

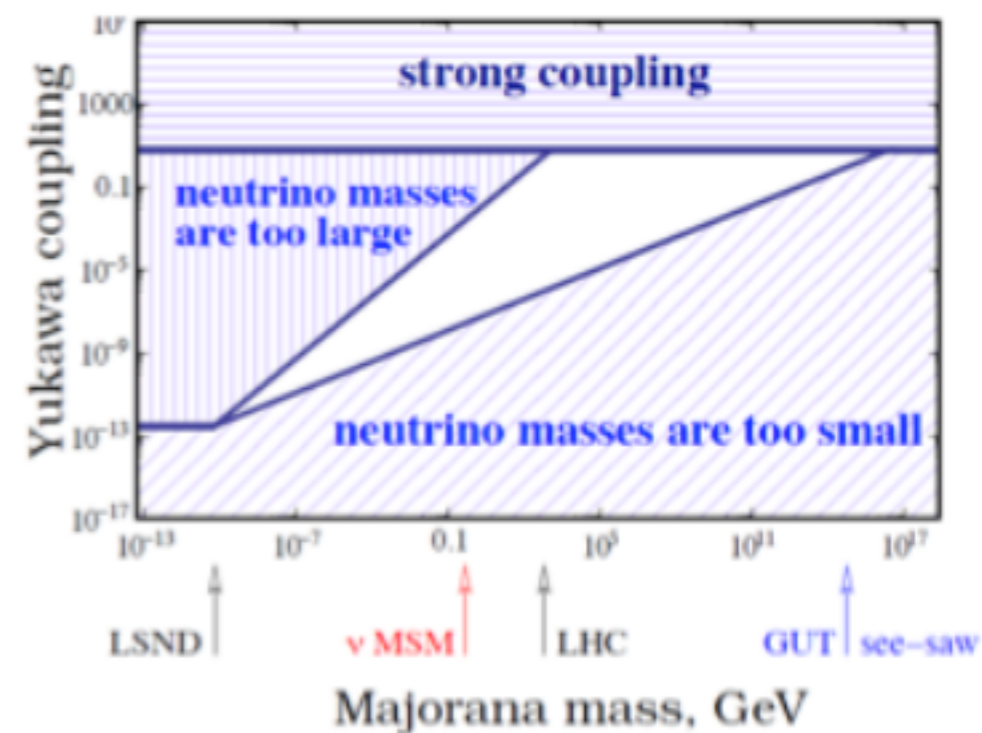
Next obvious question

Where does the ν mass come from?

—> it is clear that experimental effort (and possibly prizes...) has to be now directed in this direction!

One remarkable possibility:

see-saw mechanism (type I) with one/two/three massive and sterile Majorana-type neutrinos



$$A = \begin{pmatrix} 0 & M \\ M & B \end{pmatrix},$$

where B is taken to be much larger than M .

It has two very disproportionate **eigenvalues**:

$$\lambda_{\pm} = \frac{B \pm \sqrt{B^2 + 4M^2}}{2}.$$

The larger eigenvalue, λ_+ , is approximately equal to B , while the smaller eigenvalue is approximately equal to

$$\lambda_- \approx -\frac{M^2}{B}.$$

Thus, $|M|$ is the **geometric mean** of λ_+ and $-\lambda_-$, since the **determinant** $\lambda_+ \lambda_- = -M^2$.

If one of the eigenvalues goes up, the other goes down, and vice versa. This is the point of the name "**seesaw**" of the mechanism.

A matrice di massa dei neutrini, B componente di Majorana (O GUT), M componente di Dirac O(v.e.v. EW)

→ λ_- porta alla massa dei neutrini leggeri O(eV)

See-saw generation of neutrino masses

Most general renormalisable Lagrangian of SM particles (+3 singlets wrt SM gauge group):

$$L_{singlet} = i\bar{N}_I \partial_\mu \gamma^\mu N_I - Y_{I\alpha} \bar{N}_I^c \tilde{H} L_\alpha - M_I \bar{N}_I^c N_I + h.c$$

Yukawa term: mixing of N_I with active neutrinos to explain oscillations

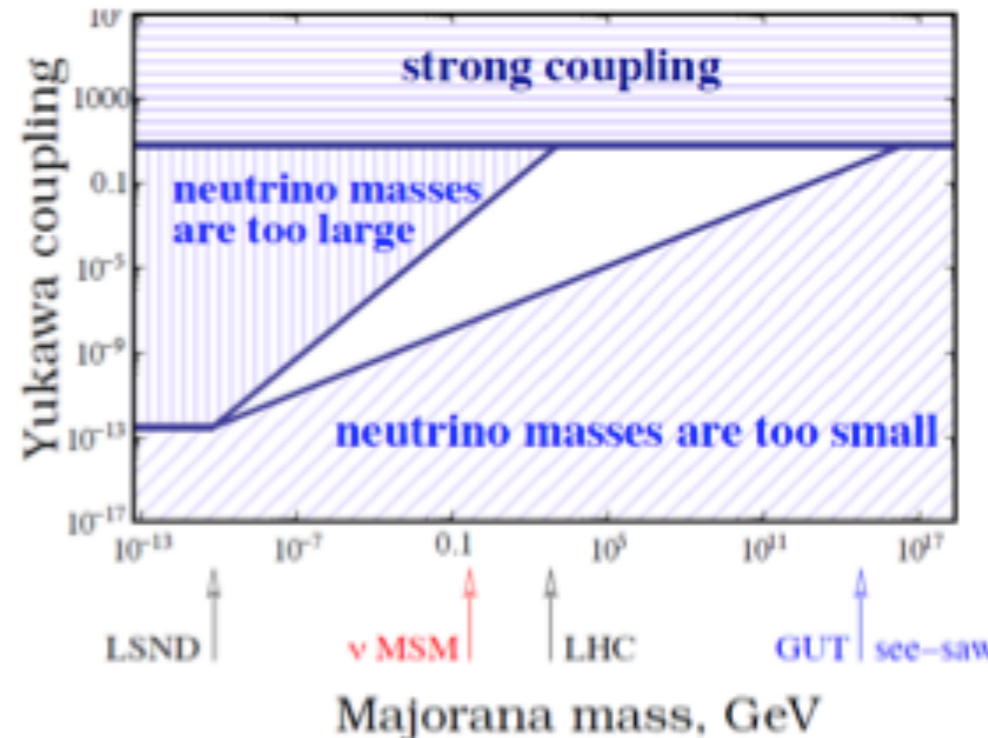
Majorana term which carries no gauge charge

The scale of the active neutrino mass is given by the see-saw formula: $m_\nu \sim \frac{m_D^2}{M} = M^2/B$
 where $m_D \sim Y_{I\alpha} v$ - typical value of the Dirac mass term

$$v \sim 246 \text{ GeV}$$

Example:

For $M \sim 1 \text{ GeV}$ and $m_\nu \sim 0.05 \text{ eV}$
 it results in $m_D \sim 10 \text{ keV}$ and Yukawa coupling $\sim 10^{-7}$





eV keV MeV GeV TeV → 10^{14} GeV

neutrino masses through seesaw

very long lifetime & warm → dark matter

baryon asymmetry (BAU) through leptogenesis

➤ Which observable effects allow to test such models?

- $M \gg \Lambda_{EW}$: “Non-unitarity effects” (indirect tests)
- $M \sim \Lambda_{EW}$: On-shell heavy neutrino effects (direct tests at colliders, various promising signature processes)
- $M < m_W$: Very sensitive searches via “displaced vertices” at colliders

The ν MSM and its fellows

3 Majorana (HNL) partners of ordinary ν , with $M_N < M_W$

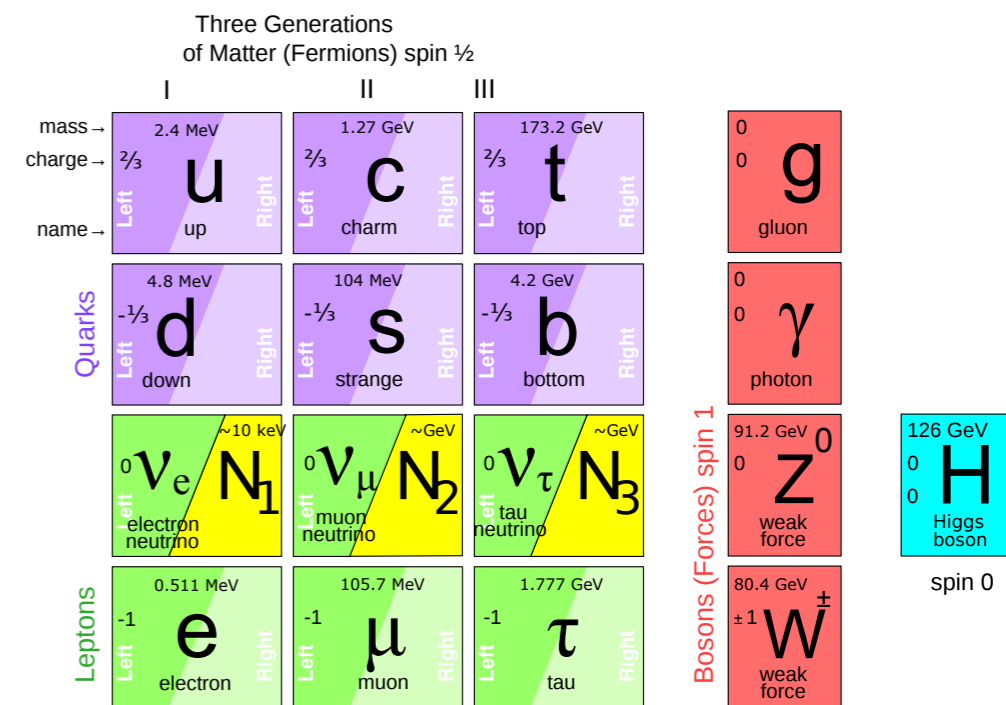
In a peculiar parameter space (N_2 and N_3 almost degenerate in mass and with $m=O(\text{GeV})$ and N_1 decoupled with $m=O(\text{keV})$), ν MSM explains:

neutrino masses (see-saw), baryogenesis (via leptogenesis) and DM (N_1)! (but, in this version of the theory, DM has to be generated outside the ν MSM, by e.g. the decay of an inflaton \rightarrow see Higgs portal)

No hierarchy problem (if also the inflaton or the NP yielding N_1 has mass below EW scale)

Naturalness of the above parameter space comes from a $U(1)$ lepton symmetry, broken at 10^{-4} level.

+ all other, less noble but less constrained, variations of this theory



ν MSM: T.Asaka, M.Shaposhnikov PL **B620** (2005) 17
M.Shaposhnikov Nucl. Phys. B763 (2007) 49

$N_{2,3}$ production

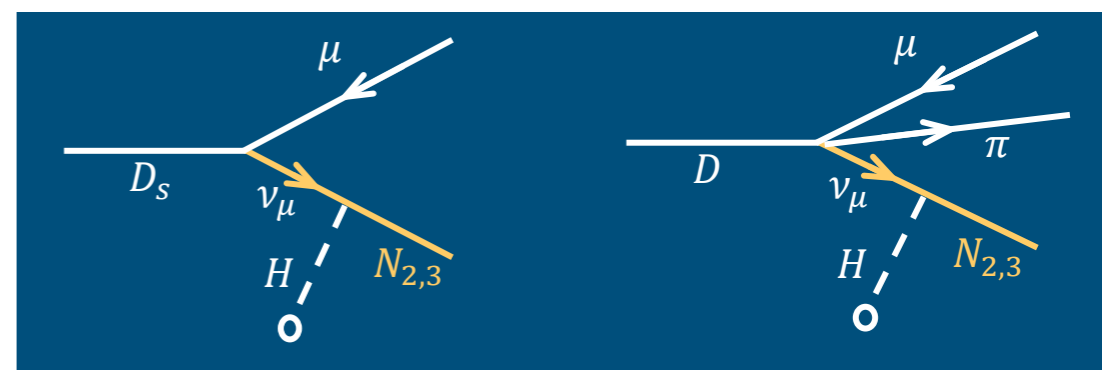
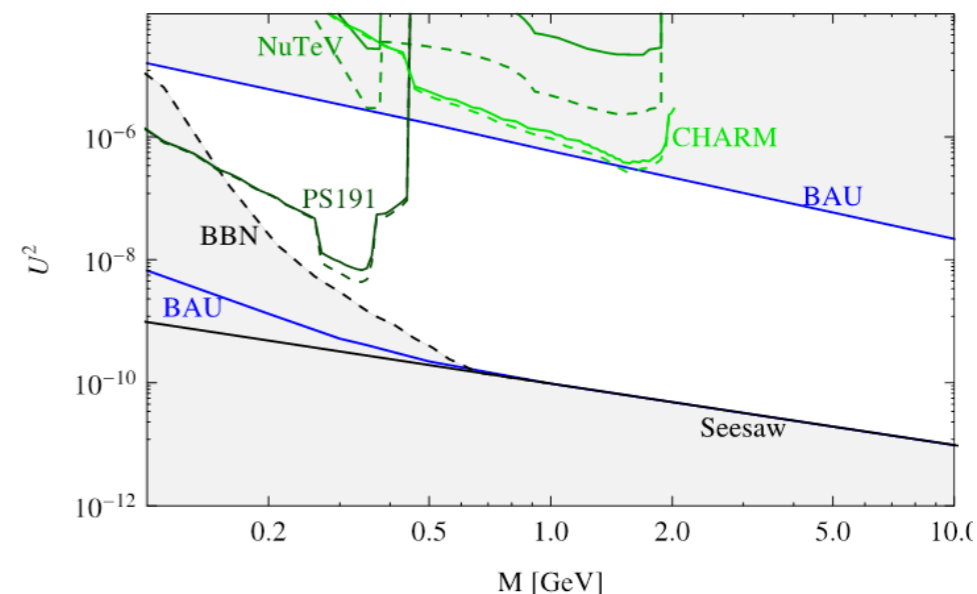
Interaction with the Higgs v.e.v.
 -> mixing with active neutrinos
 with U^2

in the ν MSSM strong limitations
 in the parameter space (U^2, m)

a lot of HNL searches in the past
 but, for $m > m_K$, with a sensitivity
 not of cosmological interest

ex. meson decays \rightarrow

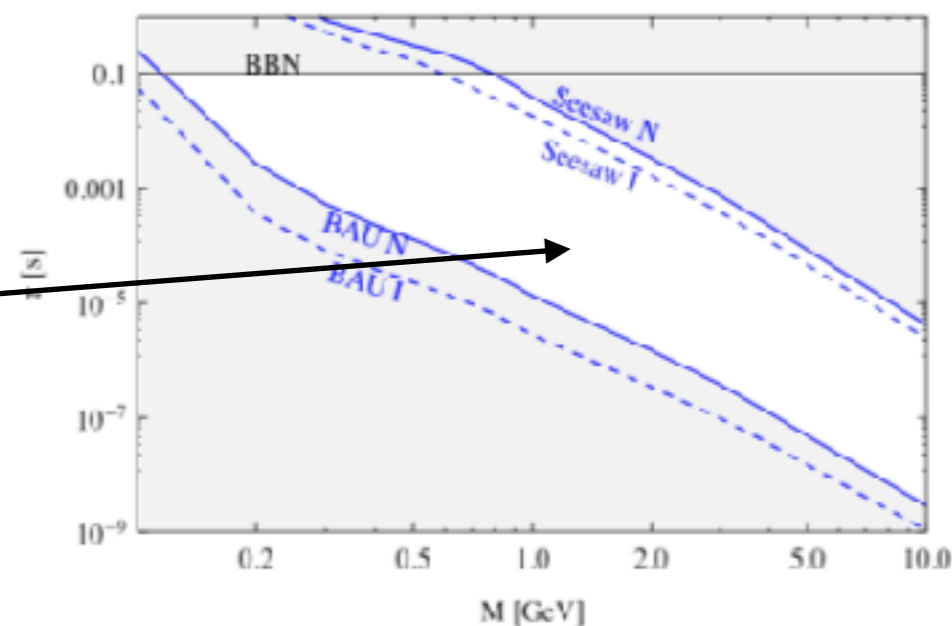
inverted mass hierarchy



$N_{2,3}$ decays

Very weak HNL-active ν \rightarrow at masses below few GeV, $N_{2,3}$ have very long lifetime

decay paths of O(km)!: for $U_{\mu}^2=10^{-7}$, $\tau_N=1.8 \times 10^{-5} \text{ s}$



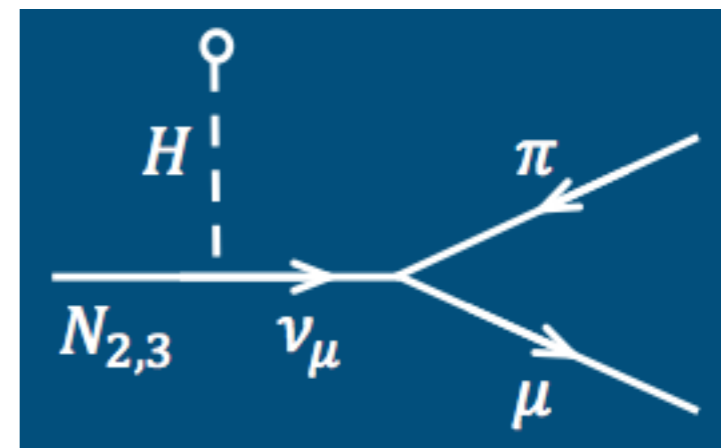
Various decay modes : the BR's depend on flavor mixing

The probability that $N_{2,3}$ decays within the fiducial volume of the experiment

$$\propto U_{\mu}^2$$

\rightarrow number of events $\propto U_{\mu}^4$ if N detected

Decay mode	Branching ratio
$N_{2,3} \rightarrow \mu/e + \pi$	0.1 - 50 %
$N_{2,3} \rightarrow \mu^-/e^- + \rho^+$	0.5 - 20 %
$N_{2,3} \rightarrow \nu + \mu + e$	1 - 10 %



Probing hidden sectors with very small couplings to SM



- High fluxes
- Displaced decays

Searches with proton beams

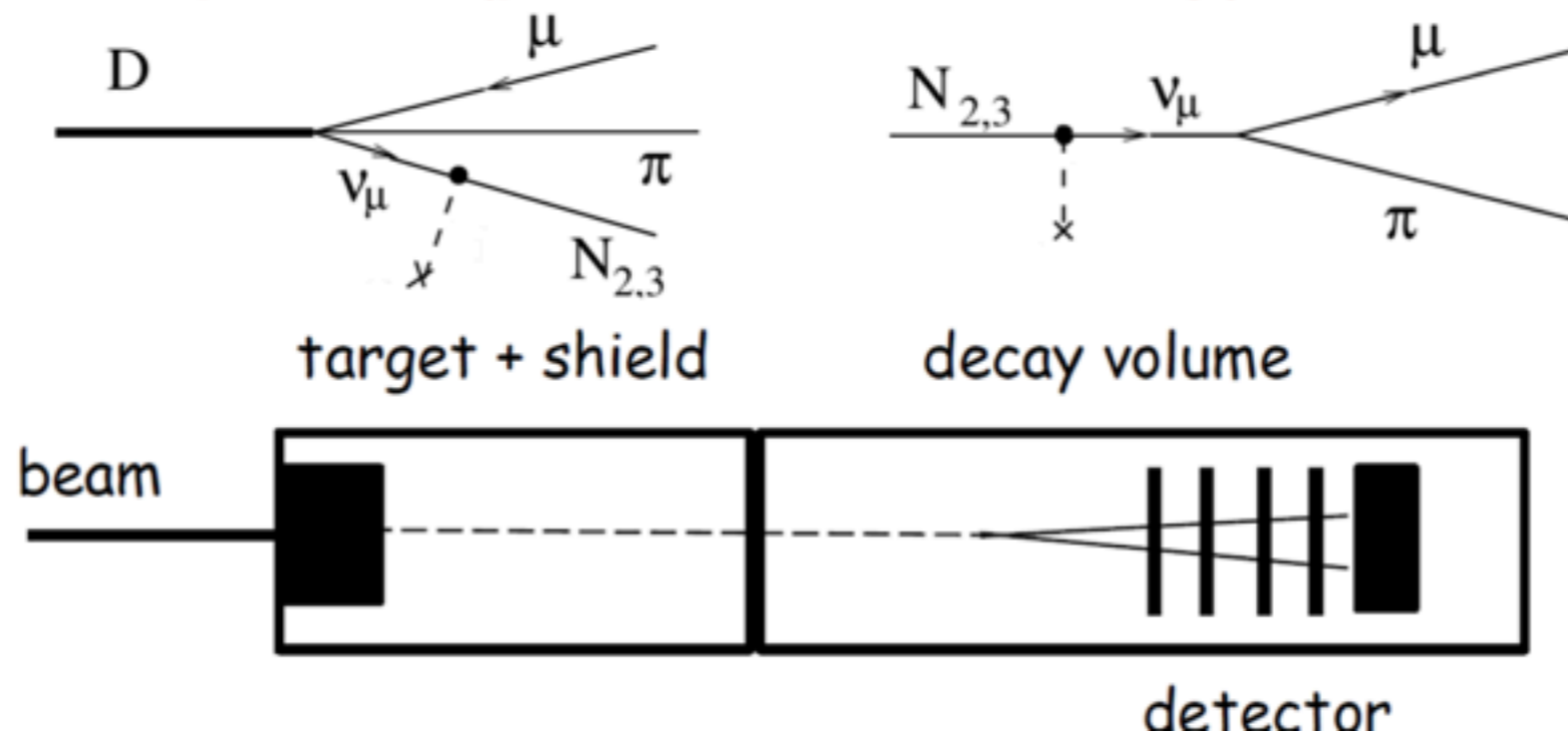
- masses up to 0.45 GeV probed through pion and kaon decays

- **PS191** *Phys. Lett. B* 203, 332 (1988)
- **NA62** in beam mode (production)



- masses up to 2 GeV probed through charmed meson decays

- **CHARM** *Phys. Lett. B* 166, 473 (1986)
- **NuTeV** *Phys. Rev. Lett.* 83, 4943 (1999)
- **NA62** in dump mode (production and decay)
- **SHiP**



PROTON BEAM DUMPS: THE PAST

Experiment	Location	approx. Date	Amount of Beam (10^{20} POT)	Beam Energy (GeV)	Target Mat.	Ref.
CHARM	CERN	1983	0.024	400	Cu	[16]
PS191	CERN	1984	0.086	19.2	Be	[17, 18]
E605	Fermilab	1986	4×10^{-7}	800	Cu	[19]
SINDRUM	SIN, PSI					
ν -Cal I	IHEP Serpukhov	1989	0.0171	70	Fe	[20–22]
LSND	LANSCCE	1994-1995	813	0.798	H ₂ O, Cu	[23]
		1996-1998	882		W, Cu	
NOMAD	CERN	1996-1998	0.41	450	Be	[18, 24]
WASA	COSY	2010		0.550	LH ₂	[25]
HADES	GSI	2011	0.32 pA*t	3.5	LH ₂ , No, Ar+KCl	[26]
		2003-2008	6.27		Be	[27]
MiniBooNE	Fermilab	2005-2012	11.3	8.9	Be	[28]
		2013-2014	1.86		Steel	[29]

+ DONUT

FNAL

3.6×10^{-3}

800

W

PROTON BEAM DUMP: (hopefully) THE FUTURE!

400GeV p, 2×10^{20} pot/5 year (already reached yearly rate with LNGS)

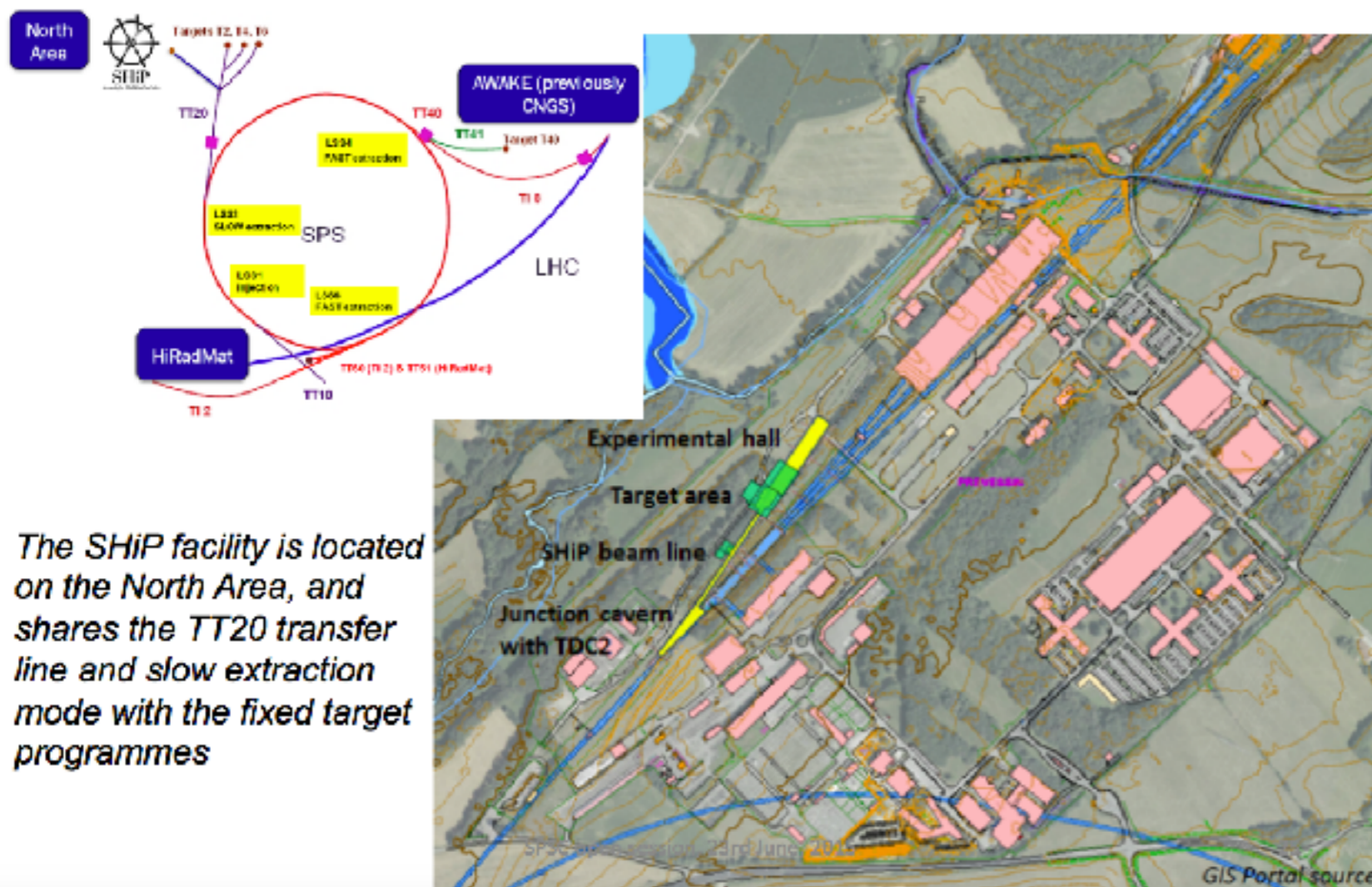


Figure of merit:

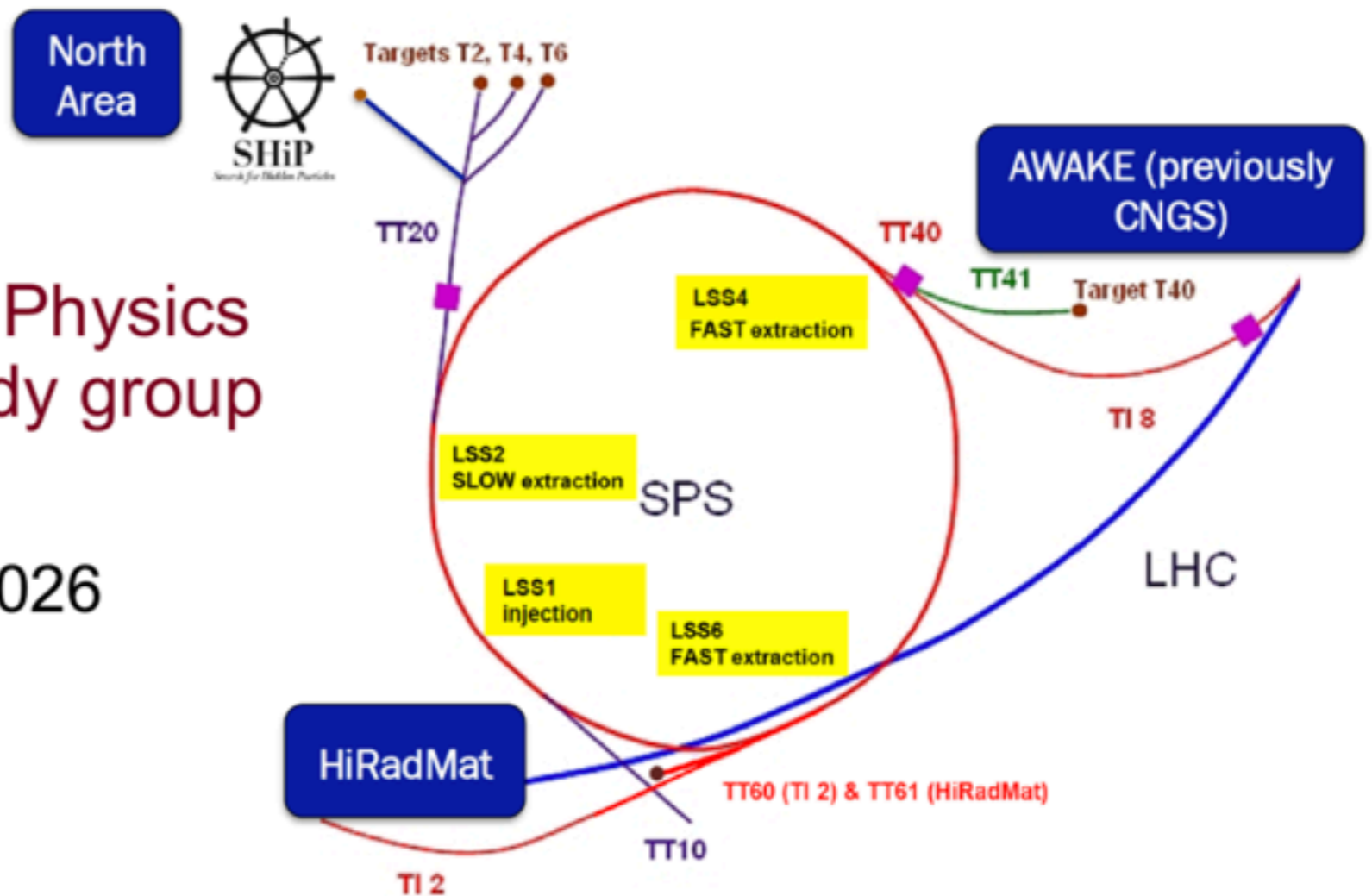
$\#(\nu_\tau)_{SHiP/DONUT}=600$

$\#(HNL)_{SHiP/CHARM}=10k$

Search for Hidden Particles (SHiP)

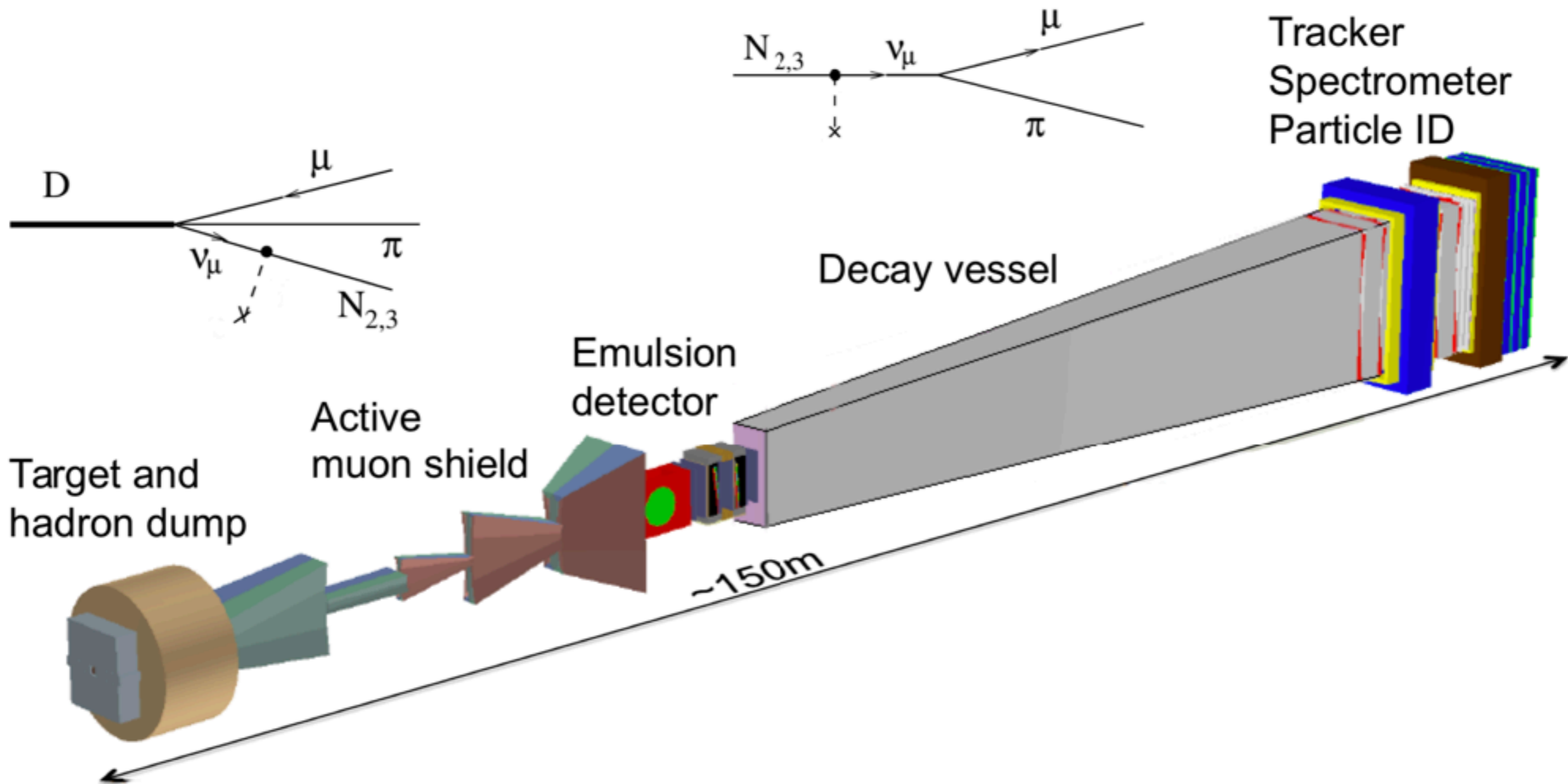
- Proposed facility: 400 GeV protons from the CERN SPS
 - New beam line and target complex
 - Aim at $2 \cdot 10^{20}$ pot in 5 years ($\rightarrow \sim 5 \cdot 10^{16}$ vs from charm decays)
- Collaboration of 250 members from 46 institutes
- Technical proposal [arXiv:1504.04956](https://arxiv.org/abs/1504.04956) (2015) Physics paper [Rep. Prog. Phys. 79](#) (2016)

- Major actor in CERN Physics Beyond Colliders study group
 - Approval ~ 2020
 - Physics runs ~ 2026



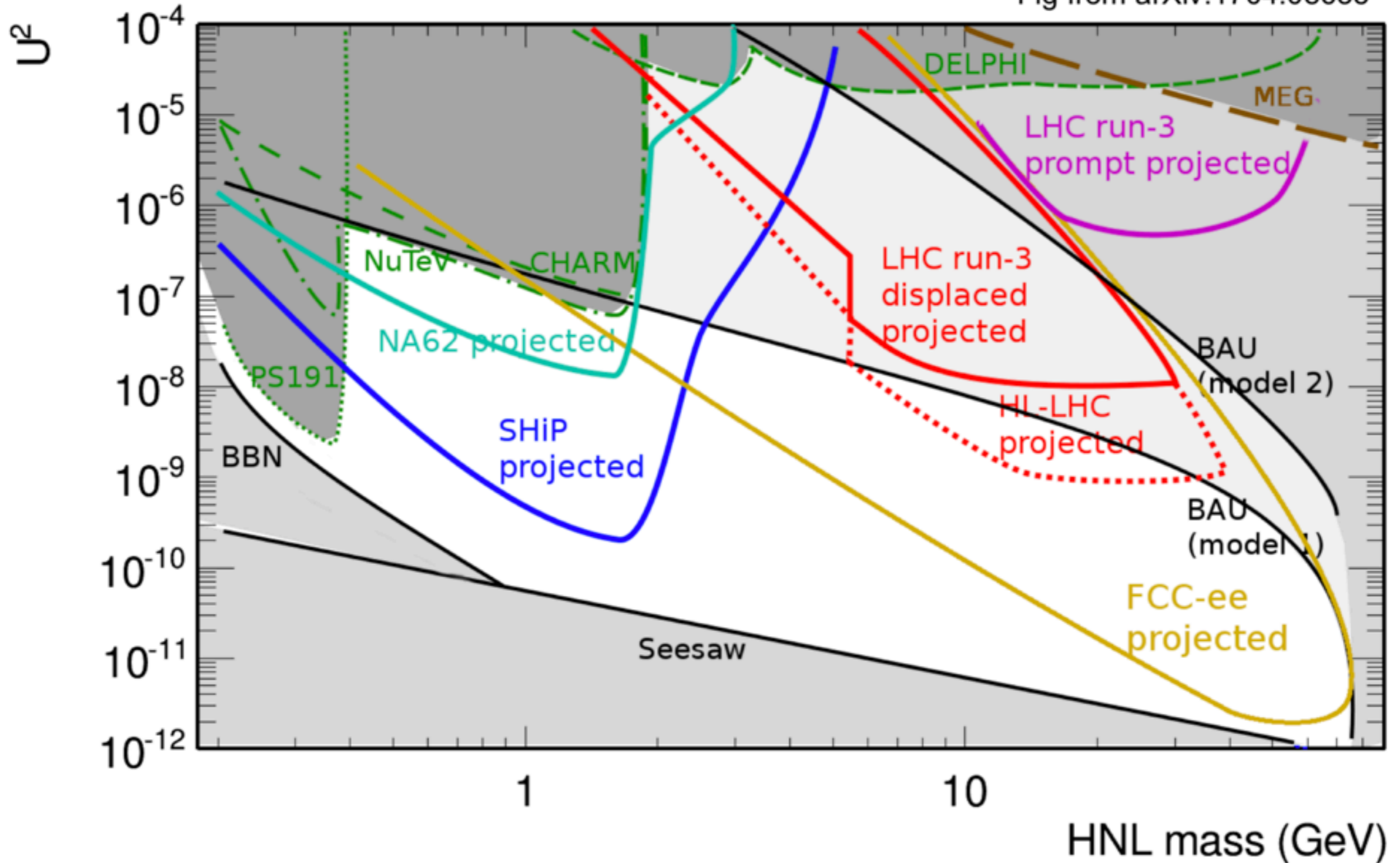
SHiP – detector

Designed for large acceptance and zero backgrounds



N at CERN in a 10-year timesecale ... and beyond

Fig from arXiv:1704.08635



The Hidden Sector



Leading SM coupling to Neutral Hidden Sector

Portals

Scalar
 $\mathcal{O}_s H^\dagger H$

Right-Handed neutrino
 $LH N_R$

U(1)
 $B_{\mu\nu} V^{\mu\nu}$

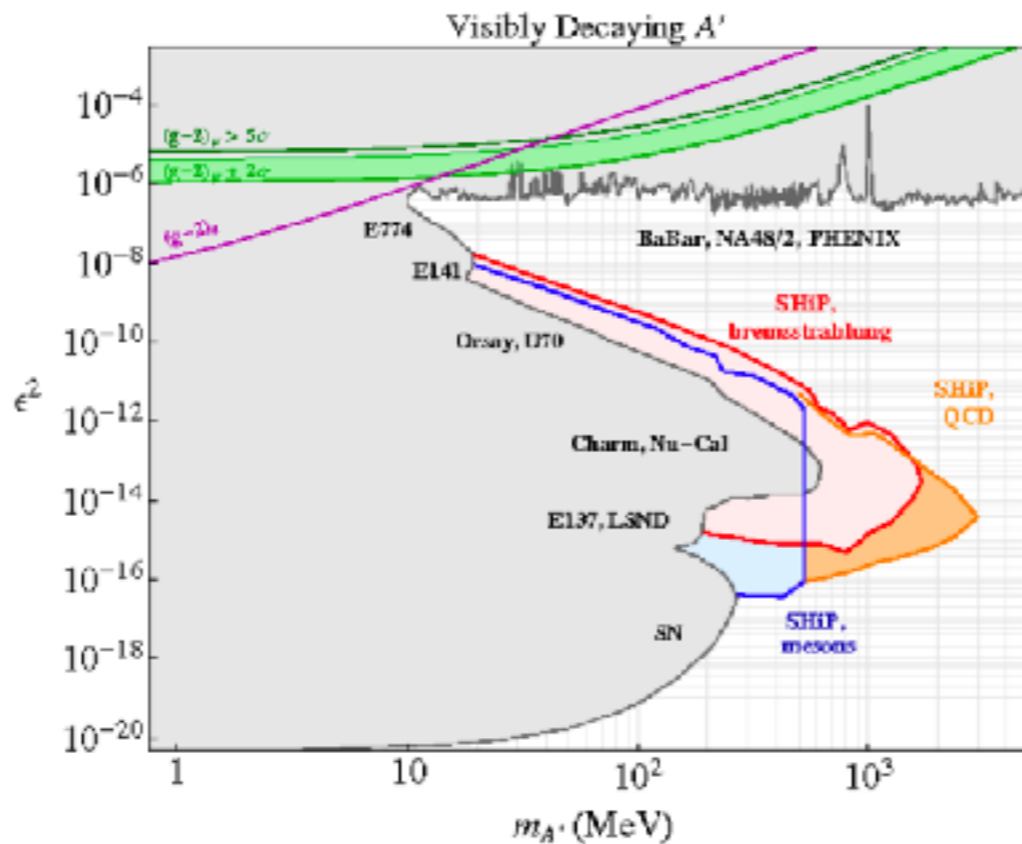
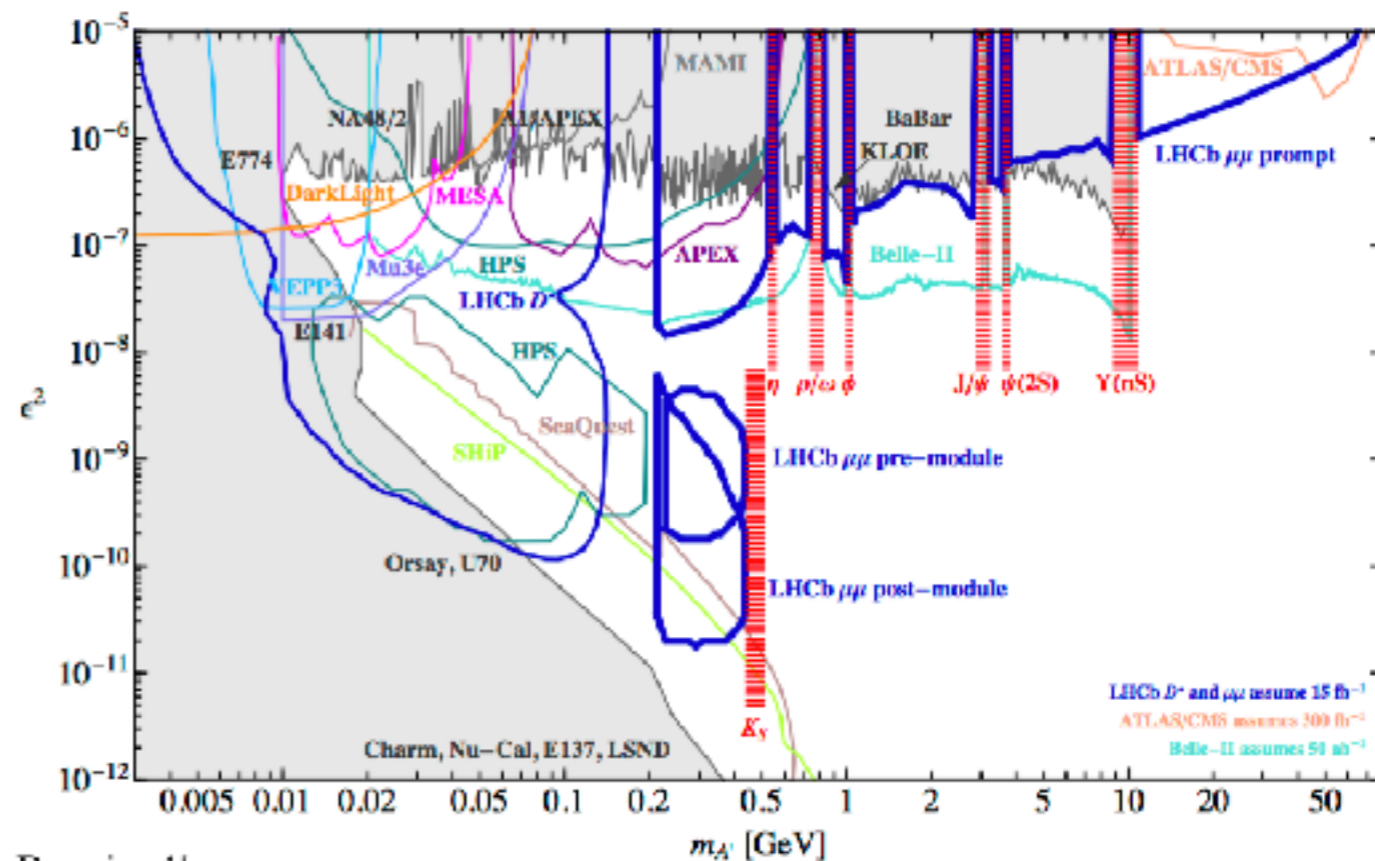
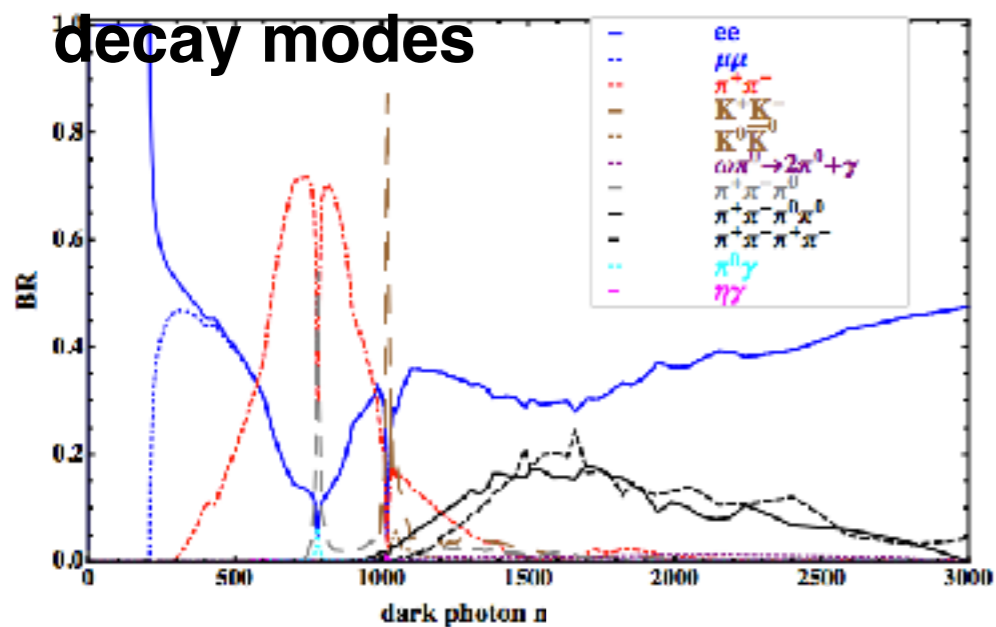
renormalizable couplings, i.e. NOT suppressed!

+other of higher dimensions (e.g. axion-like portal)

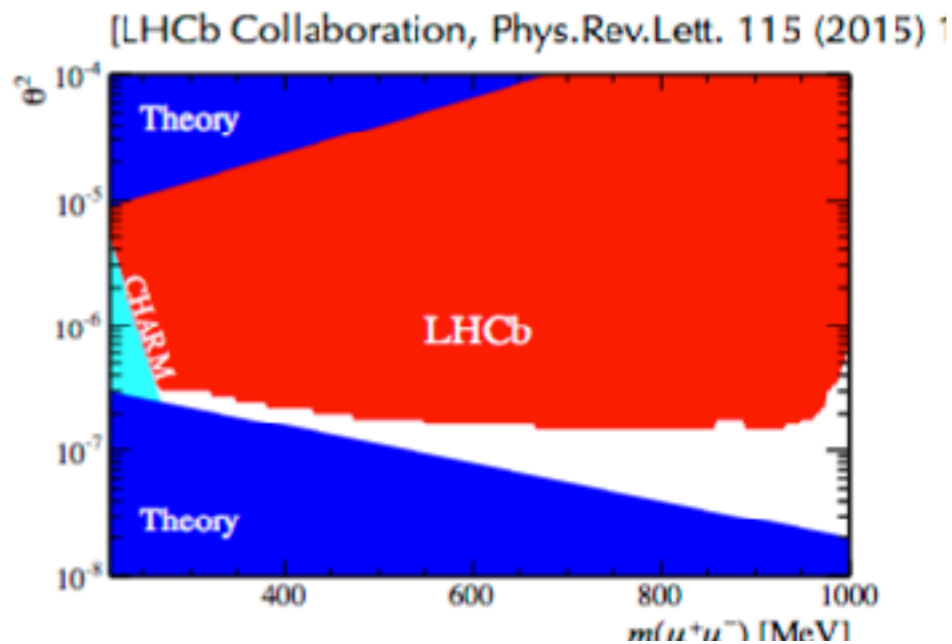
(stolen from A.Fradette, New Physics at the Intensity Frontier - Victoria, BC, Sept 2014)

Dark photon coupling to SM particles

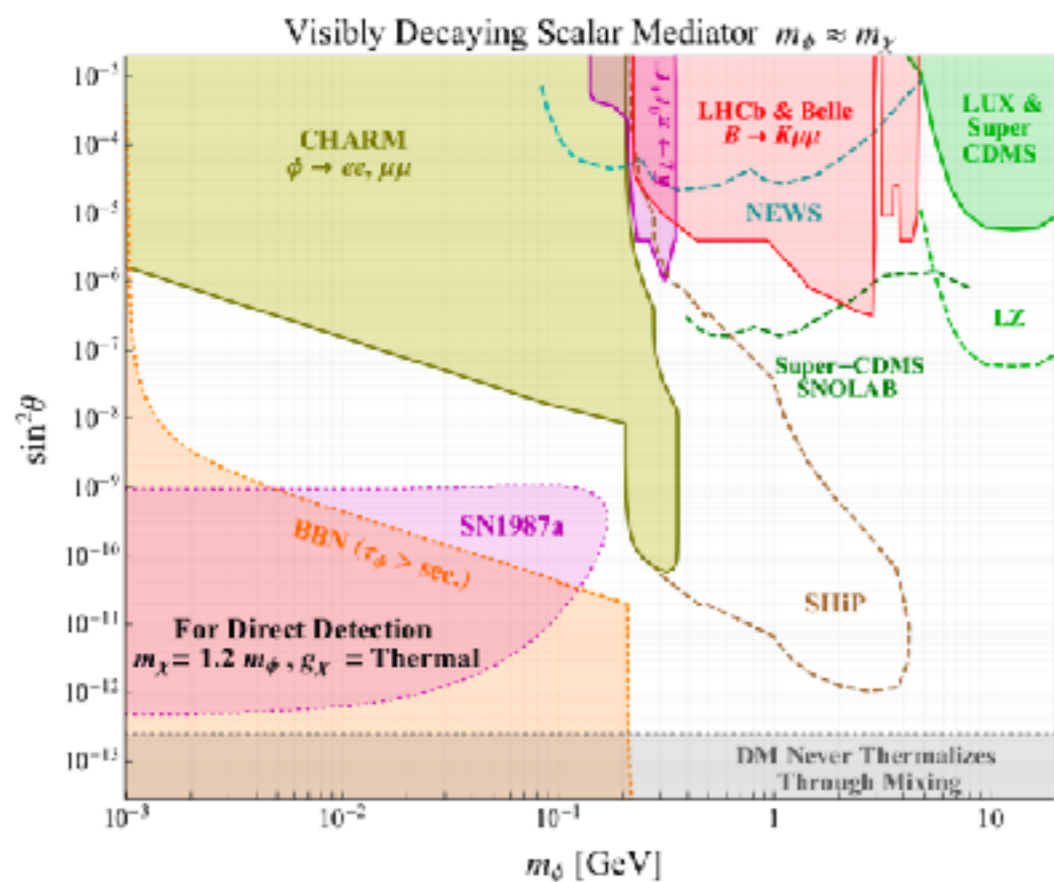
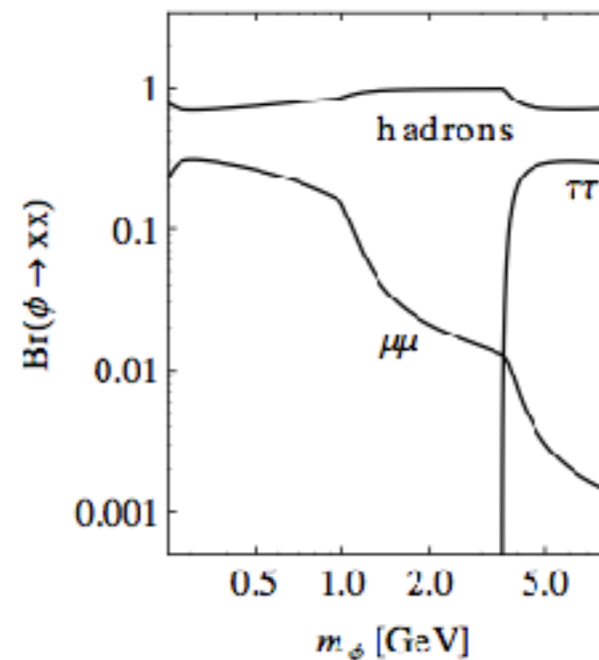
arXiv:1509.06765



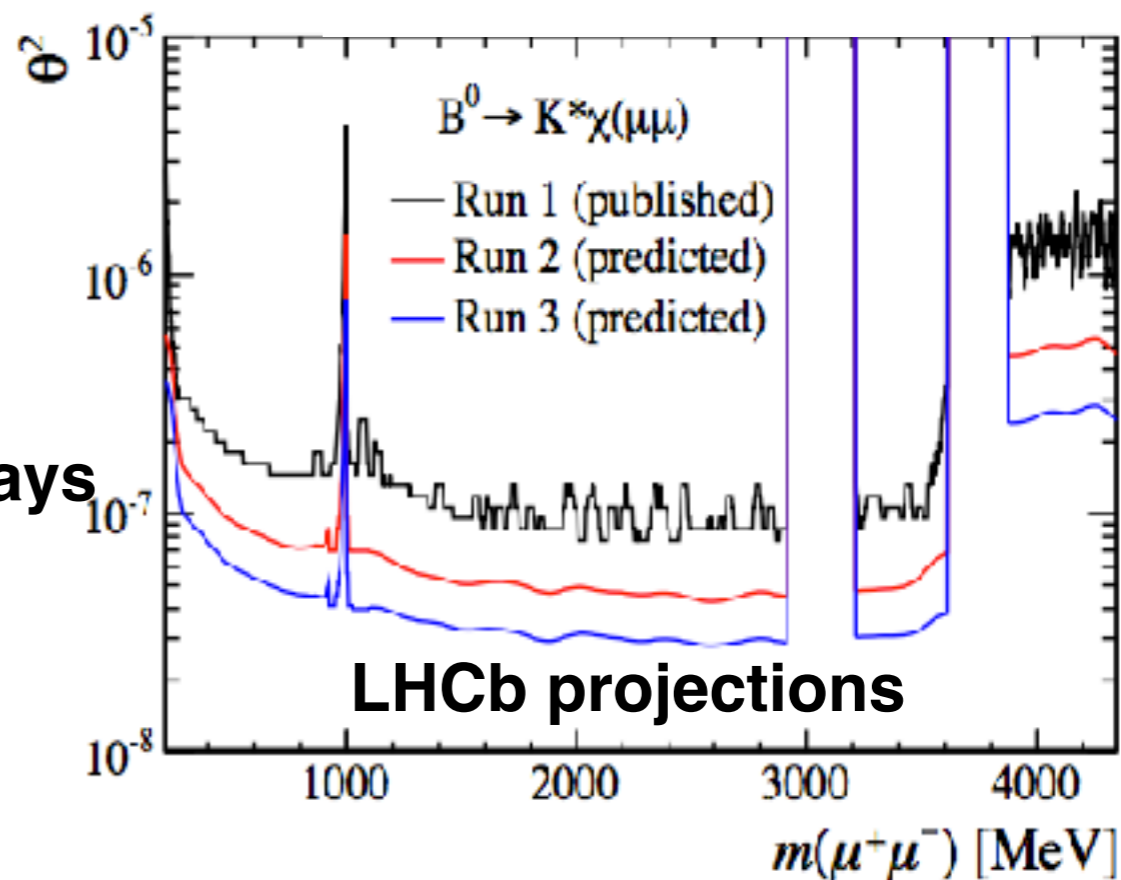
Dark Higgs



decay modes



B decays



complementary!