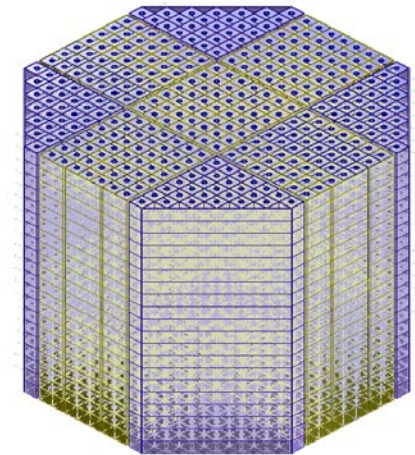


Possible readout of part of crystals with PhotoDiodes (Calocube) for calibration and extended dynamic range

Optimizing scint. / shape / readout



LYSO size: 3cm*3cm*3cm
 Detector dimension: 21
 LYSO number: ~9300
 Calo weight: ~ 1850 kg



LYSO size: 3cm*3cm*3cm
 Detector dimension: 21
 LYSO number: ~7500
 Calo weight: ~ 1500 kg

Photodiodes

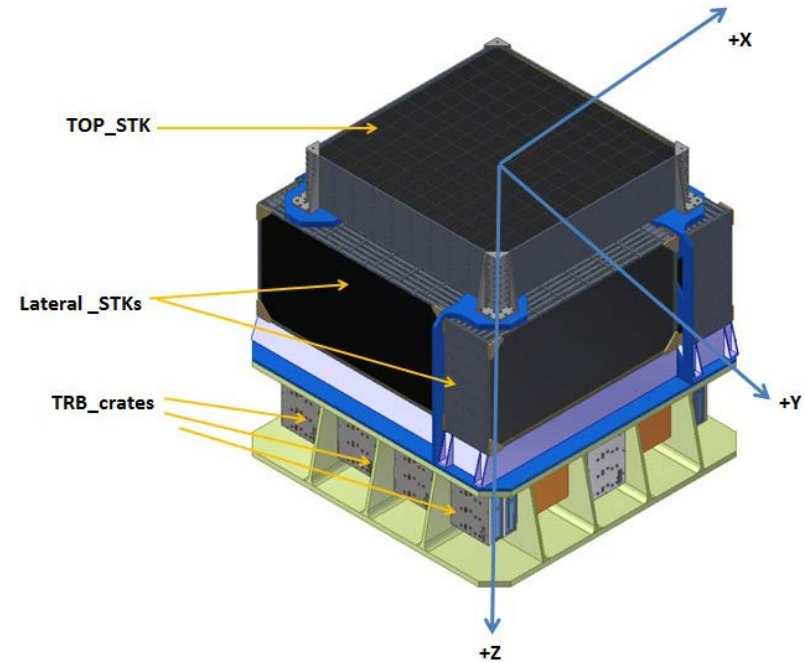


	VTH2090	VTP9412H
Active area (mm^2)	84.6	1.6
Sp.response range/peak (nm)	400÷1100 / 960	400÷1150 / 925
C_j (pF)	70 @30V	6 @15V

The HERD Si-Tracker

- CR/e trajectory
- Gamma ray conversion & tracking
- Complementary charge measurement

Item	Value
Coverage ratio	>80%
Z measurement	Z = 1 - 20 (26); 0.1-0.15 c.u
Angle resolution	0.1 deg. @10 GeV
Layers of SSD	6 X/Y (top); 3/6 X/Y (Lateral)
Active converter	1 R.L.
Dead time	<2 ms
Working mode	External trigger
Eff. Area (top)	~133 cm*133 cm
Eff. Area (lateral)	~114 cm*66.5 cm
Channels	~240,000/368,000

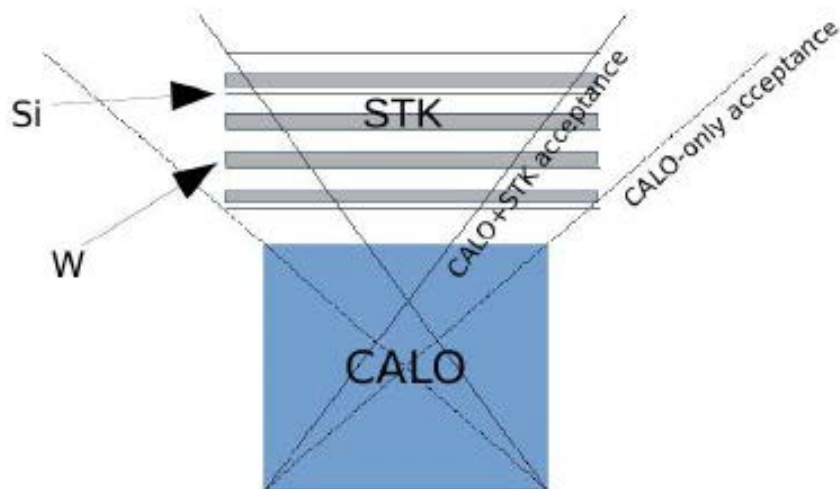


Based on the experience with AGILE, AMS-02, FERMI, DAMPE missions

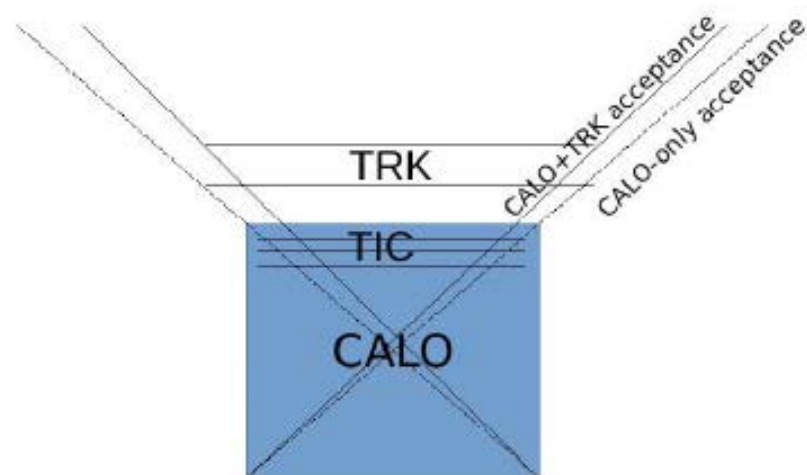
TIC: Tracker In Calorimeter

A possible TIC design is under study to:

- Optimize photon tagging and direction reconstruction
- Give multiple charge measurements (CR identification)
- Maximize calorimeter mass, i.e acceptance for the CR studies



Standard Design

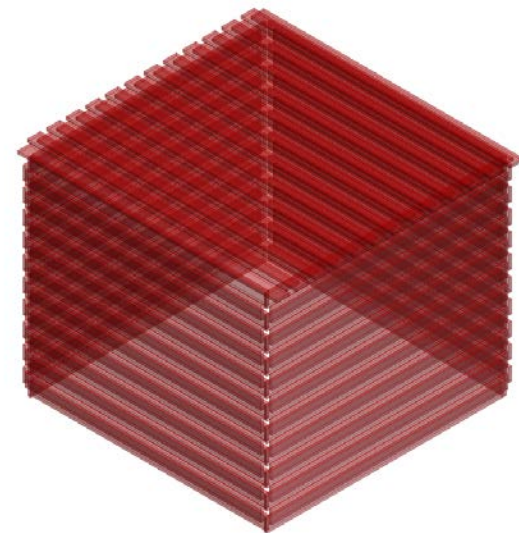


TIC Design

The HERD PSD

plastic scintillator detector

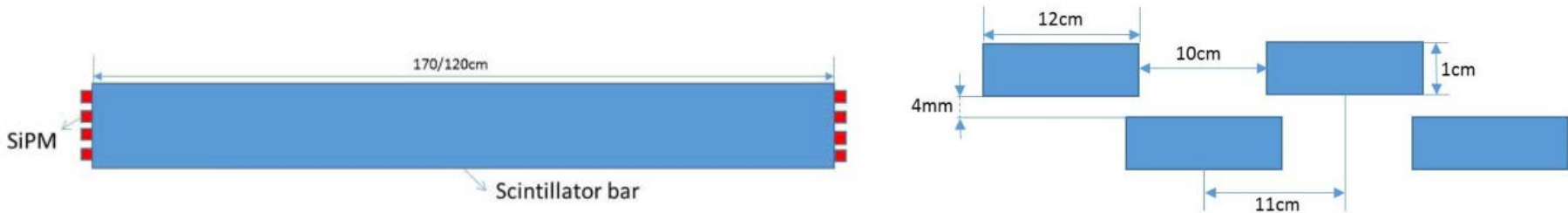
- Low energy gamma identification
- CR identification by Charge Measurement
- Design
 - 1 X/Y layer on top and 4 lateral sides
 - X layer for LE photon trigger
 - X & Y layers for Z measurement and e/gamma discrimination
 - 1 X layer on bottom side
 - SiPM + IDE3380 ASIC
 - Low & high range to cover Z=1-26
 - Redundancy SiPMs



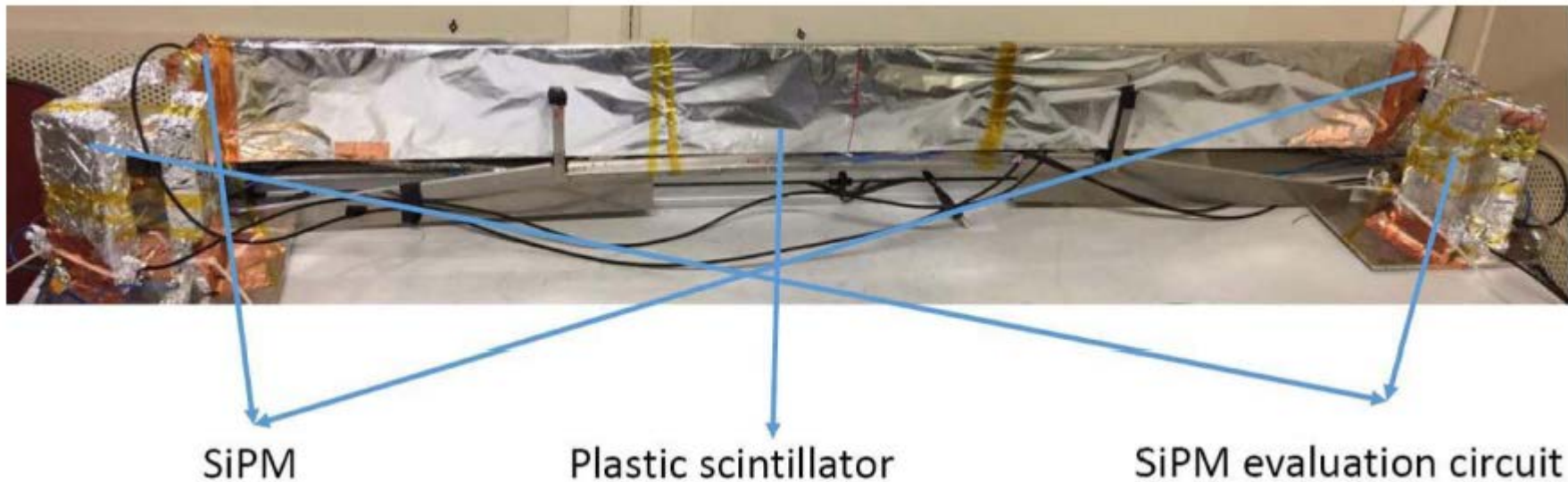
Bars vs Tiles layout
resulting from the
optimization of efficiency /
mechanics / no. channels
and backsplash effects

Alternative approach: tile geometry

The HERD PSD plastic scintillator detector

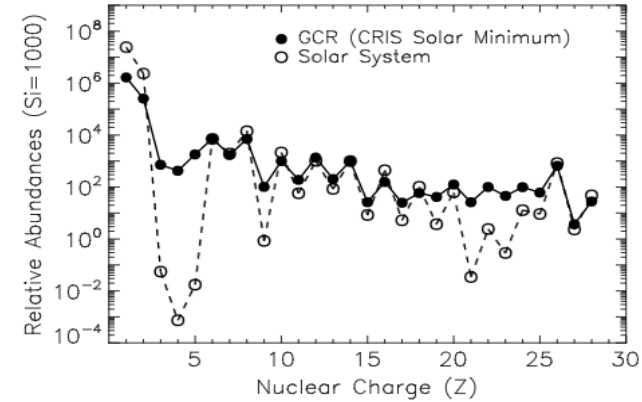
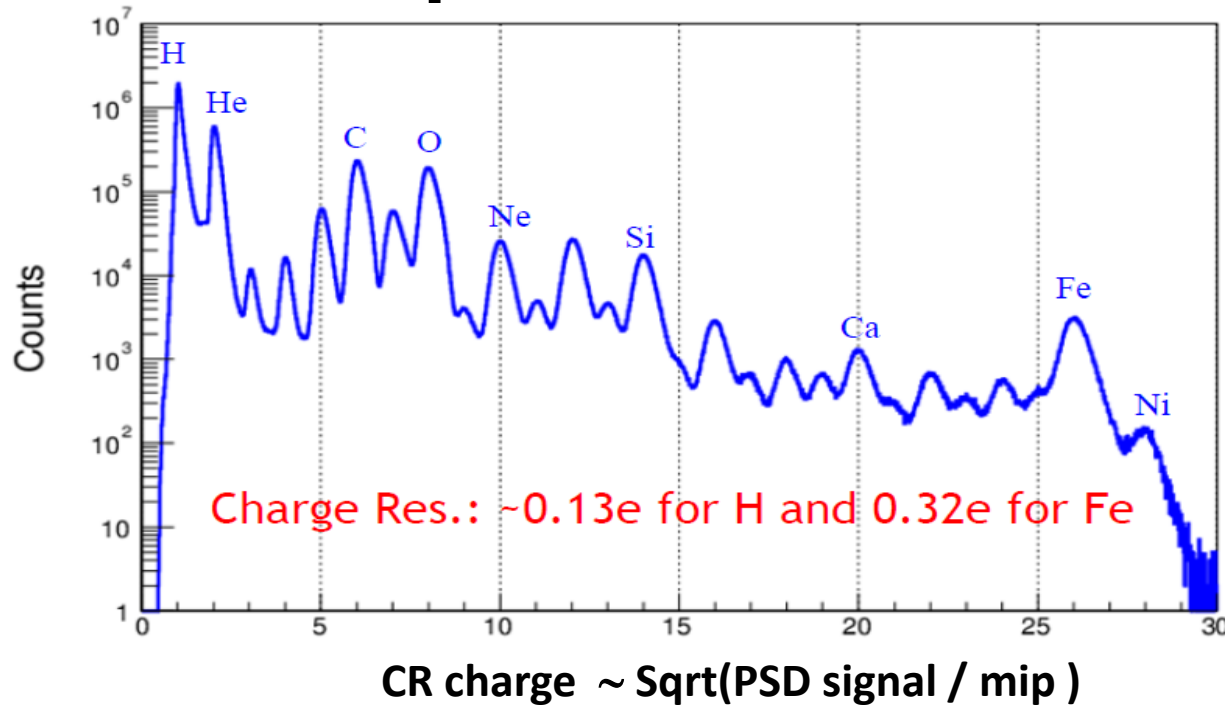


At each end of PS, 4 redundant SiPMs attached to readout as 2 low range signals and 2 high range signals

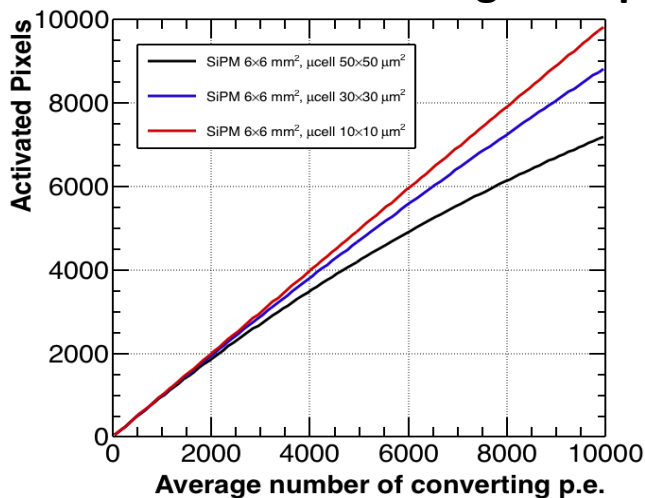


The HERD PSD

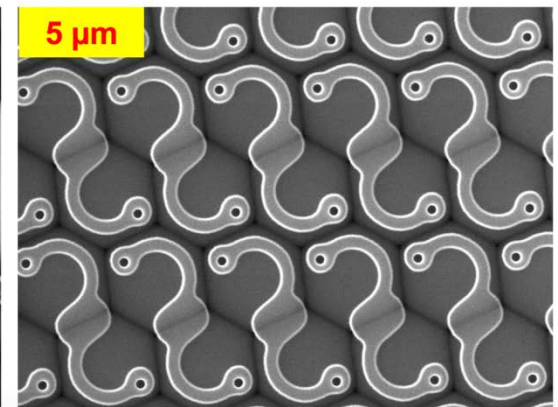
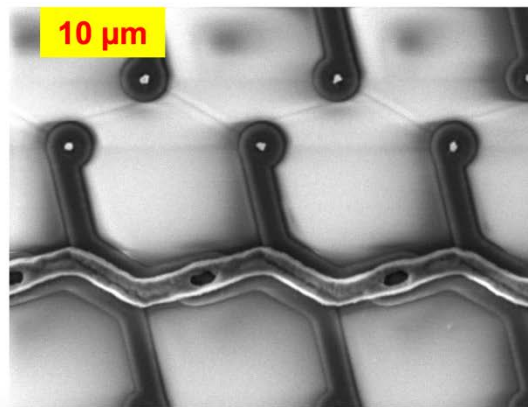
plastic scintillator detector



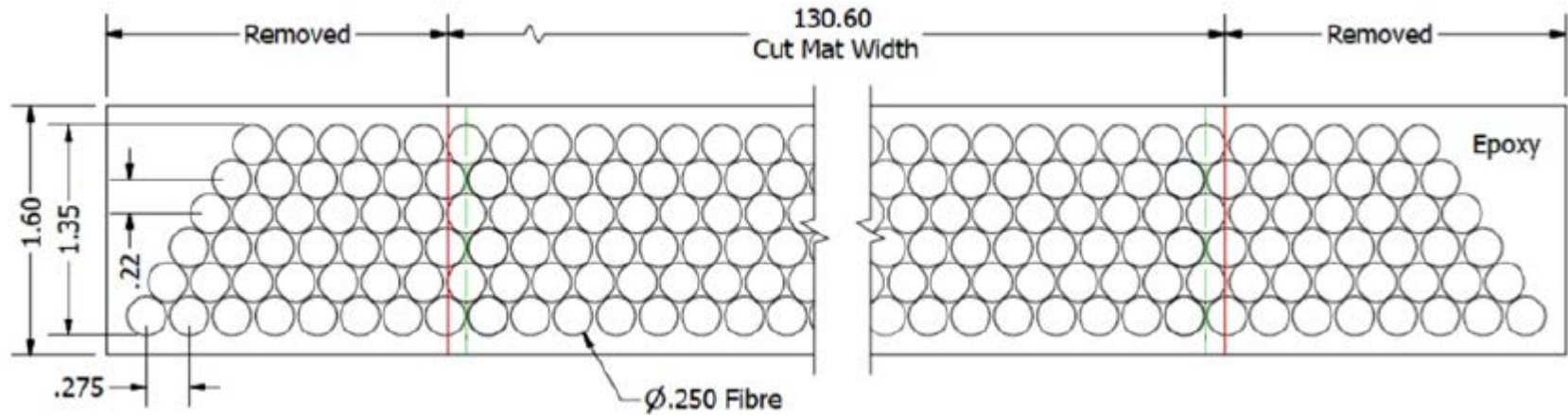
← **DAMPE preliminary results**



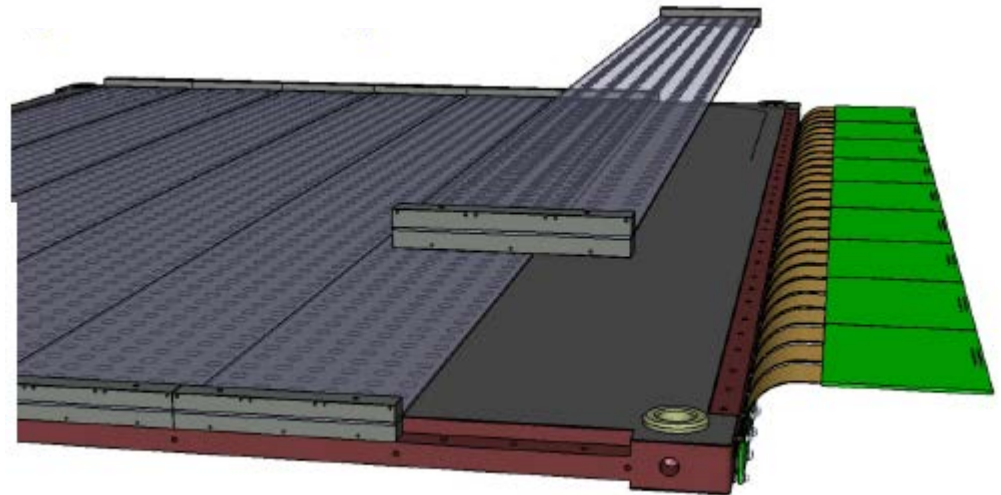
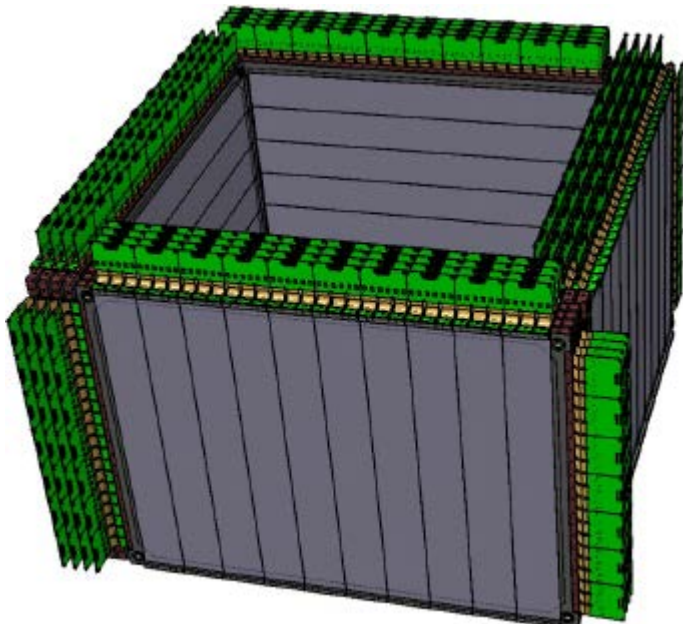
Need a dynamic range $\sim 10^3$

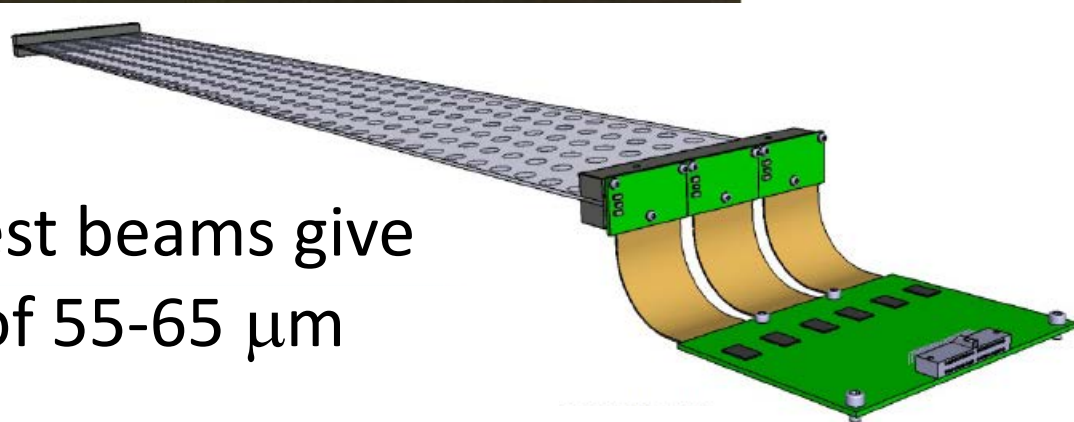
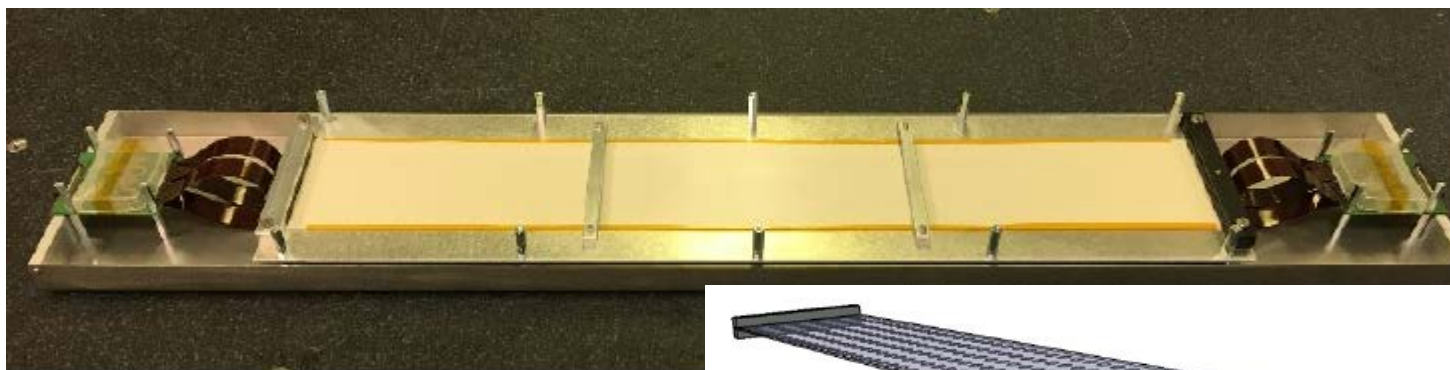


SciFi Tracker with SiPM readout

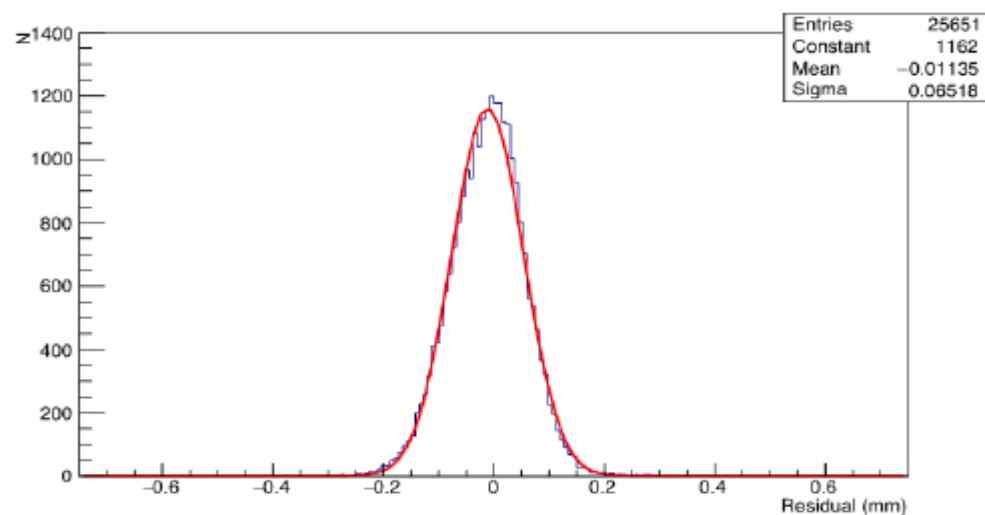
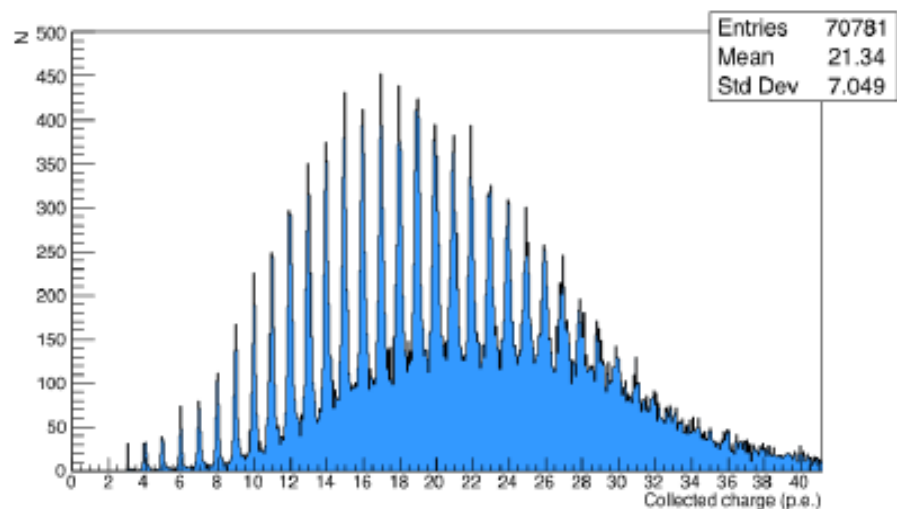


HERD mat: 97.80 mm width + 200 μ m inter-mat gap to match for 3 SiPM





Preliminary results on test beams give resolutions at the level of 55-65 μm



The HERD Collaboration

Already 6 international workshops on HERD, with the participation from several institutes:

- **China**

CSU, IHEP, XIOPM, PMO, USTC, IGG, XAO, NAOC, TSU, GXU, PKU, NJU, YNU, NBU, SYSU, University of Hong Kong (HKU), National Central University (NCU)

- **Italy**

INFN Perugia, University & INFN Firenze, University & INFN Bari, University & INFN Pisa, University & INFN Trento, University of Salento and INFN Lecce, IAPS/INAF, University & INFN Catania, University & INFN Napoli, University & INFN Trieste, Gran Sasso Science Institute

- **Switzerland:** University of Geneva

- **Sweden:** KTH

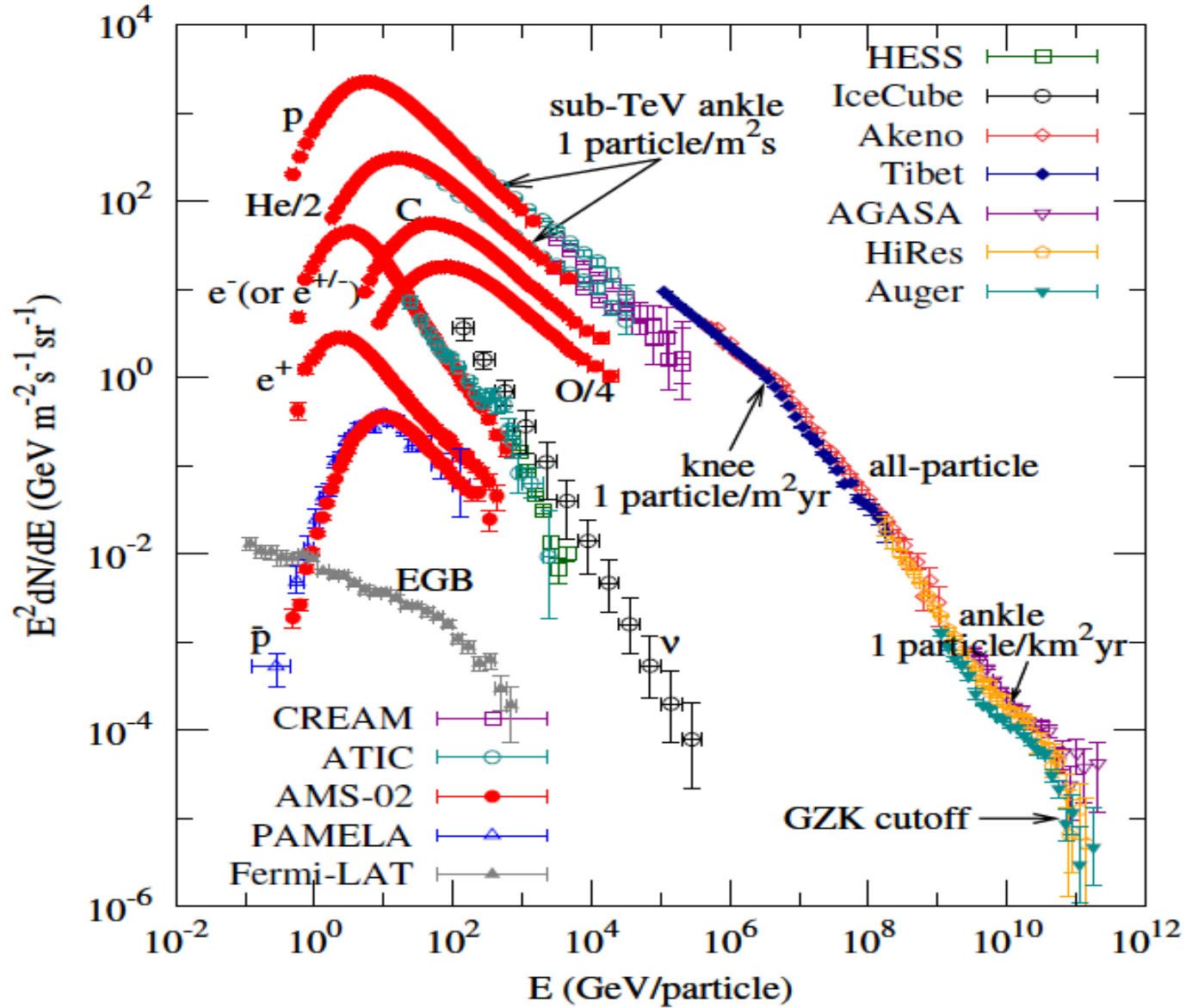
- **Spain:** CIEMAT

- **Germany:** KIT

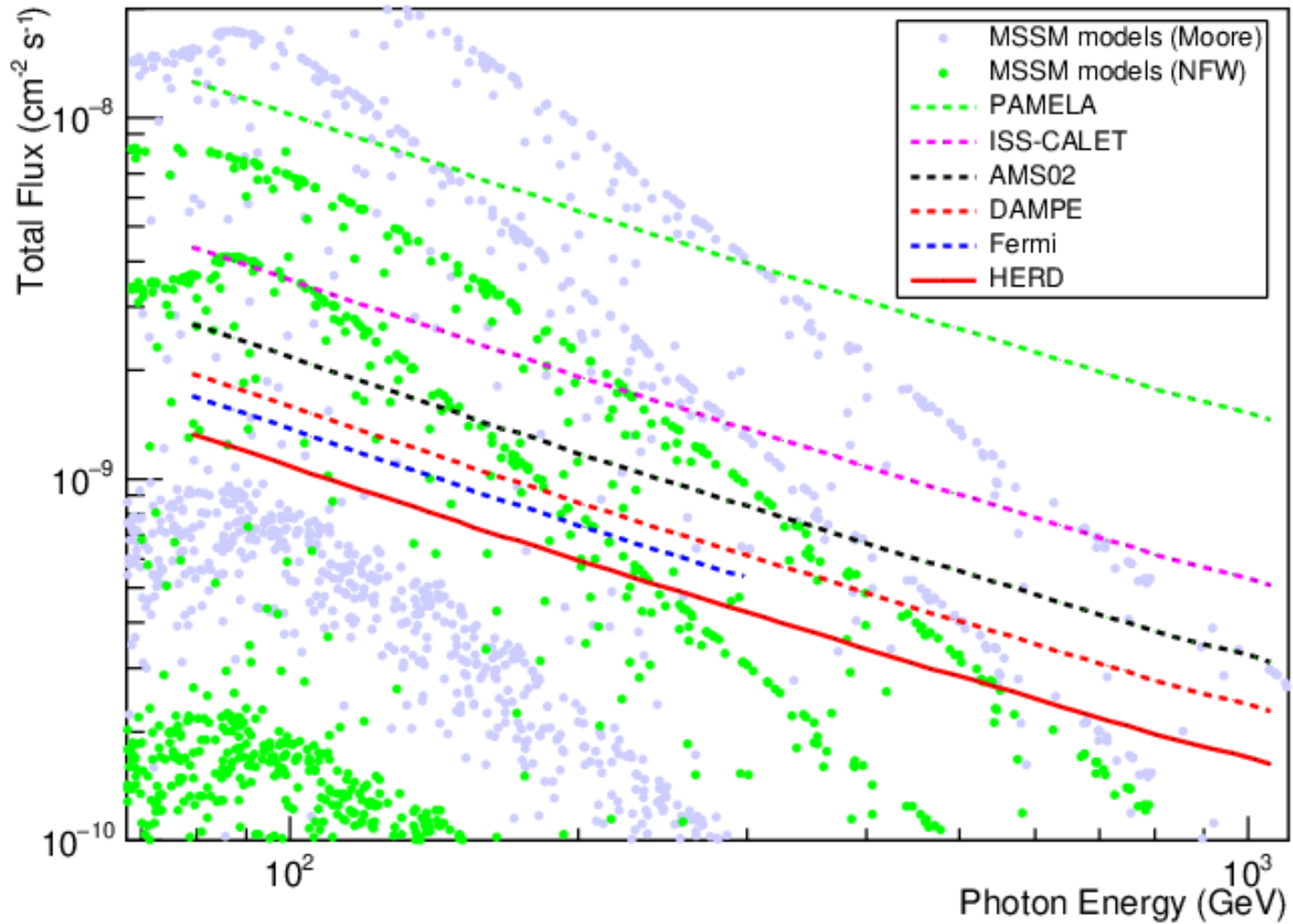
- **Russia:** Lebedev Physical Institute

- **Japan:** University of Tokyo

More Stuff

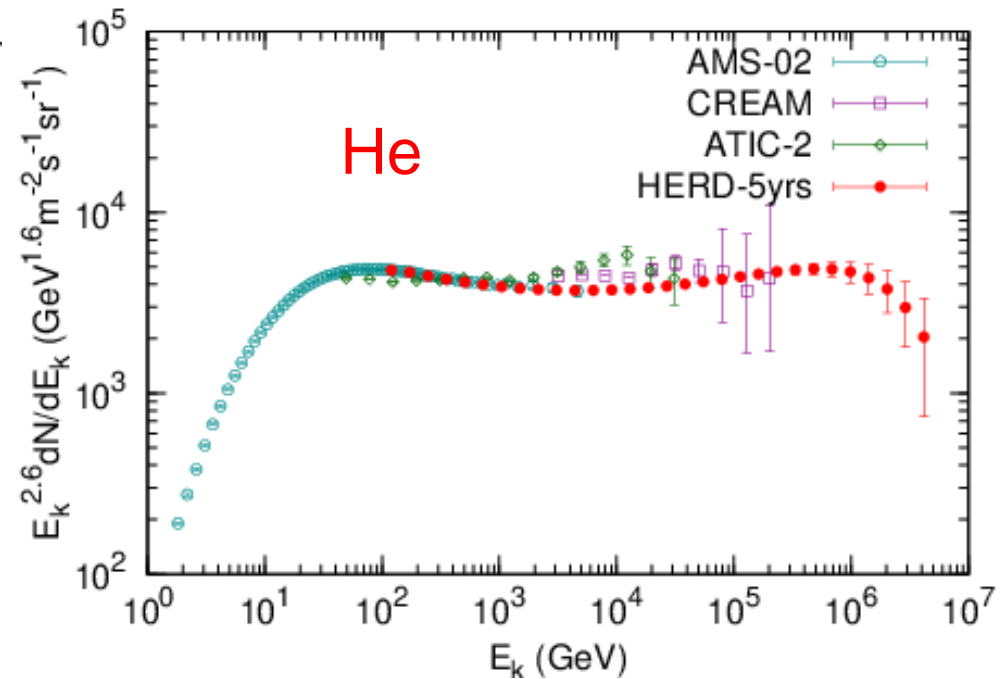
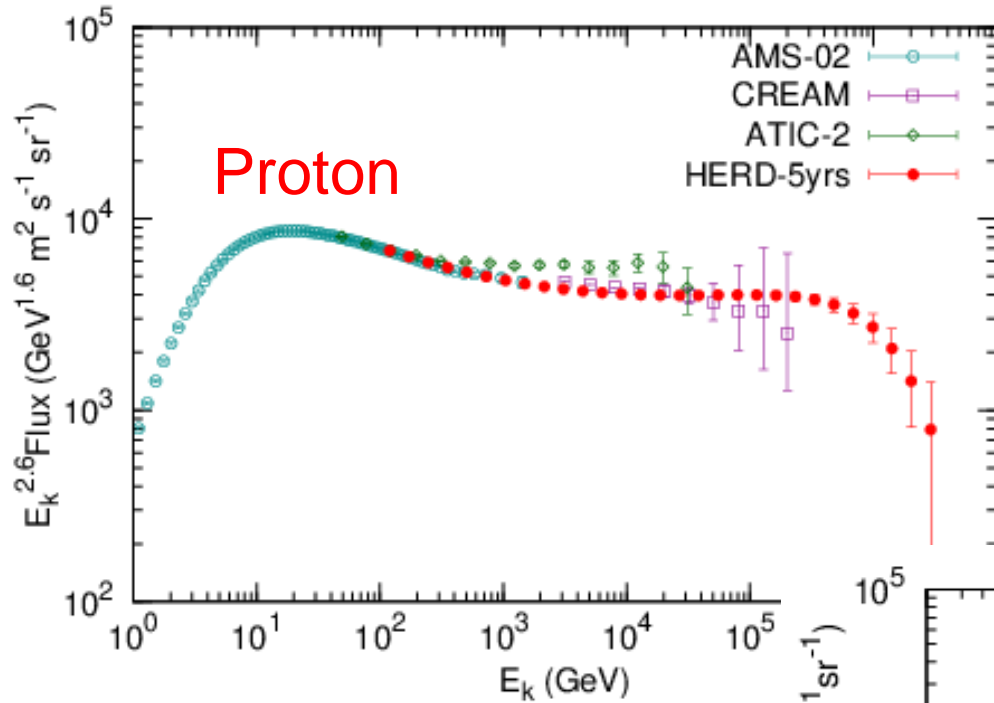


Sensitivity for gamma ray line by different experiments

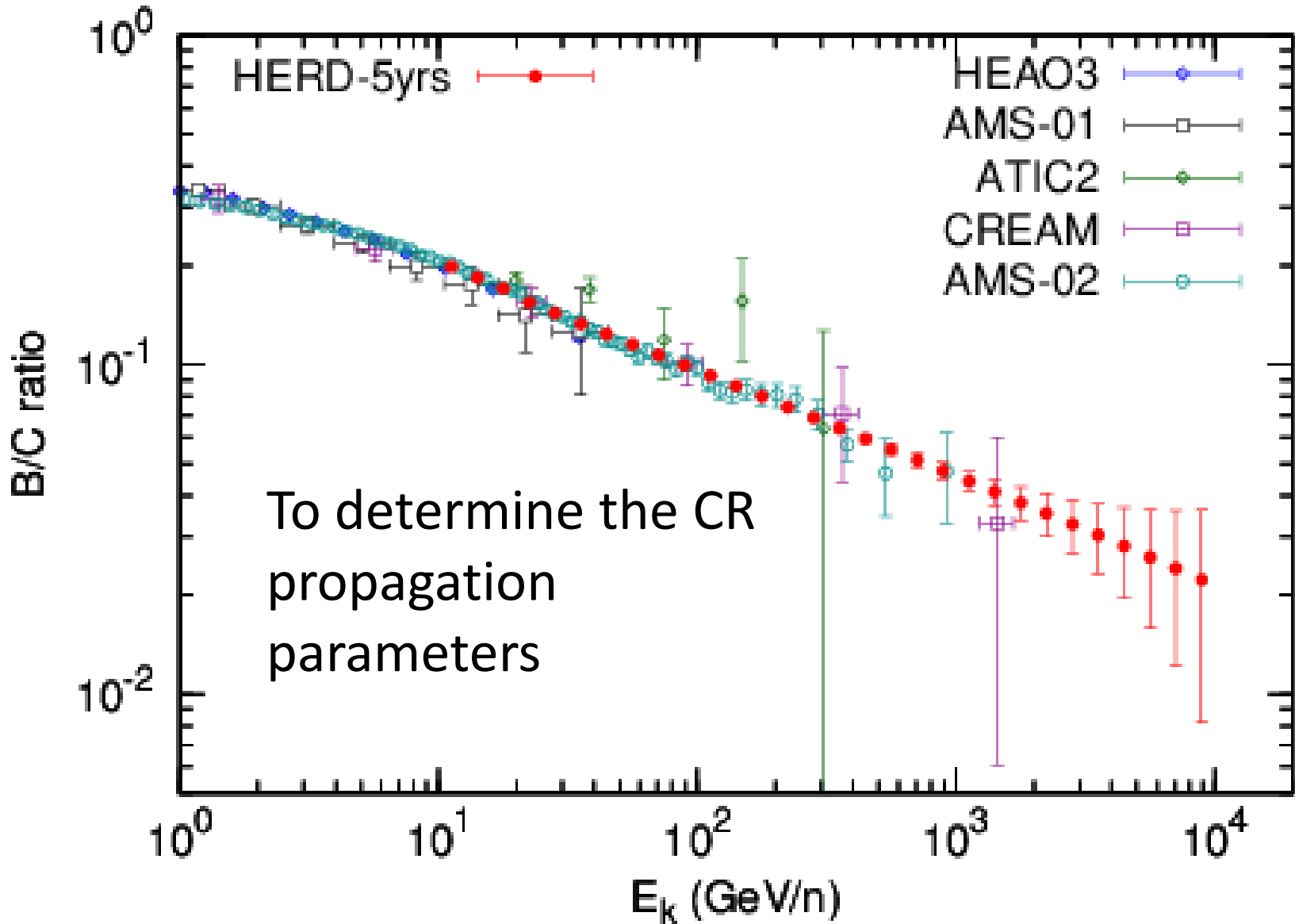


PAMELA: 2006-2016 CALET: 2015-2020; AMS: 2011-2024;
DAMPE: 2015-2020; Fermi: 2008-2018; HERD: 1 year

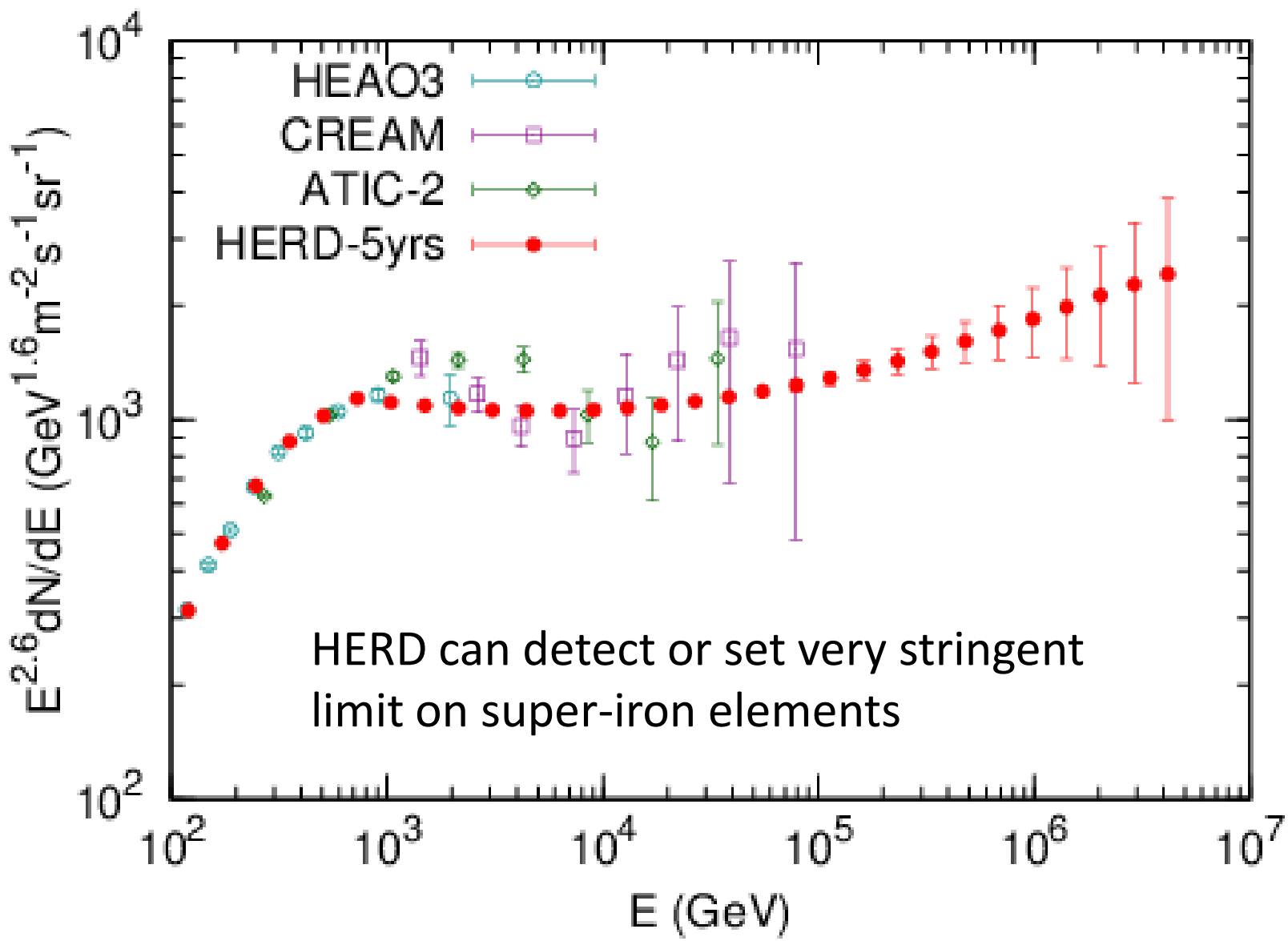
Expected HERD Proton and He Spectra



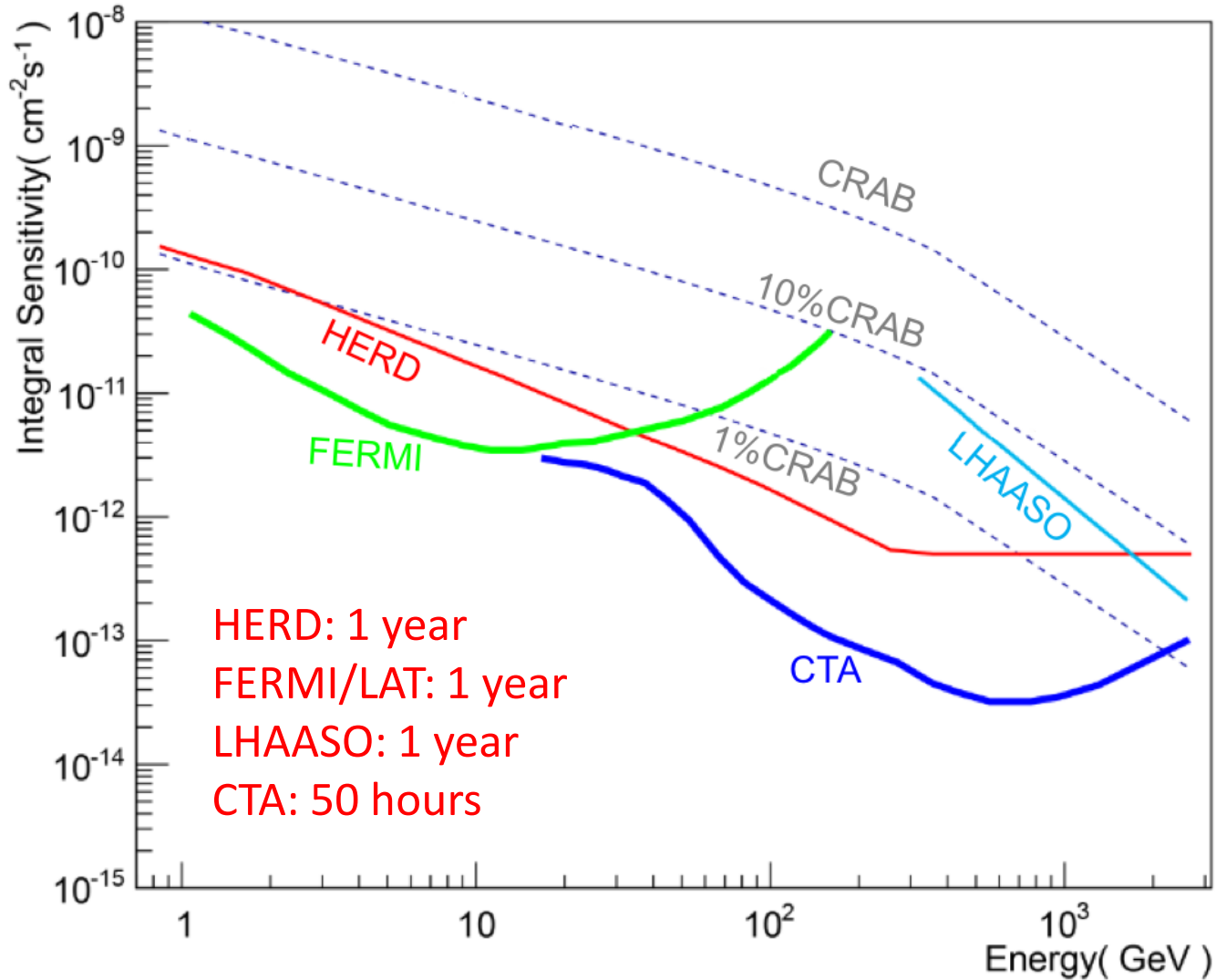
B/C measurement at HERD



Iron nucleon and super-iron elements

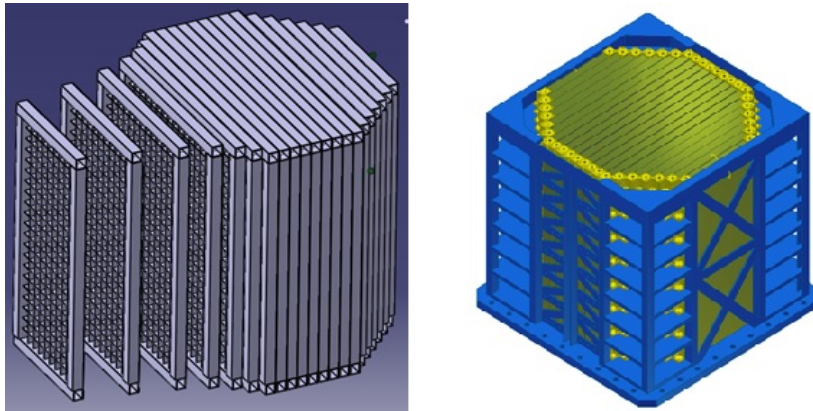
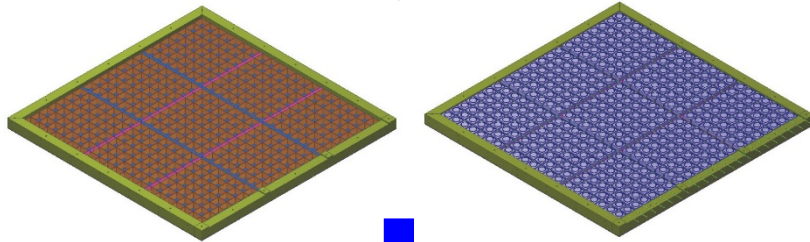
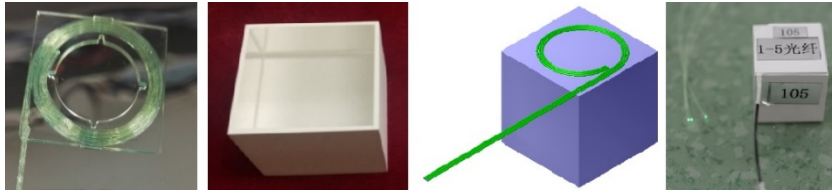


Gamma-ray sky survey

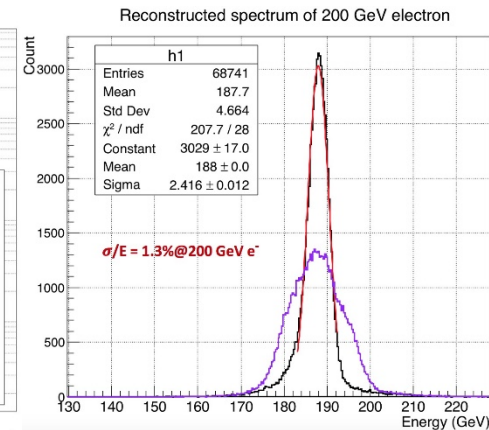
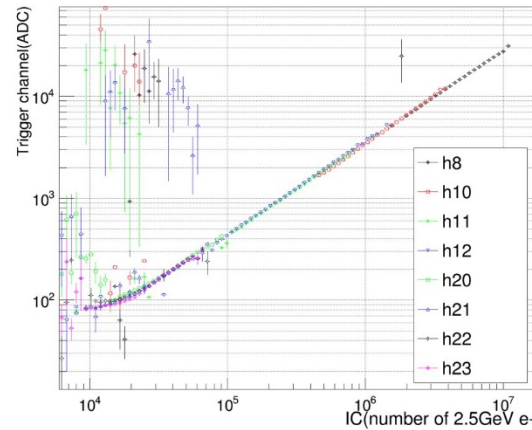


Expected HERD gamma-ray sky survey sensitivity (5σ)

Payload design - CALO



- CALOrimeter (3 N.I.L. and 55 R.L.)
 - A 3-d crystal array (~7500 LYSO)
 - IsCMOS camera
 - Trigger sub-system
- Novel readout method
 - WLSF + IsCMOS
 - Linearity of LYSO+WLSF is verified.
 - Energy measurement of WLSF + IsCMOS is verified.



Alternative approach: Photo diode readout

CALO – ISCMOS sub-system

- **IsCMOS** to collect WLSF photons
 - Faster: Global shutter; ROI readout
 - Lower noise
- Accurate energy measurement
 - 1 fiber ~ 20*20 pixels
 - Saturation effect to increase DR

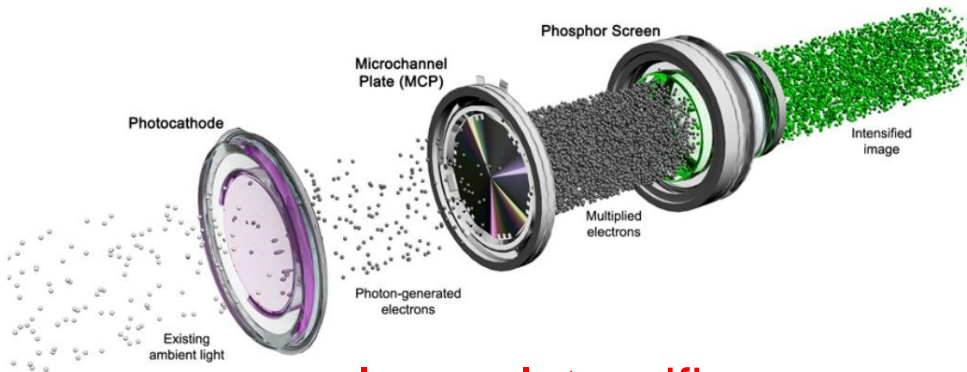
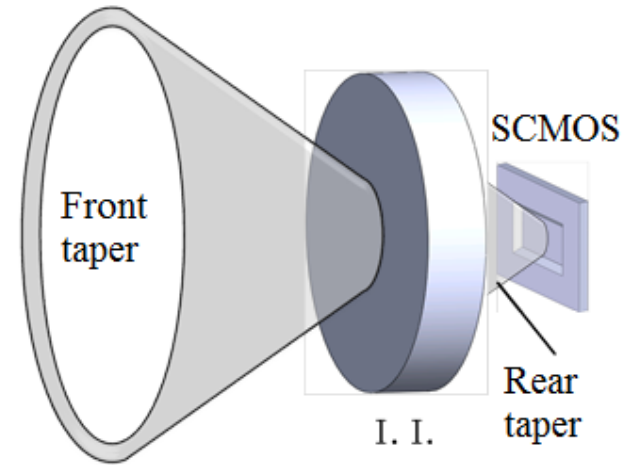
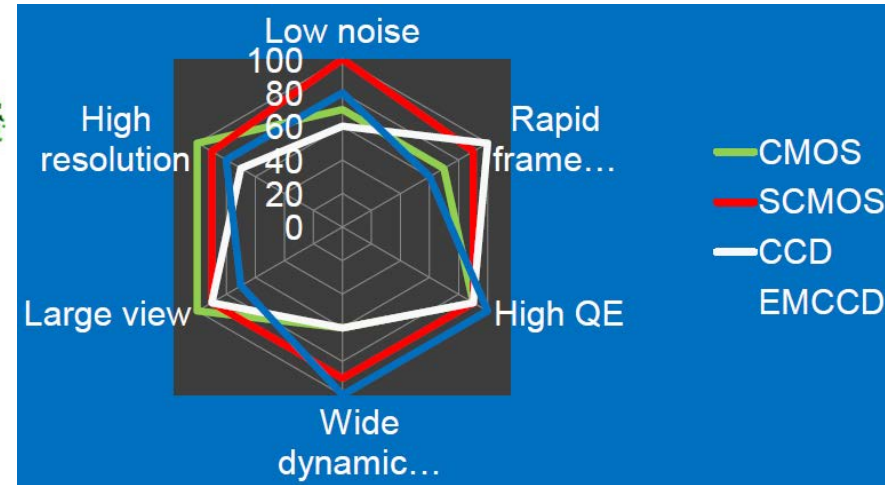
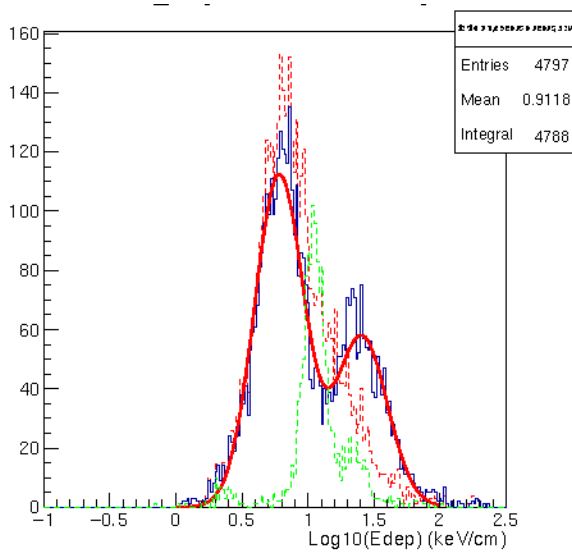


Image Intensifier

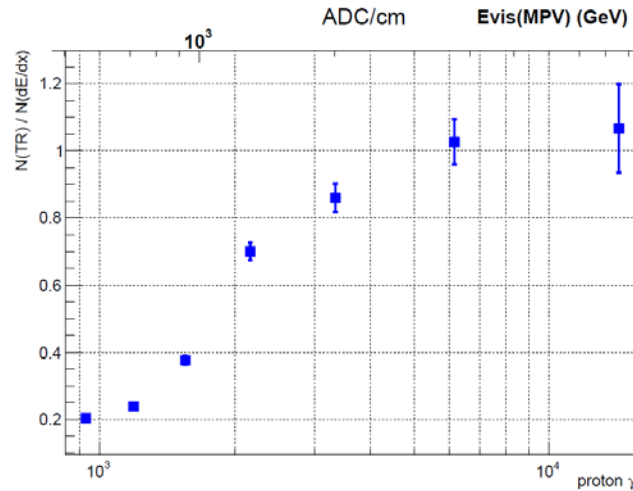


TRD payload

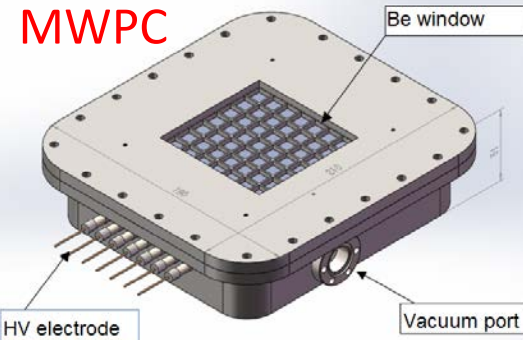
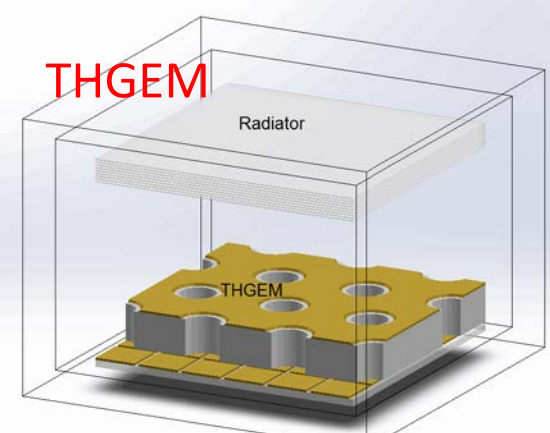
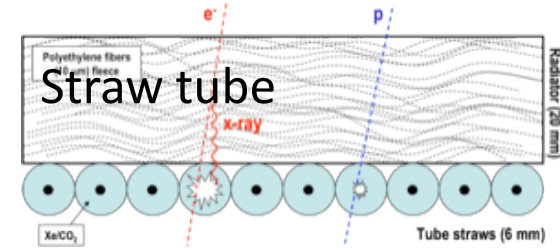
- Energy calibration of TeV protons and other nuclei
- A complete calibration in 2-3 months in-orbit operation



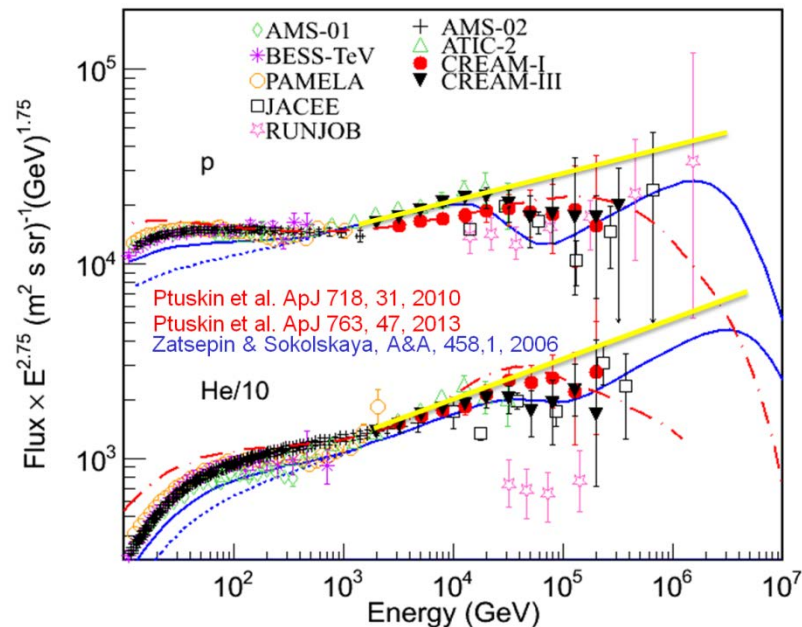
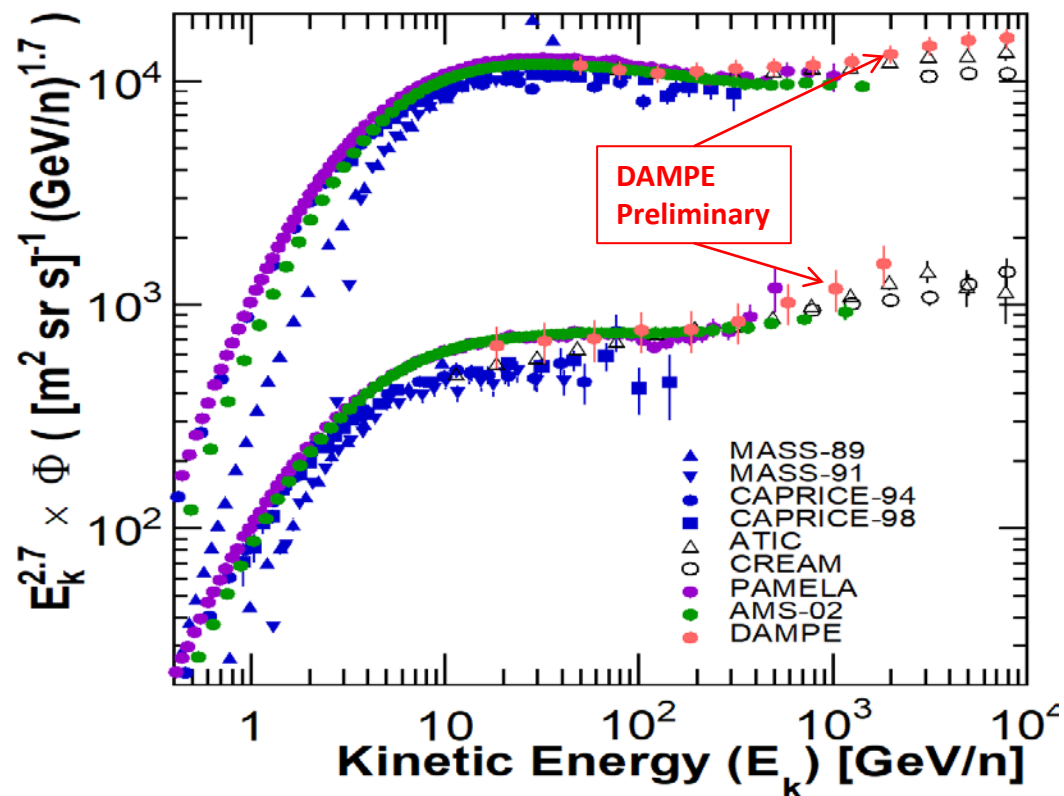
MWPC energy response to [2.25, 2.5] TeV protons



2 months simulated observation, $\sim 6300\text{cm}^2$ TRD.



Higher energies and secondaries

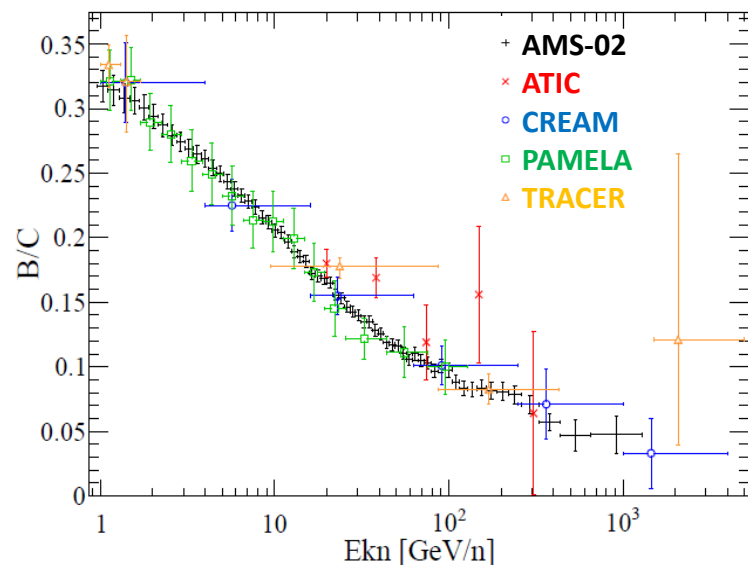


1-100 TeV

Explored by CREAM & NUCLEON

Preliminary results from DAMPE, CALET

Go to higher energies: ISS-CREAM and HERD



- **China:** CSU, IHEP, XIOPM, PMO, USTC, IGG, XAO, NAOC, TSU, GXU, PKU, NJU, YNU, NBU, SYSU, University of Hong Kong (HKU), National Central University (NCU)
- **Italy:** INFN Perugia, University & INFN Firenze, University & INFN Bari, University & INFN Pisa, University & INFN Trento, University of Salento and INFN Lecce, IAPS/INAF, University & INFN Catania, University & INFN Napoli, University & INFN Trieste, Gran Sasso Science Institute
- **Switzerland:** University of Geneva; **Sweden:** KTH; **Spain:** CIEMAT
- **Germany:** KIT; **Russia:** Lebedev Physical Institute
- **Japan:** University of Tokyo



4th HERD workshop
ASI HQs, Roma, Italy 2017.2.9